

The DAMIC-M experiment: light dark matter in a kg-scale Skipper CCD detector

Nicolás E. Avalos* on behalf of the DAMIC-M collaboration

*nicolas.avalos@lpnhe.in2p3.fr Laboratoire de Physique Nucléaire et de Hautes Energies (LPNHE - IN2P3 - CNRS)





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DArk Matter In CCDs @ Modane



DArk Matter In CCDs @ Modane



• The goal: direct detection of MeV-scale Dark Matter with a very low background, kg-scale detector



• The detection technology: Skipper CCDs

Talk by J. Tiffenberg

• The location: under the Alps at Laboratoire Souterrain de Modane (LSM), France

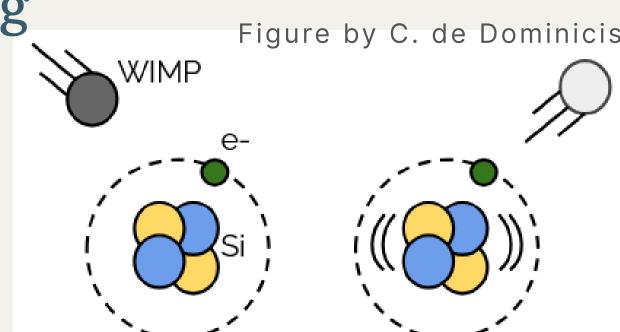
Talk by S. Scorza

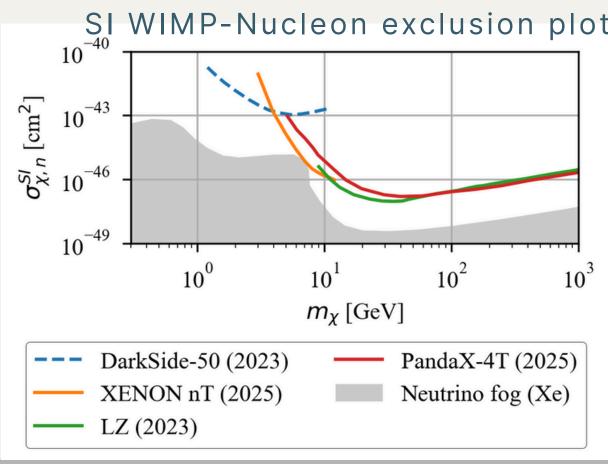
Physics goal

OKWII C.W

• WIMPs are very well-motivated and compelling candidates for explaining Dark Matter

- They may interact with silicon nucleus, thus WIMP-Nucleon interactions may be probed with CCDs
- A "Low Energy Excess" of events (below 200 eVee) prevents the WIMP search to progress
- The Dark Matter interpretation of this excess is ruled out by other experiments

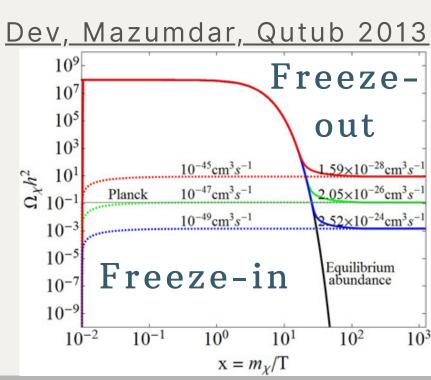




Physics goal

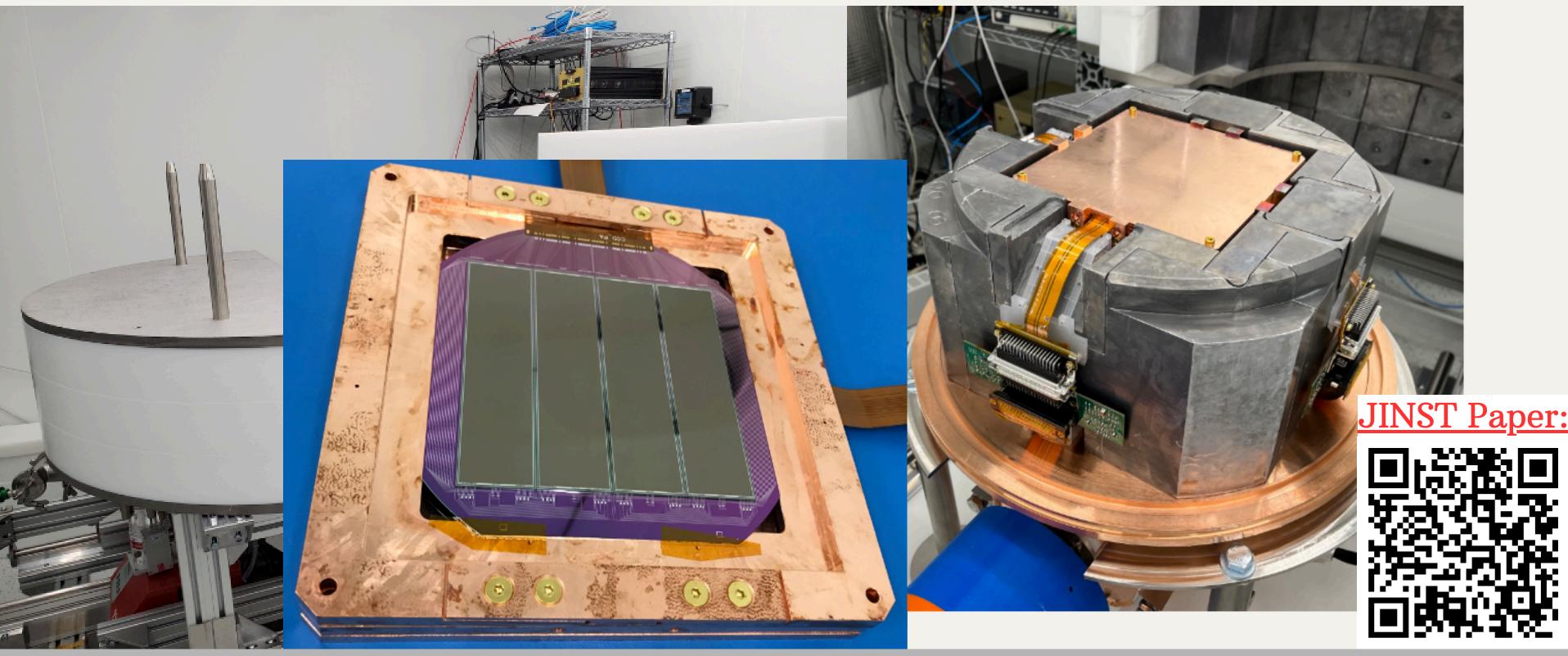
- OKWII.
- Other well-motivated candidates for DM are MeV-scale "hidden sector" particles that interact with electrons via massive "kinematically-mixed vector mediators"
- Φ Φ
 A' (dark photon)
 γ (photon)

- These would produce electron recoils in silicon
- Ultra-low sensitivity of Skipper CCDs mean they are the ideal detectors to probe these candidates
- Depending on the mediator mass, the current DM abundance may be explained via freeze-out or freeze-in



DAMIC-M prototype: the Low Background Chamber (LBC)









PRL Paper:

Event post vertical binning

2 CCD modules

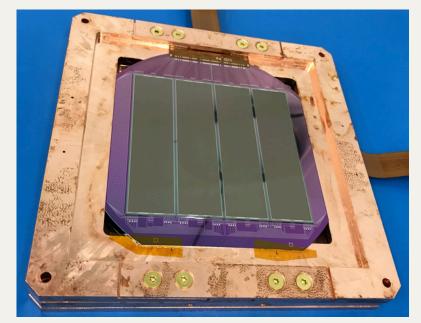
- Each with 4 9-Mpixel Skipper CCDs
- Total sensitive mass: 26 g
- Custom-designed low-noise electronics

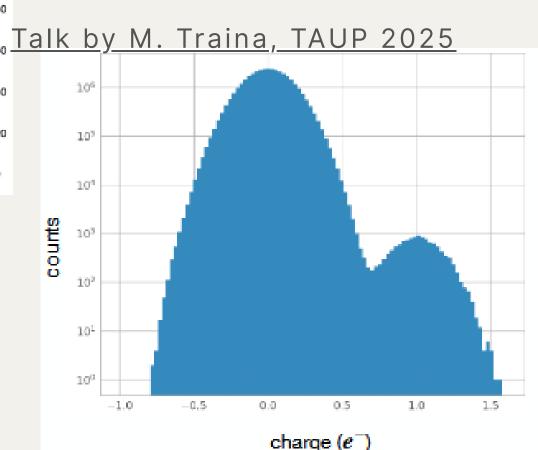
Readout:

- 1 amplifier per CCD
- Continuous readout
- 1 x 100 Hardware Binning

Performance:

- Pixel readout noise: 0.16 e- (500 "skips")
- DC: 10^{-4} e-/pix/day (~400 e-/g/day after cuts)



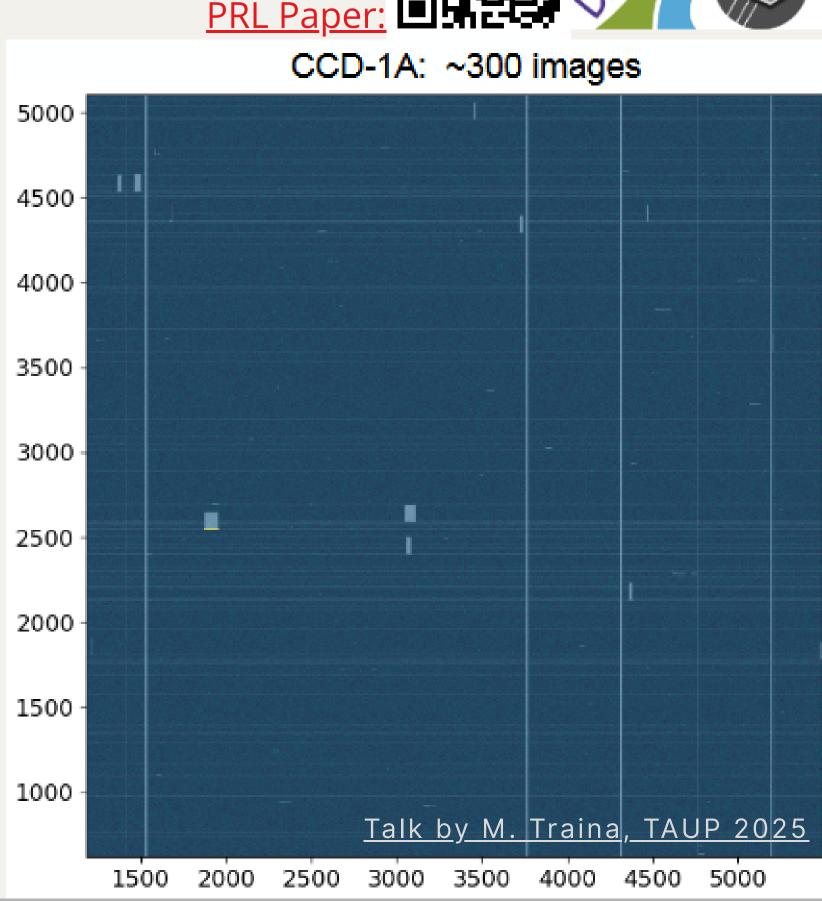




Data selection: masks

- Hot columns
- >5e- clusters
- "bleeding" (CTI):
 - Mask 100 rows above high-energy pixels (q>100e-)
- Cross-talk
 - pixels in same-module CCDs from high-energy cluster in one CCD
- Correlated noise
 - Mask pixels with charge in multiple CCDs of the same module

95% of data survive the selection cuts!



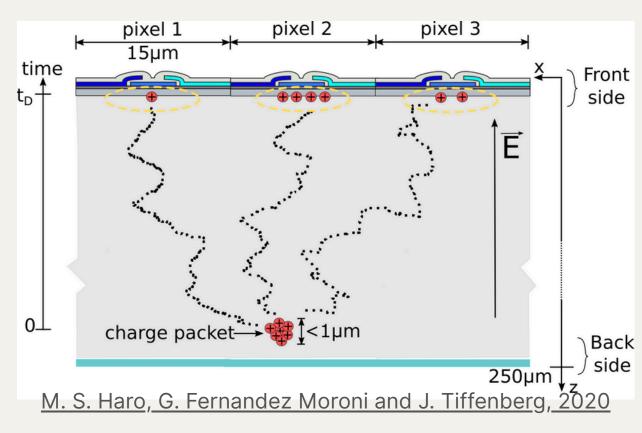
Data analysis

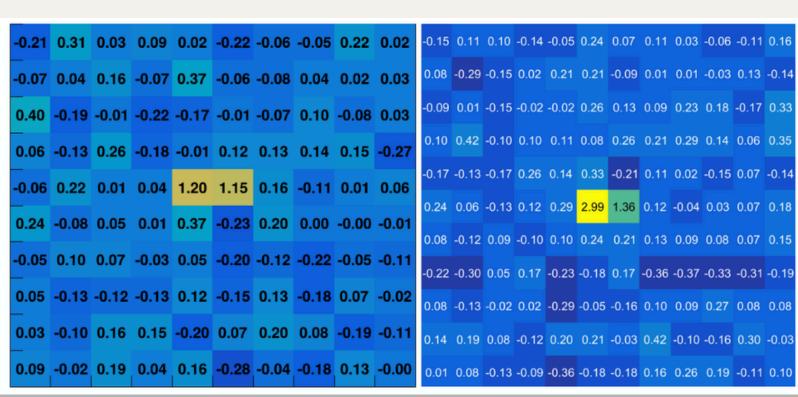
- Blinded analysis
 - o 139 g-day to establish selection cuts; 1.26 kg-day blind
- Pattern analysis
 - Excess in 2, 3 and 4 e- rate with respect to observed 1 e- rate: dark matter?
 - o If so, we would see also consecutive pixels forming patterns, because of difussion in the CCD bulk
- Background modeling
 - o Patterns due to random coincidences
 - o Patterns due to radioactive decays





PRL Paper









Results

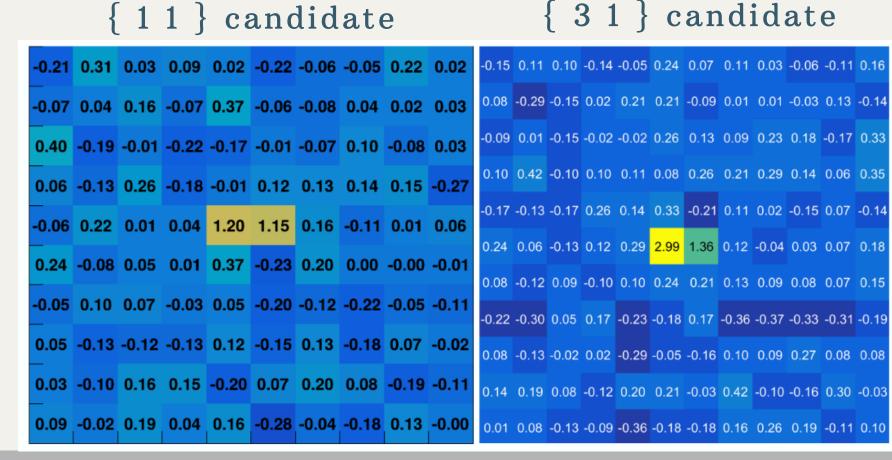
TABLE I. The number of candidates D_p in the D2 dataset, and the number expected from backgrounds due to random coincidences, $B_p^{\rm rc}$, and to radioactive decays, $B_p^{\rm rad}$.

	Pattern p		
D.,	{11} 144	{21} 0	{111} 0
D_p B_p^{rc} B_p^{rad}	141.4 0.039	0.111 0.039	0.042 0.016
	{31}	{22}	{211}
D_p B_p^{rc} B_p^{rad}	0.019 0.052	0.5×10^{-5} 0.011	0 5.8×10^{-5} 0.035

Pattern events are compatible with background-only hypothesis (even the {3 1} event)

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 \rightarrow No Dark Matter in the 2-5 errange



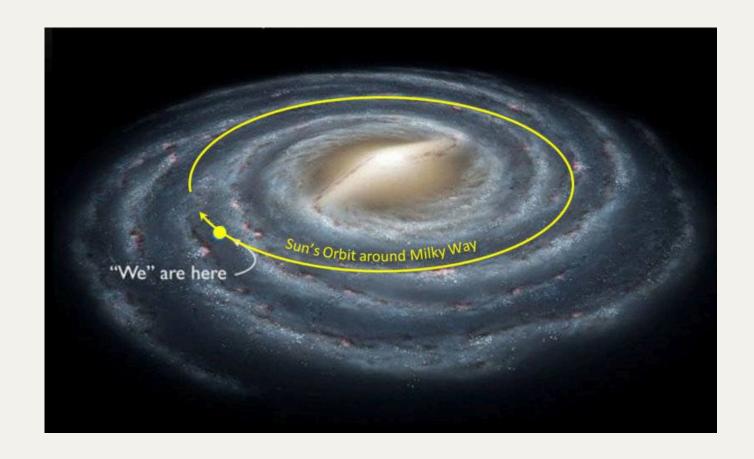
Diurnal modulation analysis

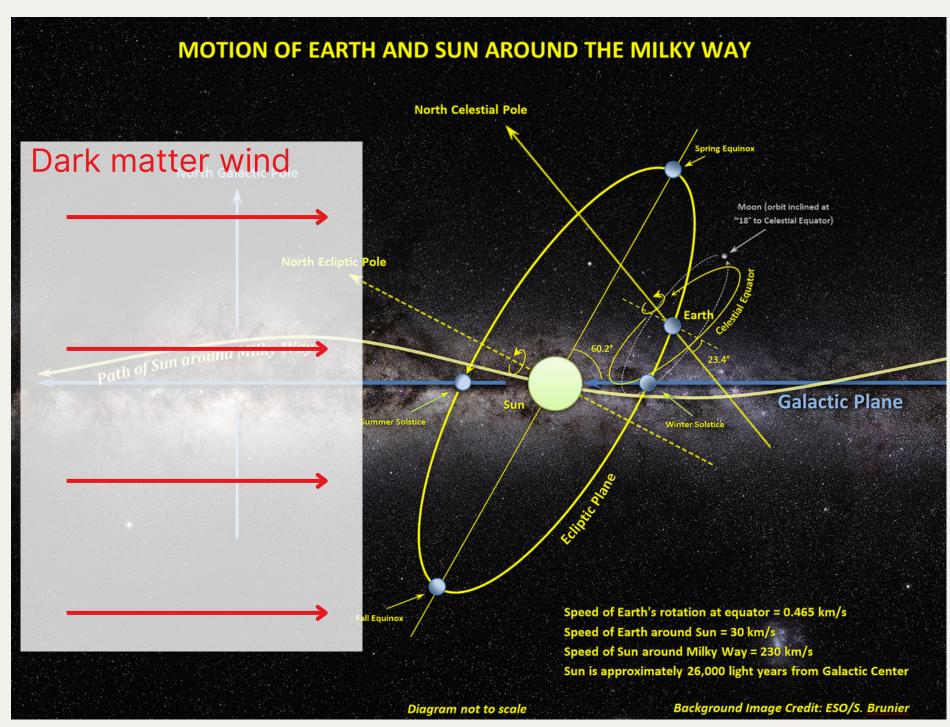
OKWII CH

What about the single-electron events? Could they be Dark Matter?

If so, they would appear in a time-dependent fashion, due to a

"diurnal modulation of events"



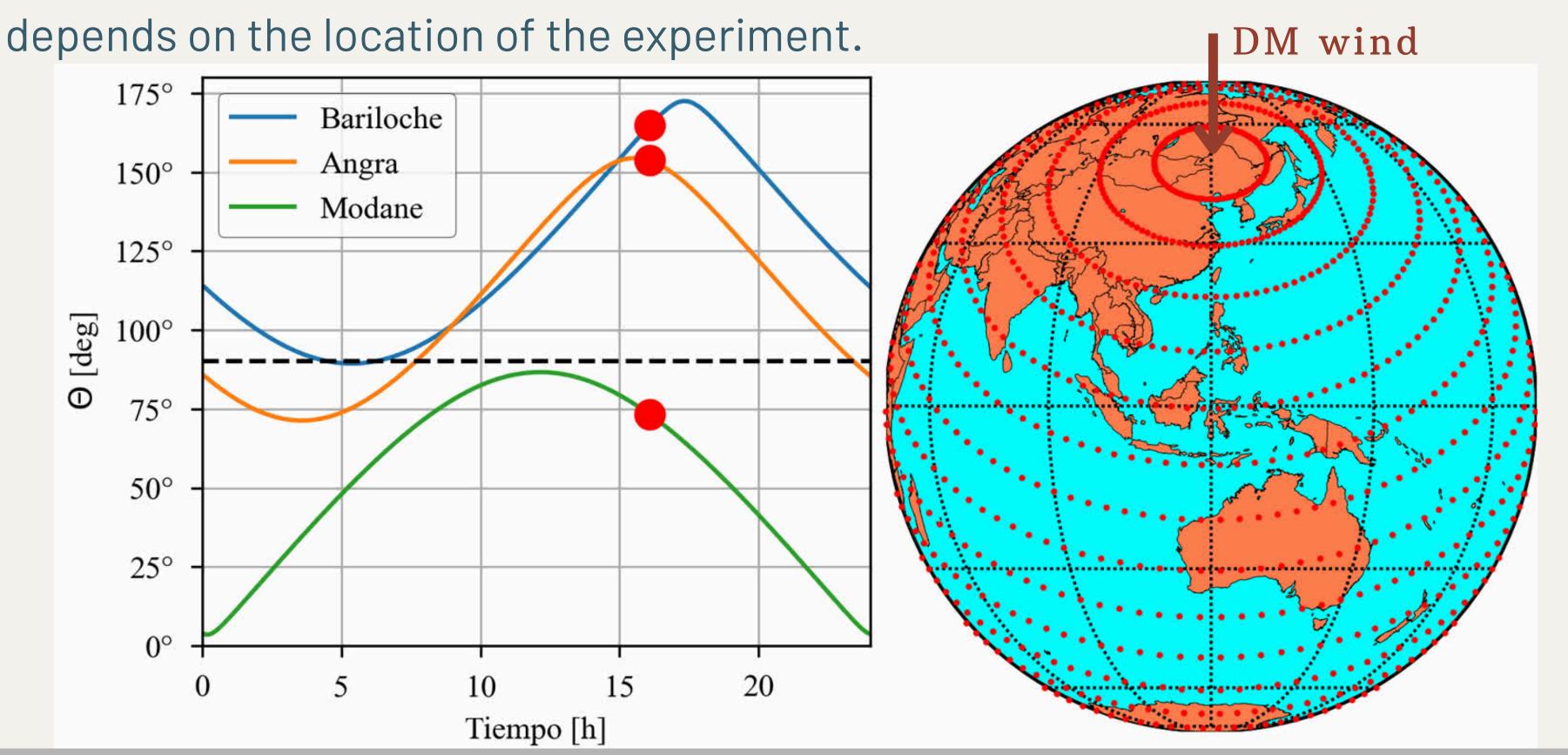


Jim slater 307, CC BY-SA 4.0, via Wikimedia Commons

Diurnal modulation analysis



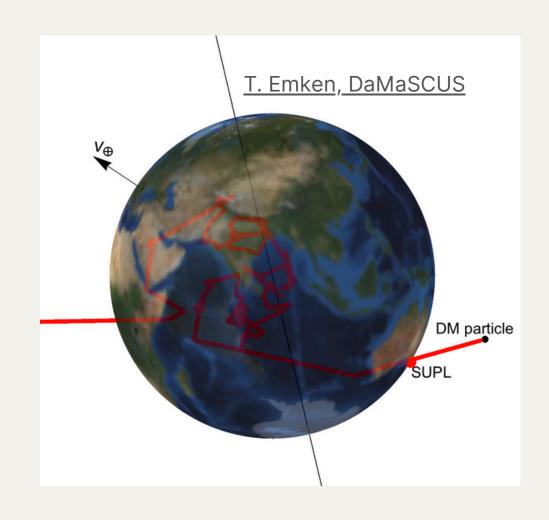
"Isodetection angle" varies with time during a sidereal day, and it also



Diurnal modulation analysis

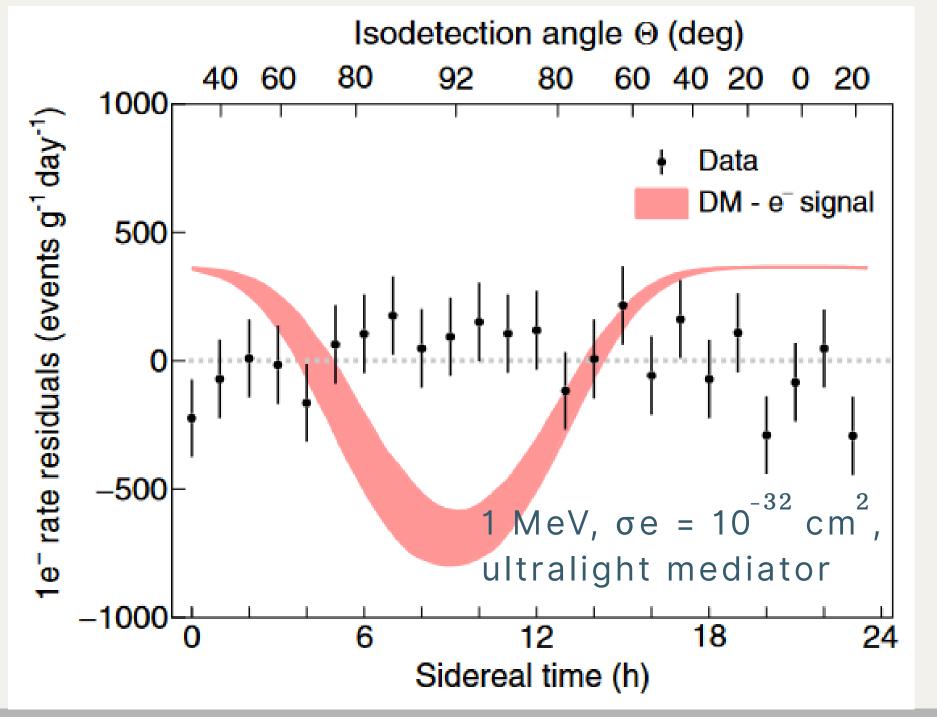
DM particles would interact with protons inside the Earth's crust, deflecting their path along their way and creating different velocity and flux distributions

at the experiment location





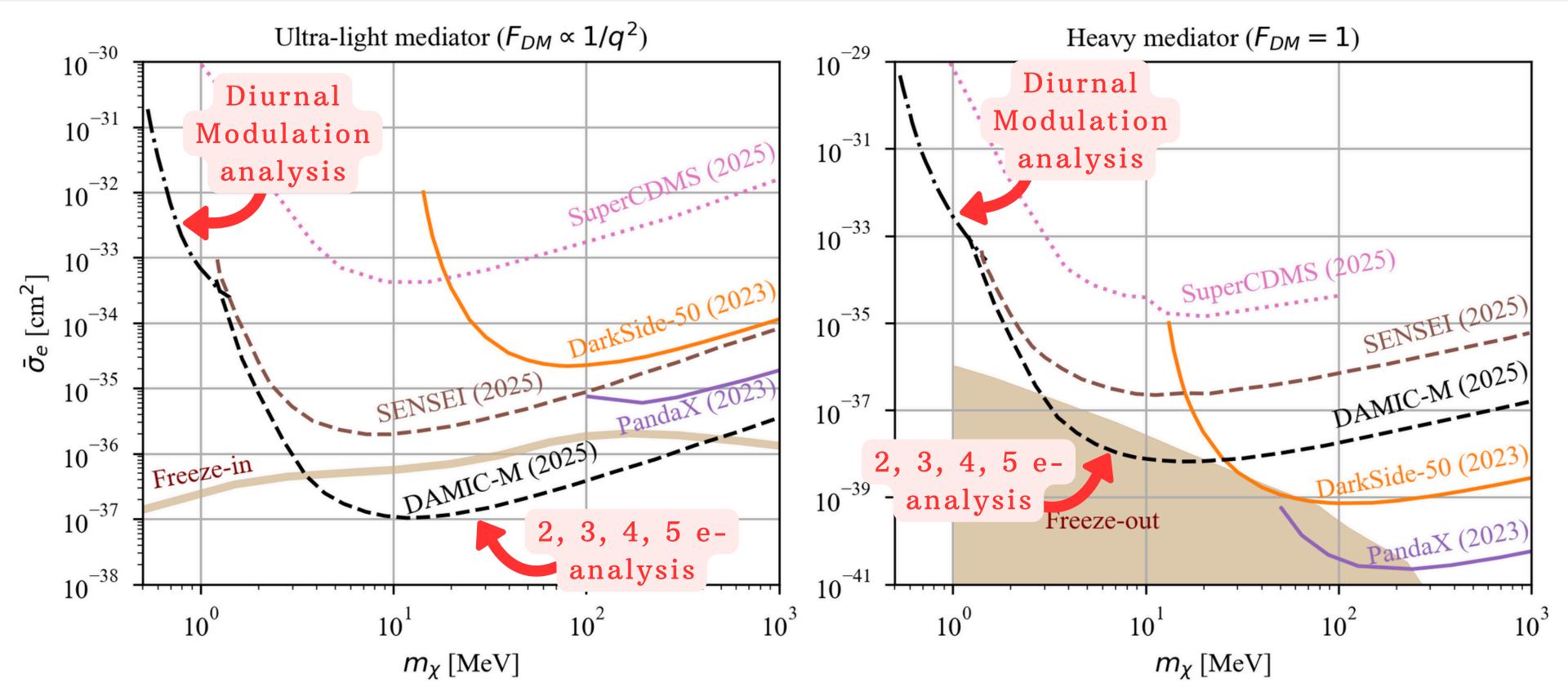
PRL Paper:



DAMIC-M LBC results

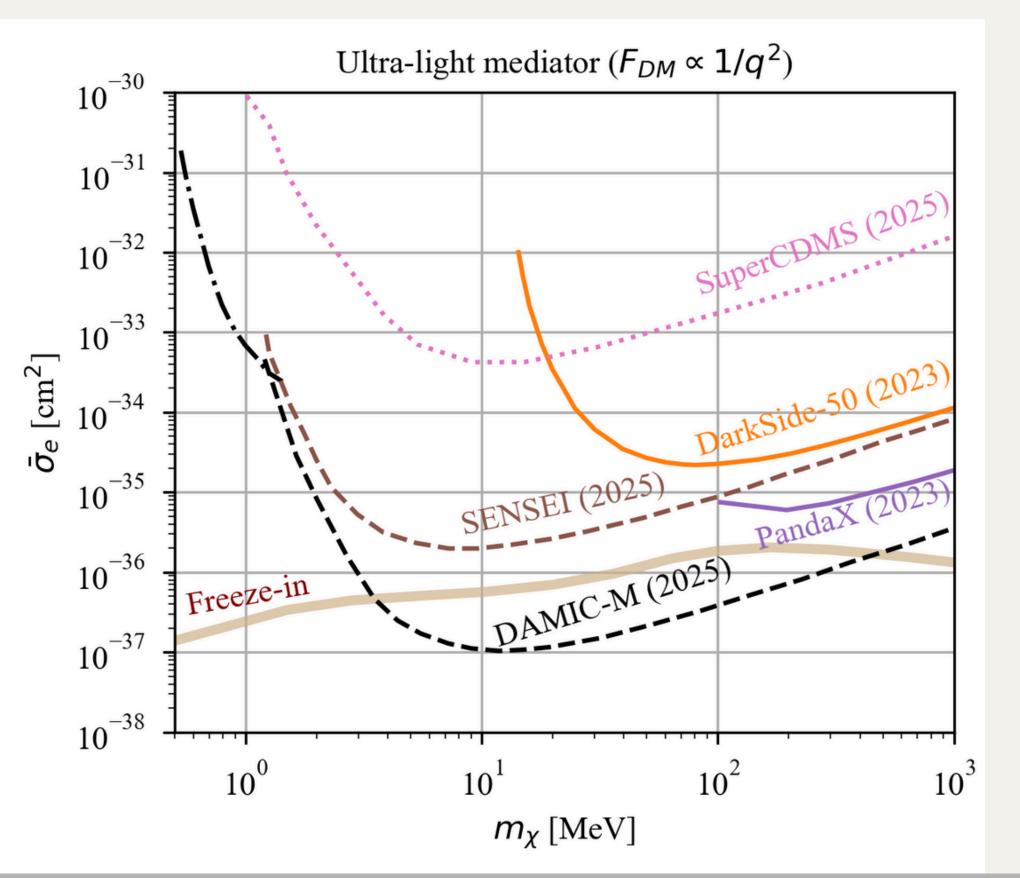


No positive signature for DM yet. We can calculate exclusion limits:



DAMIC-M LBC results





Ultra-light mediator (mA' << mX)

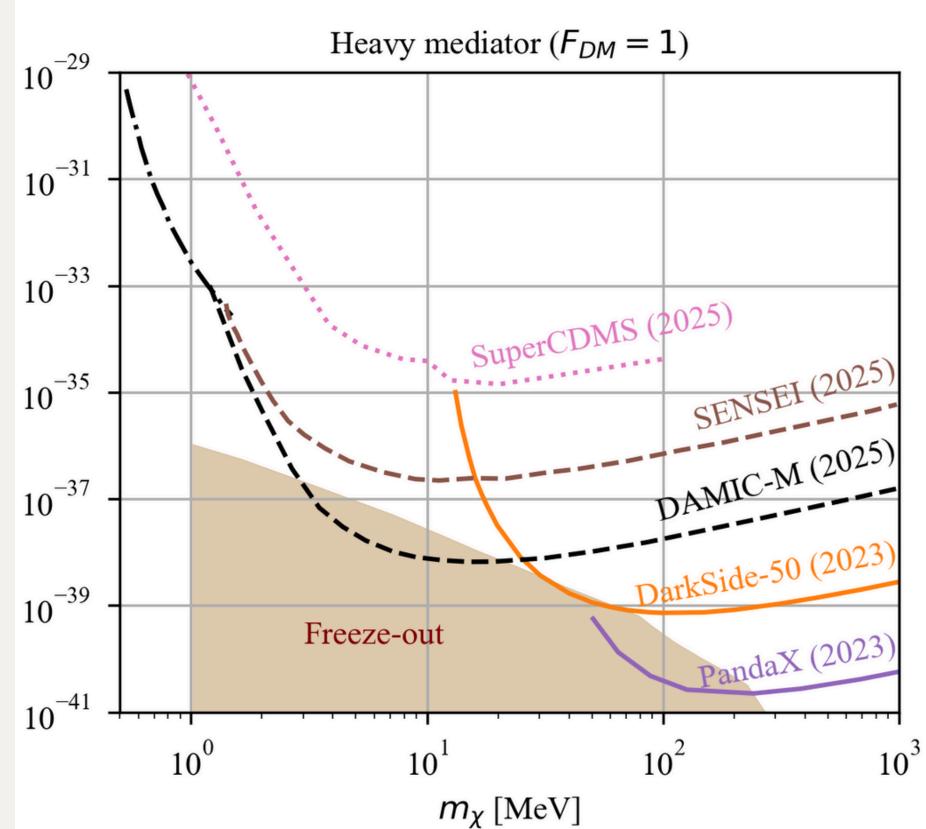
- Skipper-CCDs are the most sensitive detectors in the whole MeV mass range
- Particles with masses between 3.5-490 MeV are excluded from making up the <u>dominant</u> component of Dark Matter: cross-section would be too small, leading to an under-abundance.

DAMIC-M LBC results



Heavy mediator (mA' >> mX)

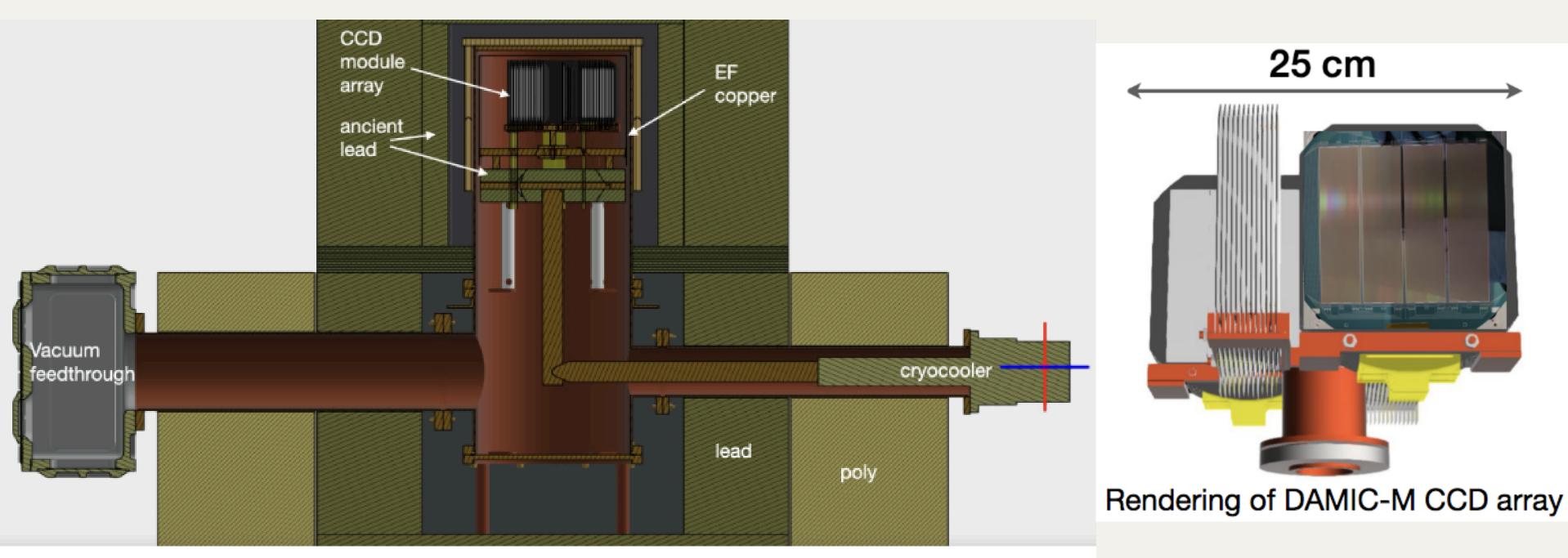
- Skipper-CCDs are the most sensitive detectors in the 1-20 MeV mass range
- Complex scalar particles with masses between 3-20 MeV are totally excluded from making up any fraction of Dark Matter: cross-section would be too small, leading to an over-abundance.
- Other candidates (e.g. fermion particles) still viable

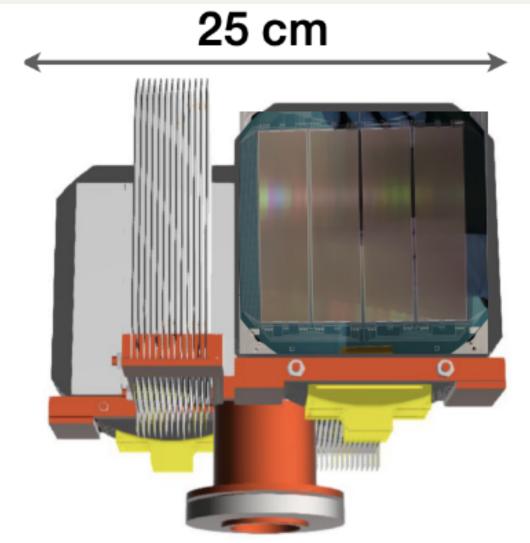


Towards DAMIC-M full installation



Next step: Installing ~350 g of Skipper CCDs at LSM





Rendering of DAMIC-M final design

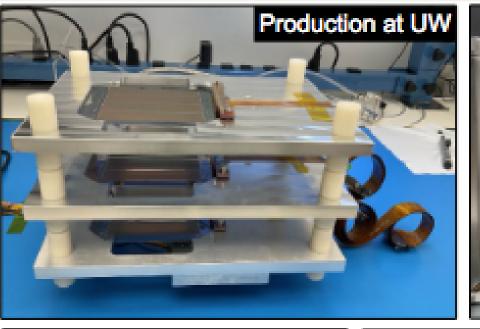
Module production preprint:



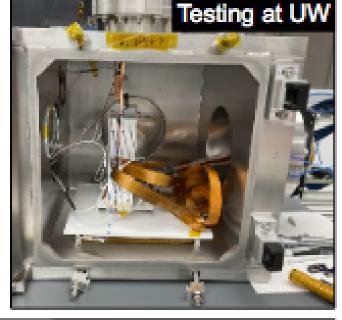


scaling the detector mass

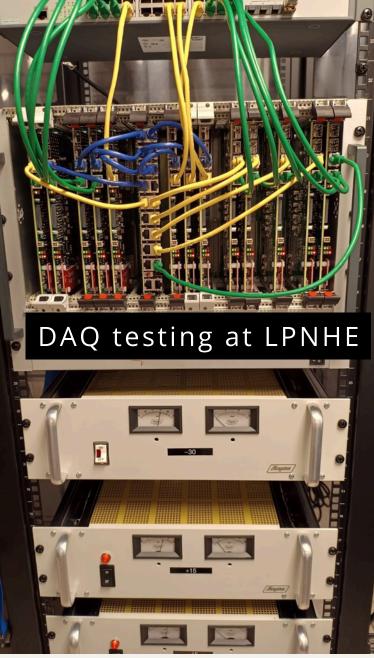
- 28 Skipper CCD modules, with 4 CCDs each, have been manufactured
- Underground testing is almost done at a specifically-designed test chamber in LSM
- All readout boards have been manufactured, DAQ stress testing ongoing at LPNHE











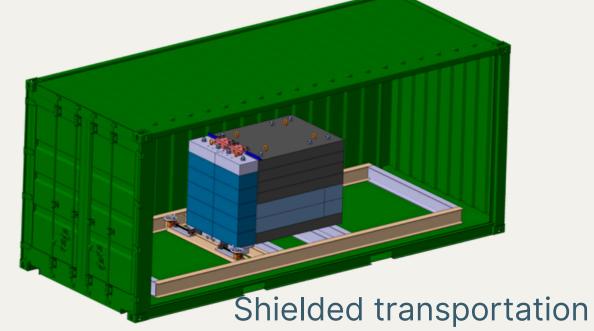
Module production preprint:





controlling our backgrounds

- Limit exposure time to cosmic rays
- Limit the detector surfaces' exposure to radon
- New materials: Electro-Formed copper, low-background cables
- Remove Si wafer surface (to reduce surface Pb210)
- Chemical treatments of Cu, Pb components to remove surface Pb210









Shield during CCD production

Summary



- DAMIC-M prototype sensor has set stringent limits to MeV-scale particles interacting with electrons
- It has ruled out that particles with masses between 3.5 and 490 MeV and ultra-light mediator are a dominant component of DM
- It has ruled out that particles with masses between 2.9 and 21.5 MeV and heavy mediator can explain any fraction of DM
- 28 Skipper CCDs will be installed at LSM during 2026 and their operation will explore deeper into the DM phase space

Stay tuned, and do not hesitate in asking further questions!

Thank you.