

# Astrophysics software development: some perspectives, suggestions, and the write-only documentation phenomenon.

*Phil Evans*

*(University of Leicester)*

# Who am I / why am I giving this talk?

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# *Phil Evans*



## Key stats:

**Name:** (See above).

**Age:**  $4.4540 \times 10^{-8}$  Gyr.

**Rest mass:**  $3.6181 \times 10^{-29} M_{\odot}$  ( $R_s = 1.0672 \times 10^{-25}$  m).

**Average redshift:** 0 (despite the t-shirt...)

**Publications:** Yes.

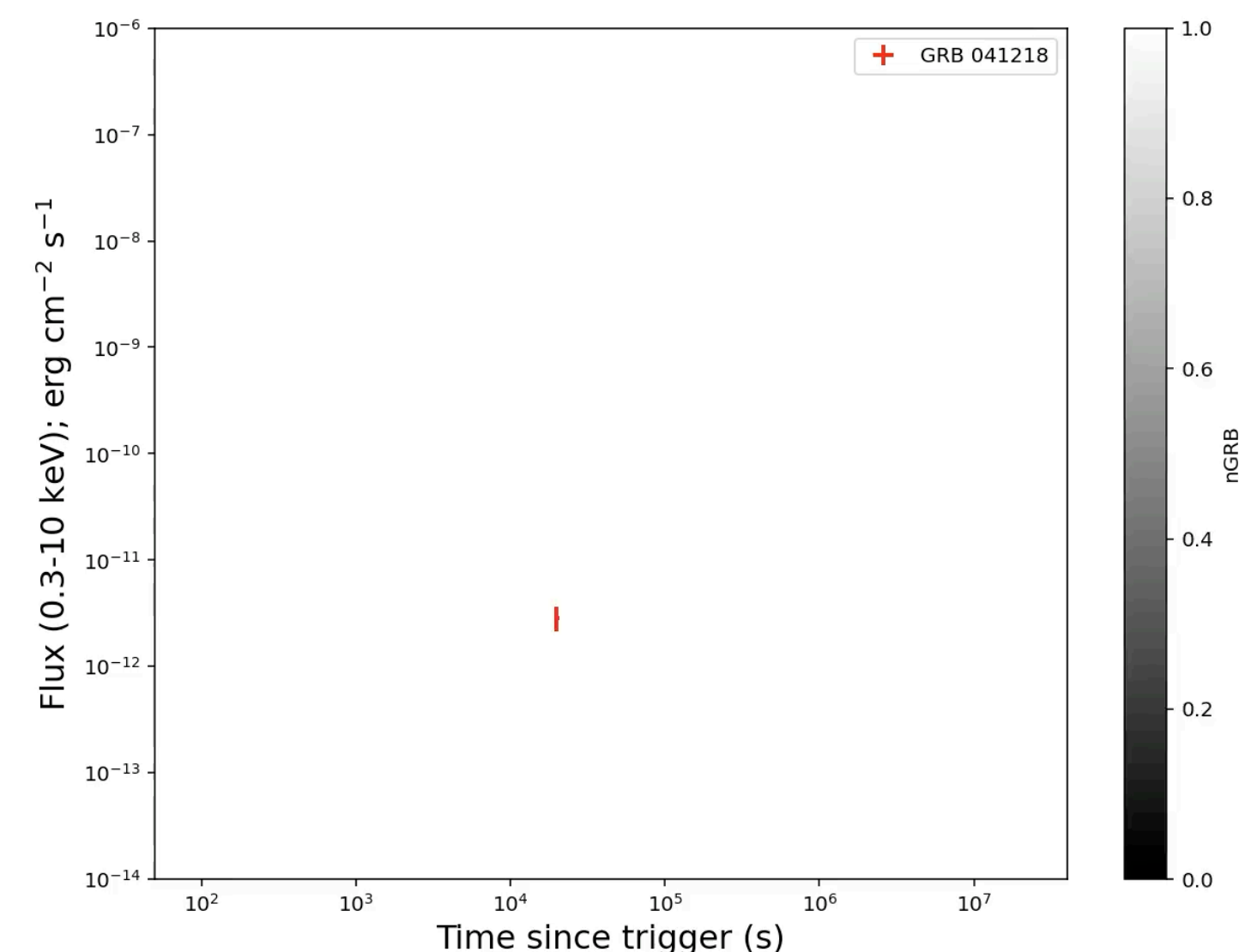
**Institute:**  UNIVERSITY OF  
LEICESTER

Responsible for the software behind [www.swift.ac.uk](http://www.swift.ac.uk)

Automatic and on-demand analysis of Swift-XRT data.

LSXPS “Living” catalogue and transient system.

Science interests: GRBs, transients, TDAMM.





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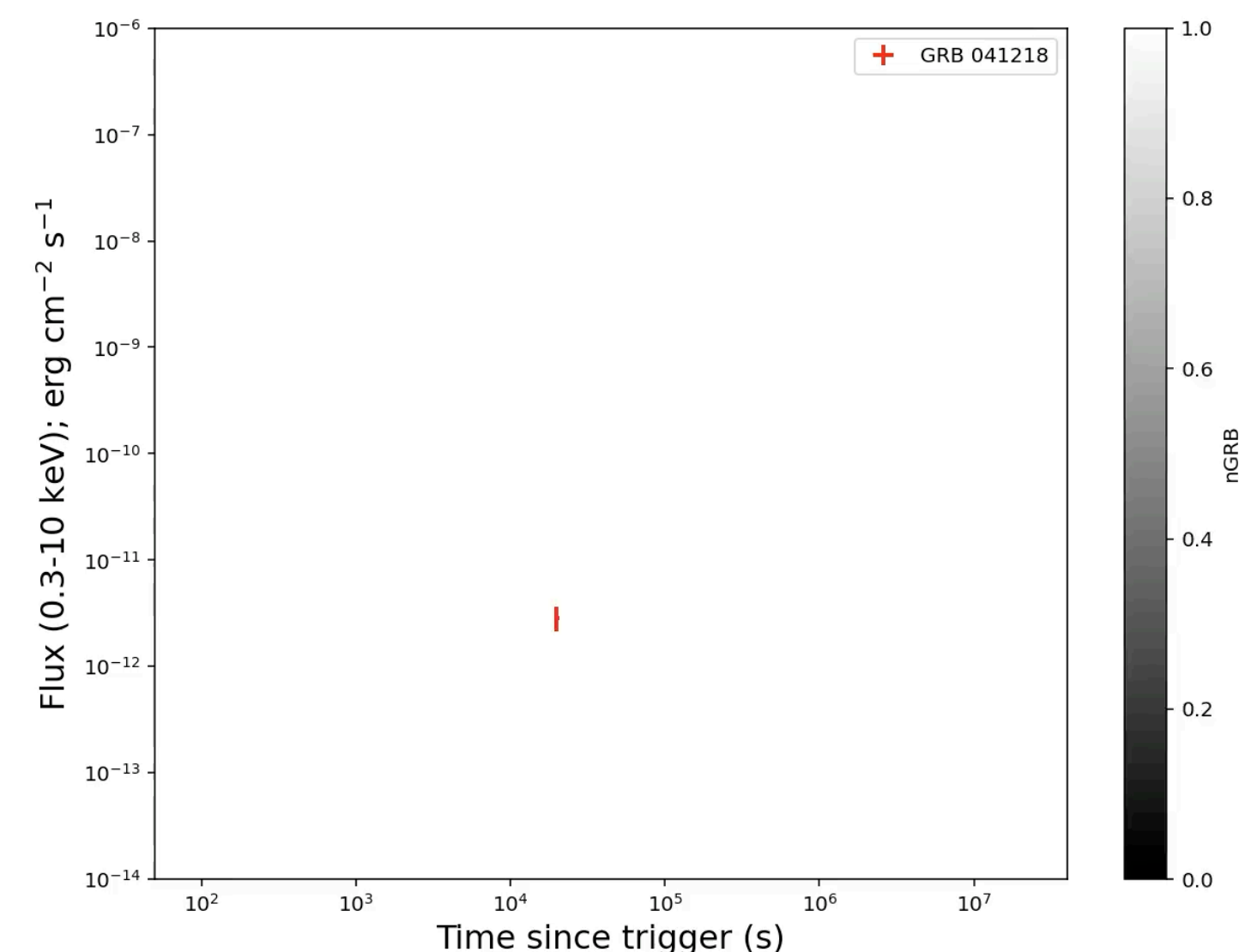
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# Who am I / why am I giving this talk?

*I'm an X-ray astronomer, who has been writing astrophysical software for 19+ years.*

Undergraduate degree in Astrophysics & Computer Science\*.

- **Swift:**

- 2006-2021: Data centre scientist for the UK *Swift* Science Data Centre.
  - From 2014 also been the sysadmin (can I put “full-stack” on my CV?).
- Since 2021: PI of the UoL Swift project.
- Since 2023: Lecturer (“Assistant Professor”), allegedly handed over UKSSDC duties.

- **NewAthena**

- Since 2017: technical lead of the *NewAthena* WFI Instrument Science Centre.
- (From 2027: lead of the above).

\* Computer science “to level 2 only”.

# My experience / context for my reflections

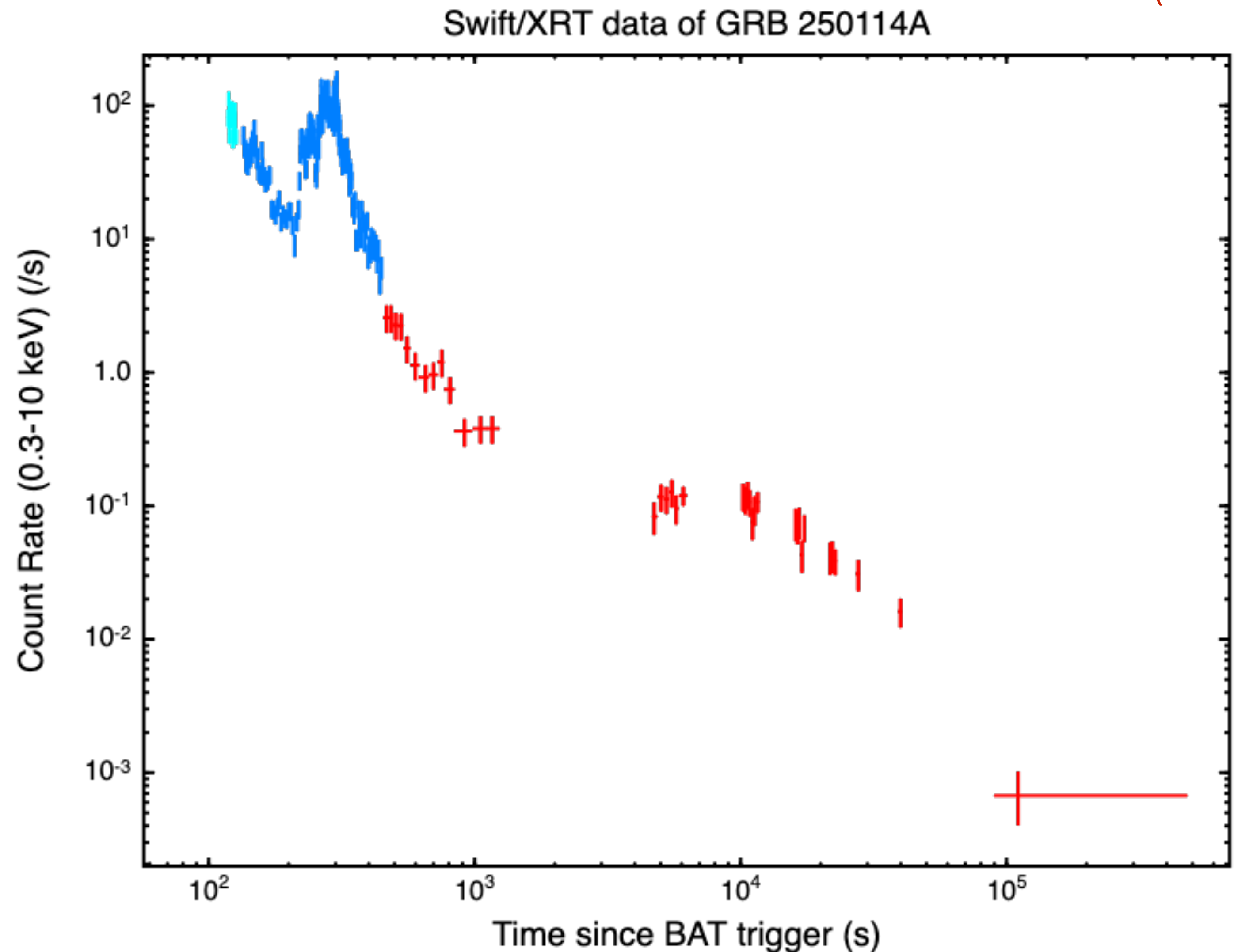
- Building automatic / on-demand analysis tools — *science ready* data products.
  - Optimised extraction, instrumental corrections applied.
  - In (most) cases, more accurate than hand-generation by experts.
- This is an increasingly-common need (e.g. VRO, SKA).



Evans+ (2007), (2009)

What we provide:

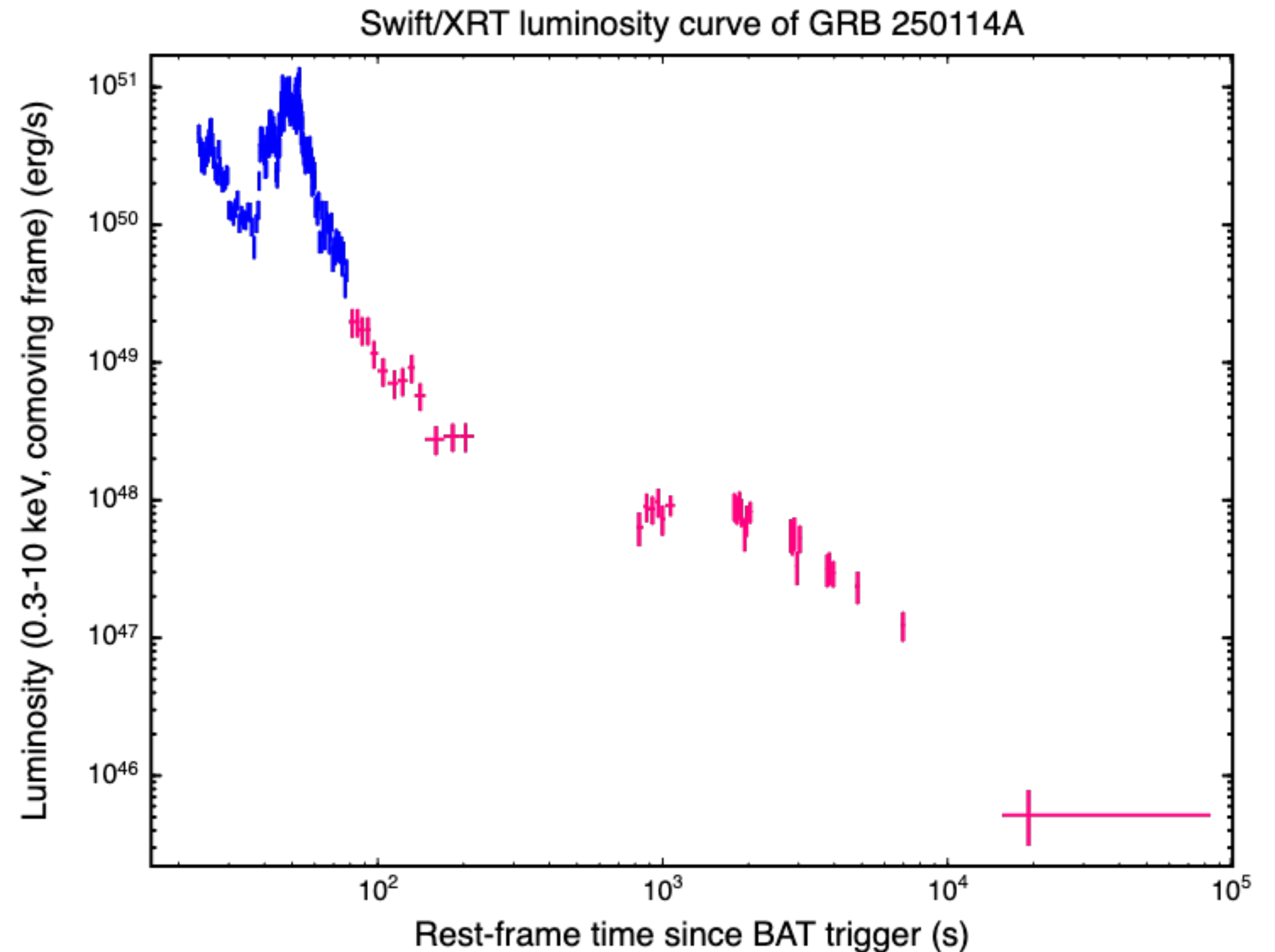
- GRBs — automated & public:
- Lightcurves



Evans+ (2007), (2009)

What we provide:

- GRBs — automated & public:
- Lightcurves  
**newish with luminosity**,  
 (when redshift known).

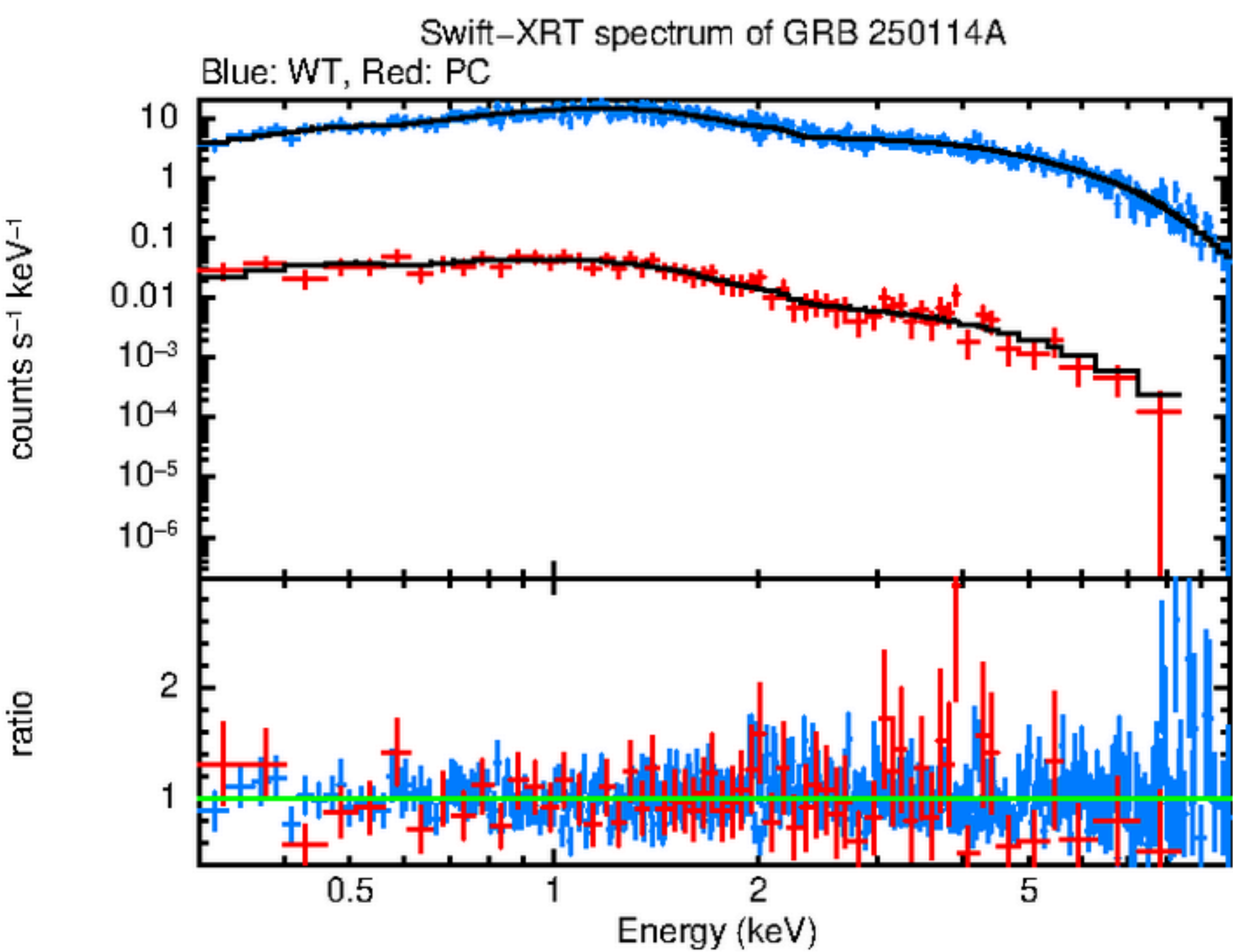




What we provide:

- GRBs — automated & public:
- Lightcurves,
- spectra (with fits)

Evans+ (2009)



WT mode. Mean photon arrival: T0+270 s

[Insert manually-determined values.](#)

N <sub>H</sub> (Galactic)	5.06 × 10 <sup>20</sup> cm <sup>-2</sup>
N <sub>H</sub> (intrinsic)	1.8 (+0.9, -0.9) × 10 <sup>22</sup> cm <sup>-2</sup>
z of absorber	4.732
Photon index	1.09 (+0.03, -0.03)
Flux (0.3-10 keV) (Observed)	2.19 (+0.06, -0.06) × 10 <sup>-9</sup> erg cm <sup>-2</sup> s <sup>-1</sup>
Flux (0.3-10 keV) (Unabsorbed)	2.28 (+0.05, -0.05) × 10 <sup>-9</sup> erg cm <sup>-2</sup> s <sup>-1</sup>

PC mode. Mean photon arrival: T0+8453 s

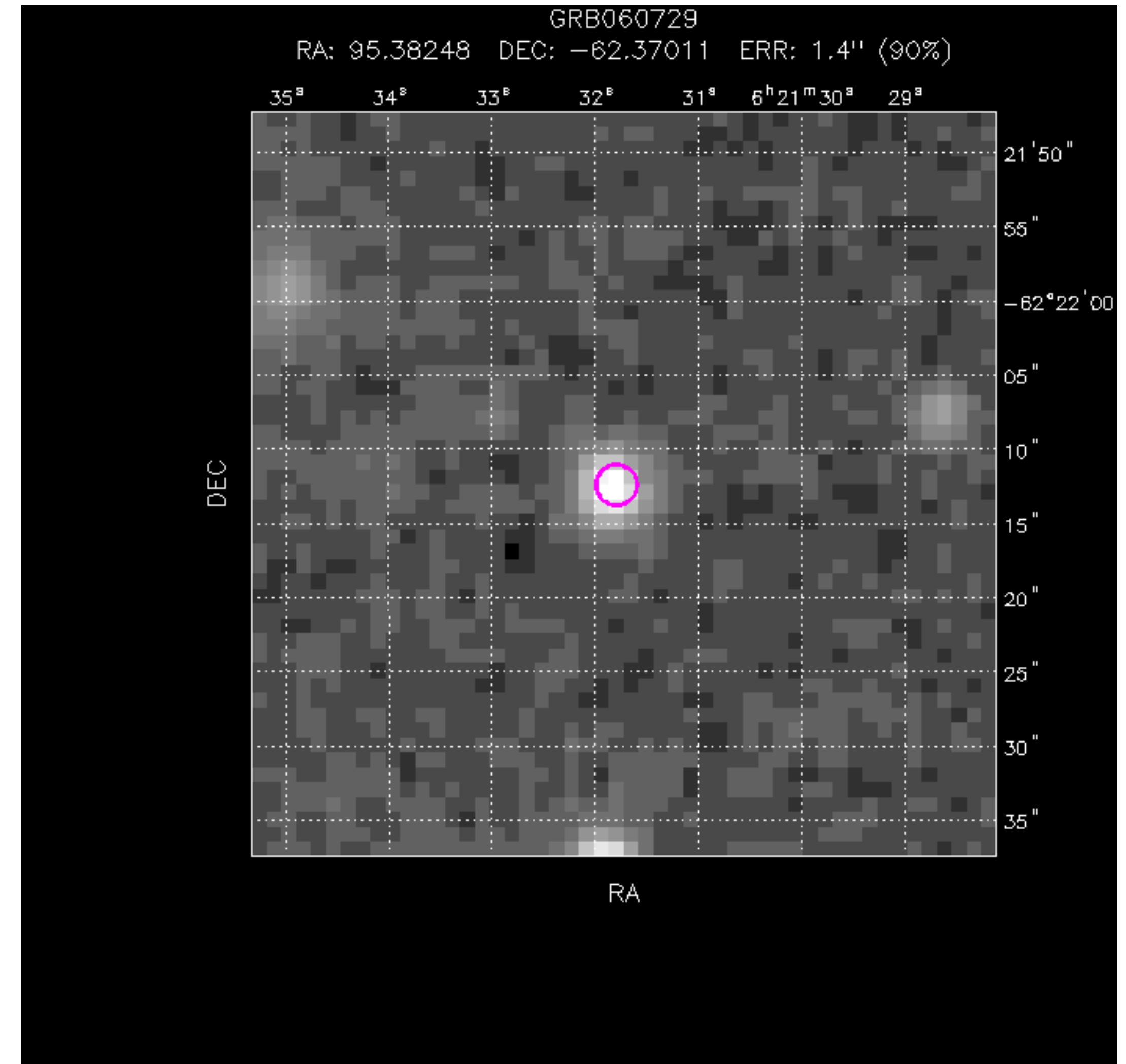
[Insert manually-determined values.](#)

N <sub>H</sub> (Galactic)	5.06 × 10 <sup>20</sup> cm <sup>-2</sup>
N <sub>H</sub> (intrinsic)	1.0 (+2.2, -1.0) × 10 <sup>22</sup> cm <sup>-2</sup>
z of absorber	4.732
Photon index	1.85 (+0.14, -0.13)
Flux (0.3-10 keV) (Observed)	4.3 (+0.4, -0.4) × 10 <sup>-12</sup> erg cm <sup>-2</sup> s <sup>-1</sup>
Flux (0.3-10 keV) (Unabsorbed)	4.9 (+0.4, -0.4) × 10 <sup>-12</sup> erg cm <sup>-2</sup> s <sup>-1</sup>

What we provide:

- GRBs — automated & public:
  - Lightcurves,
  - spectra (with fits),
  - (enhanced) positions

Goad+ (2007); Evans+ (2009)





Evans+ (2009)

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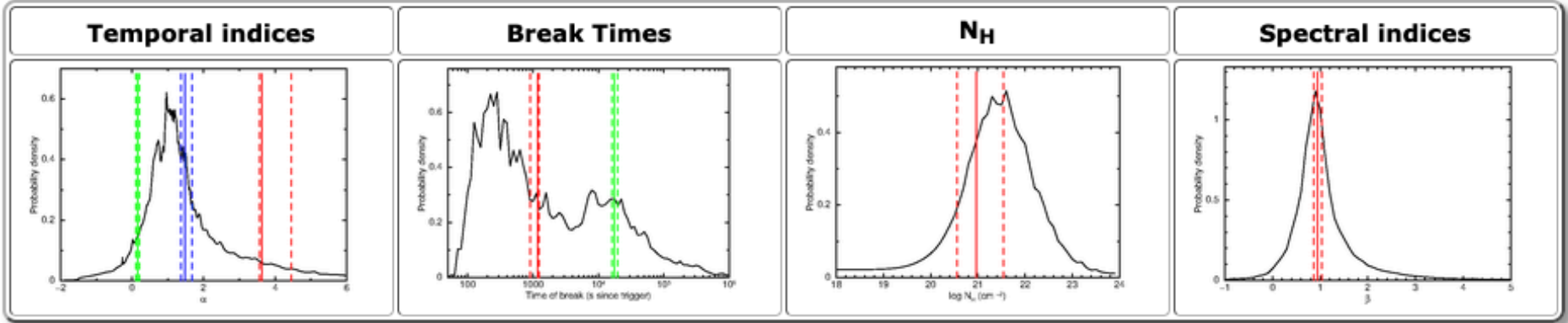
- GRBs — automated & public:
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- spectra (with fits),
- (enhanced) positions,
- live catalogue

Summary Information			
RA:	13h 14m 42.34s	LC breaks:	2
Dec:	+05° 01' 50.8''	LC type:	Canonical.
Err:	2.1''	$N_{H,intr.}$ :	$9.25 \times 10^{20} \text{ cm}^{-2}$
Gal long:	318.045	Redshift	<a href="#">2.151</a>
Gal lat:	67.217	$\Gamma$	1.9

$\alpha_1$	3.63 (+1.35, -0.10)
$T_{b,1}$	1186 (+81, -451) s
$\alpha_2$	0.16 (+0.09, -0.09)
$T_{b,2}$	$1.77 (+0.35, -0.25) \times 10^4 \text{ s}$
$\alpha_3$	1.48 (+0.32, -0.18)

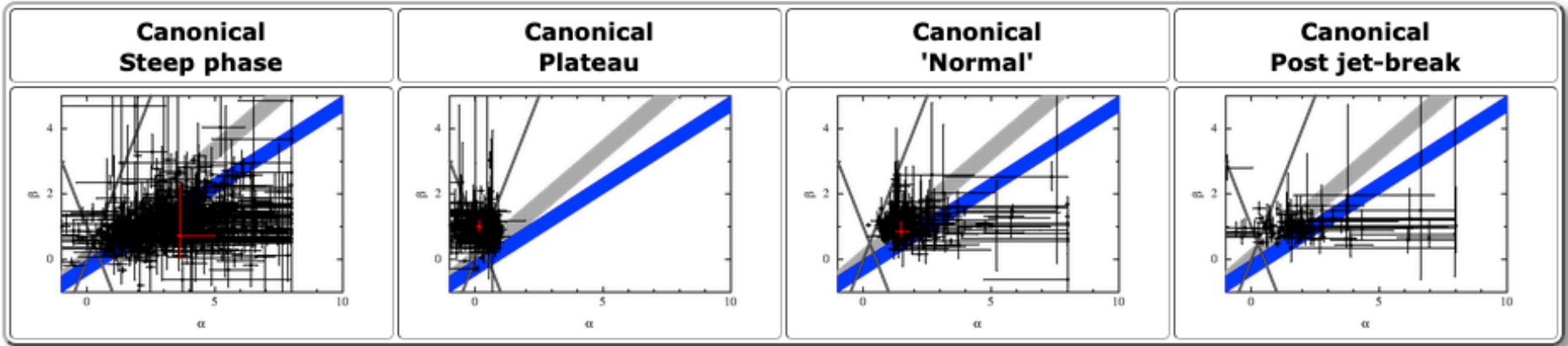
Nbreaks	$\chi^2$	dof	Comments	Image
<a href="#">0</a>	1138	47		
<a href="#">1</a>	222	45	Break required. (1.13e-14 % probability of chance improvement)	
<a href="#">2</a>	80.4	43	Break required. (3.14e-08 % probability of chance improvement). <b>This is the best fit.</b>	
<a href="#">3</a>	54.9	41	Break significance <4- $\sigma$ (0.04 % probability of chance improvement)	

GRB	RA (J2000)	Dec (J2000)	Pos. Err (", radius)	#Breaks	Light curve classification	$N_{H,gal} (\times 10^{22} \text{ cm}^{-2})$	Redshift
<a href="#">Swift J164449.3+573451</a>	16h 44m 49.95s	+57° 34' 59.7"	1.4	0	TBD	$1.75 \times 10^{-2}$	<a href="#">0.354</a>
<a href="#">SGR 1830-0645</a>	18h 30m 41.70s	-06° 45' 17.2"	1.4	0	No breaks	1.1	
<a href="#">GRB 250129A</a>	13h 14m 42.34s	+05° 01' 50.8"	2.1	2	Canonical	$2.43 \times 10^{-2}$	<a href="#">2.151</a>
<a href="#">GRB 250128B</a>	15h 25m 42.10s	-00° 32' 21.4"	2.7	0	No breaks	$8.00 \times 10^{-2}$	
<a href="#">GRB 250114B</a>	03h 15m 26.10s	-02° 17' 38.0"	7.1	0	TBD	$7.44 \times 10^{-2}$	
<a href="#">GRB 250114A</a>	11h 06m 18.29s	-15° 06' 41.5"	2.0	2	Canonical	$5.06 \times 10^{-2}$	<a href="#">4.732</a>
<a href="#">GRB 250108B</a>	13h 25m 18.38s	+25° 36' 54.8"	2.0	1	One-break	$1.16 \times 10^{-2}$	
<a href="#">GRB 250108A</a>	13h 51m 41.05s	+26° 12' 59.4"	3.5	0	No breaks	$1.36 \times 10^{-2}$	
<a href="#">GRB 250106A</a>	07h 48m 57.30s	+63° 48' 44.7"	3.0	0	No breaks	$4.48 \times 10^{-2}$	
<a href="#">GRB 250103B</a>	03h 38m 39.43s	-33° 44' 55.0"	3.7	0	No breaks	$9.99 \times 10^{-3}$	<a href="#">1.416</a>



#### Closure Relationships

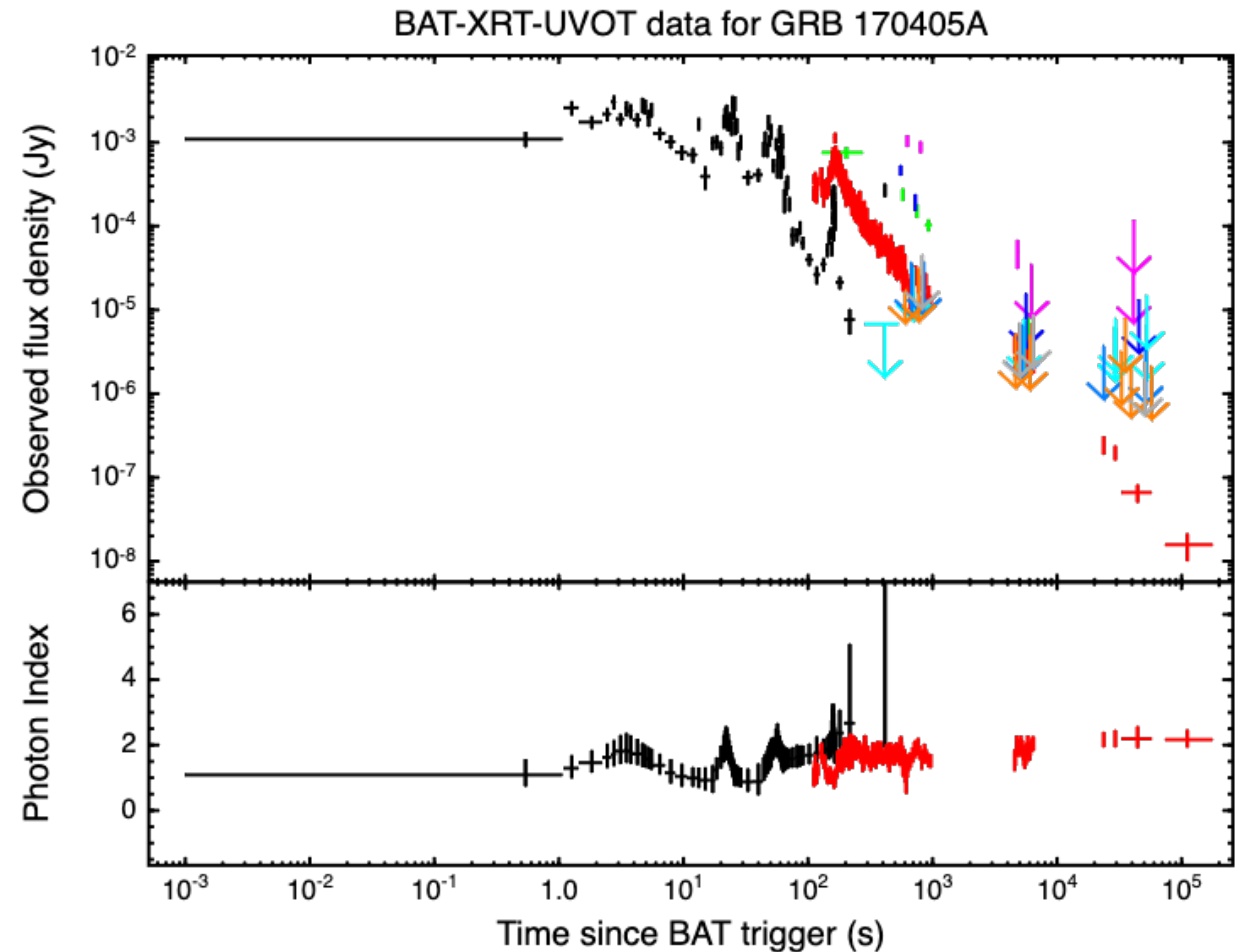
The plots below show  $(\alpha, \beta)$  scatter plots for each phase of all Canonical light curves. GRB 250129A is shown in colour; where multiple points appear in a given phase (e.g. WT and PC mode spectra for that phase), the different colours refer to different spectra.  $\alpha$  and  $\beta$  are the temporal and spectral energy power-law indices respectively. The shaded regions mark those allowed by the standard closure relationships. These figures correspond to Fig. 10 of [Evans et al. \(2009\)](#). See [the closure relationship documentation](#) for more information.



Evans+ (2010)

What we provide:

- GRBs — automated & public:
  - Lightcurves,
  - spectra (with fits),
  - (enhanced) positions,
  - live catalogue,
  - images
- burst analyser (multi-instrument)





What we provide:

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  - 1SXPS (2014)  
<https://www.swift.ac.uk/1SXPS>
  - 2SXPS (2020)  
<https://www.swift.ac.uk/2SXPS>
  - LSXPS (2023)  
<https://www.swift.ac.uk/LSXPS>

<b>Data included</b>	2005-01-01 — 2025-10-06 (408 Ms)
<b>Sky coverage</b>	5,728 deg <sup>2</sup>
<b>Detections</b>	2,760,293
<b>Unique sources</b>	346,566
<b>Variable sources<sup>2</sup></b>	146,390
<b>Uncatalogued sources<sup>2</sup></b>	131,538

Evans+ (2023a)

What we provide:

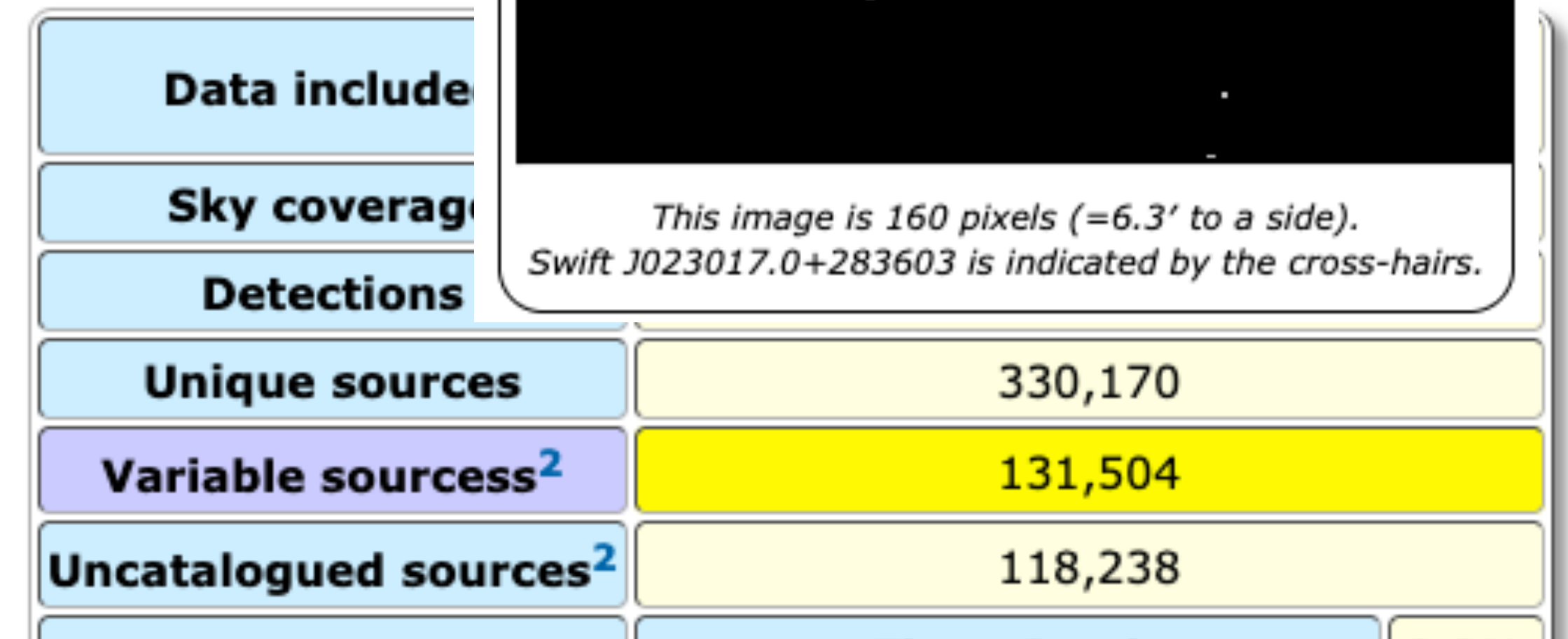
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  - LSXPS (2023) - **with real-time transient detector!**  
<https://www.swift.ac.uk/LSXPS>

Data included	2005-01-01 — 2025-01-25 (395 Ms)
Sky coverage	5,623 square degrees
Detections	2,537,606
Unique sources	330,170
Variable sources <sup>2</sup>	131,504
Uncatalogued sources <sup>2</sup>	118,238



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[https://www.swift.ac.uk/user\\_objects/](https://www.swift.ac.uk/user_objects/)  
<https://www.swift.ac.uk/API/swifttools.ukssdc> (via pypi)
- or nothing — source detection.

**Select products**

To reduce the load on our servers, please select only the independent products you require.

Light curve?☒ Spectrum?☒ Position?☐ Image?☐ Source detection?☐

---

**Object details**

\*Name:

\*Target ID:

Time zero:

All input times since this? ☐

\*Coordinates:

**Global options**

E-mail address:

Remember me ☐

Email me when complete? ☐

\*Try to centroid?

\*Centroid method:

**Light curve details**

Binning Method

\*Bin length (s):

WT:  PC:

\*Hardness ratio bin length (s):

Same as main curve ☐

WT:  PC:

Min fractional exposure

Minimum sigma:

Allow upper limits?

\* Point-sources, seen by XRT; some extreme cases may give imperfect results (although work surprisingly\*\* often).

\*\* Well, I'm always surprised when it works.



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<https://www.swift.ac.uk/API/>  
[swifttools.ukssdc](#) (via pypi)
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swifttools.ukssdc.xrt\_prods demo

```
[1]: 1 %load_ext autoreload
      2 %autoreload 2

[2]: 1 #from testtools.xrt_prods import XRTProductRequest
      2 from swifttools.xrt_prods import XRTProductRequest
      3 import pandas as pd
      4 me = 'pae9@leicester.ac.uk'
```

Submitting a new job

```
[3]: 1 myReq = XRTProductRequest(me, silent=False)
      2 myReq.setGlobalPars(getTargs=True,
      3                     centroid=True,
      4                     getCoords=True,
      5                     getT0=True,
      6                     posErr=0.1,
      7                     name='GK Per',
      8                     useSXPS=False
      9                     )
      10 #myReq.addImage(energies='0.3-10,0.3-1.5,1.51-10', whichData='all')
      11 myReq.addLightCurve(binMeth='obsid', srcrad=57)
      12 myReq.isValid()
```

Successfully created a light curve  
OK, setting binMeth = obsid  
OK, setting srcrad = 57

```
[3]: (True, '')
```

```
[4]: 1 myReq.submit()
```

Job submitted OK, with ID: 196342  
Also setting lc useObs = None, because whichData = all

```
[4]: True
```

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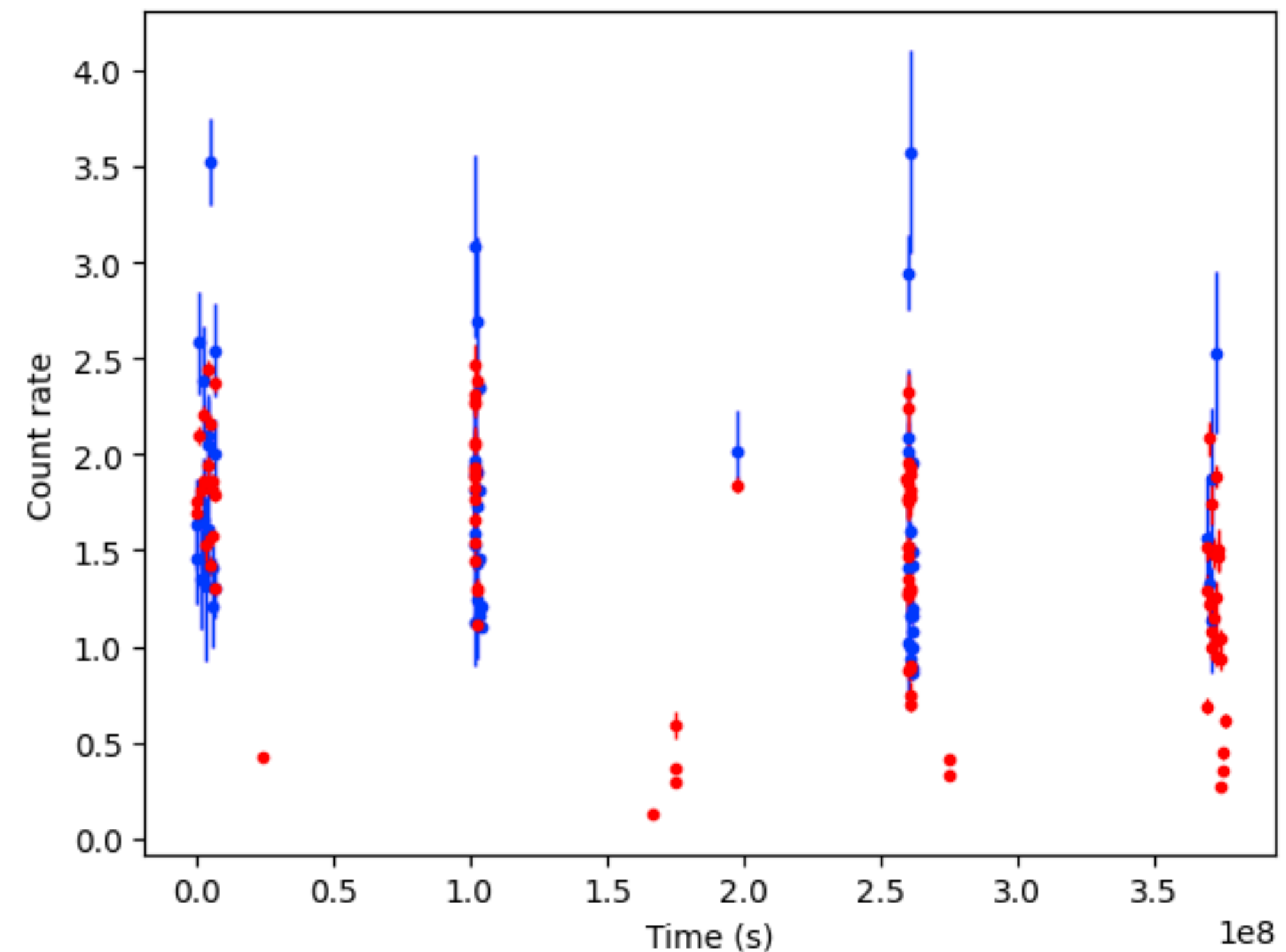
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<https://www.swift.ac.uk/API/>  
 swifttools.ukssdc (via pypi)
- or nothing — source detection.

swifttools.ukssdc.xrt\_prods demo

```
[49]: 1 myLC = r.retrieveLightCurve(returnData=True, incbad='yes', nosys='no')
      2 swifttools.ukssdc.plotLightCurve(myLC, whichCurves=("WT_incbad", "PC_incbad"), xlabel="Time (s)
      3
```



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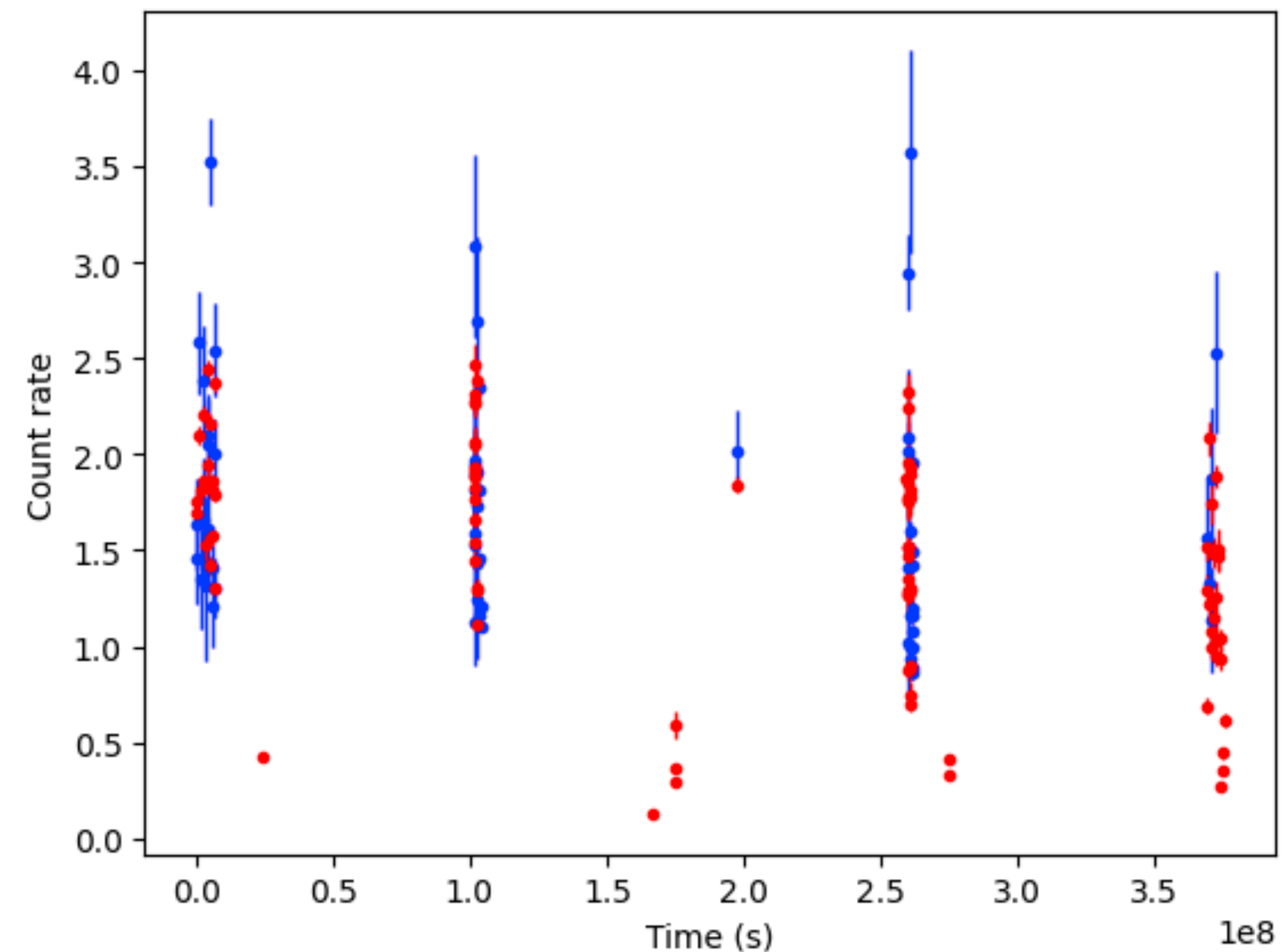
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Produces:

~1400 products/month (web)  
 ~2800 products/month (API)

swifttools.ukssdc.xrt\_prods demo

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# First warning: tech changes!

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    - “Broadband” internet was a luxury (justified by on-call support).
    - Typical server: dual-core CPU, 4GB RAM. External RAID box for the archive.
    - “Flash” and Java applets were mainstream.
    - People thought IDL was a good idea.
      - (Though some people were trying out ‘Python’)
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  - 2008: first Android phone goes on sale; Python 3.0 released.
  - 2012: 11 UK Cities get 4G mobile internet; automated storage tiering becomes mainstream.
  - 2013: Docker released. Astropy first release. Java applet support wanes.
  - 2014: Kubernetes released.
  - 2019: 6 UK Cities get 5G mobile internet; Java applets formally deprecated.
  - 2020: Python 2 end of life.
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  - 2038: *NewAthena* launches.



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2006

19 years

Now

13 years

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• Design your **code** to be **infrastructure-agnostic**.

• Use modules/classes etc. for all infrastructure-related tasks.

• e.g. so that changing from running a subprocess in a shell, to in a docker, to via K8s or via a batch processing system means changing one module, not 68 codes.

- “Broadband”
- Typical
- “Flash”
- People
- (Tho

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# Second warning: scope changes!

- On joining the UKSSDC, I was shown a grant proposal which included a pledge to write a tool to generate XRT light curves of GRBs...
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- “Could people use these things for non-GRBs...? (2009)



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- “It would be nice to apply these tools to the prompt data...”
- “Could people use these things for non-GRBs...? (2009)
- Multi-messenger things started happening... (2011)

# Unforeseen growth...





# Unforeseen growth...





# Unforeseen growth...





# Unforeseen growth...





# Unforeseen growth...





# When the science landscape changes... adapt!

---

Science can change quickly — embrace the unforeseen.

- For Swift, this started with IceCube neutrino doublets...
- ... then HE neutrino triggers,
- ... and ANTARES alerts.
- KM3NeT “tomorrow”.
- neutrinos (<https://www.swift.ac.uk/neutrino/>)






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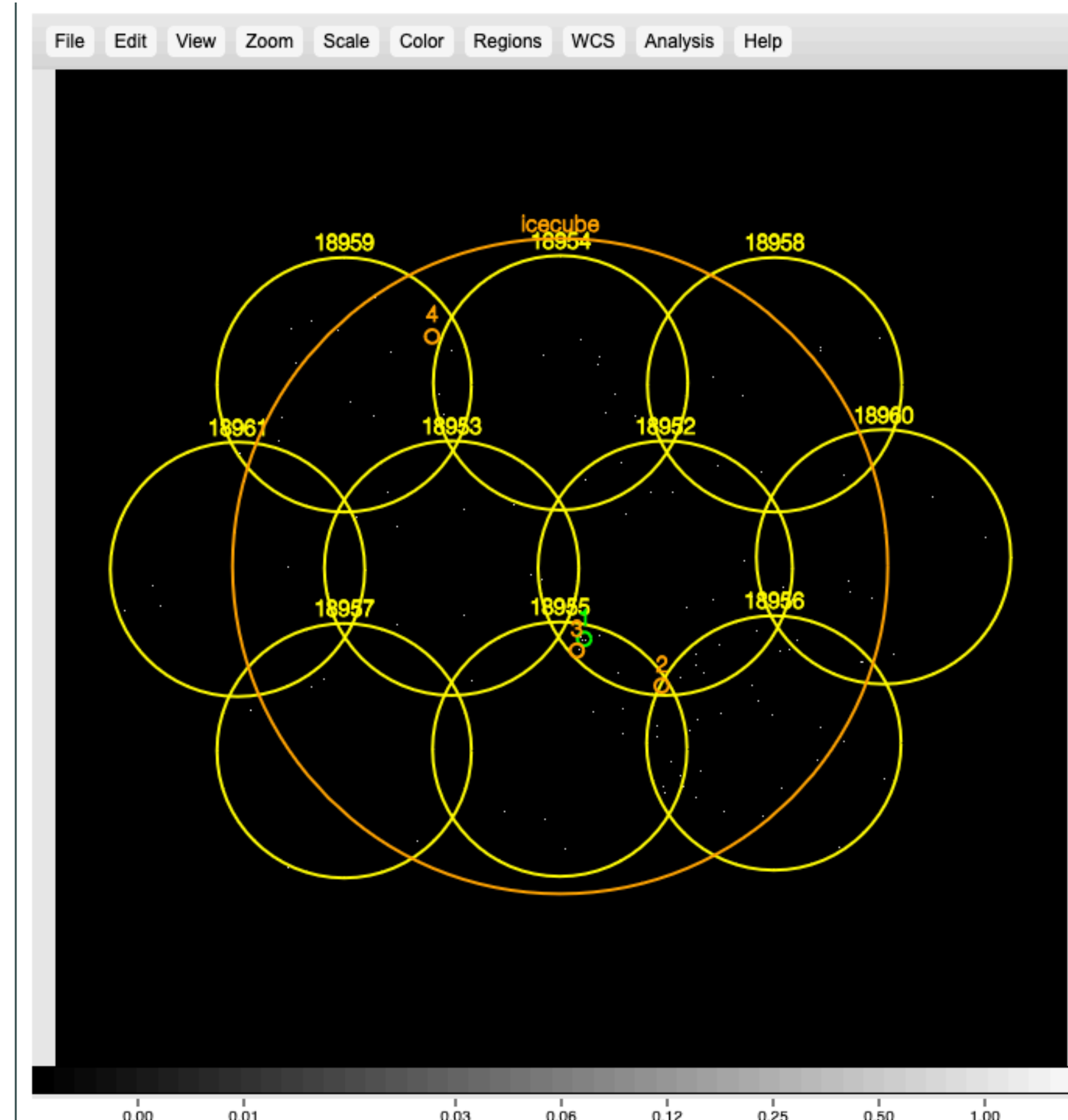
Evans+ (2015); Adrián-Martínez+ (2016);  
Aartsen+ (2017); Keivani+ (2018)

So far 4 sources have been found in the XRT data. 3 sources match catalogued sources.

Coordinate type: Sexagesimal

#	RA (J2000)	Dec (J2000)	Err <sub>90</sub> <sup>1</sup>	Detection Flag <sup>2</sup>	Rank <sup>3</sup>	Dist from IceCube pos	Exposure	Notes	Vizier
<a href="#">1</a> [details]	10h 56m 12.04s	+05° 15' 41.3"	8.4'' *	Good	4 (K)	7.5'	1.0 ks	Matches SIMBAD source: <a href="#">SDSS J105612.22+051544.6</a> , 0.8" away. Matches known X-ray source: 2SXPS J105612.0+051546 (from 2SXPS), 2.9" away.	
<a href="#">2</a> [details]	10h 55m 42.04s	+05° 11' 12.1"	7.0''	Poor	3 (U)	15.2'	1.0 ks	Matches SIMBAD source: <a href="#">NVSS J105542+051112</a> , 2.1" away.	
<a href="#">3</a> [details]	10h 56m 15.13s	+05° 14' 37.7"	8.2''	Poor	3 (U)	8.3'	1.0 ks		

Matches SIMBAD source: [PM](#)



# When the science landscape changes... adapt!

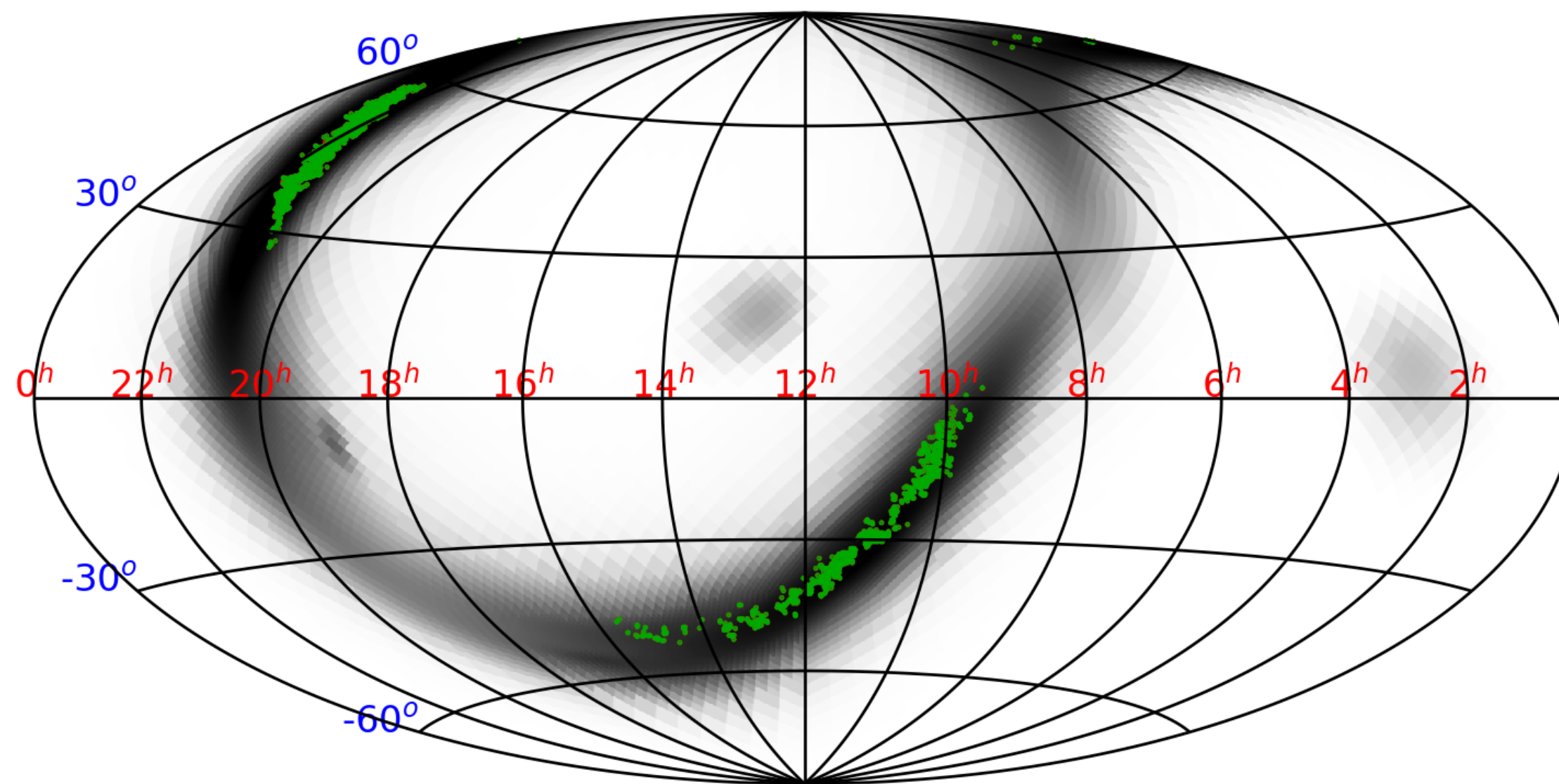
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Science can change quickly — embrace the unforeseen.

- For Swift, this started with IceCube neutrino doublets...
- ... and FRBs and then of course, Gravitational Waves.
- (<https://www.swift.ac.uk/LVC/>)

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## GW details

Trigger Date:	2017-02-27 at 18:57:31.000 UT
Trigger type:	Compact Binary Coalescence.
Distance:	193 ± 61 Mpc
P <sub>ContainsNeutronStar</sub> :	1
P <sub>EMBright</sub> :	1
False alarm rate:	1.426e-7 Hz = 1/81 d
GraceDB:	<a href="#">GraceDB</a>
P <sub>BAT</sub> (orig map)	0.111299 (raw) 0.085083 (convolved)

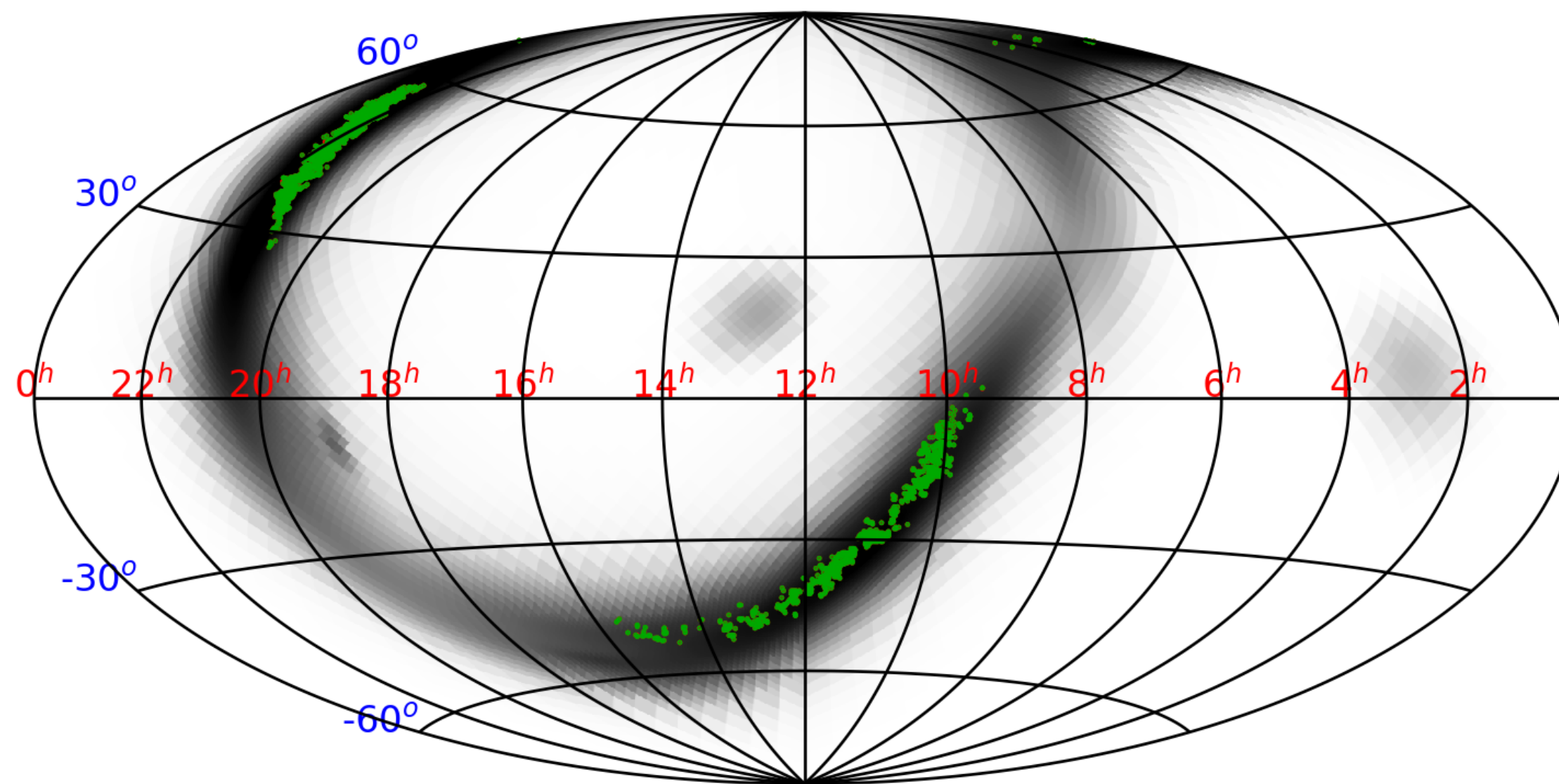
## XRT details

First observation:	T <sub>0</sub> + 264.7 min
Fields planned:	1414 [ <a href="#">Search</a> ]
Fields observed:	1408
TargetID range:	7006400—7007830
LVC Prob observed	0.089 (raw), 0.23 (conv)
LVC Prob to observe	7.11e-4 (raw), 1.35e-3 (conv) [LALInference]
XRT sources:	59
Candidate afterglows:	0
Possible afterglows:	3
Unlikely afterglows:	16
Known X-ray sources:	40



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# “Oh yay! I make some better tools”





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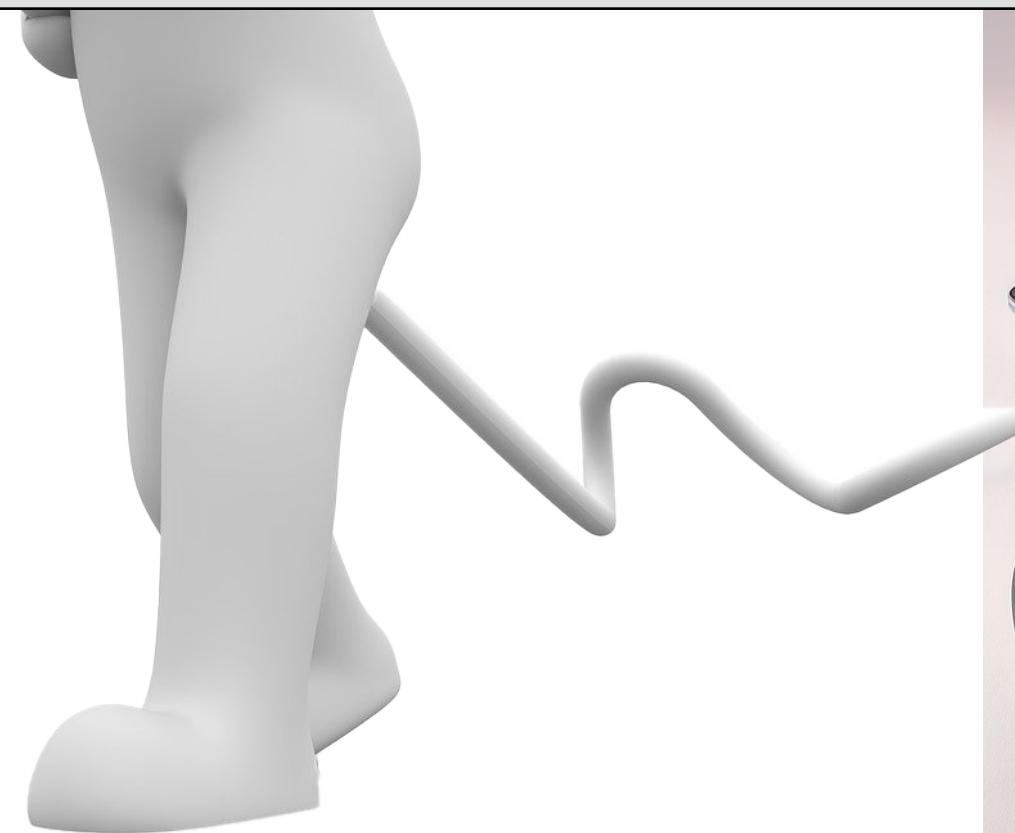




# “Oh yay! I make some better tools”



- Write good developer-facing documentation **from the start, all the time.**
- Build and maintain a **full** test harness.
- “What happens if I delete this probably obsolete code?”





# If only I could start again...

---

- Lead the UK *Swift* Science Data Centre.
- Technical lead of the *NewAthena* WFI Instrument Science Centre.



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**What have I learned?**

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## What have I learned?

- Design your *code* to be *infrastructure*-agnostic.
- Use modules/classes etc. for all infrastructure-related tasks.
- e.g. so that changing from running a subprocess in a shell, to in a docker, to via K8s or via a batch processing system means changing one module, not 68 codes.

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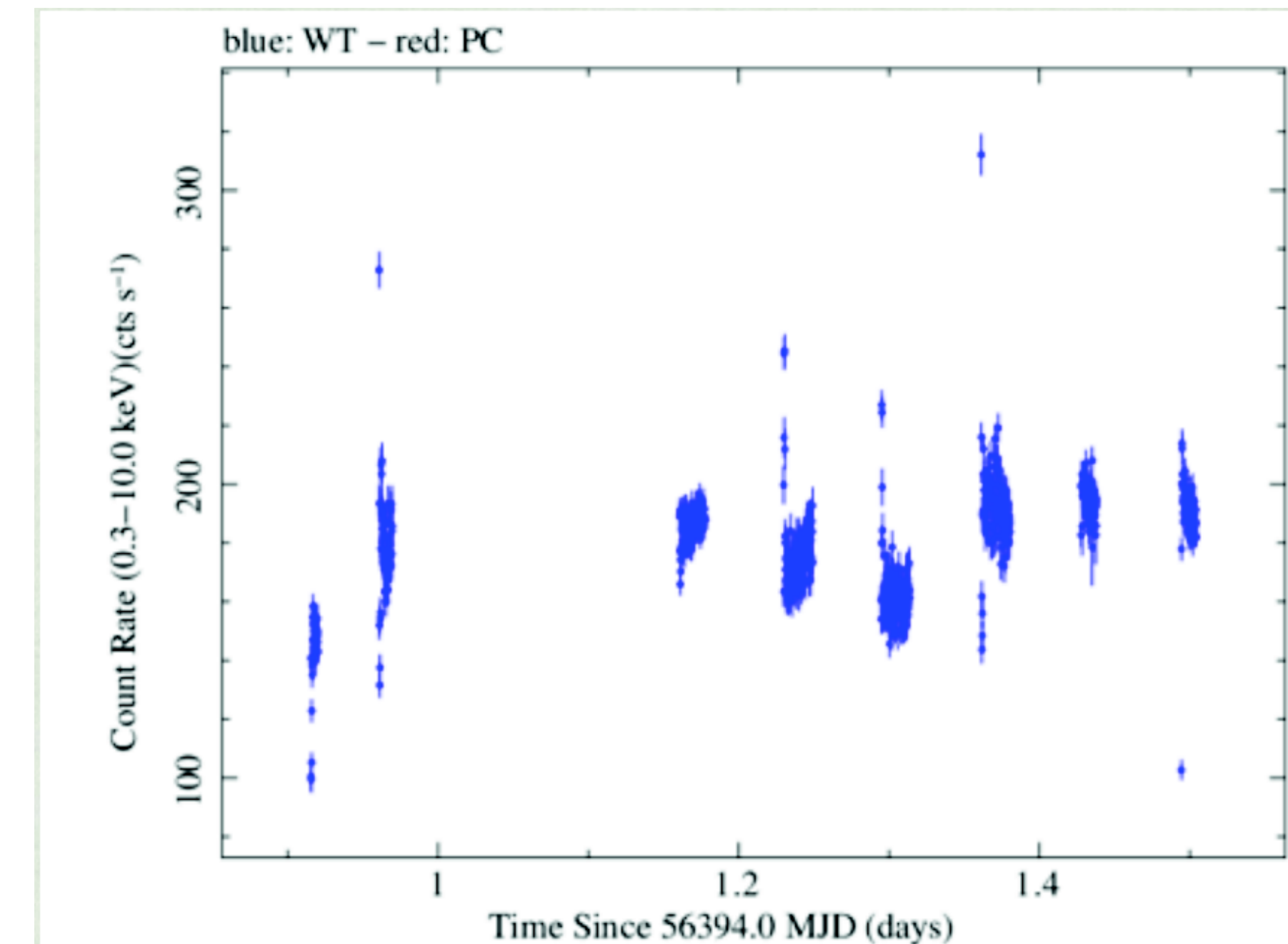
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- Maximising the science return from the project.
- **Creating software / systems that allow scientists to focus on science.**

# The starting point: what is the goal?

- Maximising the science return from the project.
- **Creating software / systems that allow scientists to focus on science.**
- Data analysis can be complex, intricate and difficult.
- It can require pretty esoteric knowledge.
- Errors are NOT always easy to spot.
- No-one wants bad analysis → bad science to be published.

We should make the bar to using our data as low as possible — **but not zero?**



# How do we achieve this?

- User focus! (I know, users are annoying...)
- “Full stack” dev-scientist is increasingly unfeasible...
  - ...but development by users is advantageous...  
(Ever tried inserting cross references into a Word document?)
  - ...and can someone please explain that to our funders?
- Agile approaches (user stories, continually working product, regular user checkins) ... but are they all agile?
- A user community regularly interacting with the developers is important.
- What users want  $\neq$  what we think they want...
- ... and  $\neq$  what each other wants.



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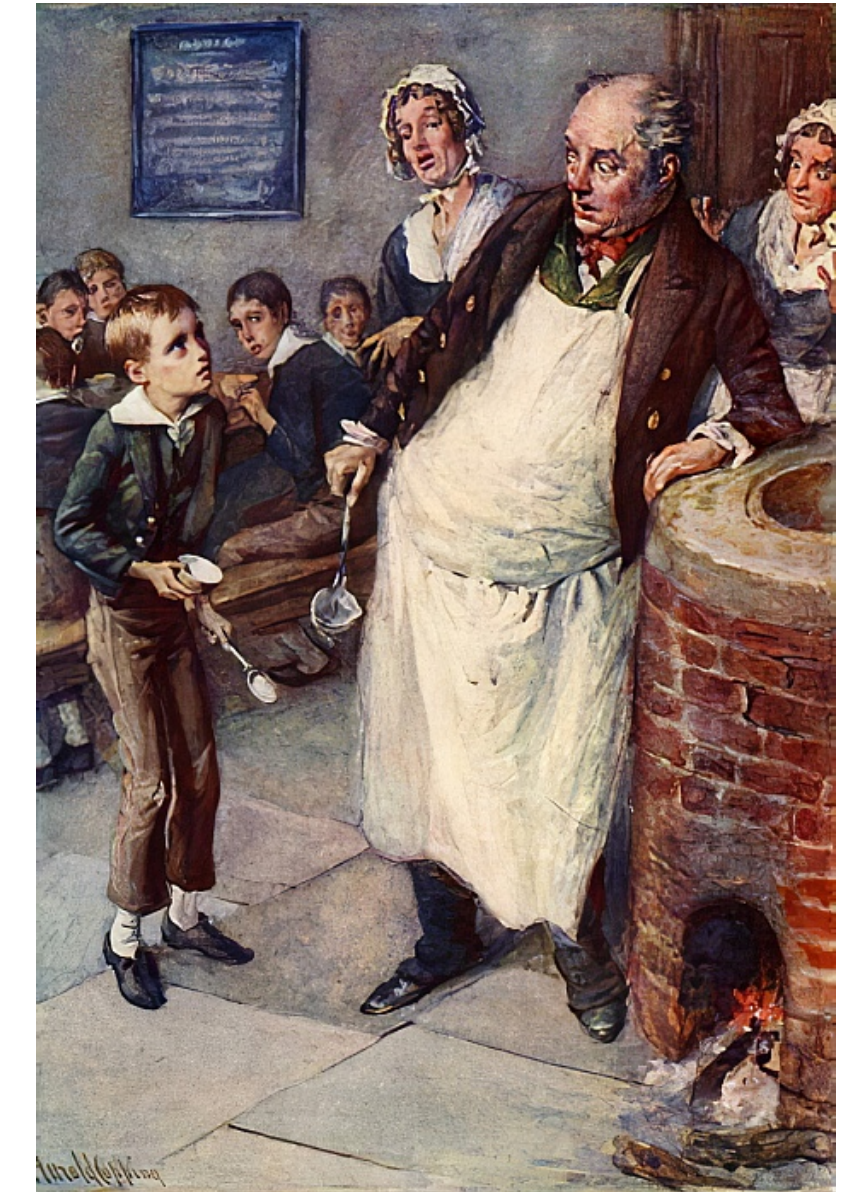


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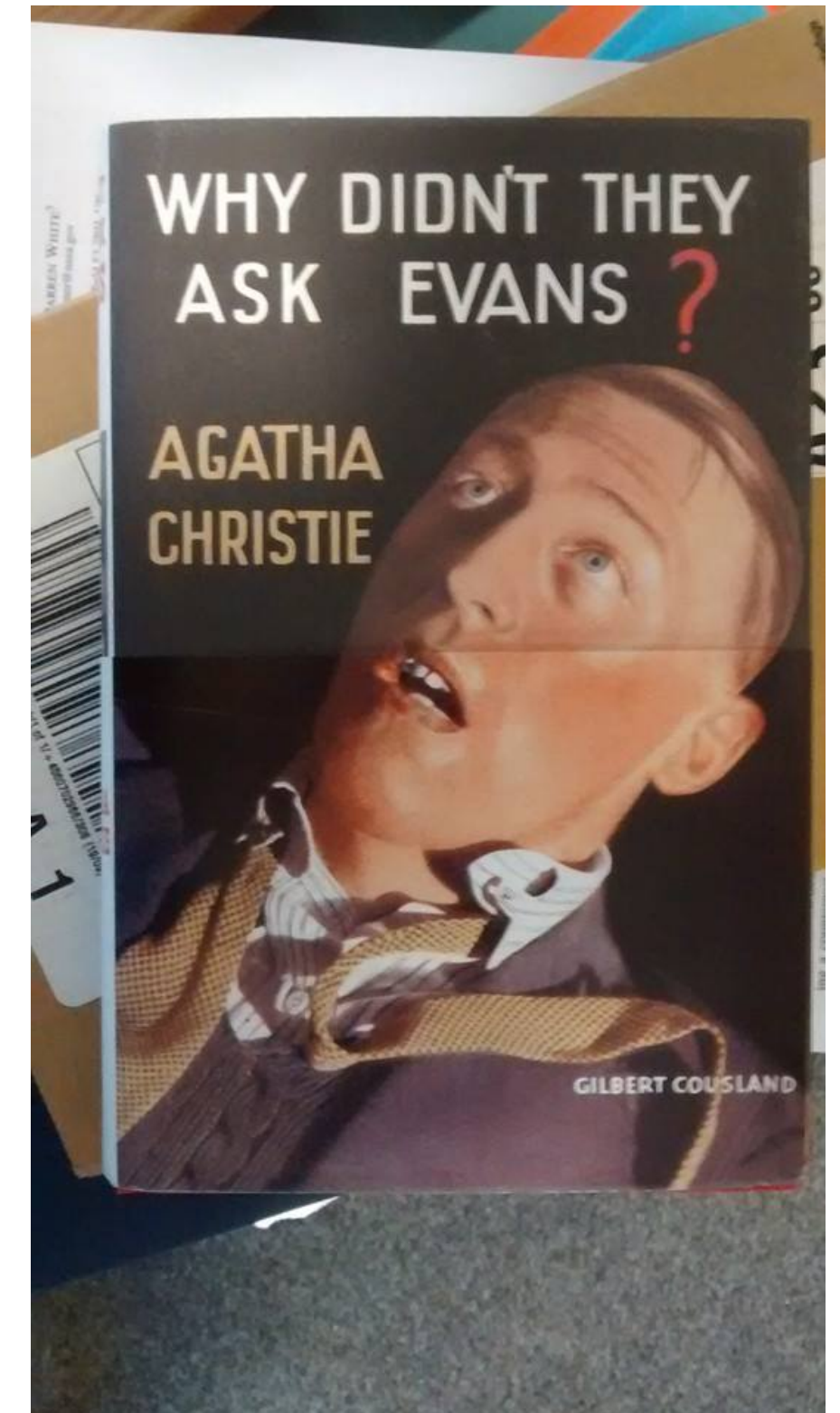
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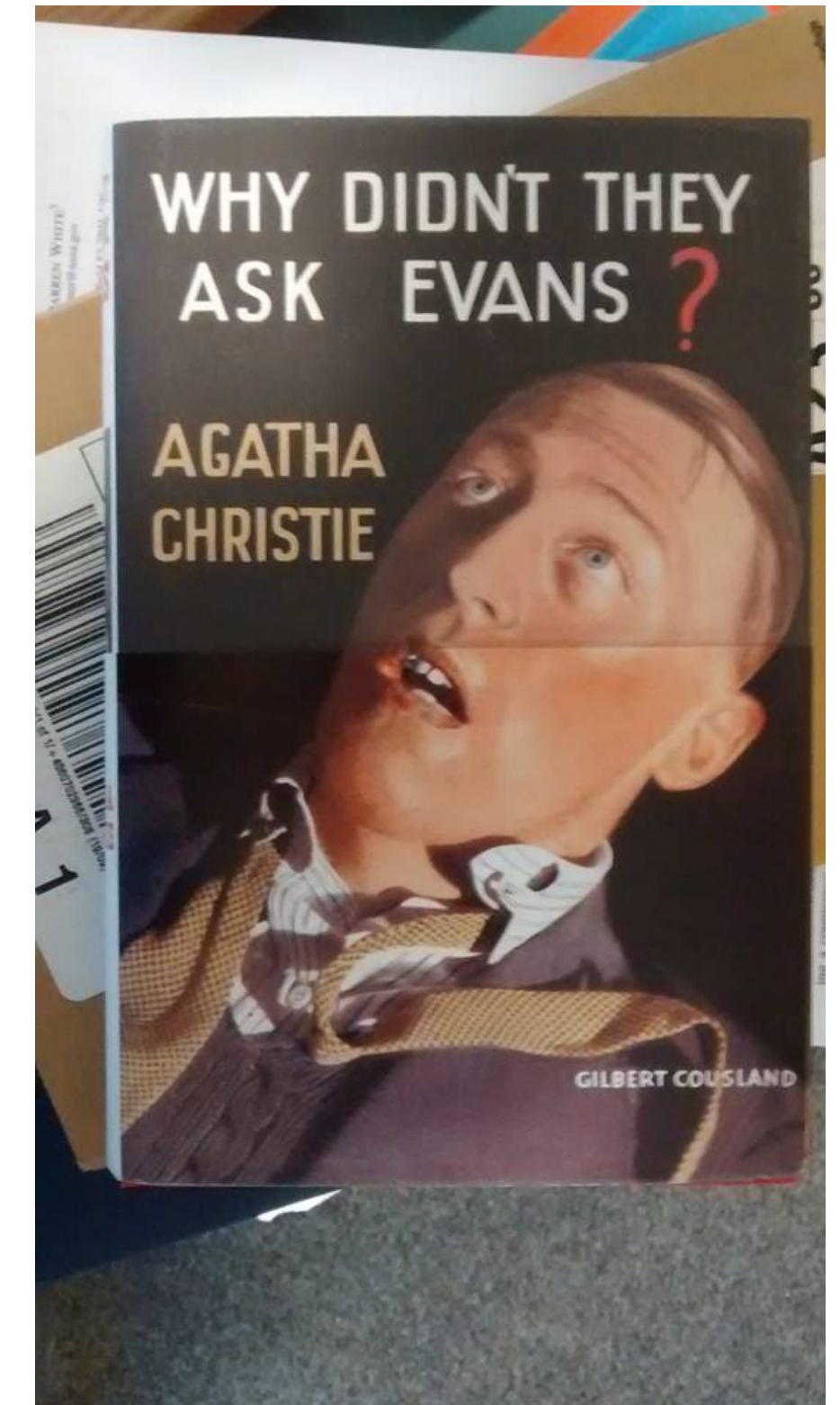
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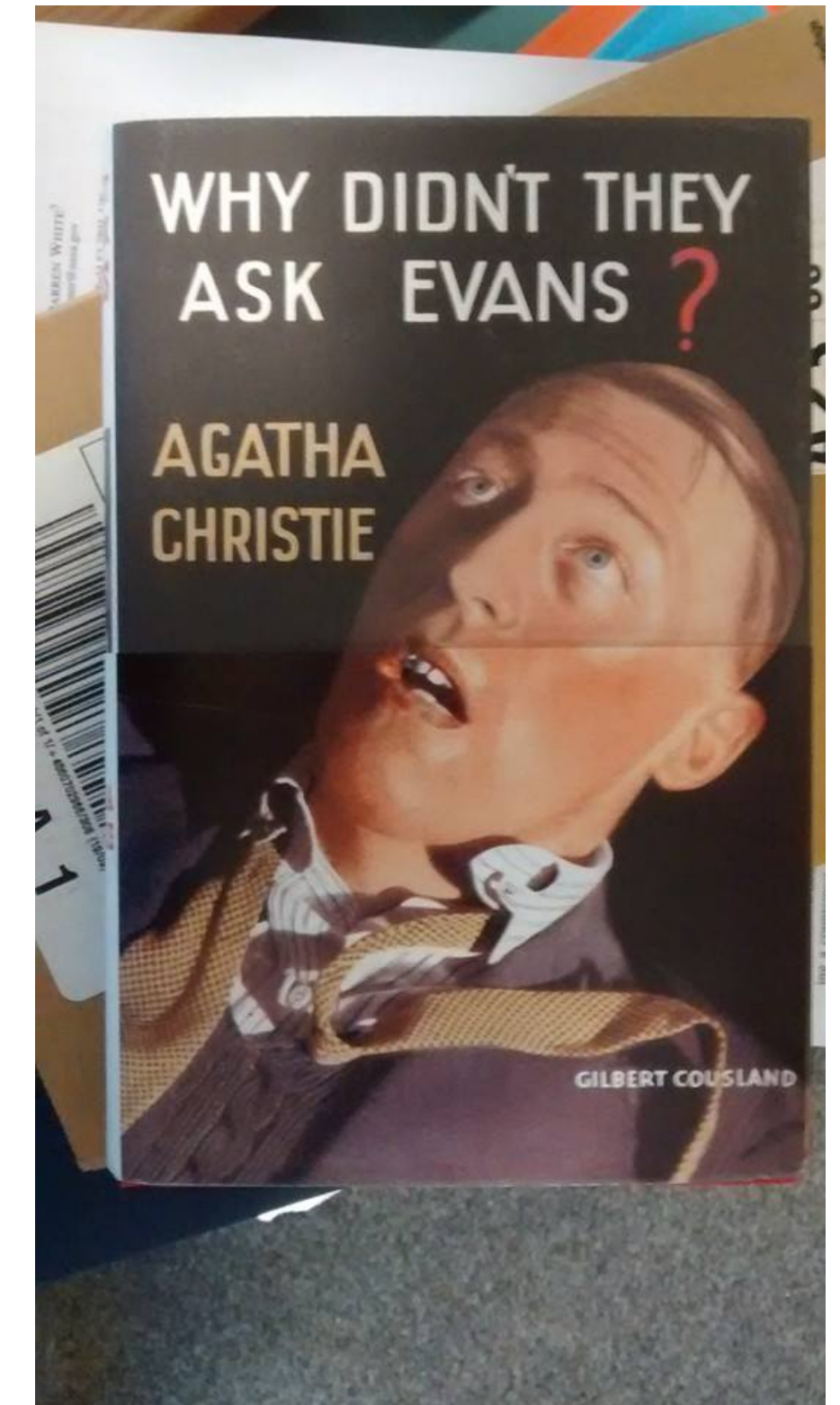
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  - Keep a stressball to hand before opening emails and never reply immediately 😏.
  - You are not a universal PhD supervisor / non-coauthor collaborator.





**Build your garden shed assuming it will one day power the world.**

Assume:

- The scope will widen, more than you expect — and then some.
- The technology it uses will change.
- You will not be the last person in your role.
- Users will do something “stupid”.

## Build your garden shed assuming it will one day power the world.

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  - ✂· “Modular” design.
  - ✂· Developer-facing docs (auto docs?).
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- The technology it uses will change.
  - Infrastructure agnosticism (separate interfaces from implementation where possible).
  - Developer-facing docs (auto docs?).
    - Design map (heterogeneous dependency and call graph).
  - Test harness.
  - Employ a sysadmin.
  - Factor “learning new stuff” in your schedule.
- You will not be the last person in your role.
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## Build your garden shed assuming it will one day power the world.

Assume:

- The scope will widen, more than you expect — and then some.
- The technology it uses will change.
- You will not be the last person in your role.
  - Did I mention developer-facing docs?
  - `git commit` messages matter\*.
  - “No siloed knowledge”.
- Users will do something “stupid”.

\* “*Work, you steaming pile of elephant dung*” is, apparently, “not a helpful record.”

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- The technology it uses will change.
- You will not be the last person in your role.
- **Users will do something “stupid”\***.
  - Build a user community to test, feedback etc. — know what your users need.
  - Lowering the bar to data use comes with drawbacks:
    - *“A common mistake that people make when trying to design something completely foolproof is to underestimate the ingenuity of complete fools.” — Douglas Adams*
  - Warning messages are helpful... and unhelpful.
  - You are responsible for the tools, users are responsible for their use of them.

\* “Stupid”: anything I didn’t expect.



- Coding is fun, science is fun; try to keep it that way.
- Force-feeding rude users their own feet is illegal in most countries\*.
  - “Accidentally” adding their IP to `/etc/httpd/conf.d/blacklist.conf` isn’t\*. 🤪
- Time pressures are real: skimping docs / design etc. costs a LOT more than it saves.
- Astrophysical software is fundamental to allowing any project to fulfil its goals.
- It is not a thankless task, honest!

\* Legal disclaimer: I am not a lawyer, any legal opinions should be treated as “completely made up but sound like they should be true.”