

Probing the Dark Sector with Electron Recoil

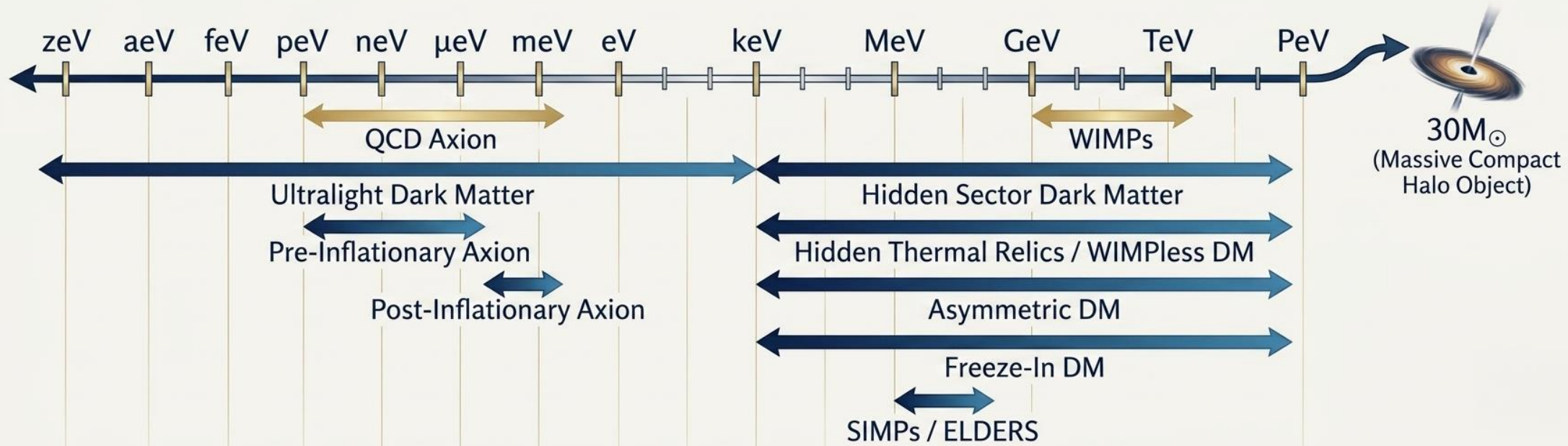
Xavier Bertou (IJCLab, IN2P3/CNRS)

1. Dark Matter and low mass searches
2. Can I use my Bonafide Xenon/Argon detector to search for Electron Recoil?
3. New technology? The Skipper CCD
4. The Last 2 years
5. The Next 2 years (?)

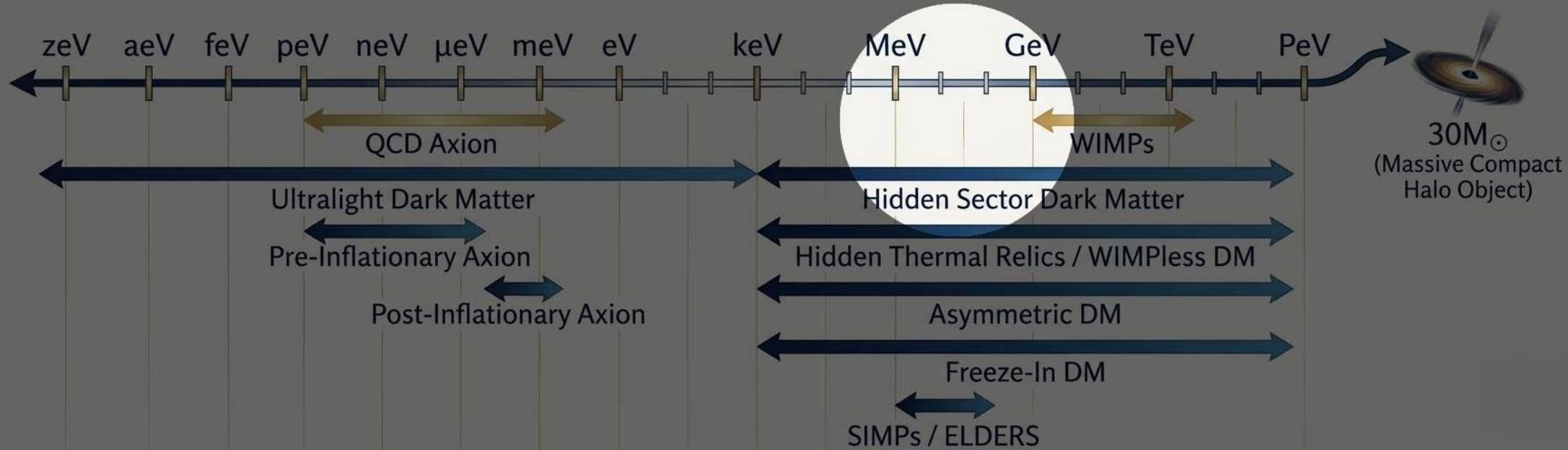
(1)

Dark Matter and low mass searches

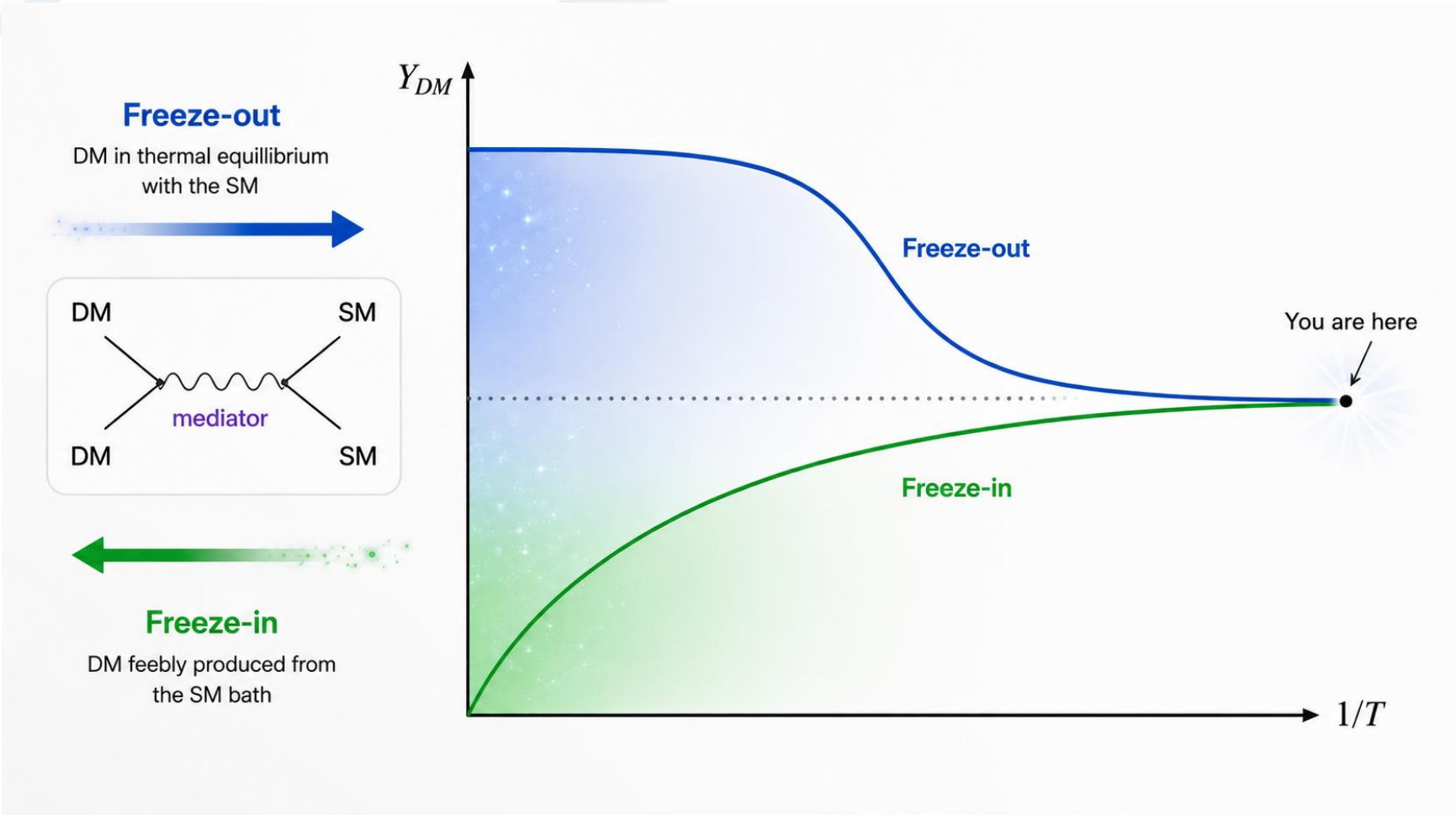
THE COSMIC DARK MATTER LANDSCAPE

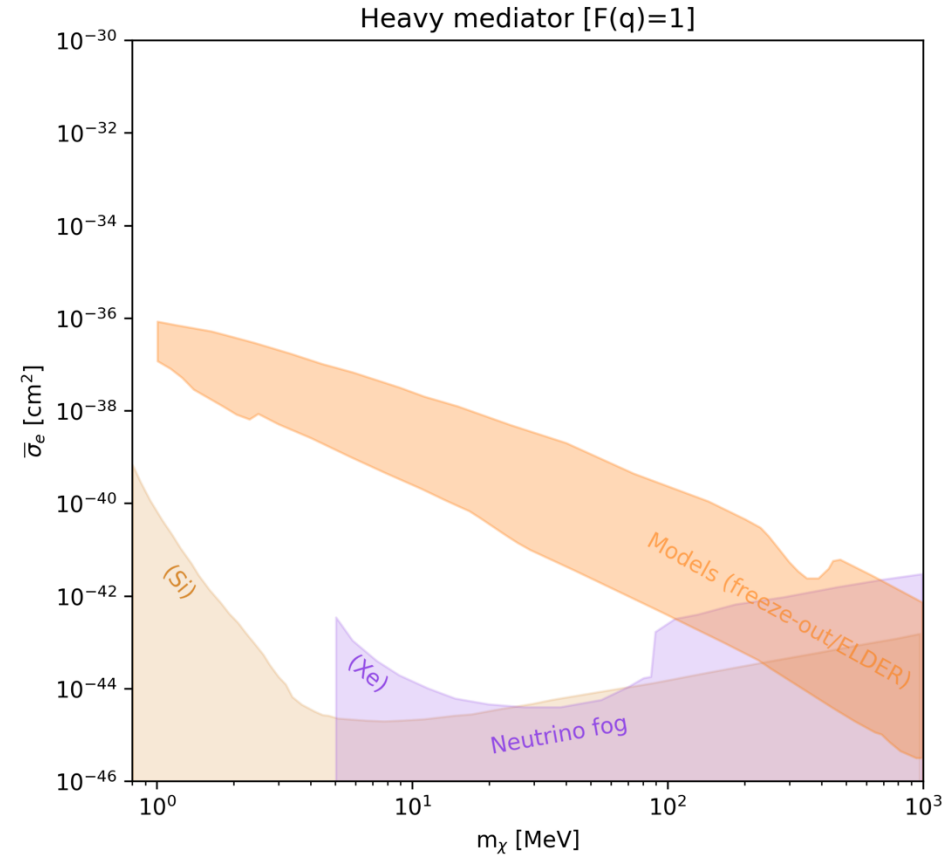
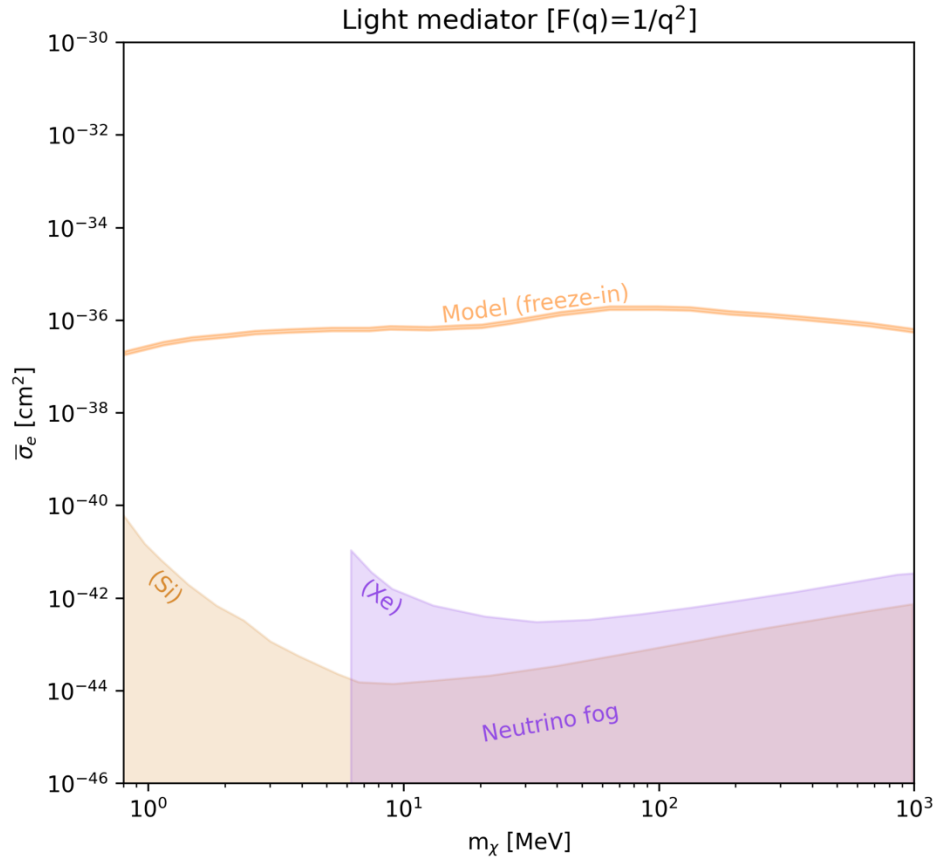


THE COSMIC DARK MATTER LANDSCAPE



- Add a $U(1)$ symmetry
 - Minimal renormalizable portal to the Standard Model
 - Gauge invariance: kinetic mixing with SM $U(1)$
 - After EWSB, dark photon couples to electric charge
 - Light DM opens the $< \text{GeV}$ scale
 - Nuclear recoils suppressed
 - Ideal to observe as electron recoil
 - Need to observe much smaller energy deposits

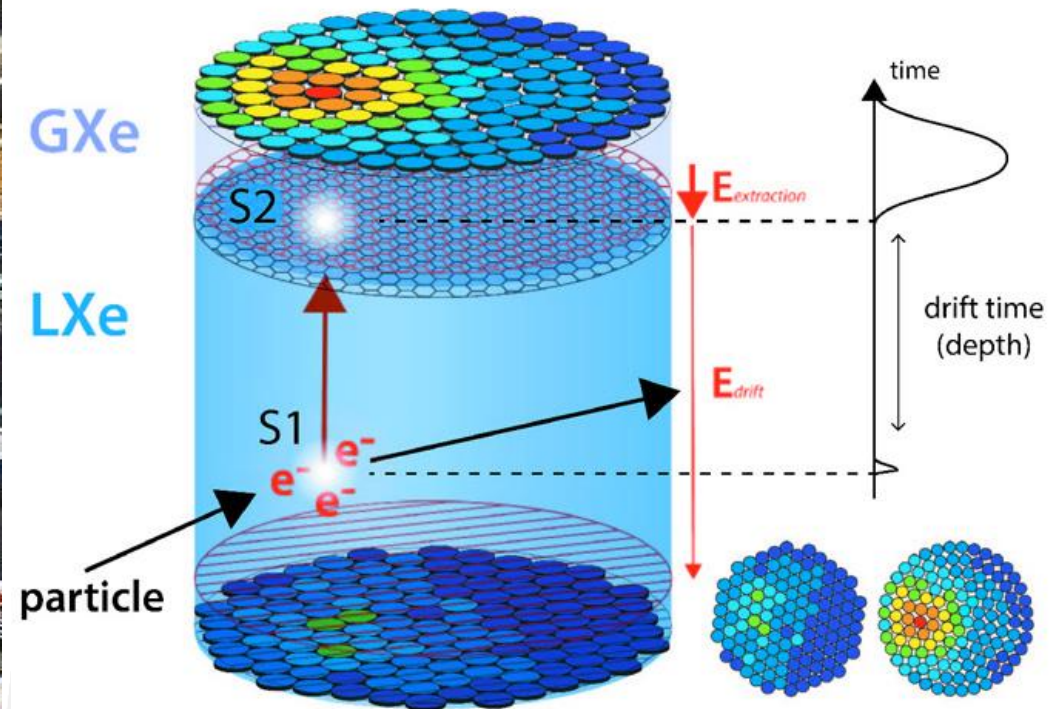




Carew et al.,
 Phys. Rev. D 109,
 083016 (2024)

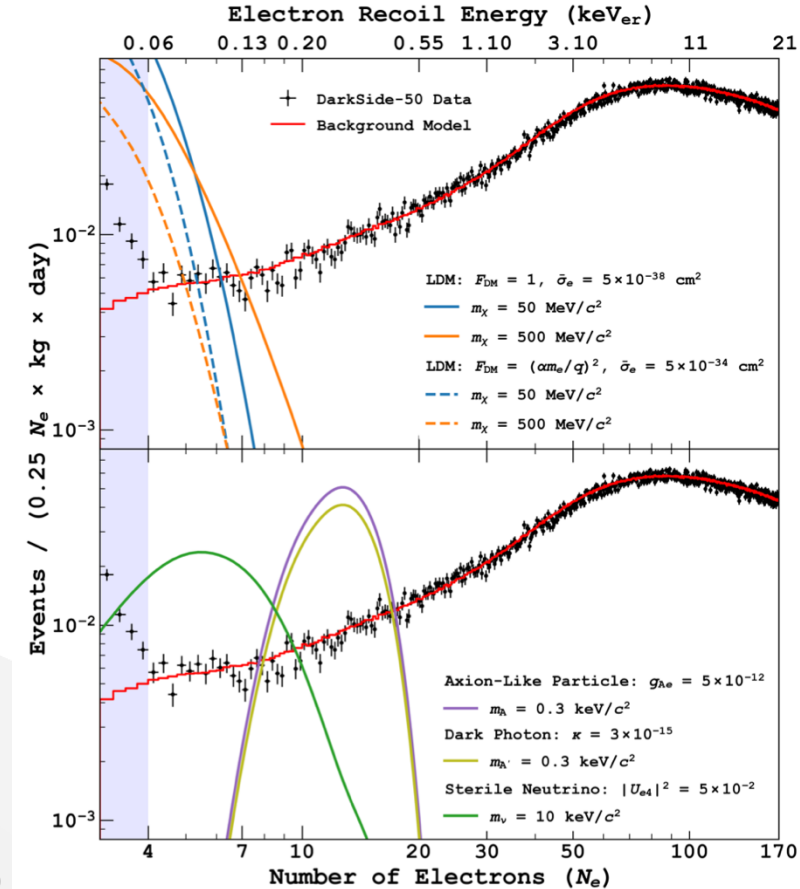
(2)

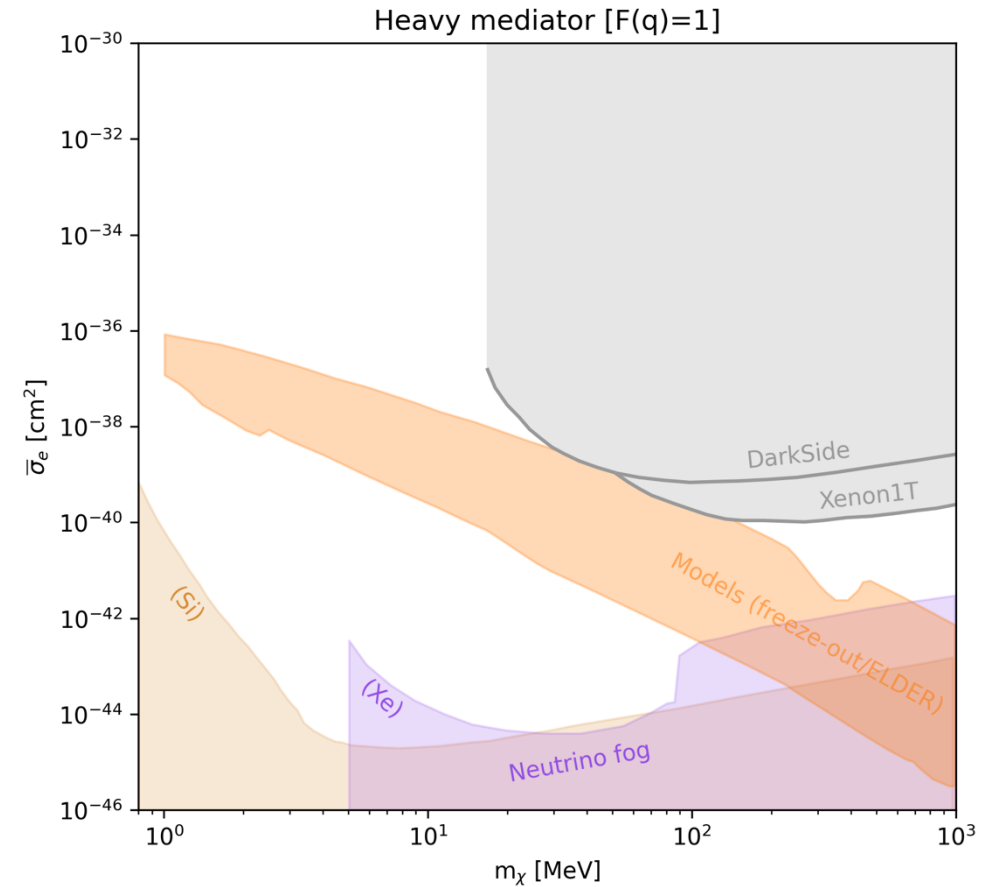
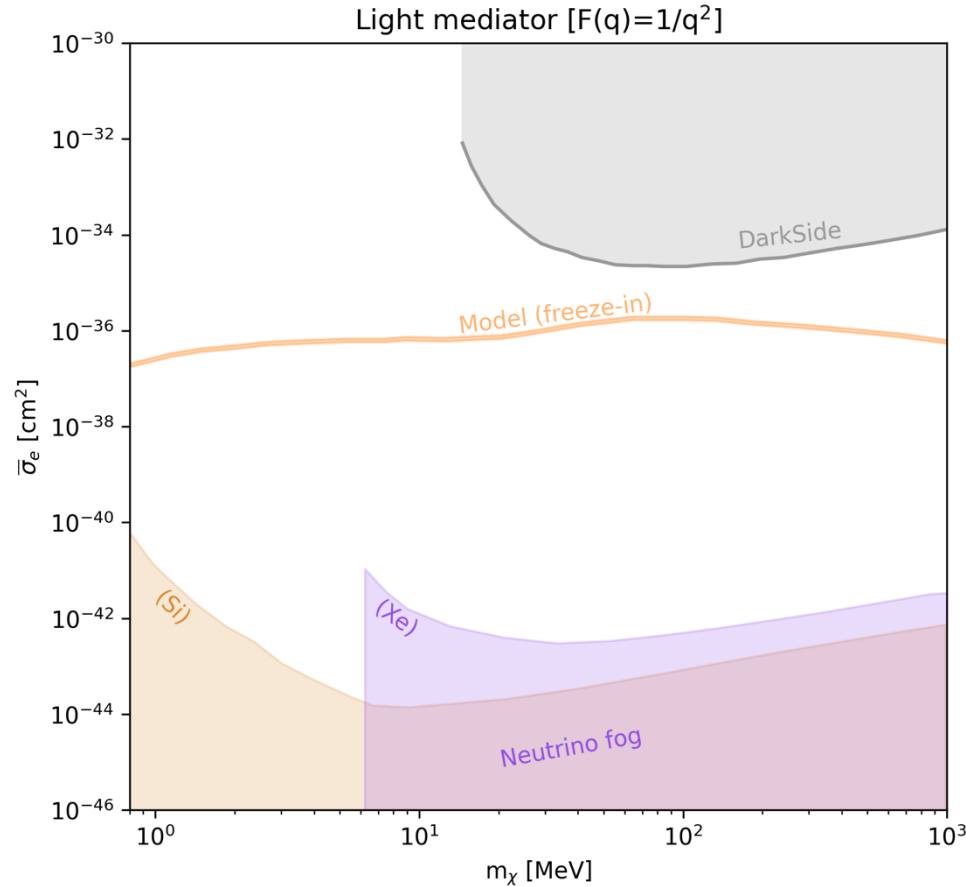
Can I use my Bonafide
Xenon/Argon detector to search
for Electron Recoil?



- Electron recoil accessible as an S2 only search
- Losing background rejection capability
 - Ok with a background model
- Usually unreducible background at a few electrons

The DarkSide-50
Collaboration
Phys.Rev.Lett. 130
(2023) 10, 101002

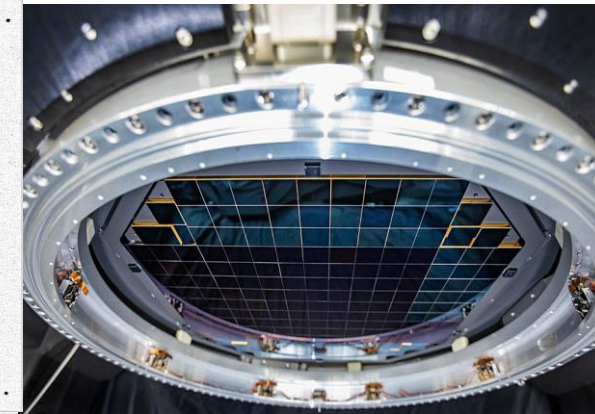
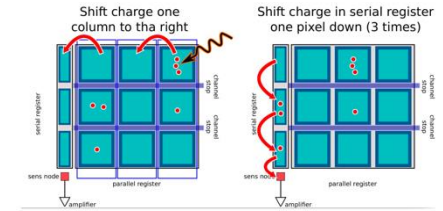
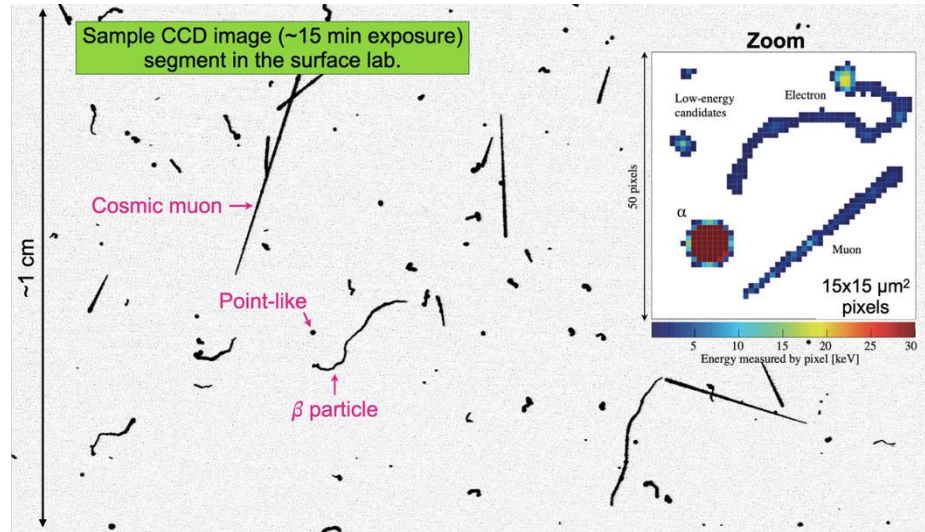
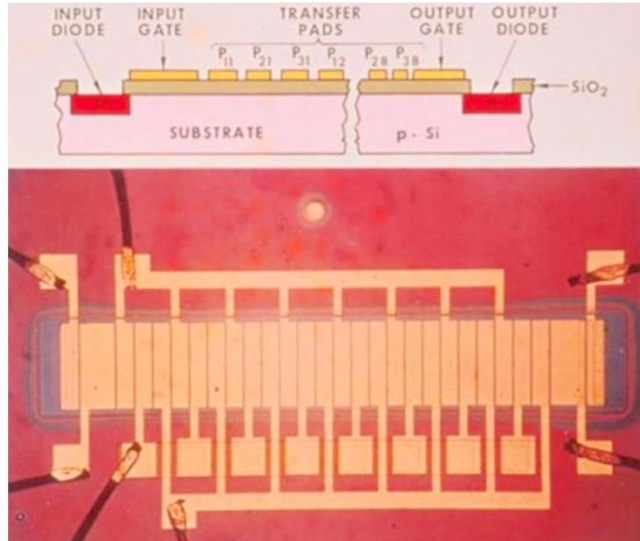


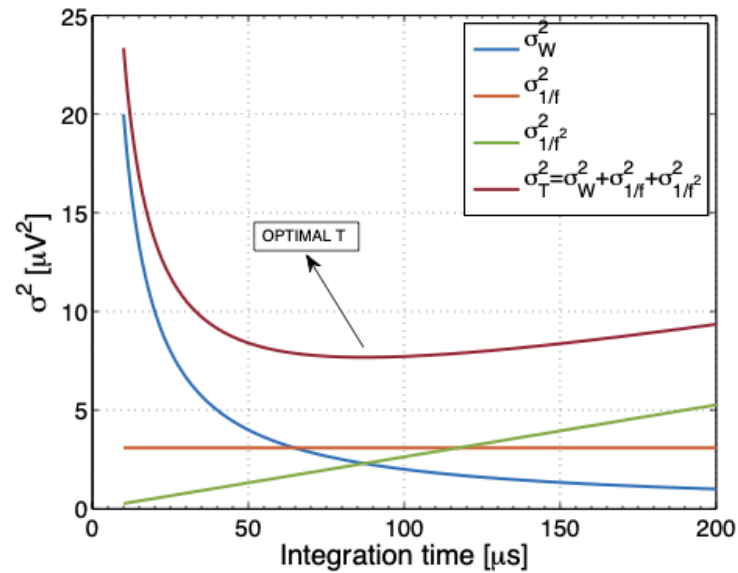
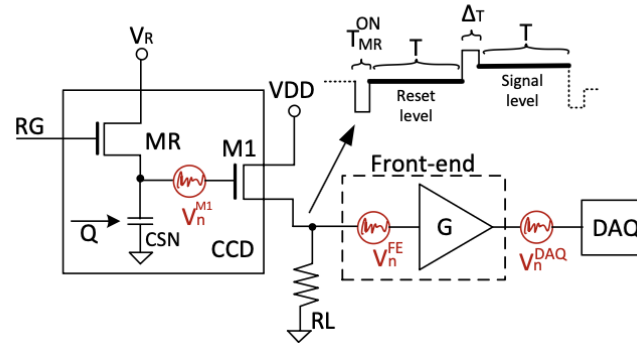


(3)

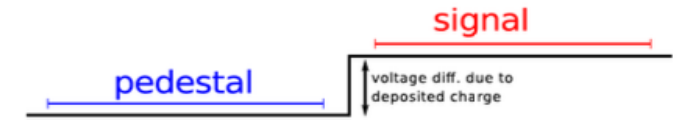
New technology? The Skipper CCD

- CCD were invented in 1969 by Boyle and Smith (Bell Labs) as a memory device. It works better as an image sensor and was quickly used in consumer devices.
- Charge is moved line by line, then pixel by pixel, read at a single sense node.

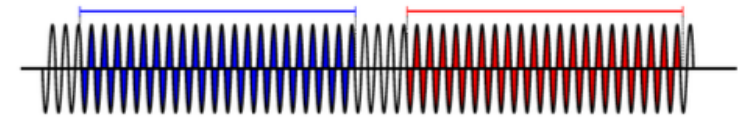




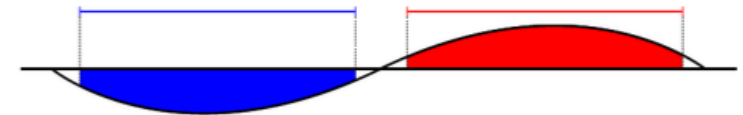
pixel charge measurement

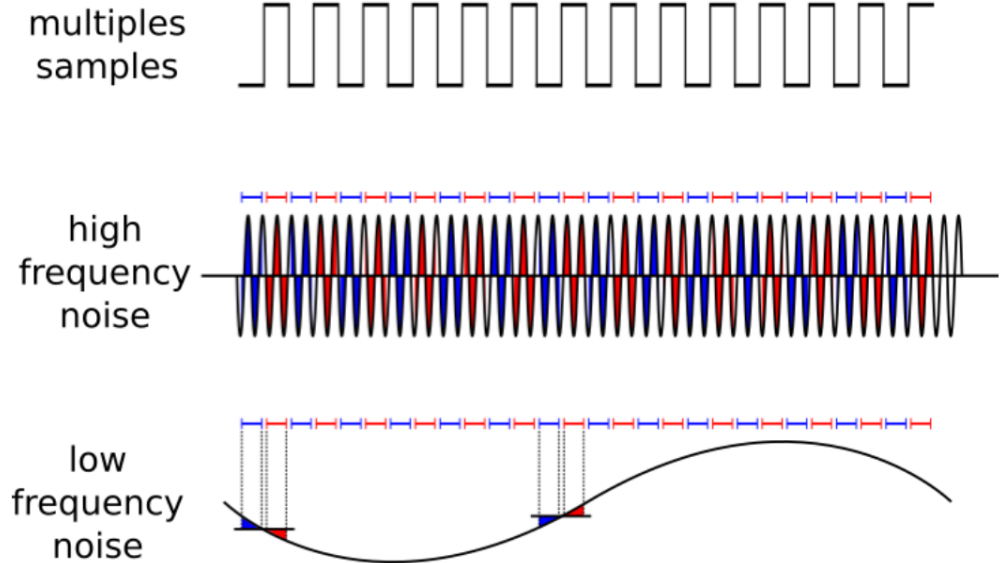
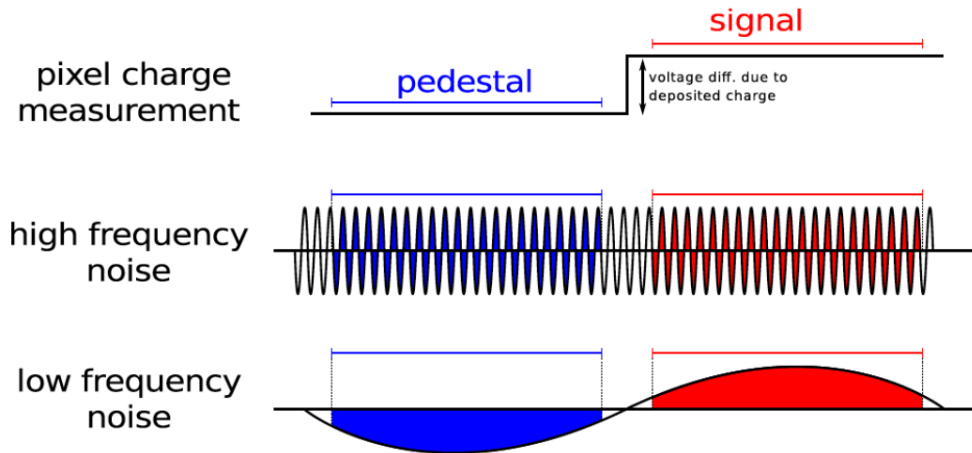


high frequency noise



low frequency noise



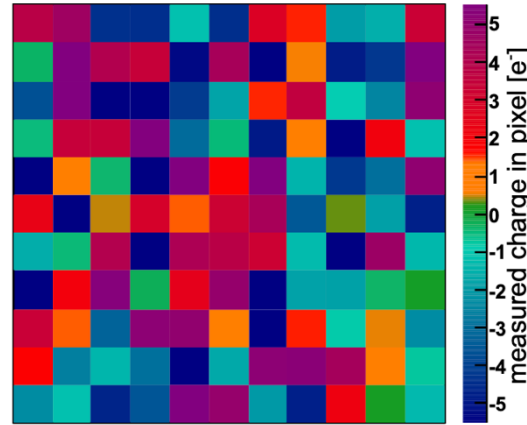


Noise goes in $1/\sqrt{N}$

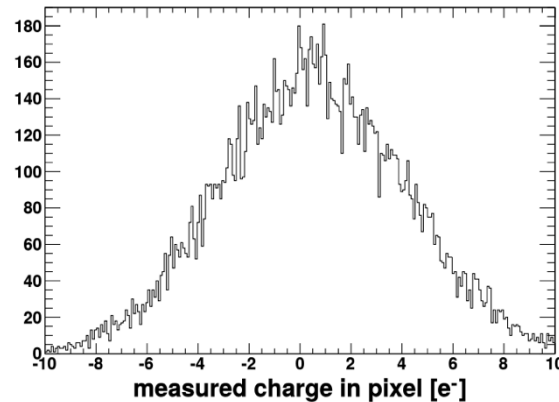
Single electron:
threshold is silicon
band gap (1.1 eV)

PRL 119, 131802 (2017)

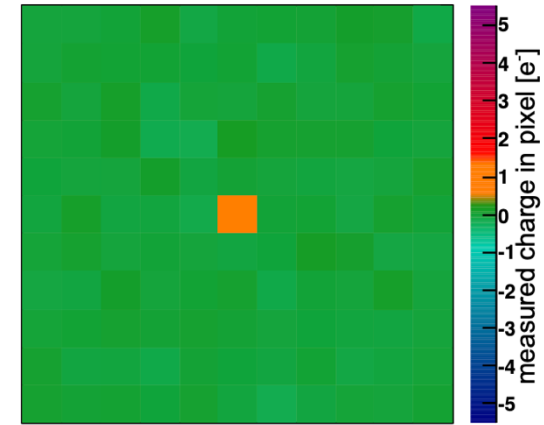
Standard CCD mode: charge in each pixel is measured once



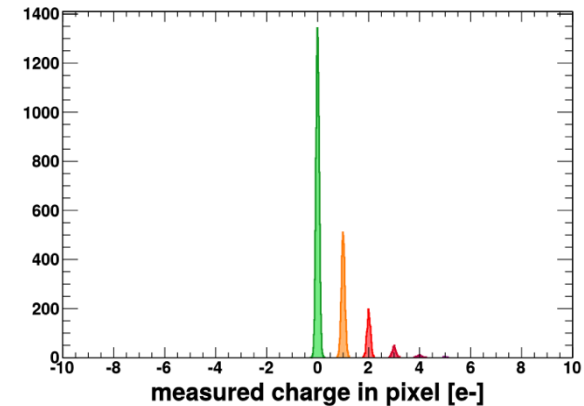
Readout-noise: 3.5 e RMS

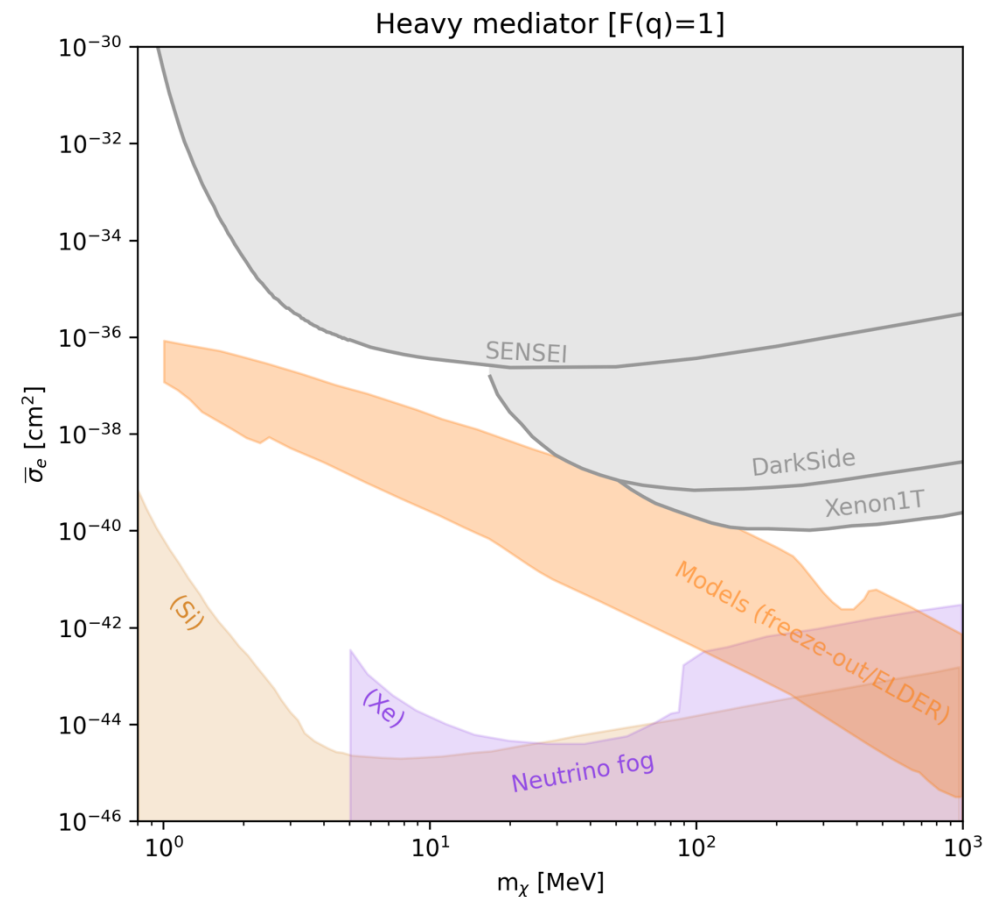
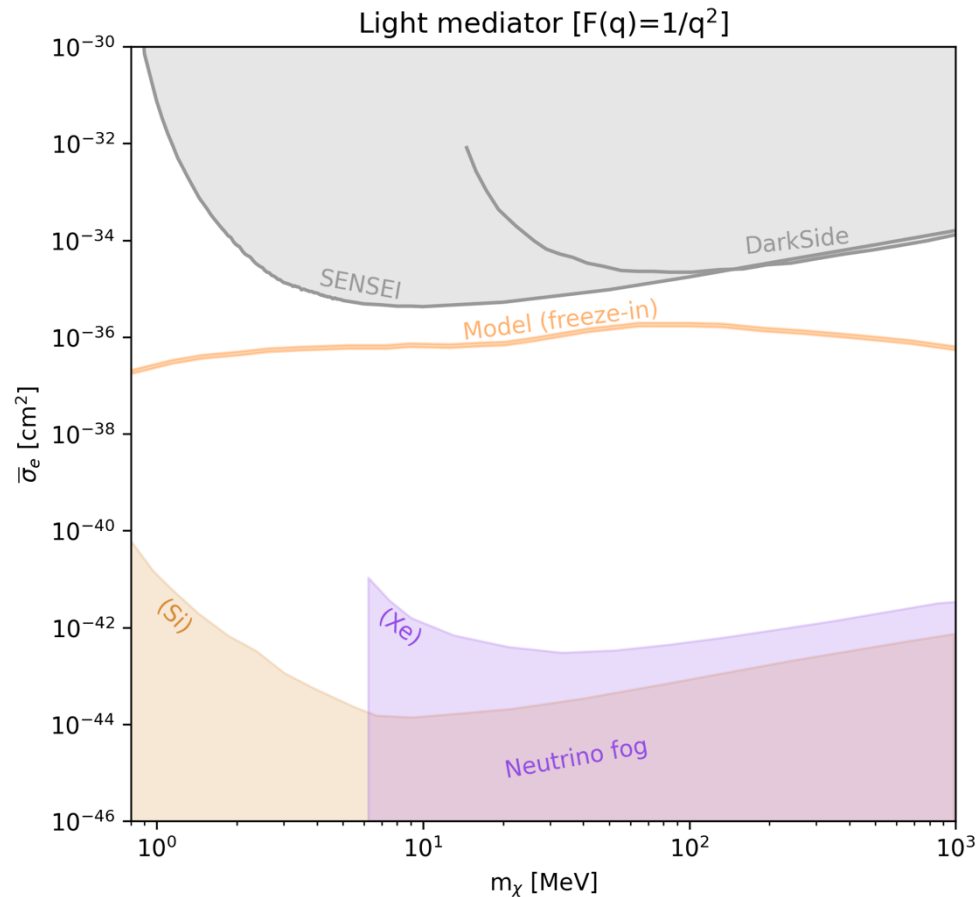


New Skipper CCD: charge in each pixel is measured multiple times

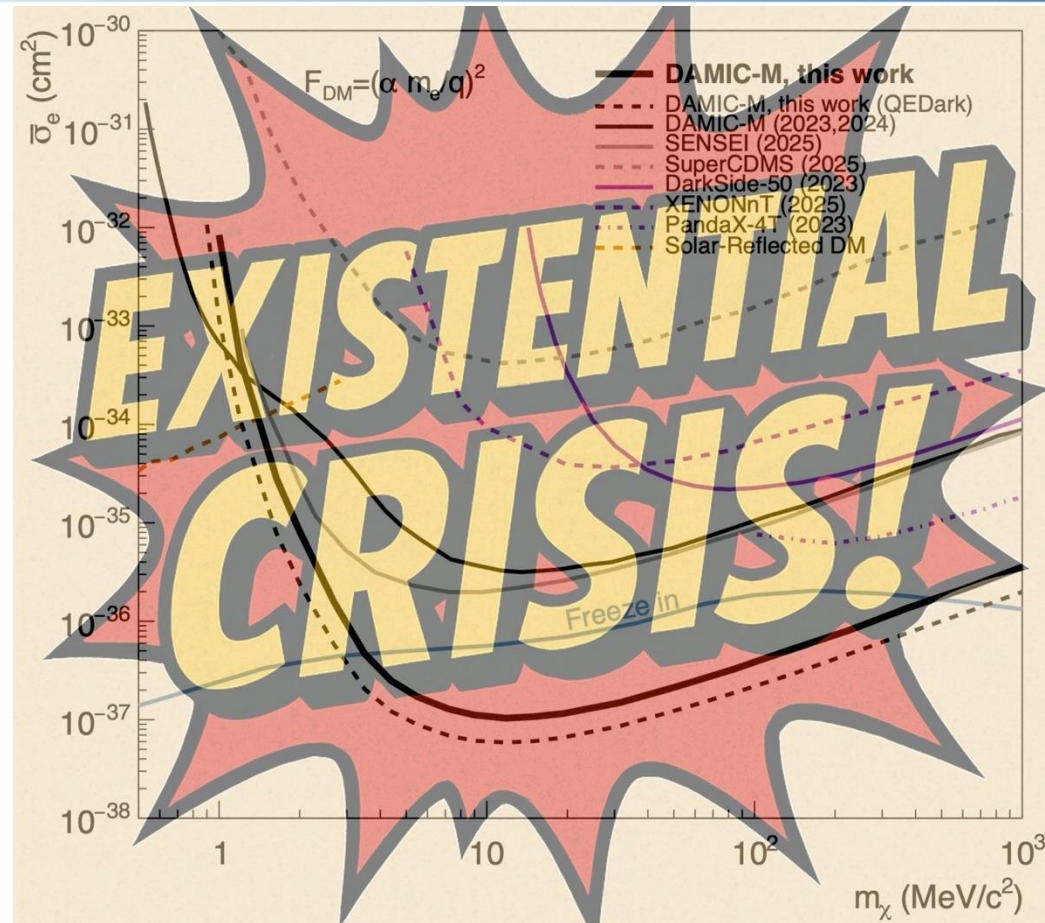


Readout-noise: 0.06 e RMS





(4) The Last 2 years



DAMIC-M
Collaboration
[2503.14617]

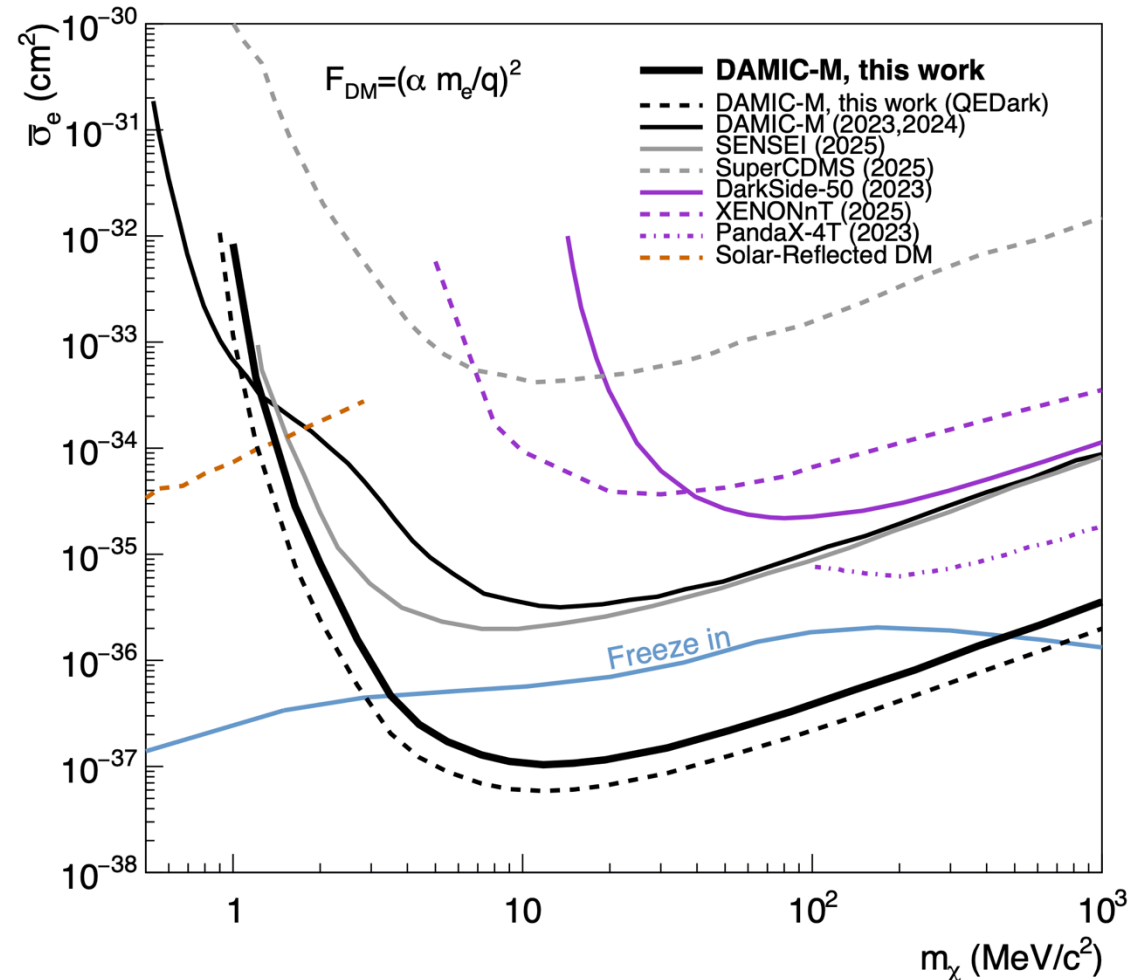
“A Moving Target: Direct Detection in the Light DM Landscape”

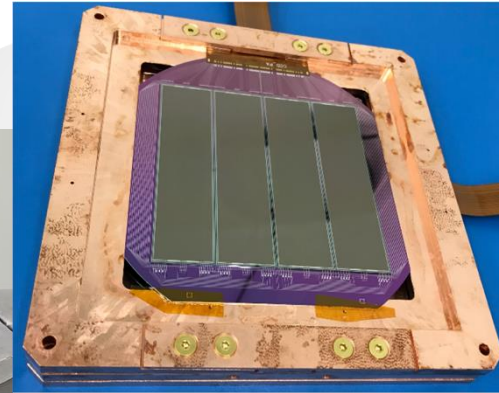
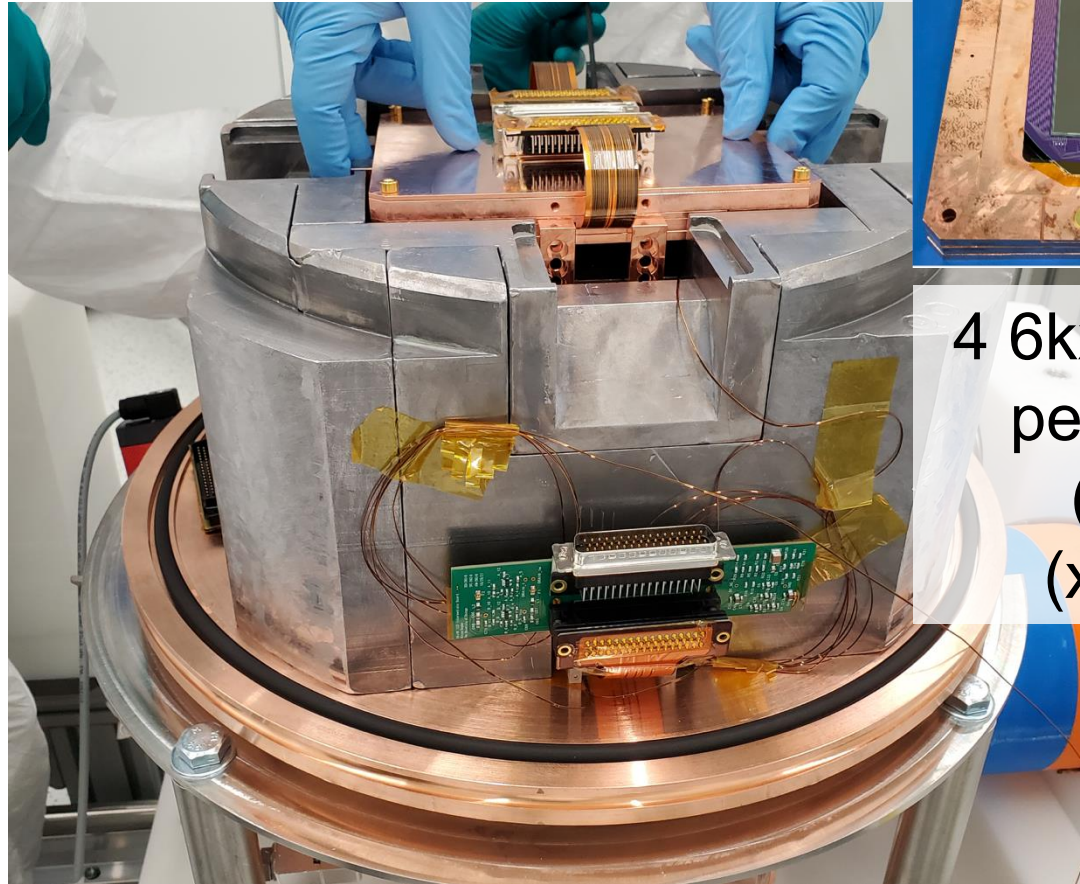
The 10th LCTP Spring Symposium : Hunting for Particle Dark Matter - <https://indico.global/event/1382/contributions/125550/>

DAMIC-M collaboration

 Phys. Rev. Lett. 135, 071002 –

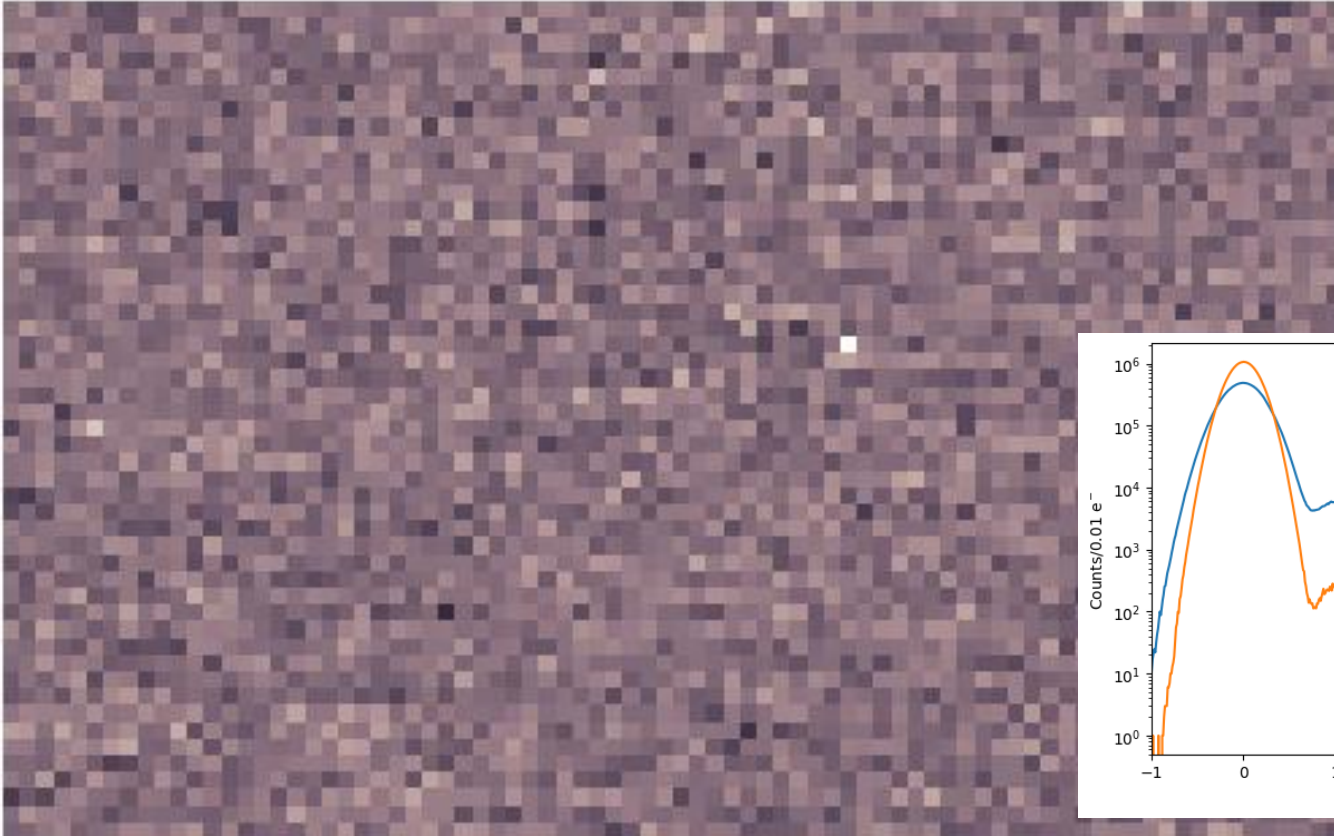
 Published 13 August, 2025





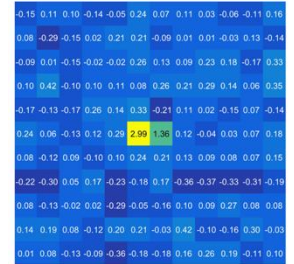
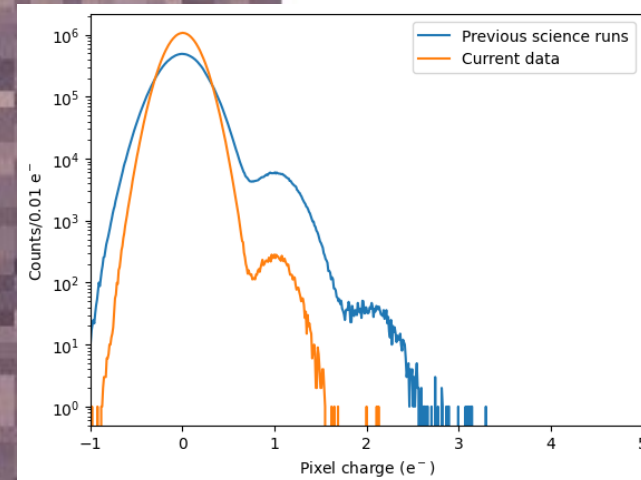
4 6kx1.5k CCD
per module
(13.4g)
(x2=27g)





	Pattern p		
	$\{11\}$	$\{21\}$	$\{111\}$
D_p	144	0	0
B_p^{rc}	141.4	0.111	0.042
B_p^{rad}	0.039	0.039	0.016

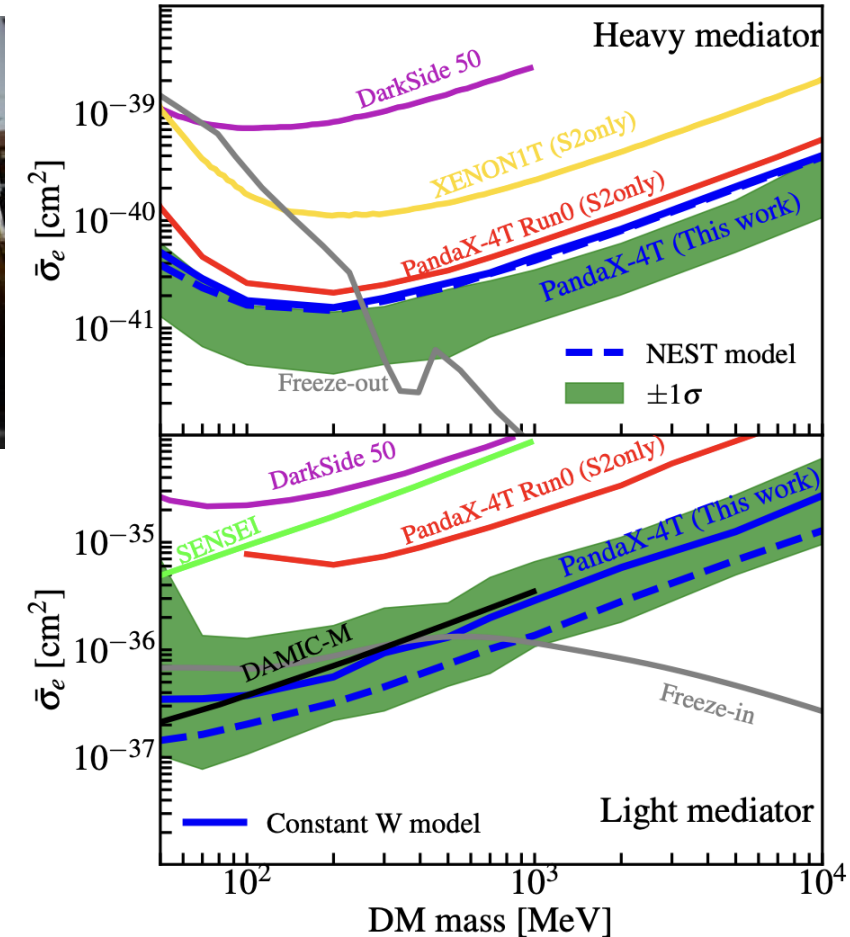
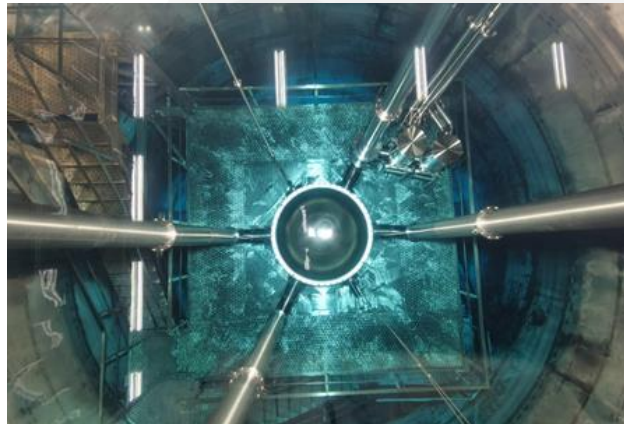
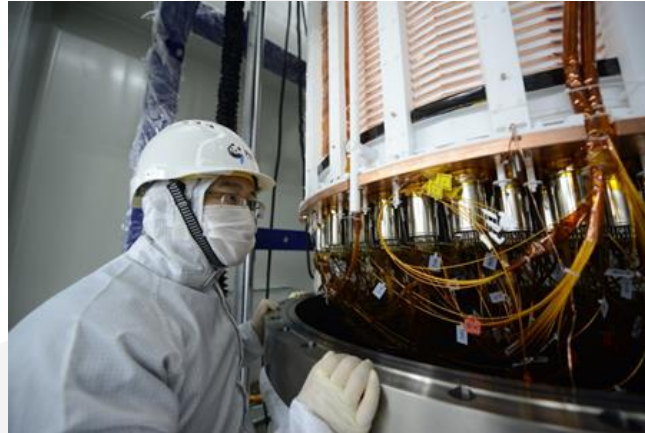
	$\{31\}$	$\{22\}$	$\{211\}$
D_p	1	0	0
B_p^{rc}	0.019	$2.5 \cdot 10^{-5}$	$5.8 \cdot 10^{-5}$
B_p^{rad}	0.052	0.011	0.035

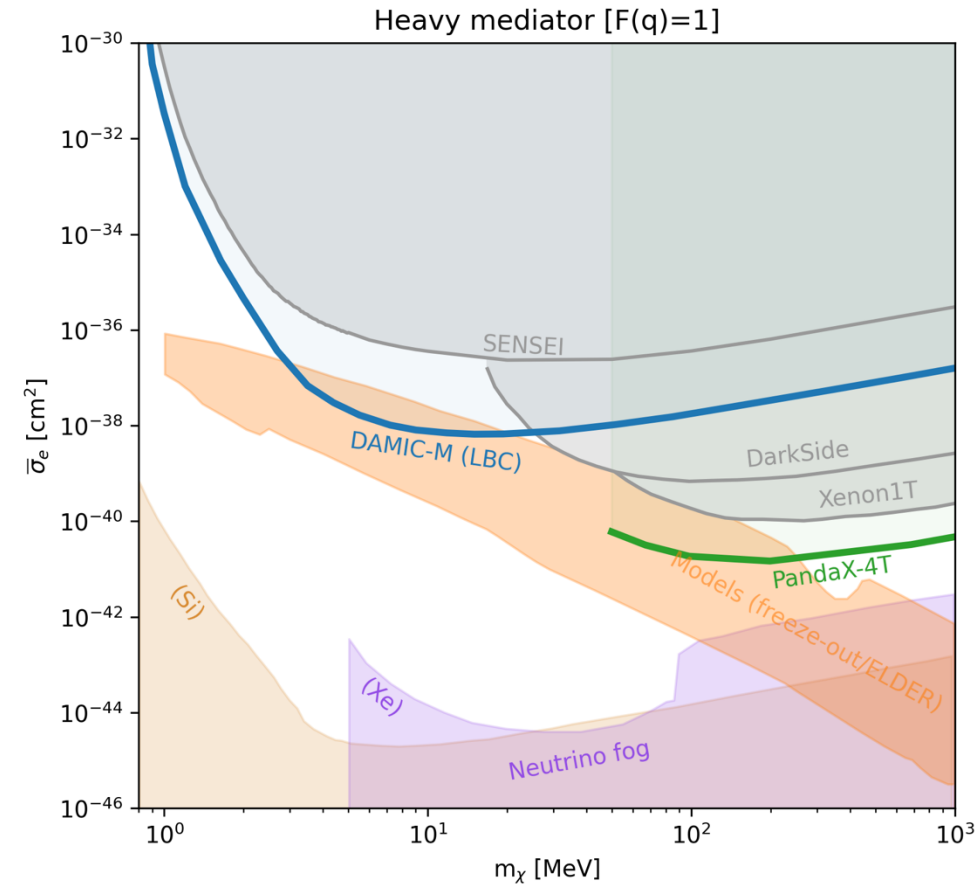
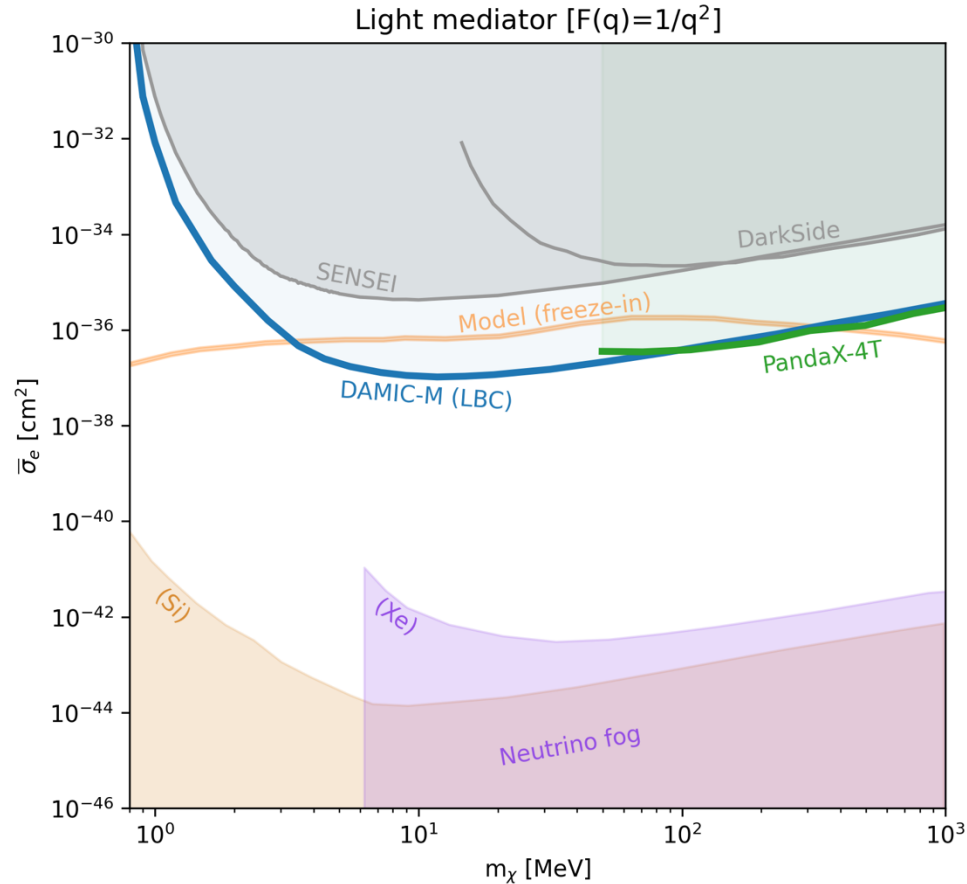


PandaX collaboration

 Phys. Rev. Lett. 135, 211001 –

 Published 18 November, 2025

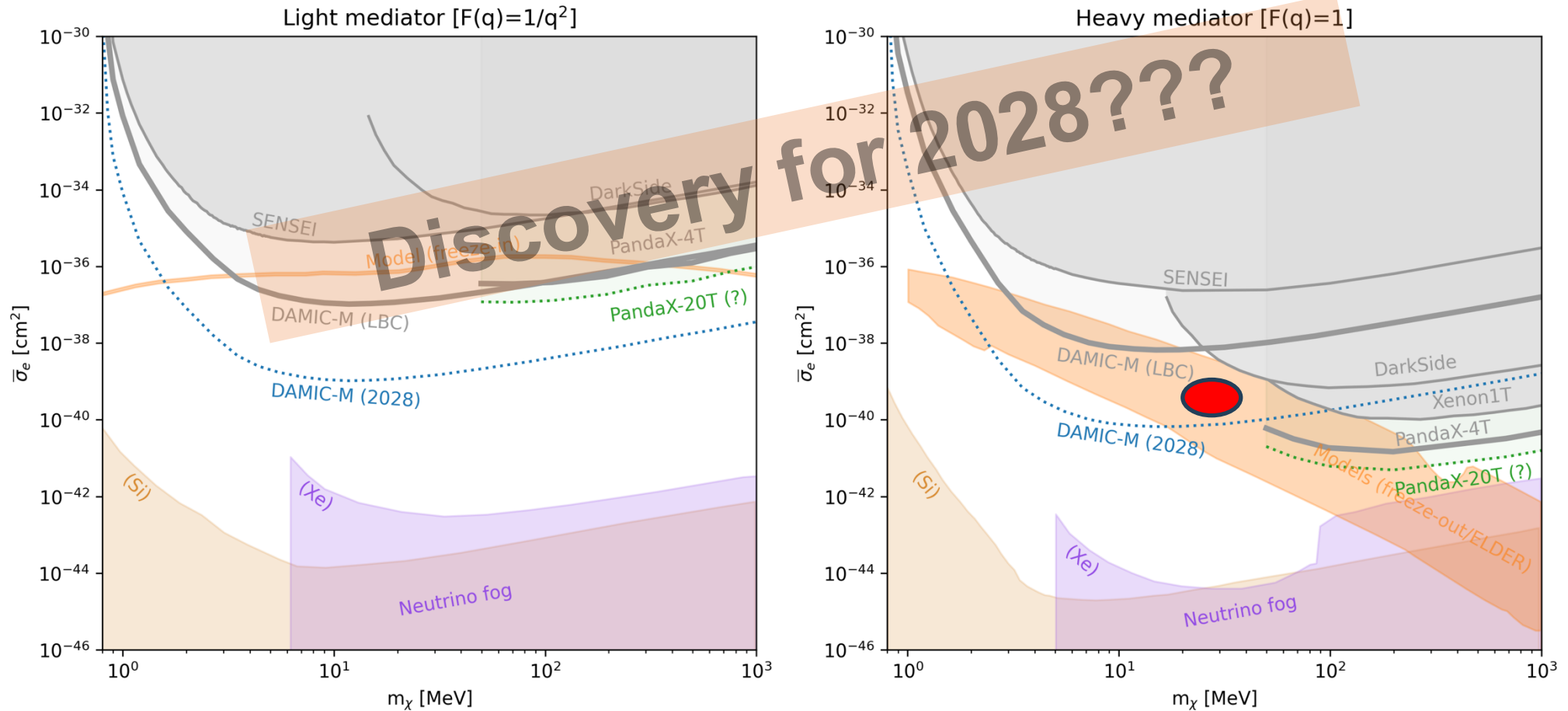




(5) The Next 2 years (?)

- **PandaX-xT:**
 - Mass increase towards 43T target (?)
 - First step at 20T (?)
 - Data taking starting in 2027 (?)
 - Sensitivity improving naively 3 to 5 (?)
- **DarkSide 20k:**
 - Huge size improvement over DarkSide-50
 - Data taking starting in 2027 (?)
 - Likely to be competitive against DAMIC-M/PandaX-20T, but maybe not by next Noirmoutier
- **Xenon/LZ/XLZD:**
 - Likely not competitive for next Noirmoutier...

- **DAMIC-M:**
 - Installation under way
 - Exposure increase: 27 g $\rightarrow$ 350 g
 - Exposure increase: 2 months $\rightarrow$ 2 years
 - Radioactive backgrounds $\downarrow$ $\sim 10\times$ (goal)
 - Dark current $\downarrow$ $\sim 2\times$ (or better?)
- **SuperCDMS:**
 - Delayed by >5 years relative to original schedule
 - Currently in commissioning
 - Original projections exceeded competing semiconductor technologies by 2–3 orders of magnitude
 - Current projected advantage over DAMIC-M is typically <1 order of magnitude
 - Ultimate sensitivity depends critically on understanding low-energy / single-charge backgrounds (LEE)



- Electron recoil experiments progressing extremely fast in the last years
 - A very significant part of the strongly motivated parameter space in the MeV-GeV region has been explored in the last year or will be explored in the next 3(?) years
- Si neutrino fog is not yet within experiment reach, Xe one isn't far
 - Directional electron recoil still unexplored
- Beside Electron Recoil, absorption processes also contribute to the search
- Dark sector can of course be more complex than a simple U(1)
 - Entering a new age for models in the MeV-GeV range
- Electron Recoil at GeV or more can also be interesting
 - Don't underestimate the electron recoil channel when designing your next Xenon experiment







Thank you