

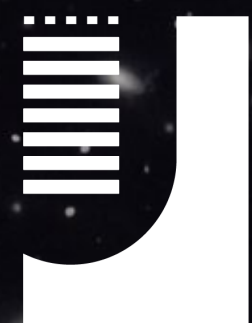
# Genetic Algorithms with Marginalised Ensembles

for model-independent reconstruction  
of cosmological quantities

In collaboration with  
Matteo Martinelli, Savvas Nesseris

**Matteo Peronaci**

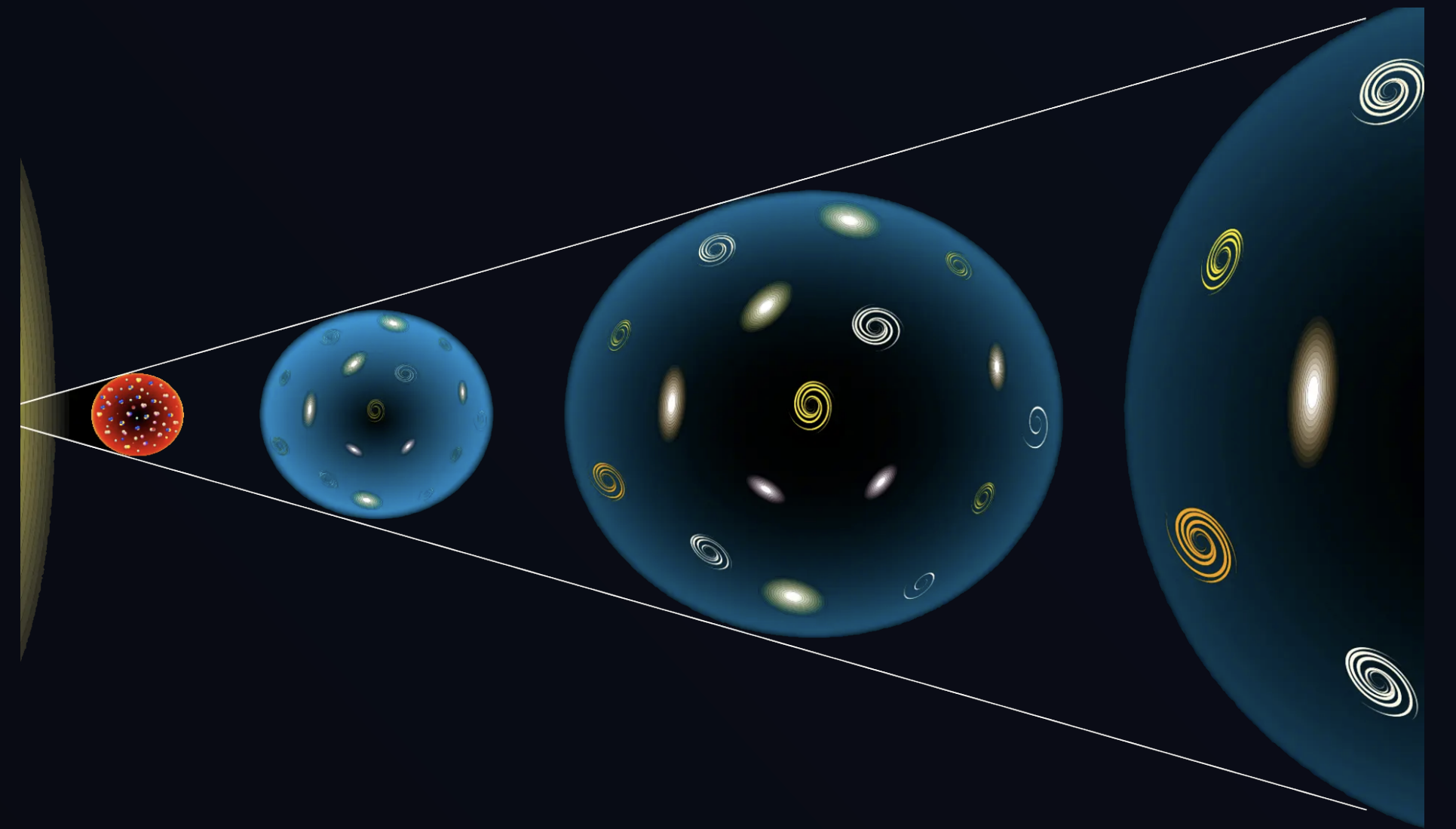
Les Rencontres de Noirmoutier, June 5, 2026



**We want to test standard cosmology**

# We want to test standard cosmology

$\Lambda$ CDM explains a lot of observations



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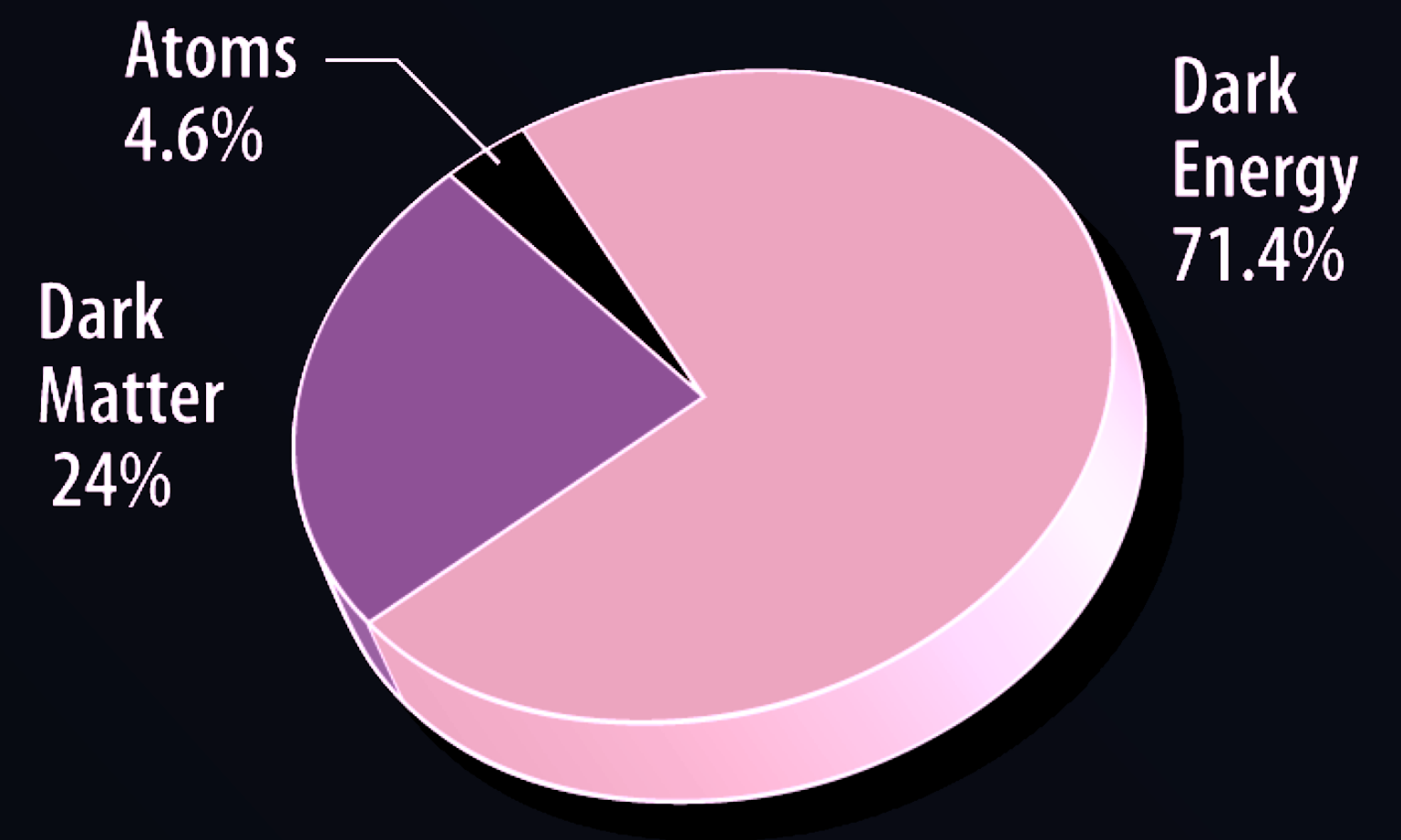
expanding universe,  
CMB, BBN,  
perturbations' theory,  
galaxy formation...

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$\Lambda$ CDM explains a lot of observations

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But there are still some open problems



# We want to test standard cosmology

$\Lambda$ CDM explains a lot of observations

expanding universe,  
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galaxy formation...

But there are still some open problems

Atoms  
4.6%

$H_0$  tension

Dark  
Energy  
71.4%

Dark  
Matter  
24%

$S_8$  tension

DE interpretation

# Dark Energy equation of state as a $\Lambda$ CDM test

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In  $\Lambda$ CDM, dark energy is a cosmological constant  $\Lambda$

$$\rho_{\Lambda} = \frac{\Lambda}{8\pi G} \quad + \quad \rho_{\text{de}} = \rho_{\text{de},0} a^{-3(1+w_{\text{de}})} \quad + \quad P_{\text{de}} = w_{\text{de}} \rho_{\text{de}}$$
$$\implies w_{\text{de}} = w_{\Lambda} \equiv -1$$

# Dark Energy equation of state as a $\Lambda$ CDM test

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$$\implies w_{\text{de}} = w_{\Lambda} \equiv -1$$

If  $w_{\text{de}} \neq -1$

$\implies$  We have a problem with  $\Lambda$ CDM  
**We need a new model!**

# Observables for cosmological evolution

We can express  $w_{\text{de}}(z)$  as a function of observables quantities

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$$\Rightarrow w_{\text{bg}}(z) = -1 + \frac{1}{3}(1+z) \frac{d}{dz} \ln \left( \frac{H(z)^2}{H_0^2} - \Omega_{\text{m},0}(1+z)^3 \right)$$

# Observables for cosmological evolution

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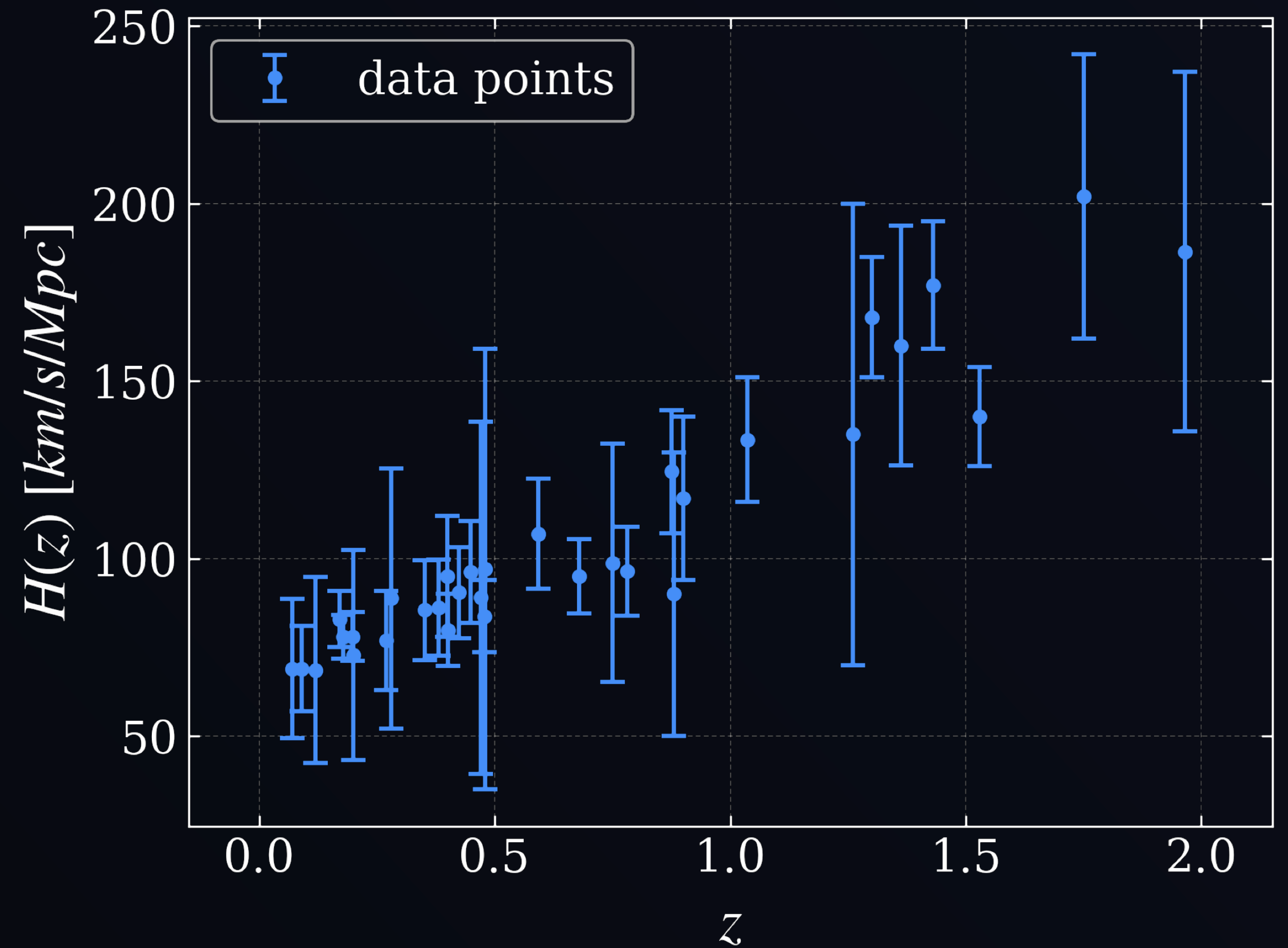
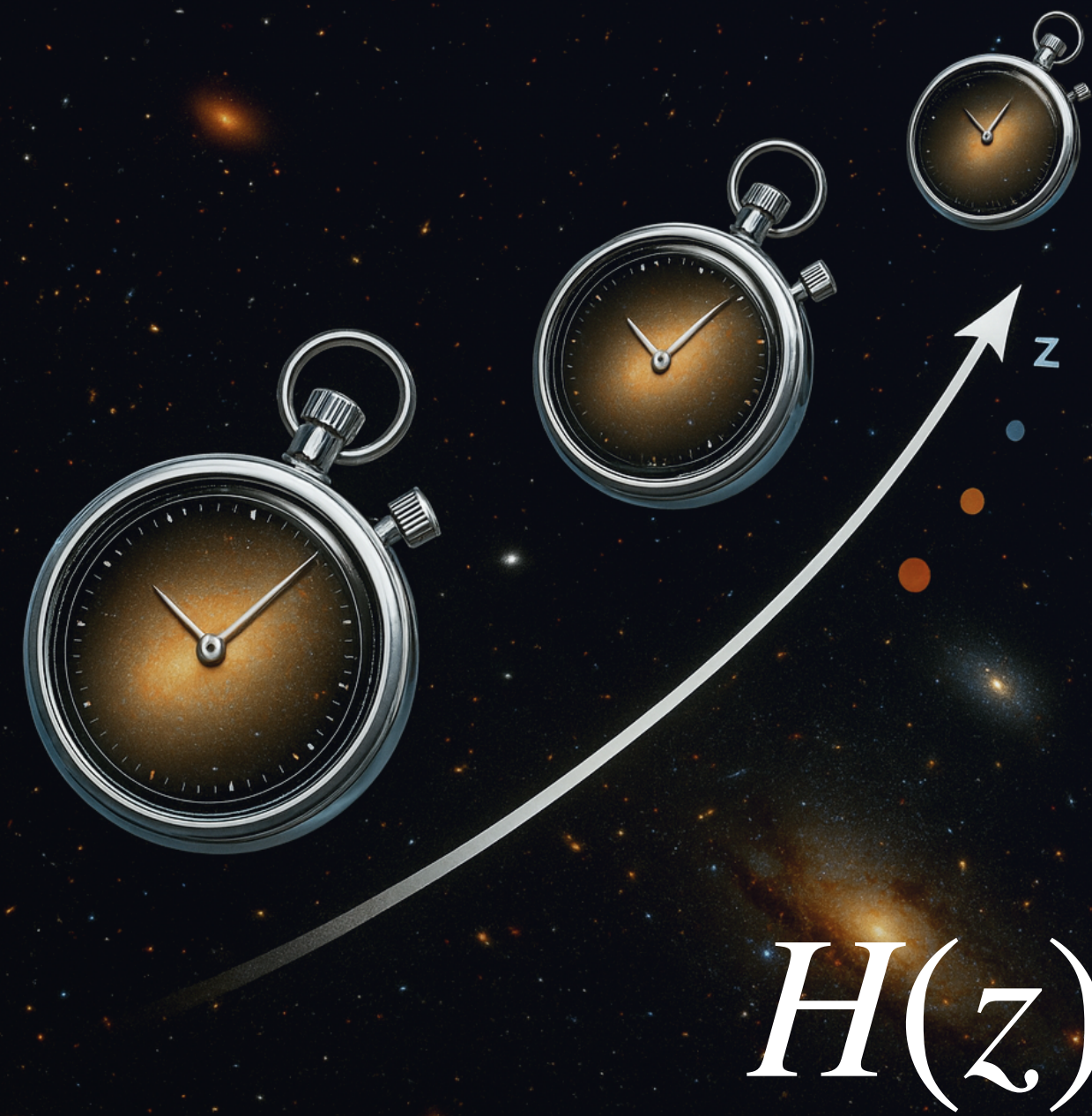


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We want to reconstruct  $H(z)$  from observational data only!

# Model-independent approach

Cosmic Chronometers



We want to reconstruct  $H(z)$  from observational data only!

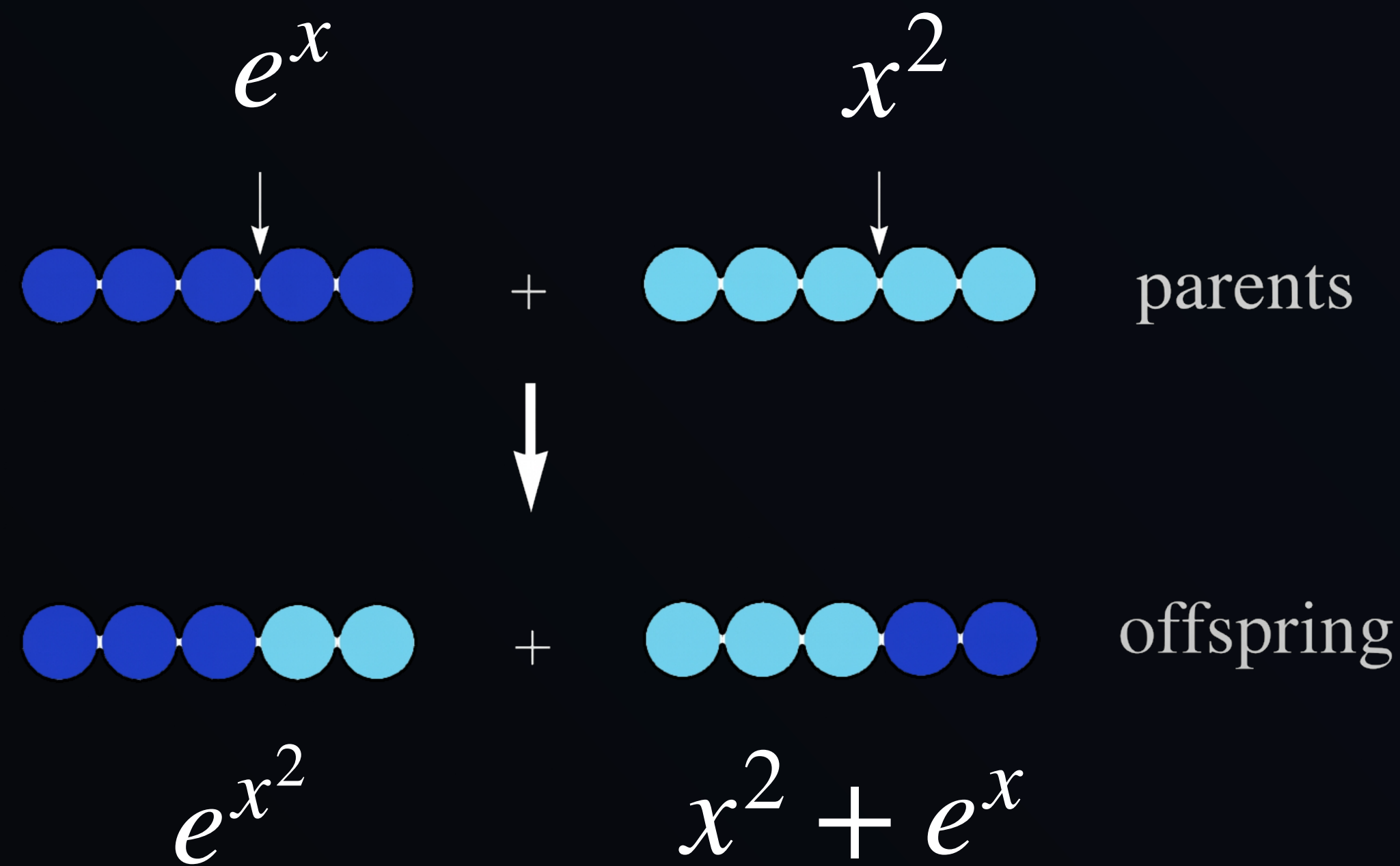
# **Model-independent approach**

Genetic Algorithms

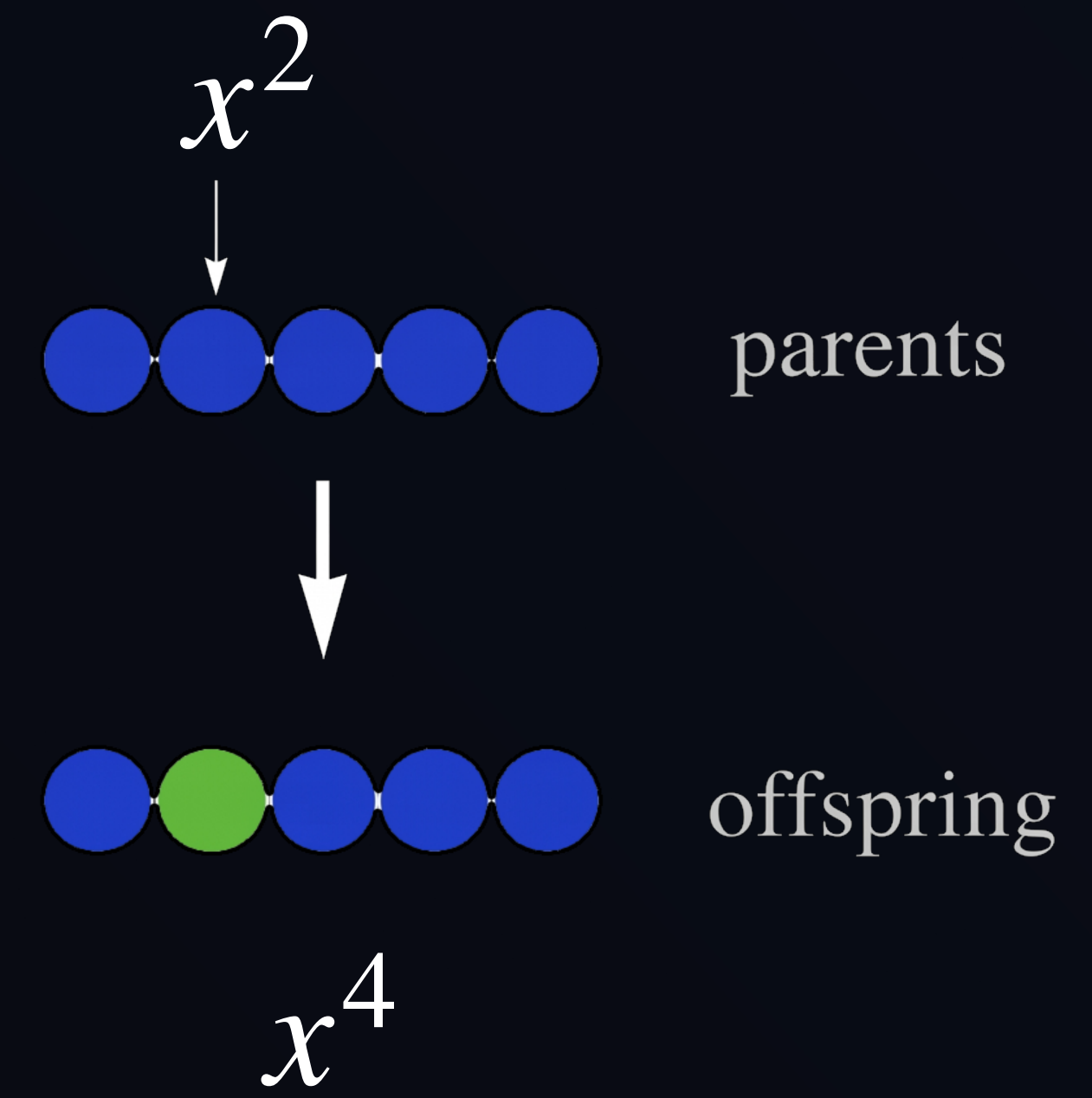
# Model-independent approach

## Genetic Algorithms

**crossover**



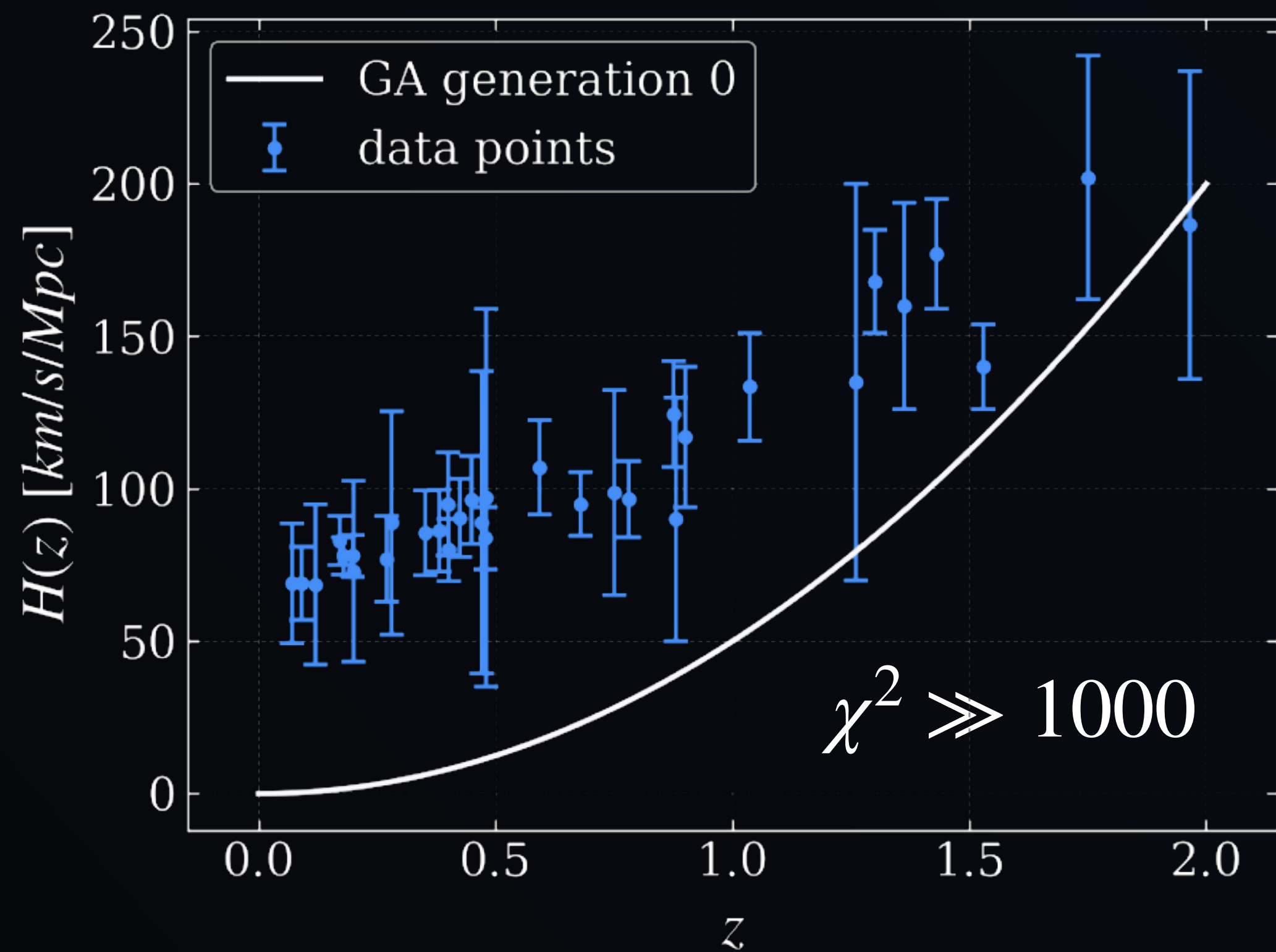
**mutation**



# Genetic Algorithms

random initial function

grammar, random seed,  $N_{\text{pop}}$

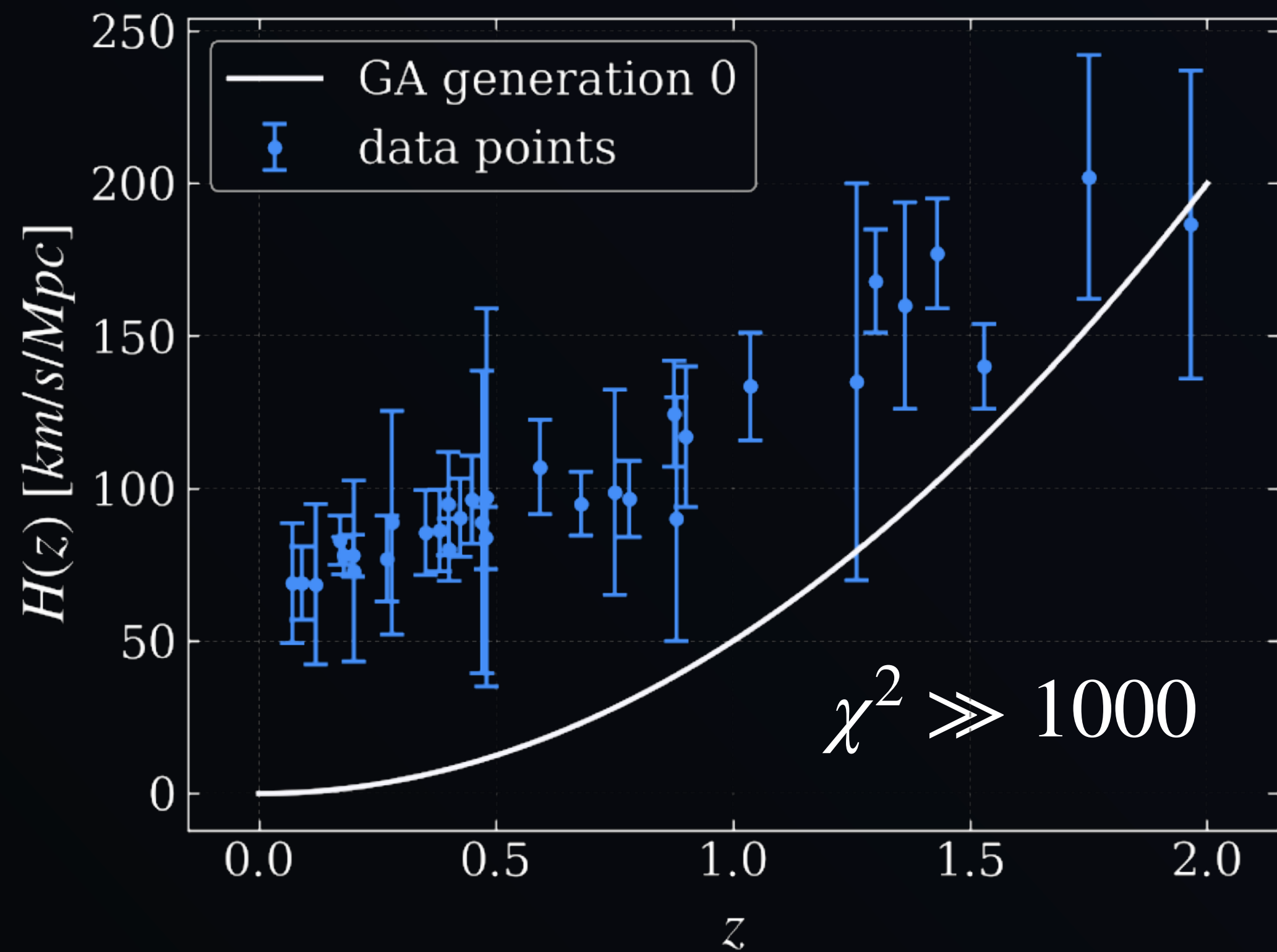


# Genetic Algorithms

random initial function  
grammar, random seed,  $N_{\text{pop}}$



improve fitness  $\chi^2$   
through GA



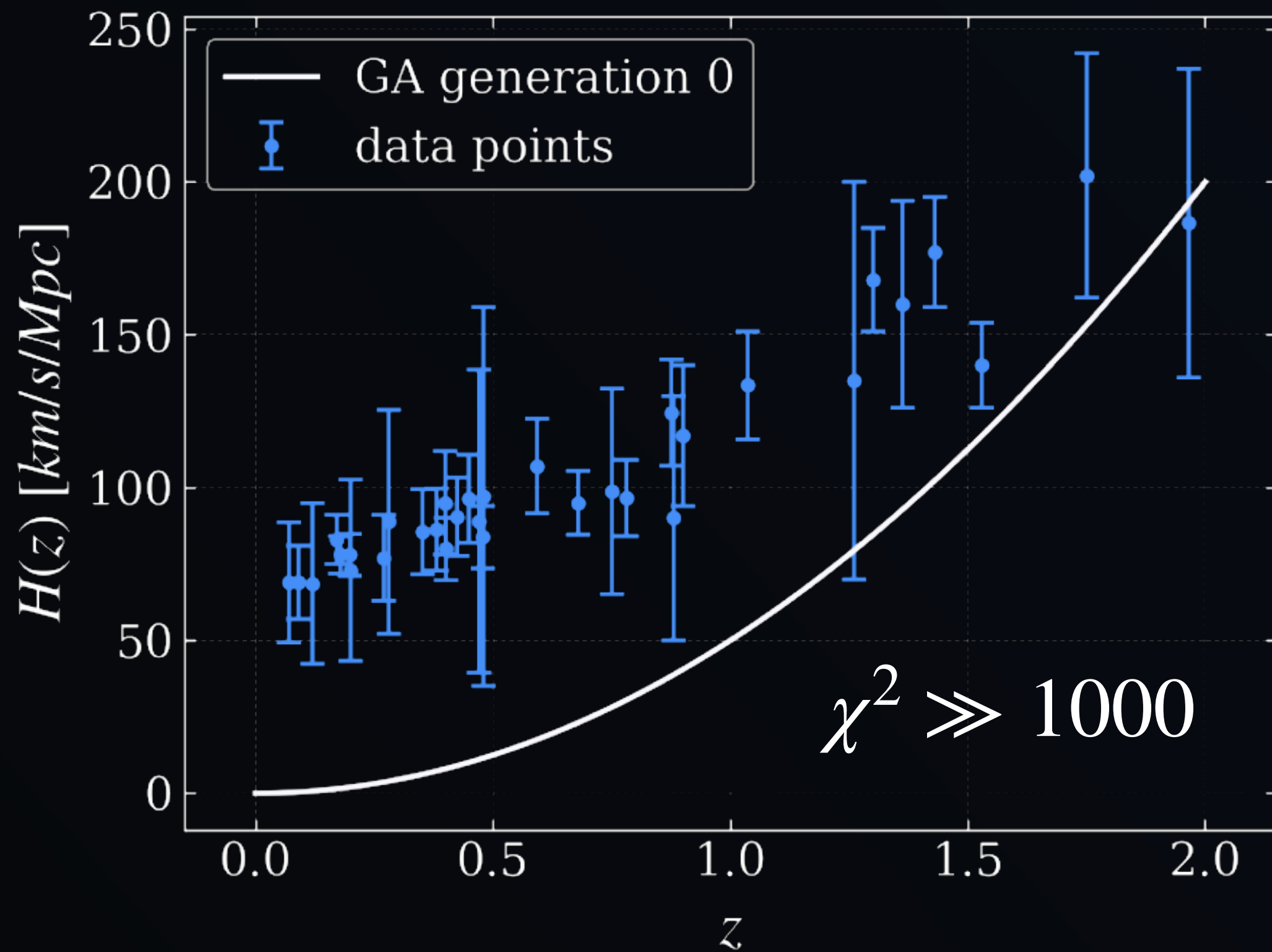
selection  
crossover  
mutation

# Genetic Algorithms

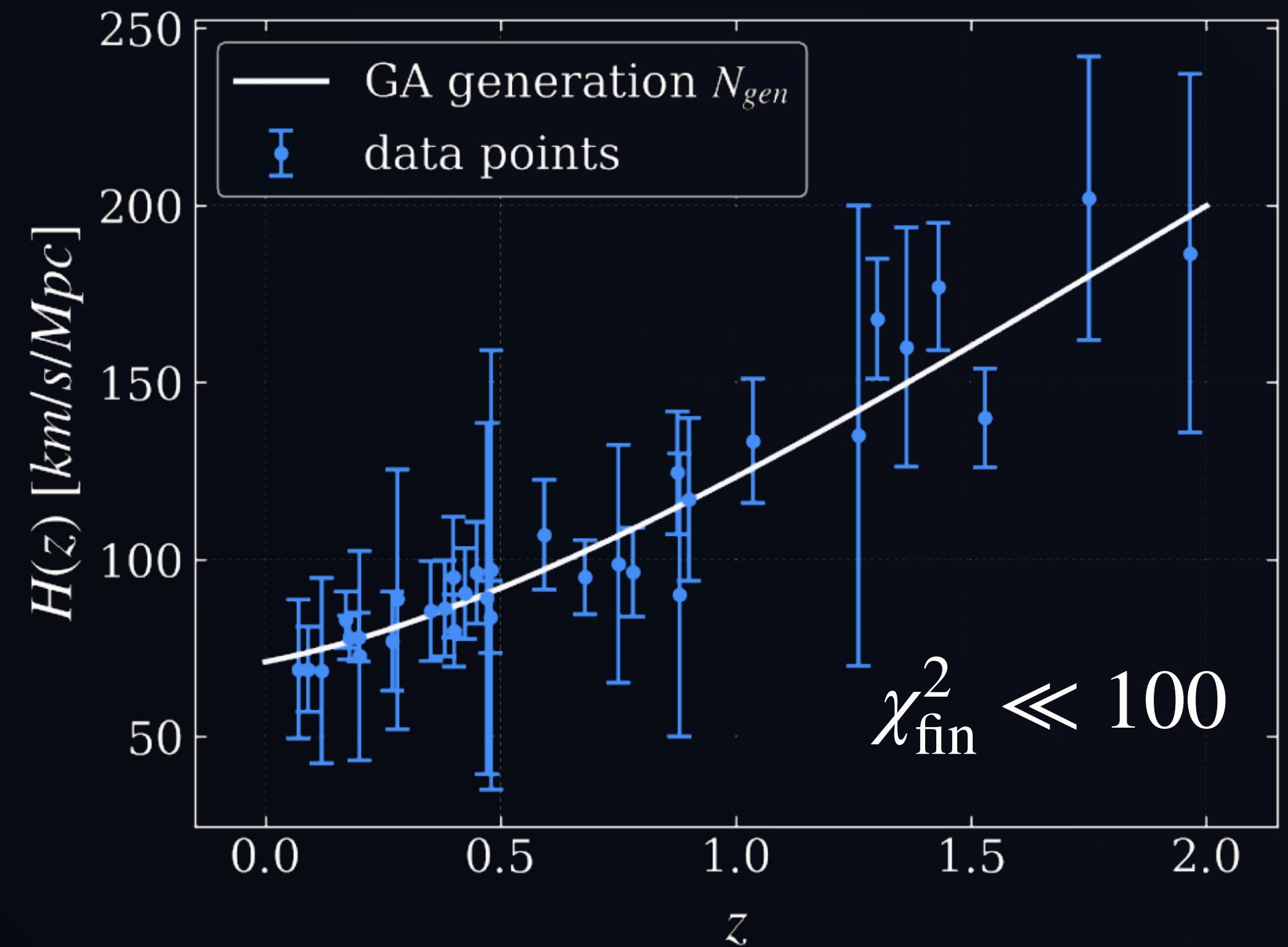
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final processed function  
number of generations



selection  
crossover  
mutation



# Genetic Algorithms

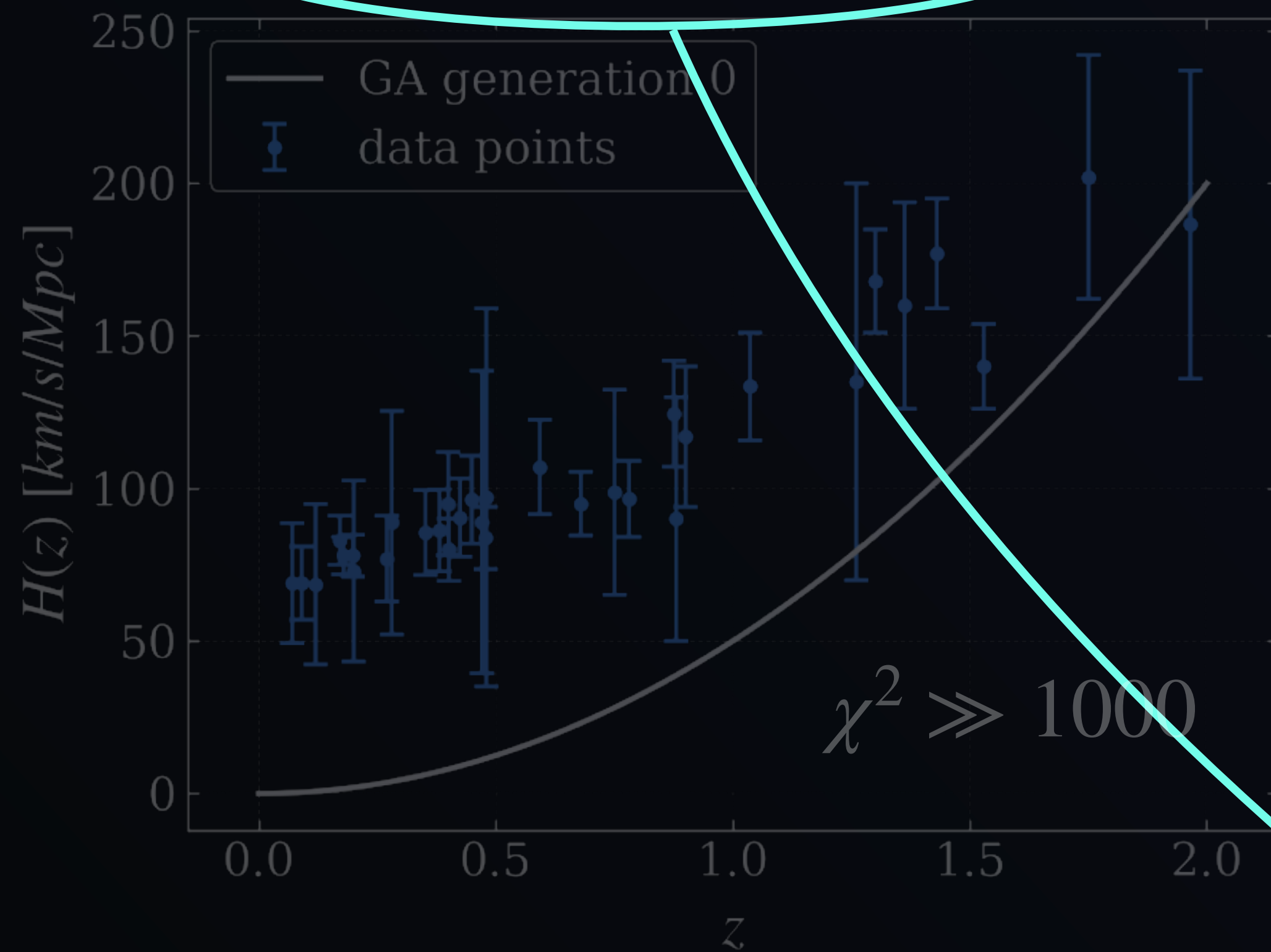
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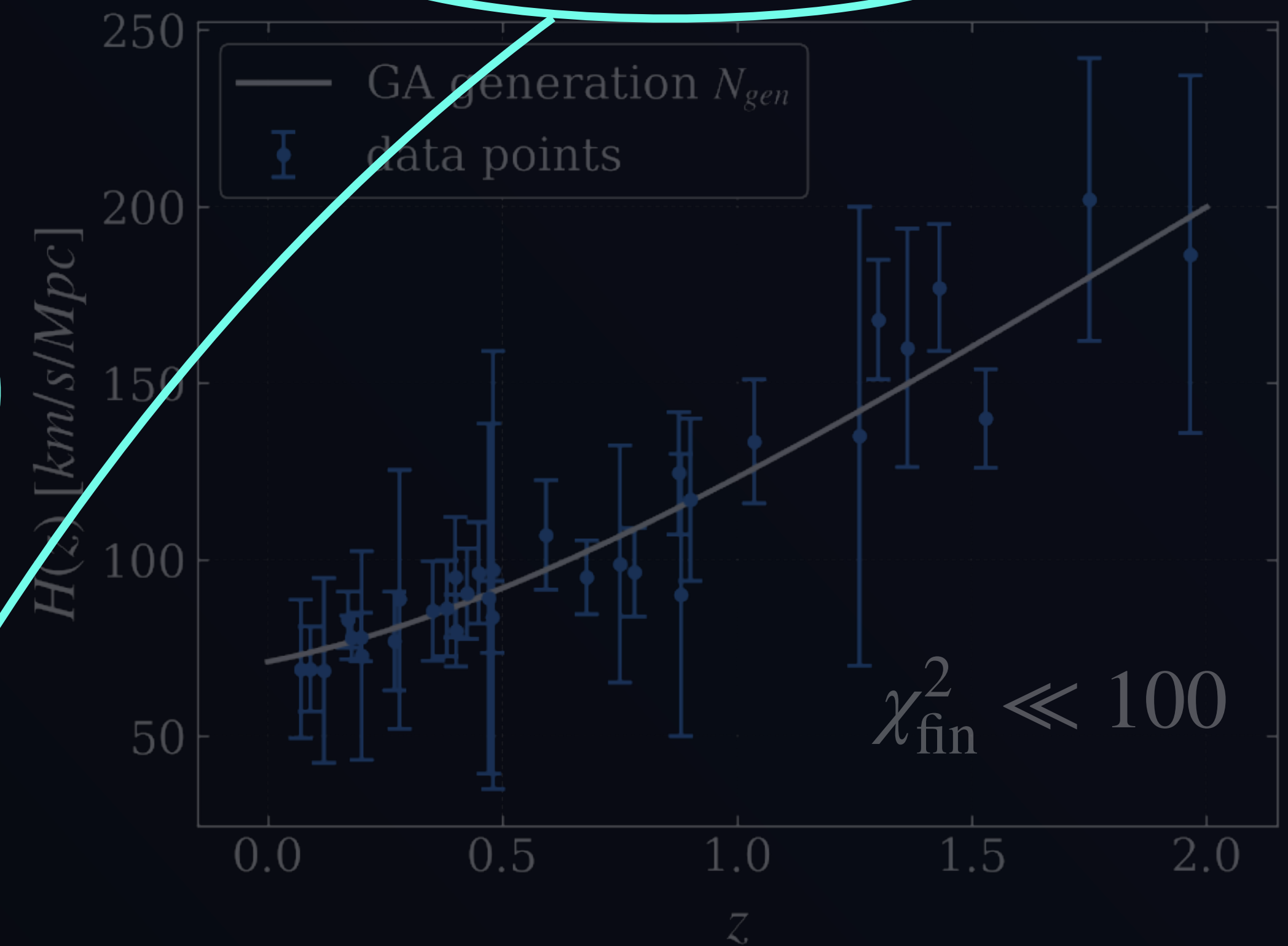
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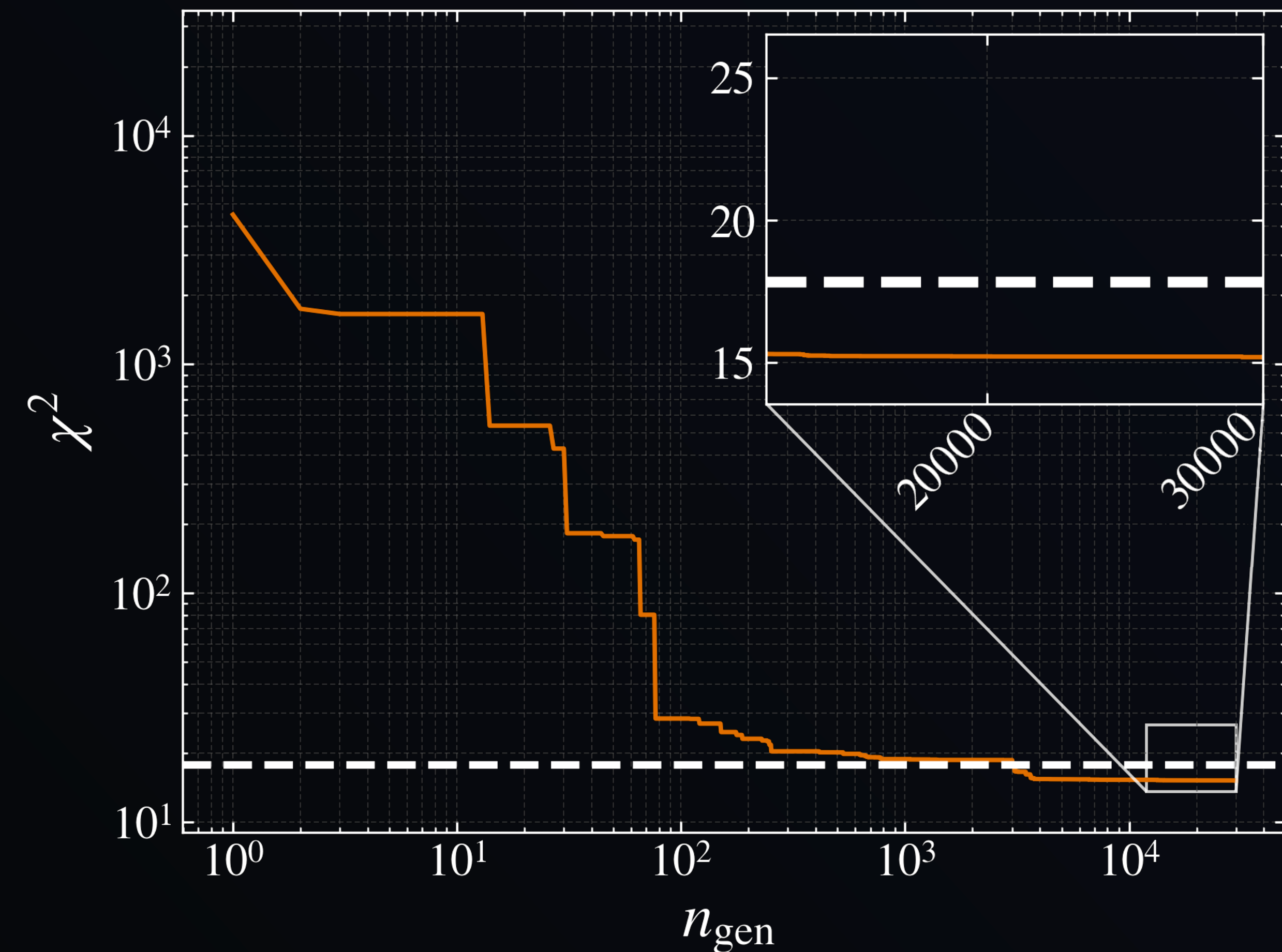


selection  
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hyperparameters

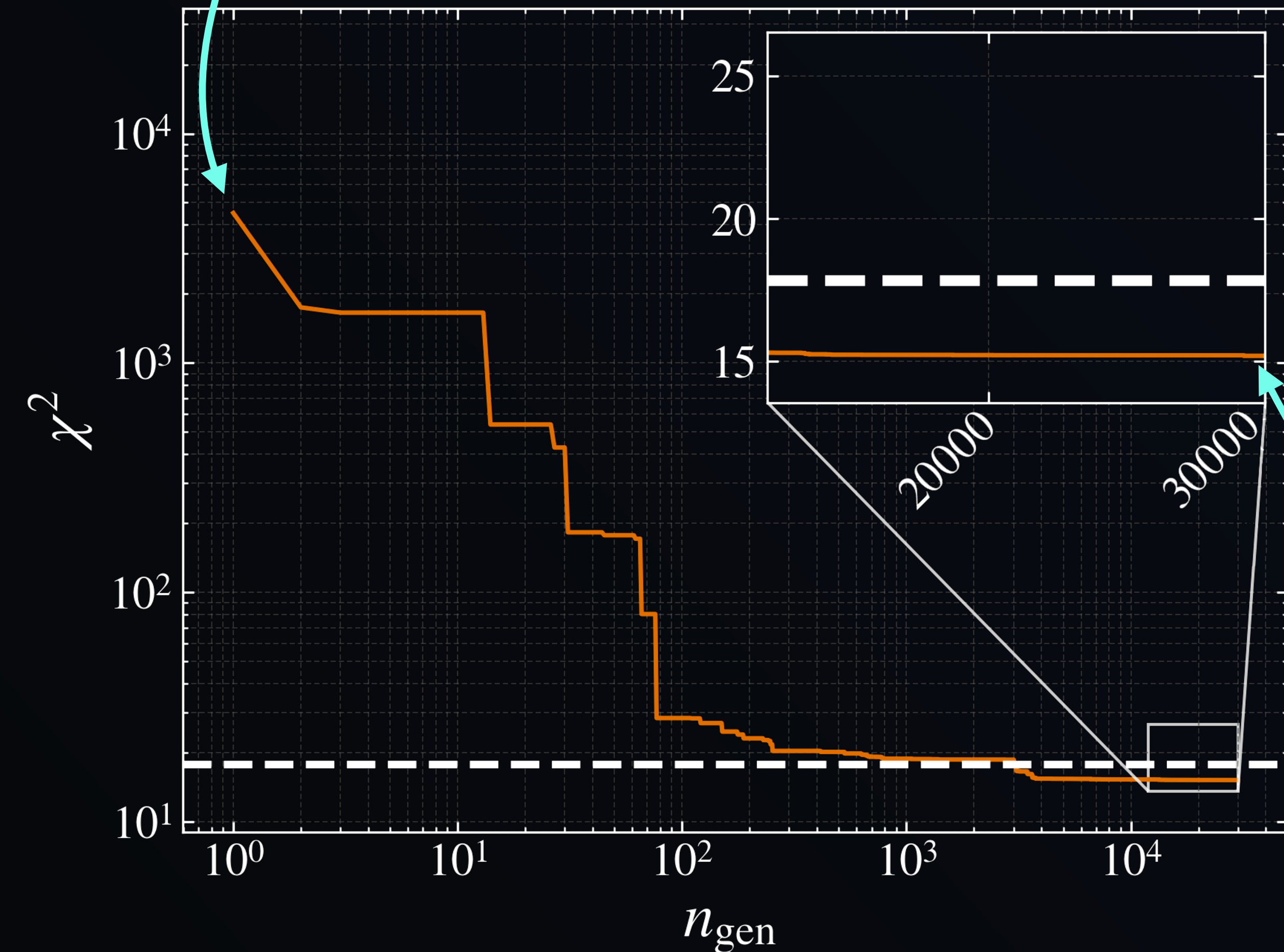
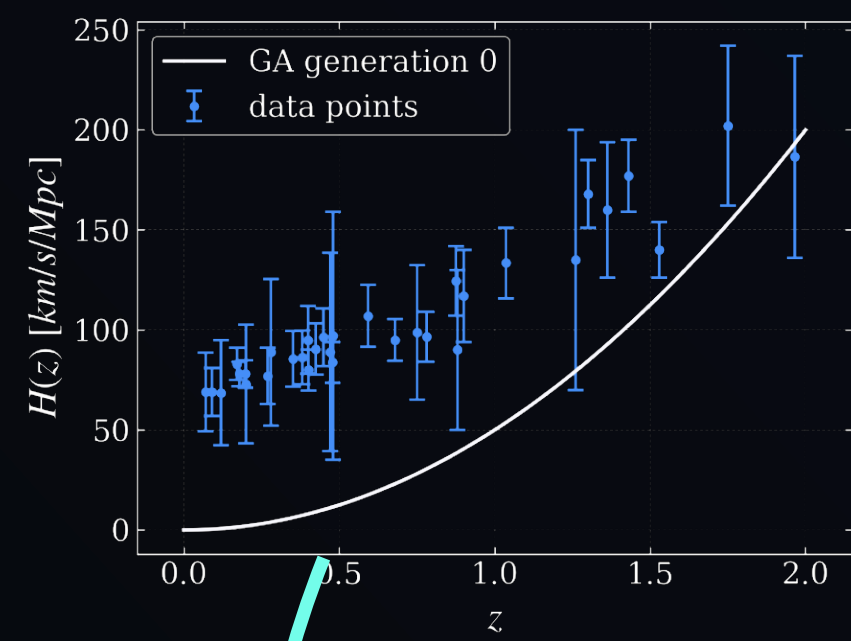
# GA hyperparameters configurations



---  $\chi^2_{thr} = 17.834$   
—  $\chi^2_{min} = 15.192$   
seed=76822, cross=0.85, mut=0.85, gramm=['poly', 'exp']

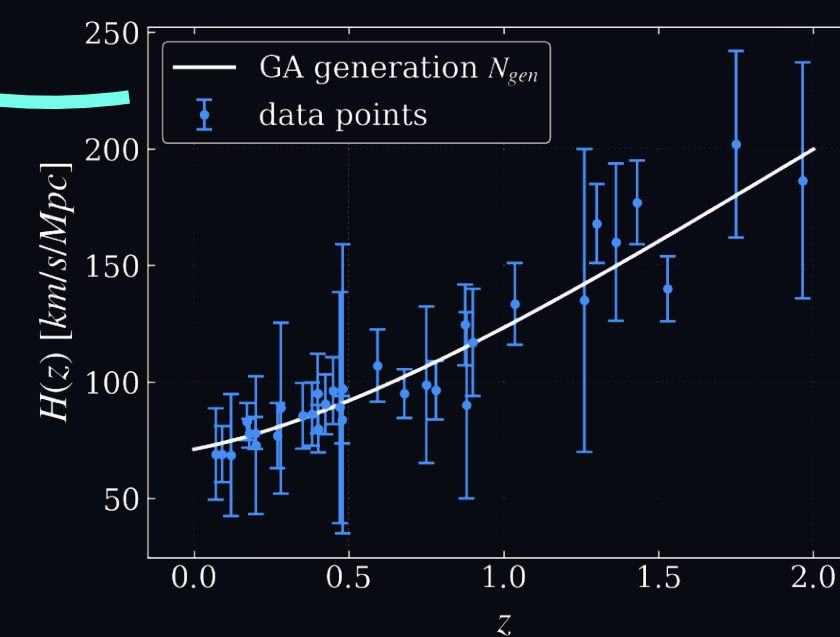
1 hyperparameters configuration...

# GA hyperparameters configurations

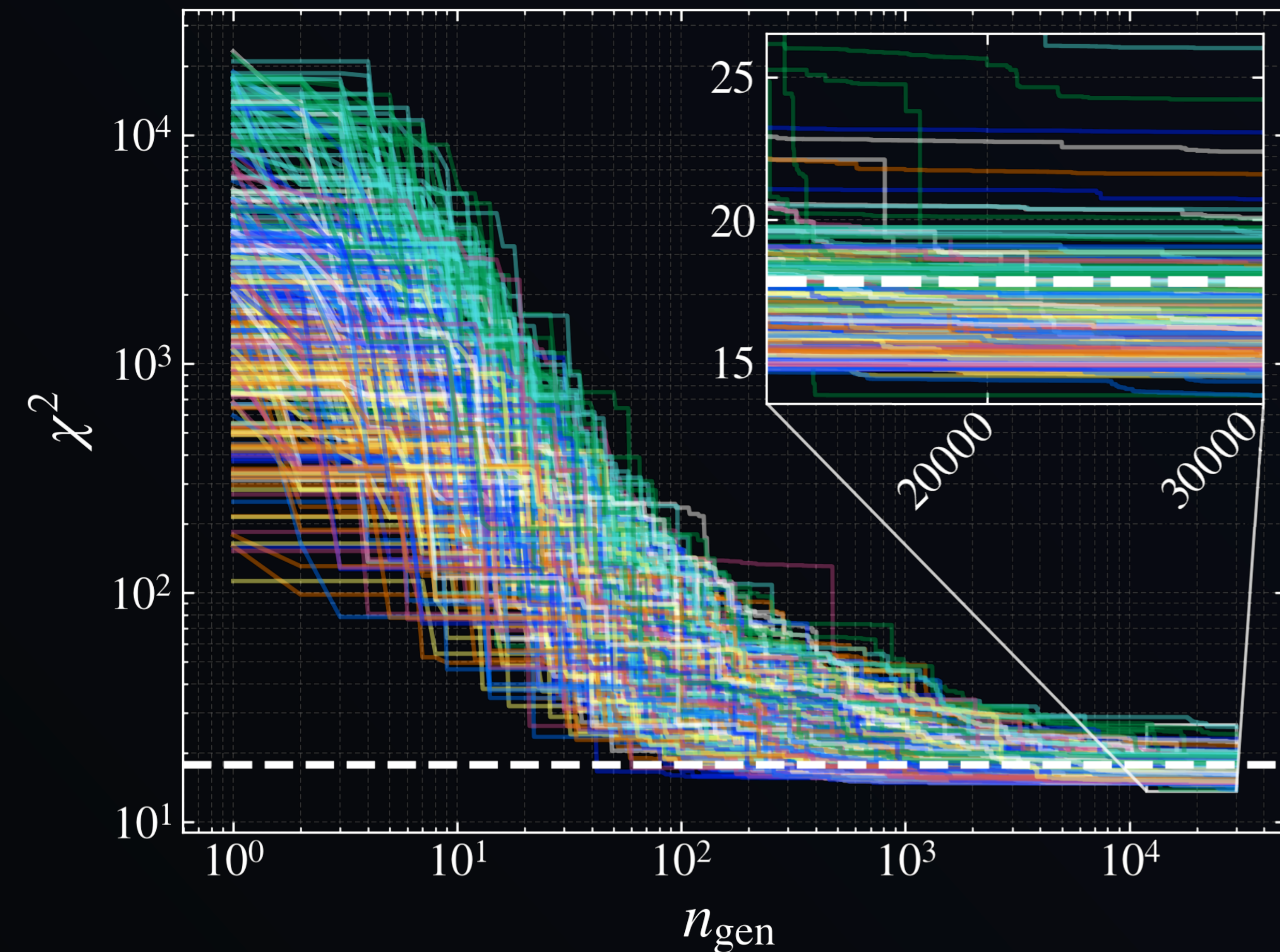


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---  $\chi^2_{thr} = 17.834$

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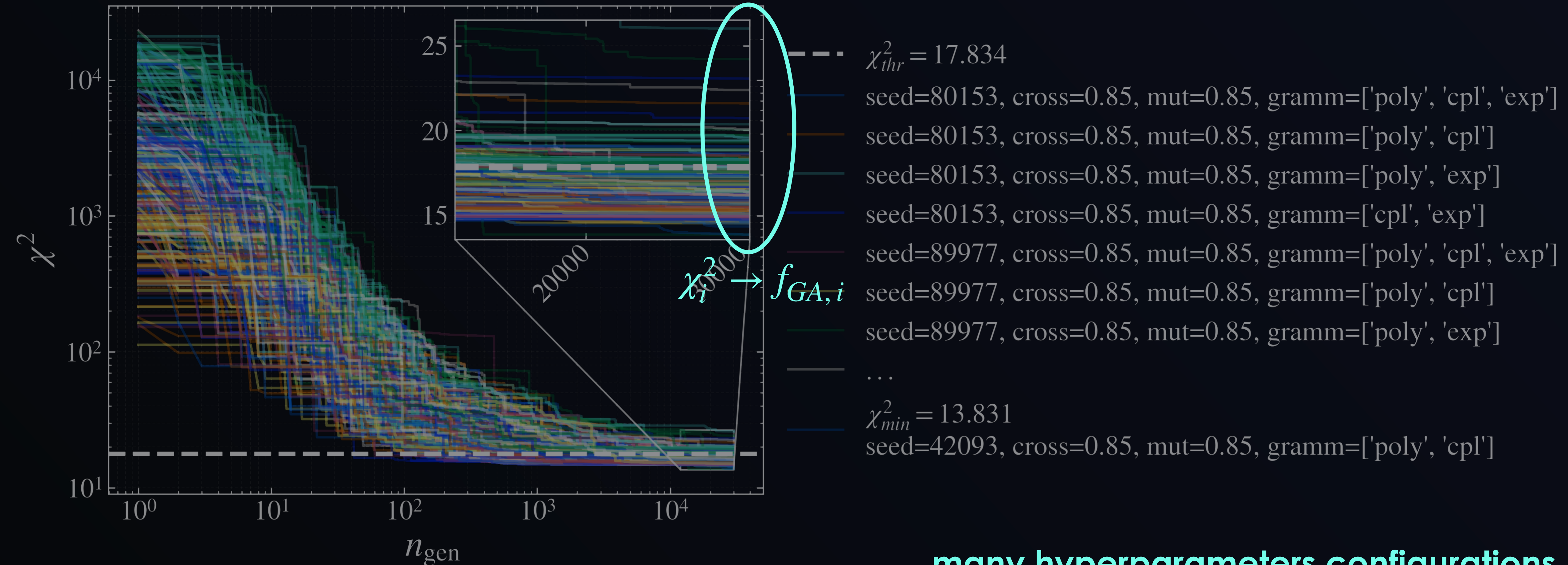
...

—  $\chi^2_{min} = 13.831$

— seed=42093, cross=0.85, mut=0.85, gramm=['poly', 'cpl']

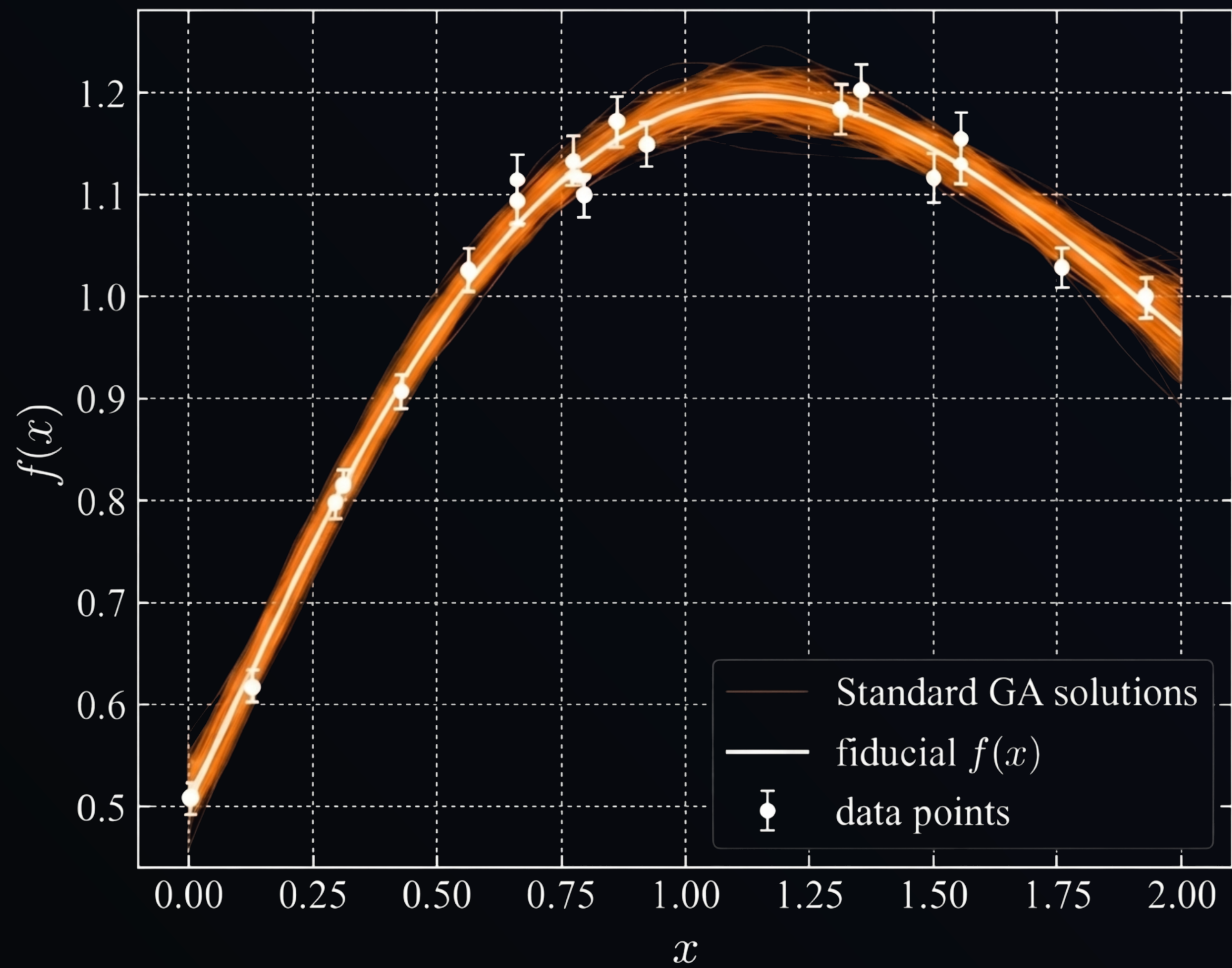
...many hyperparameters configurations

# GA hyperparameters configurations

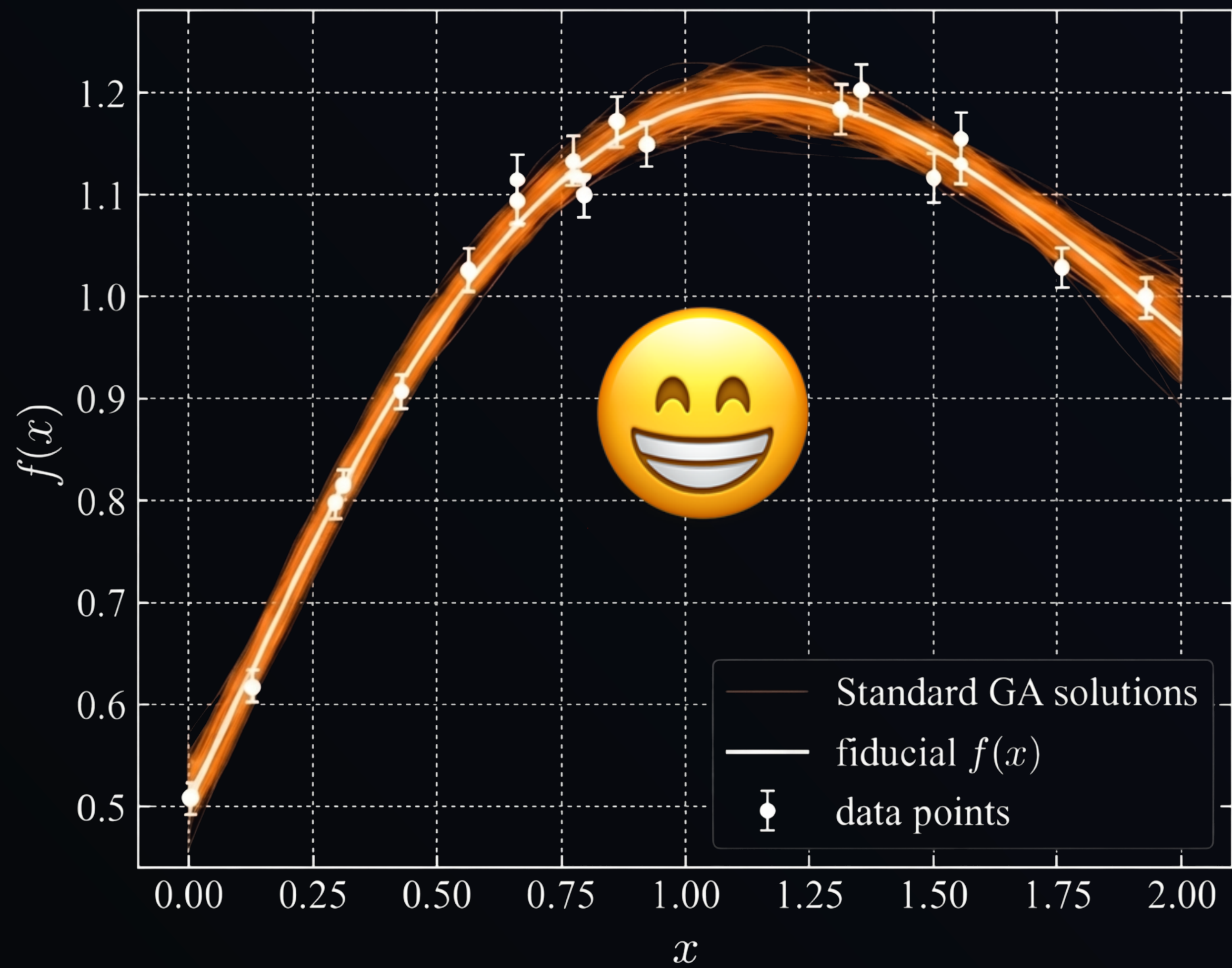


...many hyperparameters configurations

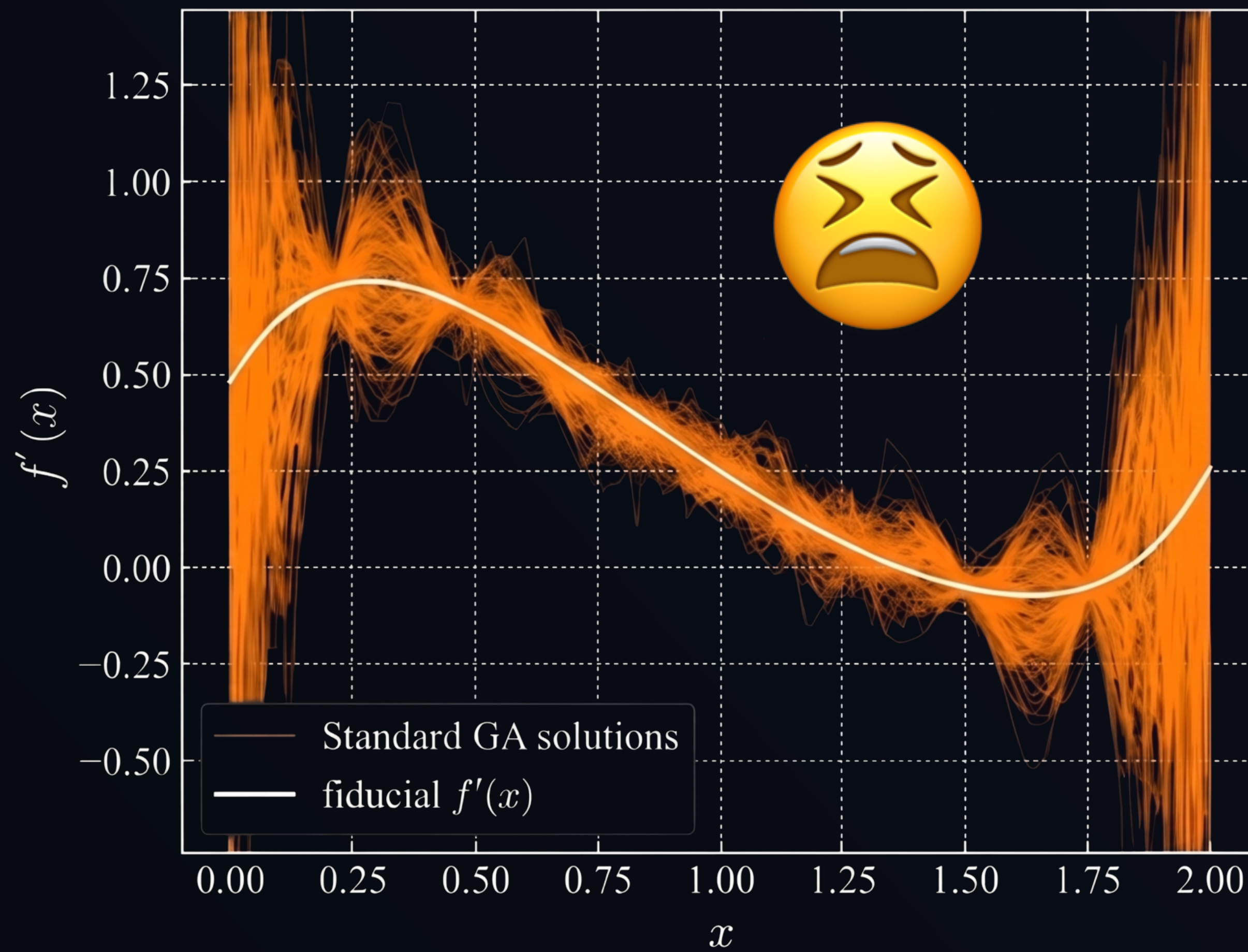
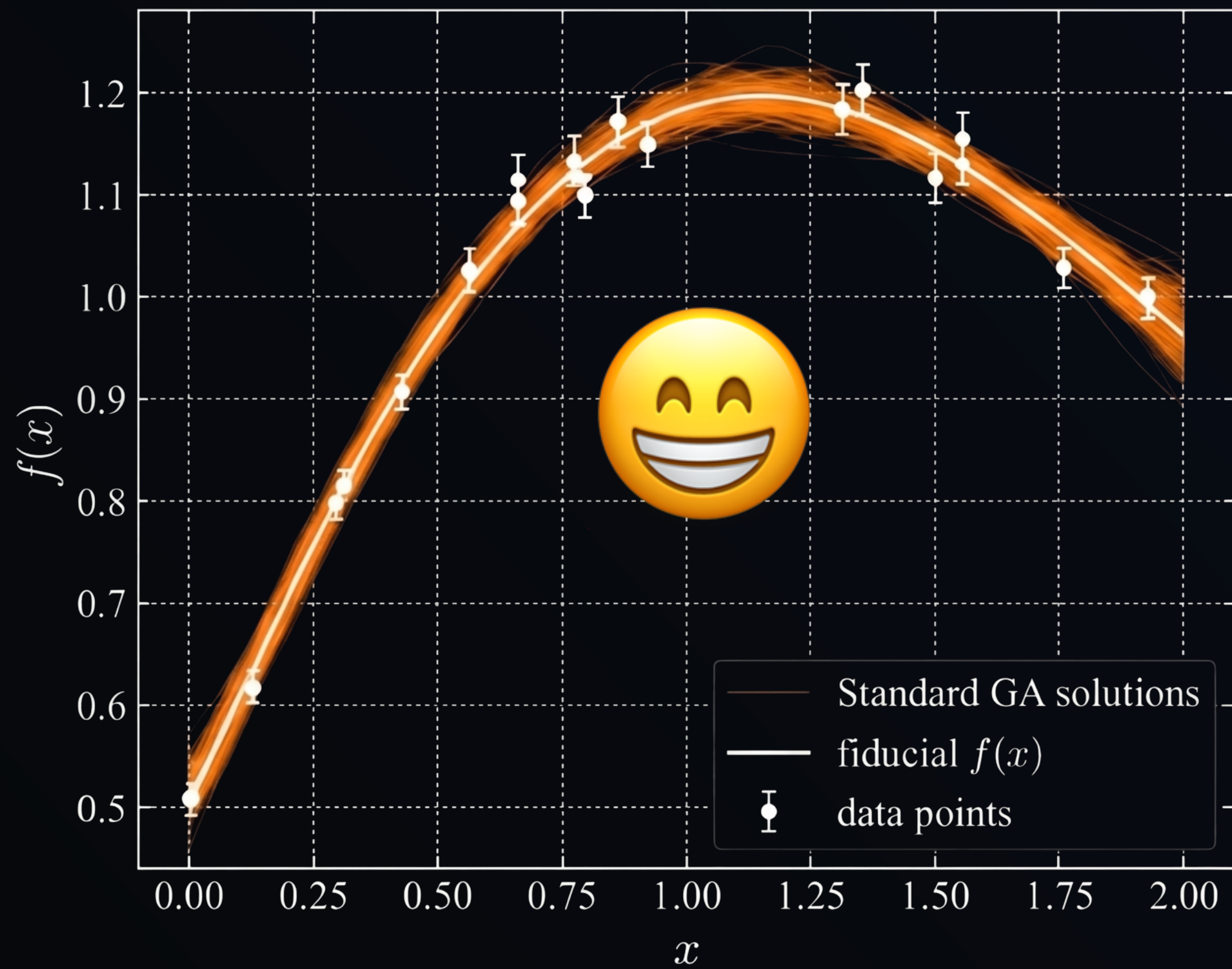
# The “Derivative” trap



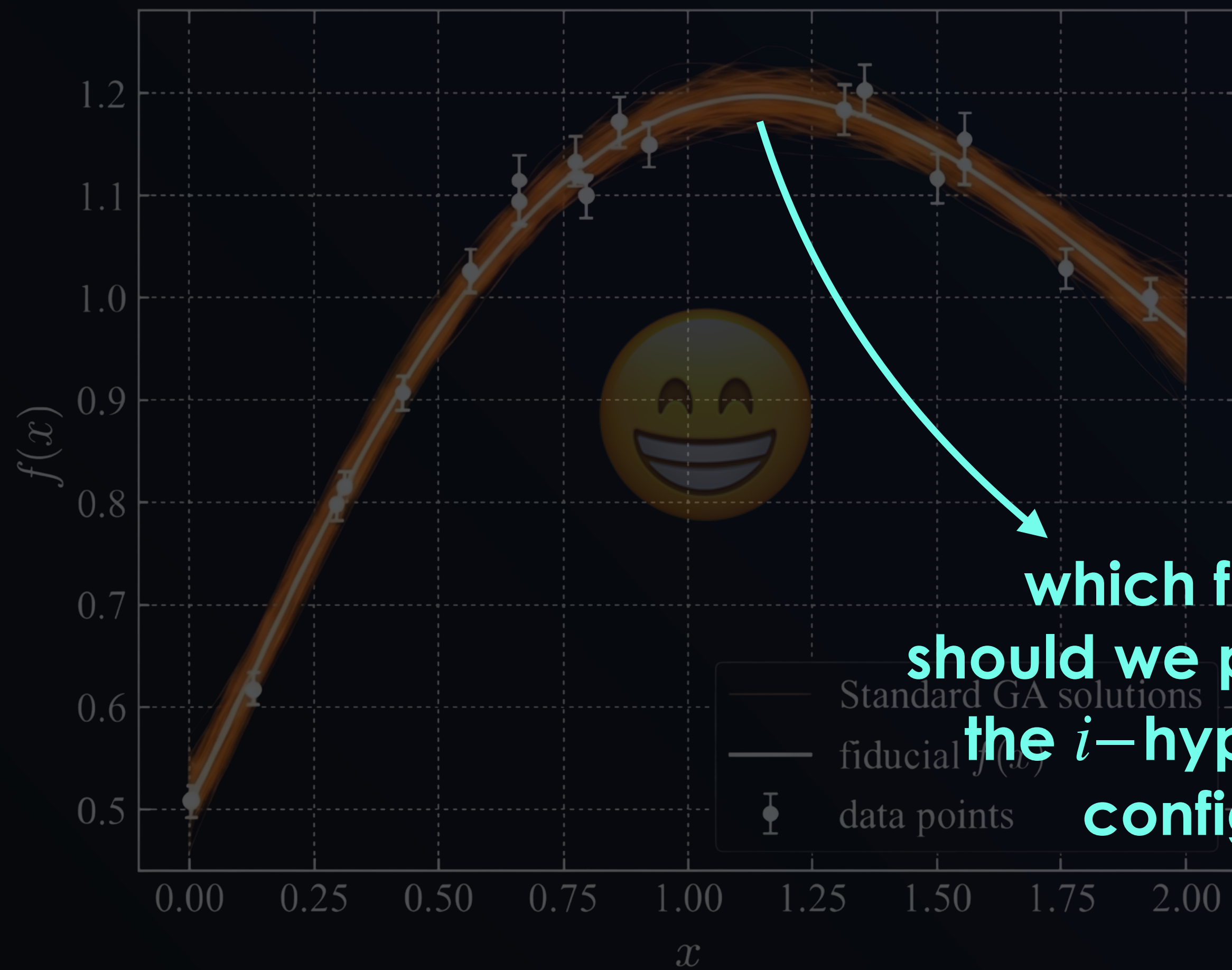
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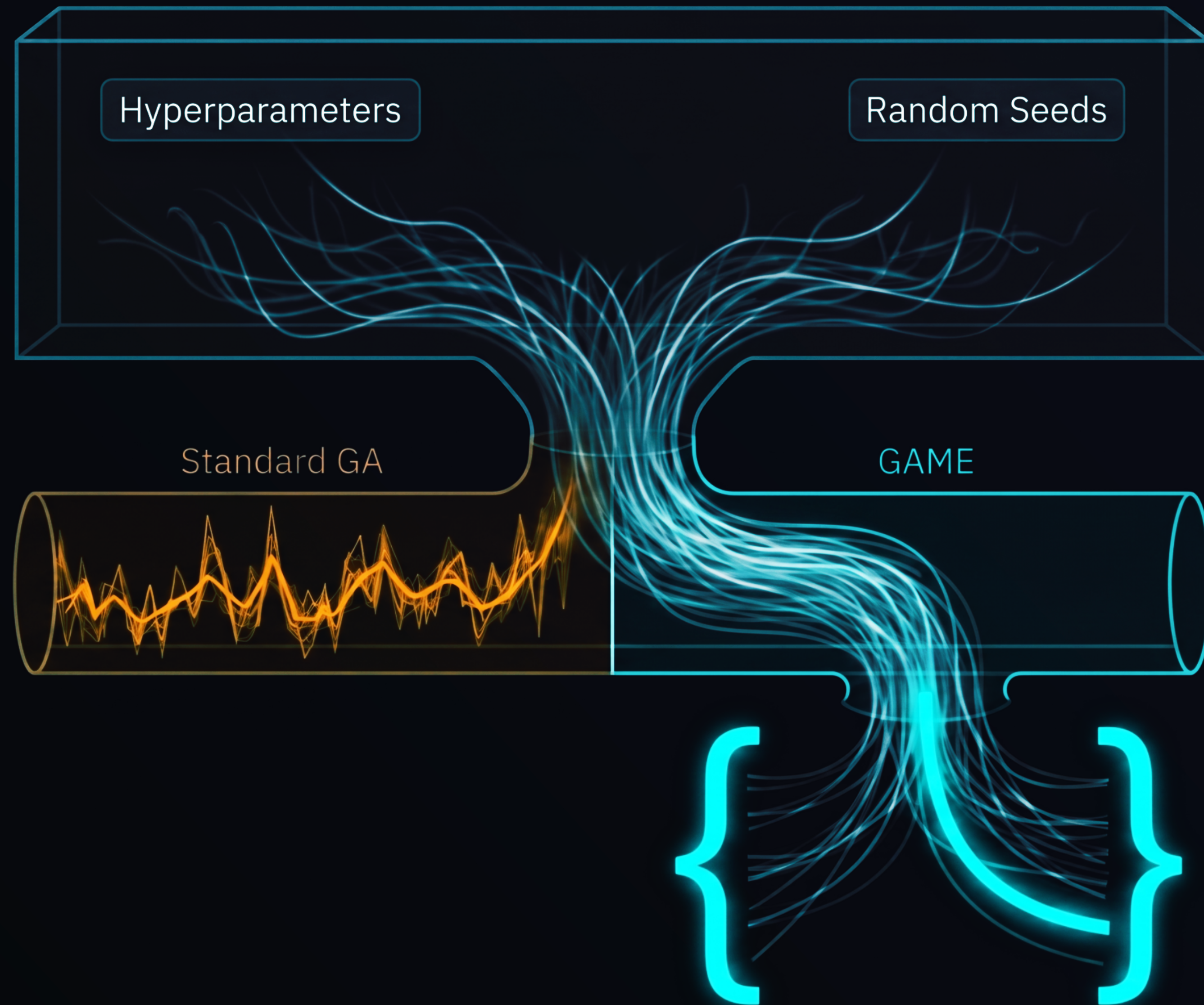
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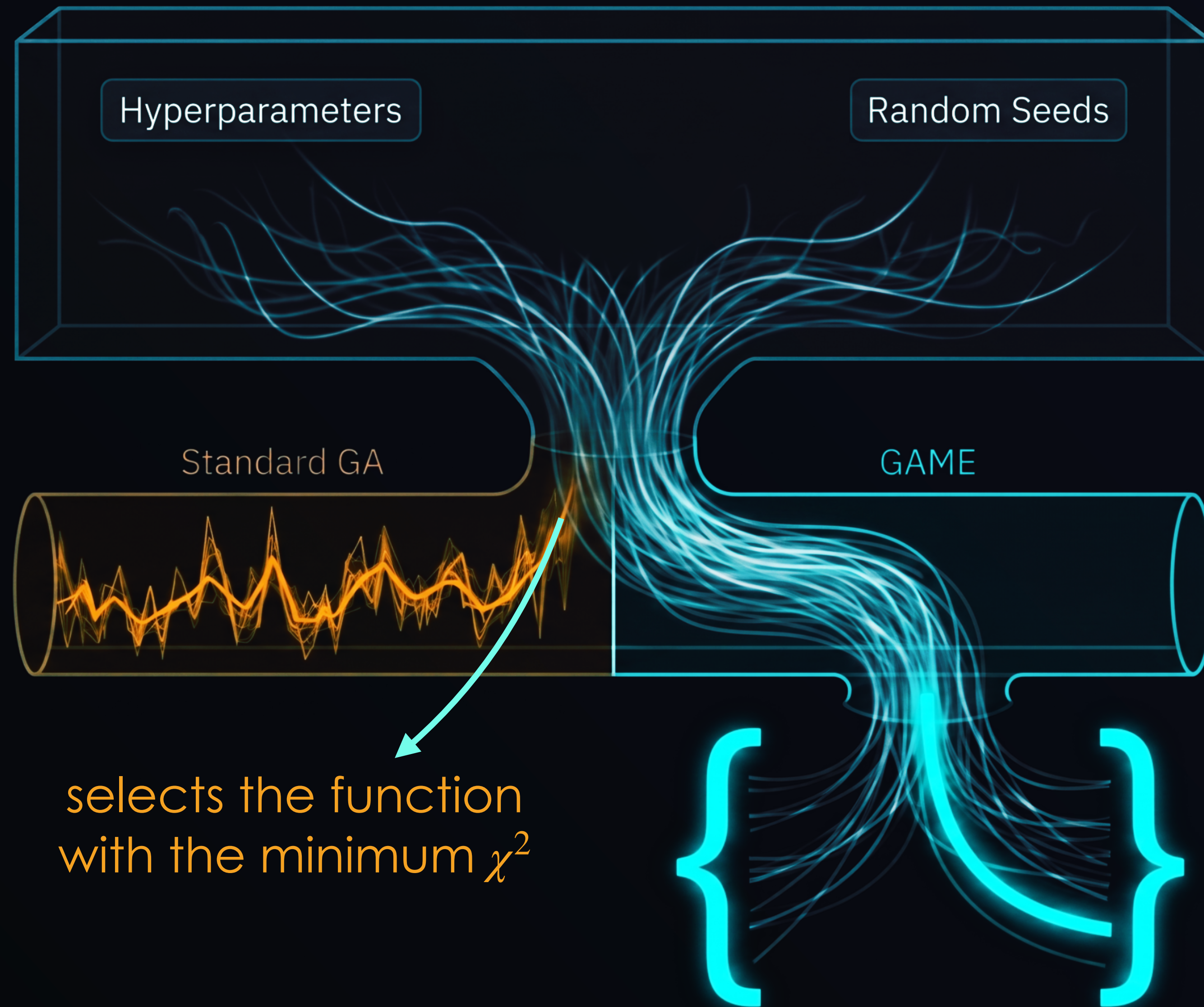
which final function  
should we pick between all  
the  $i$ -hyperparameters  
configurations?



# GAME – Genetic Algorithms with Marginalised Ensembles



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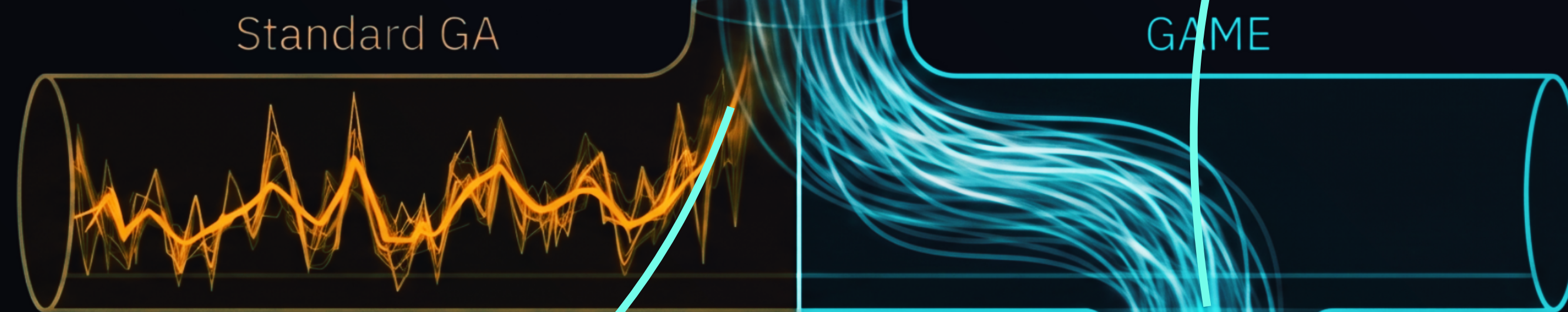


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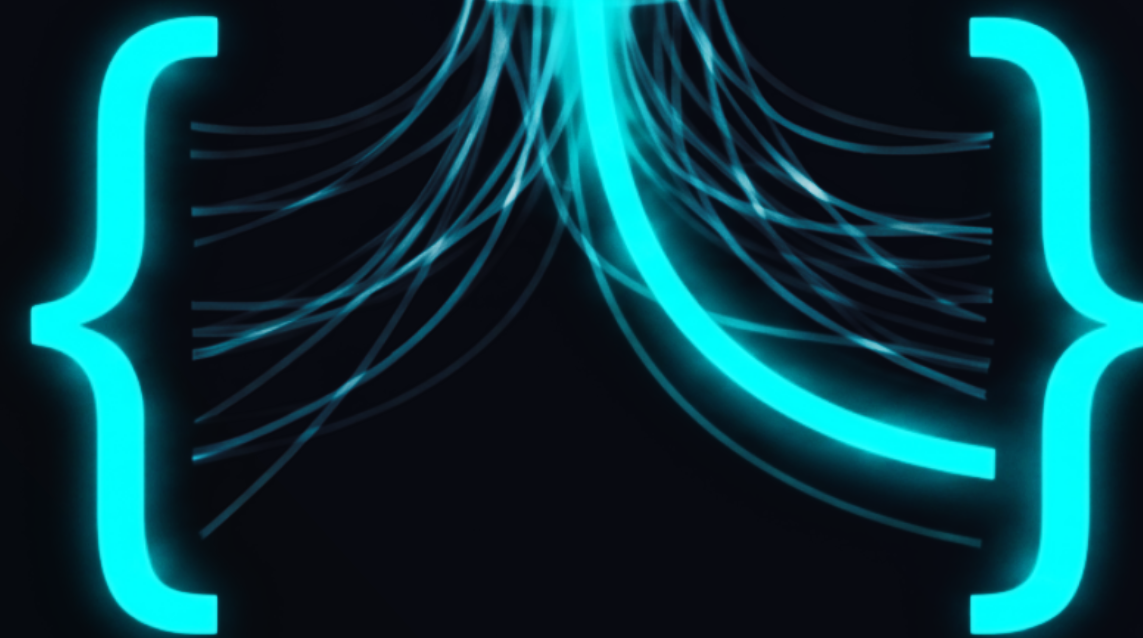


computes a weighted average of all the  $i$  – configurations functions

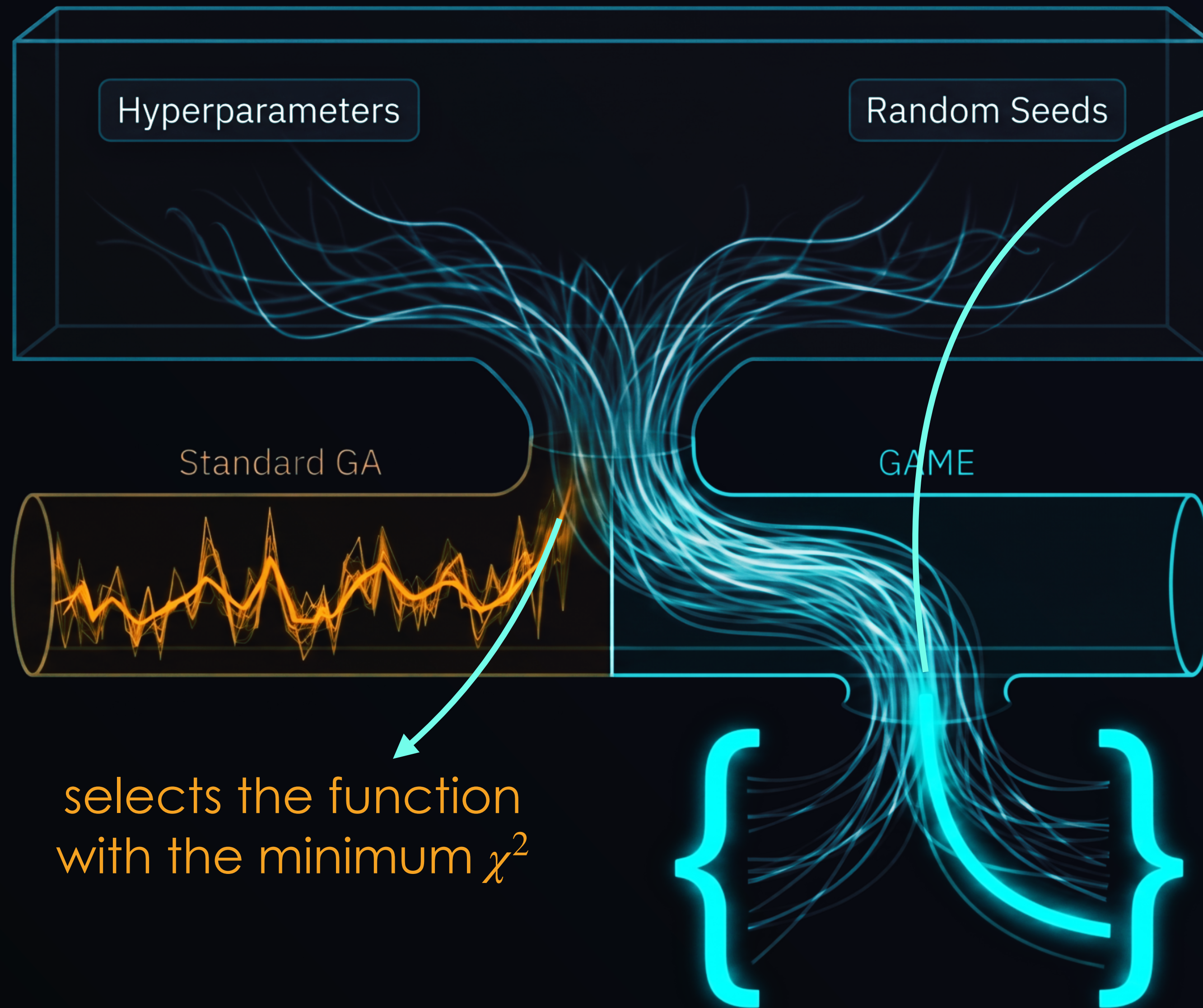
$$f_{\text{GAME}}(x) = \sum_j^{N_{\text{conf}}} w_j \cdot f_{j, \text{GA}}(x)$$



selects the function with the minimum  $\chi^2$



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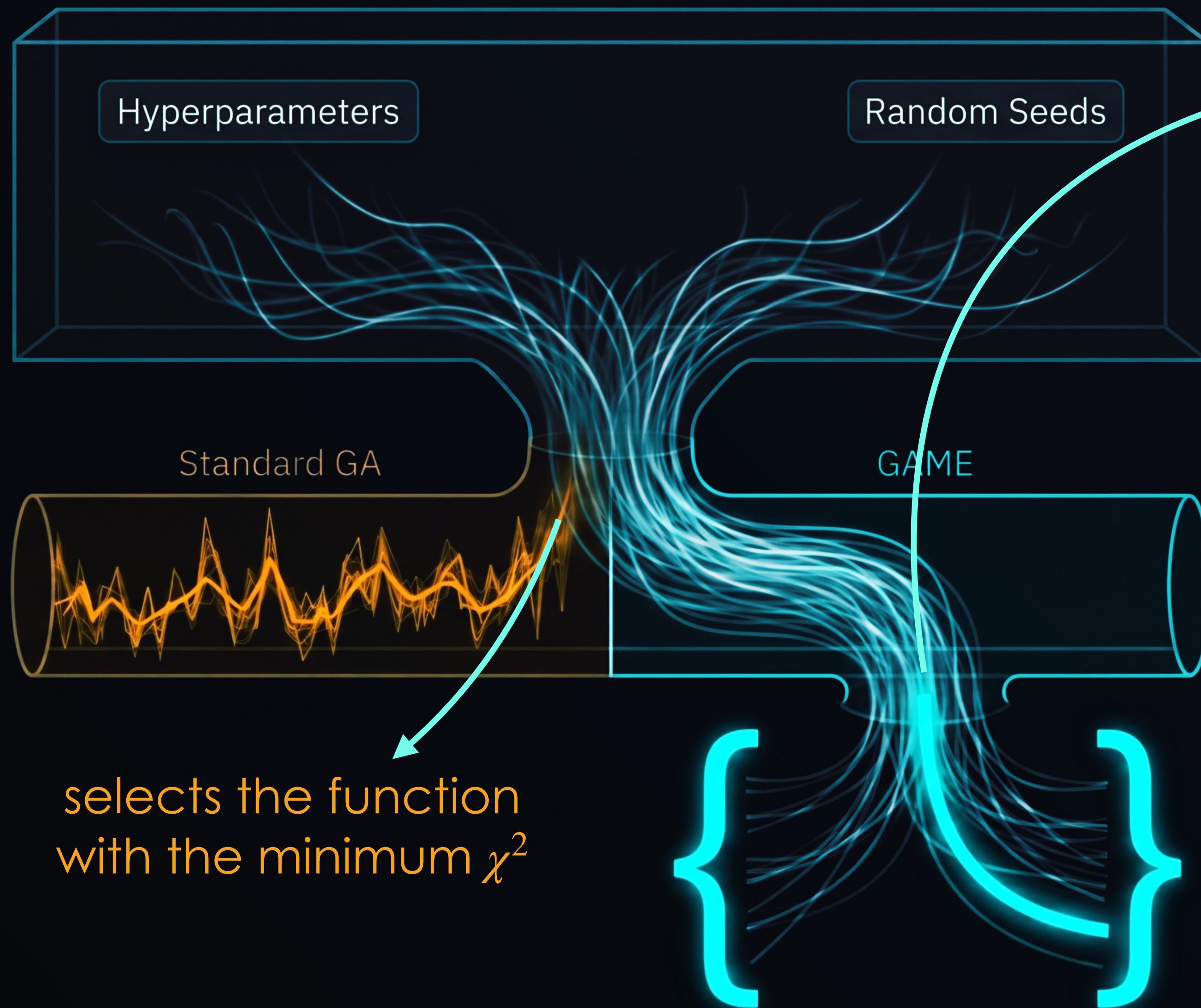
$$w_j \propto S_j = \chi_j^2 + \lambda R_j$$

and

$$R_j = \int \left| f_j''(x) \right|^2 dx$$

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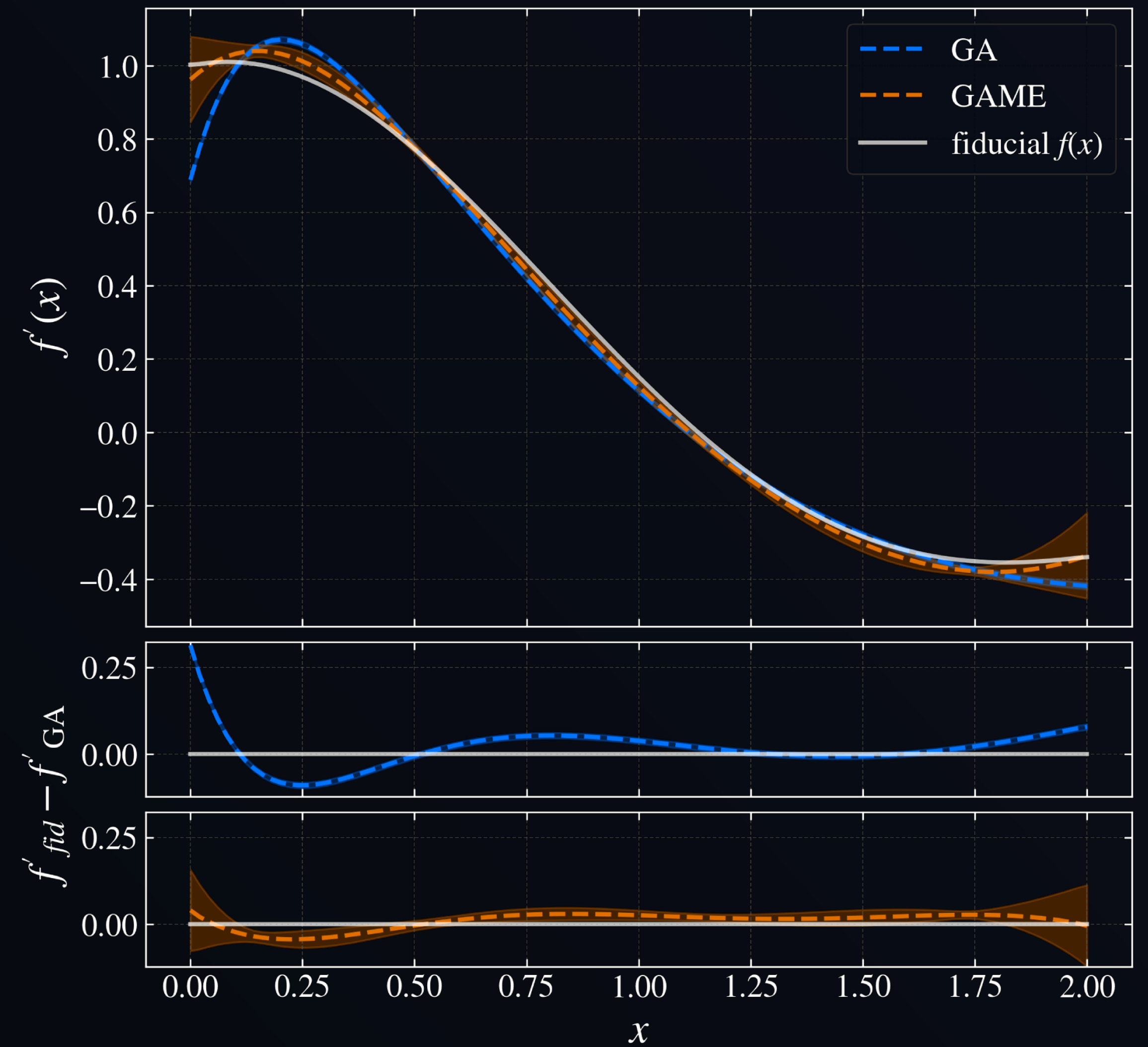
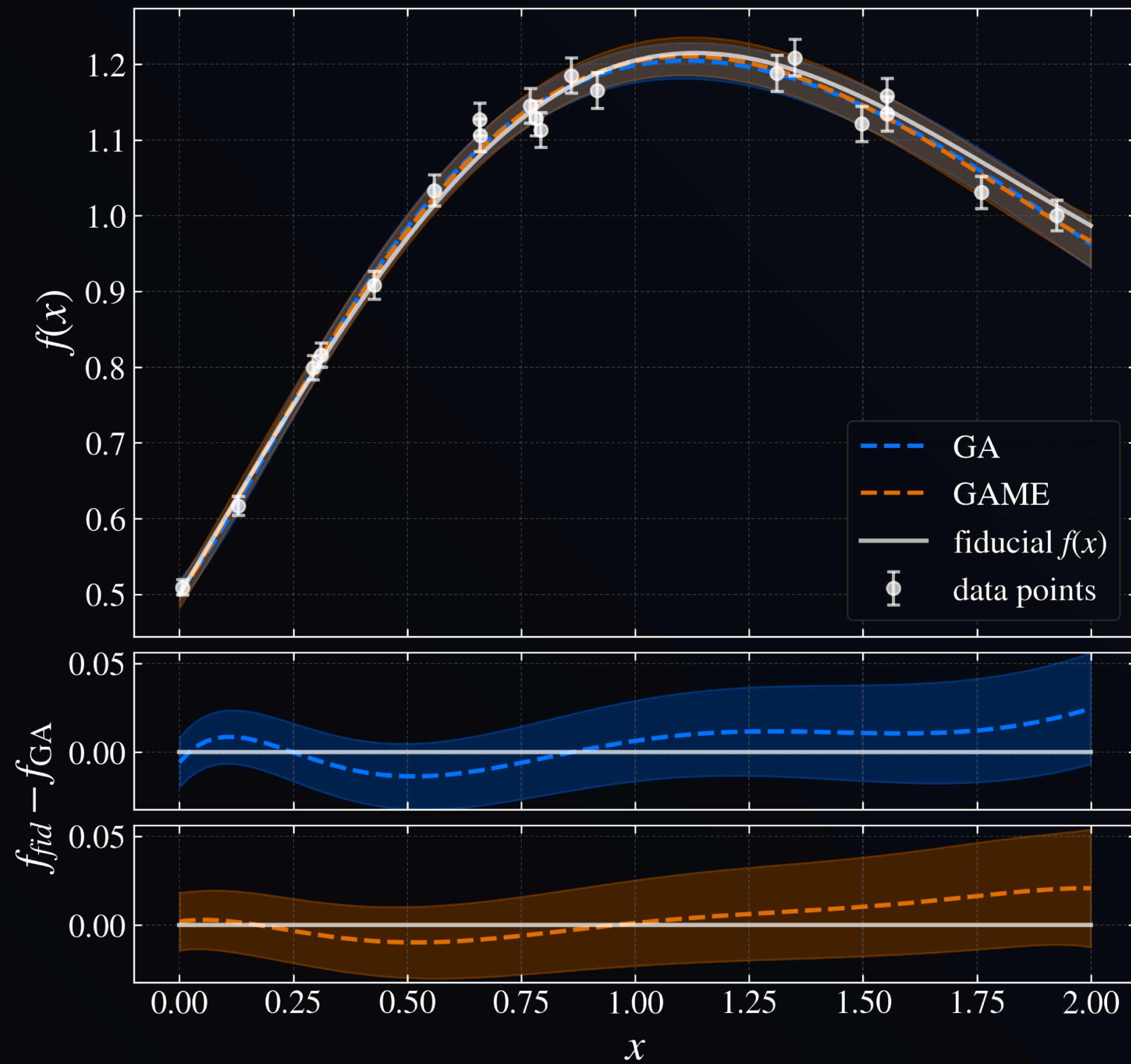
$$w_j \propto S_j = \underbrace{\chi_j^2}_{\text{noisy orange line}} + \underbrace{\lambda R_j}_{\text{smoothing}}$$

and

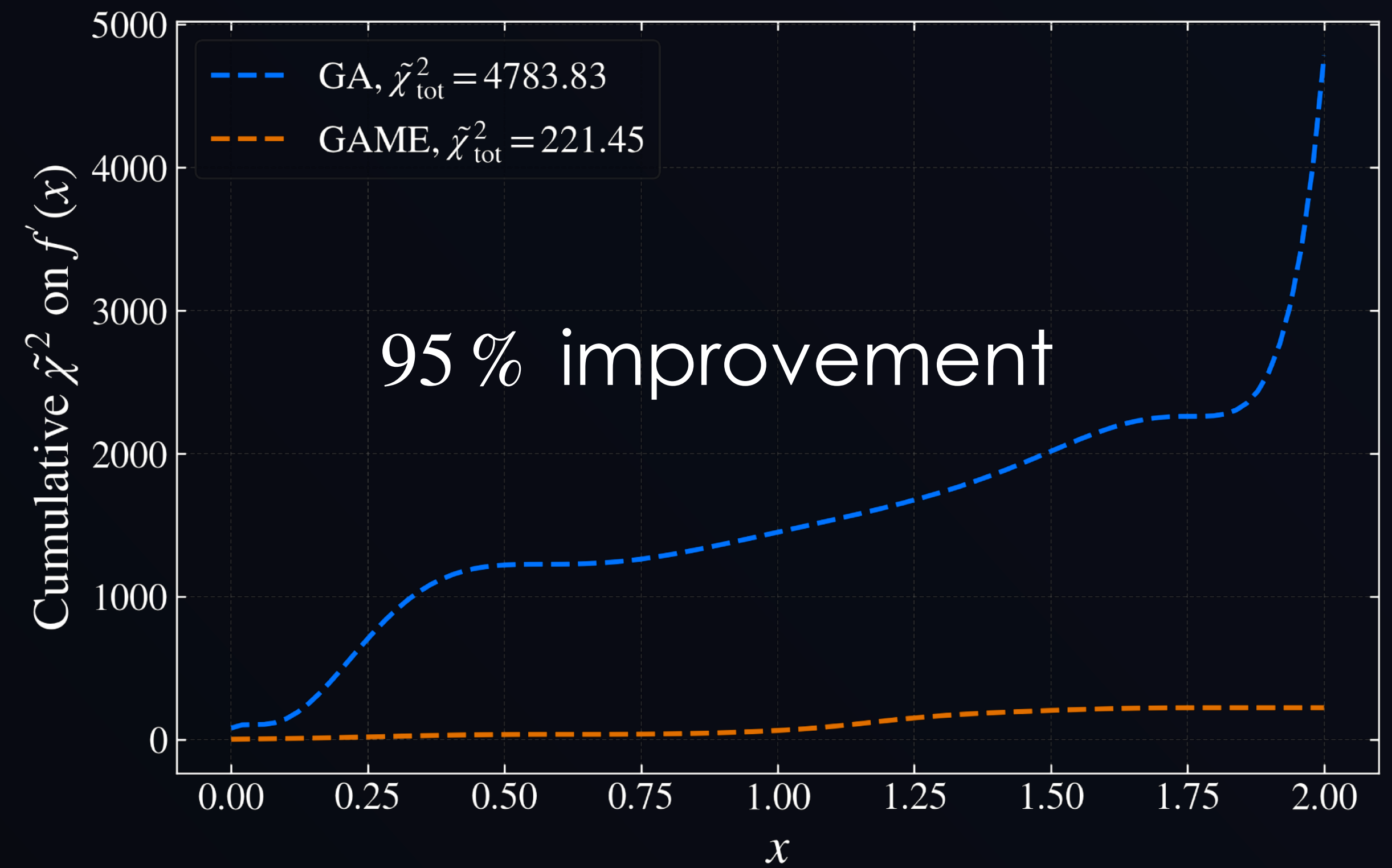
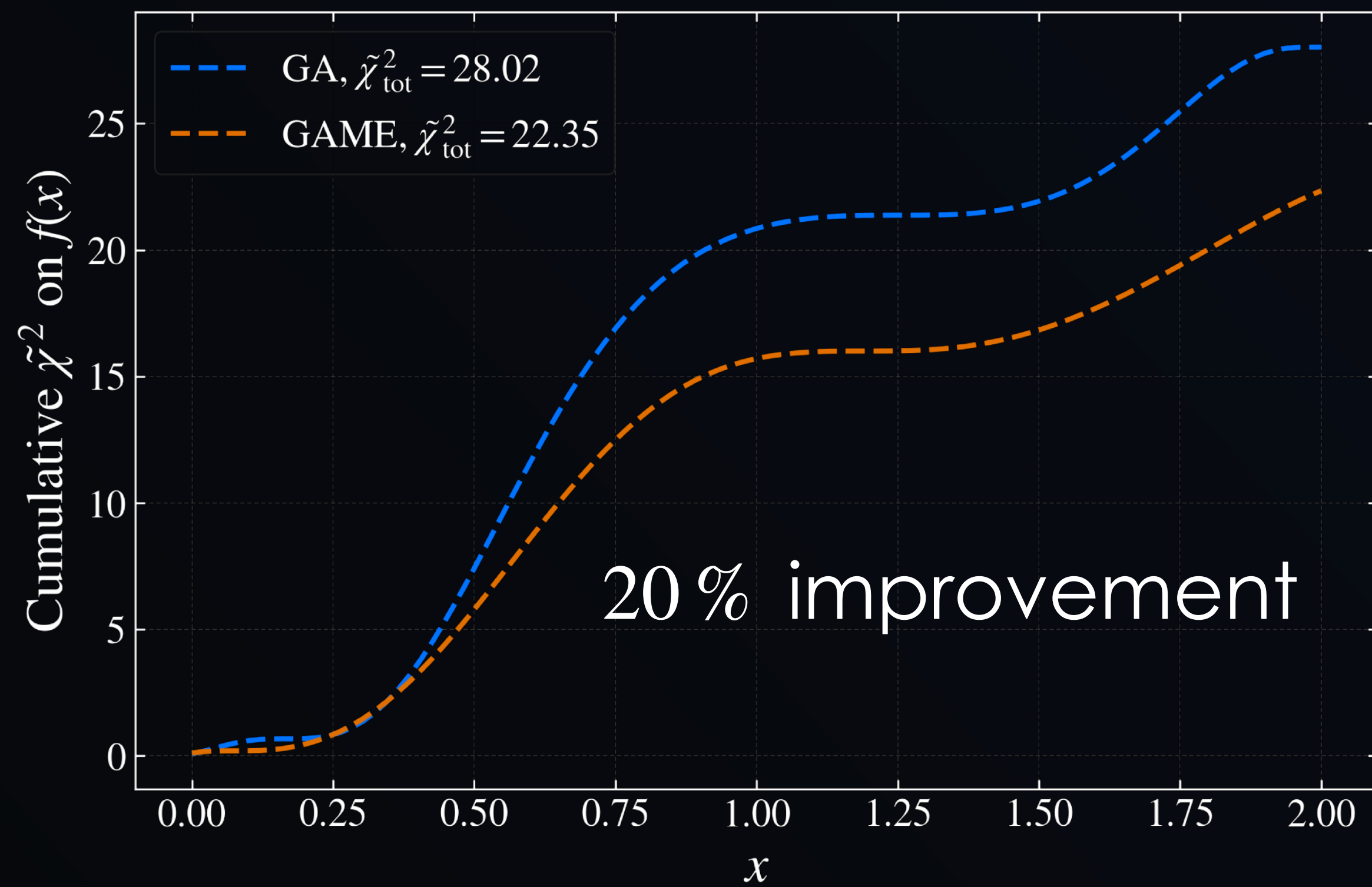
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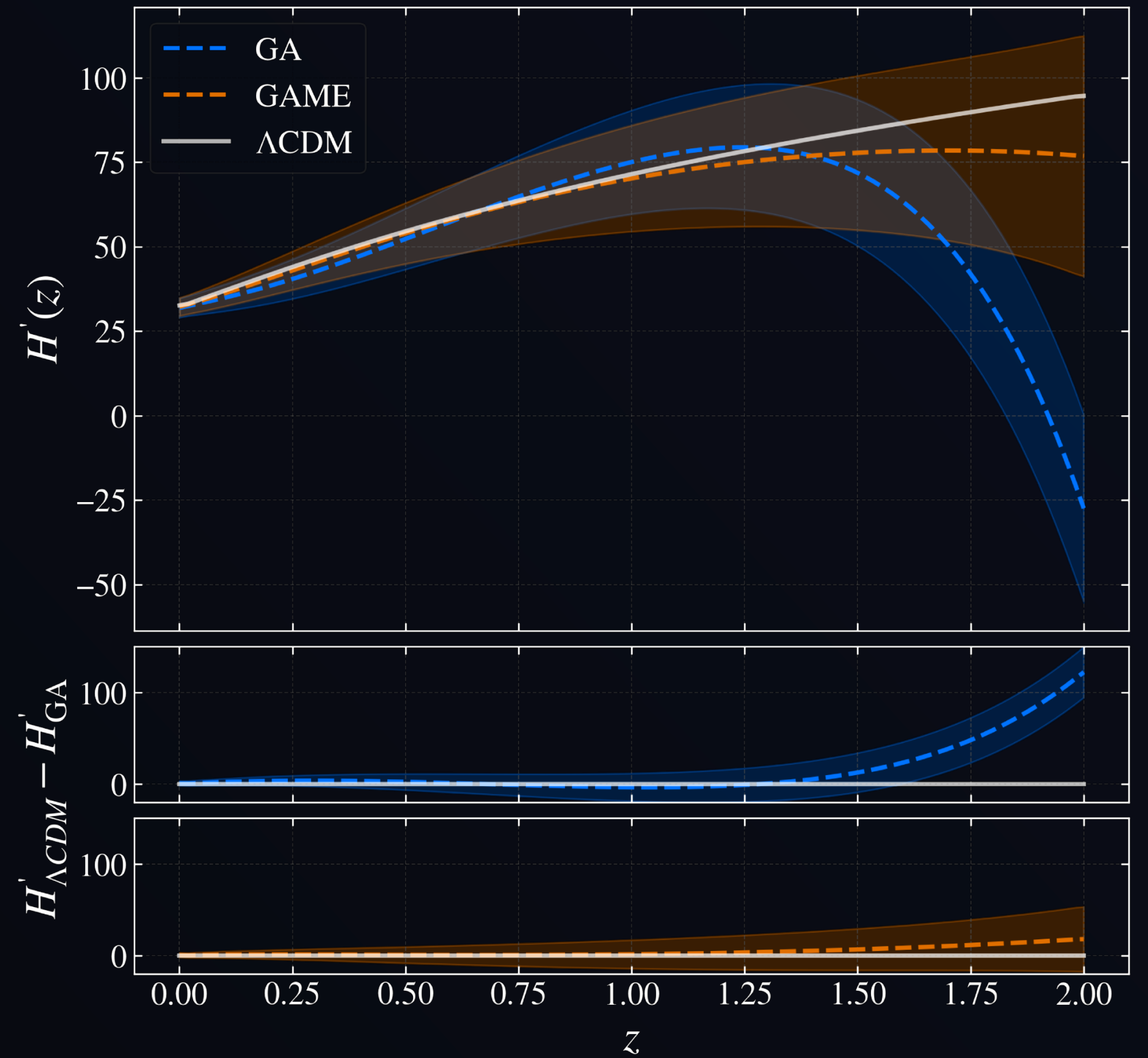
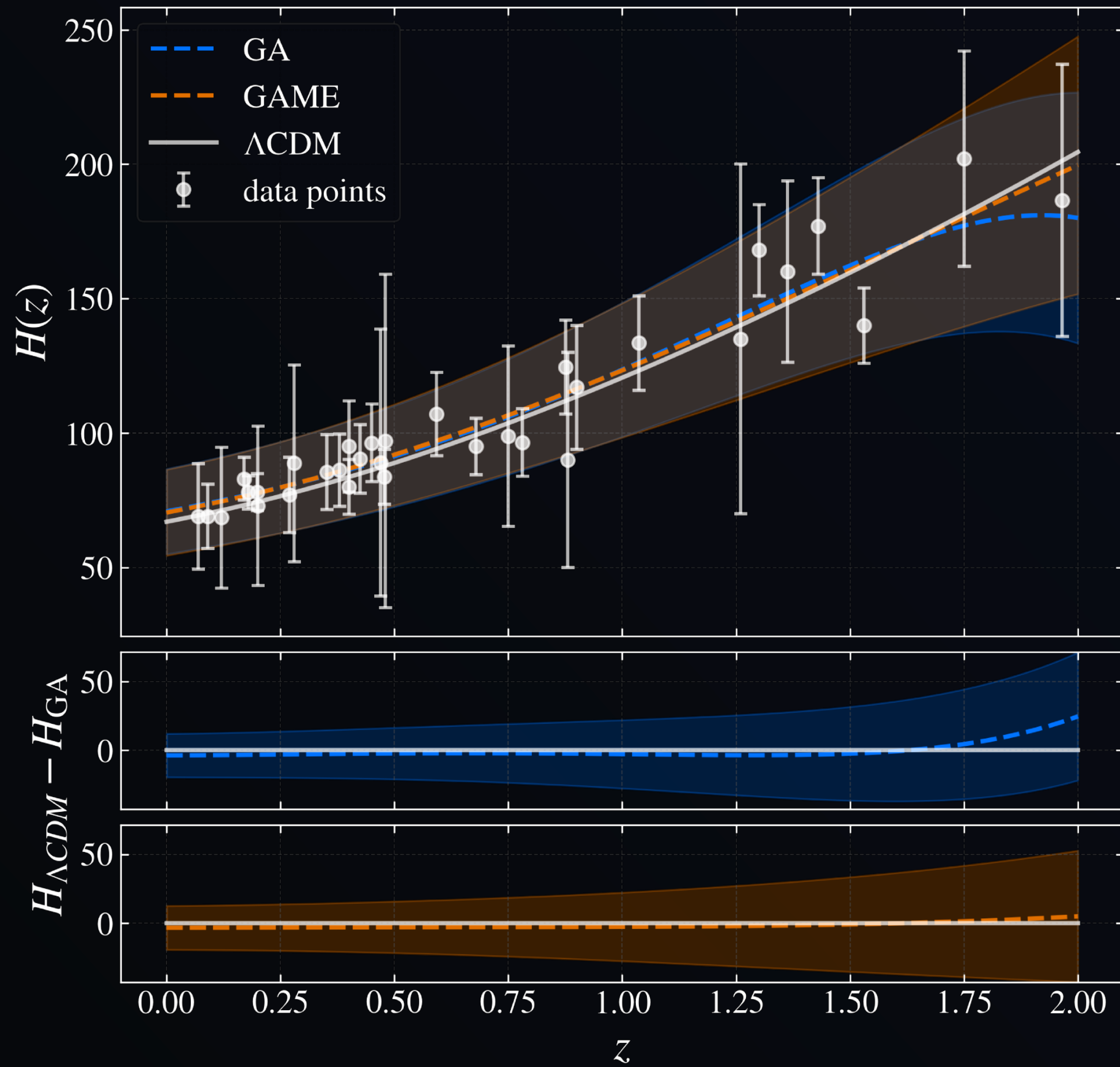
# GAME – generic test function



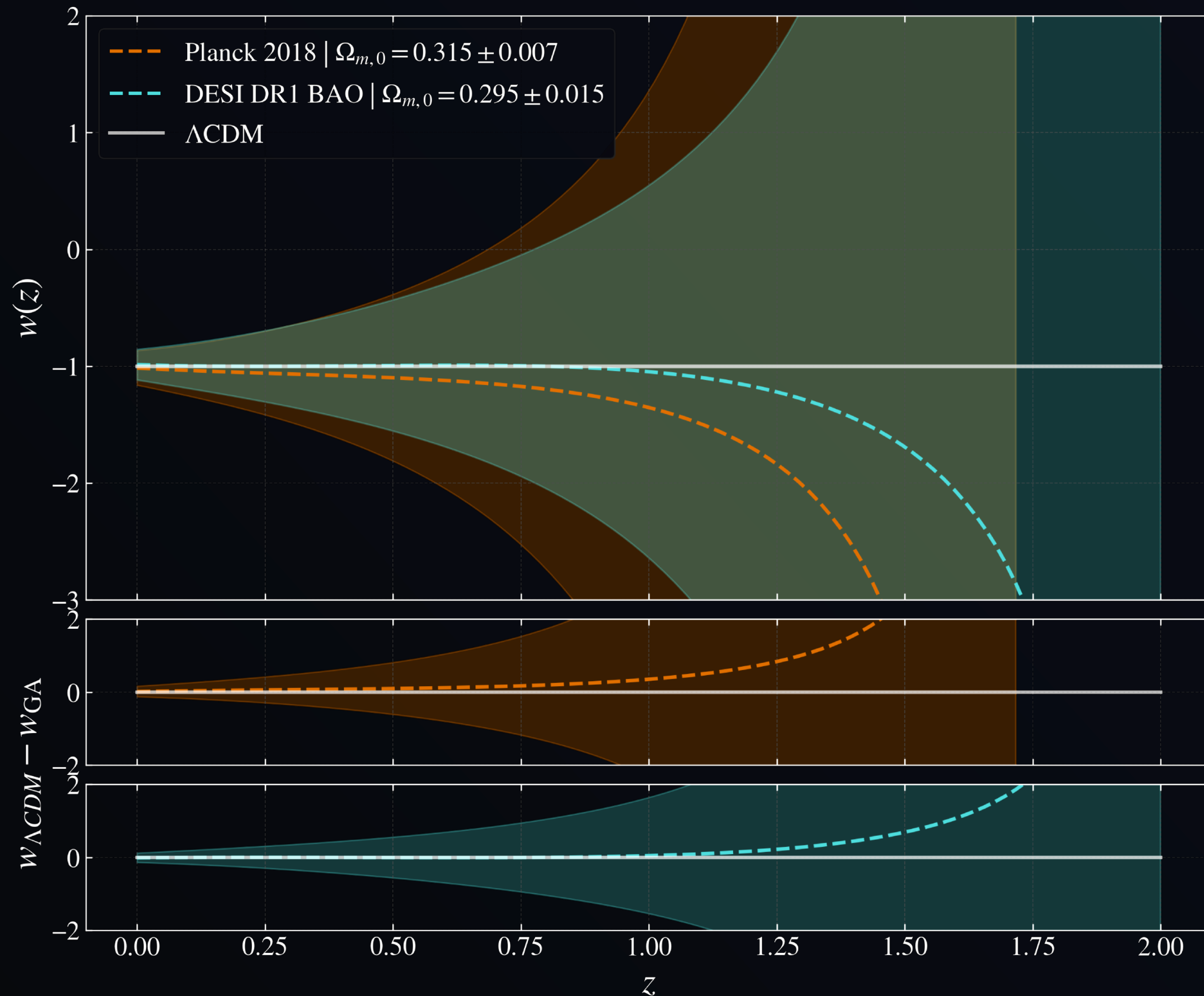
# GAME – generic test function



# Reconstruction of $H(z)$ from CC



# Reconstruction of $w_{bg}(z)$ from CC



$$w(0) = -0.986 \pm 0.132$$

**Perfectly compatible with  $\Lambda$ CDM but large uncertainties bands using current CC data**

# Conclusions and next steps

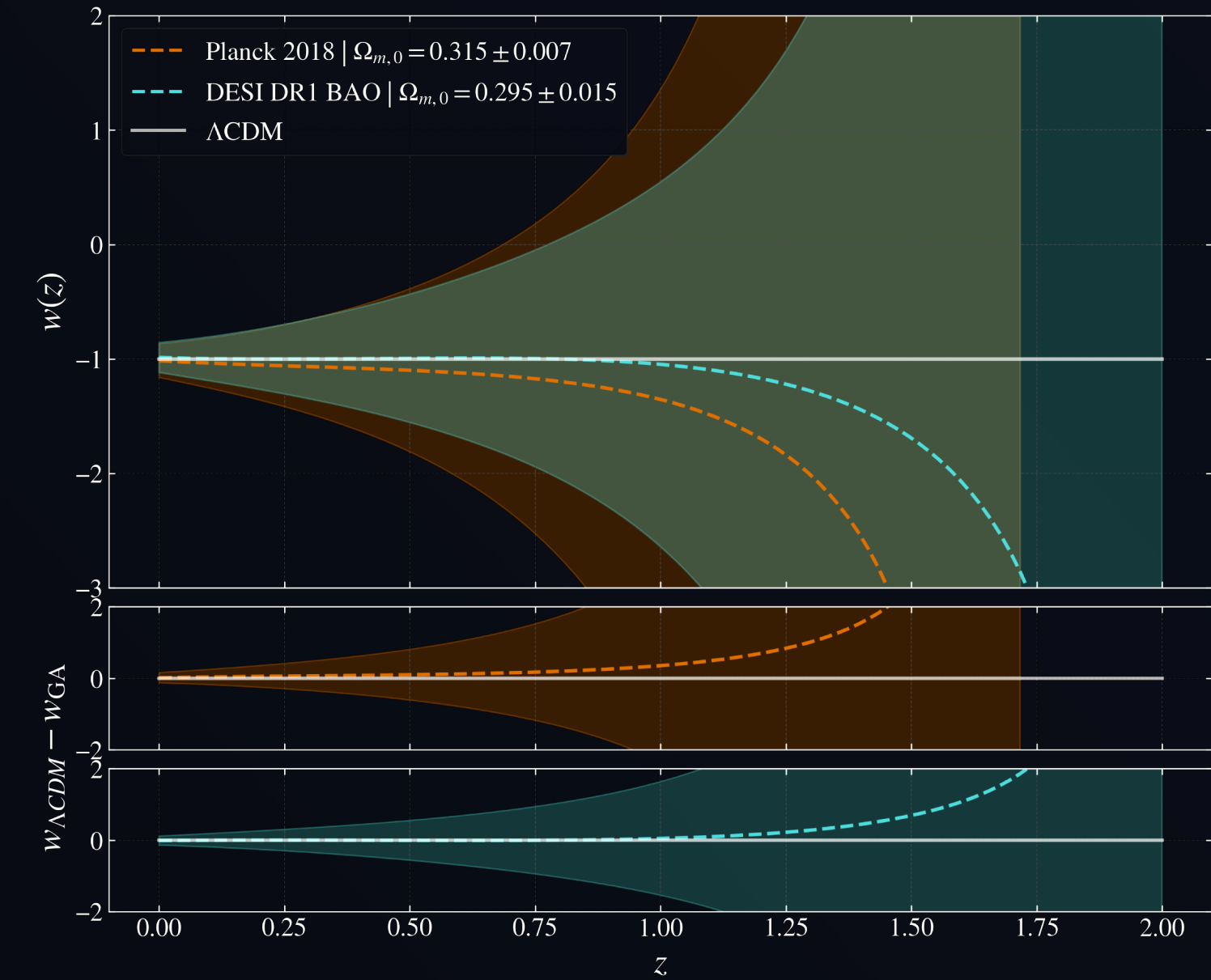
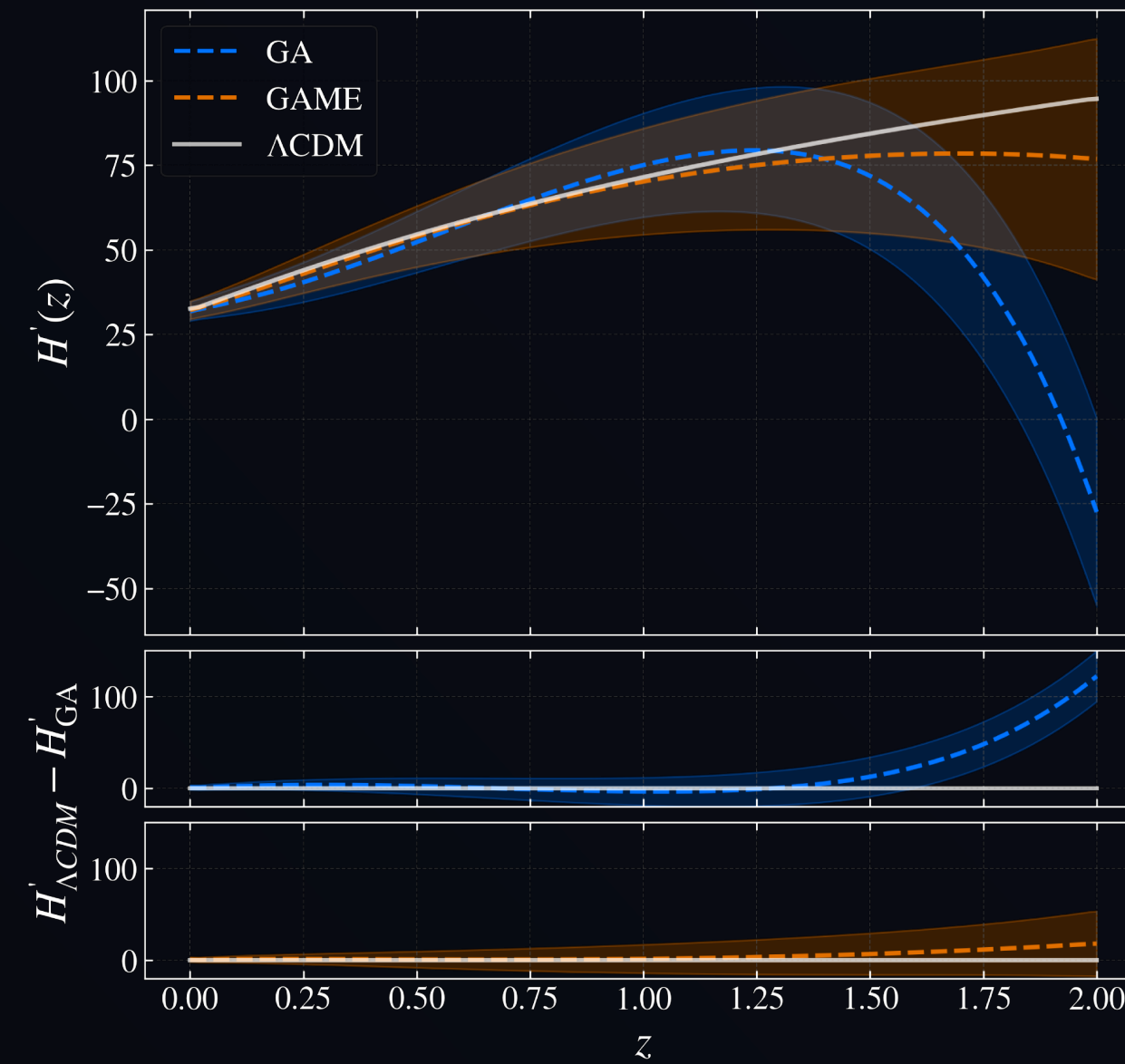
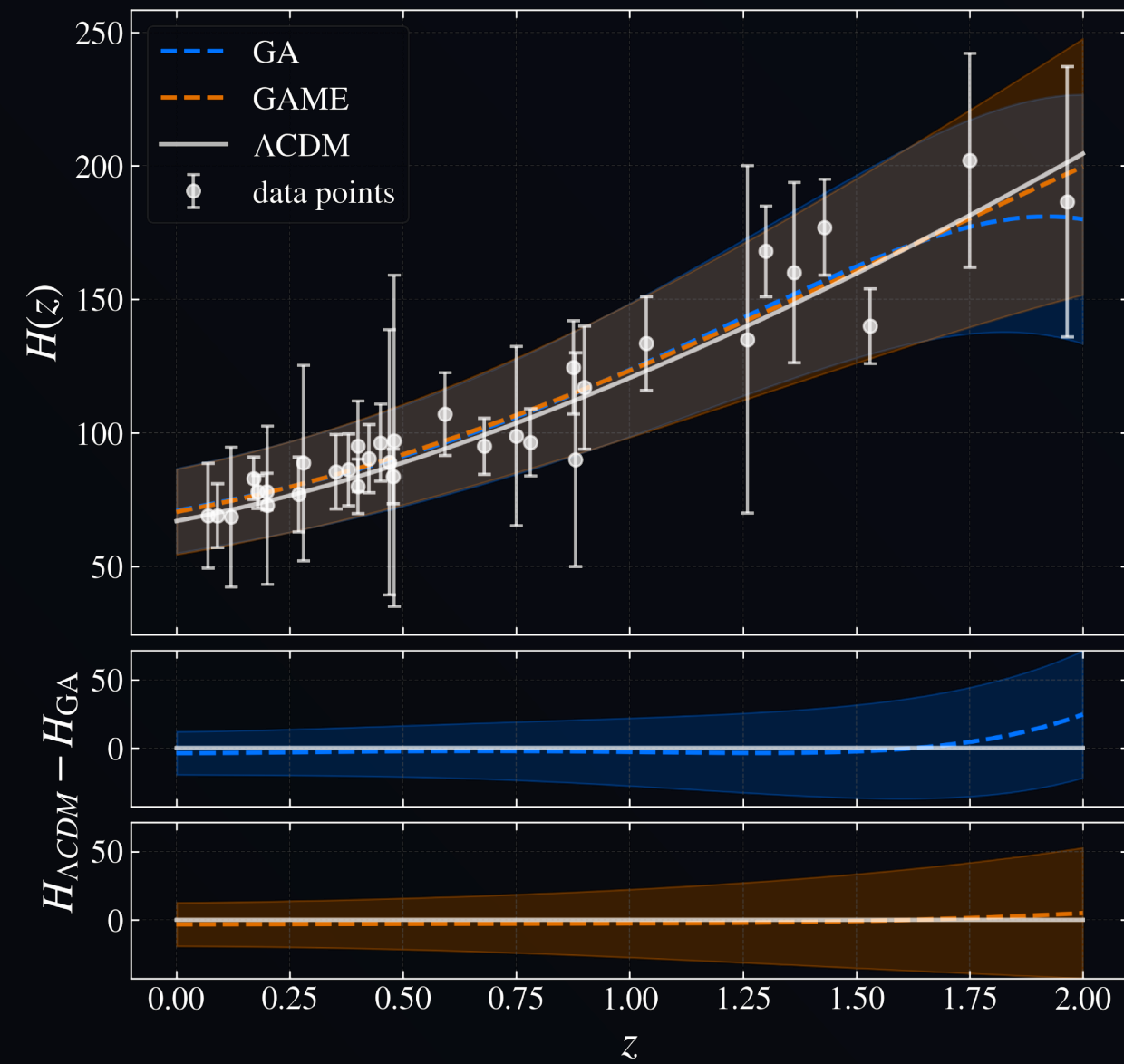
# Conclusions and next steps

- GAME is a powerful tool that can reconstruct smooth analytical functions with a particular sensitivity to derivatives, only from data.
- By applying GAME to current CC data, we get an expression for  $w_{bg}(z)$  that is perfectly compatible with  $\Lambda$ CDM, but with large uncertainties bands due to data quality.

# Conclusions and next steps

- GAME is a powerful tool that can reconstruct smooth analytical functions with a particular sensitivity to derivatives, only from data.
  - By applying GAME to current CC data, we get an expression for  $w_{bg}(z)$  that is perfectly compatible with  $\Lambda$ CDM, but with large uncertainties bands due to data quality.
1. Apply GAME to BAO and Supernovae (to obtain  $H(z) \rightarrow w(z)$ ).
  2. Derive  $w(z)$  from reconstructions of perturbations quantities (i.e  $\sigma_8(z)$ ).
  3. Compare GAME with other ML regression algorithms (i.e Gaussian Processes).
  4. Wait for new CC points to reveal potential departures from  $\Lambda$ CDM.

# GAME for the reconstruction of cosmological quantities



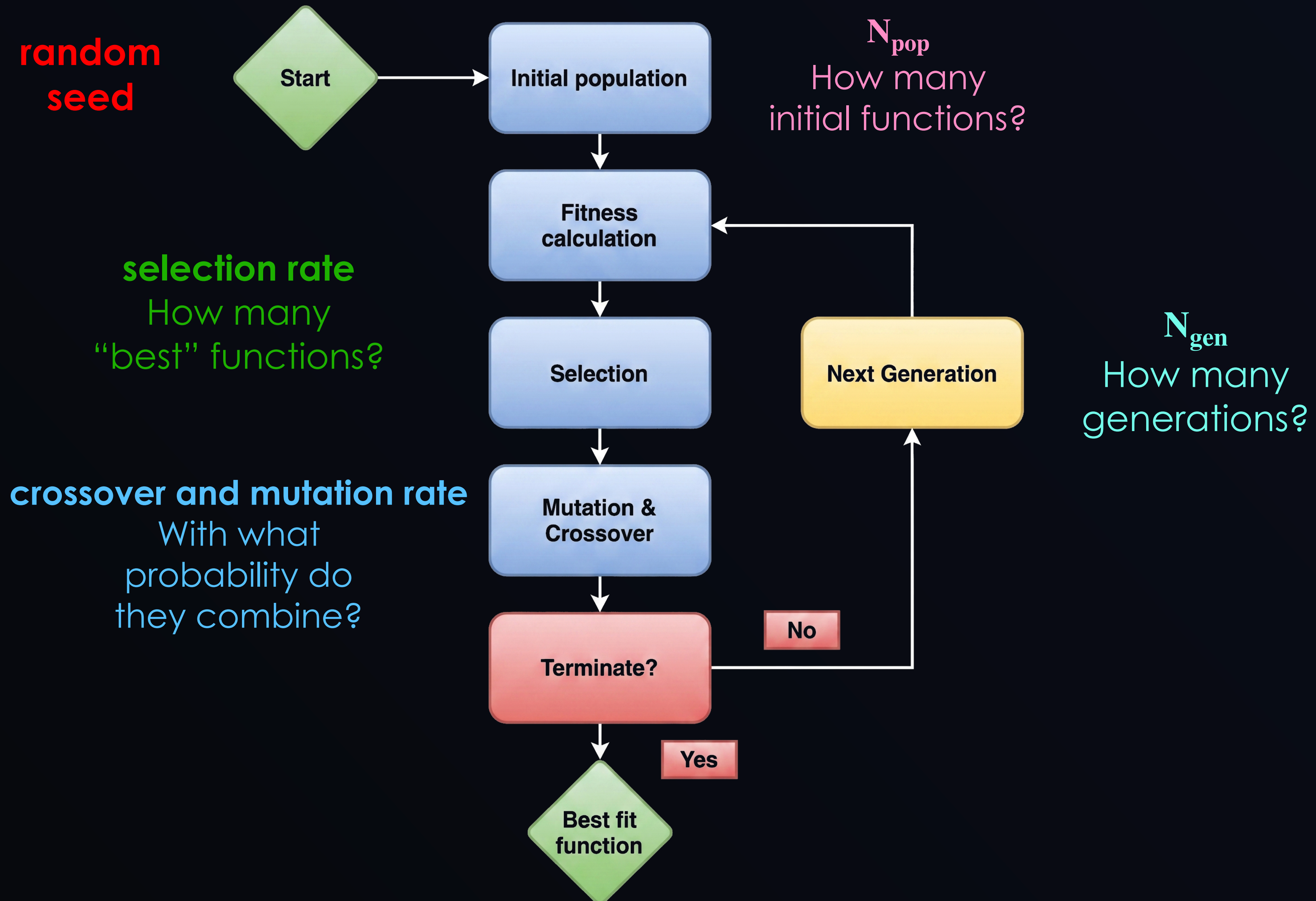
contact

matteo.peronaci@uniroma1.it



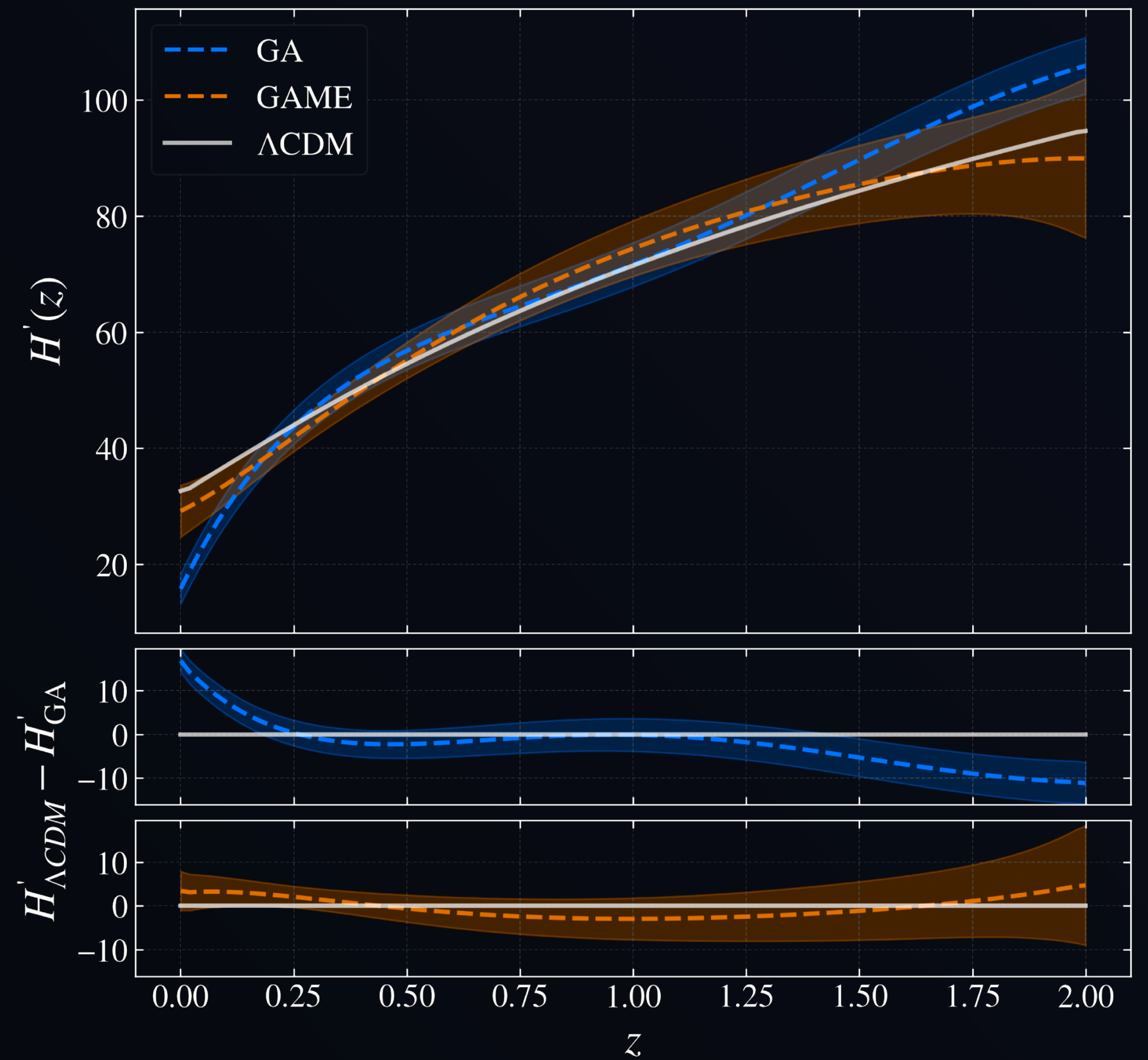
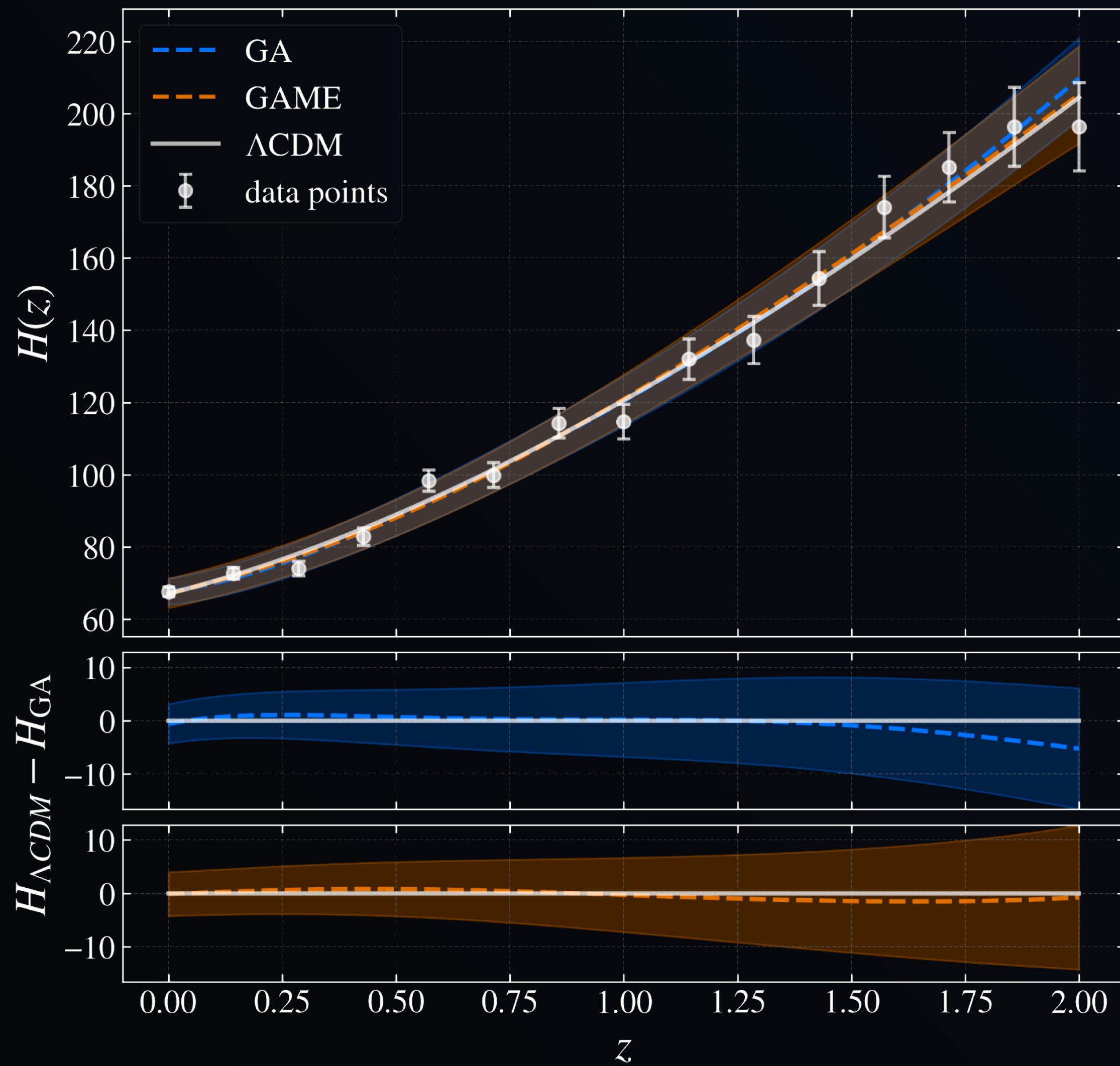
**Backup slides**

# Genetic Algorithms hyperparameters



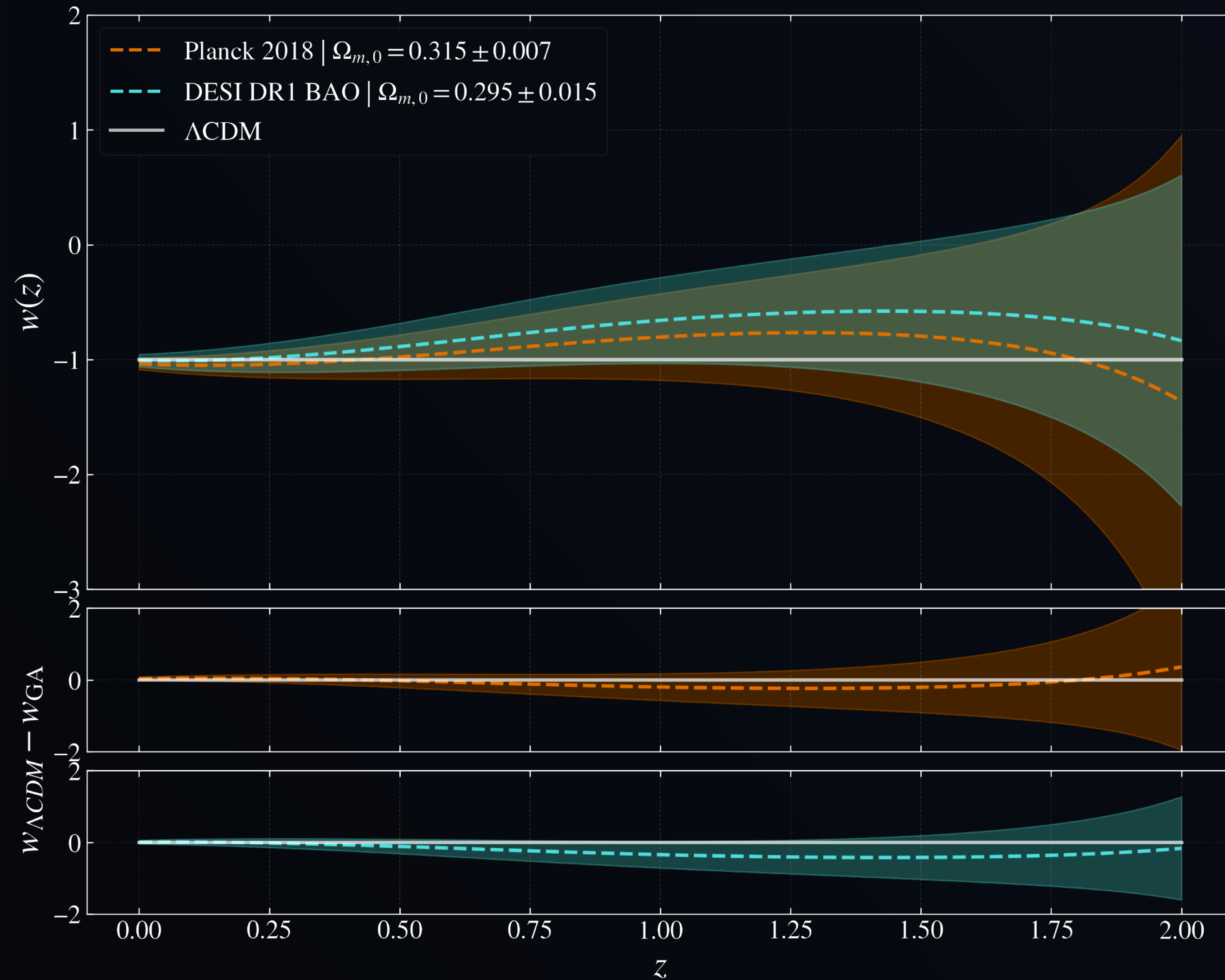
# GAME consistency test with mock data

Reconstruction of  $H(z)$  from mock data

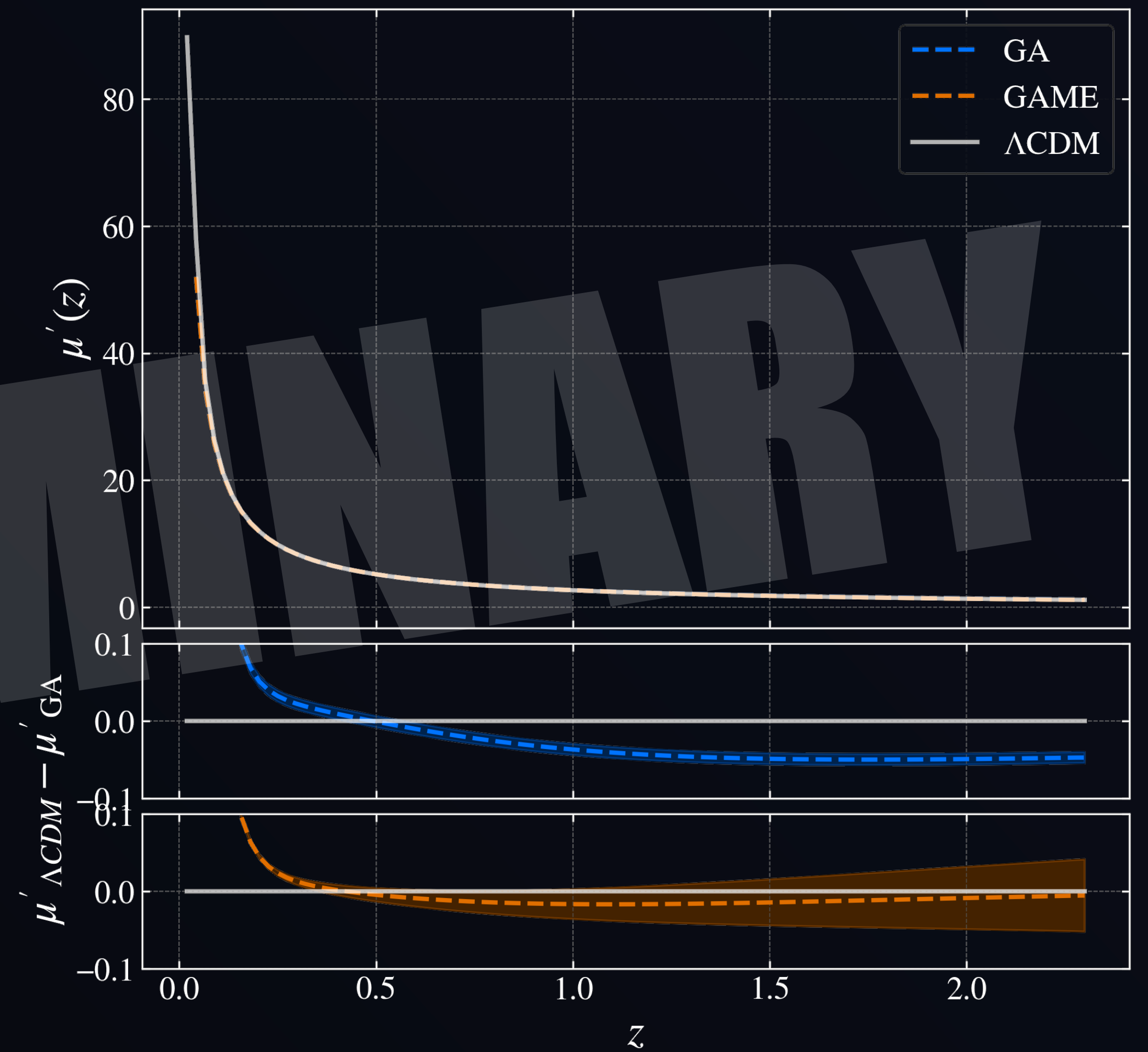
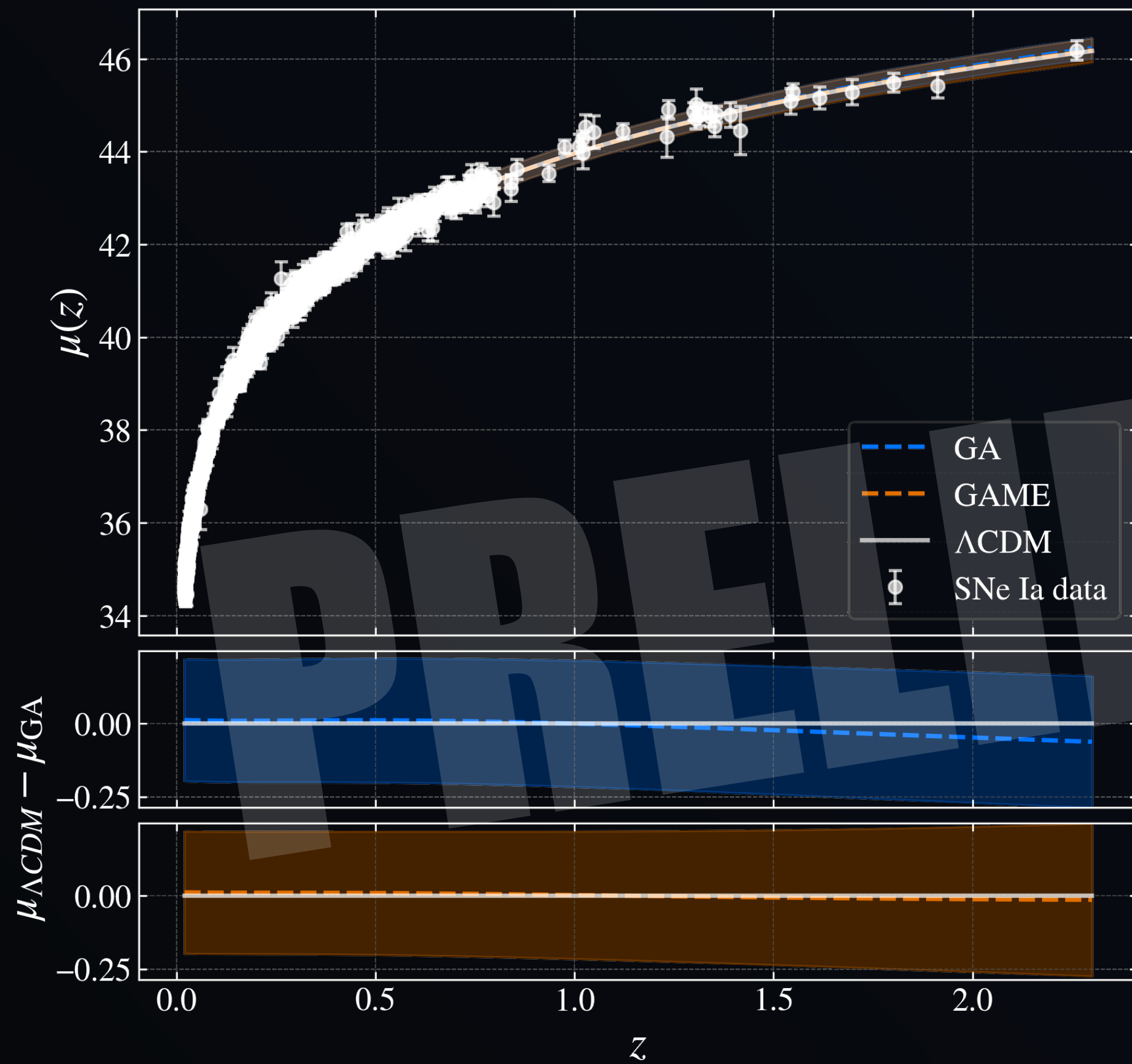


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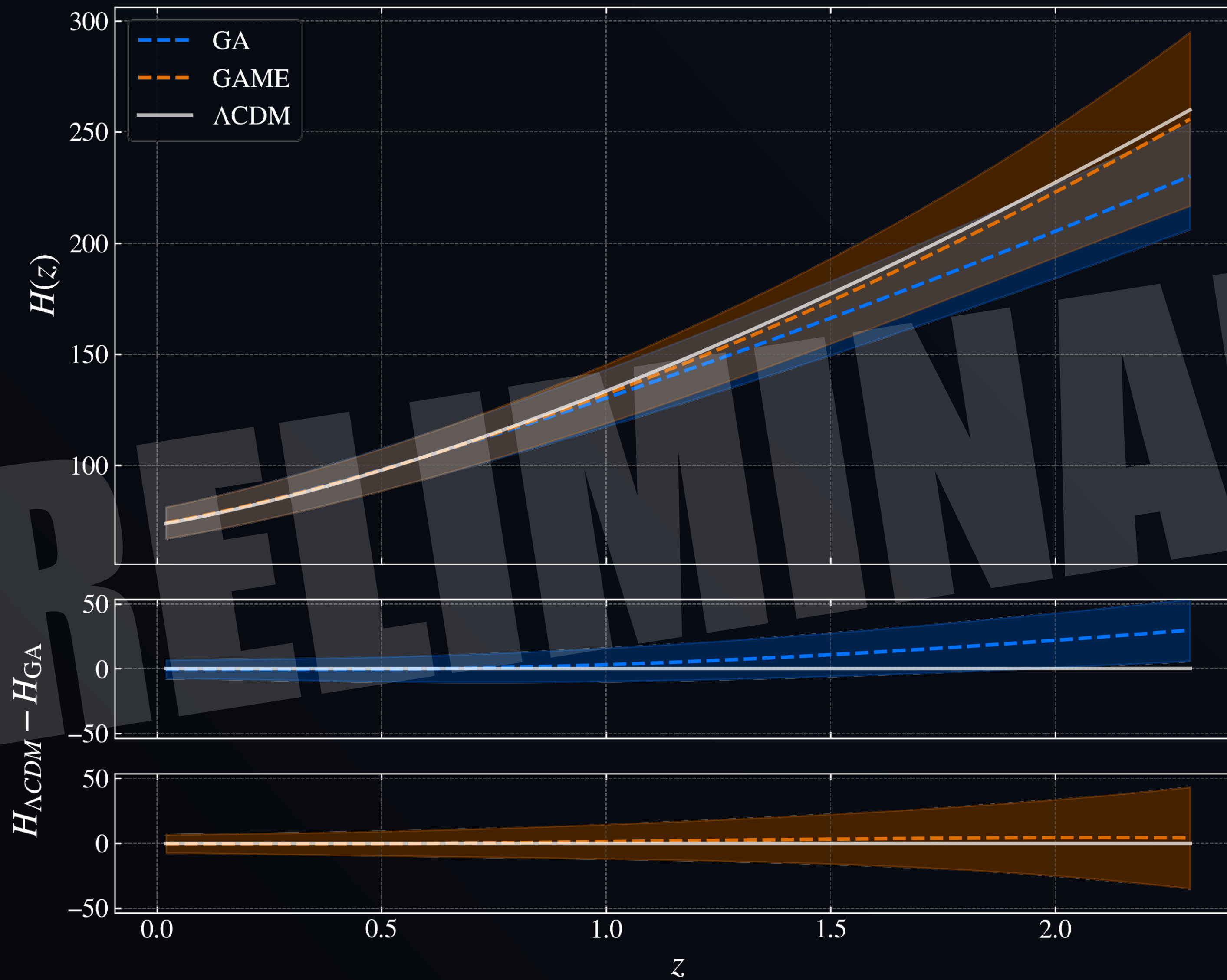
Reconstruction of  $w_{bg}(z)$  from mock data



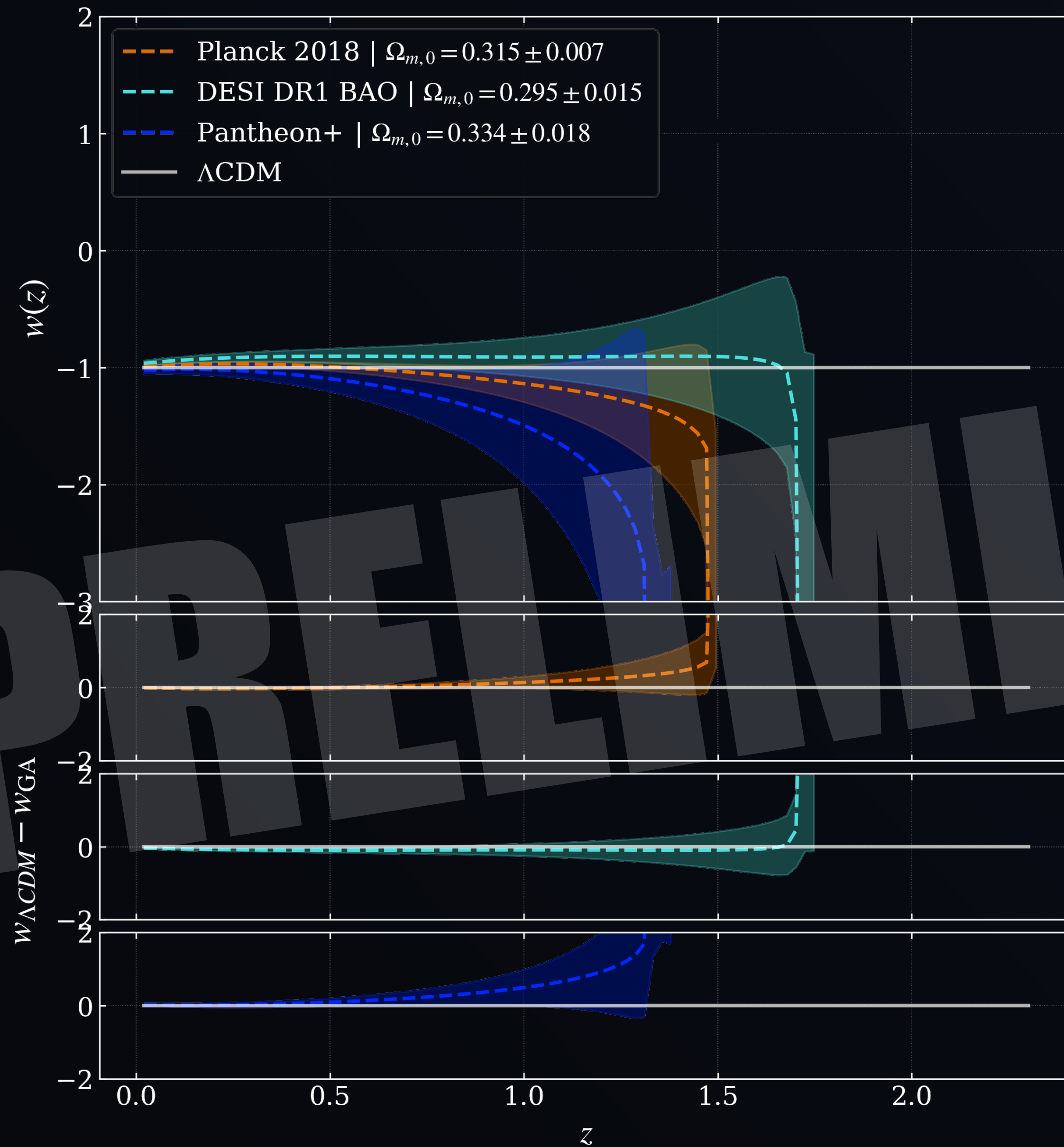
# Reconstruction of $\mu(z)$ from SNe Ia



# Reconstruction of $H(z)$ from SNe Ia



# Reconstruction of $w_{bg}(z)$ from SNe Ia

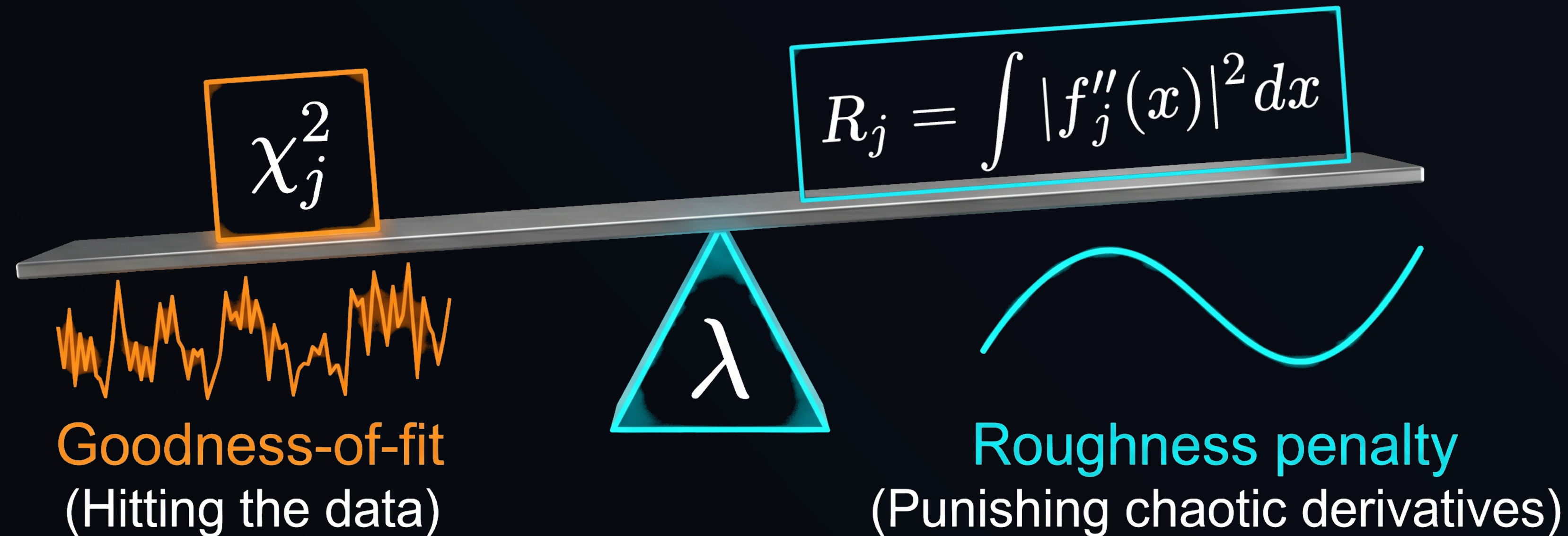


$$w(0.02) = -0.995 \pm 0.024$$

Perfectly compatible  
with  $\Lambda$ CDM using  
current SNe Ia data

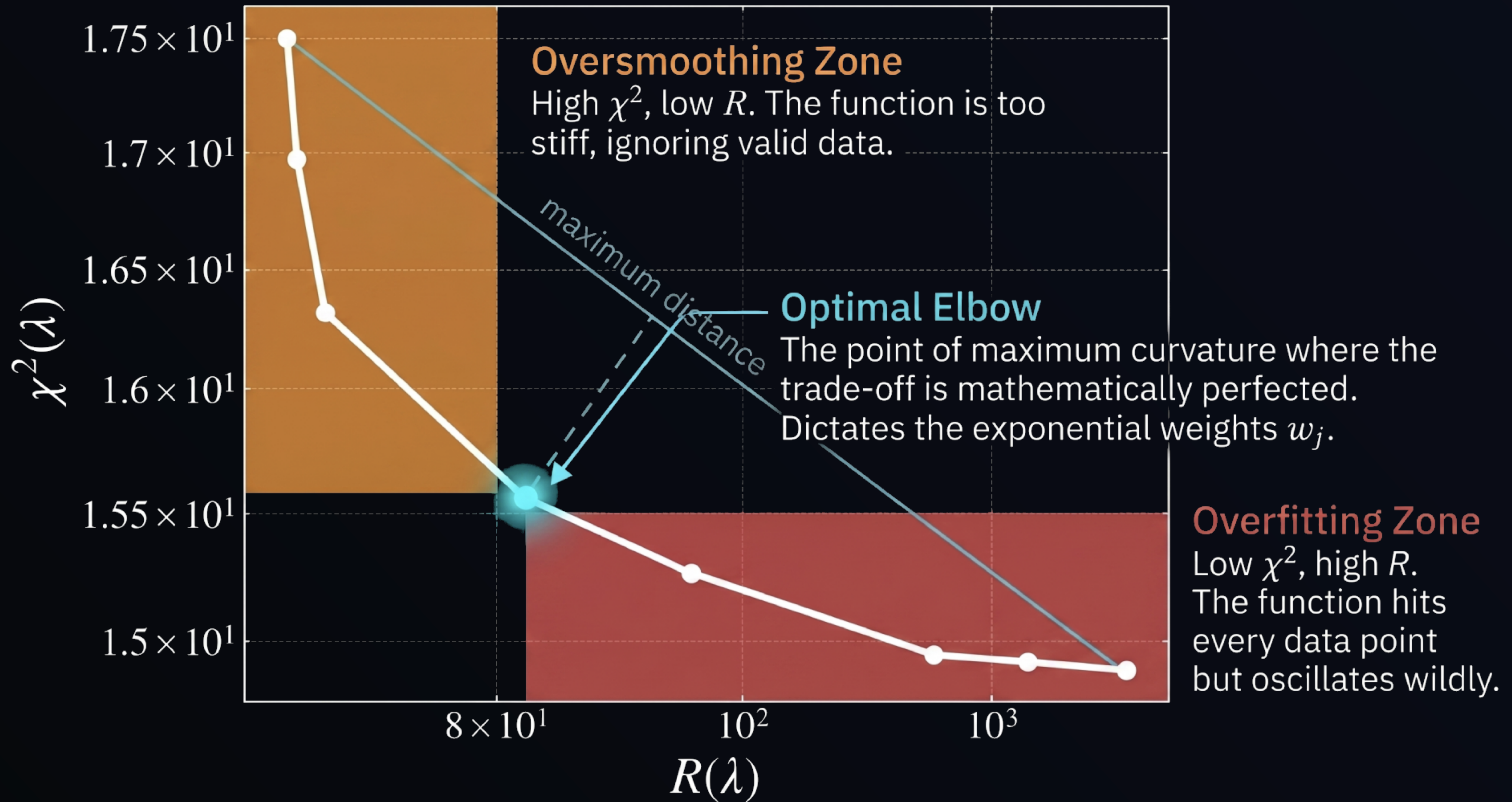
# Elbow point estimation for $\lambda$

$$S_j = \chi_j^2 + \lambda R_j$$

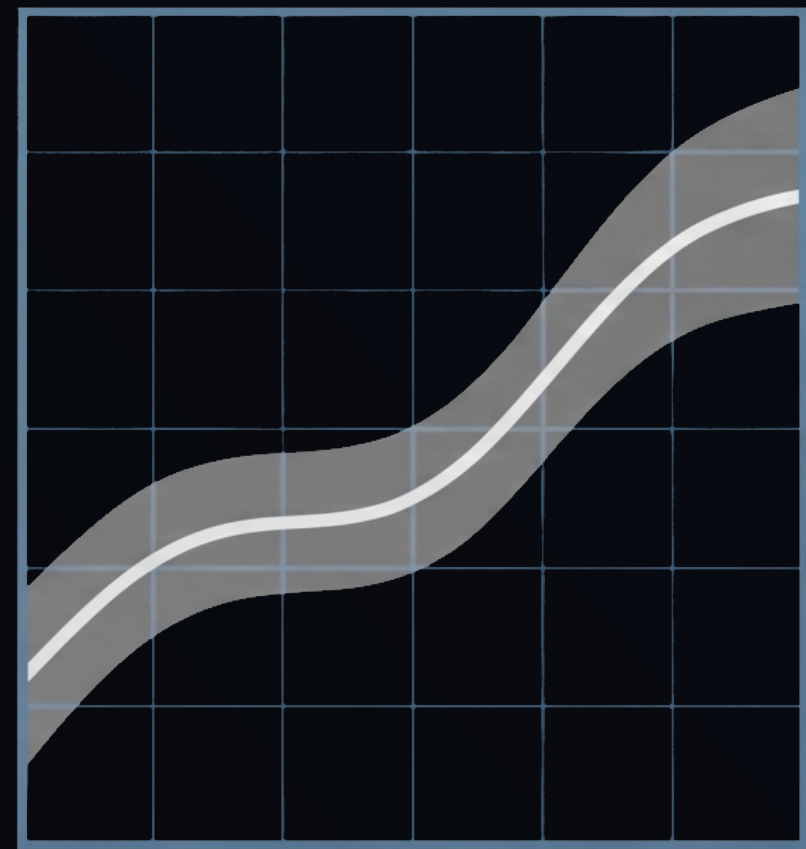


By replacing pure  $\chi^2$  with  $S_j$ , the algorithm stops chasing random noise and starts hunting for the true, smooth underlying physical model.

# Elbow point estimation for $\lambda$

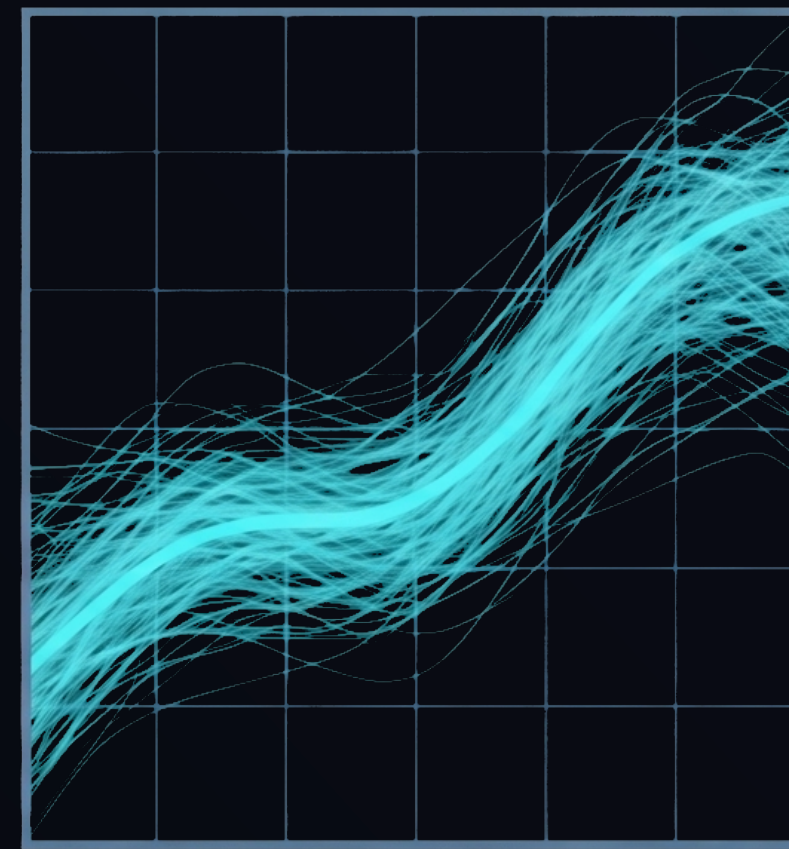


# Uncertainties estimation



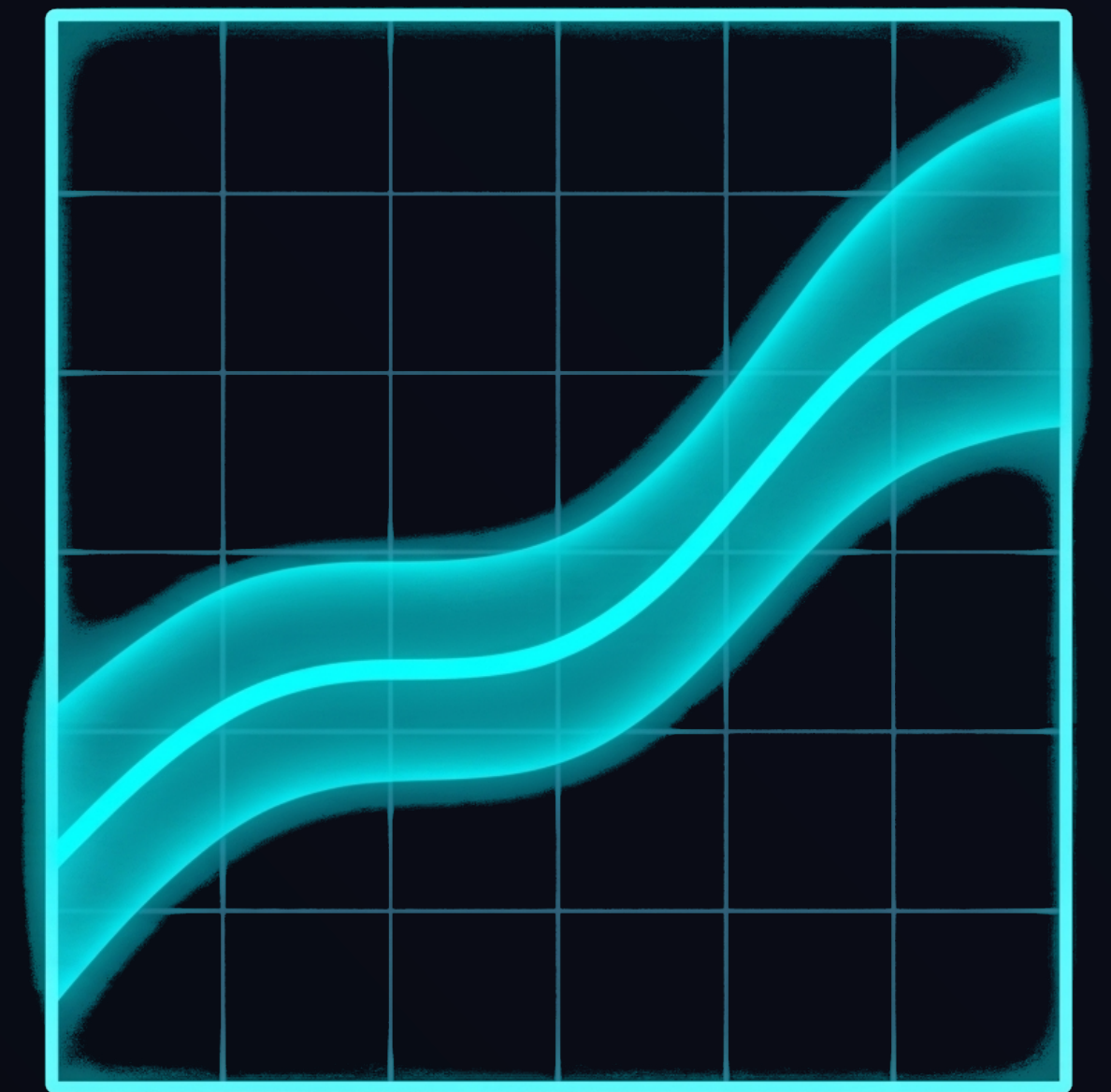
$$\delta f_{PI}$$

**Statistical Uncertainty.**  
Computed via Path-Integral,  
representing standard  
measurement noise.



$$\sigma_{ens}$$

**Configuration Uncertainty.**  
The weighted variance across  
the algorithmic ensemble.



$$\sigma_{tot}$$

**Total GAME Confidence.**

$$\sigma_{tot}^2 = \delta f_{PI}^2 + \sigma_{ens}^2$$