

Studying structure near the neutron emission threshold using the detector TETRA at ALTO

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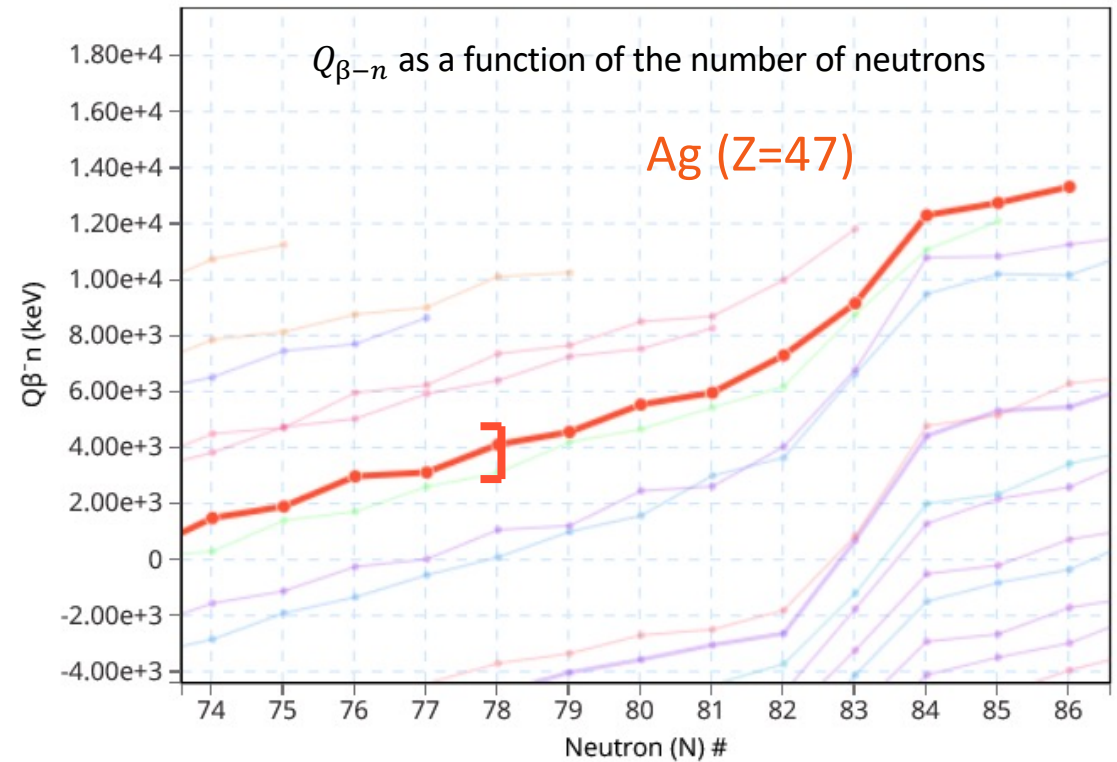
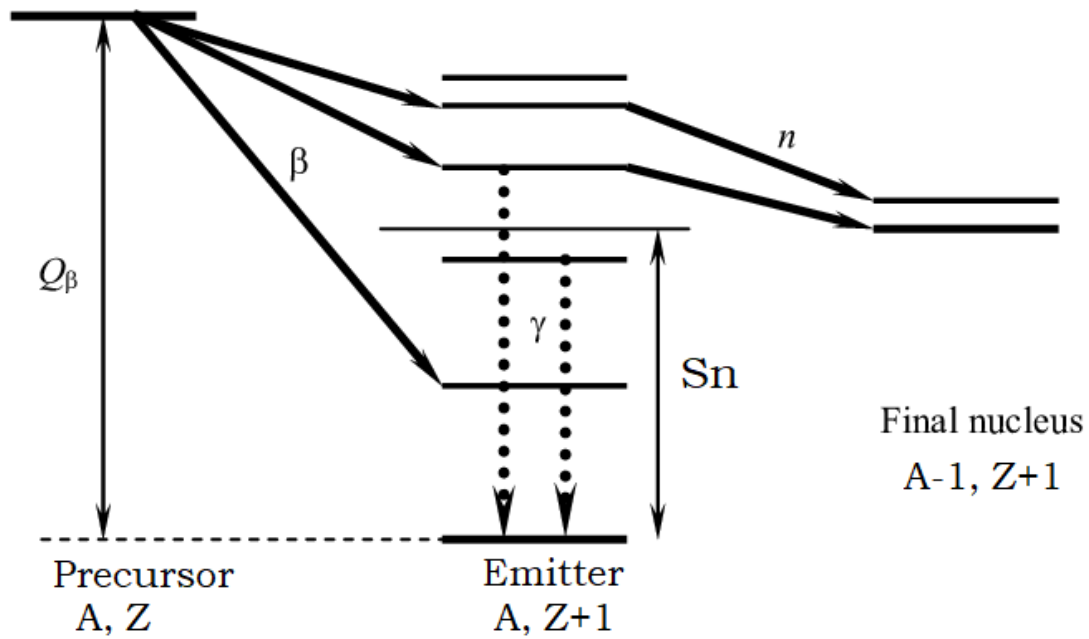
Team FIIRST



- I. Structure near the neutron emission threshold**
- II. Experimental setup**
- III. Pn measurement of ^{84}Ga**
- IV. Ring ratio method to measure mean neutron energy**



Beta-delayed neutron emission



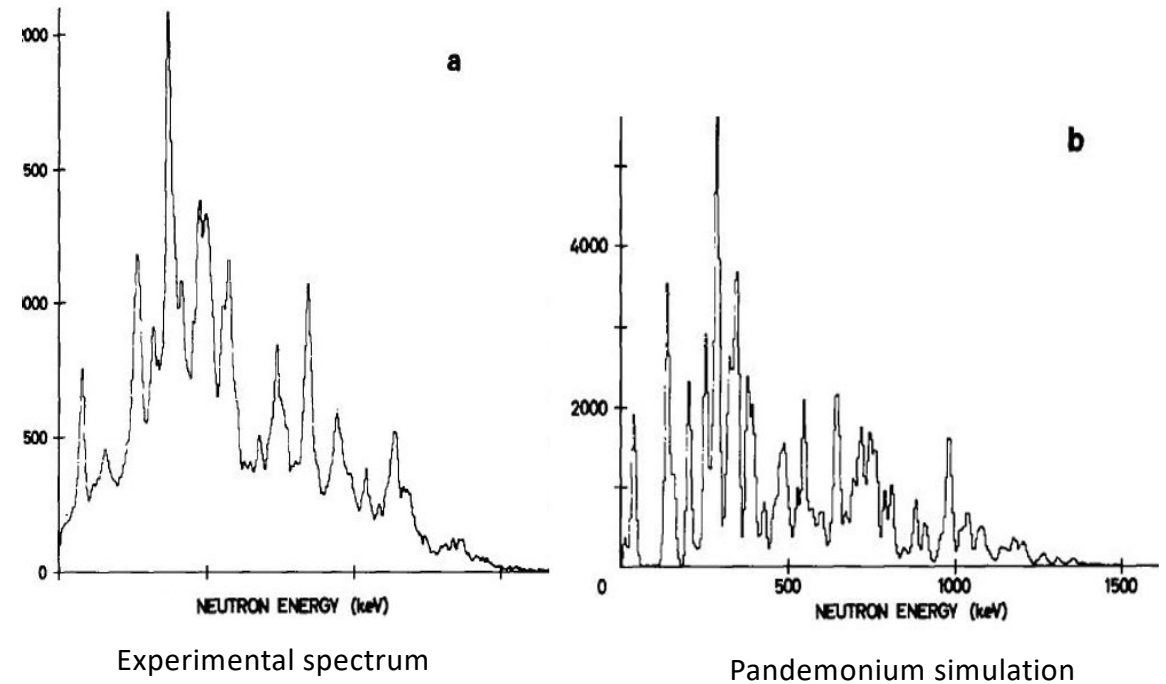
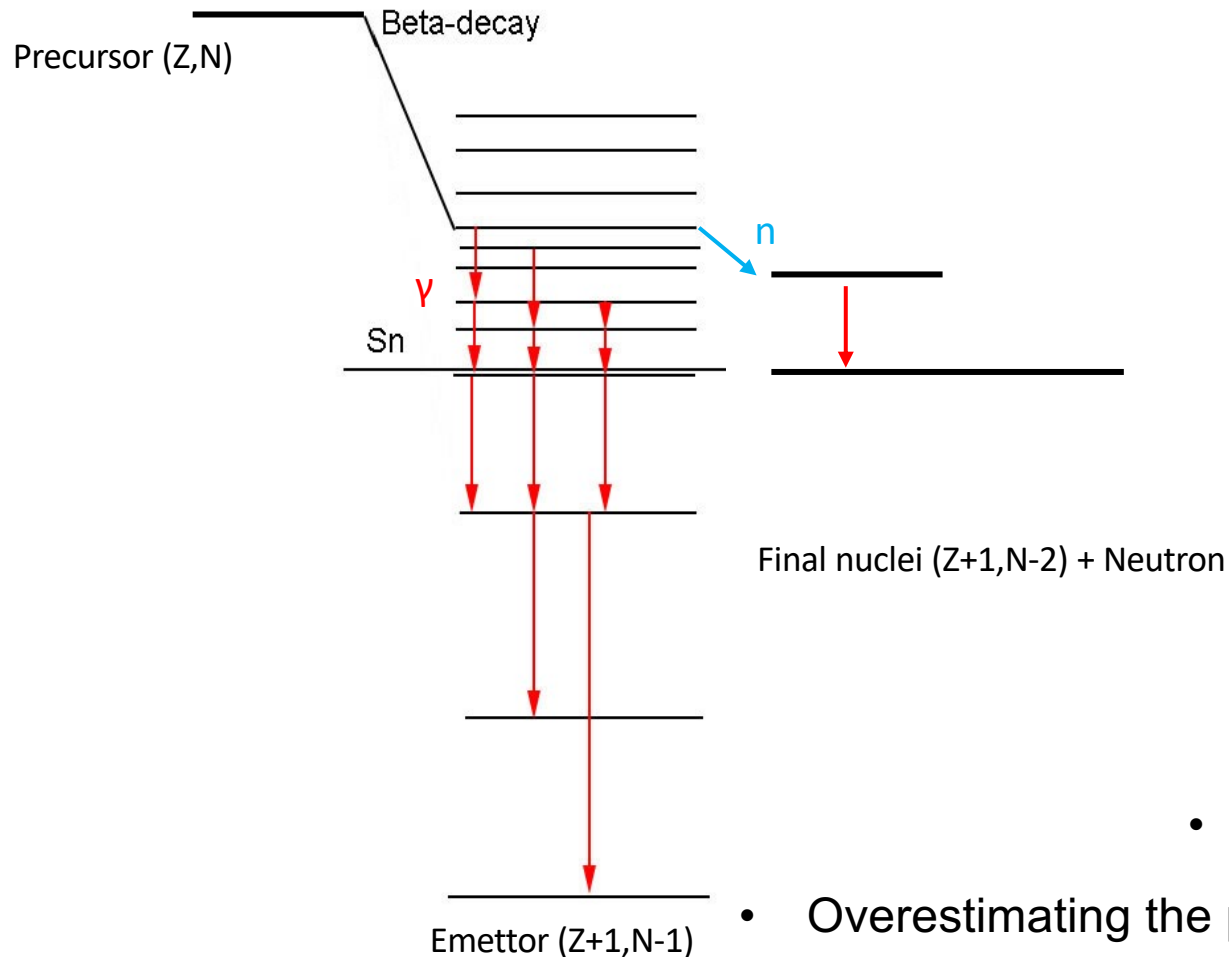
- $Q_{\beta-n} = Q_{\beta} - S_n$ \rightarrow Available energy for neutrons
- P_n : Probability for the daughter nucleus to emit at least one neutron after the beta decay

$$P_n = P_{1n} + P_{2n} + \dots \neq \langle n \rangle = P_{1n} + 2P_{2n} + \dots$$



The statistical models

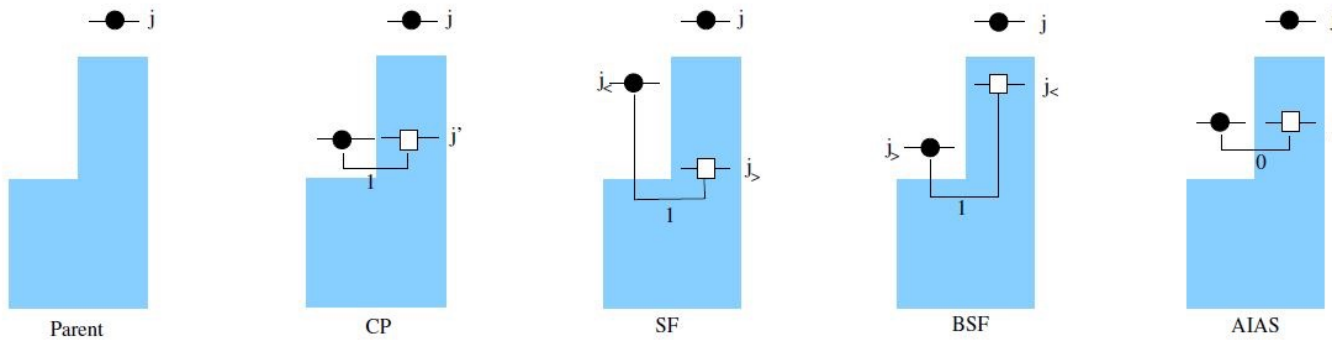
Hardy et al. "The essential decay of pandemonium: β -Delayed neutrons" 1977



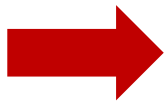
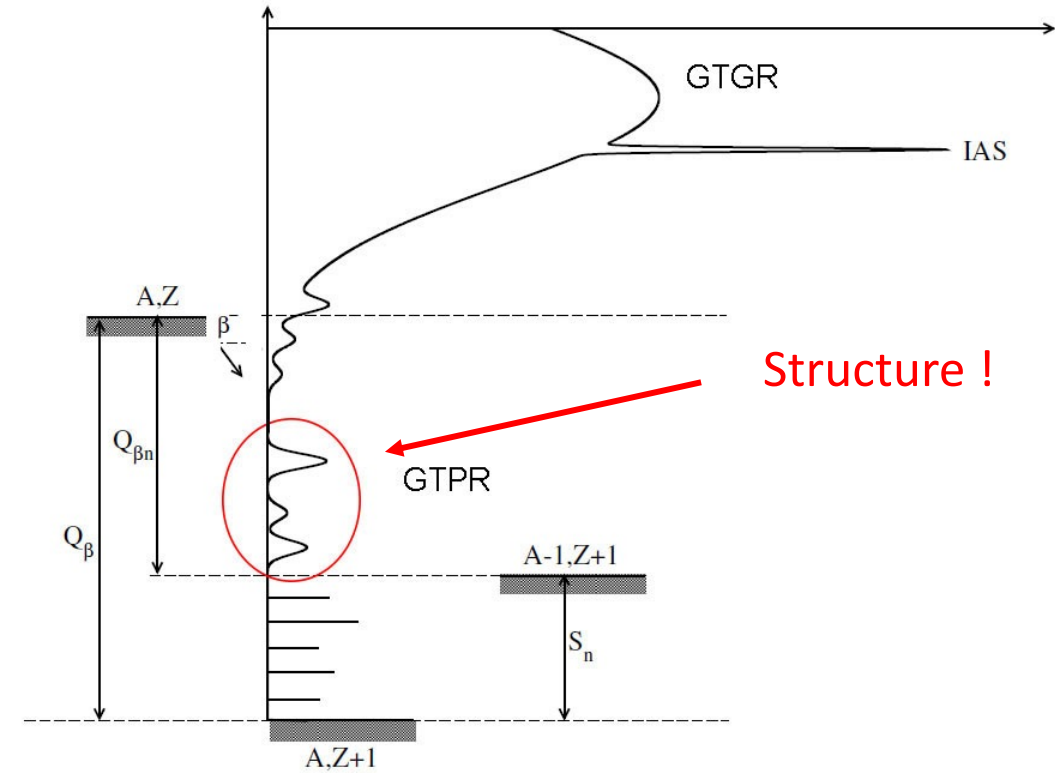
- Discret spectrum but bad prediction of P_n and $T_{1/2}$
- Overestimating the probability to populate levels under S_n : Pandemonium effect



Gamow-Teller « Doorway » transitions



$$\Delta L = 0 \quad \Delta J = 0, 1$$

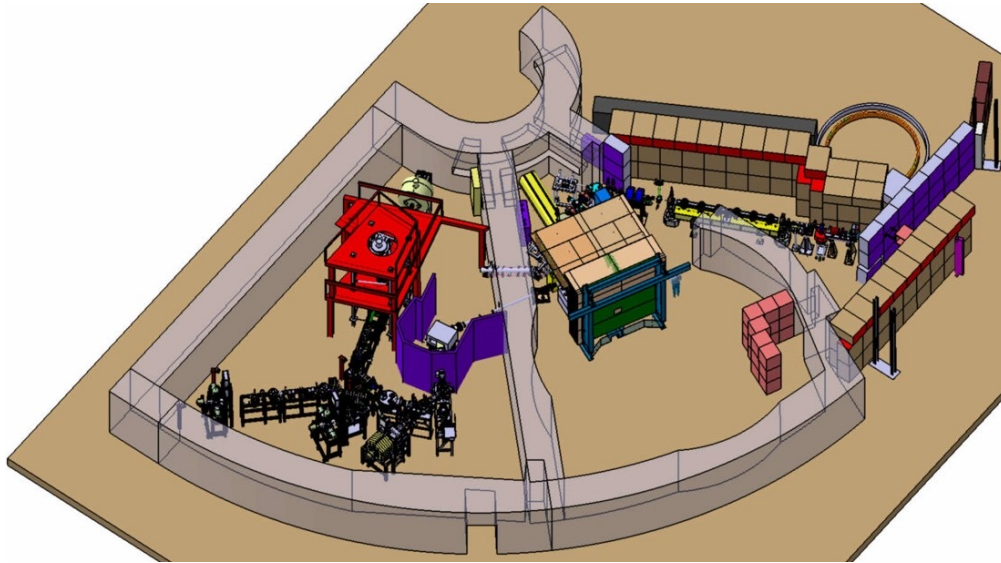


- Better results for integrated properties in $N = 50$ region
- No overestimation of the population of levels under S_n

Statistical and Non Statistical Models for Delayed Neutron Emission : Application to nuclei near $A=90$ Z.M. De Oliveira (1980)



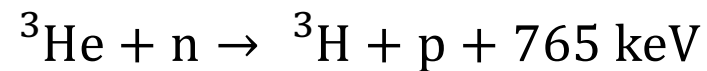
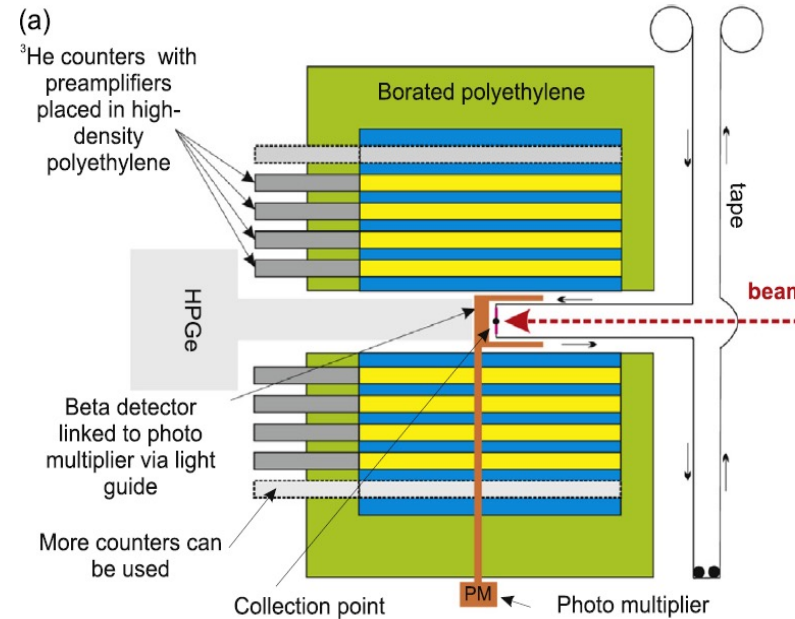
ALTO - LEB



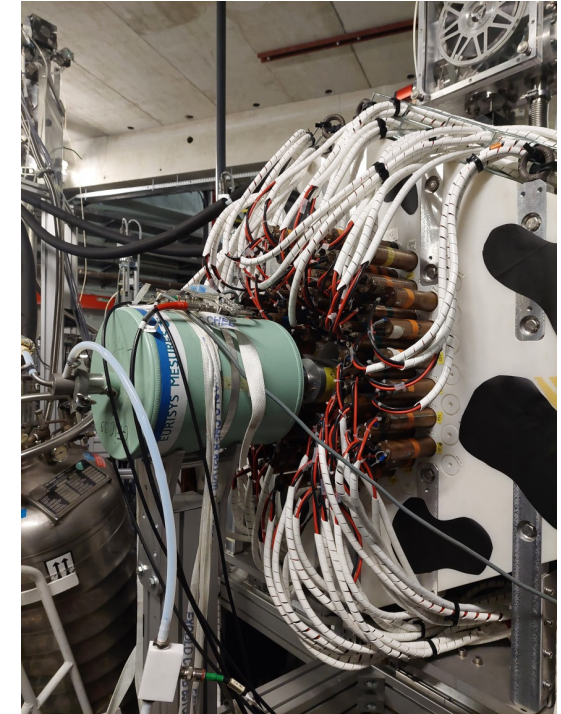
Beam production using ISOL technique

- Photofission
- Laser ionization
- Mass separation

Neutron counter : TETRA

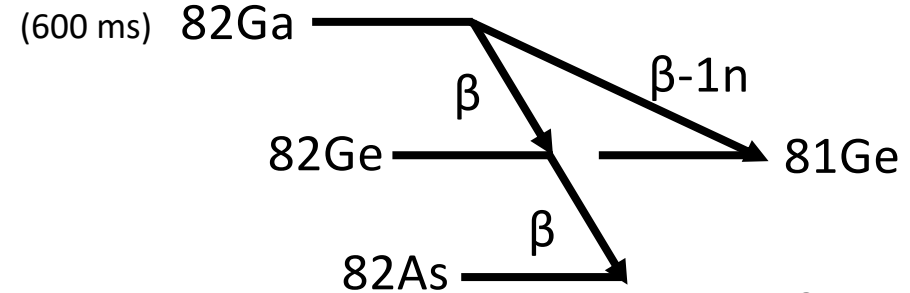
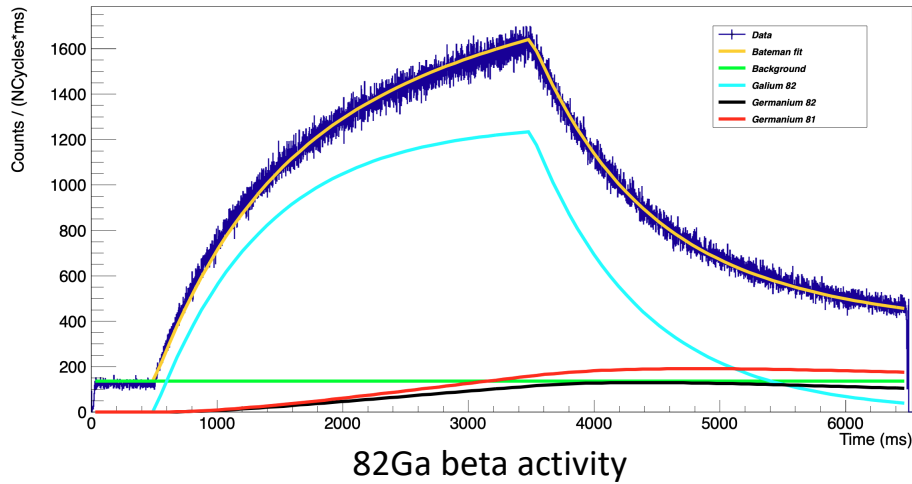


Around 5400 barns at thermal energies



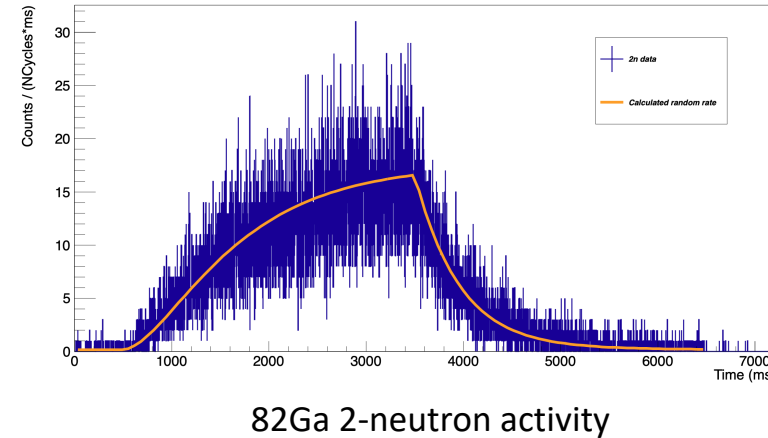
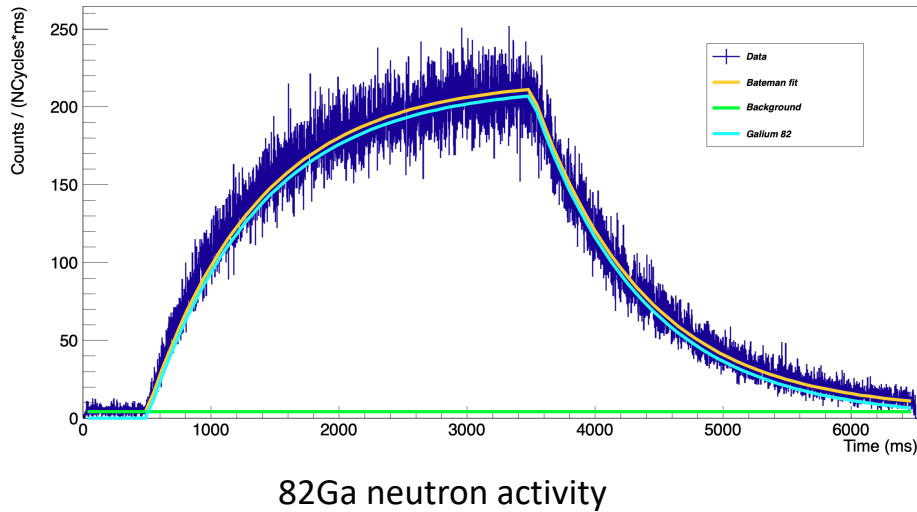


Calibration isotope : 82Ga



$$Q_{\beta-1n} = 5.290 \text{ MeV}$$

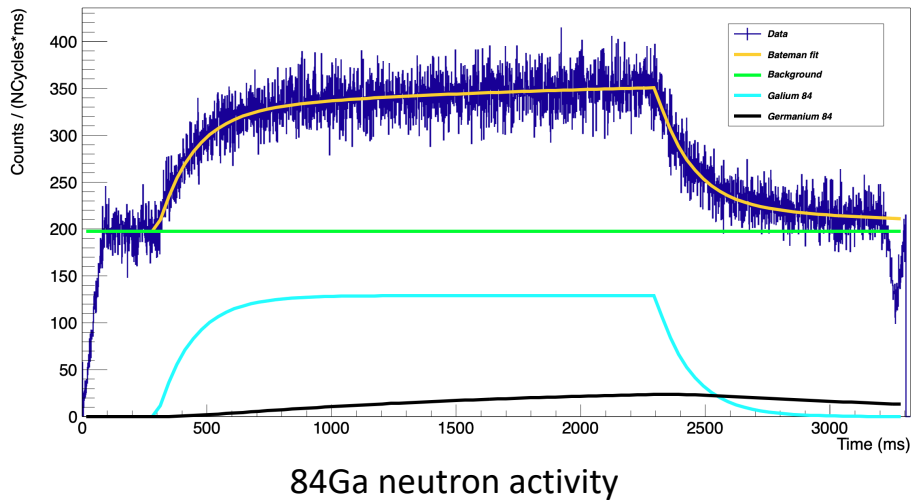
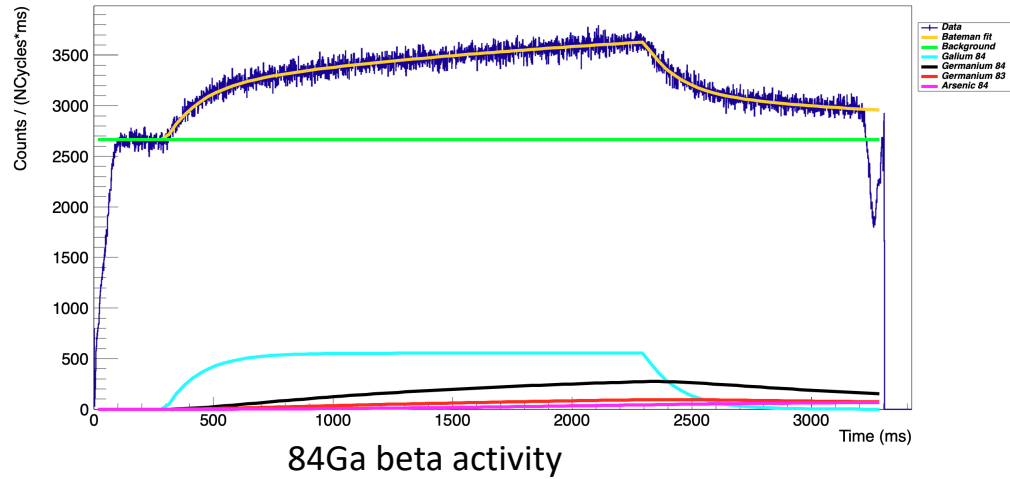
$$Q_{\beta-2n} < 0 \text{ MeV}$$



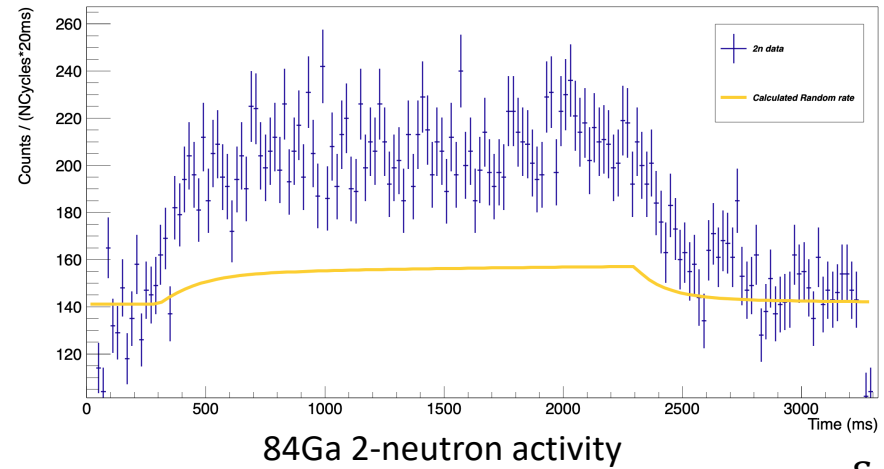
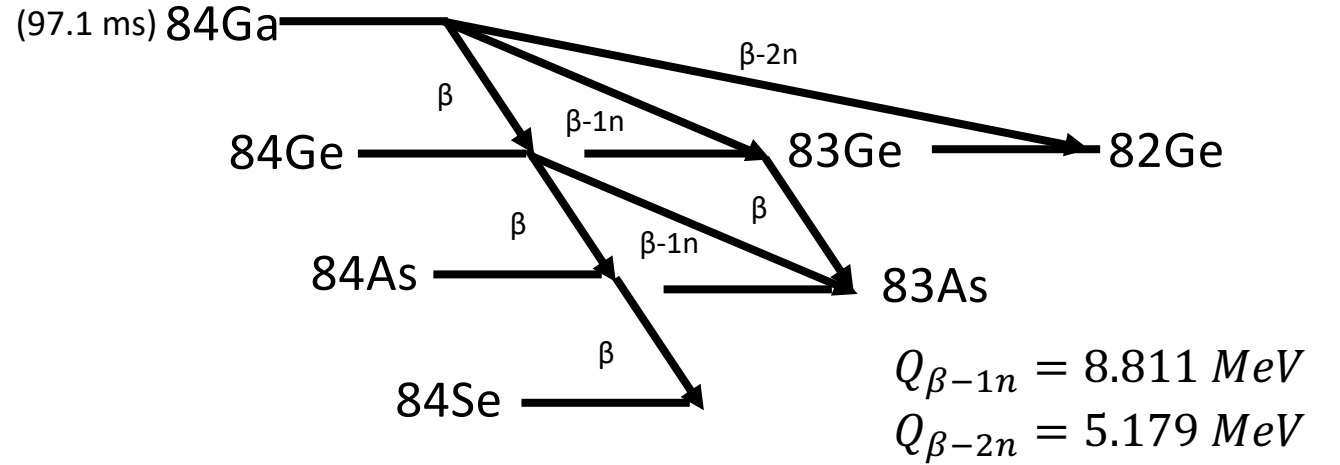
$$P_n = \frac{N_n}{N_\beta} \times \frac{\epsilon_\beta}{\epsilon_n} = 22.6\% \quad \rightarrow \text{Only random 2n counting}$$



84Ga data analysis



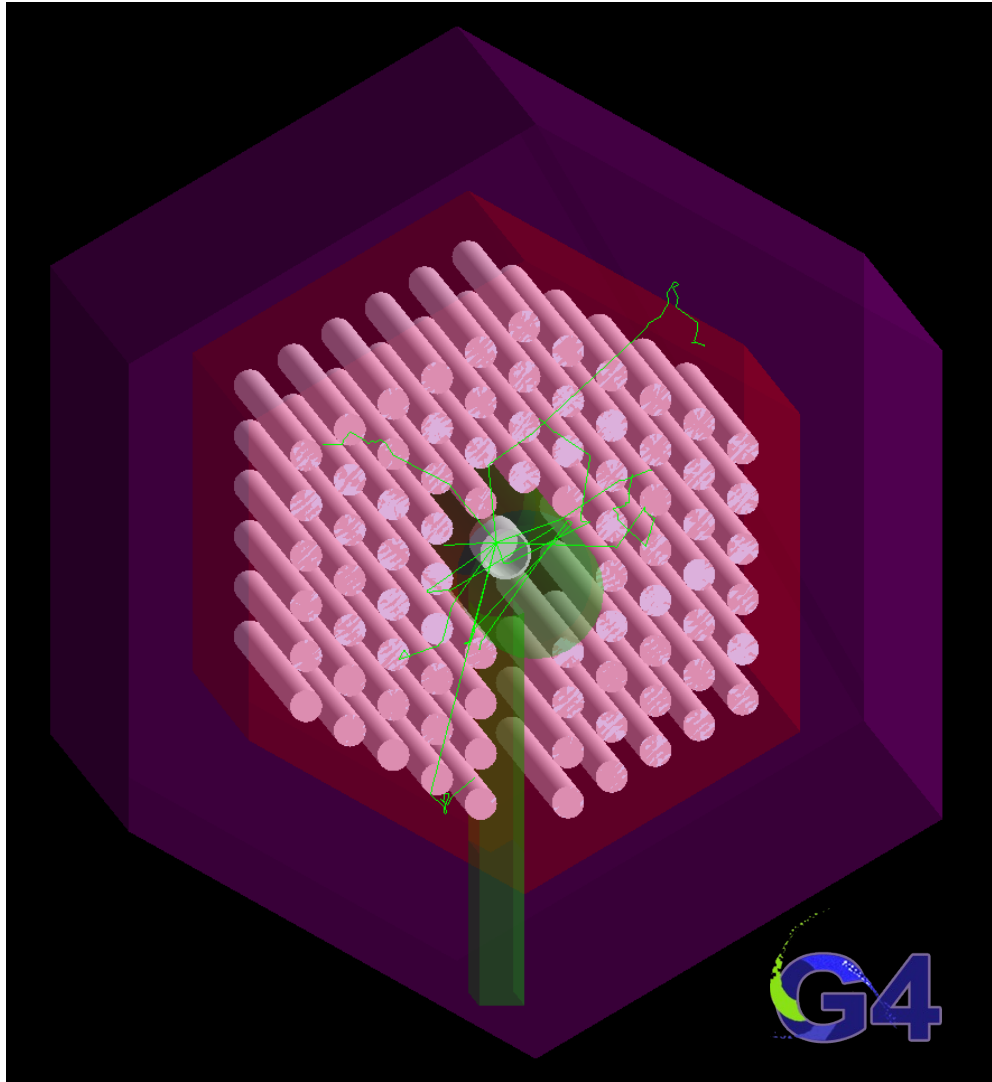
$$\langle n \rangle^* = P_{1n} + (1 + \epsilon_n)P_{2n} = 37.5 \%$$



$$P_{2n} = \frac{\epsilon_{\beta}}{\epsilon_n^2} \times \frac{N_{2n}}{N_{\beta}} = 1.2 \%$$

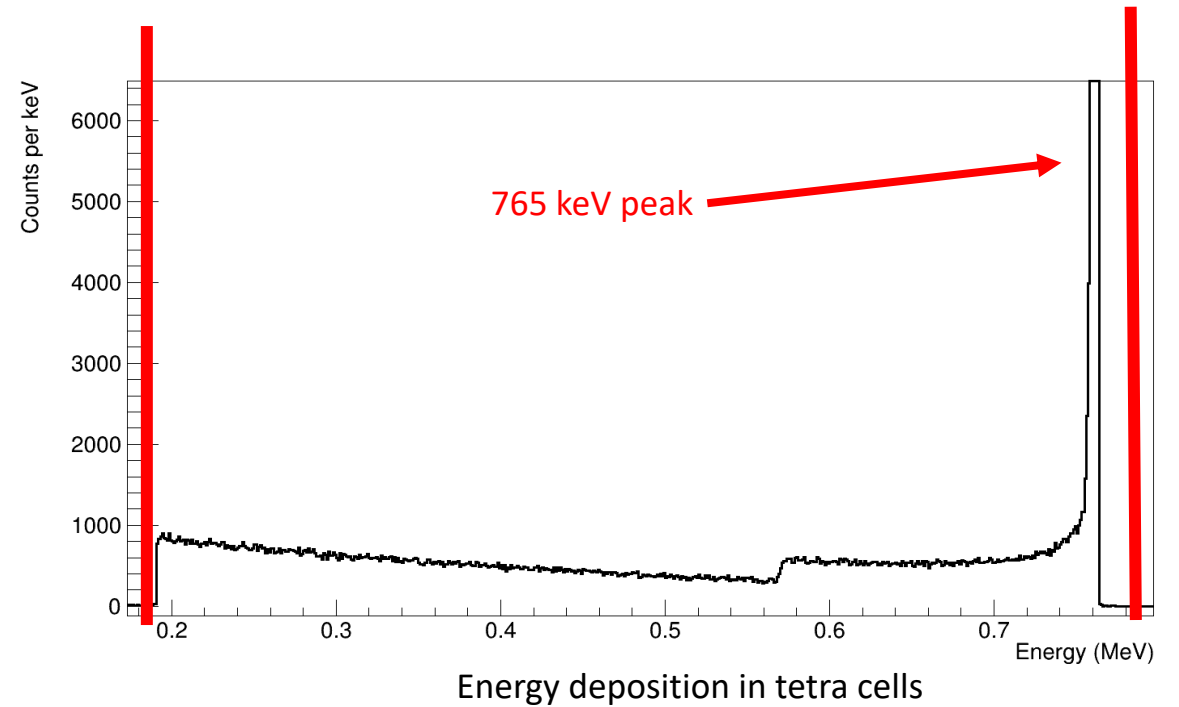


Geant4 TETRA simulation



High energy threshold : Sparks, alphas, neutrons

Low energy threshold: gamma pile-up, noise



Low energy tail → Events were the secondary particles left the cell before depositing all their energy

*Thermal neutrons physics list included



Different energy distributions

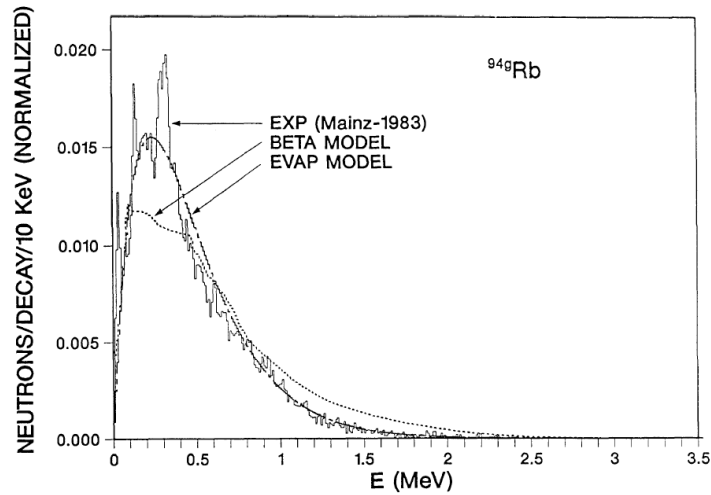


Fig. 58. Comparison of spectra for ^{94}Rb .

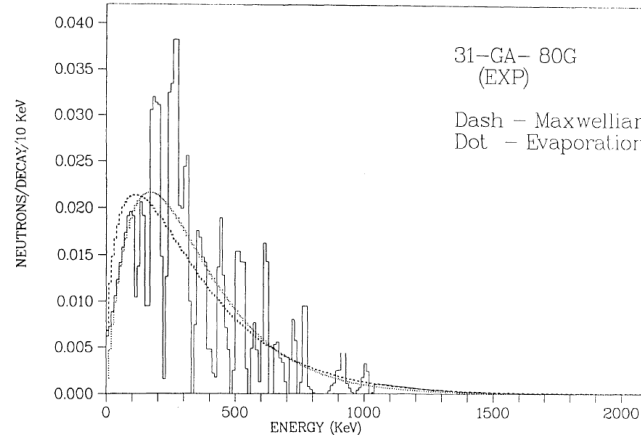


Fig. 47. Comparison of model spectra for ^{86}Ga .

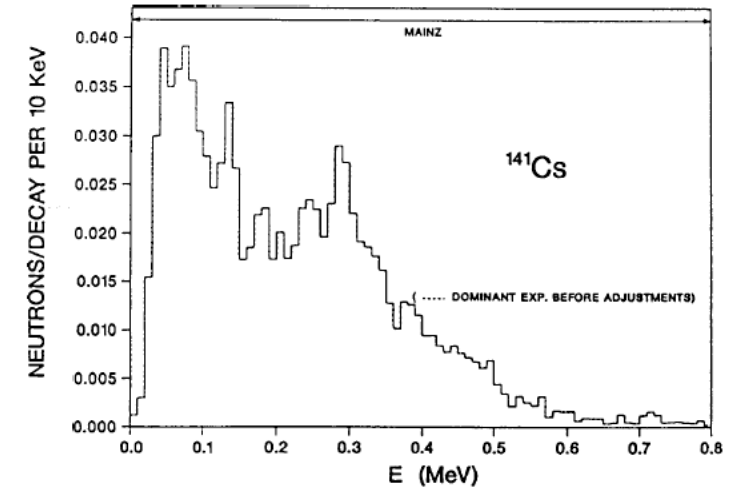
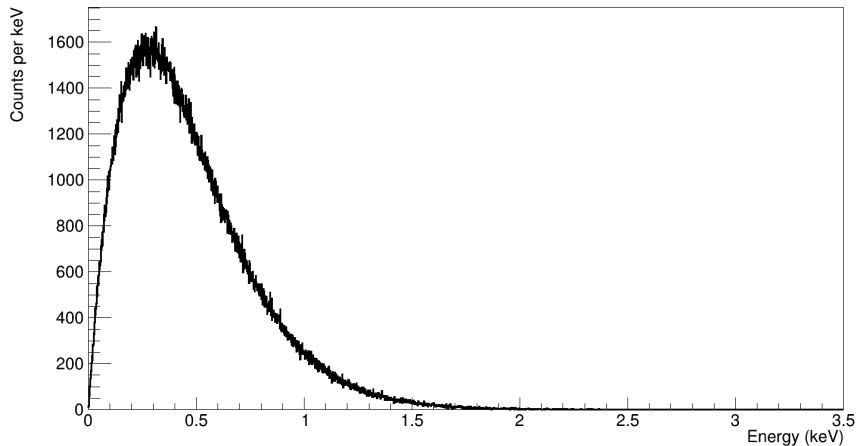


Fig. 39. Normalized delayed neutron spectra for ^{141}Cs .

« Evaluation and Application of Delayed Neutron Precursor Data » Michaele Clarice Brady

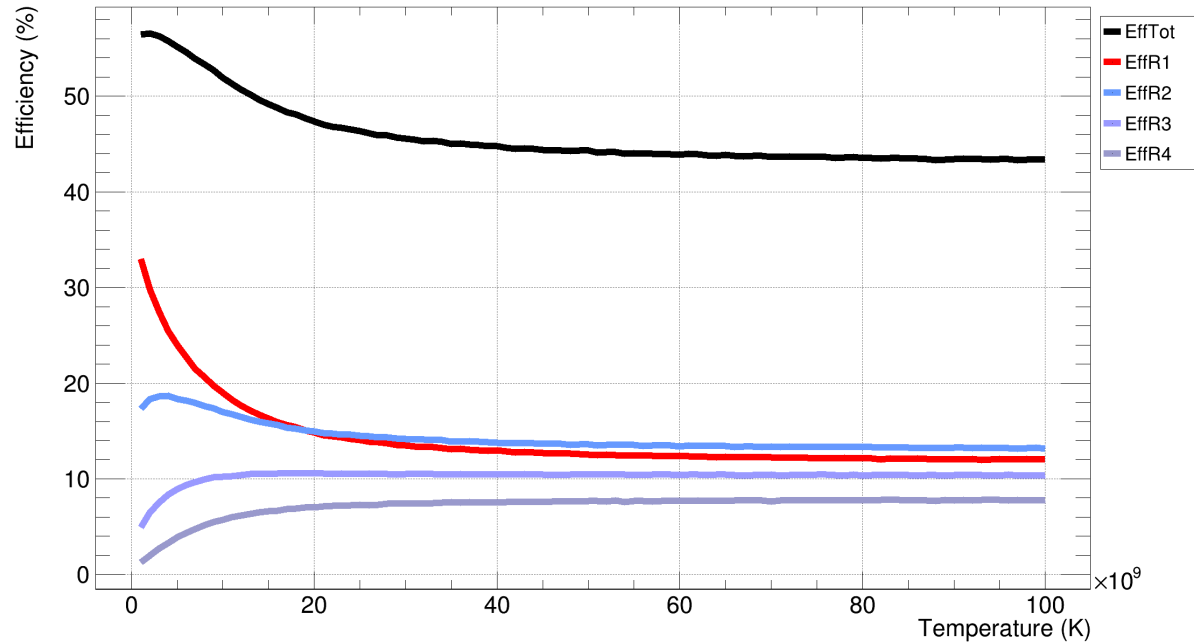


Maxwellian approximation -> Parameter : Temperature

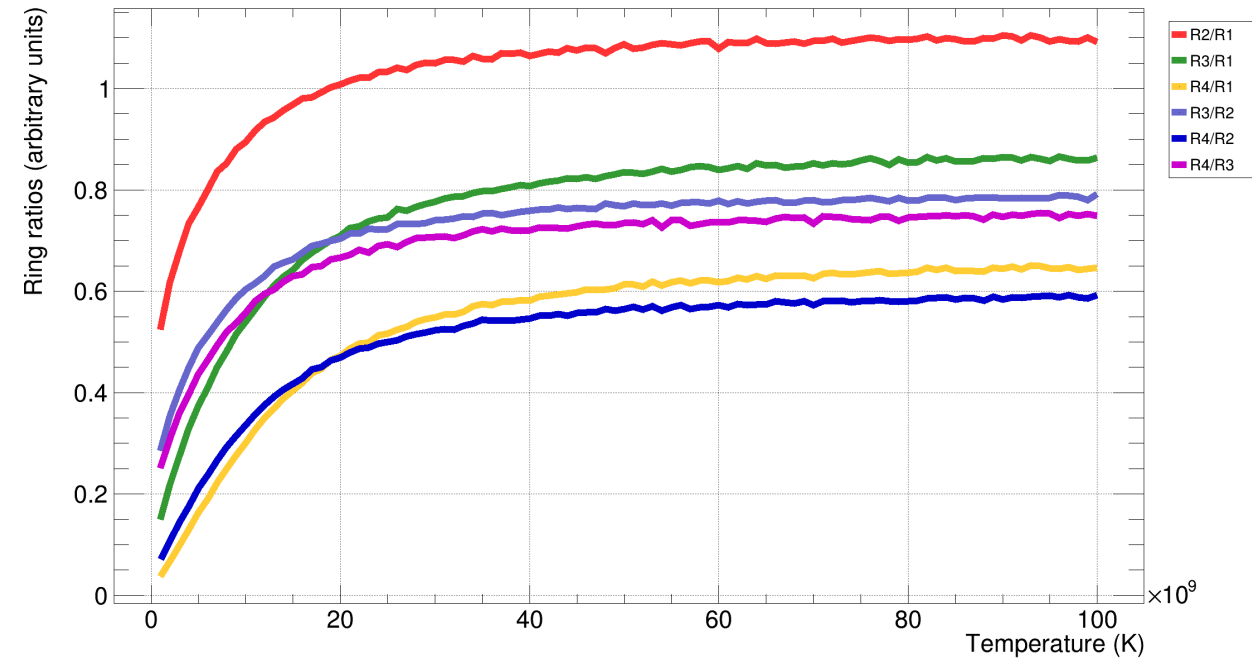
*S.-T. Park, "Neutron Energy Spectra of ^{252}Cf , Am-Be source and of the $\text{D}(d,n)^3\text{He}$ Reaction,"



TETRA efficiency per ring



Efficiency as function of Maxwellian temperature parameter

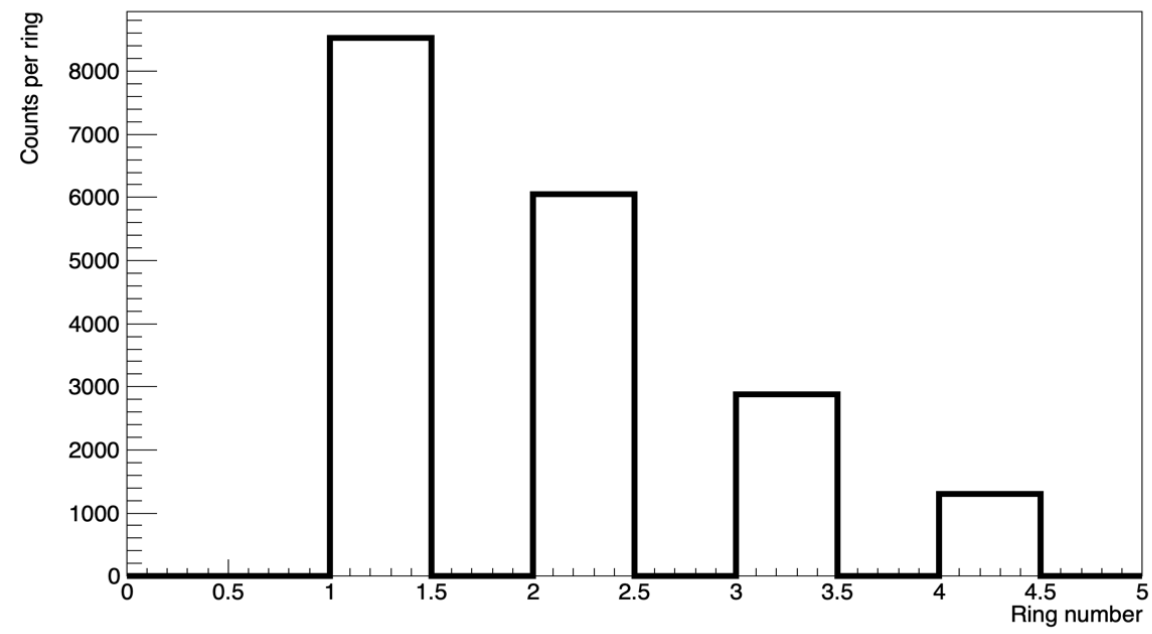


Ring ratios as function of Maxwellian temperature parameter

- The efficiency per ring changes with the neutron energy
- Can a link between ring efficiency and mean neutron energy be made ?

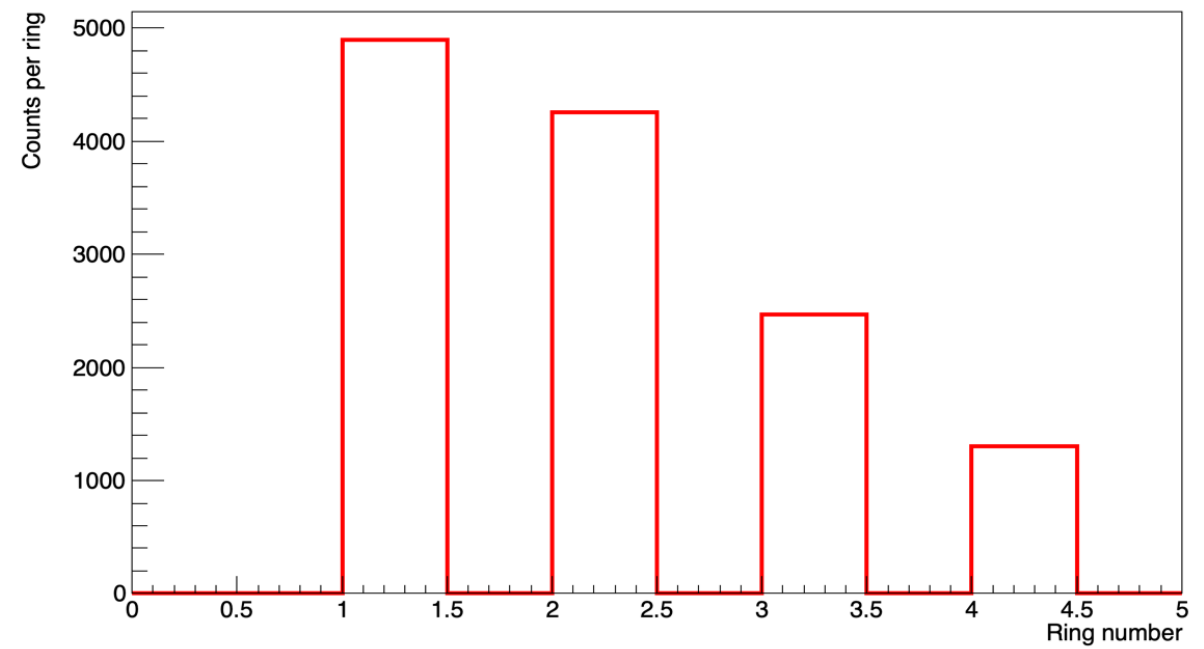


Different energy distributions



82Ga Ring Ratios

0.93 MeV mean energy



84Ga Ring Ratios

2.1 MeV mean energy



Coming up

MONSTER @ ALTO

- > Structure installation in Juin 2024
- > Experiment MONSTER + BEDO planned in Autumn 2024





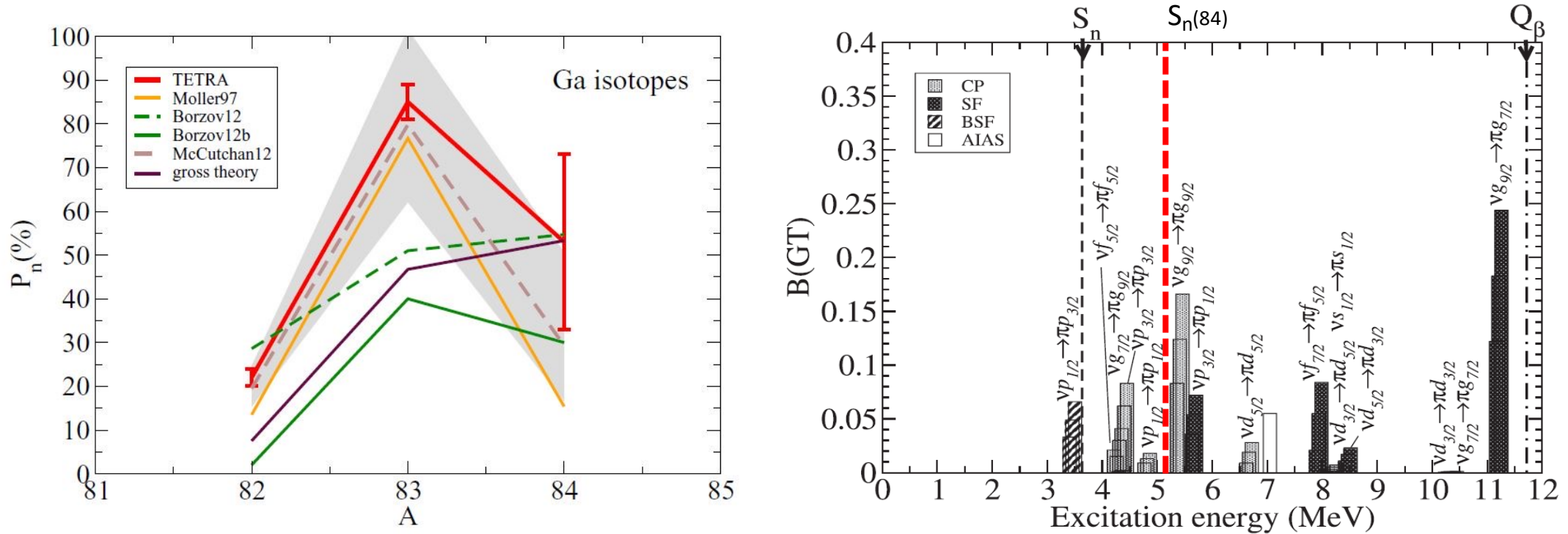
Thank you for your
attention



Back up slides



Galium (Z=31) systematics



Verney et al. "Pygmy Gamow-Teller resonance in the N = 50 region: New evidence from staggering of β -delayed neutron-emission probabilities" 2017



Structure revealed