

# Lessons from XRISM analysis: spectral models and diagnostics

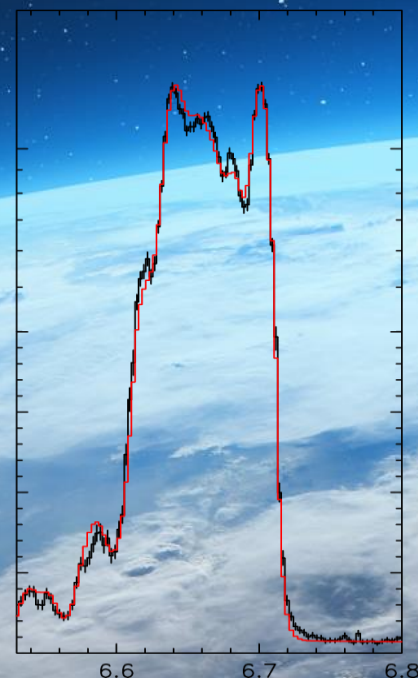
Liyi Gu

**SRON**

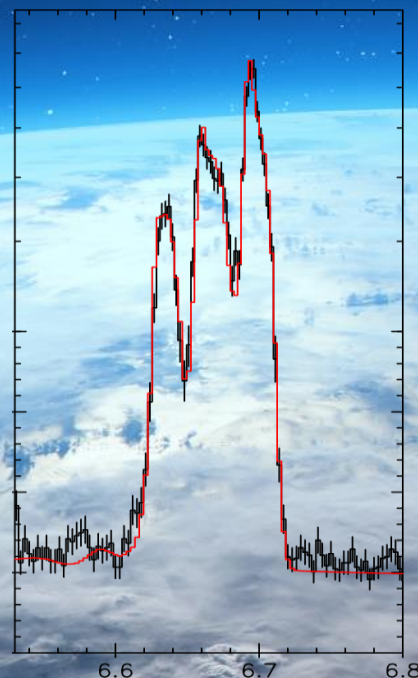
SPACE  
RESEARCH  
ORGANISATION  
NETHERLANDS

Jan 2026 Lumière Paris-Saclay

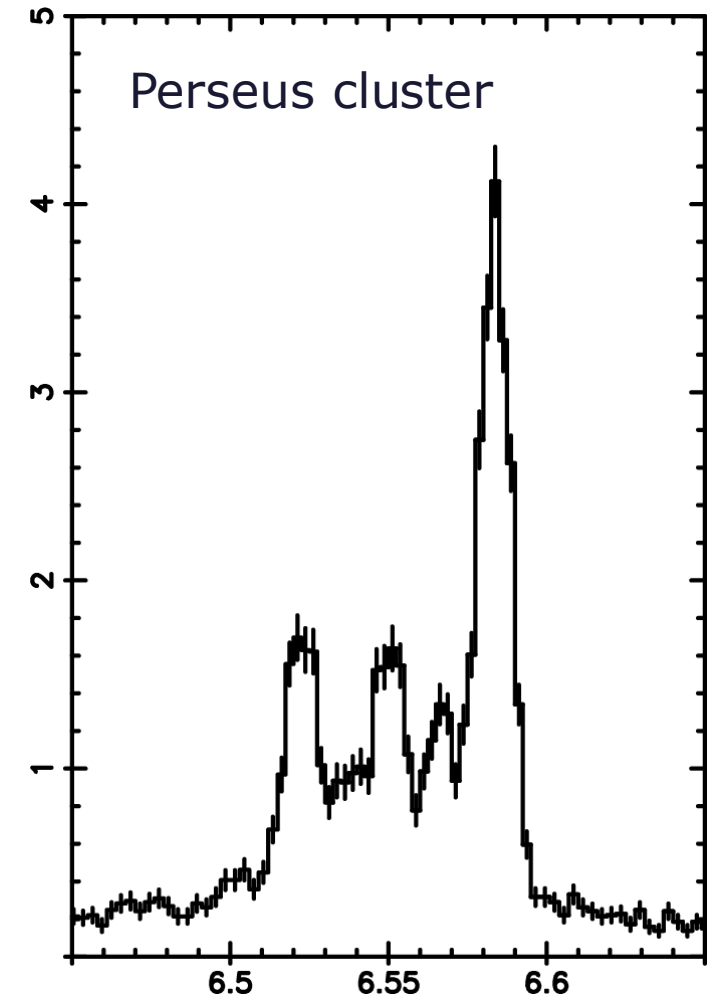
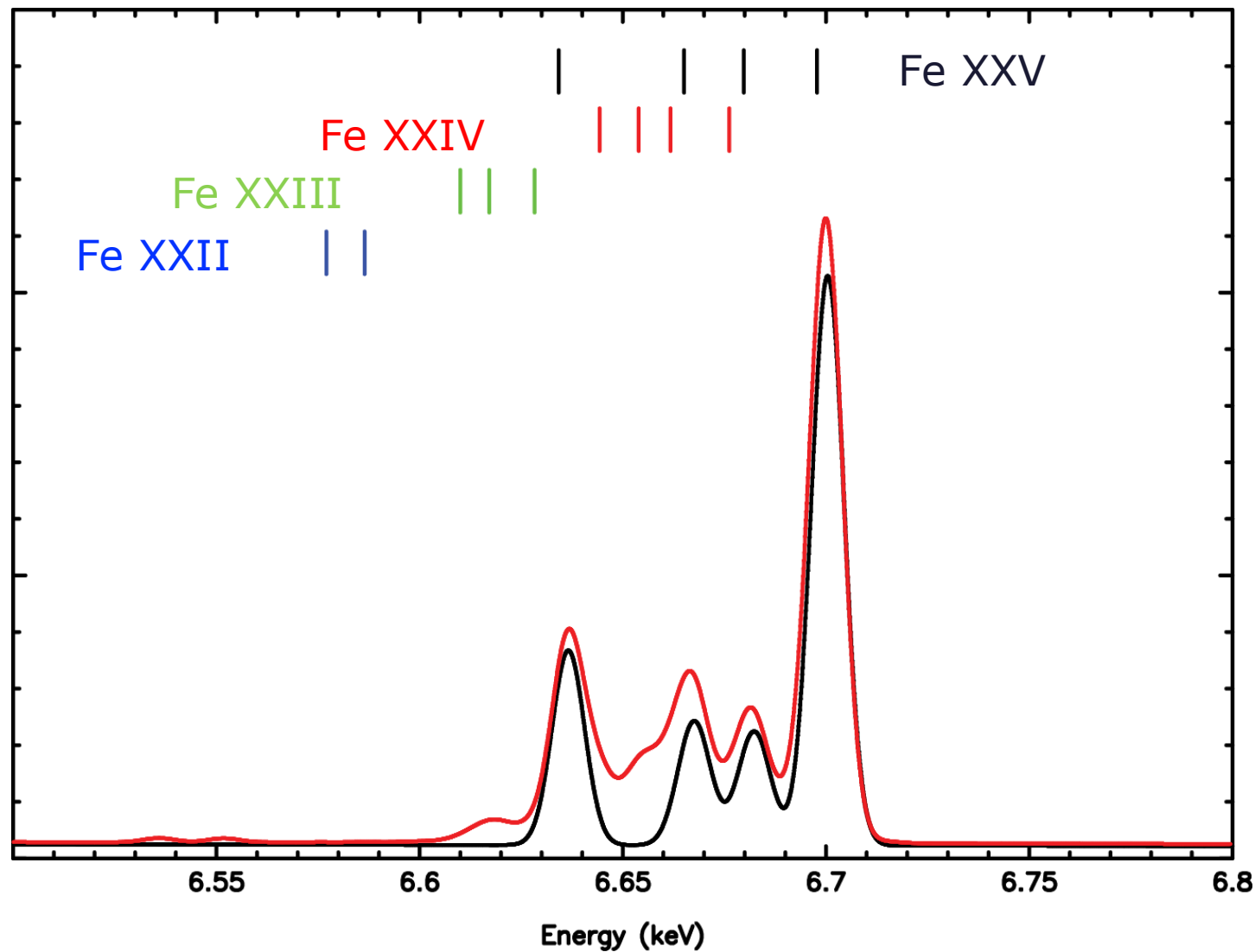
Collisional



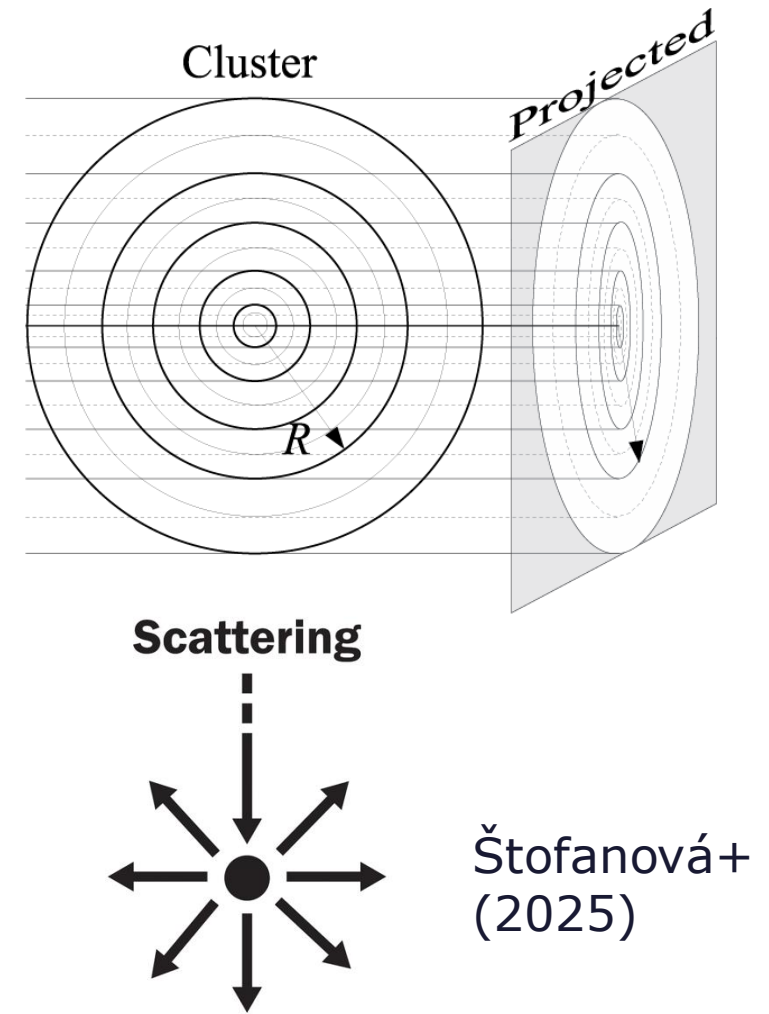
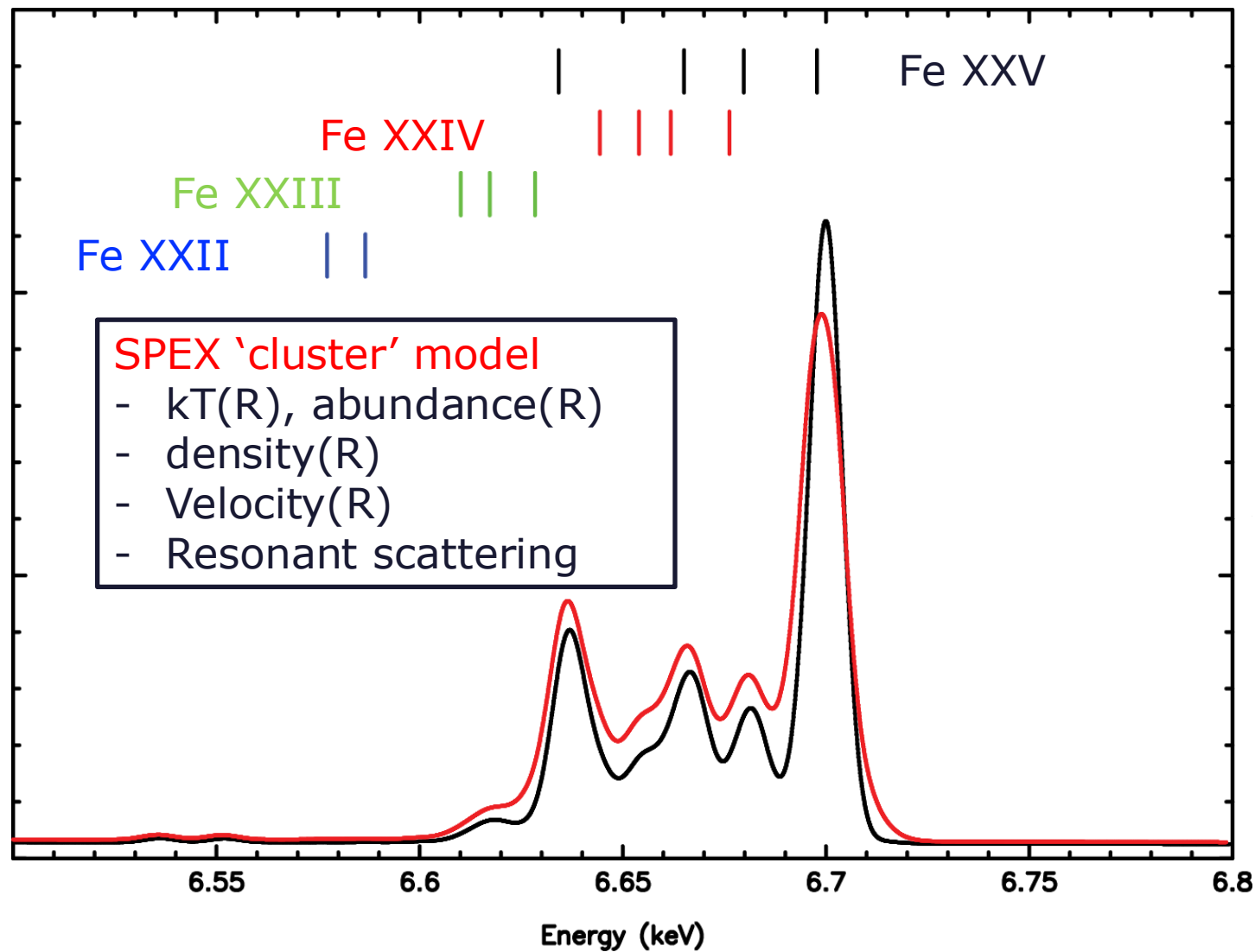
Photoionized



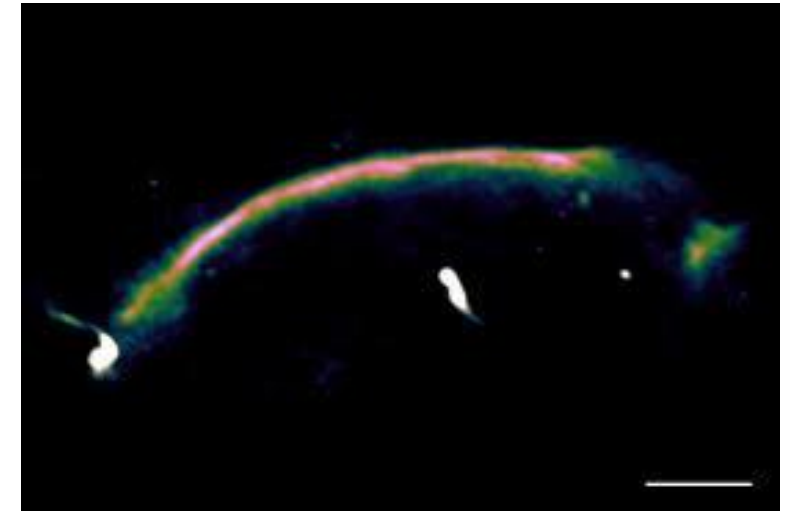
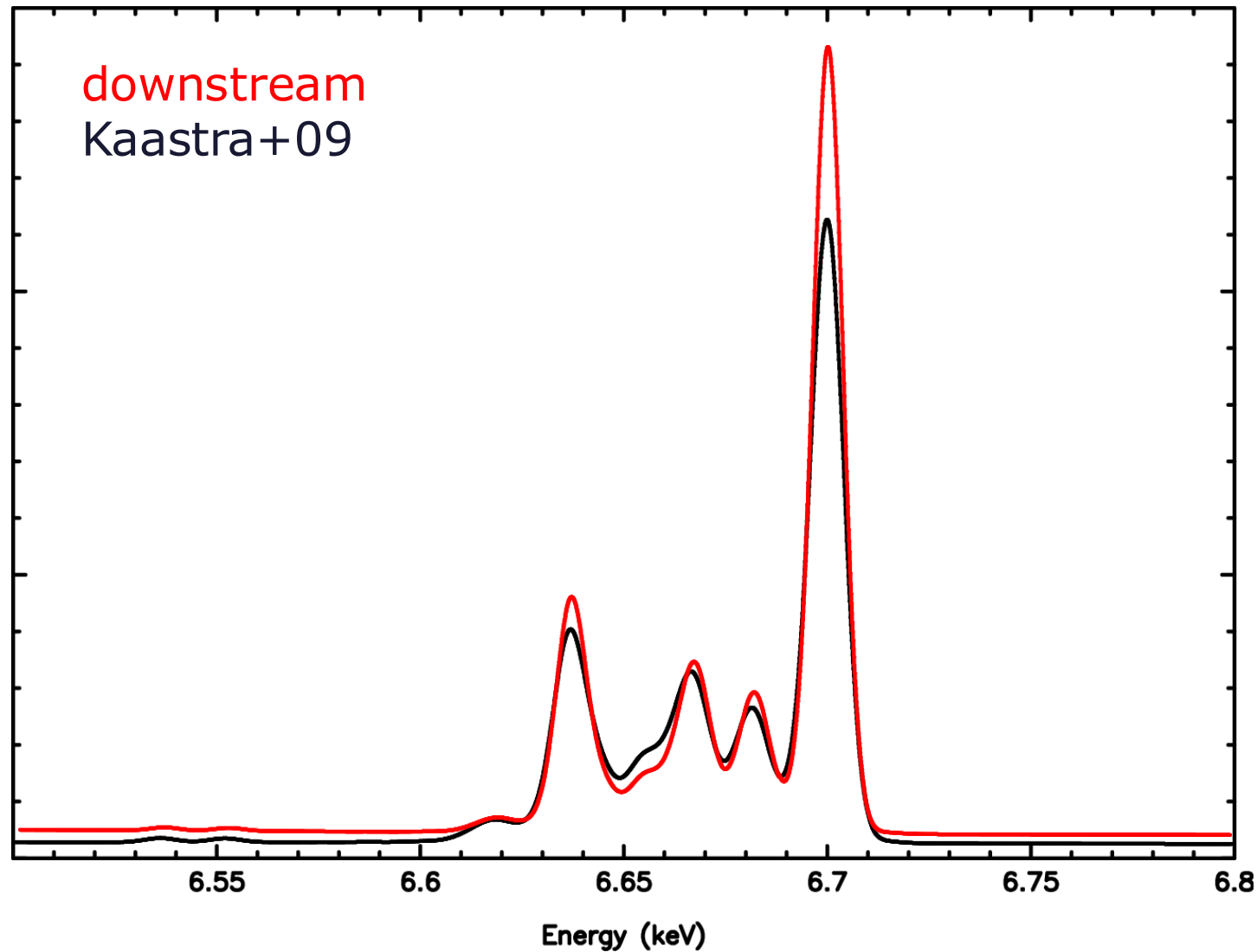
## Start with the basic: a CIE model with $kT = 4$ keV



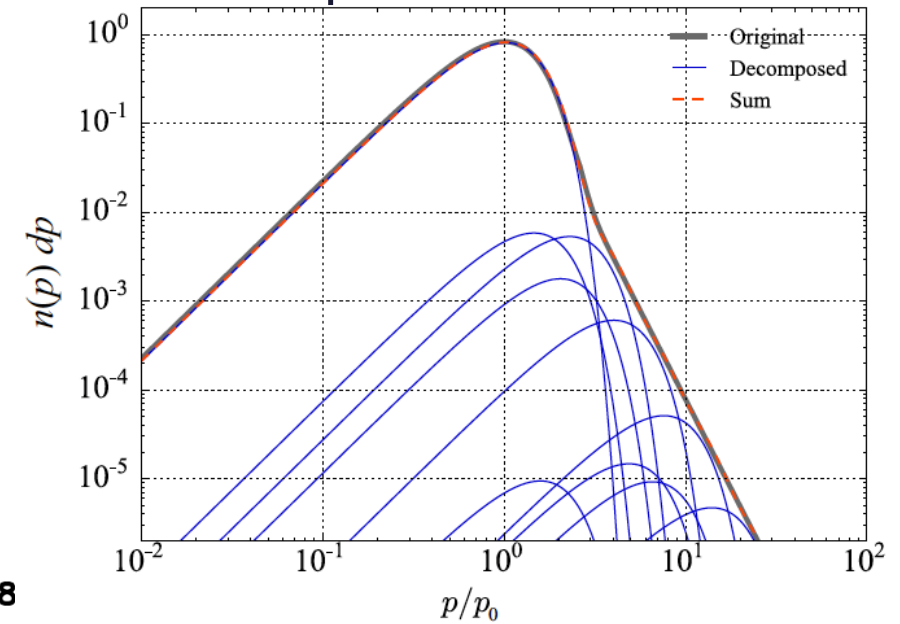
# Adding gradients and photon scattering



# Adding non-thermal component



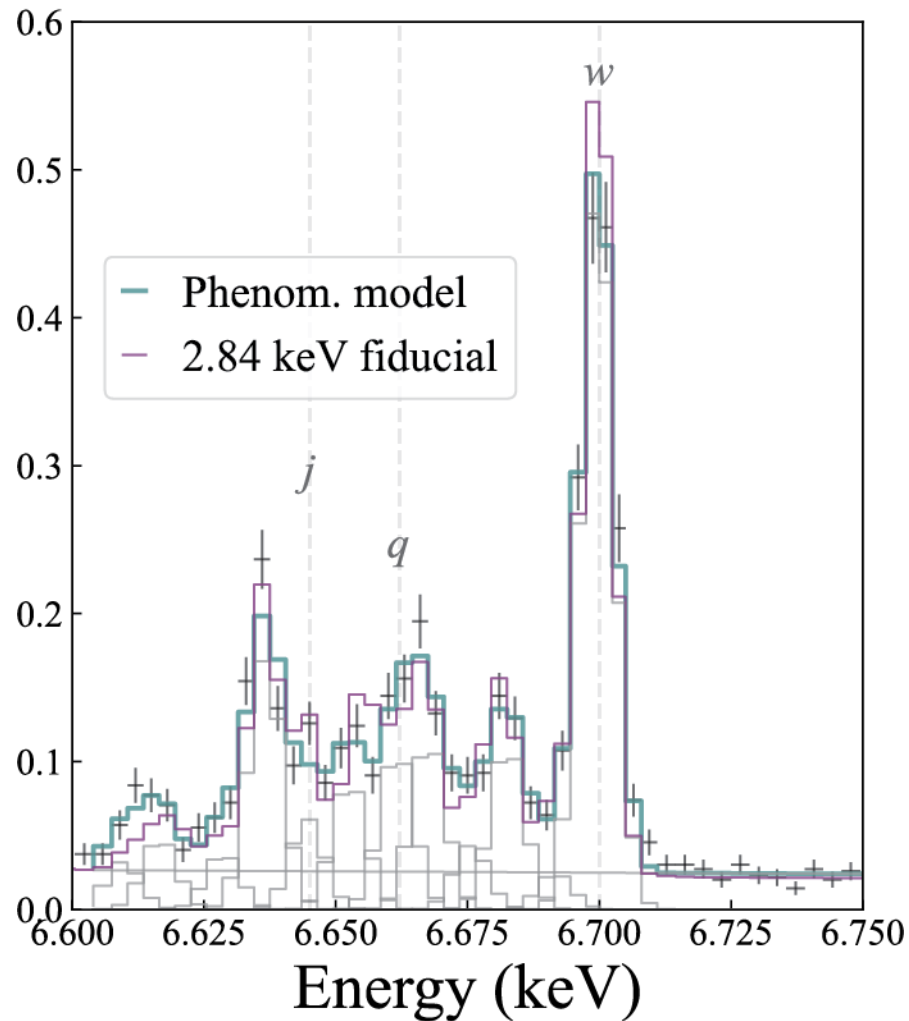
## Decomposition



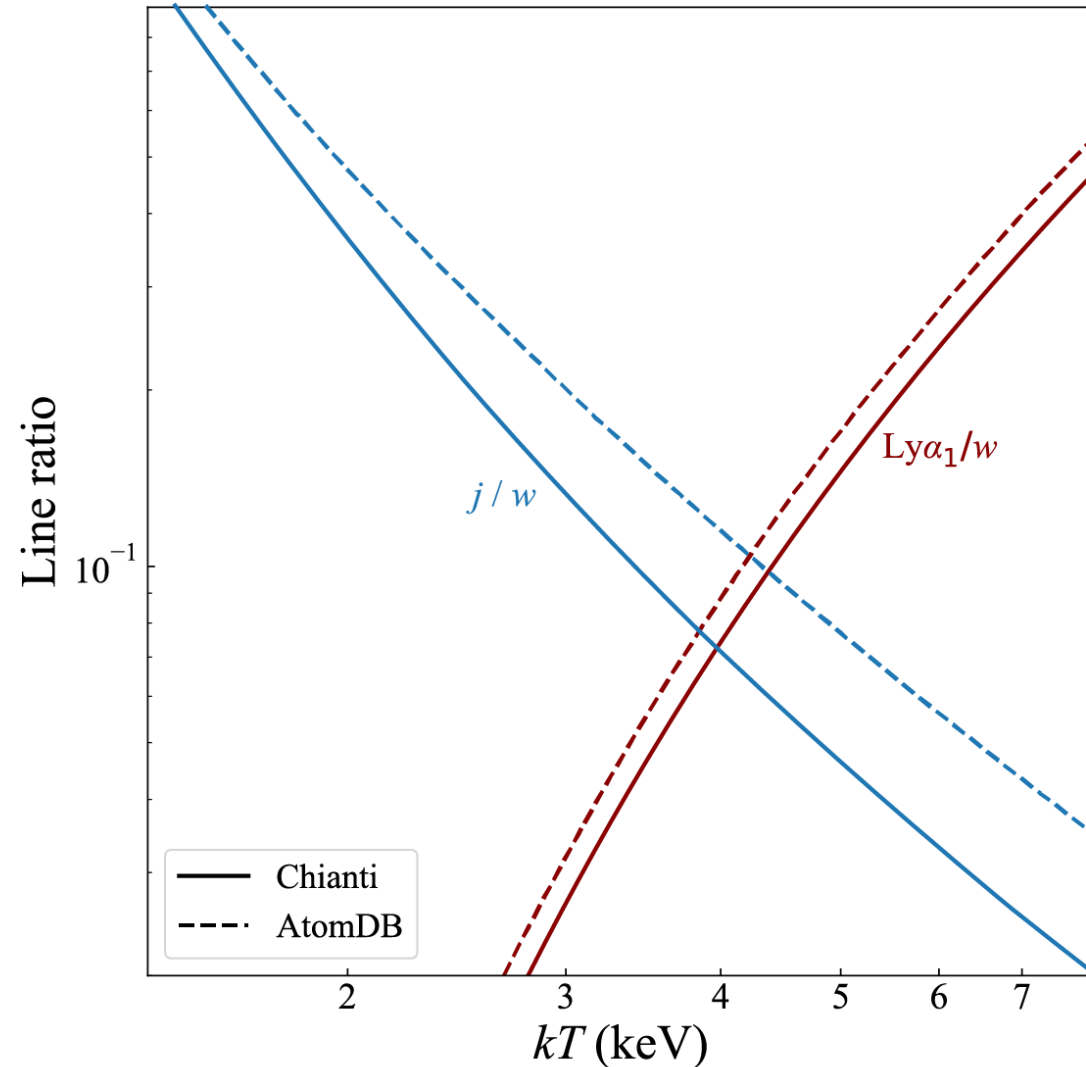
# Line diagnostics

GT Mus (Resolve)

Kurihara+25



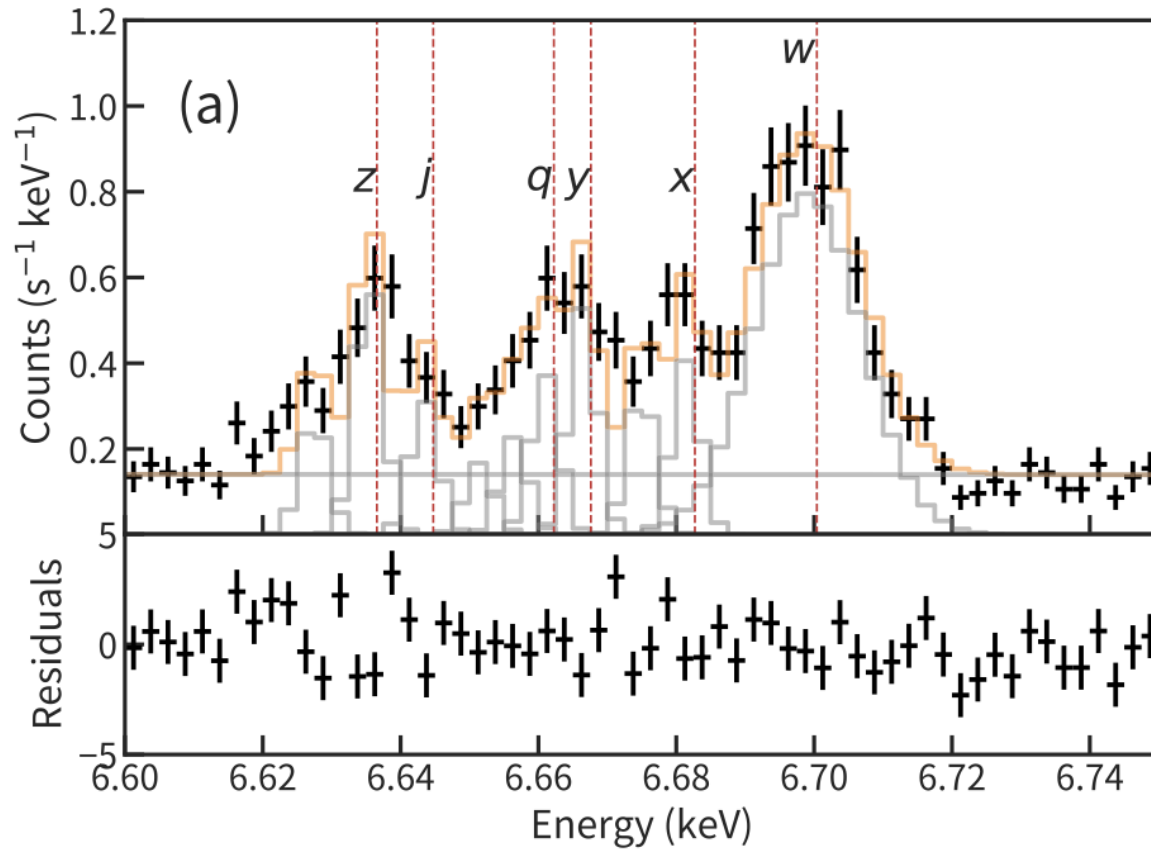
DR lines (such as  $j$ ) will be suppressed at high  $kT$ .



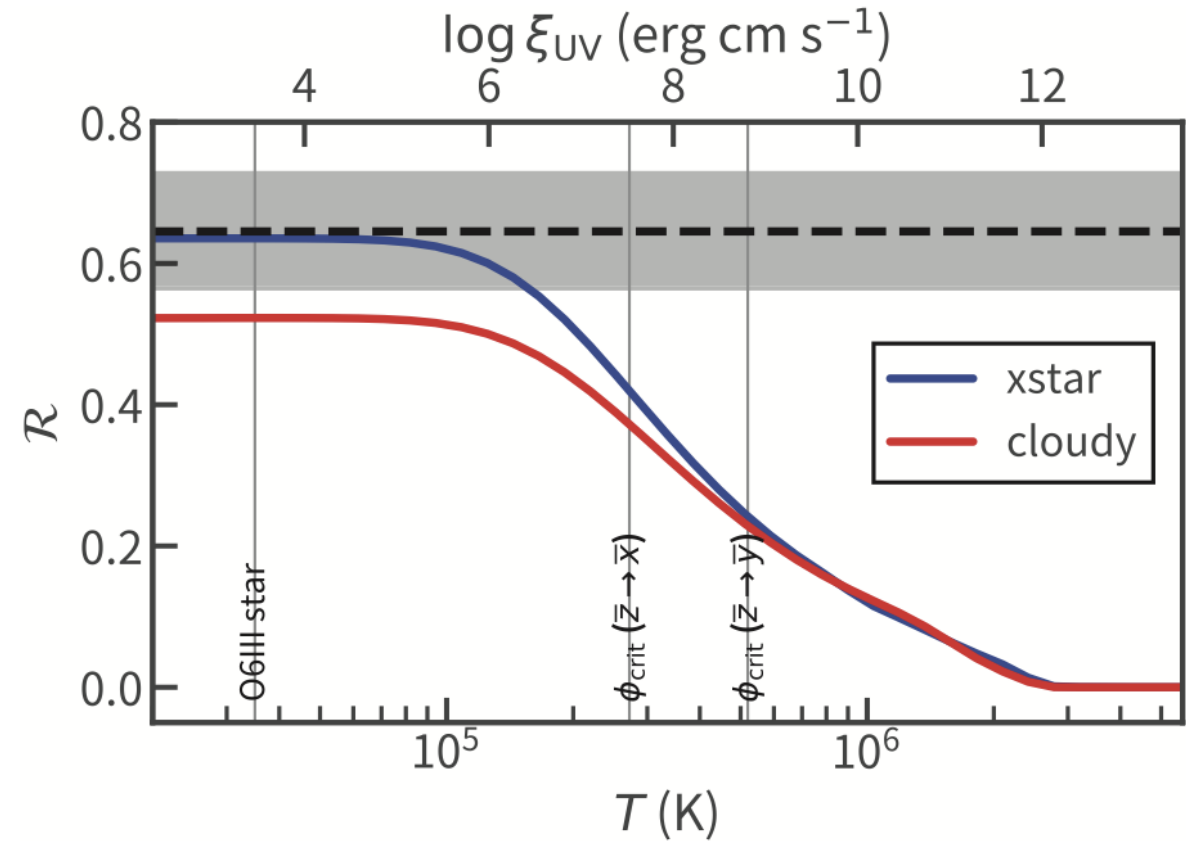
# Line diagnostics

Cen X-3

Mochizuki+25



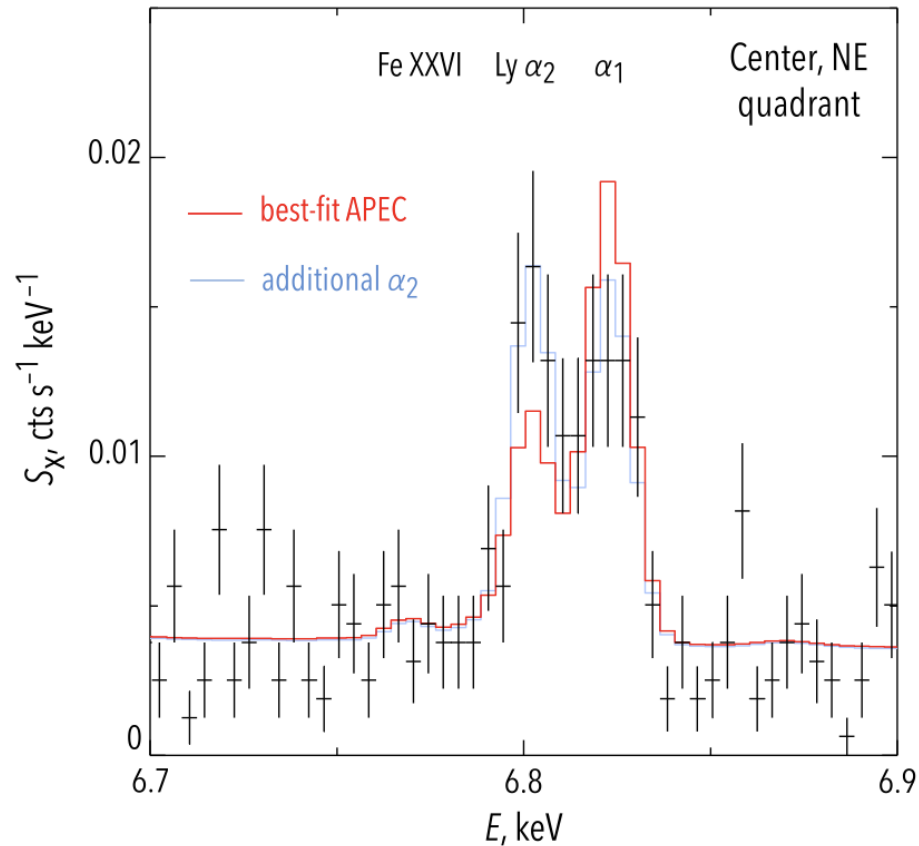
z line is suppressed at high density or high UV flux



Forbidden level can be mixed with upper levels via photon or collision excitation

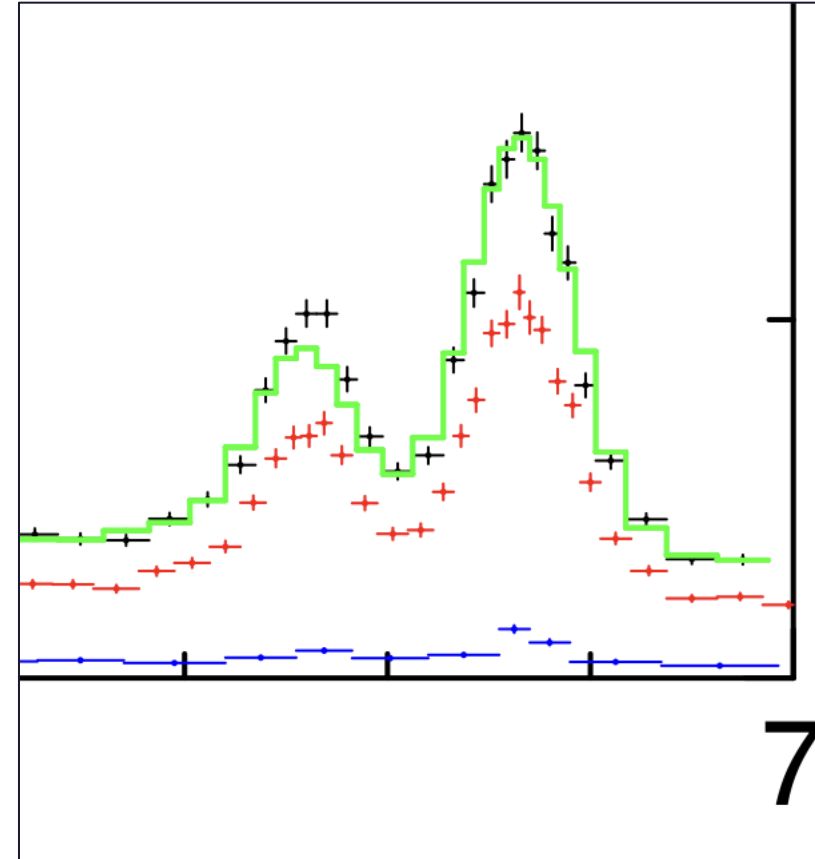
# Fe XXVI Lyman $\alpha$ lines

Coma cluster  
XRISM+25



By theory,  $L_{\alpha 2}/L_{\alpha 1} = 0.5$   
Several clusters indicate  $> 0.5$

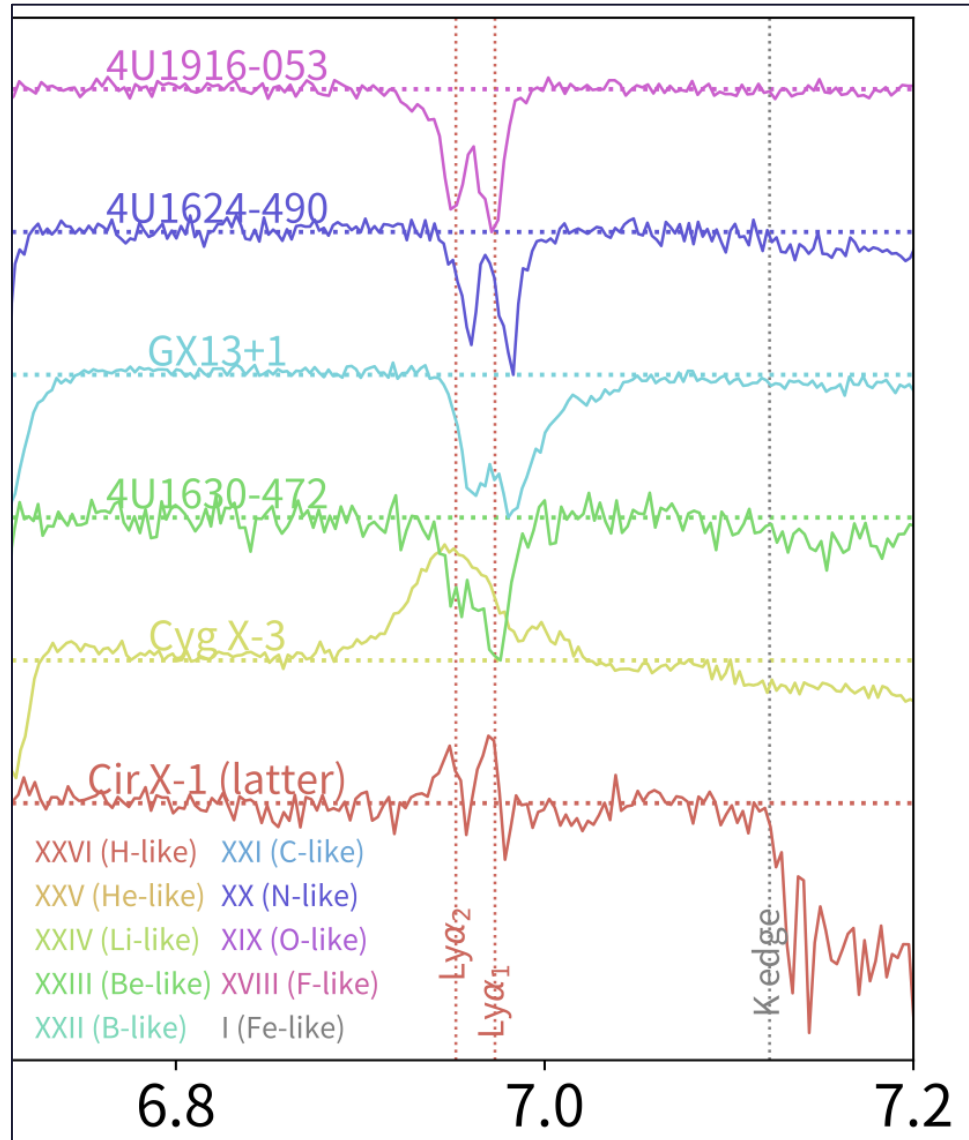
Stacked 10 clusters  
XRISM+25



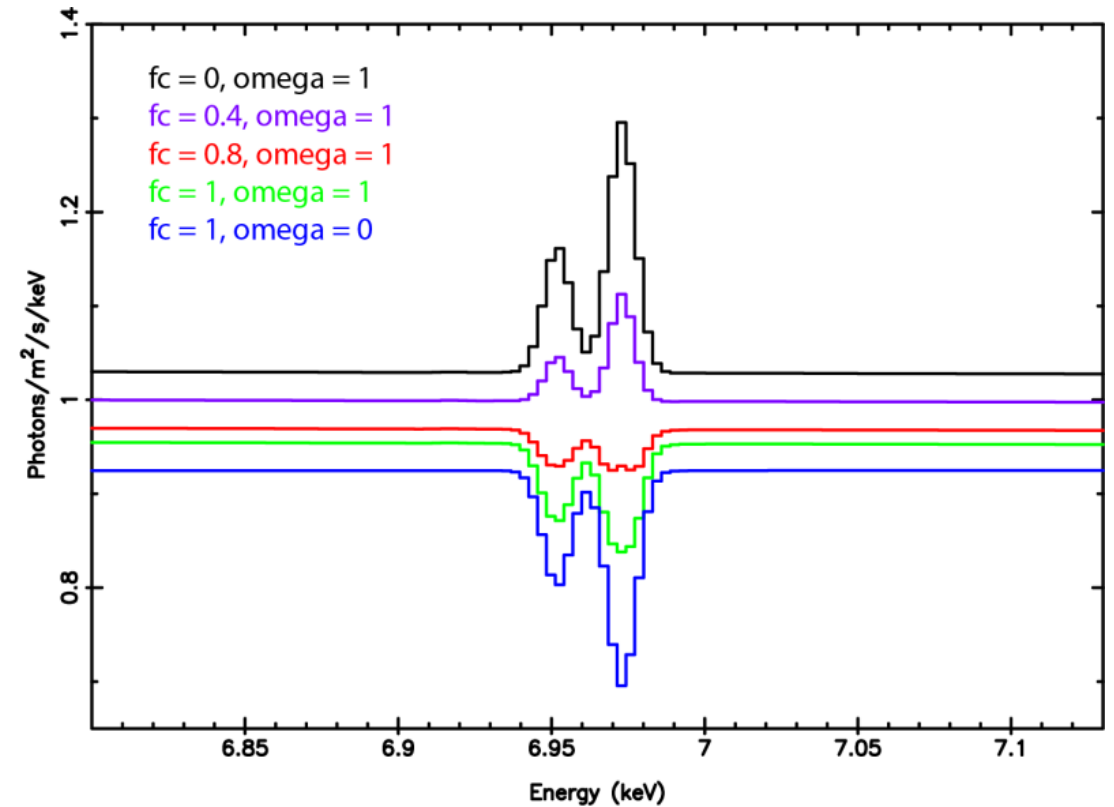
Potential diagnostics: resonant scattering,  
blending M1 transition, charge exchange...

# Fe XXVI Lyman $\alpha$ lines

X-ray binaries Tsujimoto+25



## Pion calculation

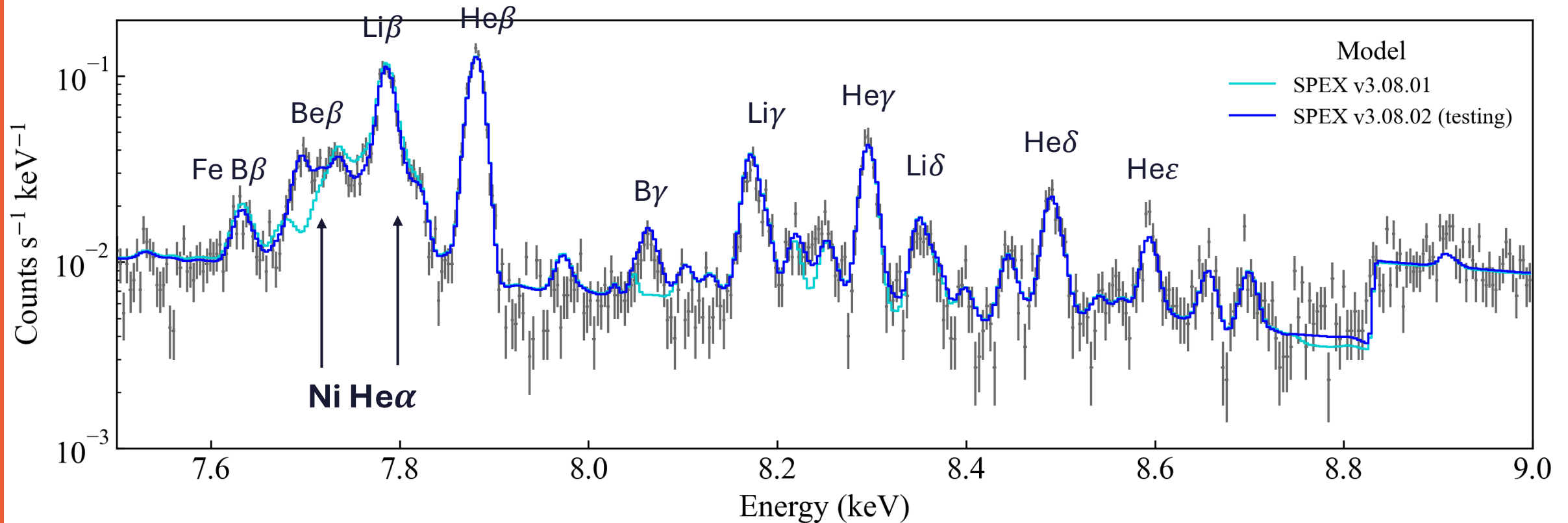


Potential diagnostics: saturation/partial covering, scattered emission, high density, high UV flux, polarization...

# Fe $K\beta$ lines

W49B

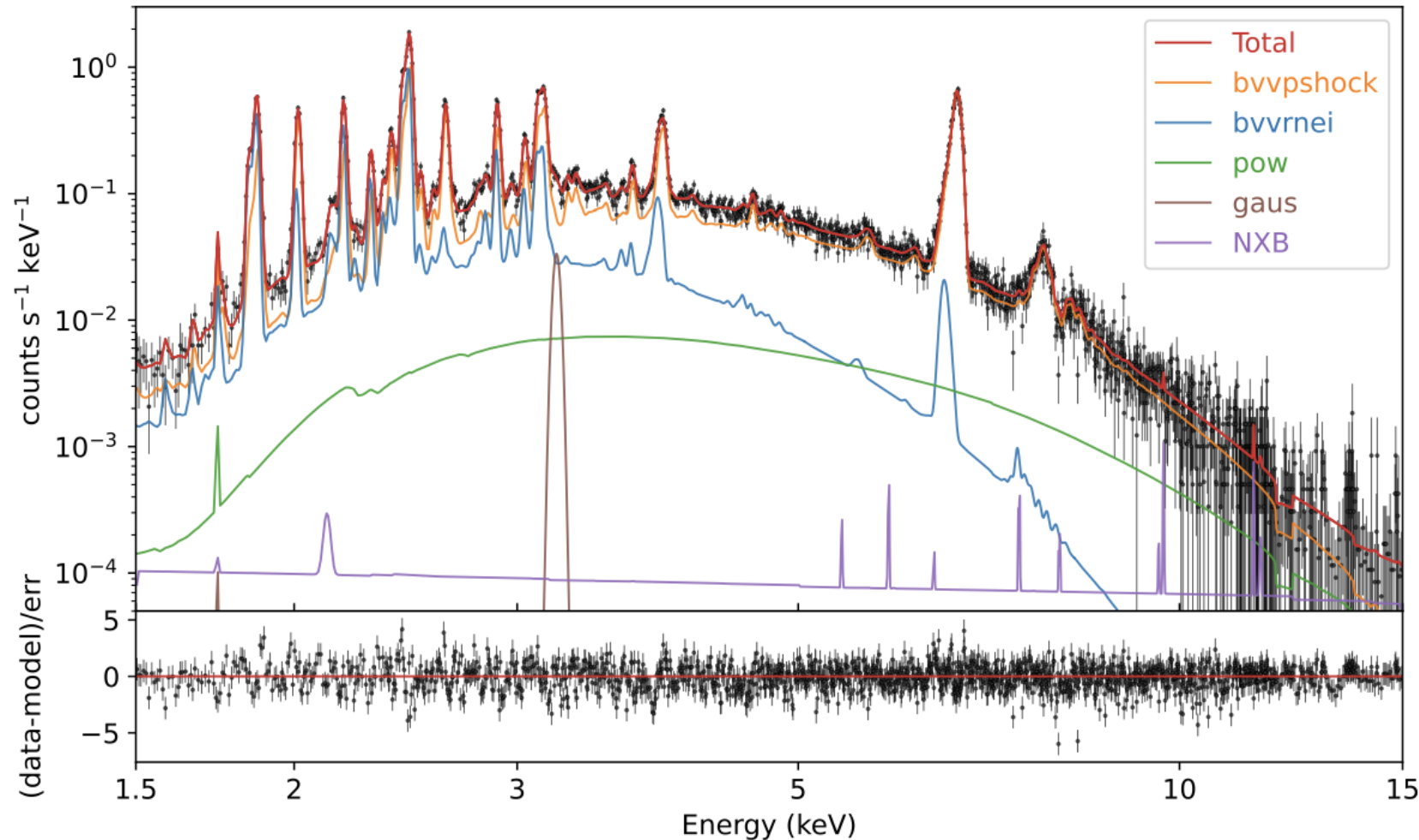
XRISM+25



- Improved atomic data: innershell excitation and DR data for C-, B-, Be-, and Li-like ions
- $K\beta$  to  $>K\epsilon$  lines are resolved with XRISM
- These data on Si, S, and Fe are more essential for measuring Ni, P, Cl, and K abundances due to line blending

# Fe $K\beta$ and Ni line complex

Cas A Plucinsky+25



Ni/Fe = 1 (AtomDB  
3.1.0)

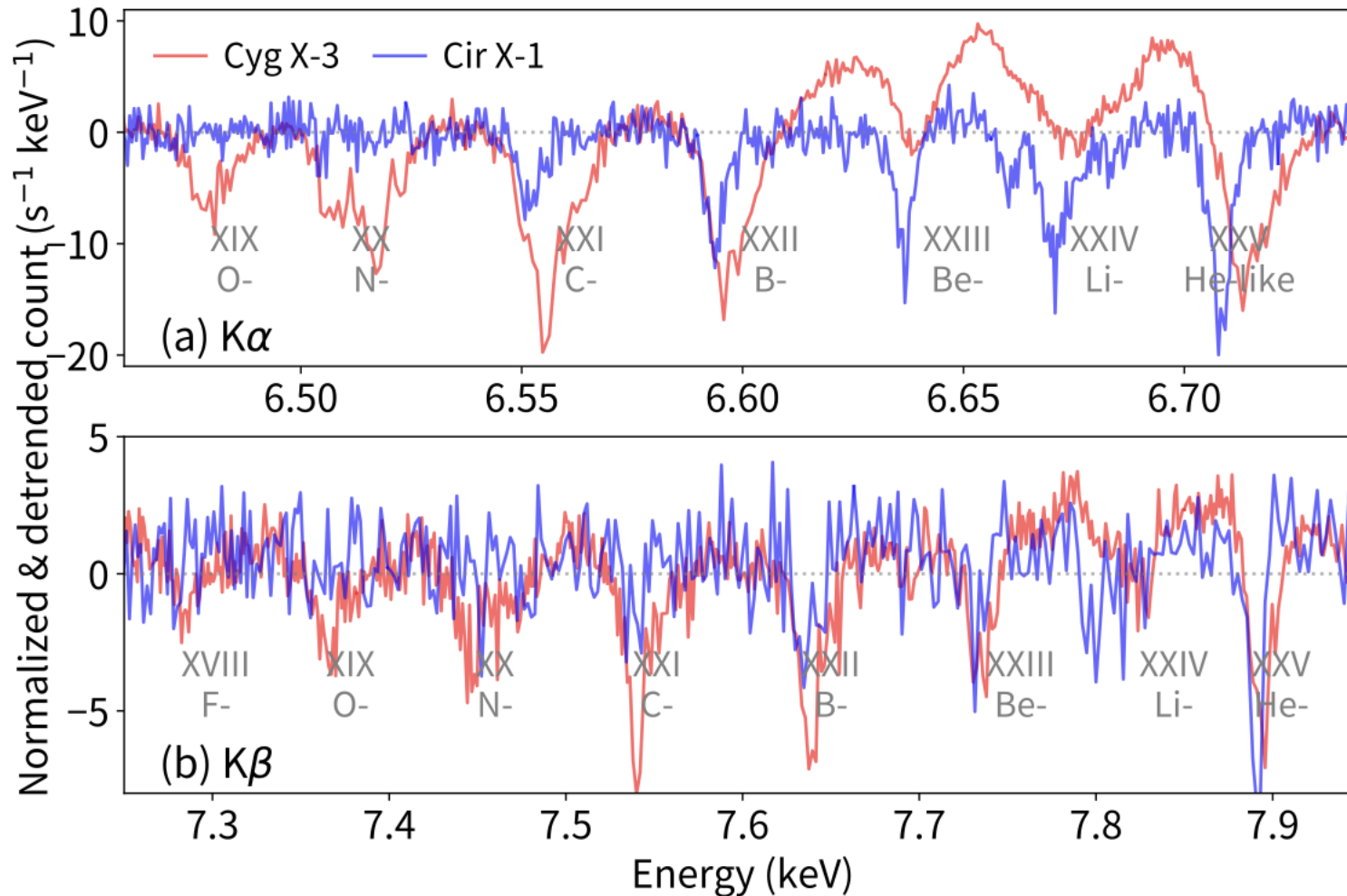
= 0.6 (SPEX  
3.08.01)

Both Fe and Ni are in  
intermediate charge  
states.

# Fe $K\beta$ lines in X-ray binaries

X-ray binaries

Tsujimoto+25

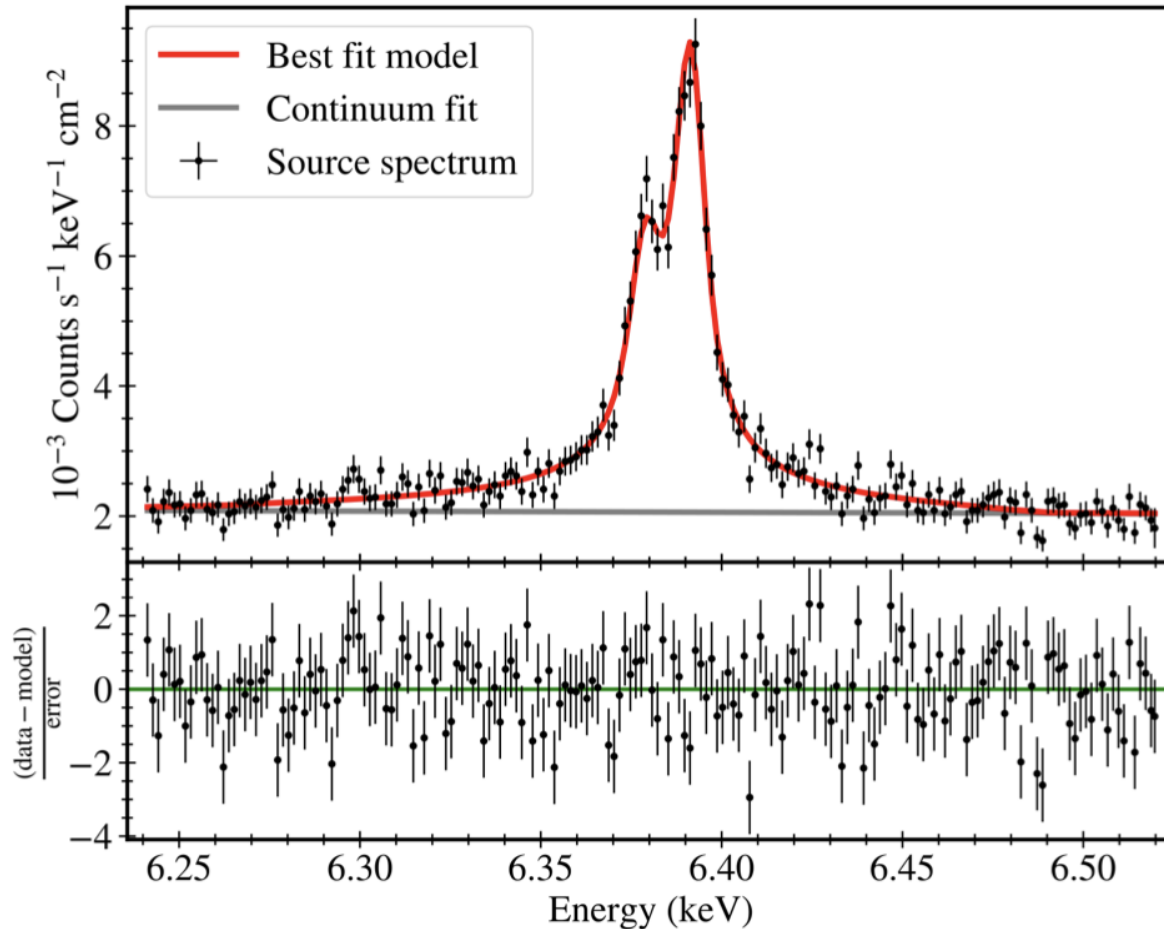


- Fe  $K\beta$  lines are better diagnostics of outflow velocity and optical depth, as  $K\alpha$  can be saturated or blended with scatter emission
- Atomic data still need to be improved on  $K\beta$

# ~Neutral Fe K $\alpha$ lines

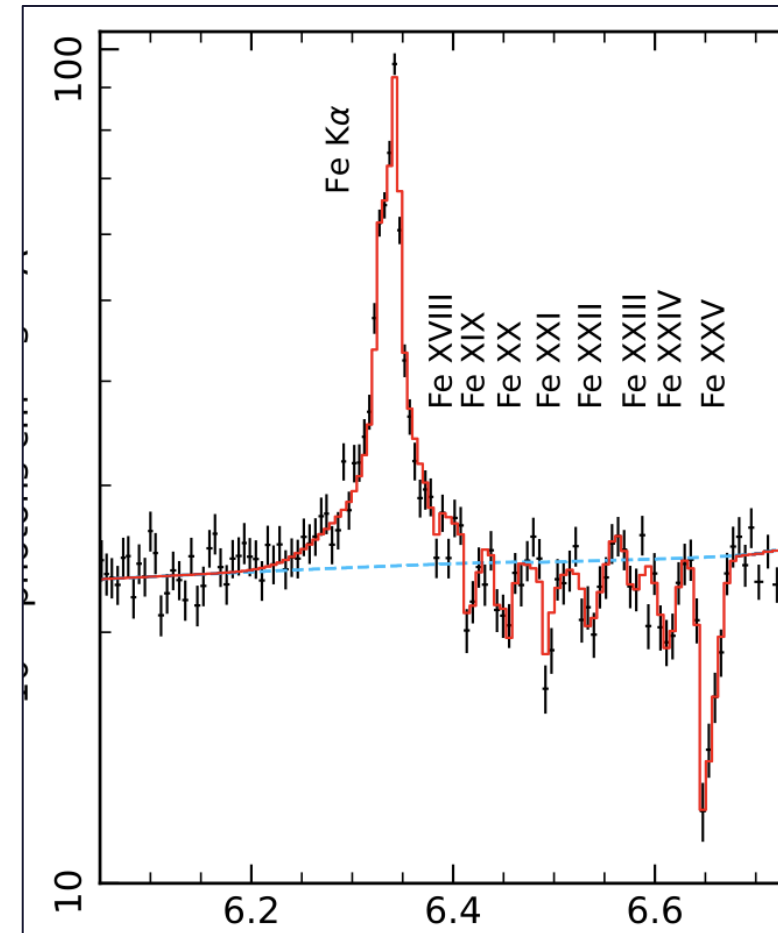
Cen A

Bogensberger+25



NGC 3783

Mehdipour+25

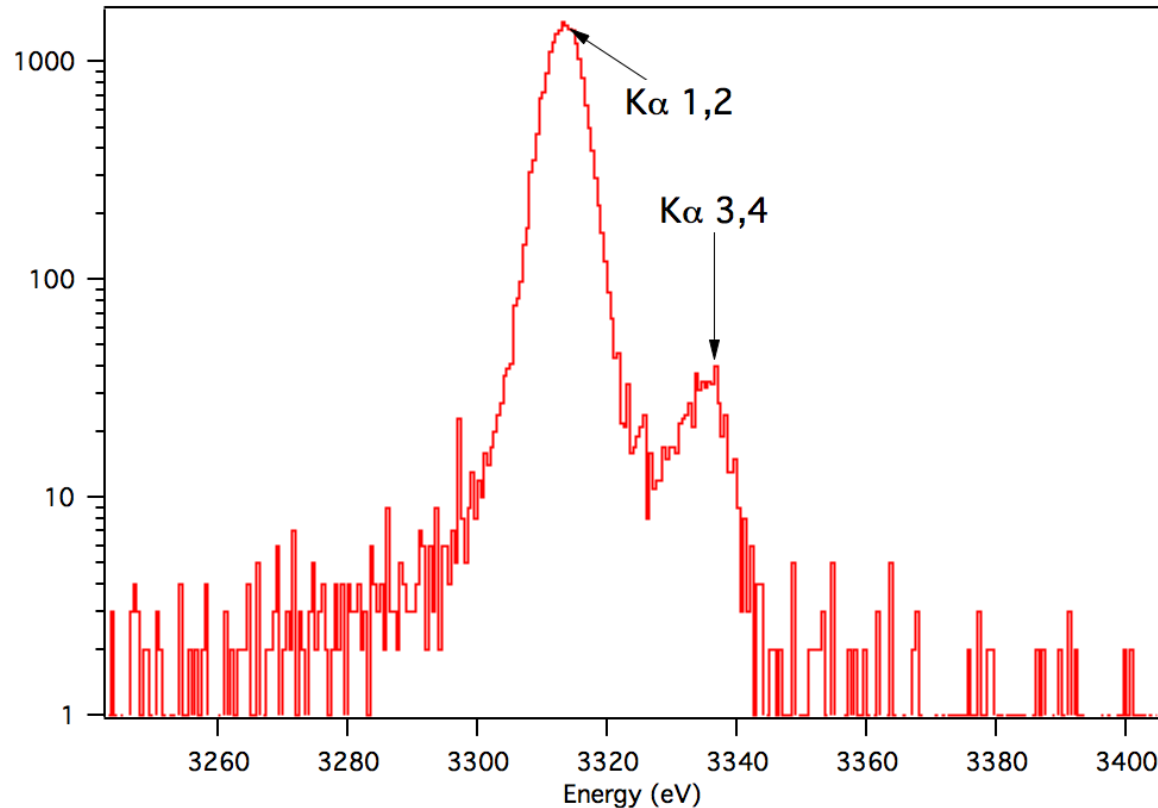


- Fe K $\alpha$  main lines are complexes of large number transitions forming two main peaks
- No existing atomic model. For now, use Holzer+1997 lab model

# ~Neutral Fe $K\alpha$ lines - satellites

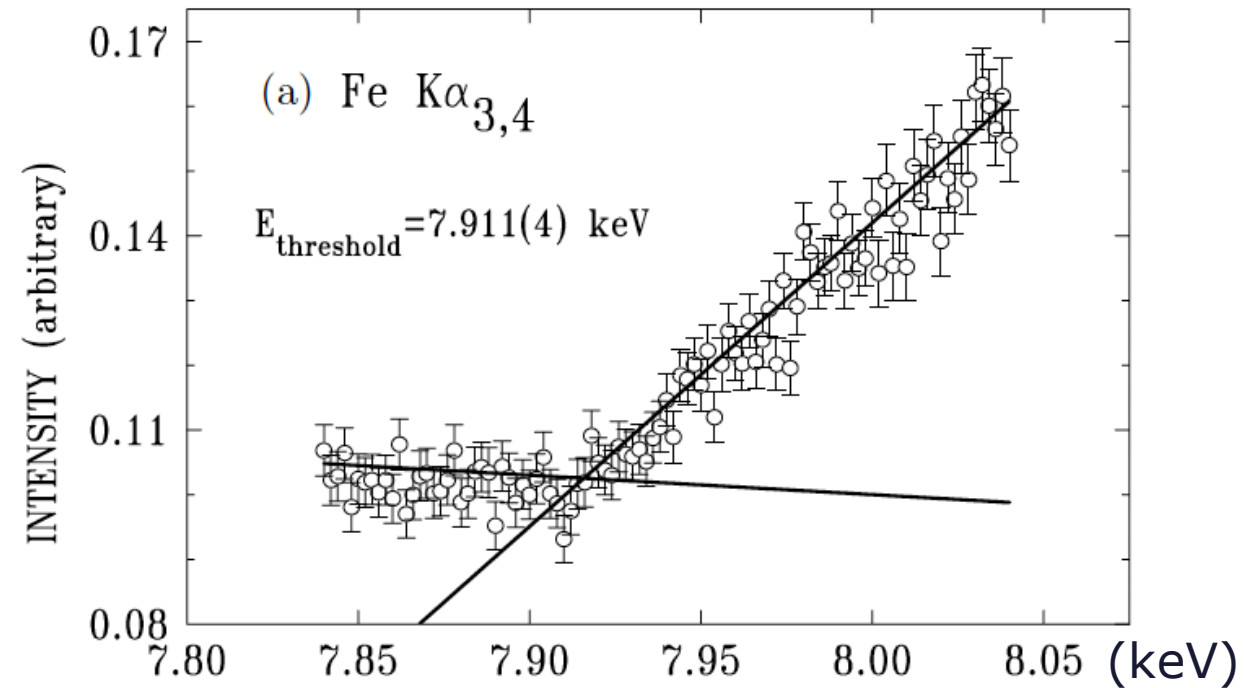
Lab data of K

Porter+



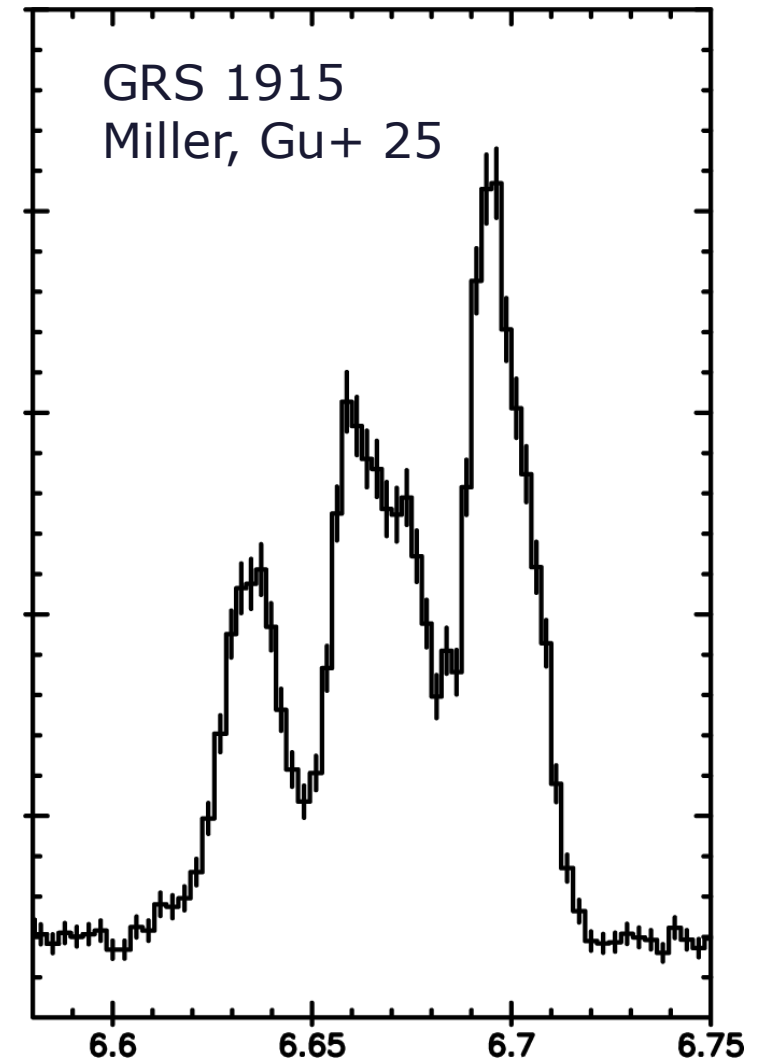
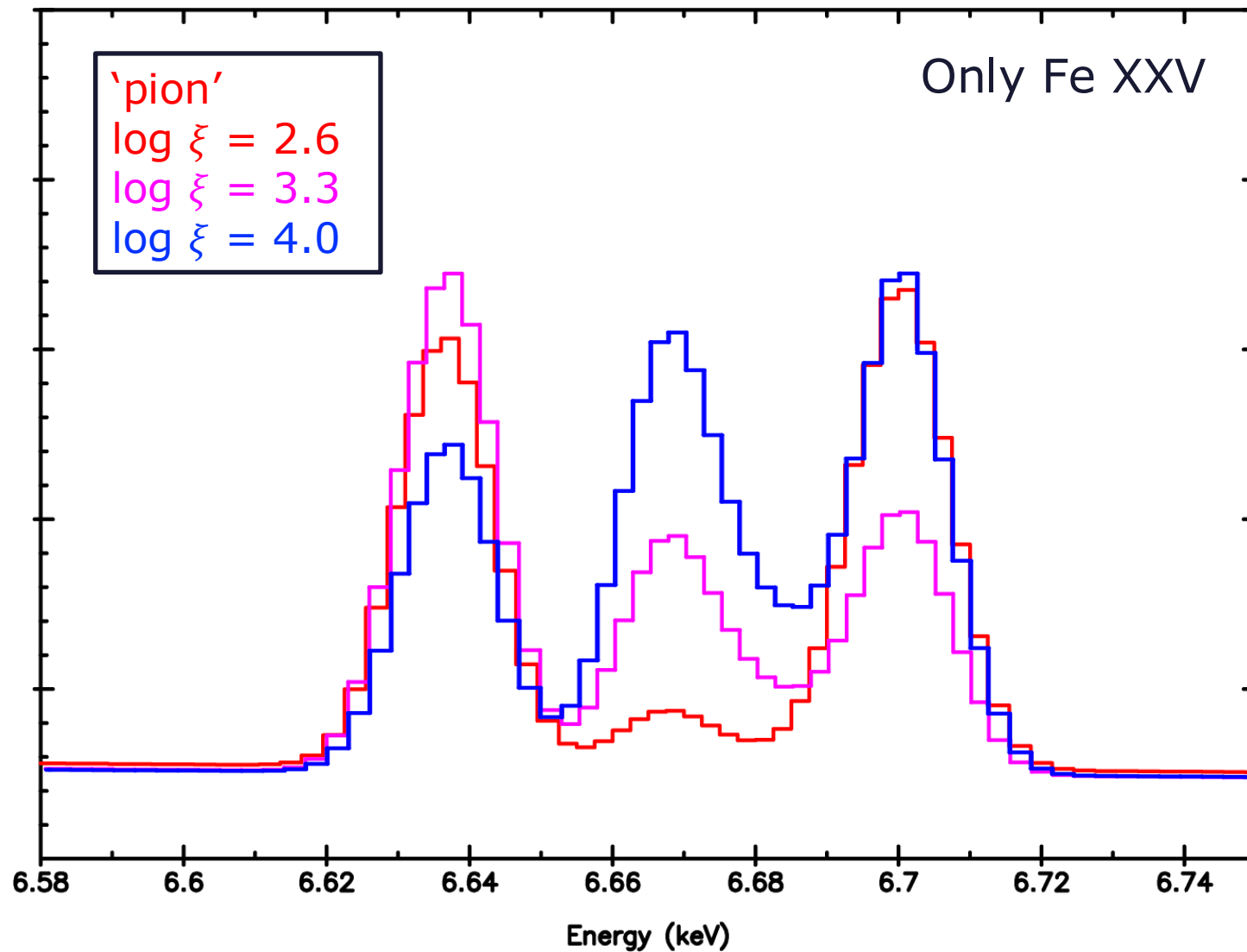
Lab data of Fe

Diamant+05

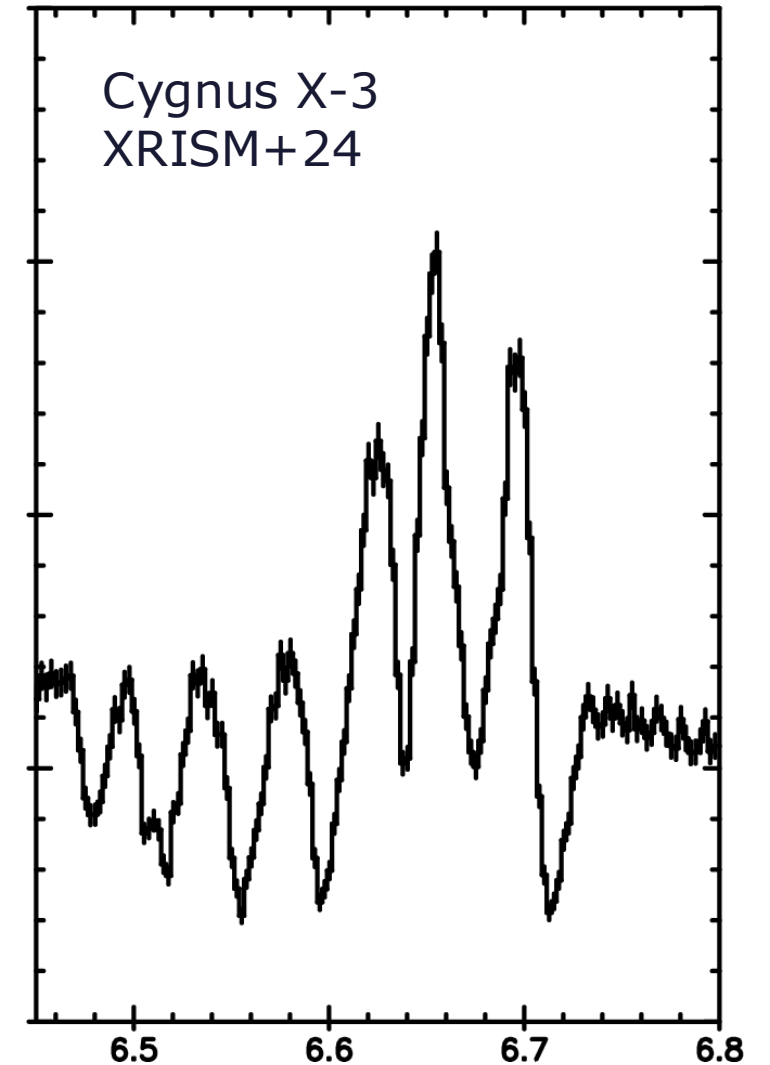
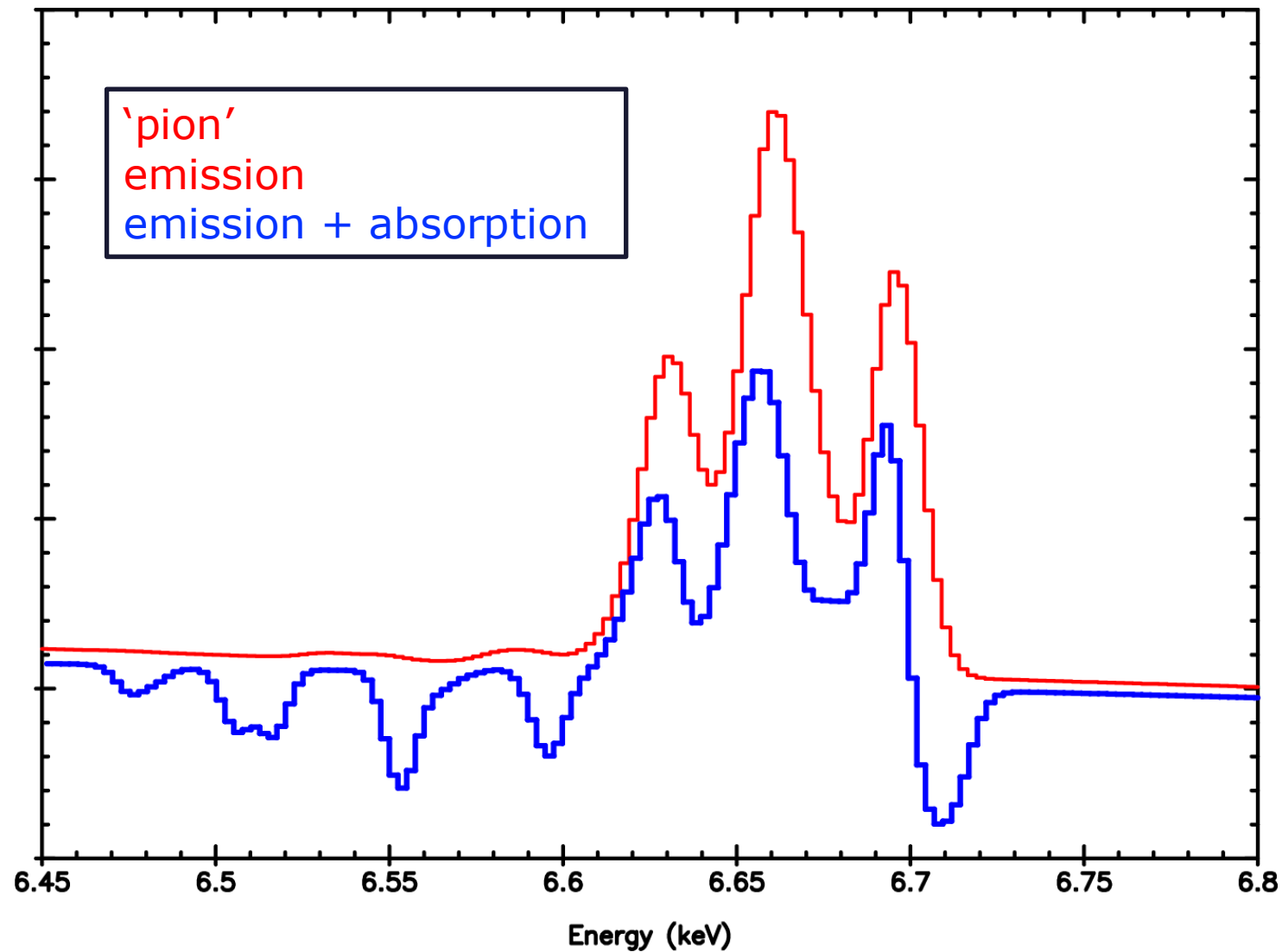


- Satellites are detected in a few objects with Resolve
- Diagnostics of incident photon spectrum on the reflector (Vander Meulen+)

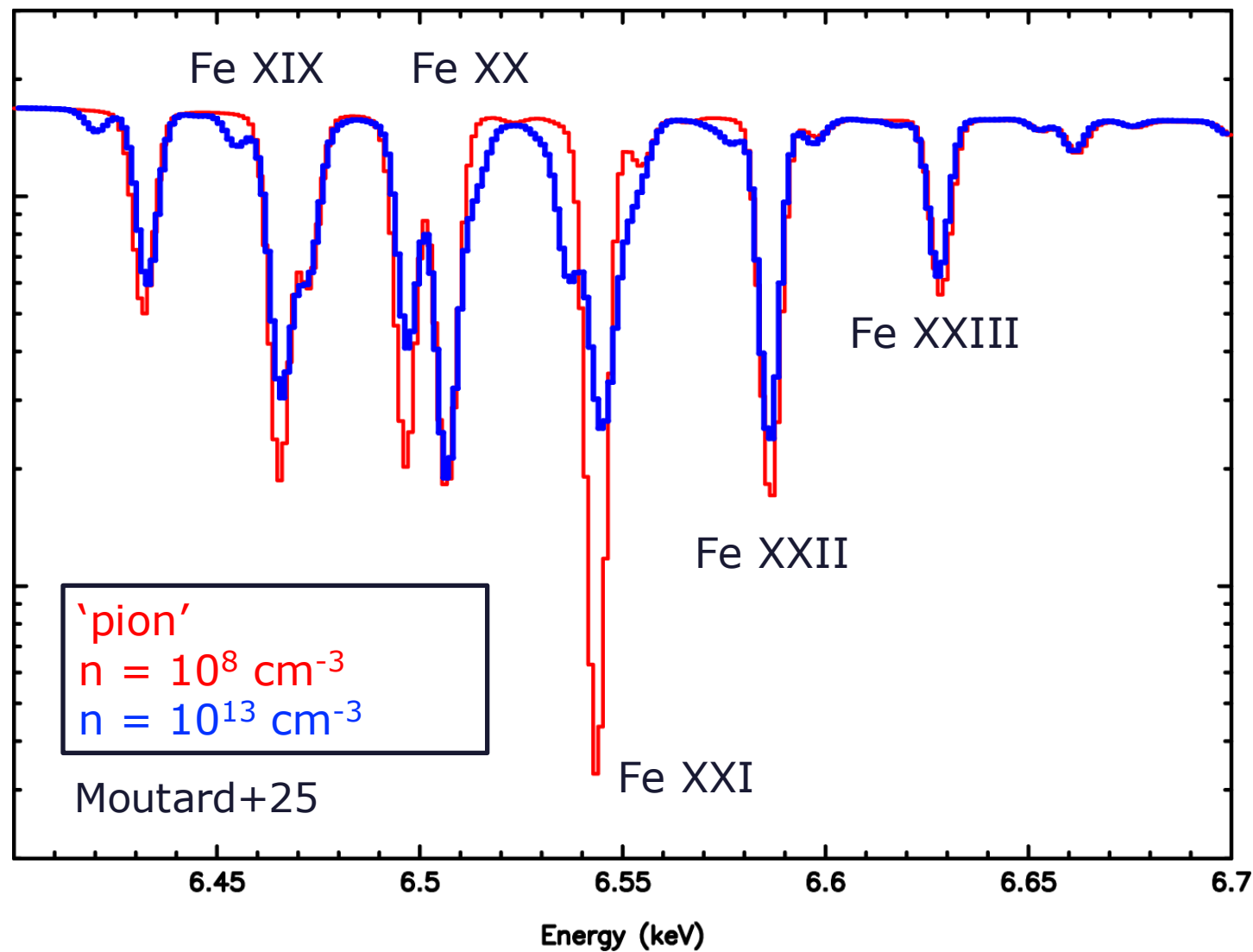
# Photonionized emission



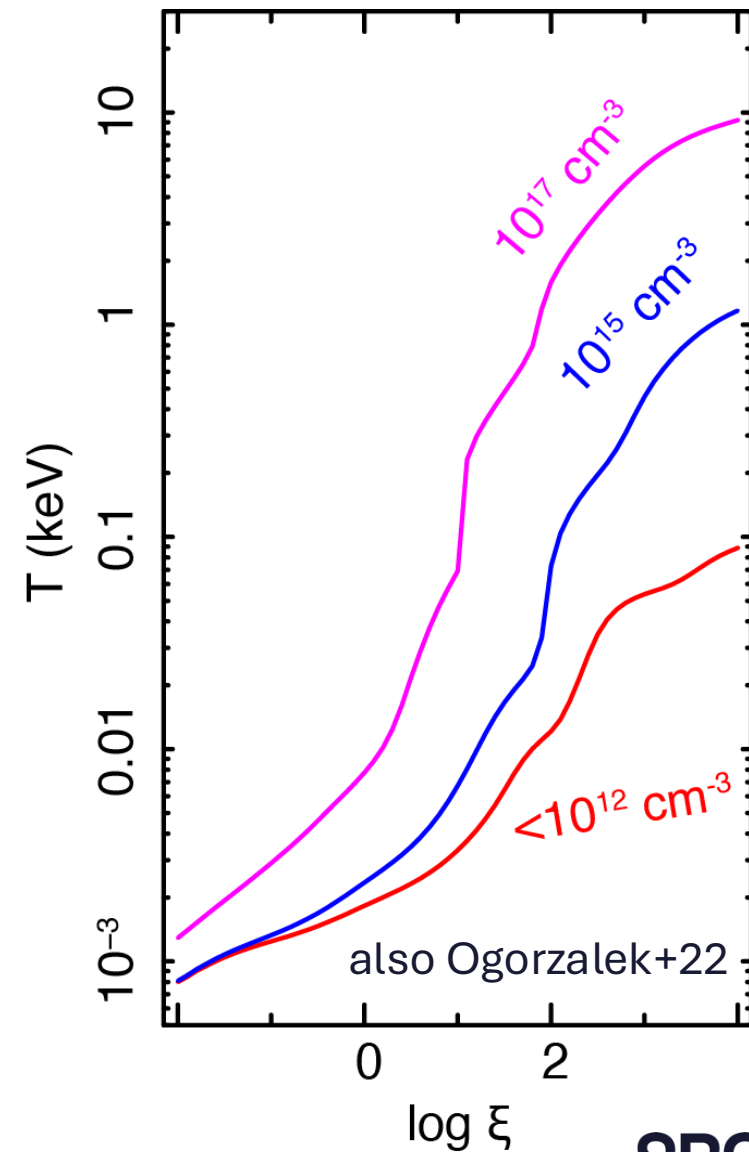
# Both emission and absorption present



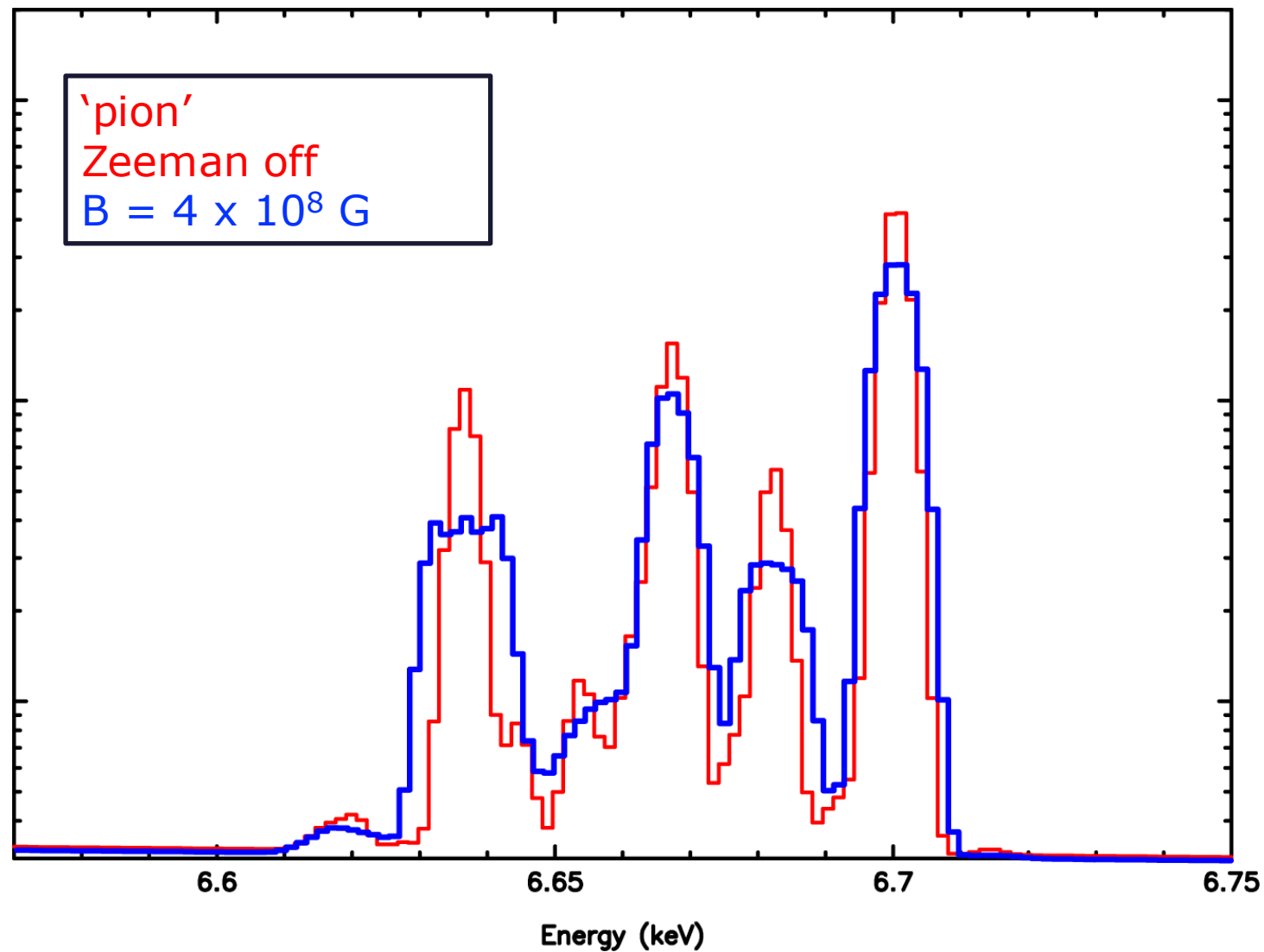
# Effects from high density plasmas



Density effect on Ion fraction



# Effects from strong magnetic field



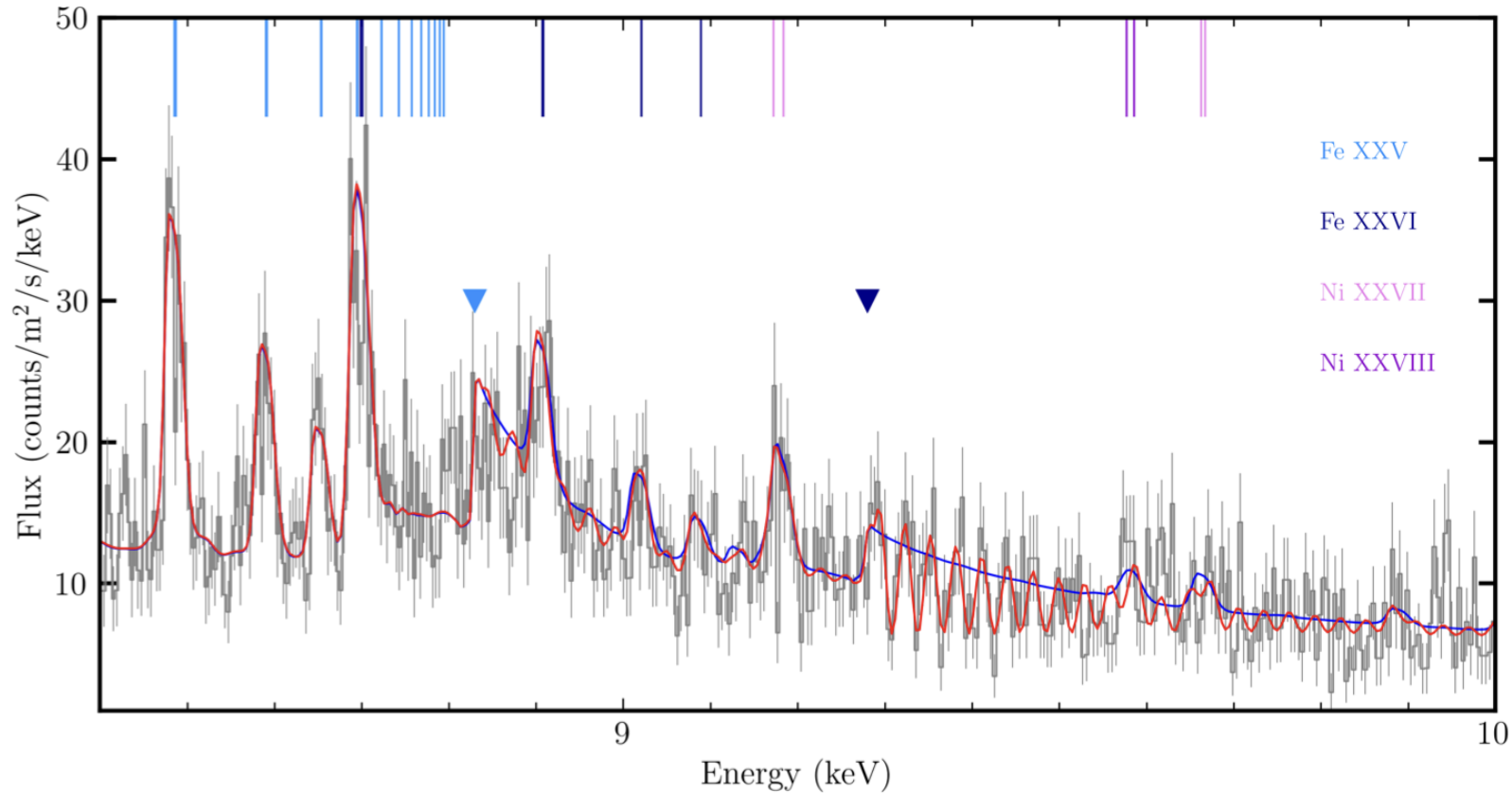
Zeeman effect



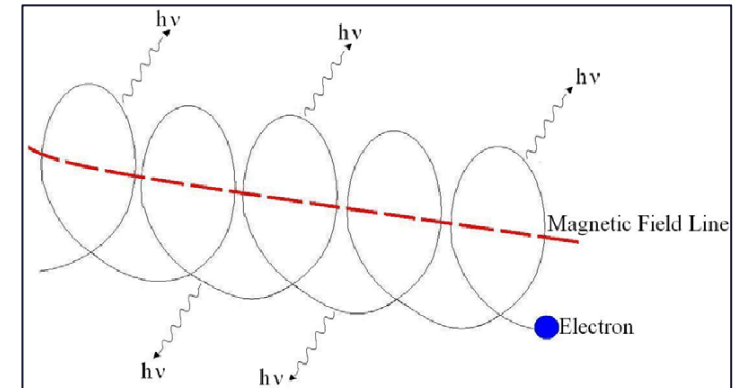
# Landau quantum effect

GRS 1915

Miller+25



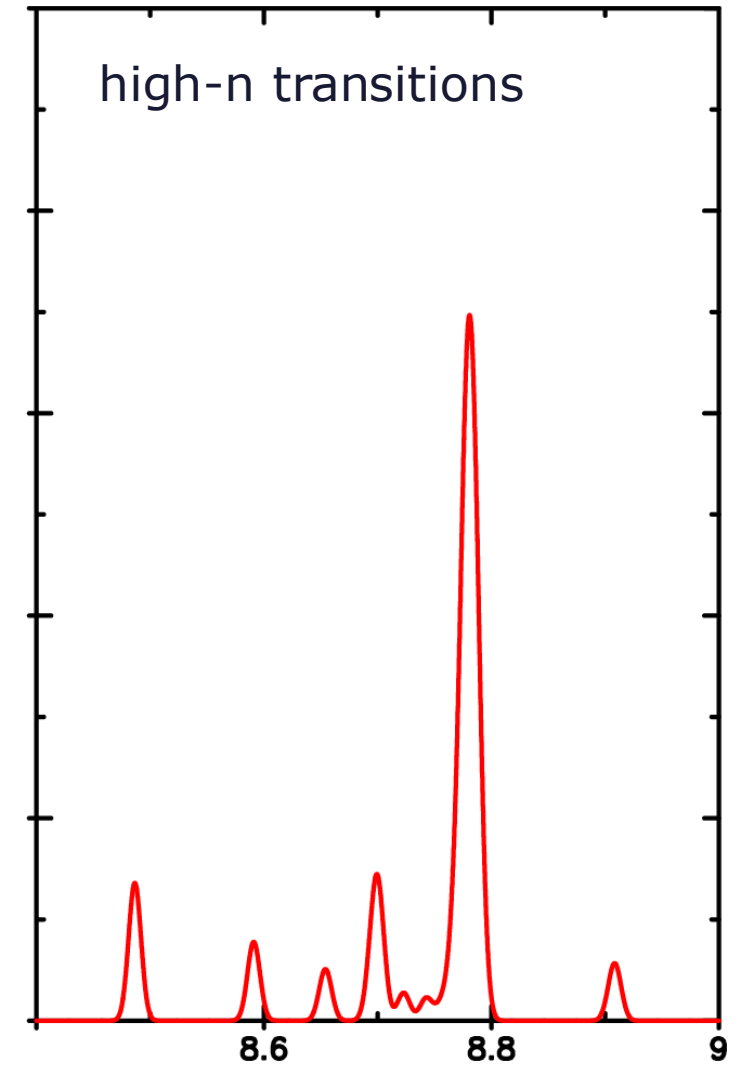
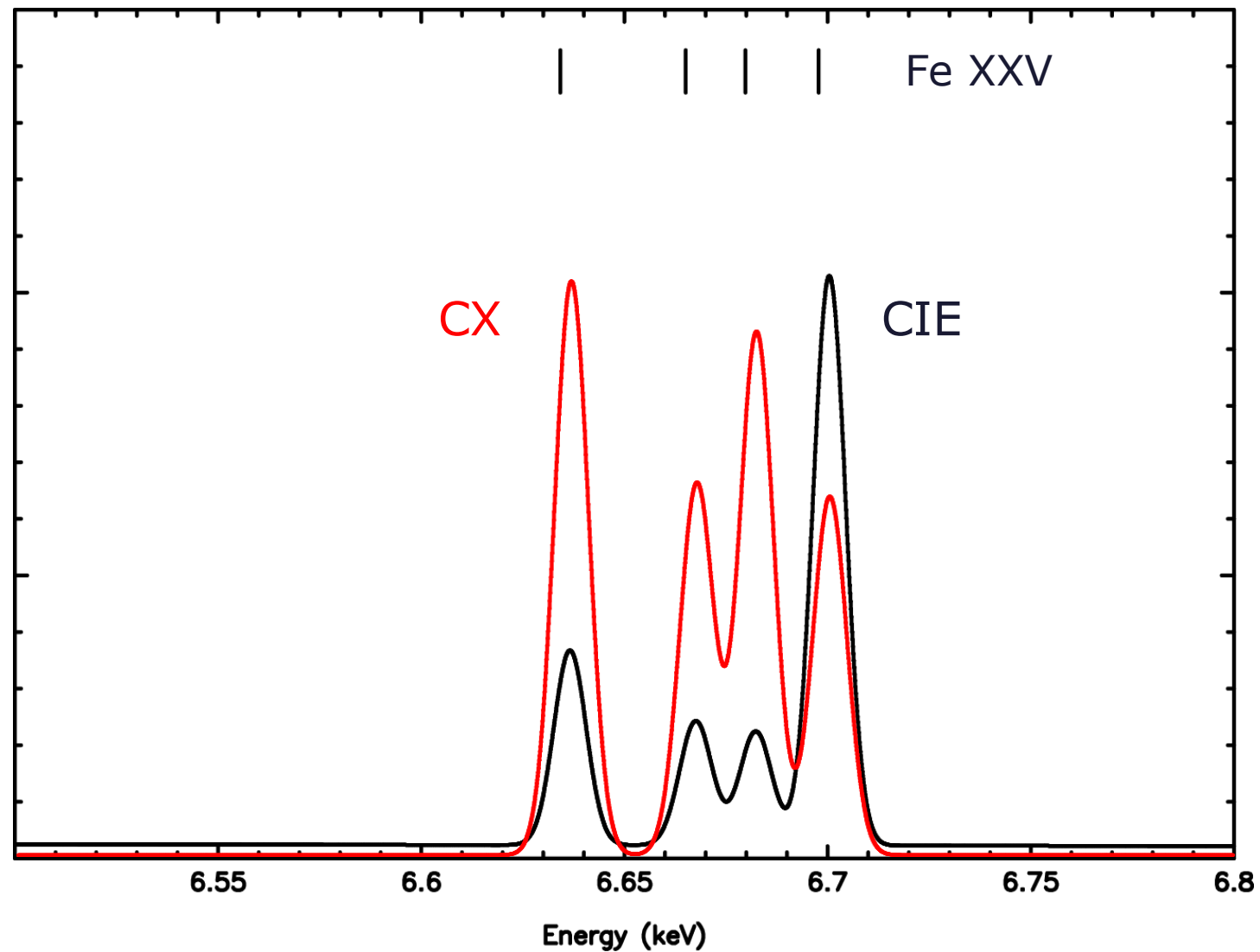
Quantum levels from  
cyclotron of free electrons



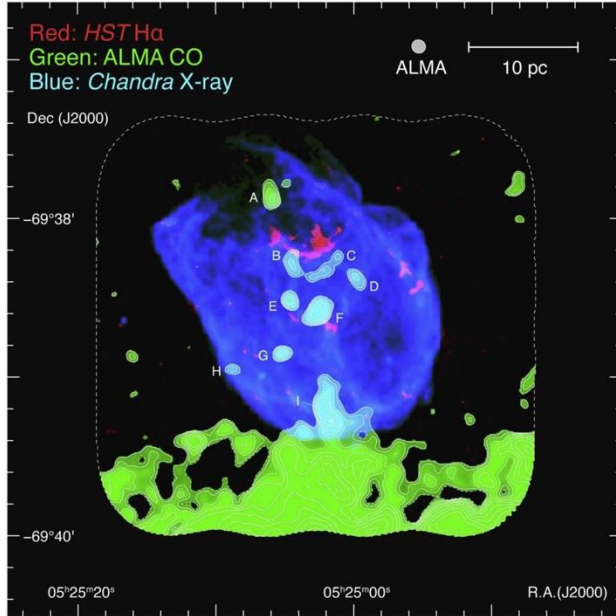
Available in SPEX v3.08.02

Either  $2.5 \times 10^9$  G in the accretion disk  
Or  
A set of well-structured UFOs

# Charge exchange emission



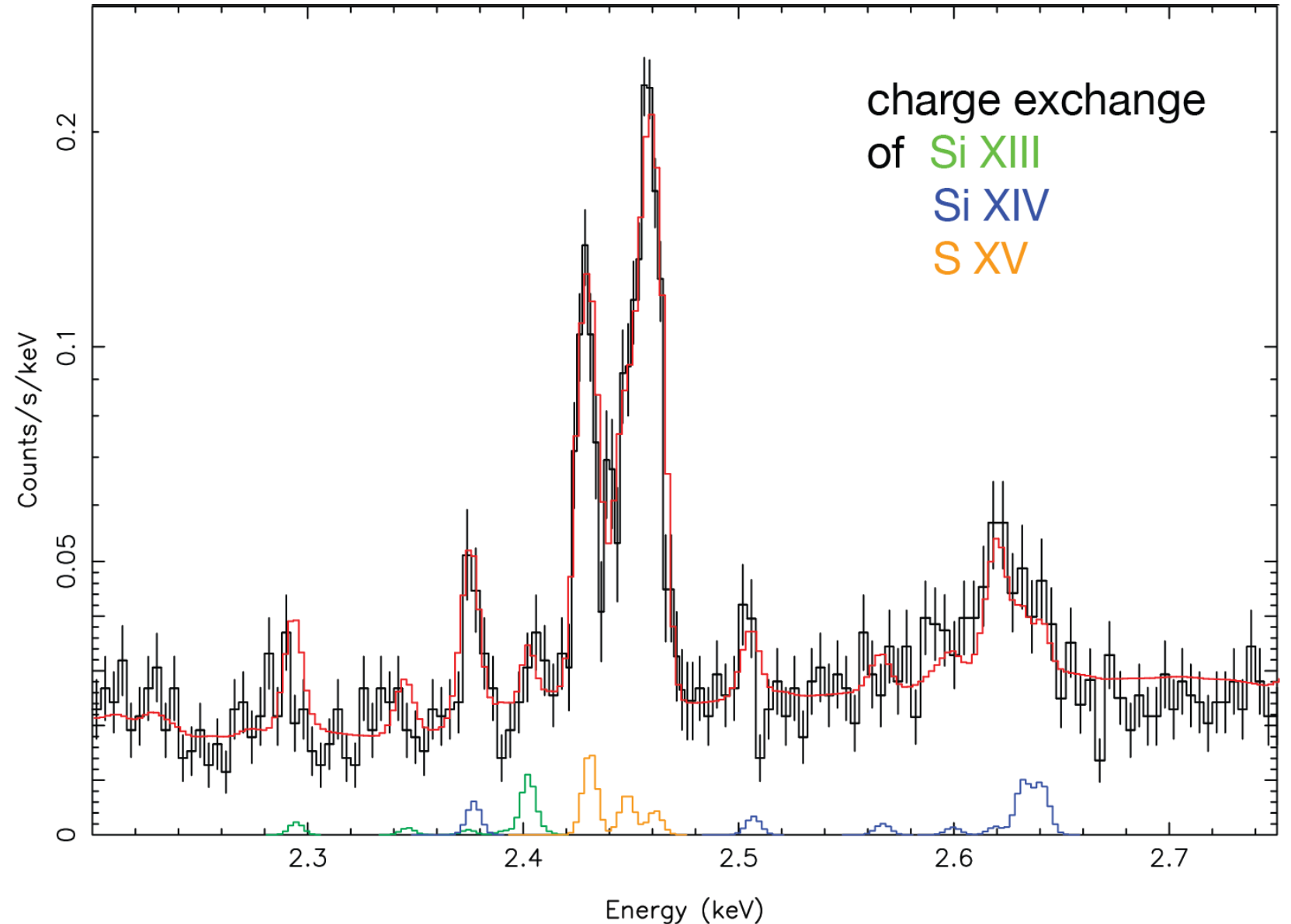
# N132D: potential charge exchange



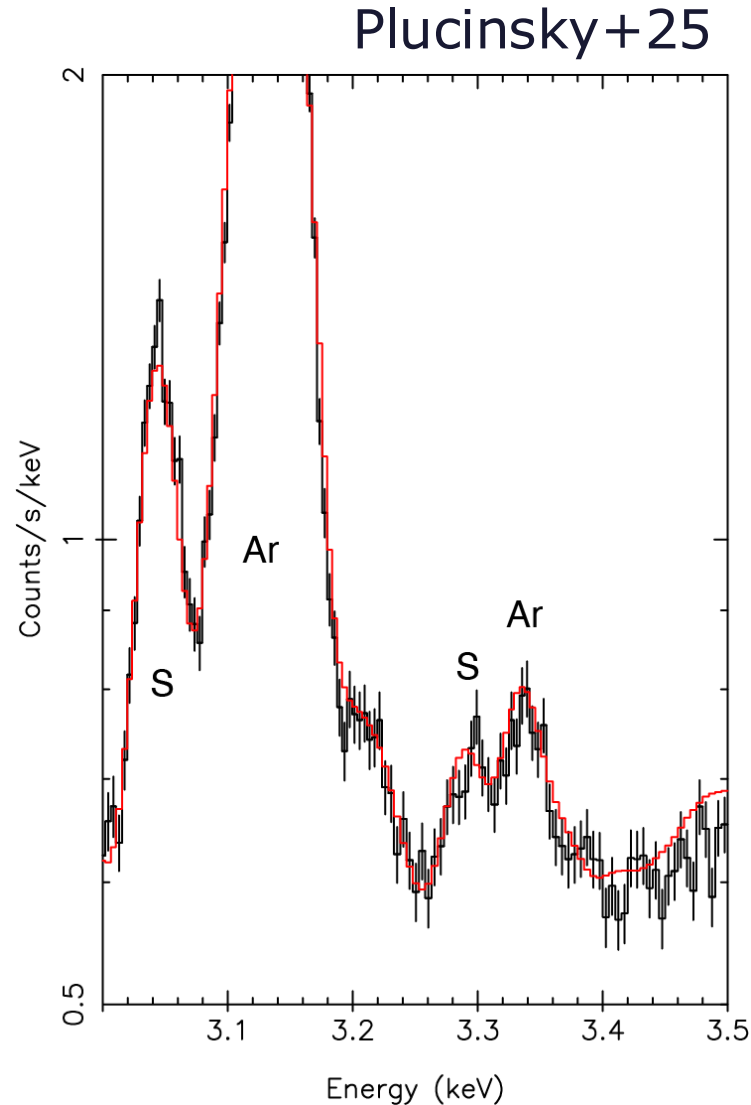
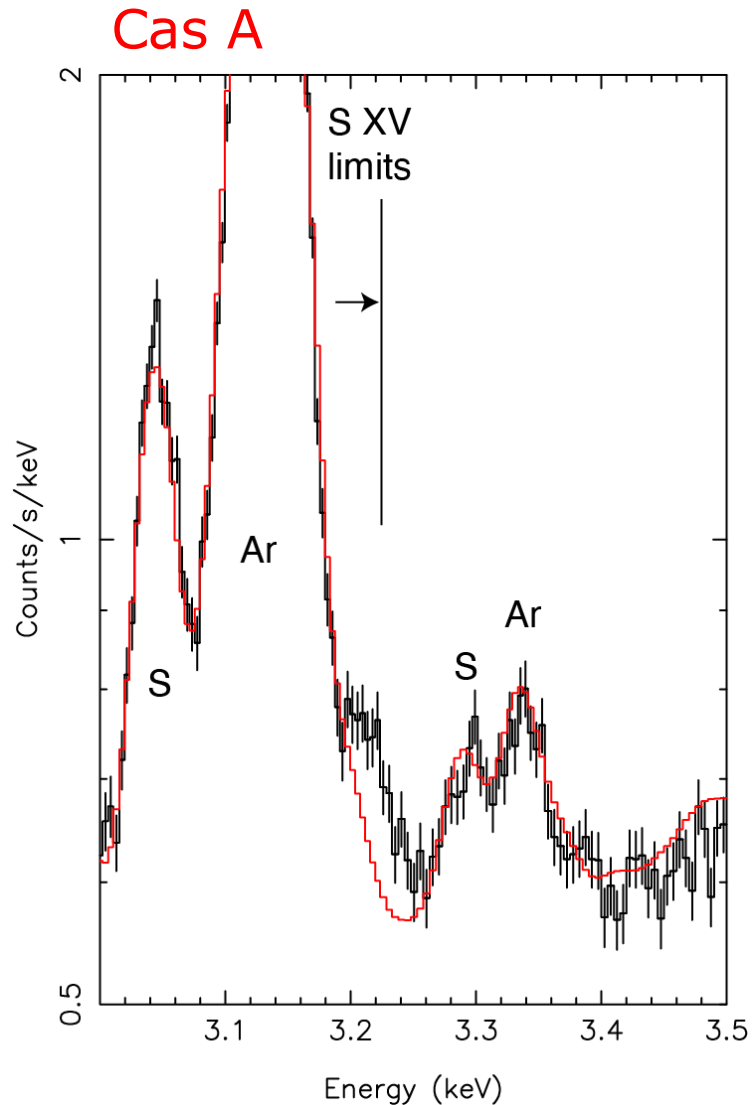
Shock-neutral  
interaction

Diagnostics of freeze-in  
temperature

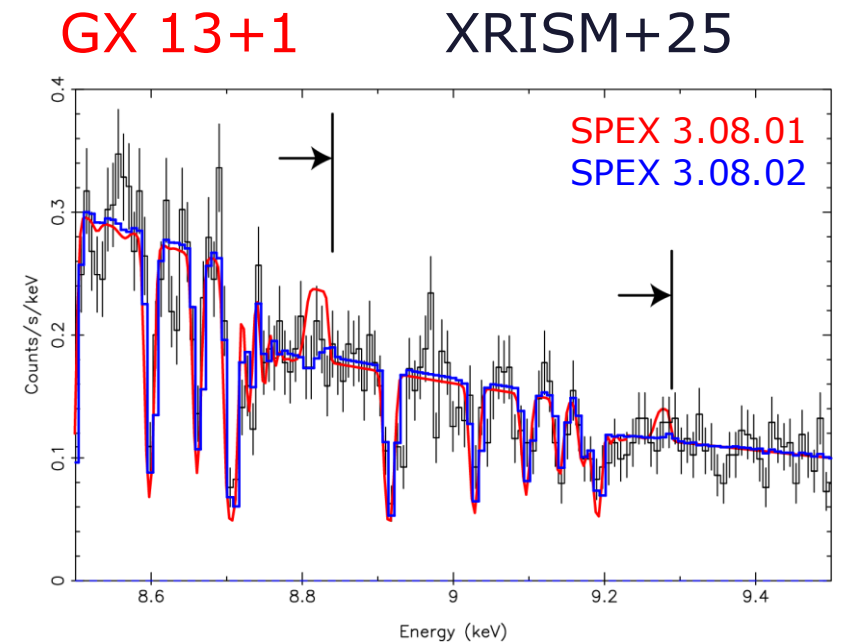
Gu+2025



# Rydberg series up to very high n states



- Data up to **n=52** for Si, S, and Fe by FAC
- Essential for identifying RRC components in supernova remnants, or for measuring density based on Debye screening from ejecta of X-ray binaries



# Summary

- Plasma model has demonstrated the power analyzing XRISM data from collisional and photoionized plasmas, both in and out of equilibrium
- New SPEX provides tools for unusual conditions, such as non-thermal, high density, or strong magnetic fields
- XRISM and XIFU are propelling lab astro forward

SRON FDM/TES  
detector ( $\sim 2.5$  eV)  
on Heidelberg EBIT  
since 2024 summer

