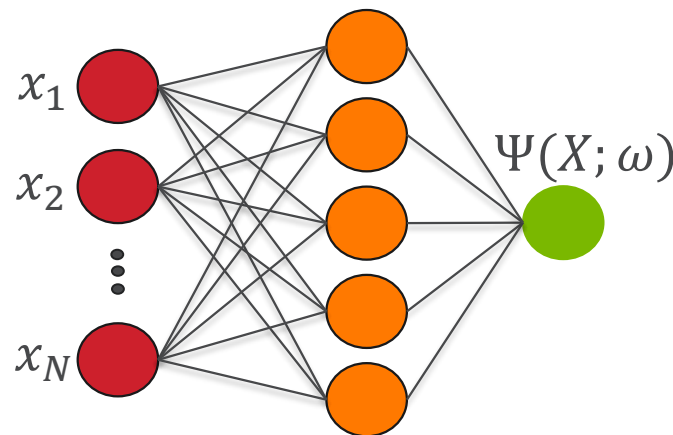


MODELING LOW-DENSITY NUCLEAR MATTER WITH NEURAL QUANTUM STATES

BRYCE FORE



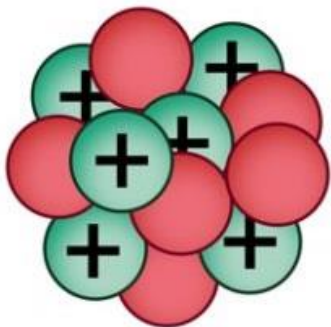
QUANTUM MANY-BODY METHODS

THE NUCLEAR MANY-BODY PROBLEM

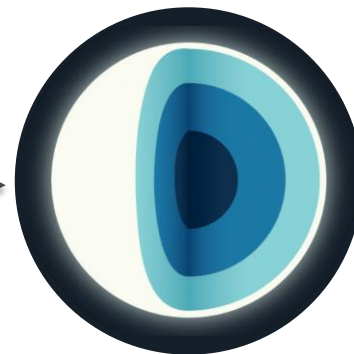
Many-body Schrödinger equation

$$\left(-\sum_i \frac{\nabla_i^2}{2m_N} + V \right) |\Psi_0\rangle = E_0 |\Psi_0\rangle$$

Nuclei



Neutron stars

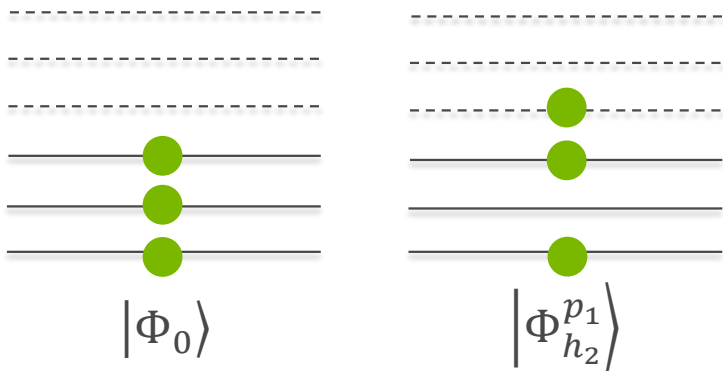


THE NUCLEAR MANY BODY METHODS

Configuration-interaction

$$|\Psi_0\rangle = \sum_{h_1, \dots, p_1, \dots} c_{h_1 \dots}^{p_1 \dots} |\Phi_{h_1 \dots}^{p_1 \dots}\rangle$$

$$|\Phi_{h_1 \dots}^{p_1 \dots}\rangle = a_{p_1}^\dagger \dots a_{h_1} \dots |\Phi_0\rangle$$



Quantum Monte Carlo

$$|\Psi_V\rangle = \sum_n c_n |\Psi_n\rangle$$

$$H|\Psi_n\rangle = E_n |\Psi_n\rangle$$

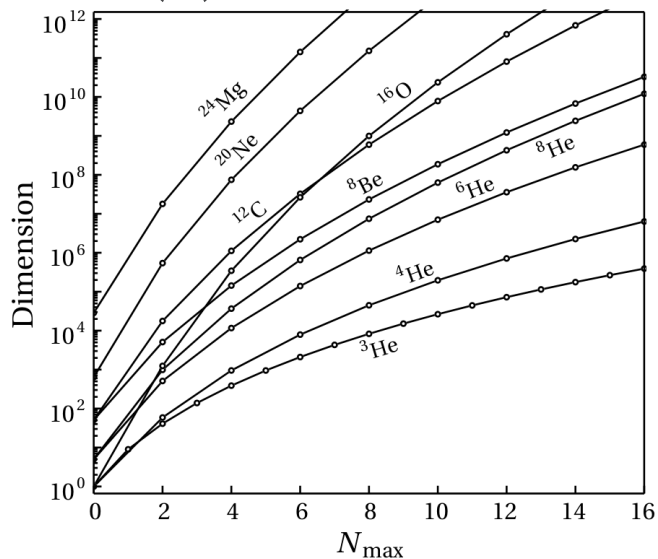
$$\lim_{\tau \rightarrow \infty} e^{-(H-E_0)\tau} |\Psi_V\rangle = c_0 |\Psi_0\rangle$$

CURSE OF DIMENSIONALITY

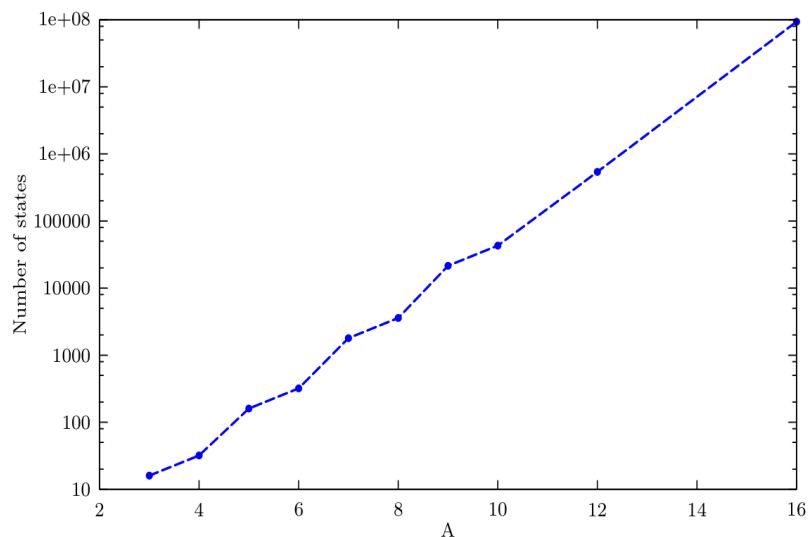
Configuration-interaction

Green's function Monte Carlo

$$\binom{N}{A} = \frac{N!}{(N-A)! A!}$$



$$\lim_{\tau \rightarrow \infty} e^{-(H-E_0)\tau} |\Psi_V\rangle = c_0 |\Psi_0\rangle$$

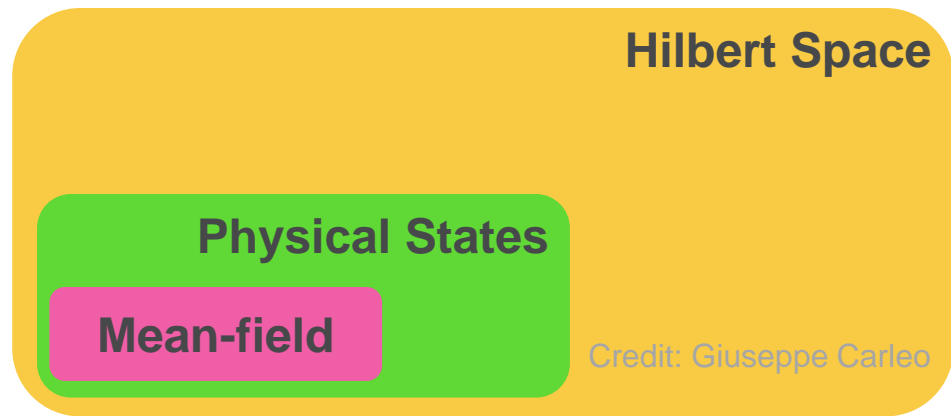
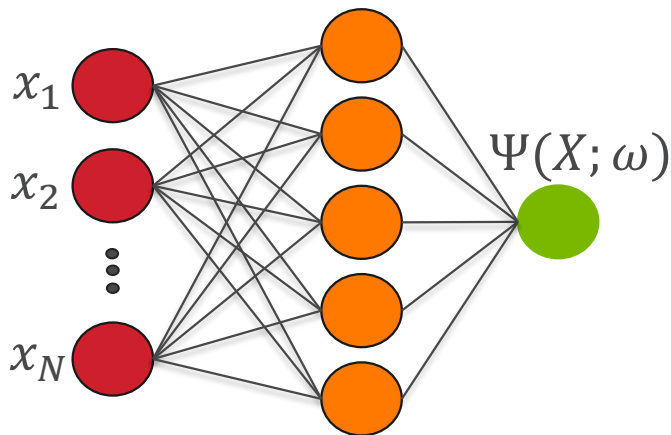


Credit: Patrick Fasano

Credit: Alessandro Lovato

NEURAL NETWORK QUANTUM STATES

- Artificial neural networks compactly represent complex high-dimensional functions
- Most quantum states of interest have distinctive features and intrinsic structures



Credit: Giuseppe Carleo

SCALING AND COMPUTATIONAL PERFORMANCE

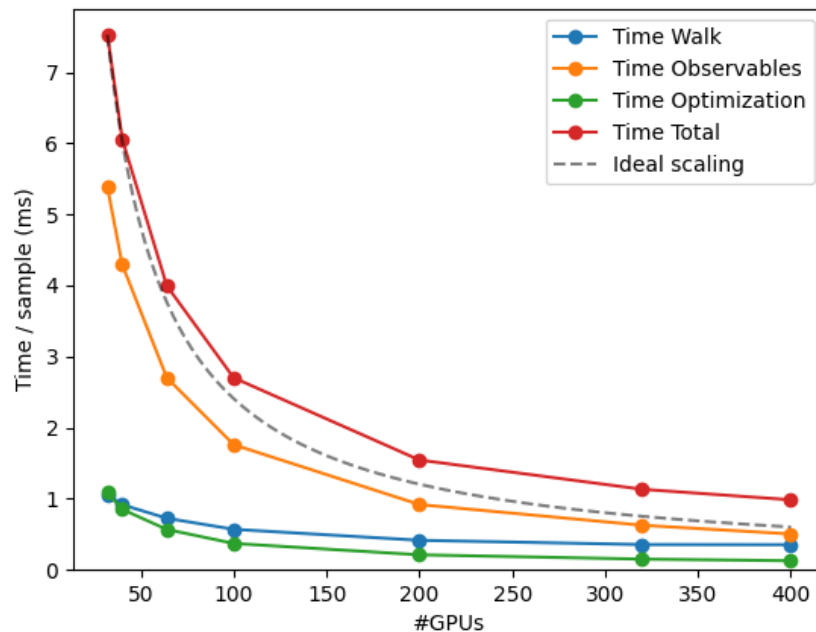
Scaling with system size

Conventional QMC: $O(2^A)$

Neural quantum states: $O(A^5)$

A = Number of particles in system

Scaling with resources



VARIATIONAL MONTE CARLO WITH NQS

1. Create the NQS wavefunction

$$\Psi_V(R, S; \omega) = e^{U(R, S; \omega)} \Phi(R, S; \omega)$$

2. Compute its energy (Metropolis-Hastings algorithm)

$$\frac{\langle \Psi_V | O | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} \approx \frac{1}{N_{conf}} \sum_{\{R, S\}} O_L(R, S)$$

3. Minimize the energy w.r.t parameters, ω , to reach the ground state

$$E_0 \leq E_V = \frac{\langle \Psi_V | \hat{H} | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle}$$

FERMIONIC WAVEFUNCTIONS

$$\Psi_V(X) = e^{U(X)} \Phi(X)$$

$$\Psi(\dots, x_i, \dots x_j \dots) = -\Psi(\dots, x_j, \dots x_i \dots)$$

$$U(\dots, x_i, \dots x_j \dots) = U(\dots, x_j, \dots x_i \dots)$$

$$\Phi(\dots, x_i, \dots x_j \dots) = -\Phi(\dots, x_j, \dots x_i \dots)$$

- Build in fermion antisymmetry for network compactness
- Permutation-invariant Jastrow function improves ansatz flexibility
- Build U and Φ functions from fully connected, deep neural networks

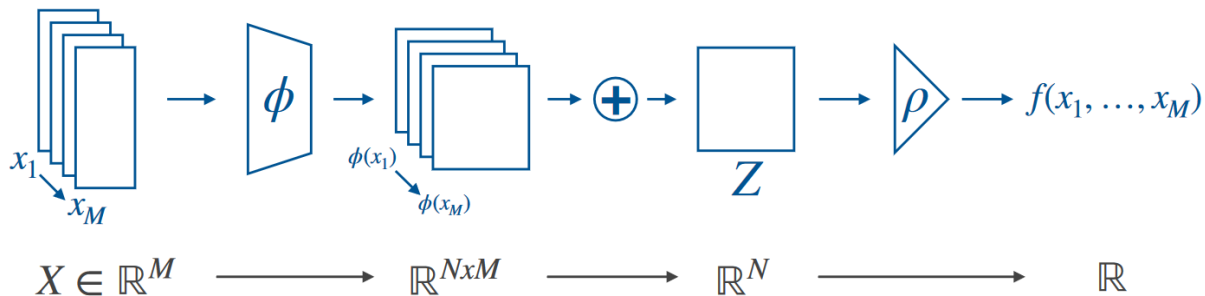
DEEP SET ARCHITECTURE

- Generic permutation invariant function

$$U(X) = \rho \left(\sum_i \vec{\phi}(x_i) \right)$$

$$\vec{\phi}: \mathbb{R}^5 \rightarrow \mathbb{R}^N$$

$$\rho: \mathbb{R}^N \rightarrow \mathbb{R}$$



Wagstaff et al., arXiv:1901.09006 (2019)

Zaheer et al., arXiv:1703.06114 (2017)

NEURAL SLATER-JASTROW ANSATZ

- Slater determinants used in other methods to represent many-body states
- Slater determinant to enforce antisymmetry
- Single particle wavefunctions represented by neural networks

$$\Psi_V(X) = e^{U(X)} \Phi(X)$$

$$\Phi(X) = \begin{vmatrix} \phi_1(x_1) & \phi_1(x_2) & \dots & \phi_1(x_N) \\ \phi_2(x_1) & & & \vdots \\ \vdots & & & \\ \phi_n(x_1) & \dots & & \phi_n(x_N) \end{vmatrix}$$

NEURAL PFAFFIAN ANSATZ

- Slater determinant \rightarrow Pfaffian
- Build in antisymmetry through a single pairing network rather than determinant of single-particle states

$$\det(M) = pf(M)^2$$

$$\phi(x_i, x_j) = \eta(x_i, x_j) - \eta(x_j, x_i)$$

$$\Phi(X) = pf \begin{bmatrix} 0 & \phi(x_1, x_2) & \cdots & \phi(x_1, x_N) \\ \phi(x_2, x_1) & 0 & \cdots & \phi(x_2, x_N) \\ \vdots & \vdots & \ddots & \vdots \\ \phi(x_N, x_1) & \phi(x_N, x_2) & \cdots & 0 \end{bmatrix}$$

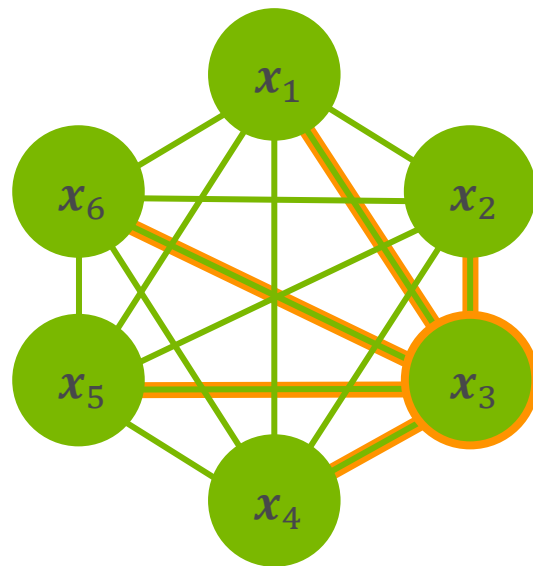
J. Kim, *Commun Phys* **7**, 148 (2024)

MESSAGE PASSING NEURAL NETWORK

Backflow transformation representable by updates to fully connected graph

$$x_i \rightarrow f(x_i; \{x_j\}_{j \neq i})$$

- Update edge values based on vertices
- Update vertex values based on connected edges
- Retains ordering information for antisymmetry operation



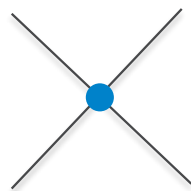
NQS RESULTS

PIONLESS EFT HAMILTONIAN

- Pionless-EFT Hamiltonian

$$H_{LO} = - \sum_i \frac{\vec{\nabla}_i^2}{2m_N} + \sum_{i<j} v_{ij} + \sum_{i<j<k} V_{ijk}$$

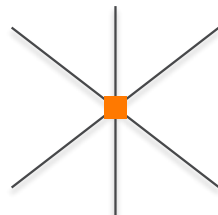
- NN potential fit to
 - np scattering lengths
 - effective radii
 - deuteron binding energy



$$v_{ij}^{\text{CI}} = \sum_{p=1}^4 v^p(r_{ij}) O_{ij}^p$$

$$O_{ij}^{p=1,4} = (1, \tau_{ij}, \sigma_{ij}, \sigma_{ij}\tau_{ij})$$

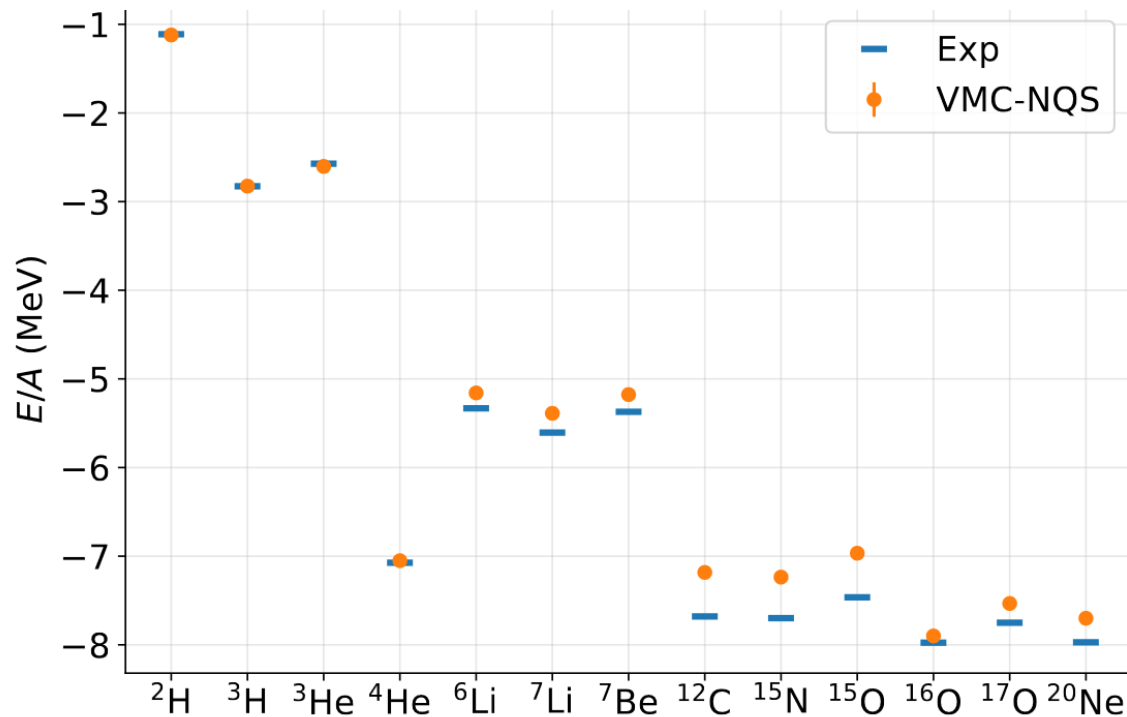
- Three body potential adjusted to reproduce ^3H binding



$$V_{ijk} = \tilde{c}_E \sum_{\text{cyc}} e^{-(r_{ij}^2 + r_{jk}^2)/R_3^2}$$

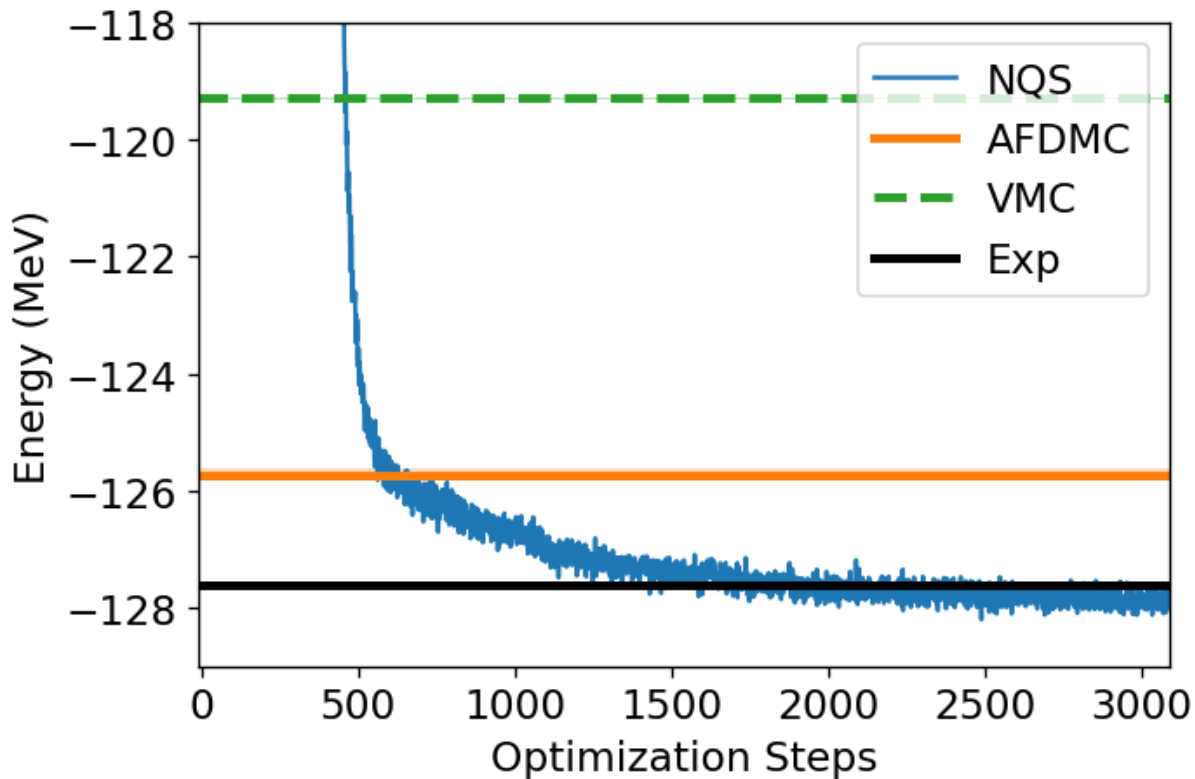
R. Schiavilla, PRC 103, 054003(2021)

NEURAL QUANTUM STATE RESULTS IN NUCLEI

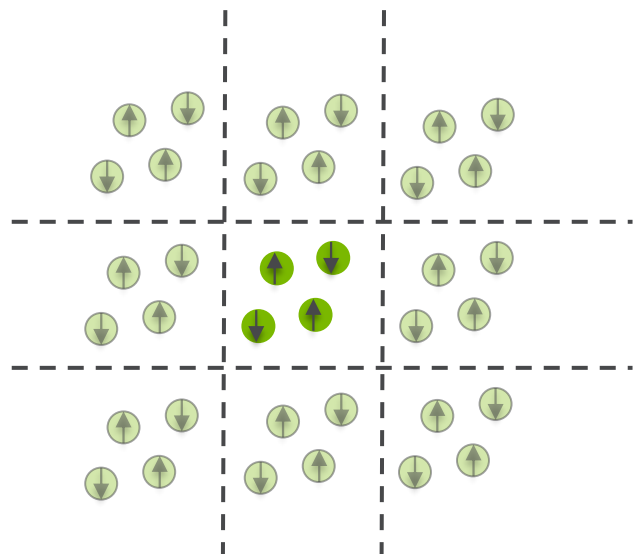


A. Gnech, Phys. Rev. Lett. 133, 142501

RESULTS IN 160



BULK NUCLEAR MATTER SETUP



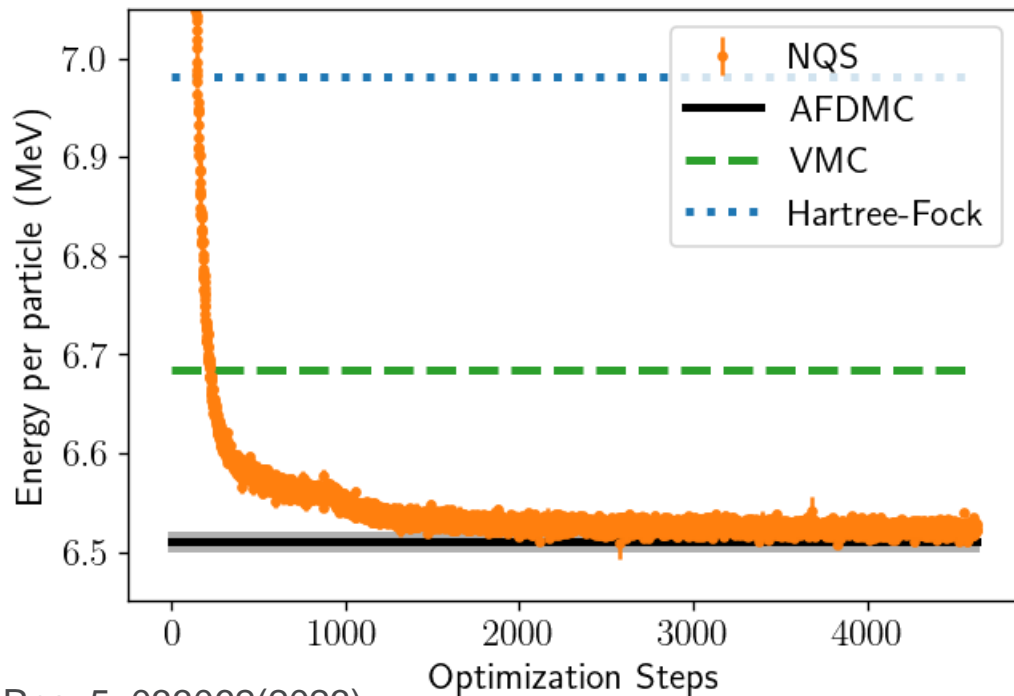
- Periodic boundary conditions and coordinate system

$$\mathbf{r}_i \longrightarrow \tilde{\mathbf{r}}_i = \left\{ \sin \left(\frac{2\pi}{L} \mathbf{r}_i \right), \cos \left(\frac{2\pi}{L} \mathbf{r}_i \right) \right\}$$

- Potential energy contribution from particle images
- Remove Coulomb potential

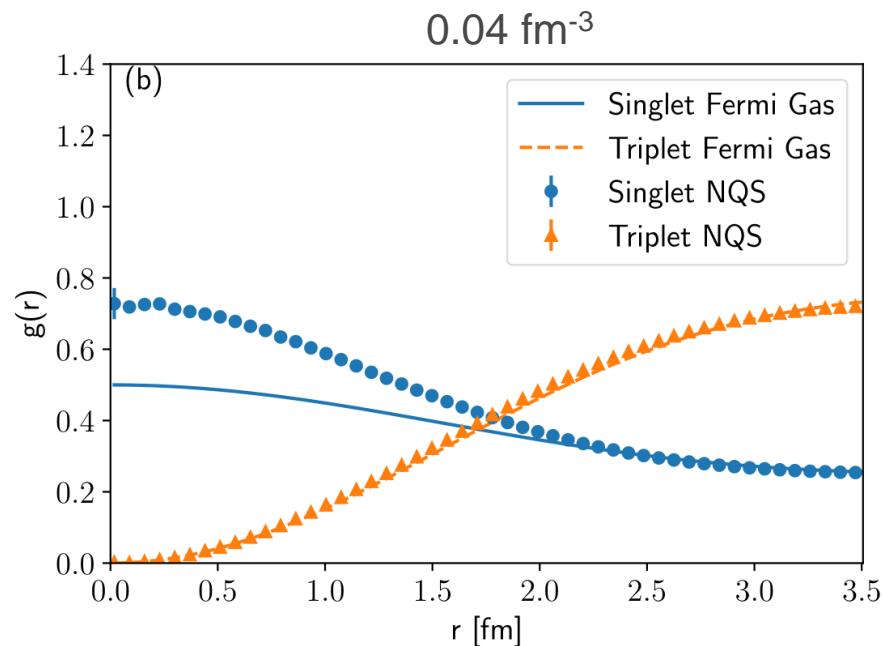
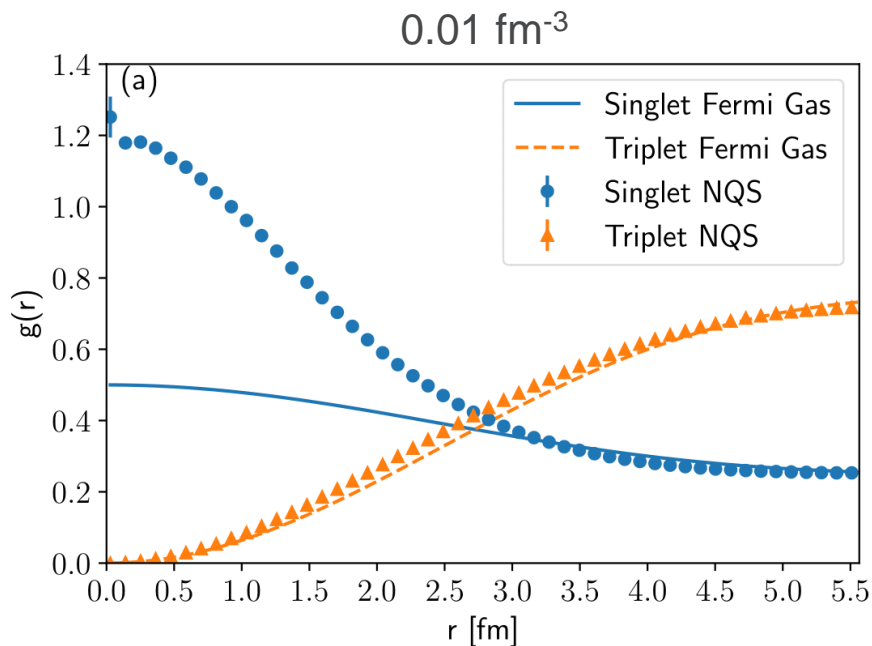
PURE NEUTRON MATTER

Neutron Matter (14 particles, 0.04 fm^{-3})



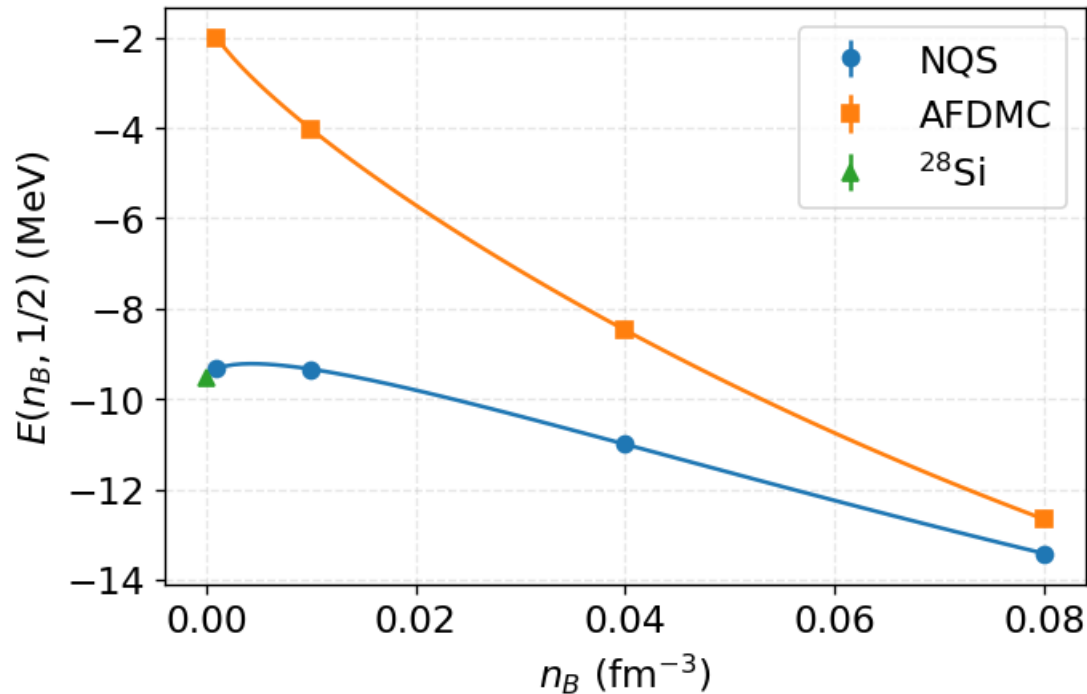
B. Fore, Phys. Rev. Res. 5, 033062(2023)

TWO-BODY PAIR DISTRIBUTIONS



B. Fore, Phys. Rev. Res. 5, 033062(2023)

SYMMETRIC NUCLEAR MATTER

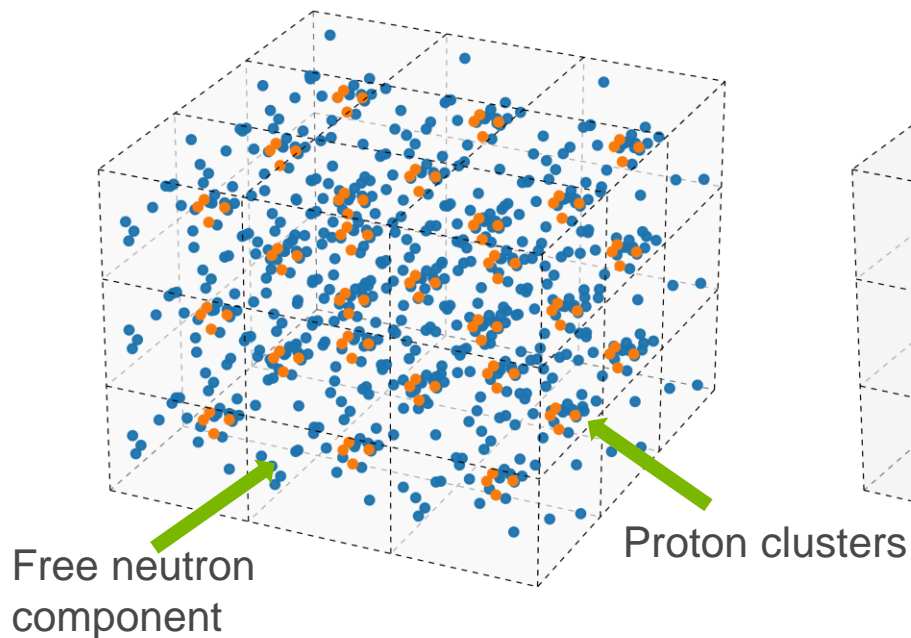


B. Fore, arXiv:2407.21207

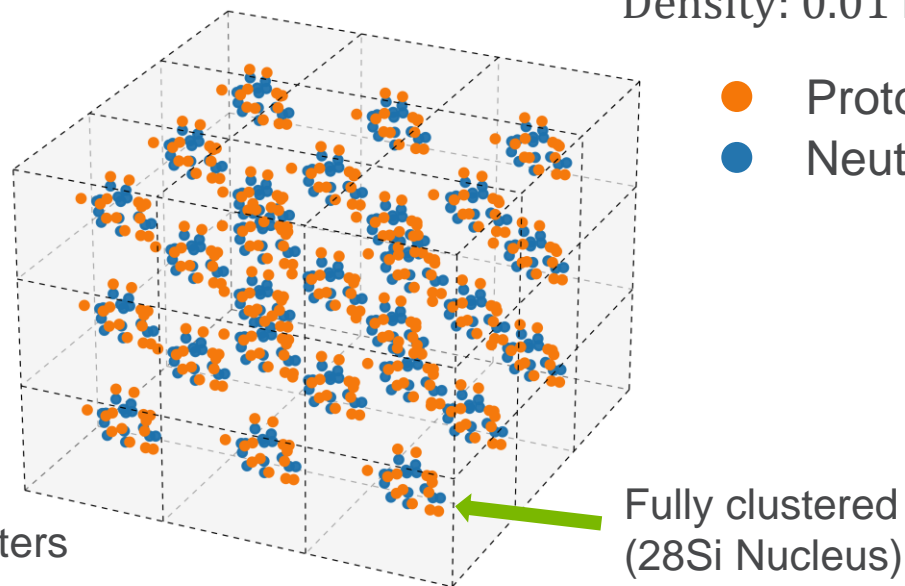
SELF EMERGING CLUSTERING

Density: 0.01 fm^{-3}

- Proton
- Neutron



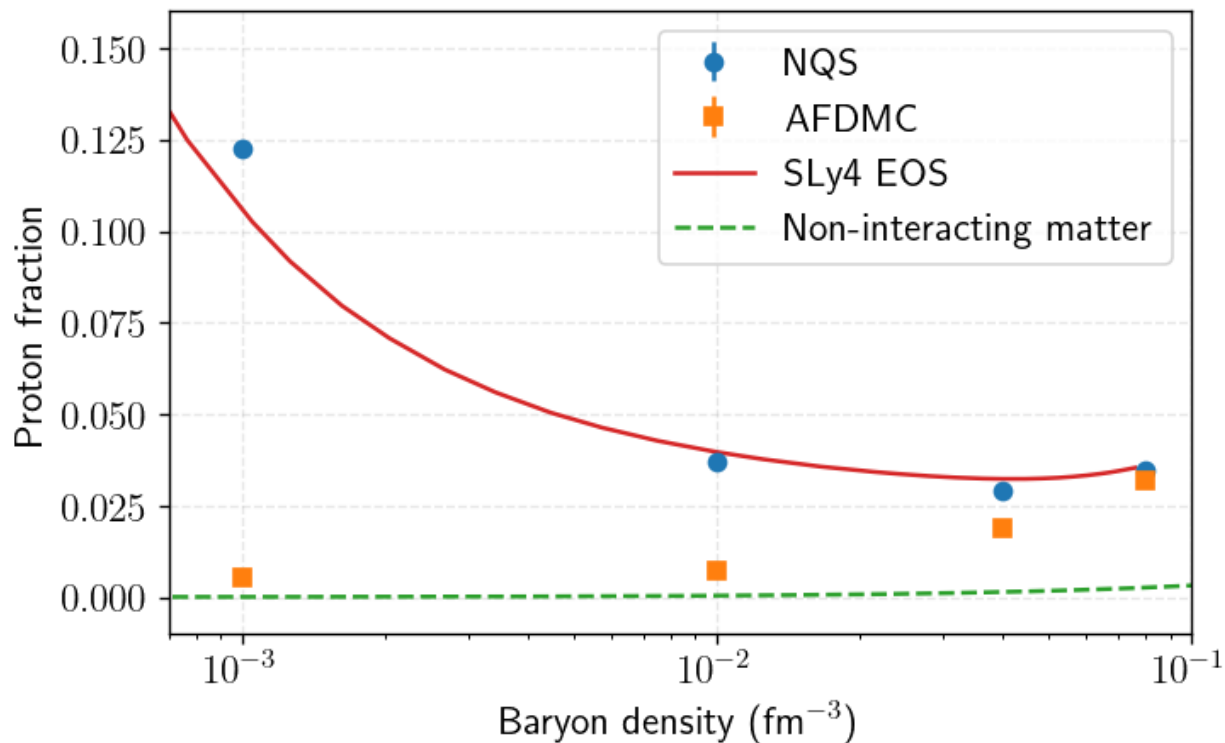
4 protons; 24 neutrons



14 protons; 14 neutrons

B. Fore, arXiv:2407.21207

NUCLEAR MATTER PROTON FRACTION



Assumptions:

- Charge neutrality
 $n_p = n_e$
- Beta equilibrium
 $\mu_e = \mu_n - \mu_p$

B. Fore, arXiv:2407.21207

CONCLUSIONS AND NEXT STEPS

- Conclusions:
 - Favorable scaling with number of fermions
 - Universal and accurate approximations for fermion wavefunctions
 - Scaling to leadership-class computers
 - NQS can model a variety of phases of nuclear matter material
- Next steps:
 - Improved nuclear potential including tensor term
 - Expanding to larger nuclei and larger periodic systems to avoid finite size effects
 - Excited state calculations

COLLABORATORS



A. Lovato, A. Tropiano



J. Kim



M. Hjorth-Jensen

THANK YOU

METROPOLIS-HASTINGS SAMPLING

Sampling algorithm:

- Randomly sample coordinates, R' , and spins, S'

$$P_R = \frac{|\Psi_V(R', S)|^2}{|\Psi_V(R, S)|^2} \quad P_S = \frac{|\Psi_V(R, S')|^2}{|\Psi_V(R, S)|^2}$$

- If P is greater than uniform random variable from 0 to 1, accept new values
- Observables are estimated by taking averages over sampled configurations

$$\frac{\langle \Psi_V | O | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} = \frac{\sum_S \int dR |\Psi_V(R, S)|^2 O_L(R, S)}{\sum_S \int dR |\Psi_V(R, S)|^2} \approx \frac{1}{N_{conf}} \sum_{\{R, S\}} O_L(R, S)$$

$$O_L = \frac{\langle RS | O | \Psi_V \rangle}{\langle RS | \Psi_V \rangle}$$

NQS OPTIMIZATION

Improve wavefunction by minimizing energy expectation value

$$E_0 \leq E_V = \frac{\langle \Psi_V | \hat{H} | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle}$$

Gradient of energy ($G_i = \frac{dE_V}{d\omega_i}$)

$$G_i = 2 \left(\frac{\langle \partial_i \Psi_V | \hat{H} | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} - E_V \frac{\langle \partial_i \Psi_V | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} \right)$$

Parameters at step s are updated as

$$\omega^{s+1} = \omega^s - \eta G$$

STOCHASTIC RECONFIGURATION

Improve wavefunction by minimizing energy expectation value

$$E_0 \leq E_V = \frac{\langle \Psi_V | \hat{H} | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle}$$

Gradient of energy ($G_i = \frac{dE_V}{d\omega_i}$), supplemented by Quantum Fisher Information S_{ij}

$$G_i = 2 \left(\frac{\langle \partial_i \Psi_V | \hat{H} | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} - E_V \frac{\langle \partial_i \Psi_V | \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} \right); \quad S_{ij} = \frac{\langle \partial_i \Psi_V | \partial_j \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle} - \frac{\langle \partial_i \Psi_V | \Psi_V \rangle \langle \Psi_V | \partial_j \Psi_V \rangle}{\langle \Psi_V | \Psi_V \rangle \langle \Psi_V | \Psi_V \rangle}$$

Parameters at step s are updated as

$$\omega^{s+1} = \omega^s - \eta(S + \Lambda)^{-1}G$$

HIDDEN FERMIONS

$$\Psi_V(X) = \begin{vmatrix} \phi_1(x_1) & \phi_1(x_2) & \phi_1(x_3) & \phi_1(x_4) & \phi_1(y_1) & \phi_1(y_2) & \phi_1(y_3) & \phi_1(y_4) \\ \phi_2(x_1) & \phi_2(x_2) & \phi_2(x_3) & \phi_2(x_4) & \phi_2(y_1) & \phi_2(y_2) & \phi_2(y_3) & \phi_2(y_4) \\ \phi_3(x_1) & \phi_3(x_2) & \phi_3(x_3) & \phi_3(x_4) & \phi_3(y_1) & \phi_3(y_2) & \phi_3(y_3) & \phi_3(y_4) \\ \phi_4(x_1) & \phi_4(x_2) & \phi_4(x_3) & \phi_4(x_4) & \phi_4(y_1) & \phi_4(y_2) & \phi_4(y_3) & \phi_4(y_4) \\ \chi_1(x_1) & \chi_1(x_2) & \chi_1(x_3) & \chi_1(x_4) & \chi_1(y_1) & \chi_1(y_2) & \chi_1(y_3) & \chi_1(y_4) \\ \chi_2(x_1) & \chi_2(x_2) & \chi_2(x_3) & \chi_2(x_4) & \chi_2(y_1) & \chi_2(y_2) & \chi_2(y_3) & \chi_2(y_4) \\ \chi_3(x_1) & \chi_3(x_2) & \chi_3(x_3) & \chi_3(x_4) & \chi_3(y_1) & \chi_3(y_2) & \chi_3(y_3) & \chi_3(y_4) \\ \chi_4(x_1) & \chi_4(x_2) & \chi_4(x_3) & \chi_4(x_4) & \chi_4(y_1) & \chi_4(y_2) & \chi_4(y_3) & \chi_4(y_4) \end{vmatrix}$$

Visible wavefunctions
on visible coordinates

Visible wavefunctions
on hidden coordinates

Hidden wavefunctions
on visible coordinates

Hidden wavefunctions
on hidden coordinates

Comparison to
Neural SJ Ansatz

$$\Psi_V(X) = e^{U(X)} \Phi(X)$$

J. R. Moreno, PNAS 119 (32) e2122059119

DETERMINANT VS PFAFFIAN

Defined for $n \times n$ matrices

$$\det(A) = \sum_{\sigma \in S_n} \text{sgn}(\sigma) \prod_{i=1}^n a_{i,\sigma(i)}$$

$$\det(A^T) = \det(A)$$

$$\det(A) \det(B) = \det(AB)$$

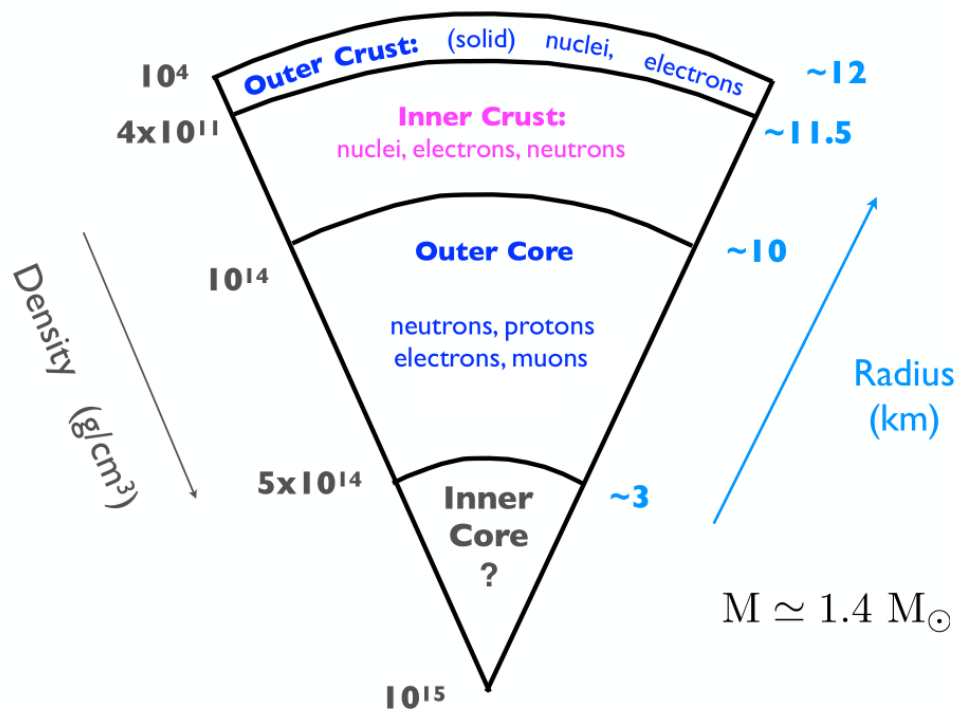
Defined for $2n \times 2n$ skew-symmetric matrices

$$\text{pf}(A) = \frac{1}{2^n n!} \sum_{\sigma \in S_{2n}} \text{sgn}(\sigma) \prod_{i=1}^n a_{\sigma(2i-1), \sigma(2i)}$$

$$\text{pf}(A^T) = (-1)^n \text{pf}(A)$$

$$\text{pf}(A) \text{pf}(B) = \exp\left(\frac{1}{2} \text{tr} \log(A^T B)\right)$$

NEUTRON STAR STRUCTURE



- Mostly neutrons but composition varies with density
- Nuclei in crust are squeezed into uniform matter in core
- Likely neutron superfluid in inner crust and outer core
- Calculations currently focus on inner crust

ASYMMETRIC MATTER ENERGY

- Fit $E(n_B, 0)$ and $E(n_B, 1/2)$

$$E(n_B, x) = a_0 + a_{2/3} \left(\frac{n_B}{n_0} \right)^{2/3} + a_1 \left(\frac{n_B}{n_0} \right) + a_2 \left(\frac{n_B}{n_0} \right)^2$$

- Model energy by symmetry energy expansion

$$S(n_B) = E(n_B, 0) - E(n_B, 1/2)$$

$$E(n_B, x) = E(n_B, 1/2) + (1 - 2x)^2 S(n_B)$$

