



ARC CENTRE OF EXCELLENCE FOR
GRAVITATIONAL WAVE DISCOVERY

Radio variability with ASKAP and LOFAR MW follow-up of radio transient and vice versa

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ASTRO-COLIBRI Workshop 2025



Outline

Follow-up of radio transients

- slow
- fast

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Radio follow-up of

- optical
 - multi messenger
 - high-energy
- transients

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Identifying and studying

- Synchrotron transients
- Long-period radio transients

&

Radio follow-up of

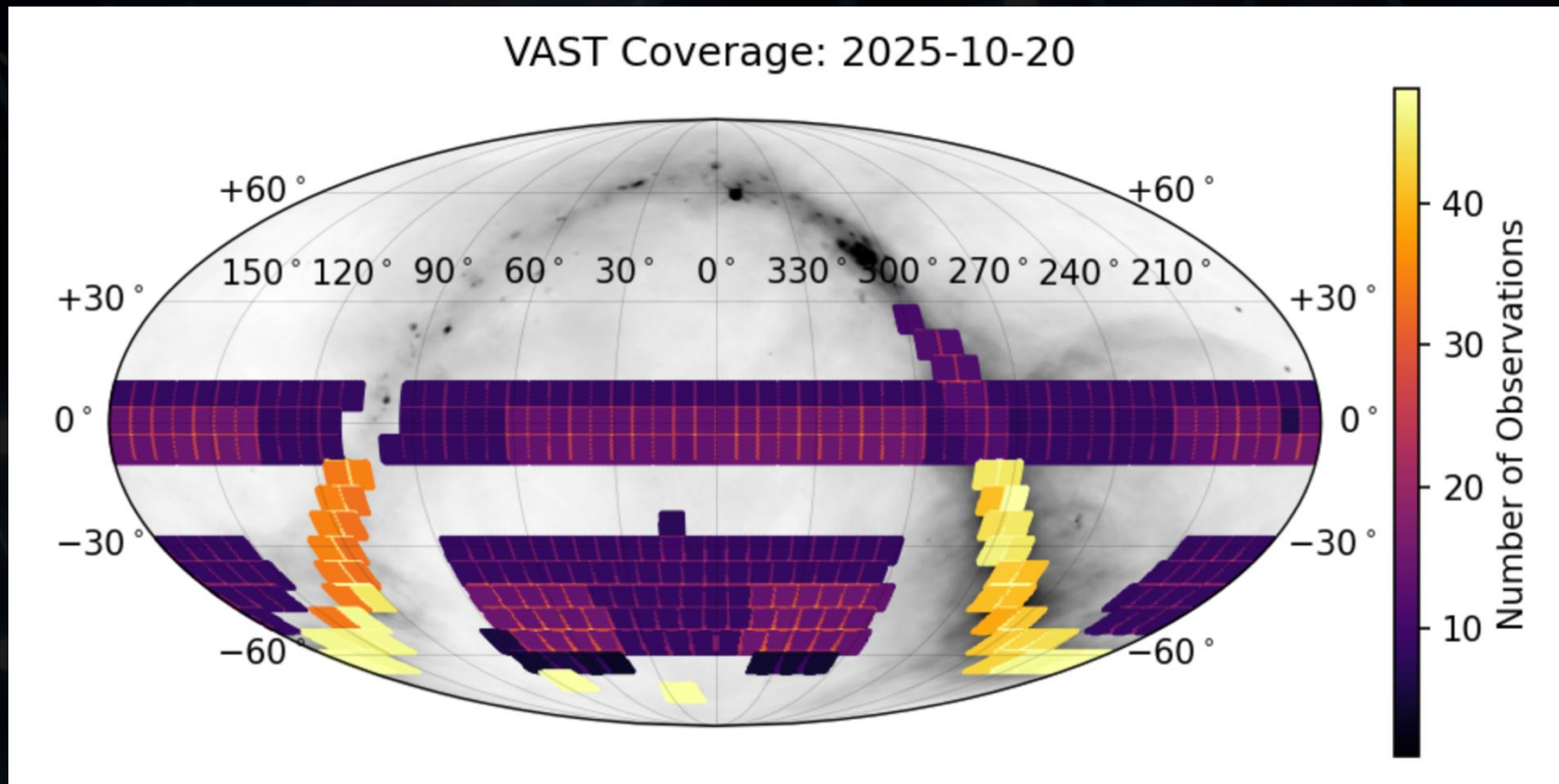
- Recurrent novae
- GW170817
- Gamma-ray bursts

Identifying radio transients (1) dedicated surveys

Variables and Slow Transients (VAST)
survey science project on ASKAP (Murphy et al. 2013/2021)



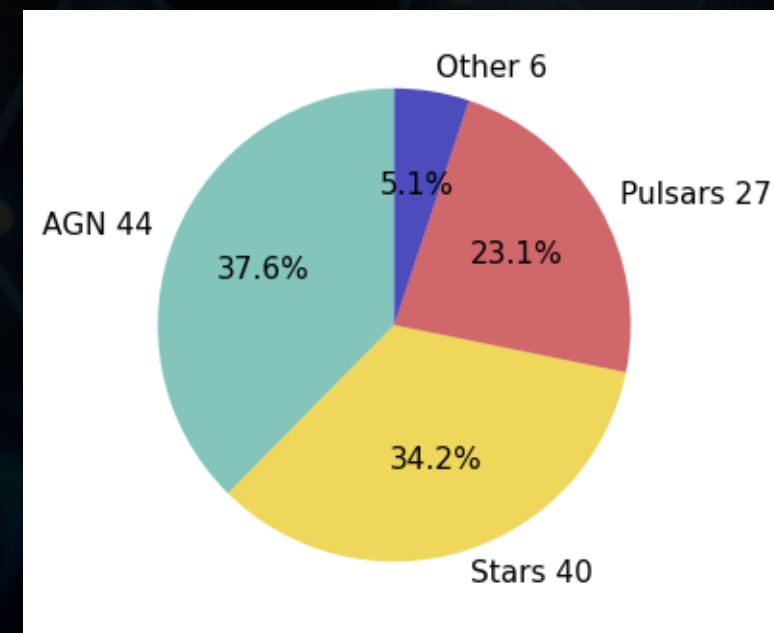
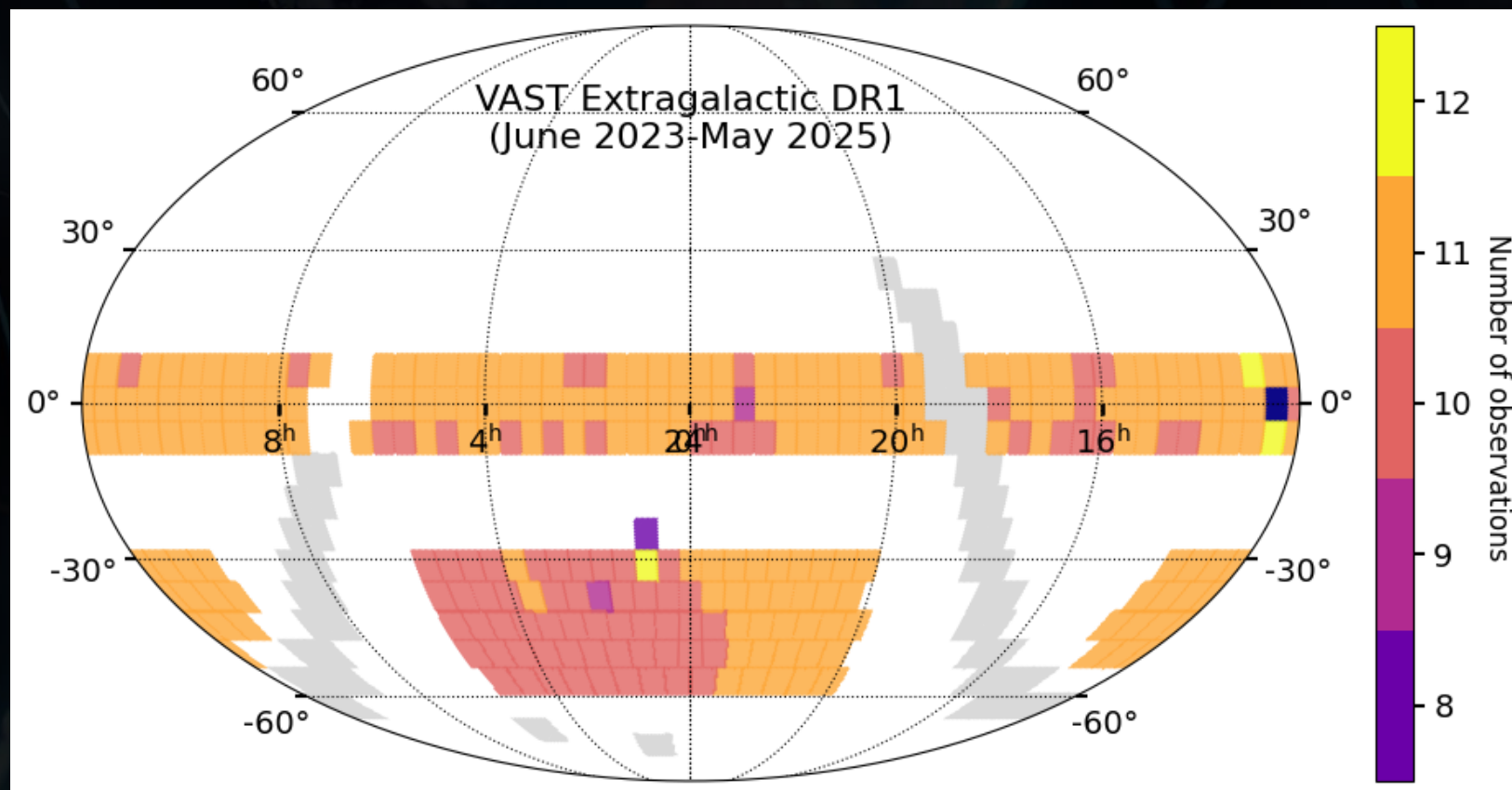
(36x12 m, 6 km)



Identifying radio transients (1)

VAST Extragalactic DR1

Full survey started in 2023 -> 10-12 observations per field



de Ruiter et al. (in prep)

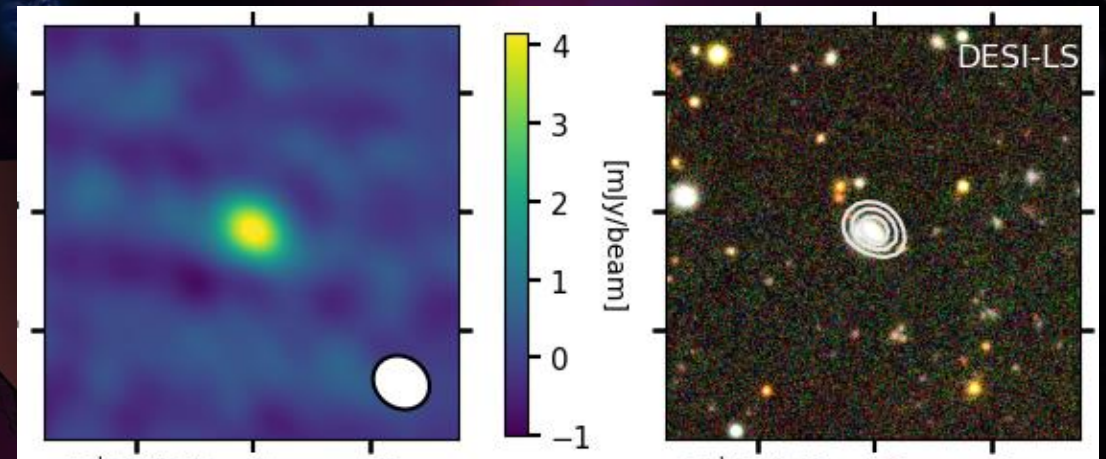
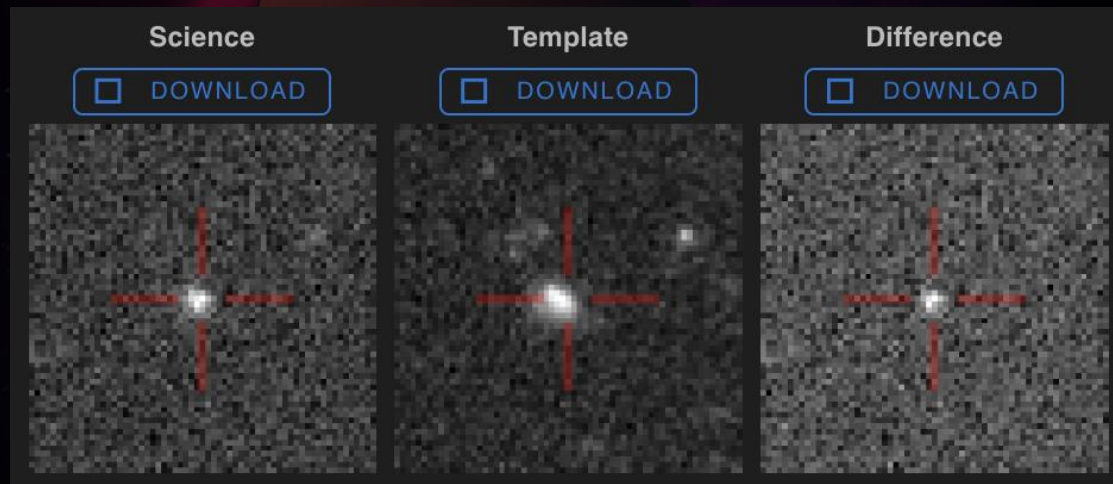
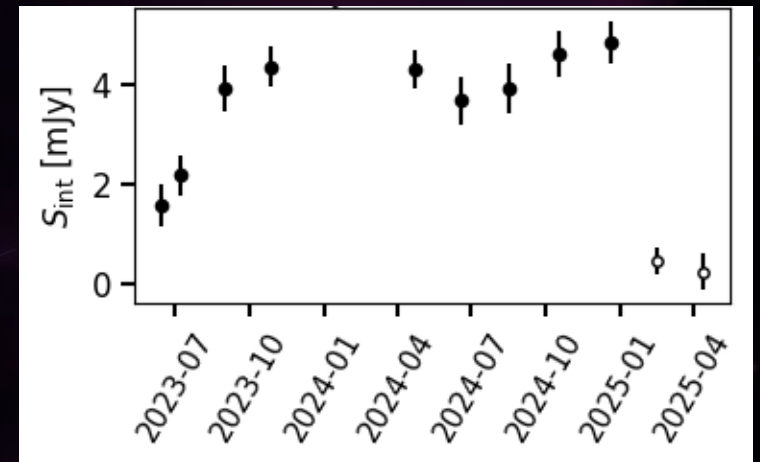
Slow transients: supernova candidate

Highly variable radio source in VAST

Dwarf galaxy at $z_{\text{phot}} = 0.034 \pm 0.015$

Peak radio luminosity $\sim 10^{29} \text{ erg s}^{-1} \text{ Hz}^{-1}$

AleRCE ZTF explorer



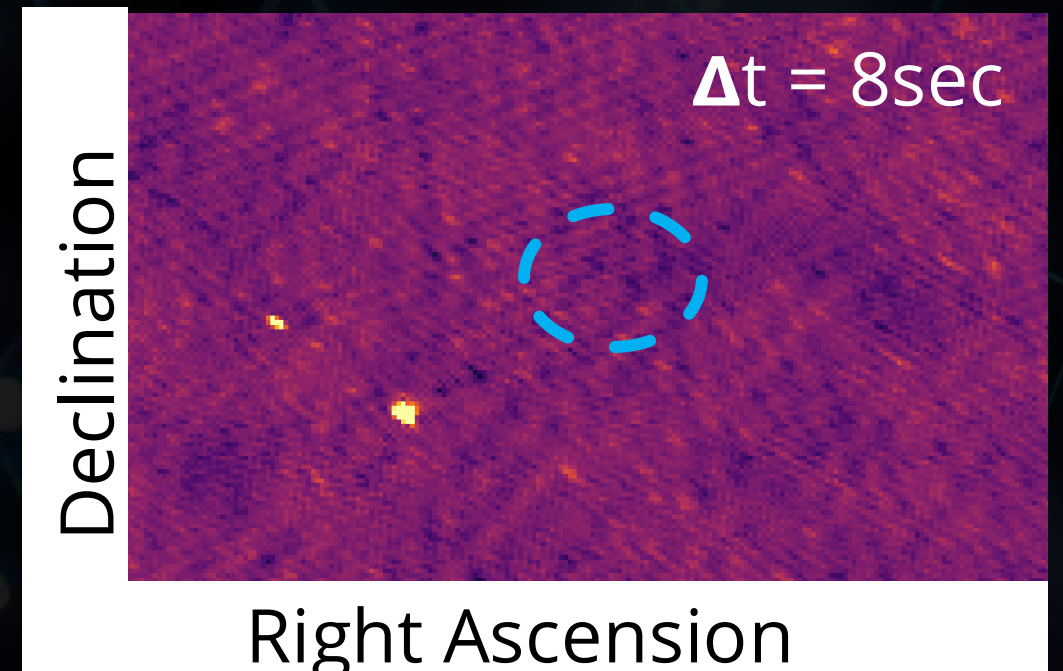
Identifying radio transients (2) commensal searches

Single observation – multiple science cases

High time resolution imaging is a recent development

(de Ruiter et al. 2025)

Complementary to
beamformed searches (FRBs)



Current commensal, high-time resolution imaging searches

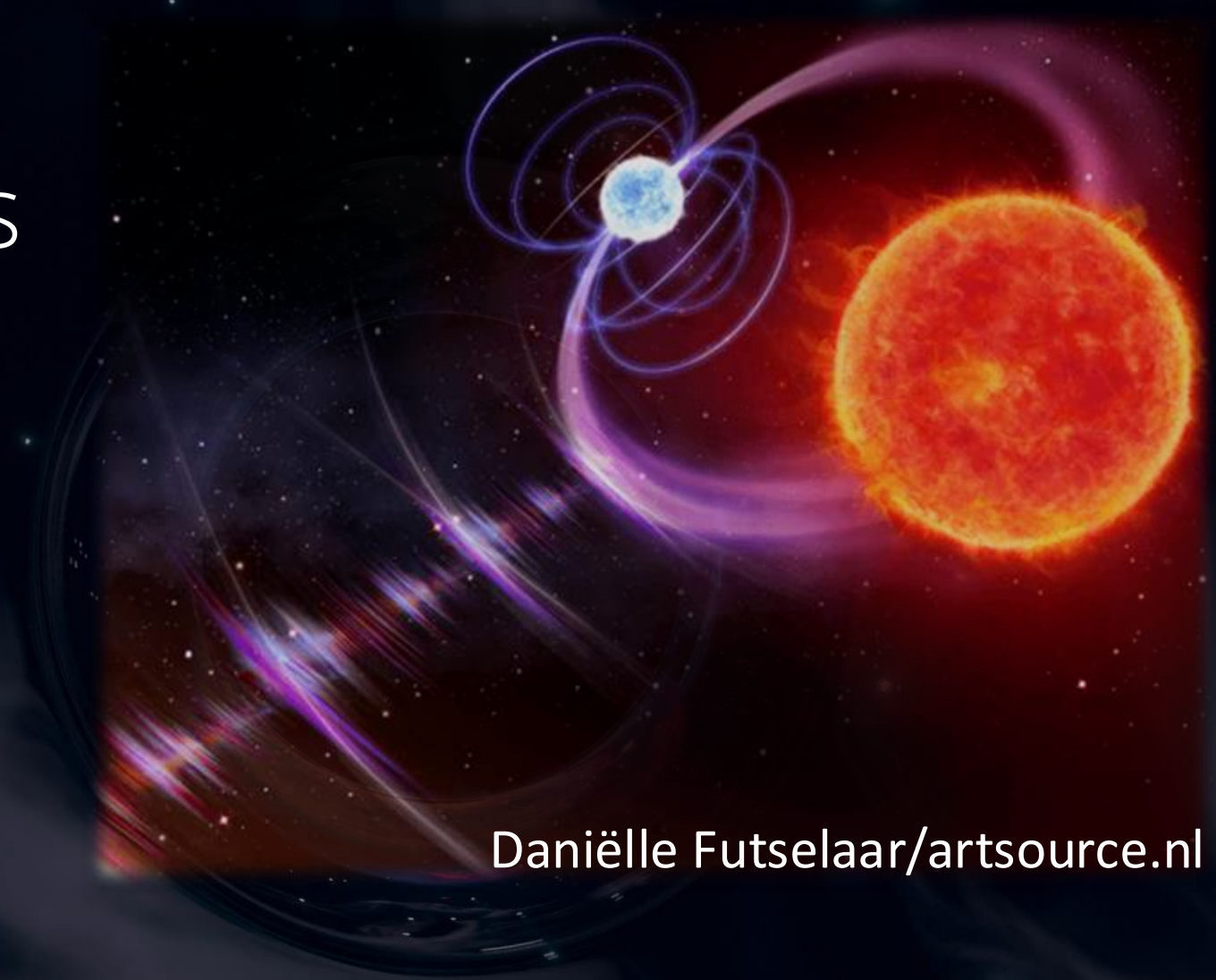
Custom pipelines for:

- MeerKAT – TRON (Smirnov et al. 2025)
- MWA (Horváth et al. 2025)
- ASKAP – VASTER (Wang et al. 2023)
- LOFAR (de Ruiter et al. 2024)

Fast transients: long-period radio transients

- Bright second to minute radio bursts
- Periods of minutes to hours

Slowly spinning pulsars?



Daniëlle Futselaar/artsource.nl

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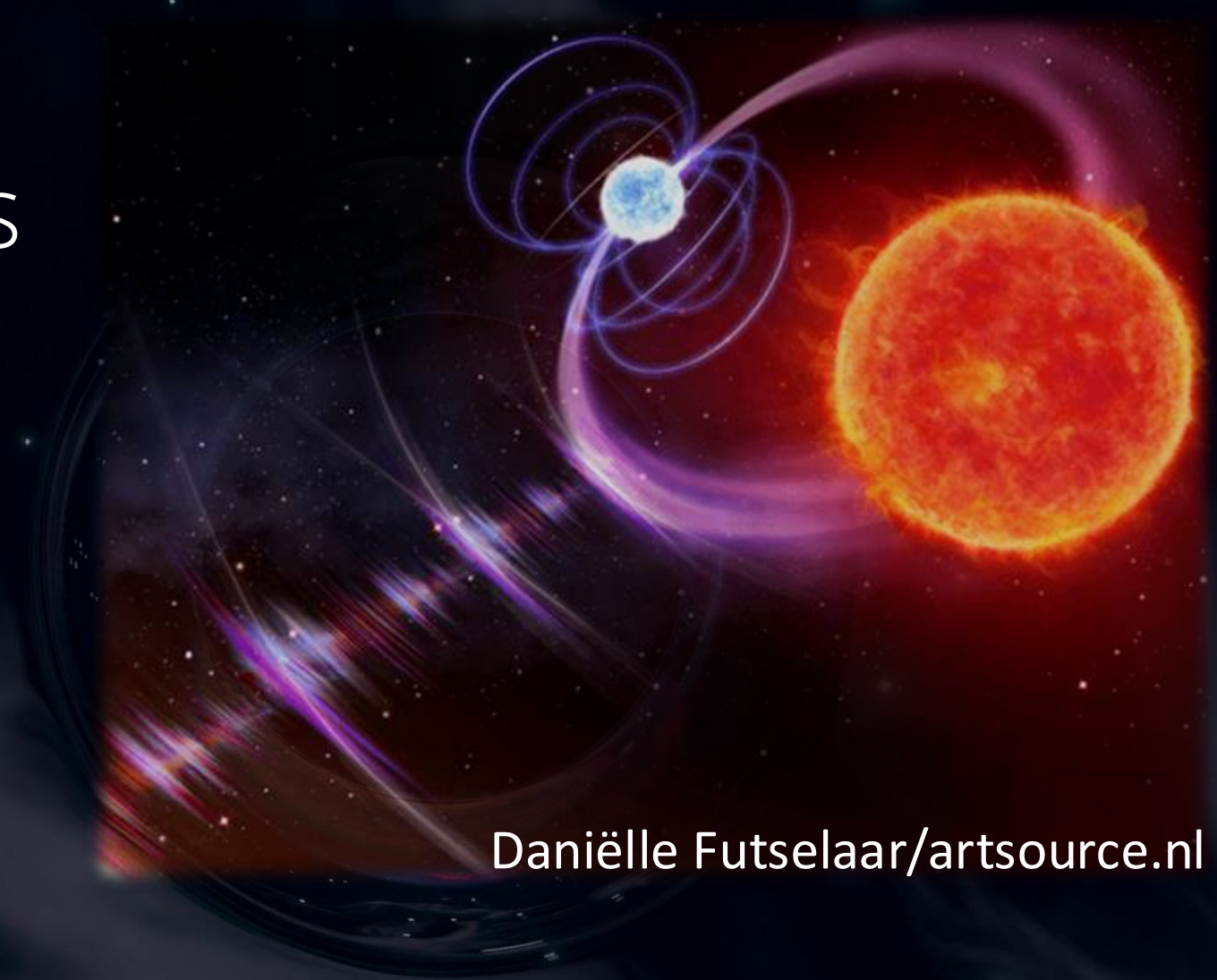
Slowly spinning pulsars?

ILTJ1101+55 (de Ruiter et al. 2025)

- Optical counterpart
- White dwarf – M dwarf binary

Other wavelengths

- Two X-ray detected LPTs (Wang et al. 2025, Anumalapudi et al. 2025)
- Optical variability (Marsh et al. 2016; Pelisoli et al. 2023)



Daniëlle Futselaar/artsource.nl



and vice versa

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‘Traditional’ approach
for synchrotron
transients

RS Ophiuchi

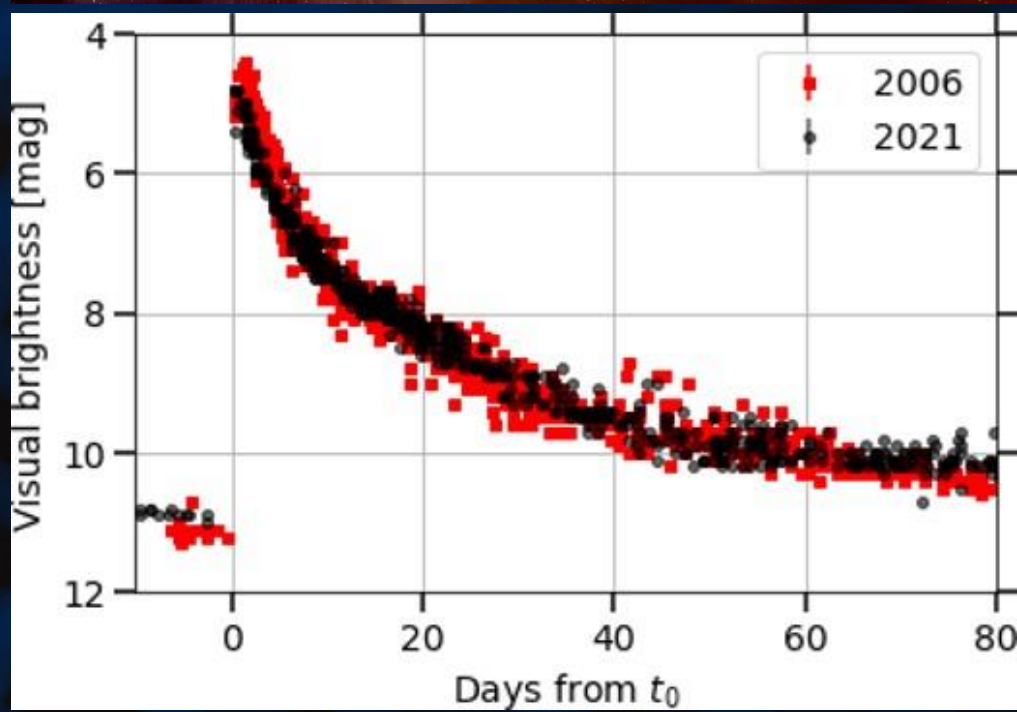
Recurrent nova system

White dwarf accreting from companion

Type 1A Supernova?

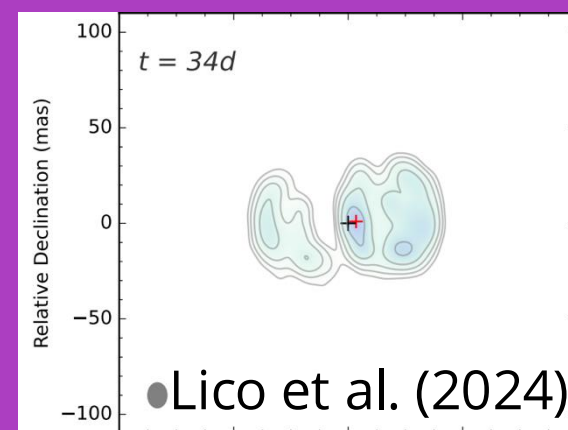
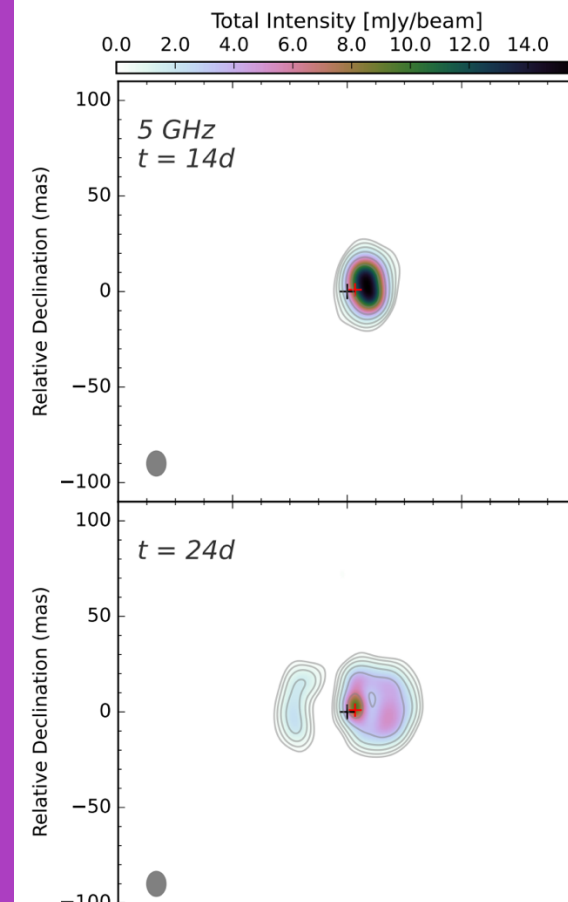
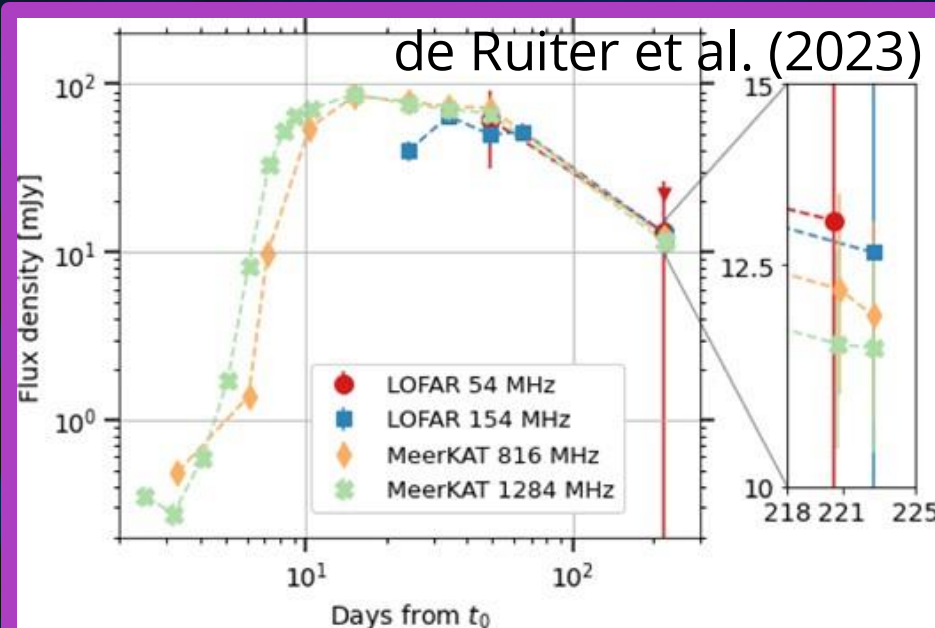
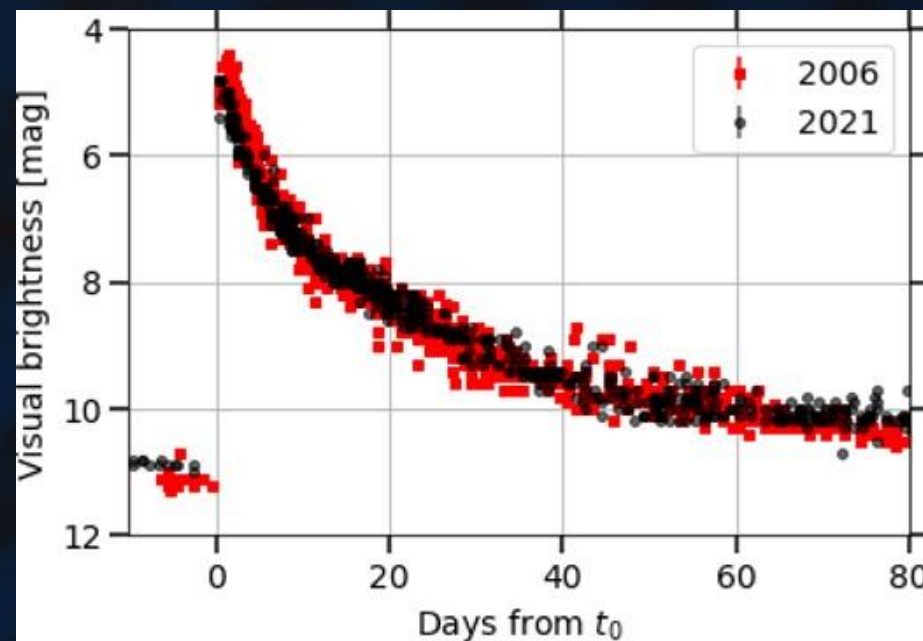
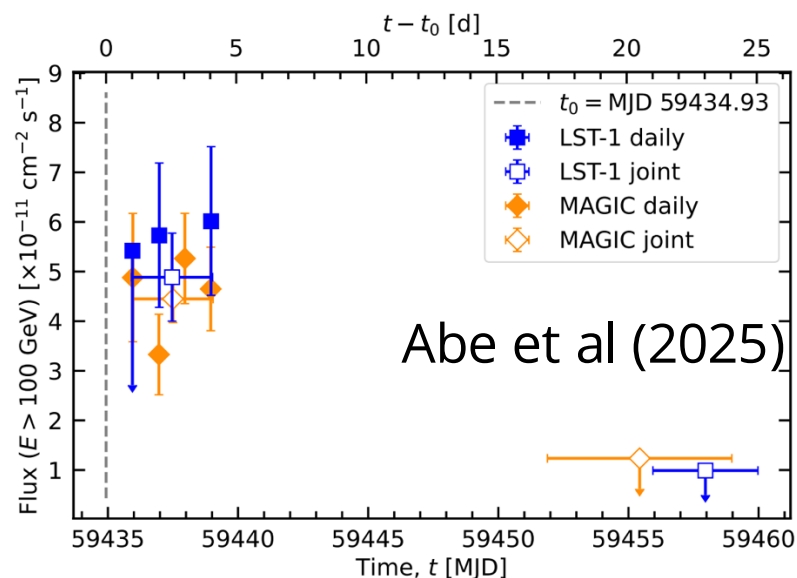
Maximum white dwarf mass?

David A. Hardy/PPARC



RS Ophiuchi

2021 outburst
A wide range of
multi-wavelength
follow-up

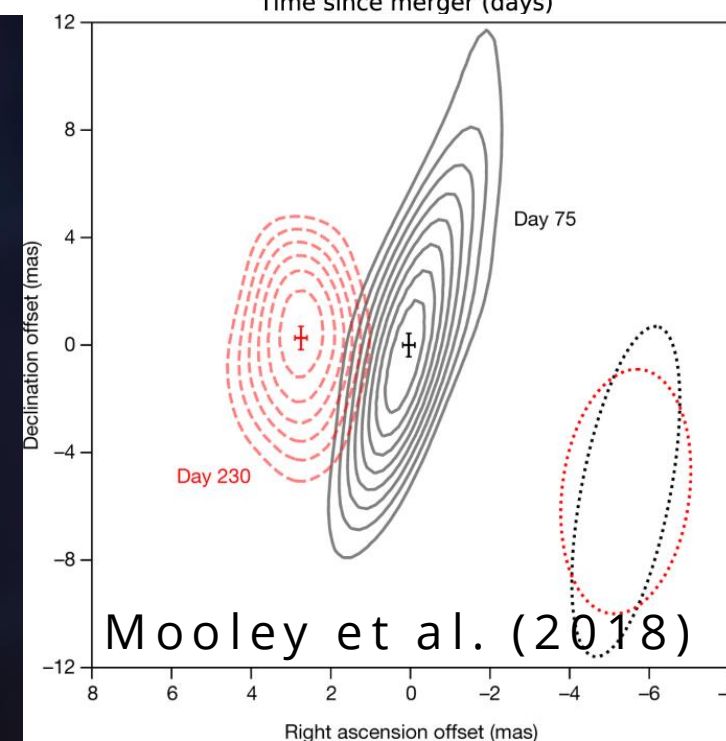
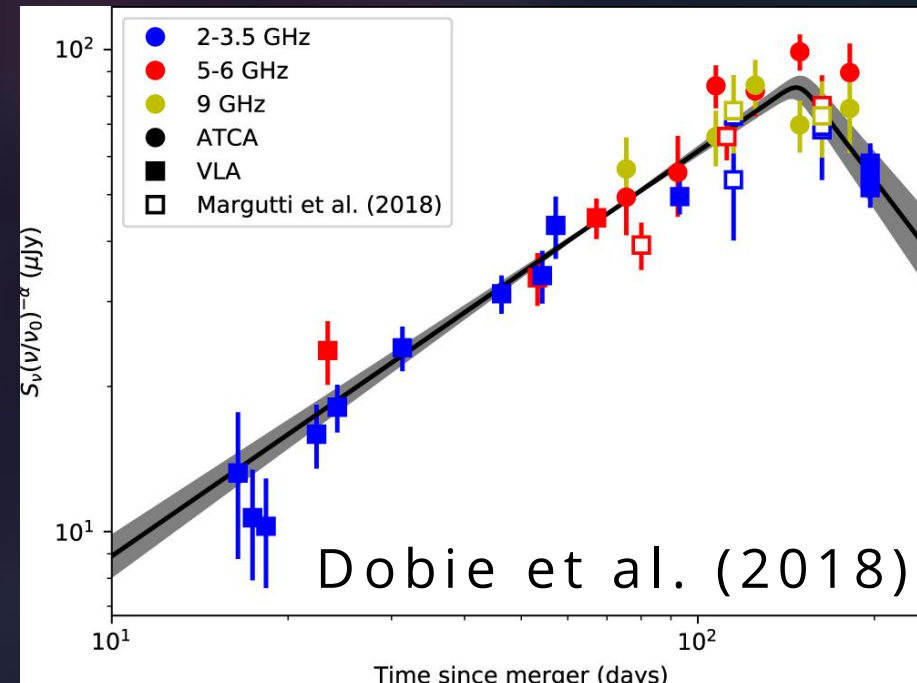


GW170817

GW170817 associated with a sGRB
(Abbott et al. 2018)

Unique constraints on:

- Energetics
- Circumbinary material
- Jet structure (cocoon/top-hat)



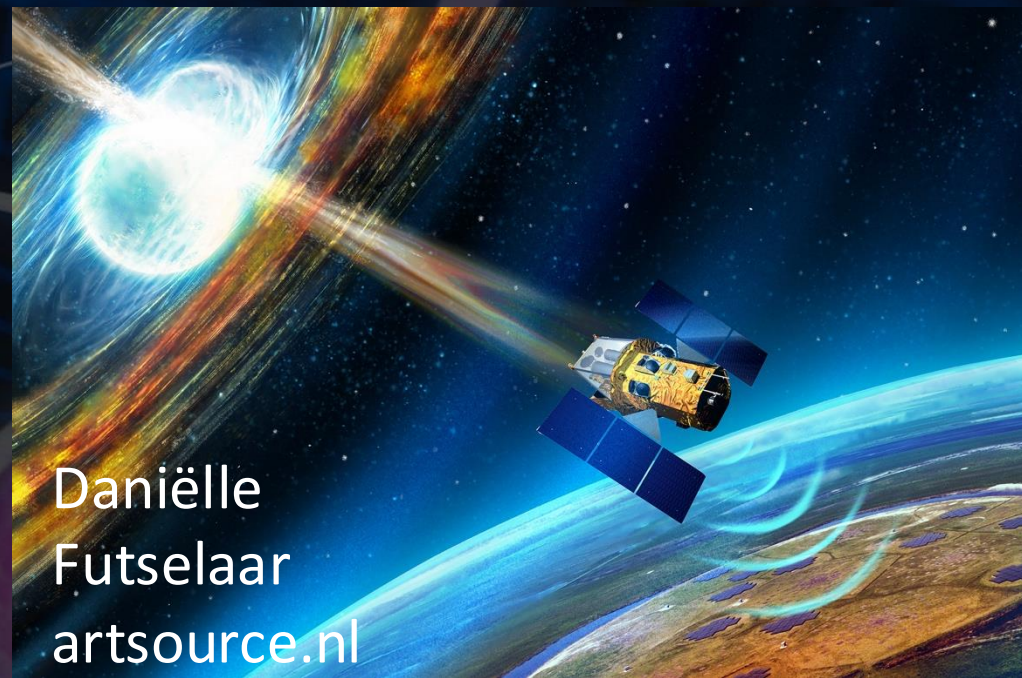
(Short) GRBs

Radio probes the afterglow emission

Prompt radio burst

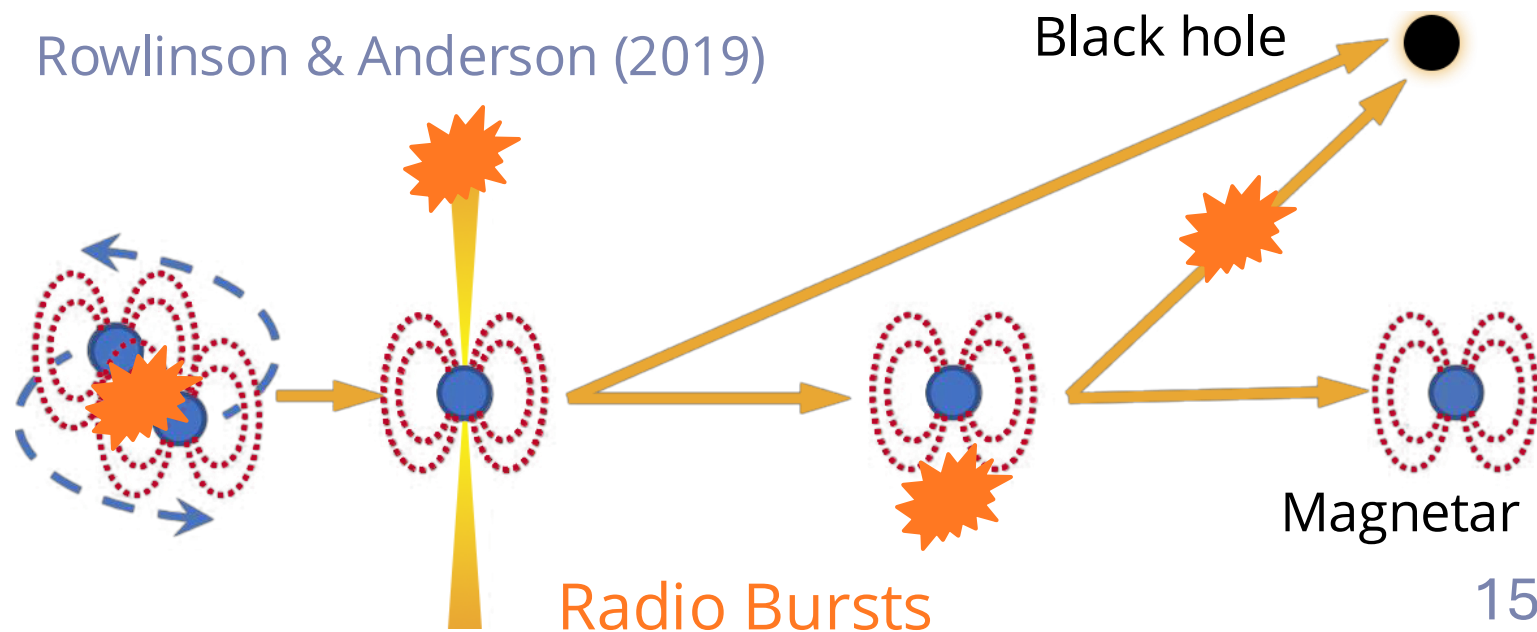
Rowlinson et al. (2024)

- Rapid response modes on LOFAR, ATCA, MWA
- All-sky monitors



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Futselaar
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Rowlinson & Anderson (2019)



Future

MW follow-up of radio transients:
Exploiting the next generation radio
telescopes (LOFAR 2.0, SKAO, ngVLA,
DSA-2000) requires robust algorithms
with short turn around between
detection and transient alert.

SKAO



Future

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Radio follow-up of MW transients:
Limited radio follow-up facilities, how to identify the most promising events?

SKAO



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Excited to learn from and collaborate with transient communities at other wavelengths!

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SKAO



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