

# $0\nu\beta\beta$ : Nuclear Physics and BSM measurements with $2\nu\beta\beta$

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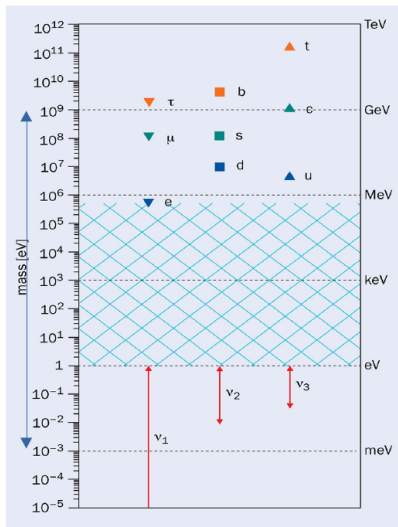


The Standard Model can not explain:

- What is **dark matter**?
  - What is **dark energy**?
  - Prevalence of **matter** over **antimatter**
- 
- The original SM that we all learned in school predicts that neutrinos are strictly massless
  - **We know that neutrinos have mass** from neutrino flavor oscillations

# How are $\nu$ masses generated?

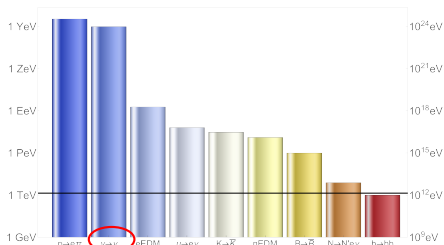
12 orders of magnitude



- Dirac neutrino masses: the neutrino mass is generated via the standard Higgs mechanism in the same way as leptons and quarks  $\rightarrow \nu$ 's Yukawa couplings more than 6 orders of magnitude smaller.
- **Majorana mass term**

# Neutrinos : Break on through to the other side

- **Effective Field Theory**: the Standard Model  $\mathcal{L}_{SM} = \mathcal{L}_{D=4}$ . The Lagrangian is expanded in powers of  $1/\Lambda$ .
- The first correction has a single dimension 5 operator, which gives rise to a Majorana mass term :  $\mathcal{L} = -\frac{1}{2} \sum_i m_i \bar{\nu}_i \nu_i \propto \frac{1}{\Lambda}$
- **New physics scale  $\Lambda$**  explored by  $0\nu\beta\beta$  searches with high sensitivity



A. Falkowski lecture at CP2023 (<https://indico.in2p3.fr/event/27893>)

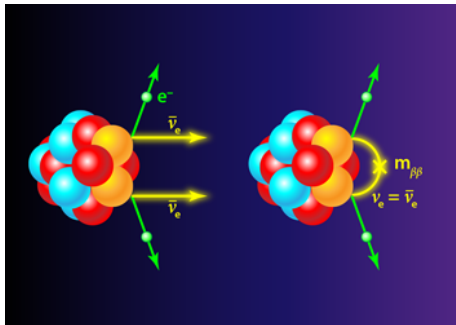
- The smallness of neutrino masses comes up naturally for  $\Lambda$  of the order of the GUT scale (see-saw model)
- Baryon-antibaryon asymmetry explained through leptogenesis

$2\nu\beta\beta$

Lepton number is conserved

$0\nu\beta\beta$

Lepton number not conserved,  
creation of matter



Engels and Volker, Physics 11, 30, 2018

# $0\nu\beta\beta$ decay half-life

Experimental observable:  $0\nu\beta\beta$  decay rate :  $\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1}$

$0\nu\beta\beta$  can happen by several mechanisms. Assuming the *light-neutrino exchange scenario*:

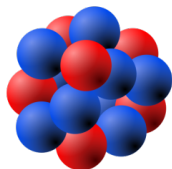
$$\left(T_{1/2}^{0\nu\beta\beta}\right)^{-1} = g_A^4 \cdot \underbrace{G_{0\nu} \cdot |M_{0\nu}|^2}_{\text{nuclear}} \cdot \overbrace{\langle m_{\beta\beta} \rangle^2 / m_e^2}^{\text{new physics}}$$

We measure  $T_{1/2}^{0\nu\beta\beta}$  and interpret it in terms of  $m_{\beta\beta}$  with inputs from theory:

- $g_A$  is the coupling to the nucleon (axial vector coupling constant). For a bare nucleon  $g_A \approx 1.27$ .
- $G_{0\nu}$  is the Phase-space factor represents the kinematic of the decay. Depends on  $Q_{\beta\beta}$ ,  $Z$ , electrons. **Calculated with good accuracy.**

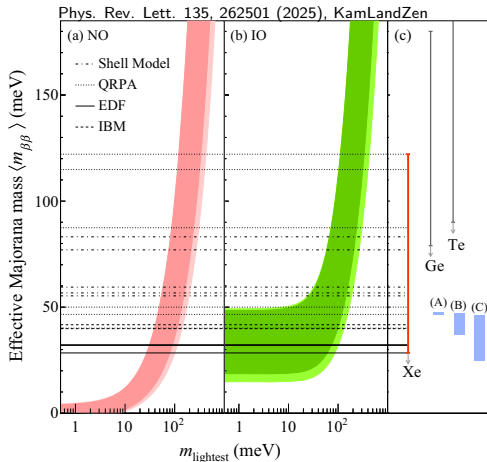
# Nuclear Matrix elements

- $M_{0\nu}$ , the nuclear matrix element, NME, relates to the nuclear structure of the initial and final state nuclei  $\rightarrow$  we need to consider a many body problem  $\rightarrow$  **the hard part**



- Difficult to calculate accurately. Differences  $\sim$  factor 3 - 4 depending on the nuclear models

# Effective Majorana mass

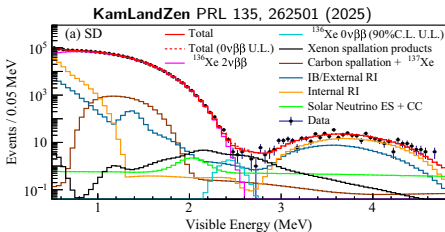
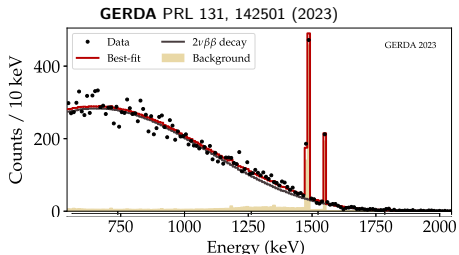
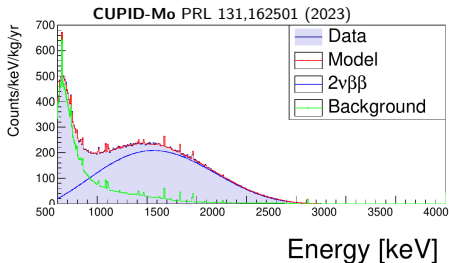
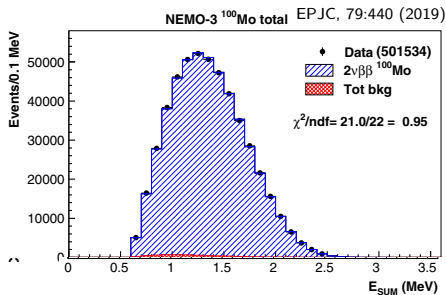


$$m_{\beta\beta} = \left| \sum_{i=1}^3 |U_{ei}^2| e^{i\phi_i} m_i \right|$$

- NME theoretical uncertainties introduce large uncertainties in the interpretation of  $T_{1/2}^{0\nu}$  in terms of  $m_{\beta\beta}$
- Prevent to determine accurately the expected sensitivities of future experiments

**Experimental inputs to constrain nuclear models are critical to assess  $m_{\beta\beta}$**

# $2\nu\beta\beta$ becomes the signal



## $2\nu\beta\beta$ rate and the quenching issue

$$\Gamma^{2\nu} = \frac{1}{T_{1/2}^{2\nu}} \simeq G_{2\nu} \cdot \underbrace{|\mathcal{M}_{2\nu}|^2}_{M_{2\nu} \cdot g_{2\nu}^2}$$

- Single  $\beta$  and  $2\nu\beta\beta$  theoretical rates systematically overpredict experimental rates  $\rightarrow \mathcal{M}_{2\nu}$  is overestimated
- Quenching of  $g_A$  ( $g_{2\nu}$ ) ?
- The situation is more clear today:
- *Ab initio* or first principles nuclear structure methods, exploiting increasing computational performances, could describe  $\beta$  decay half-lives of light nuclei without any adjustments  $\Rightarrow$
- There is a deficiency in the other many-body approaches due to approximations
- So far limited benchmarking with experimental data

$$\Gamma^{2\nu} \simeq G_0 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4$$

$$\Gamma^{2\nu} \simeq G_0 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 + \xi_{31} G_2 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4$$

- Taylor expansion in the lepton energies. [Phys. Rev. C 97, 034315 (2018)]
- $\xi_{31} = \frac{M_{GT-3}}{M_{GT}}$
- subleading  $2\nu\beta\beta$  matrix elements

$$\Gamma^{2\nu} \simeq G_0 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 + \xi_{31} G_2 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 + \frac{1}{3} \xi_{31}^2 G_{22} |M_{GT}^{2\nu}|^2 (g_A^{eff})^4$$

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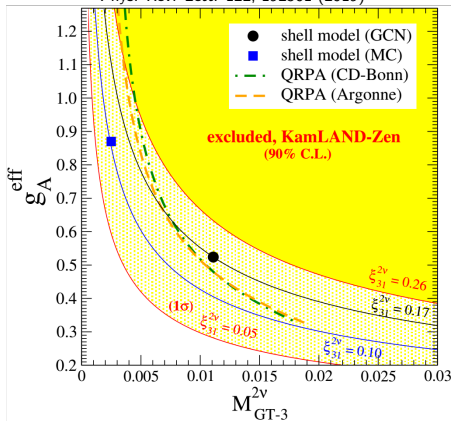
$$\begin{aligned}\Gamma^{2\nu} \simeq & G_0 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 + \xi_{31} G_2 \cdot |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 \\ & + \frac{1}{3} \xi_{31}^2 G_{22} |M_{GT}^{2\nu}|^2 (g_A^{eff})^4 \\ & + \left( \frac{1}{3} \xi_{31}^2 + \xi_{51} \right) G_4 |M_{GT}^{2\nu}|^2 (g_A^{eff})^4\end{aligned}$$

- Taylor expansion in the lepton energies. [Phys. Rev. C 97, 034315 (2018)]
- $\xi_{31} = \frac{M_{GT-3}}{M_{GT}}$
- $\xi_{51} = \frac{M_{GT-5}}{M_{GT}}$
- subleading  $2\nu\beta\beta$  matrix elements

# Experimental inputs to nuclear models : KamLand Zen

The shape of the  $2\nu\beta\beta$  spectrum constrains  $\xi_{31} = \frac{M_{GT-3}}{M_{GT}}$

Phys. Rev. Lett. 122, 192501 (2019)



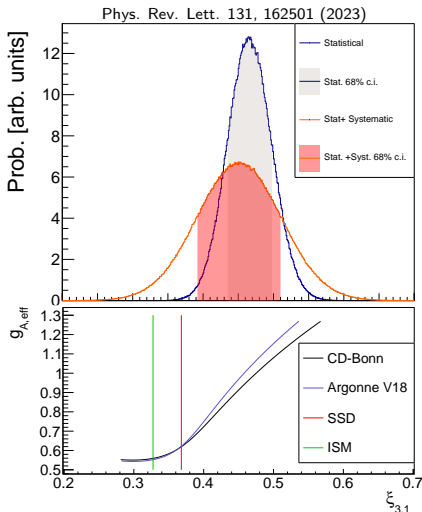
$$\xi_{31} < 0.26 \text{ at } 90 \% \text{ C.L.}$$

- Shell model predictions well consistent with 90% upper limit
- QRPA predictions consistent with upper limit, small region in  $g^{eff}$  is excluded

# Experimental inputs to nuclear models : Cupid-Mo

$$\xi_{3,1} = 0.45 \pm 0.03(\text{stat.}) \pm 0.05(\text{syst.})$$

- Under HSD  $\xi_{3,1} = 0$ .  
Disfavoured by  $8 \sigma$
- Under SSD  $\xi_{3,1} = 0.368$ .  
Compatible at  $1.4 \sigma$
- Shell Model mildly incompatible ( $\sim 2.1 \sigma$ )
- pn-QRPA compatible if the value of  $g_A^{\text{eff}}$  is moderately quenched ( $> 0.8$ ) or unquenched.

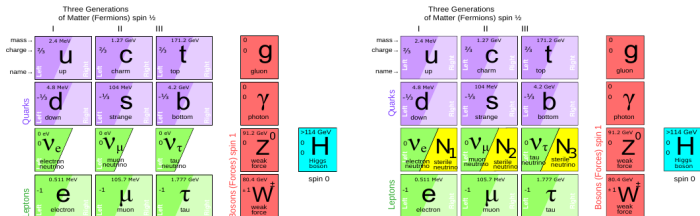


First **measurement** of  $\xi_{3,1}$  and  $g_{A,\text{eff}}$  based on  $2\nu\beta\beta$  spectral shape study

# Beyond Standard Model measurements with $2\nu\beta\beta$

# Sterile neutrinos

- Neutrinos in the SM are massless **left-handed** fermions
- Simple extension of the SM: add spin 1/2 **right-handed** neutrinos, or sterile neutrinos, with mass  $M_{\text{heavy}}$



M. Shaposhnikov, J.Phys.Conf.Ser. 408 (2013) 012015

- Type-I seesaw mechanism can simultaneously explain
  - Smallness of the SM neutrino mass
  - baryogenesis via leptogenesis
  - dark matter if one RH neutrino has a mass in the keV range

(for ex: Phys. Rev. D 100, 075029 (2019), Int. J. Mod. Phys. E 22, 1330019 (2013))

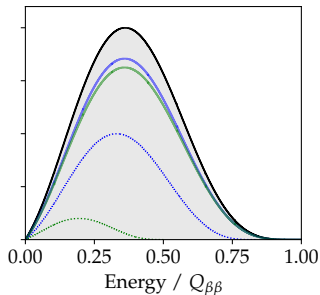


# Sterile neutrinos and $2\nu\beta\beta$

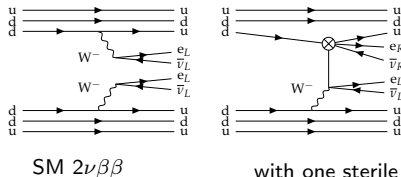
- Sterile neutrinos can be produced in any decay involving  $\nu$ 's
- Kinematically possible in  $2\nu\beta\beta$  if sterile neutrino mass  $M_N < Q_{\beta\beta}$
- It would distort the  $2\nu\beta\beta$  shape

.....  $\nu N\beta\beta$  ( $m_N = 0.5$  MeV)

.....  $\nu N\beta\beta$  ( $m_N = 1.5$  MeV)



dotted: pure  $\nu N\beta\beta$  decay, solid:  $\sin^2\theta = 0.1$



$$\frac{d\Gamma^{tot}}{dE} = \cos^4\theta \frac{d\Gamma_{2\nu, SM}}{dE} + 2\cos^2\theta\sin^2\theta \frac{d\Gamma_{\nu N}}{dE}$$

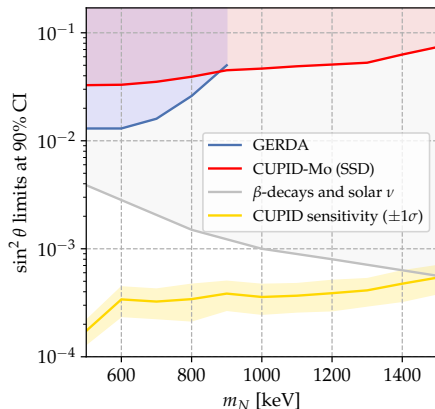
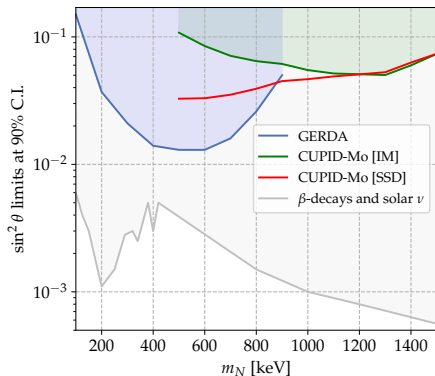
$\sin\theta \rightarrow$  active-sterile mixing

$$\sin\theta \simeq \frac{m_D}{M} \quad \mathbf{M} = \begin{vmatrix} \frac{m_D^2}{M} & 0 \\ 0 & M \end{vmatrix}$$

$$\frac{m_D^2}{M} = m_{light}, \quad M = M_{heavy}$$

# Sterile neutrinos limits from $2\nu\beta\beta$ in Cupid-Mo and GERDA

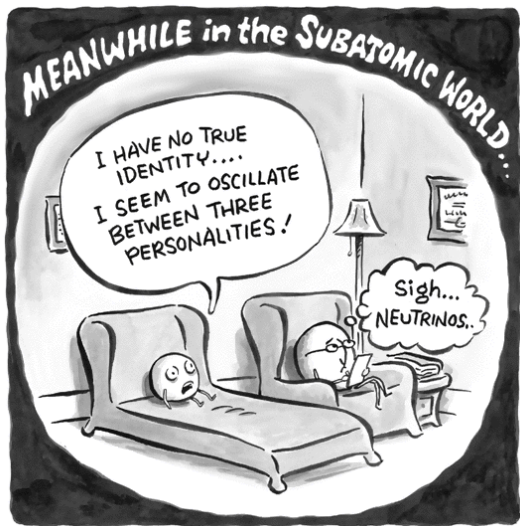
- $2\nu\beta\beta$  can explore the region in the sub-GeV range
- BSM results are affected by the standard  $2\nu\beta\beta$  spectrum we use.



- The uncertainty in the nuclear matrix elements introduces large uncertainties in  $T_{1/2}^{0\nu\beta\beta}$  interpretation in terms of  $m_{\beta\beta}$ .
- On the other hand,  $\mathcal{M}_{2\nu}$  is overestimated ("quenching problem"). *Ab initio* methods provides proper answer, still:

→ Need for experimental inputs to constrain or tune nuclear models

- Thanks to large statistics of  $2\nu\beta\beta$  collected by  $0\nu\beta\beta$  experiments in the last 10 years, we can use  $2\nu\beta\beta$  spectrum to give inputs to nuclear models.
- $2\nu\beta\beta$  spectral shape also can be used to put limits on beyond standard model processes, like sterile neutrinos.
  - Potential for competitive results from next generation  $0\nu\beta\beta$  experiments



# The see-saw mechanism

## The See-Saw mechanism

After diagonalization:  $\mathbf{M}_{diag} = \begin{pmatrix} m_{light} & 0 \\ 0 & M_{heavy} \end{pmatrix}$

Eigenvalues ("Physical" masses):  $\begin{cases} m_{light} \approx \frac{m_D^2}{M} \\ M_{heavy} \approx M \end{cases}$



By plugging in a **top-quark-like** Dirac mass (**100 GeV**) and a GUT-scale ( **$10^{14} - 10^{16}$  GeV**) Majorana mass into the **seesaw formula**, we obtain light neutrinos in the **sub-eV range**, consistent with experimental data!

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Eigenvectors ("Physical" neutrino states):

|       |   |   |   |   |
|-------|---|---|---|---|
| Light | → | { | $\Psi_1 \approx \underbrace{(\nu_L + \nu_L^c)}_{Active} - \underbrace{\frac{m_D}{M} (\nu_R^c + \nu_R)}_{Sterile}$ | Familiar Standard Model<br>light neutrinos    |
| Heavy | → | { | $\Psi_2 \approx \underbrace{(\nu_R^c + \nu_R)}_{Sterile} + \underbrace{\frac{m_D}{M} (\nu_L + \nu_L^c)}_{Active}$ | Exotic never-observed<br>very heavy neutrinos |

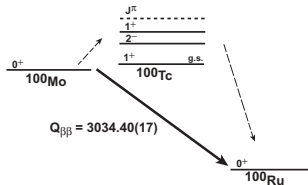
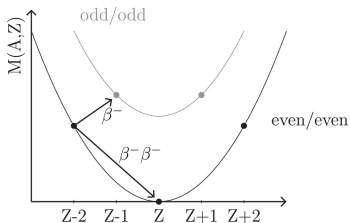
**All physical neutrino states (light and heavy) are Majorana particles!**

Neutrinos and antineutrinos are just different spin states of the same fundamental particle

$$(\nu_L + \nu_L^c)^c = \nu_L^c + \nu_L$$

# $2\nu\beta\beta$ decay

- $(A, Z) \rightarrow (A, Z+2) + 2e^- + 2\bar{\nu}_e$
- $2\nu\beta\beta$  is observable in nuclei where single  $\beta$  decay is forbidden energetically due to nuclear pairing interaction

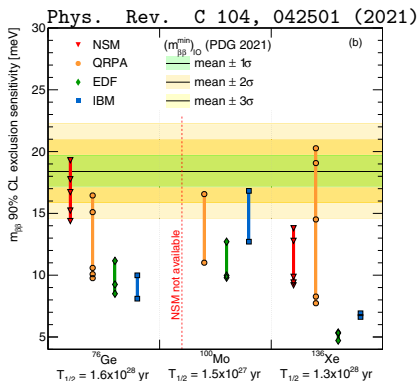


- Directly observed, with  $T_{1/2}^{2\nu\beta\beta}$  between  $10^{19}$  -  $10^{21}$  years

|                   |               |                   |
|-------------------|---------------|-------------------|
| $^{48}\text{Ca}$  | $\rightarrow$ | $^{48}\text{Ti}$  |
| $^{76}\text{Ge}$  | $\rightarrow$ | $^{76}\text{Se}$  |
| $^{82}\text{Se}$  | $\rightarrow$ | $^{82}\text{Kr}$  |
| $^{96}\text{Zr}$  | $\rightarrow$ | $^{96}\text{Mo}$  |
| $^{100}\text{Mo}$ | $\rightarrow$ | $^{100}\text{Ru}$ |
| $^{116}\text{Cd}$ | $\rightarrow$ | $^{116}\text{Sn}$ |
| $^{130}\text{Te}$ | $\rightarrow$ | $^{130}\text{Xe}$ |
| $^{136}\text{Xe}$ | $\rightarrow$ | $^{136}\text{Ba}$ |
| $^{150}\text{Nd}$ | $\rightarrow$ | $^{150}\text{Sm}$ |

# Uncertainty in $0\nu\beta\beta$ sensitivity comes from NME's

- $M_{0\nu}$  relates to the nuclear structure of the initial and final state nuclei  
→ we need to consider an N-body problem
- Difficult to calculate accurately. Differences  $\sim$  factor 3 depending on the nuclear models

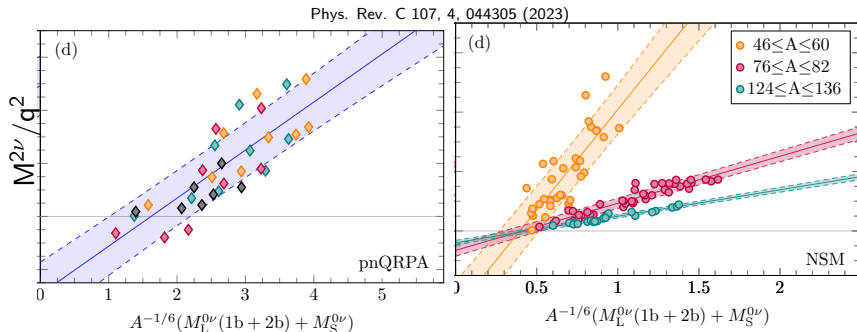


- NME theoretical uncertainties introduce large uncertainties in the interpretation of  $T_{1/2}^{0\nu}$  in terms of  $m_{\beta\beta}$
- Prevent to determine accurately the expected sensitivities of future experiments

Experimental inputs to constrain nuclear models are critical to

# What can we learn from $2\nu\beta\beta$ ?

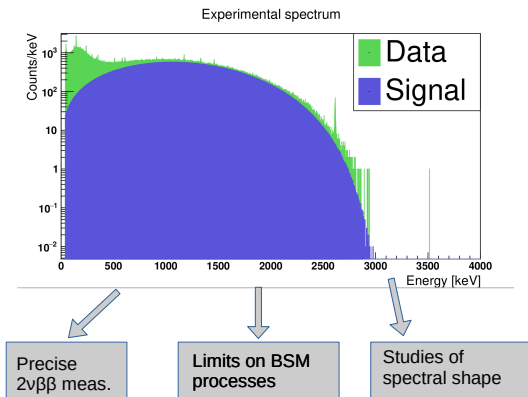
$0\nu\beta\beta$  and  $2\nu\beta\beta$  involve same initial and final states, but permitted momentum transfers are larger in  $0\nu\beta\beta$



We can use  $2\nu\beta\beta$  data to constraint  $0\nu\beta\beta$  NME's!

# $2\nu\beta\beta$ becomes the signal

- Starting in  $\sim 2015$   $0\nu\beta\beta$  experiments exploiting higher exposures, lower backgrounds and improvements in the measurement techniques start to measure  $2\nu\beta\beta$  spectra with high precision.



# Theoretical issues : $g_A$ quenching?

$2\nu\beta\beta$  rate:

$$\left(T_{1/2}^{2\nu}\right)^{-1} = G_{2\nu} \cdot \underbrace{|M_{2\nu}|^2 \cdot g_{2\nu}^4}$$

- **Single  $\beta$**  and  $2\nu\beta\beta$  theoretical rates overpredict experimental rates  $\rightarrow$   $|M_{2\nu}|^2 \cdot g_{2\nu}^4$  overestimated

What is the impact for  $0\nu\beta\beta$  decay?

- Quenching of  $g_A$  w.r.t bare nucleon = 1.27?
- Too large NME's?
- Both are related, we should consider  $g_A \times$  NME



$g_{2\nu} \neq g_{0\nu}$

- Possible interpretation: quenching of  $g_A$
- **Generally speaking, uncertainties in nuclear models** introduce large uncertainties in  $0\nu\beta\beta$  predictions



$g_{2\nu} \neq g_{0\nu}$

$2\nu\beta\beta$  spectrum is useful to extract observables to constrain nuclear models

## $2\nu\beta\beta$ theoretical description

Here I would like to start with the general expression :

$$\frac{d\Gamma}{dE_1 dE_2} \propto F_0(Z, E_1) F_0(Z, E_2) p_1 p_2 \int_0^{E_i - E_f - E_1 - E_2} E_{\nu_1}^2 E_{\nu_2}^2 A^{2\nu} dE_{\nu_1}, \quad (1)$$

where  $F_0(Z, E_1)$  is the Fermi function,  $E_1, E_2, p_1, p_2$  are the two electron energies and momenta,  $E_{\nu_2}, E_{\nu_1}$  are the two neutrino energies and  $E_i, E_f$  the energies of the initial and final states.  $A^{2\nu}$  involves matrix elements which depend on the lepton energies and on the sum over possible states of the intermediate nucleus:

$$A^{2\nu} = \frac{1}{4} |M_{GT}^K + M_{GT}^L|^2 + \frac{1}{12} |M_{GT}^K - M_{GT}^L|^2. \quad (2)$$

Here  $M_{GT}^{K,L}$  are the Gamow Teller matrix elements given by:

$$M_{GT}^{K,L} = m_e \sum_n \langle 0_F^+ | |\tau_m^+ \sigma_m^+| | 1_n^+ \rangle | \langle 1_n^+ | |\tau_m^+ \sigma_m^+| | 0_I^+ \rangle \frac{C_n}{C_n^2 - \epsilon_{K,L}^2}, \quad (3)$$

where the sum goes over all possible states of the intermediate nucleus.

## $2\nu\beta\beta$ theoretical description

In full generality,

$$\frac{d\Gamma^{2\nu}}{dE_1 dE_2} \propto \int_0^{E_i - E_f - E_1 - E_2} E_{\nu_1}^2 E_{\nu_2}^2 A^{2\nu} dE_{\nu_1}$$

$A^{2\nu}$  is a sum of the matrix elements, relating Final and Initial nuclear states and all possible states of the intermediate nucleus:

$$M_{GT}^{K,L} = m_e \sum_n \langle Final | |\tau_m^+ \vec{\sigma}_m| |1_n^+ \rangle | \langle 1_n^+ | |\tau_m^+ \vec{\sigma}_m| | Initial \rangle \frac{C_n}{C_n^2 - \epsilon_{K,L}^2} \quad (6)$$

And  $C_n = E_n - (E_i + E_f)/2$ .

This is a problem. A factorization is possible under some approximations:

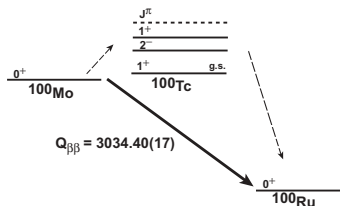
- SSD, Single State Dominance: decay goes through the lowest  $1^+$  state of the intermediate nucleus.
- HSD, High State Dominance: lepton energies are neglected w.r.t energies of intermediate states

$$\frac{d\Gamma}{dE_1 dE_2} \propto \overbrace{|M_{GT}^{2\nu}|^2}_{\text{matrix element}} \int_0^{E_i - E_f - E_1 - E_2} E_{\nu_1}^2 E_{\nu_2}^2 dE_{\nu_1} \quad (7)$$

## $2\nu\beta\beta$ theoretical description

$$\left(T_{1/2}^{2\nu}\right)^{-1} = G_{2\nu} \cdot |M_{2\nu}|^2 \cdot g_{2\nu}^4$$

- This factorization is possible only under some approximations:
  - SSD, Single State Dominance: decay goes through the lowest  $1^+$  state of  $^{100}\text{Tc}$
  - HSD, High State Dominance: lepton energies are neglected w.r.t energies of intermediate states of  $^{100}\text{Tc}$  nucleus
  - The choice affects the theoretical spectral shapes



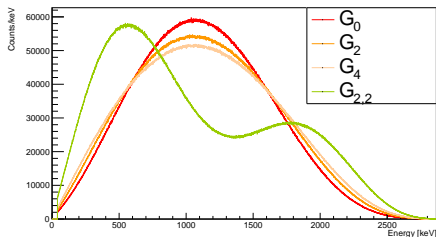
## $2\nu\beta\beta$ improved model

- Improved model from Šimkovic *et al* (PRC97 (2018) 034515), Taylor expansion in the lepton energies of  $2\nu\beta\beta$  decay rate:

$$\frac{d\Gamma}{dE} = (g_A^{\text{eff}})^4 |M_{GT-1}|^2 \left( \frac{dG_0}{dE} + \xi_{31} \frac{dG_2}{dE} + \frac{1}{3} \xi_{31}^2 \frac{dG_{22}}{dE} + \left( \frac{1}{3} \xi_{31}^2 + \xi_{51} \right) \frac{dG_4}{dE} \right)$$

- $g_A^{\text{eff}}$  effective axial vector coupling constant, tuned to reproduce experimental  $T_{1/2}^{2\nu}$
- $\xi_{31} = \frac{M_{GT-3}}{M_{GT-1}}$ ,  $\xi_{51} = \frac{M_{GT-5}}{M_{GT-1}}$

Our motivation is to extract the values of novel observables to compare to theoretical predictions



$dG_i/dE$  come from theory  
shape (produced by Jenni  
Kotila)

- We fit experimental data, with a Gaussian prior on  $\xi_{51}/\xi_{31}$ , that can be accurately calculated:
  - SSD: 0.367
  - pn-QRPA: 0.364 – 0.368
  - Shell: 0.349

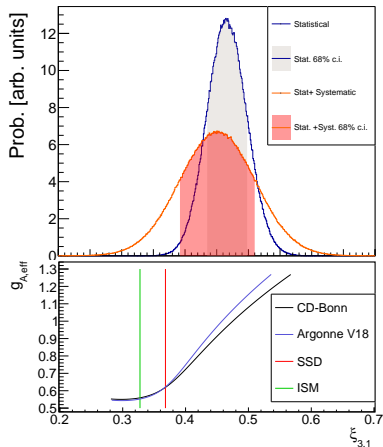
# CUPID-Mo $2\nu\beta\beta$ spectral shape results

Phys. Rev. Lett. 131, 162501 (2023)

$$\xi_{3,1} = 0.45 \pm 0.03(\text{stat.}) \pm 0.05(\text{syst.})$$

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Disfavoured by  $8\sigma$
- Under SSD  $\xi_{3,1} = 0.368$ .  
Compatible at  $1.4\sigma$
- Shell Model mildly incompatible  
( $\sim 2.1\sigma$ )
- Using  $g_{A,\text{eff}}$  vs  $\xi_{3,1}$  (from F. Šimkovic):

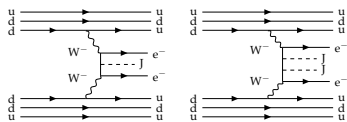
$$g_{A,\text{eff}}(\text{pn-QRPA}) = 1.0 \pm 0.1 (\text{stat.}) \pm 0.2 (\text{syst.})$$



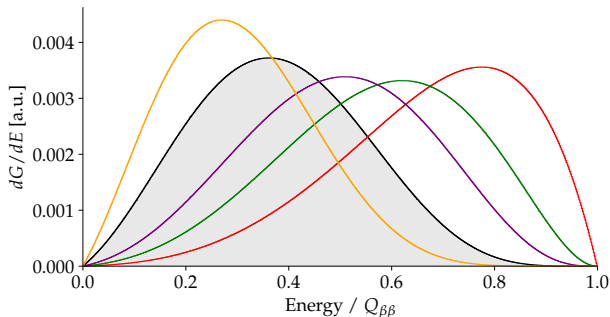
First measurement of  $\xi_{3,1}$  and  $g_{A,\text{eff}}$  based on  $2\nu\beta\beta$  spectral shape study

# Majorons

- $0\nu\beta\beta$  decay in the light neutrino exchange may occur with the emission of one or two Majorons (bosons).
- This would distort the  $2\nu\beta\beta$  spectrum.
- Different Majoron models predict different energy distributions

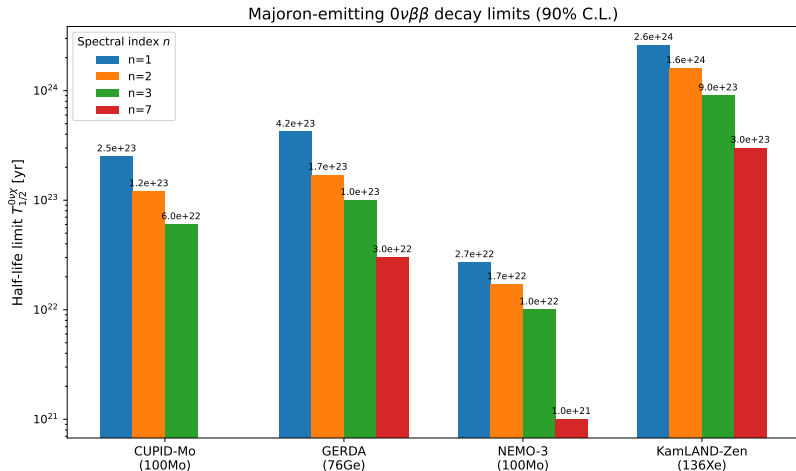


— Standard Model  $2\nu\beta\beta$     
 —  $\beta\beta\chi_0$  ( $n = 2$ )    
 —  $\beta\beta\chi_0\chi_0$  ( $n = 7$ )  
—  $\beta\beta\chi_0$  ( $n = 1$ )    
 —  $\beta\beta\chi_0\chi_0/\beta\beta\chi_0$  ( $n = 3$ )



$$\frac{dG}{dE} \sim (Q_{\beta\beta} - E)^n$$

# Limit to $T_{1/2}$ of $0\nu\beta\beta$ with Majoron emission



- The BSM decay with Majorons is completely unrelated to the SM  $2\nu\beta\beta \rightarrow$  the lower the  $2\nu\beta\beta$  decay rate, the higher the sensitivity ( $^{136}\text{Xe}$ )



