

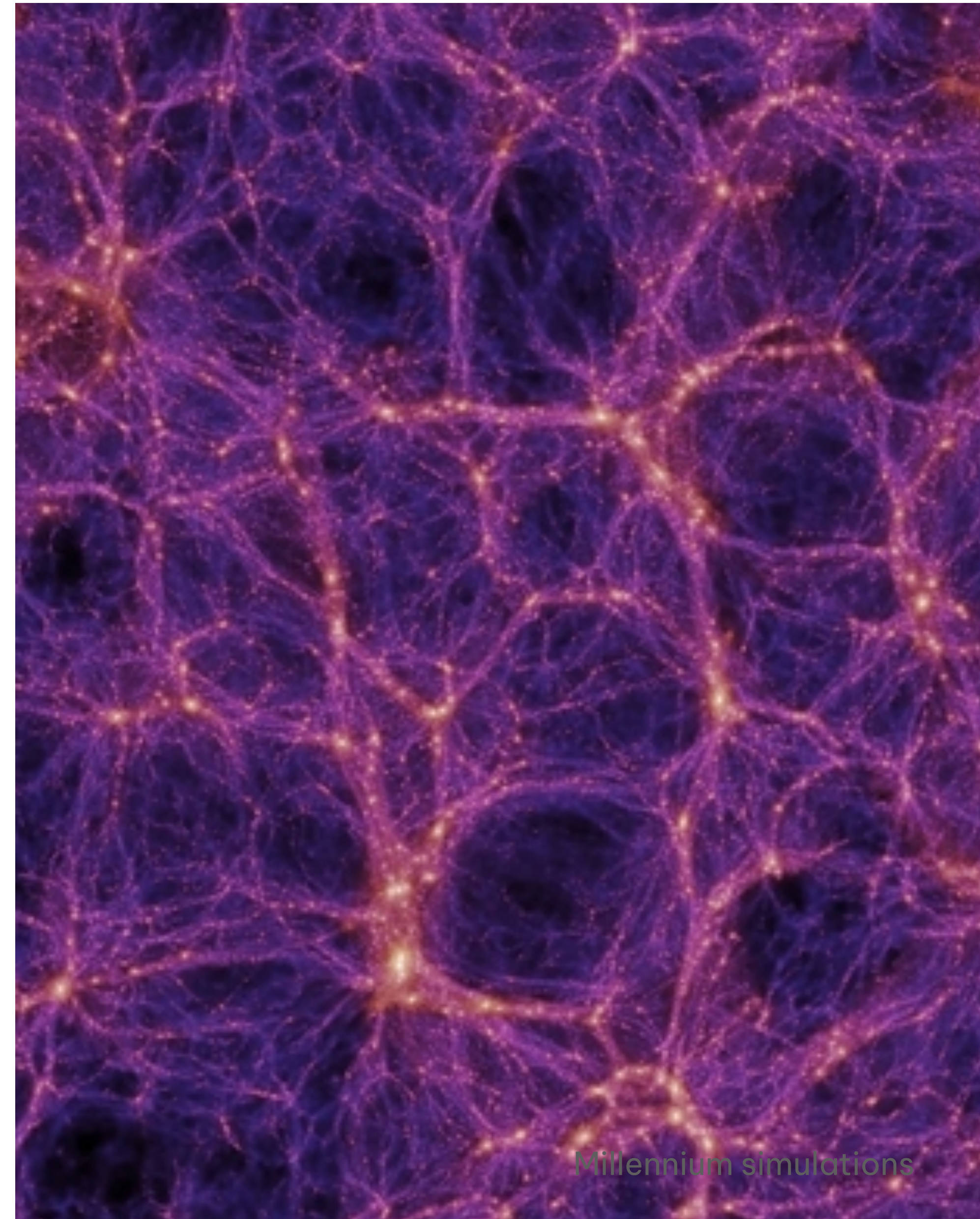
Improving the signal from cosmic voids using reconstruction

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CPPM - Aix Marseille Université



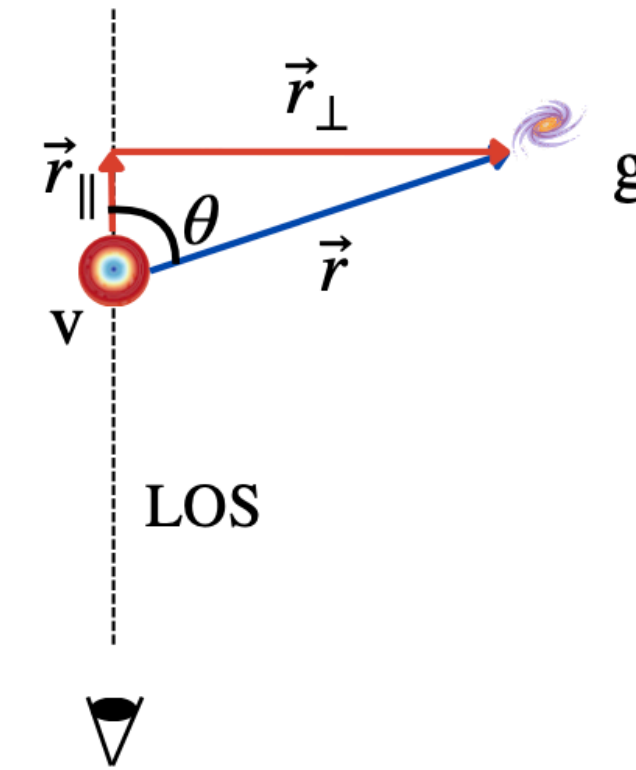
Cosmic voids

- Large underdense structures in the Universe
- Multiscale sensitivity: $10 - 100 h^{-1}\text{Mpc}$
- Interpretable as spatially distinct components of the cosmic web
- Most linear structures in the Universe: simple dynamic
- Dominated by dark energy
- Sensitive to diffuse components
- Sensitive to theories of relativity



Void-galaxy cross-correlation function

- Probability to find a galaxy at a certain distance r and angle $\mu \equiv \cos \theta$ from the center of a void (respect to a random distribution of galaxies)
- Probing the density profile of tracers inside the void region

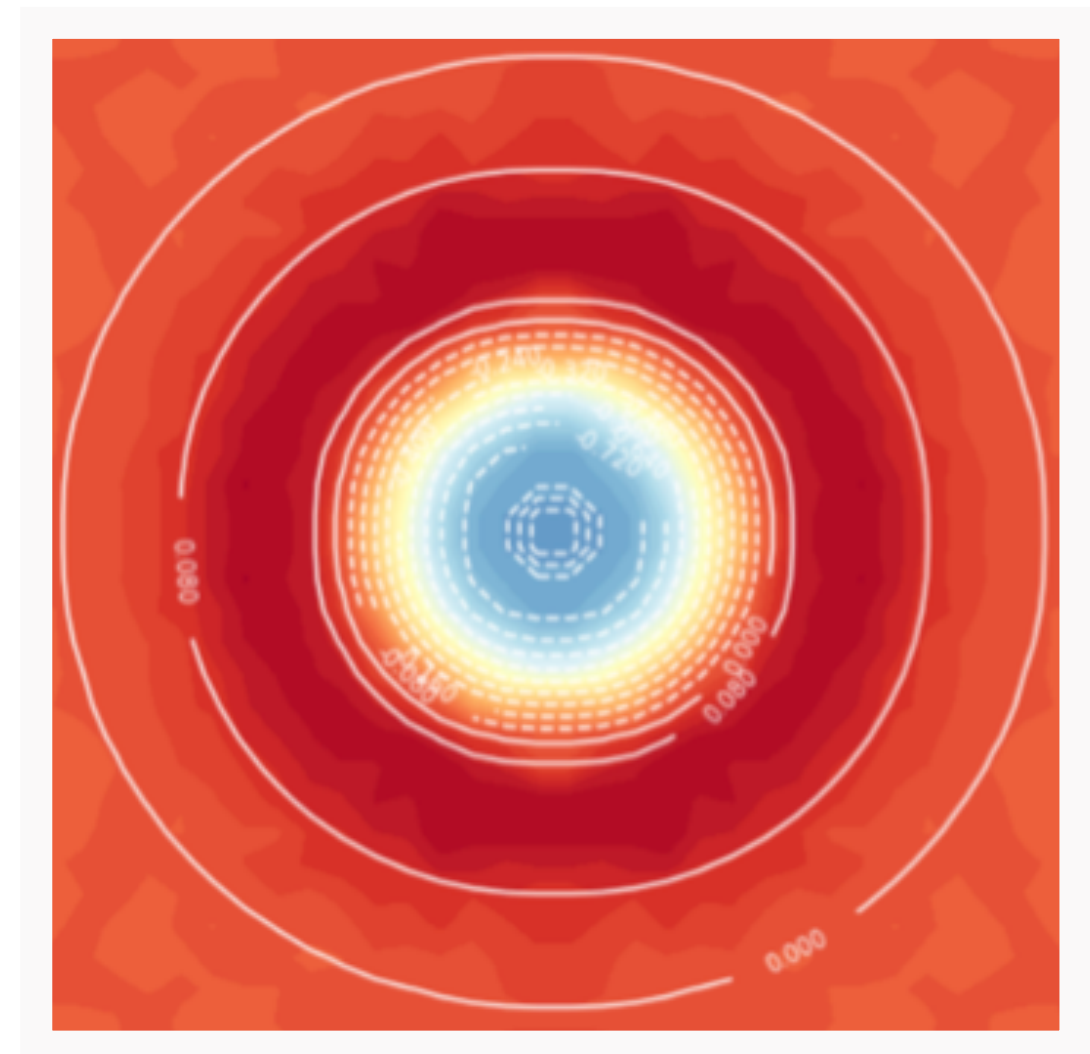


Davis-Peebles estimator:

$$\xi_{vg}^{DP}(r, \mu) = \frac{n_R D_v D_g(r, \mu)}{n D_v R_g(r, \mu)} - 1$$

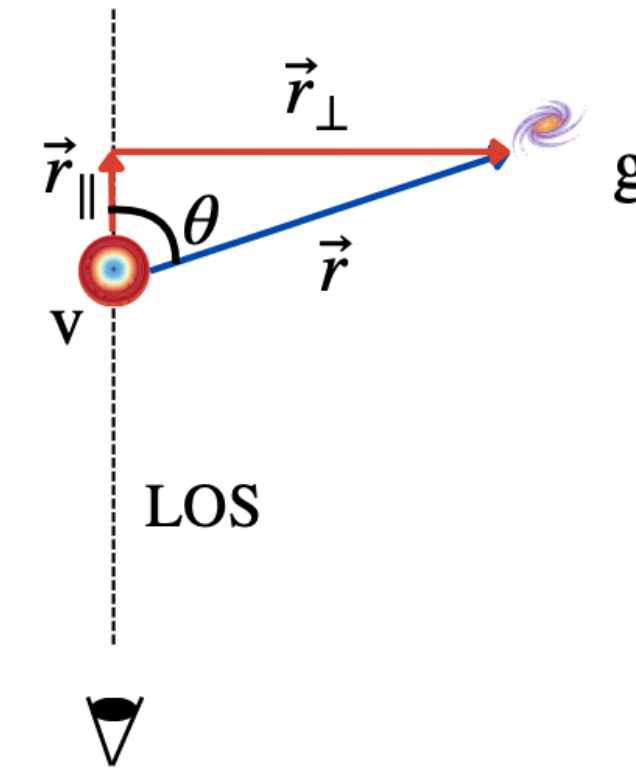
Multipoles:

$$\xi_{\ell}(r) = (2\ell + 1) \int_0^1 \xi(r, \mu) P_{\ell}(\mu) d\mu$$



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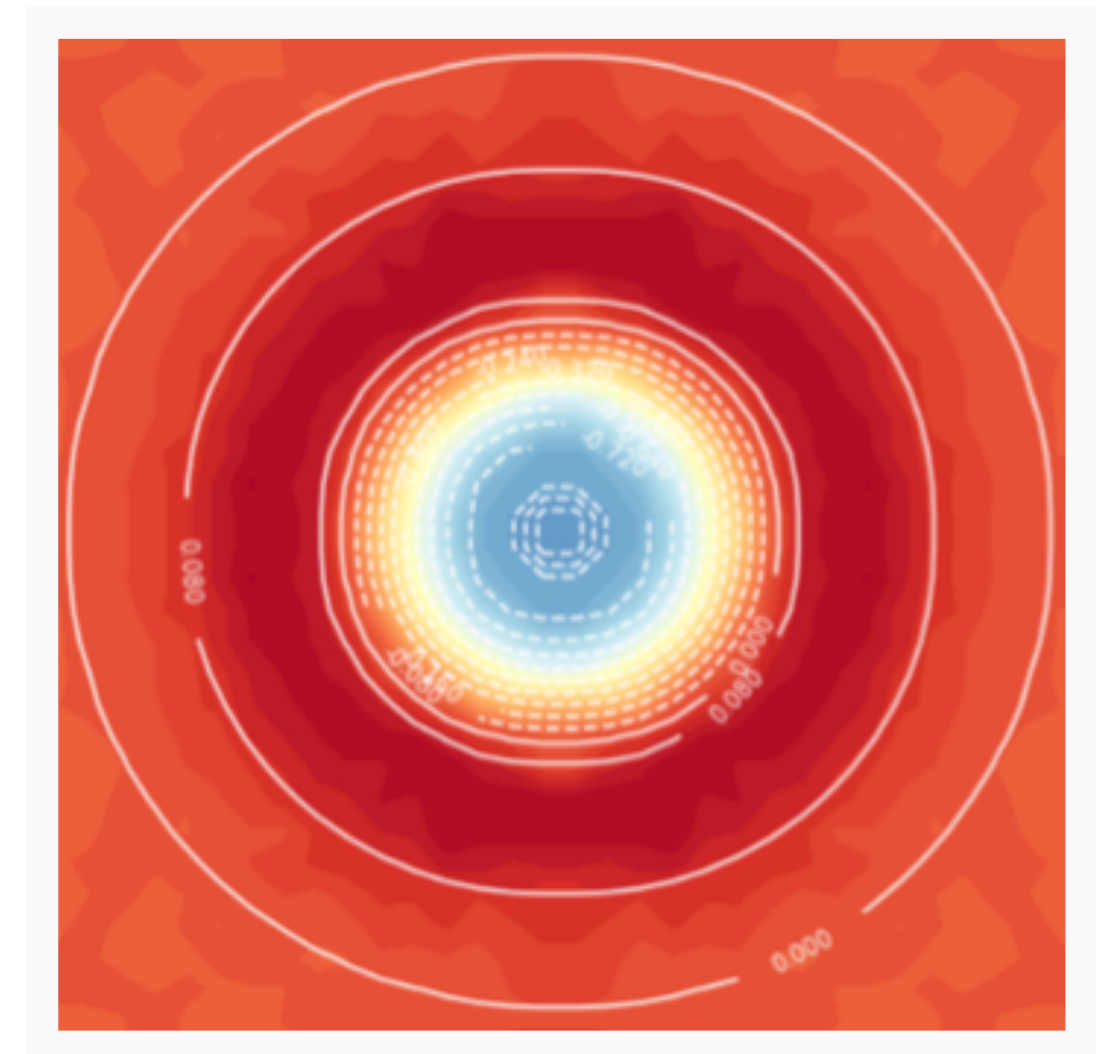


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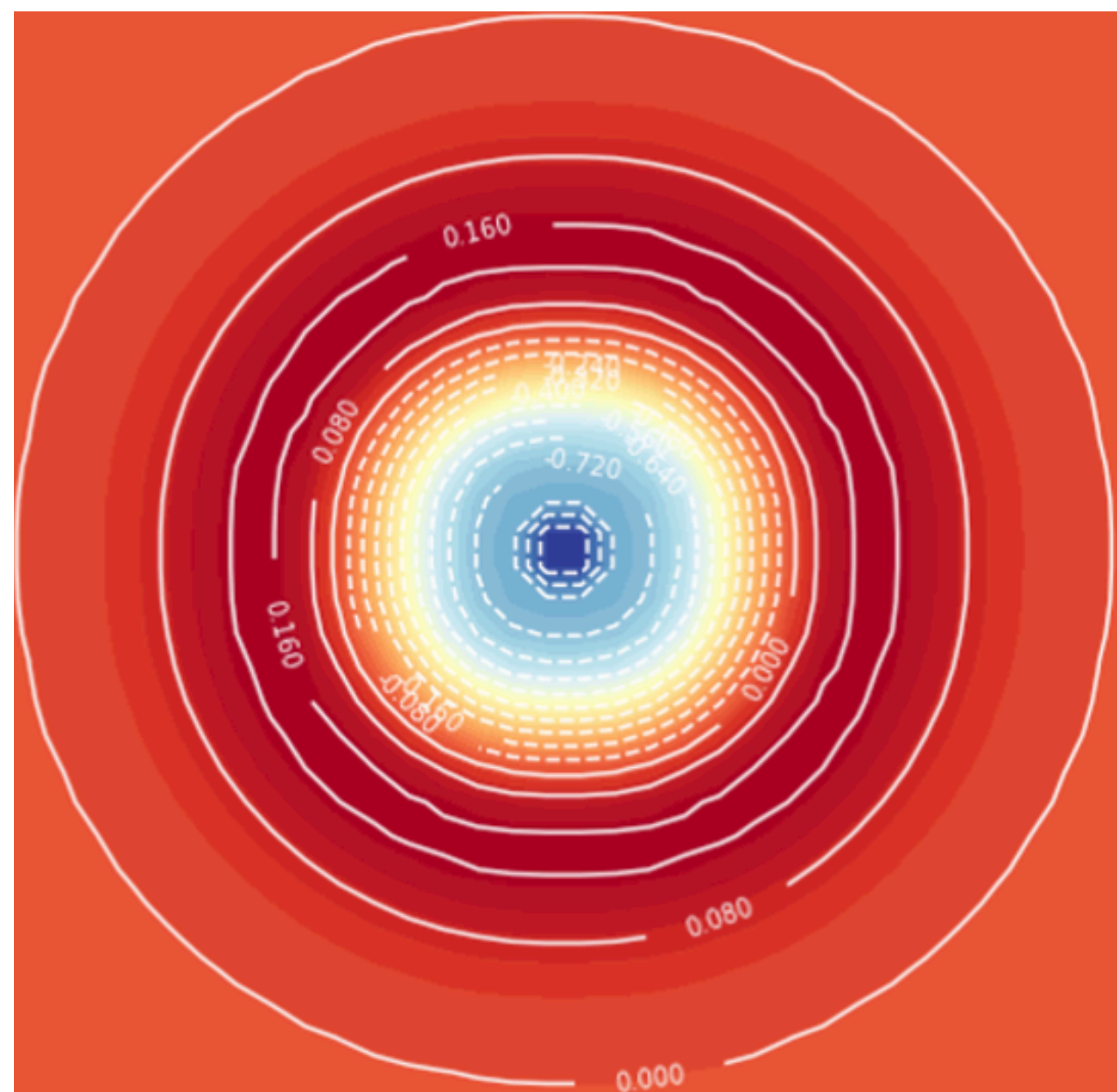
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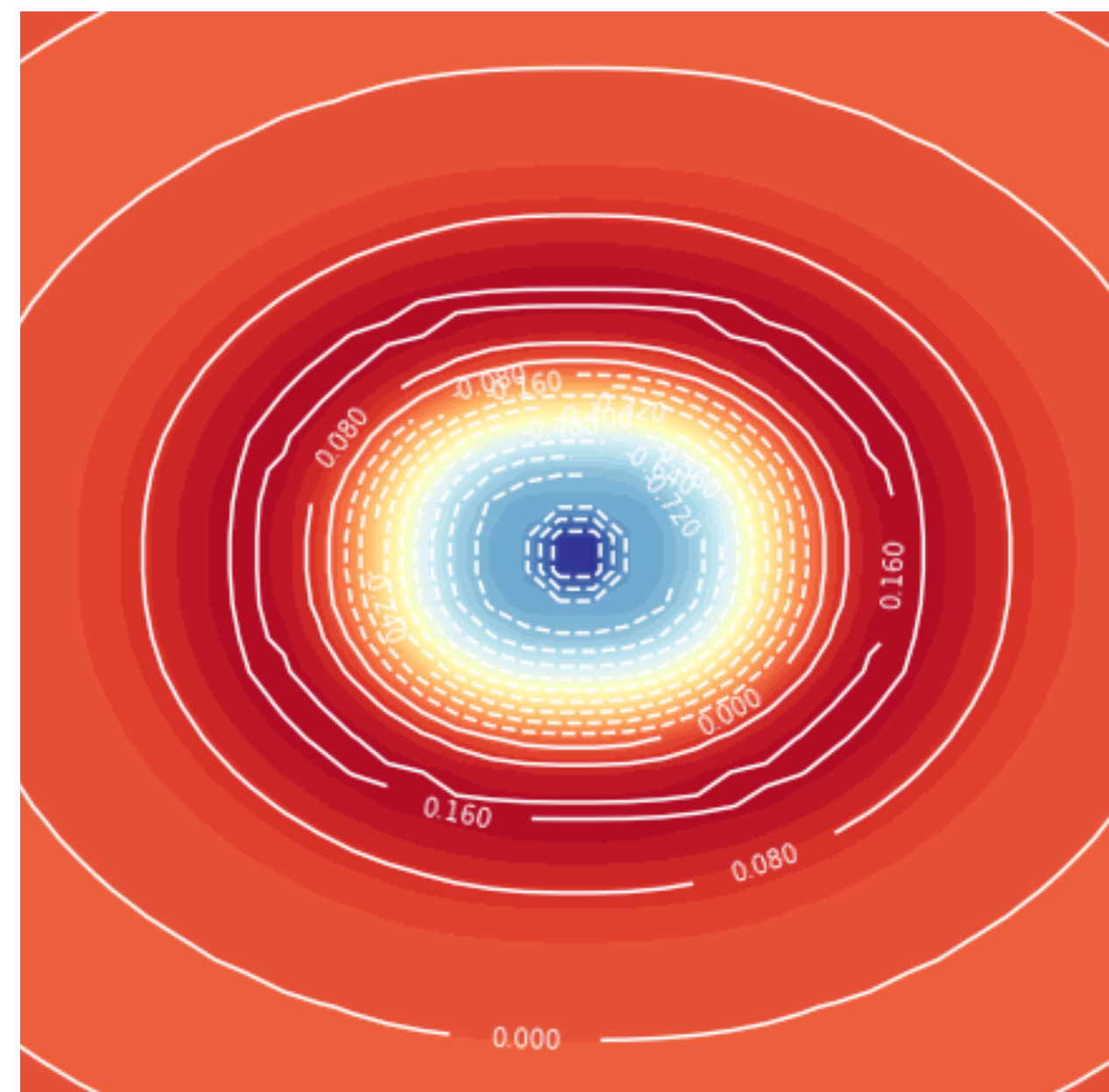


Important property : void are on average spherical !

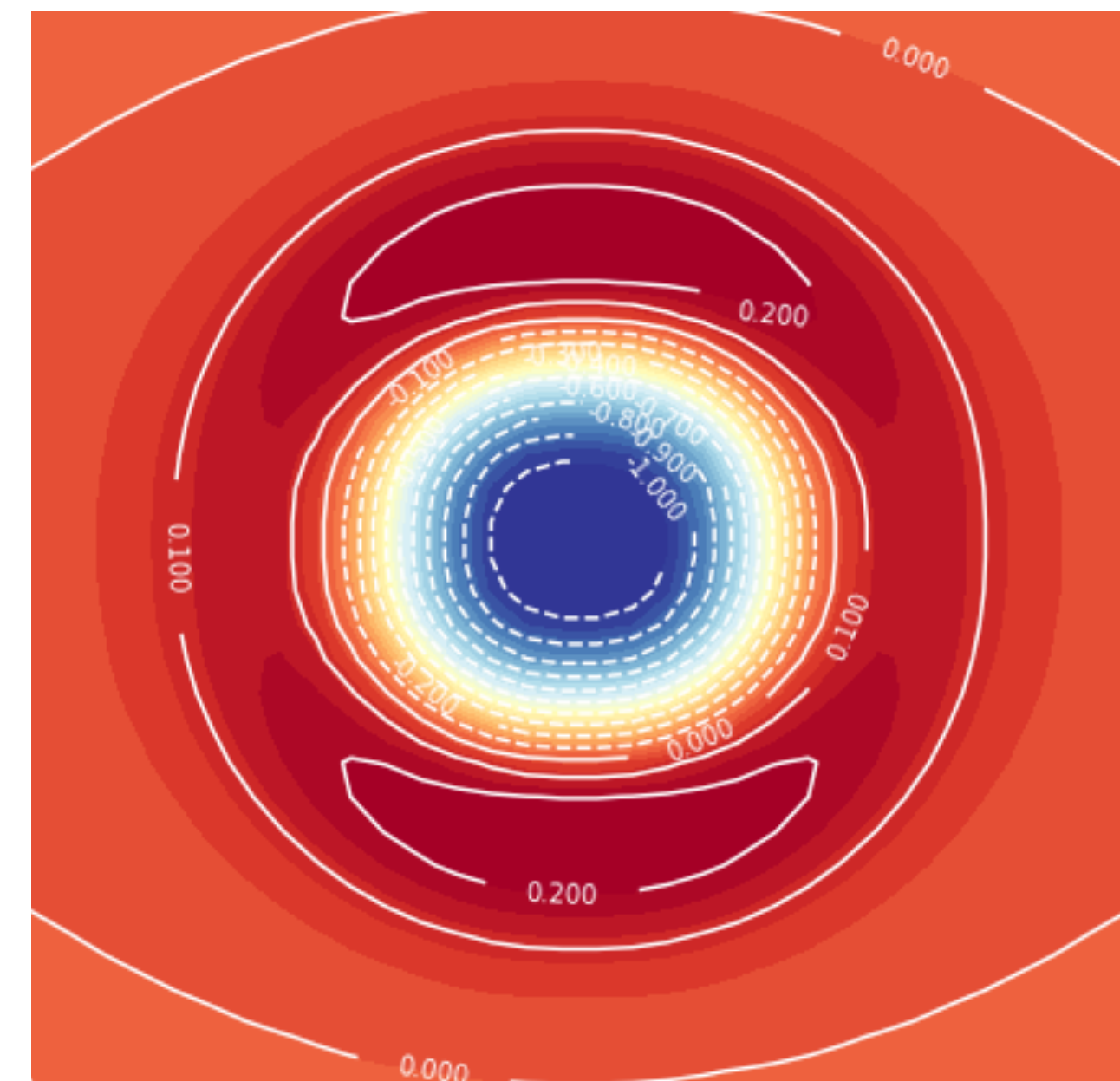
**Real Space
(no distortions)**



AP distortions



**Redshift space
distortions**



Alcock-Paczynski test

A key ingredient to cosmological inference

- Distances estimated from measured redshift

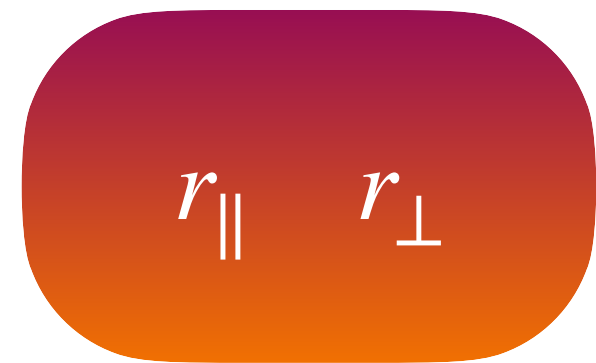
$$d(z) = \int_0^z \frac{cdz'}{H(z')}$$



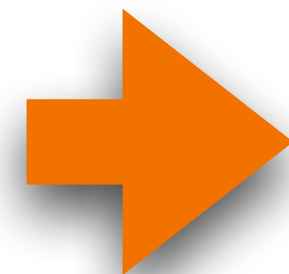
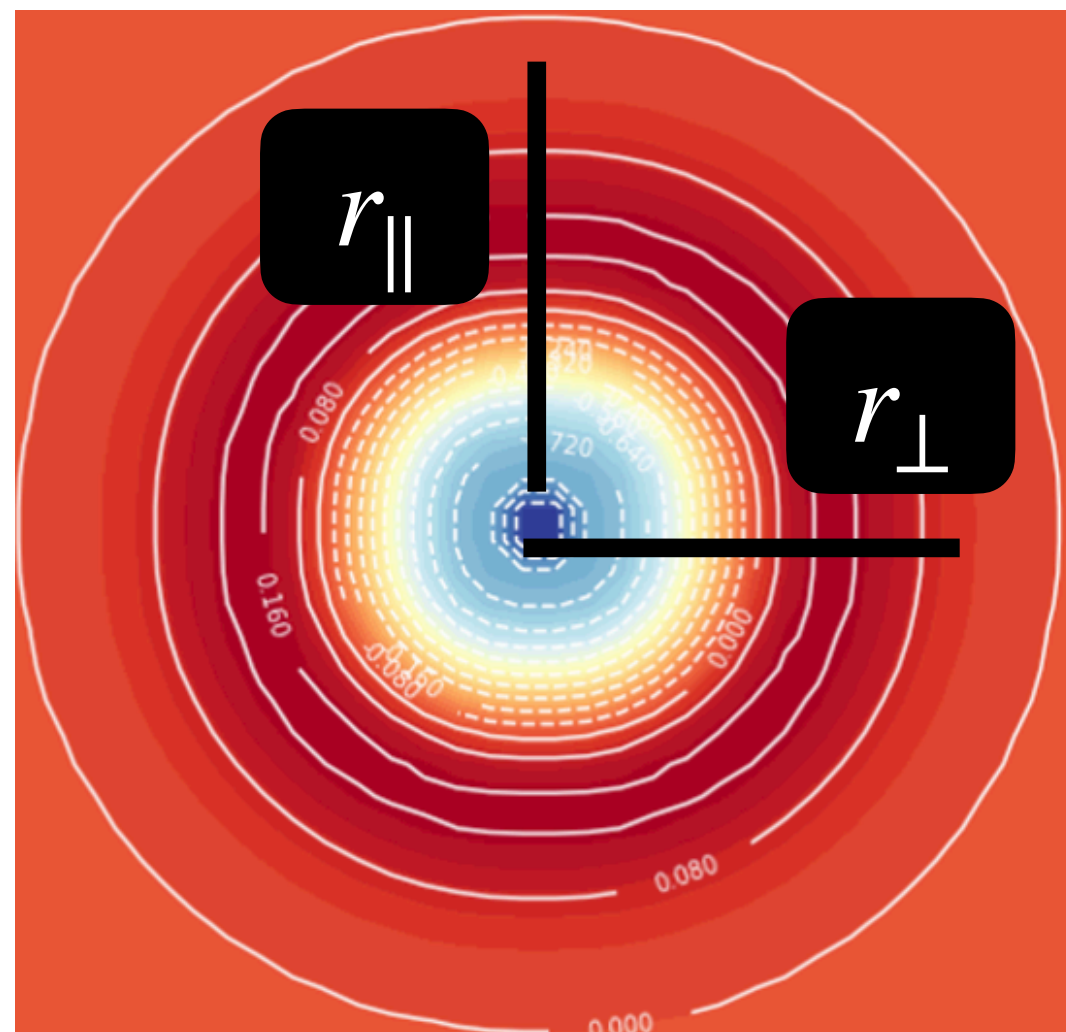
Assuming a flat Universe

- Λ CDM: $H(z) = H_0 \sqrt{\Omega_m(1+z)^3 + \Omega_\Lambda}$
- $w_0 w_a$ CDM: $H(z) = H_0 \sqrt{\Omega_m(1+z)^3 + \Omega_{de}(1+z)^{3(1+w_0+w_a)} e^{-3w_a z/(1+z)}}$

True cosmology



Real Space

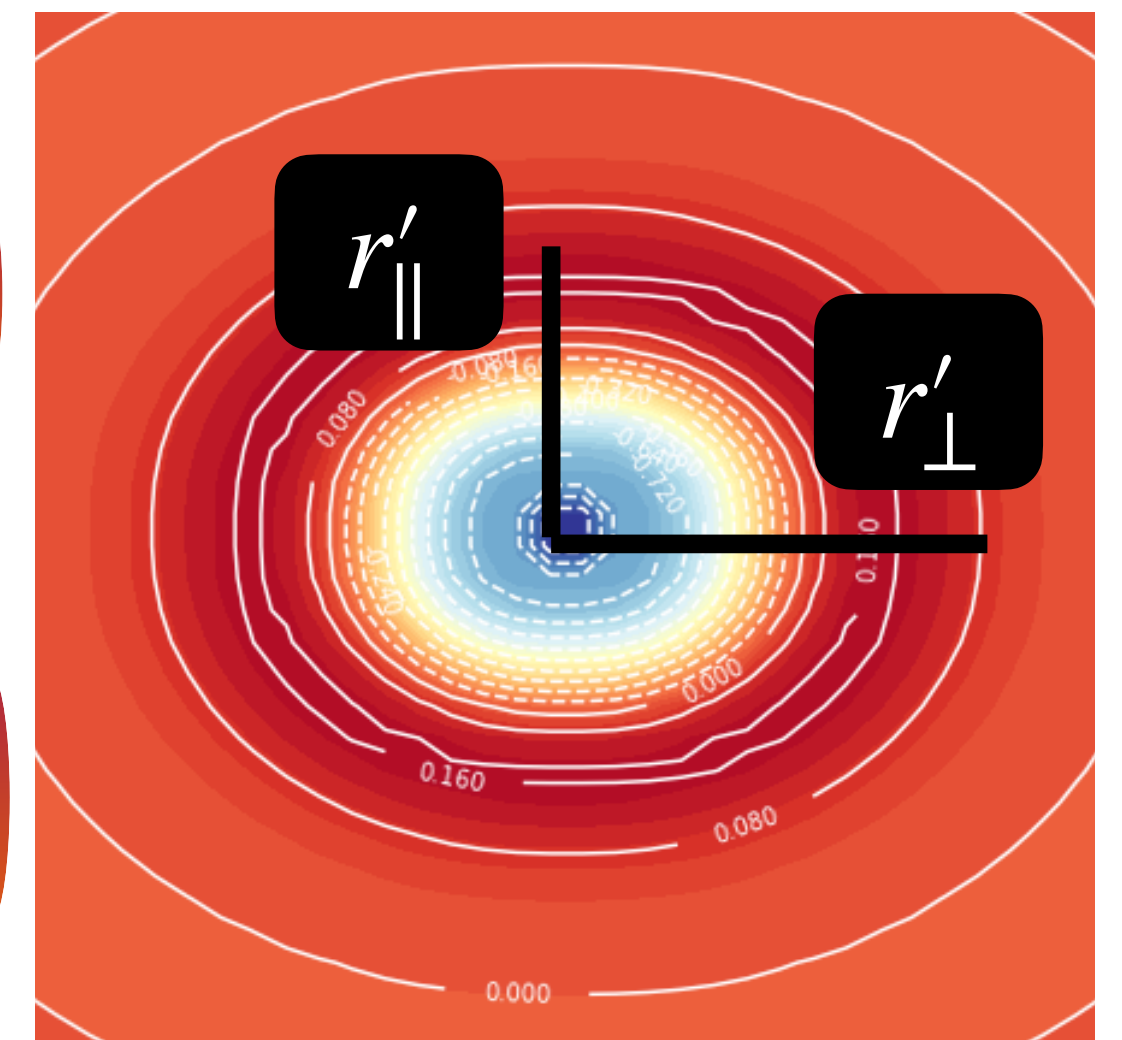


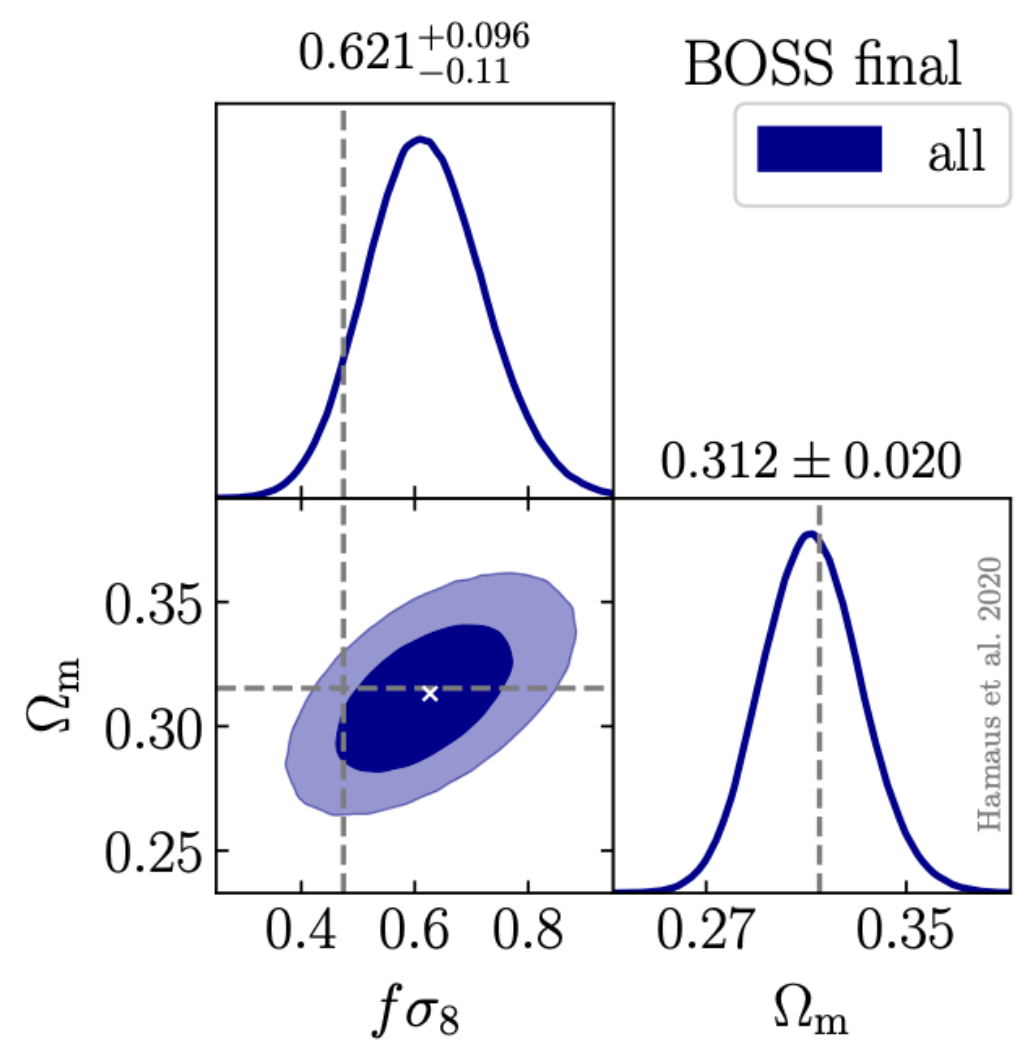
Fiducial cosmology

$$r_{\parallel} = q_{\parallel} r'_{\parallel} \quad r_{\perp} = q_{\perp} r'_{\perp}$$

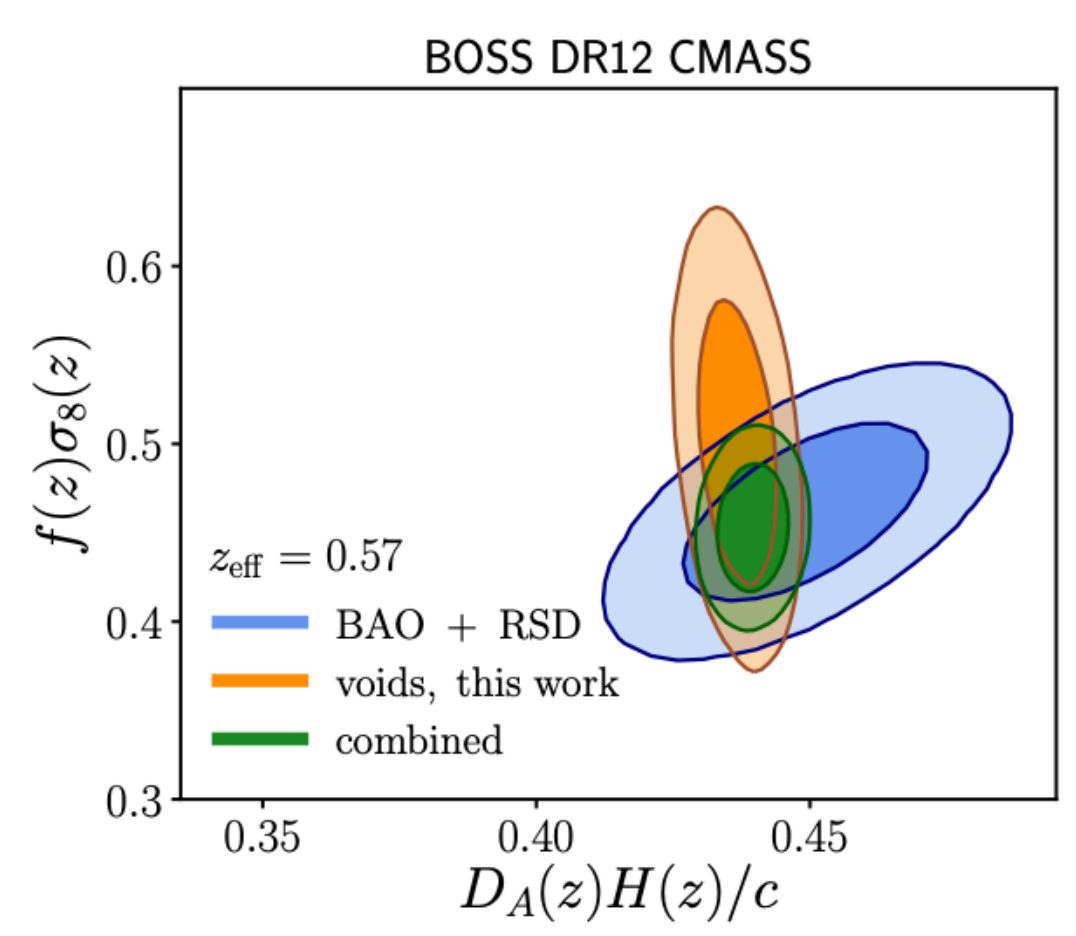
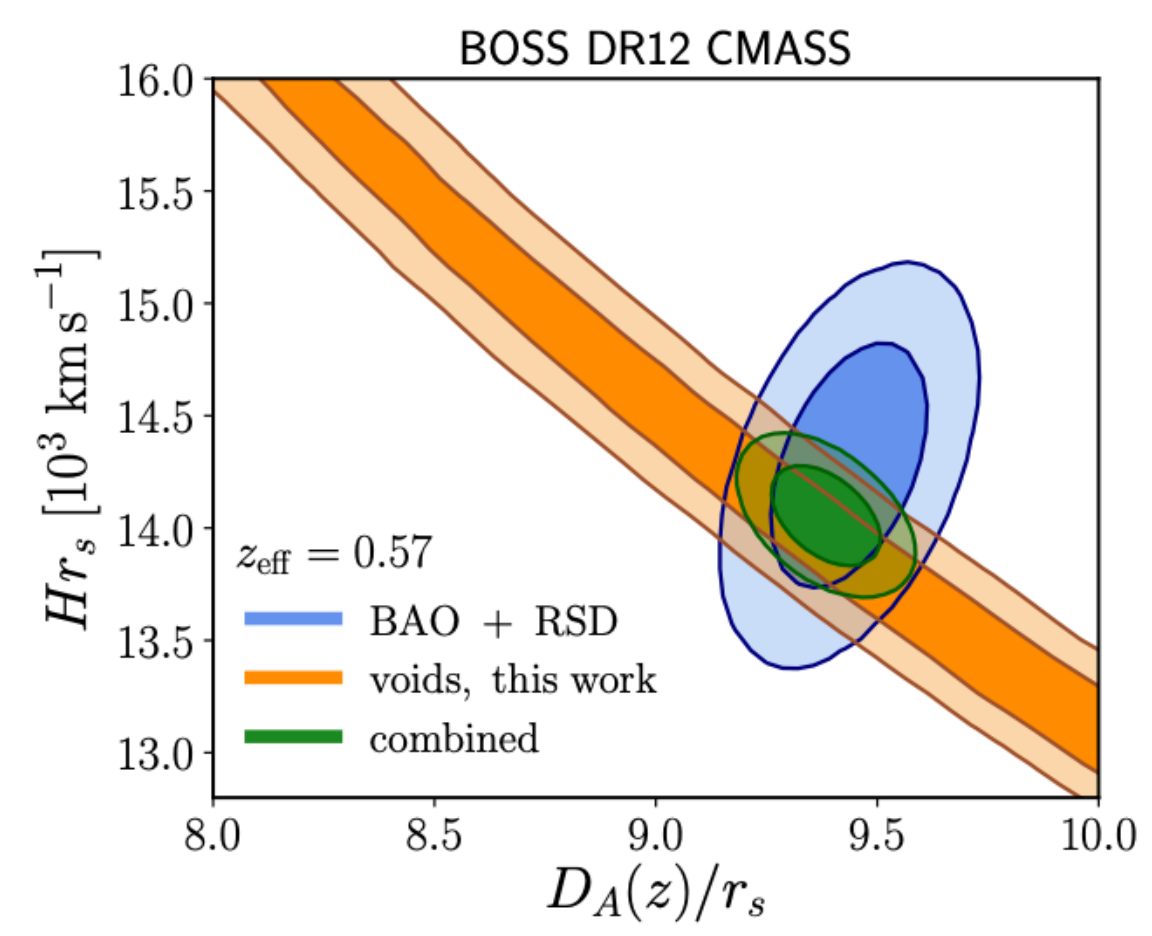
$$\varepsilon = \frac{q_{\perp}}{q_{\parallel}} = \frac{D_A(z)H(z)}{D'_A(z)H'(z)}$$

AP distortions

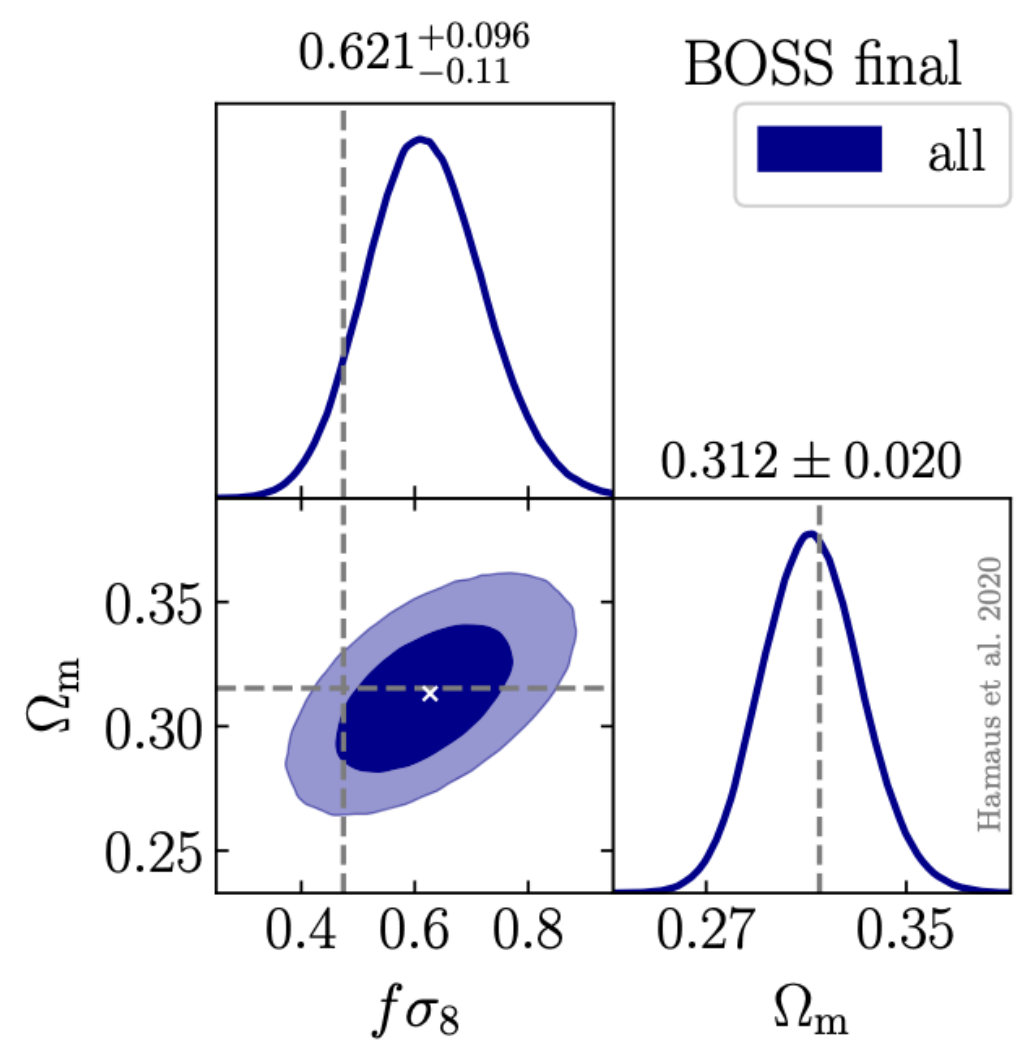




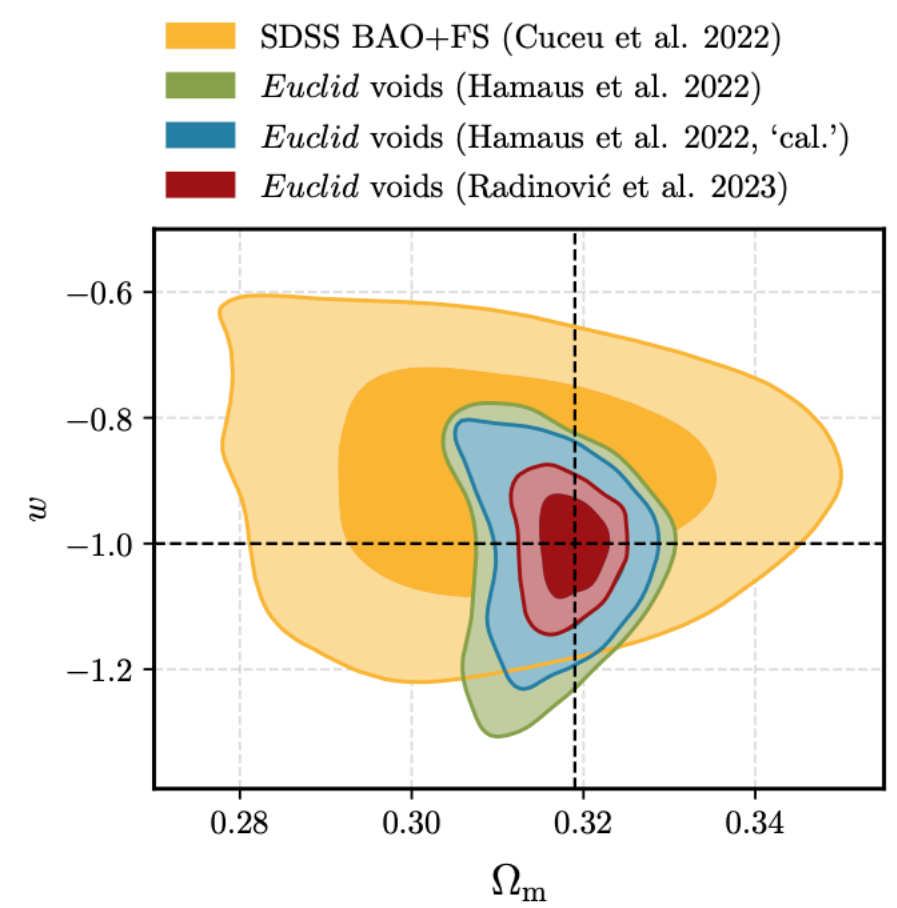
Hamaus et al. 2020



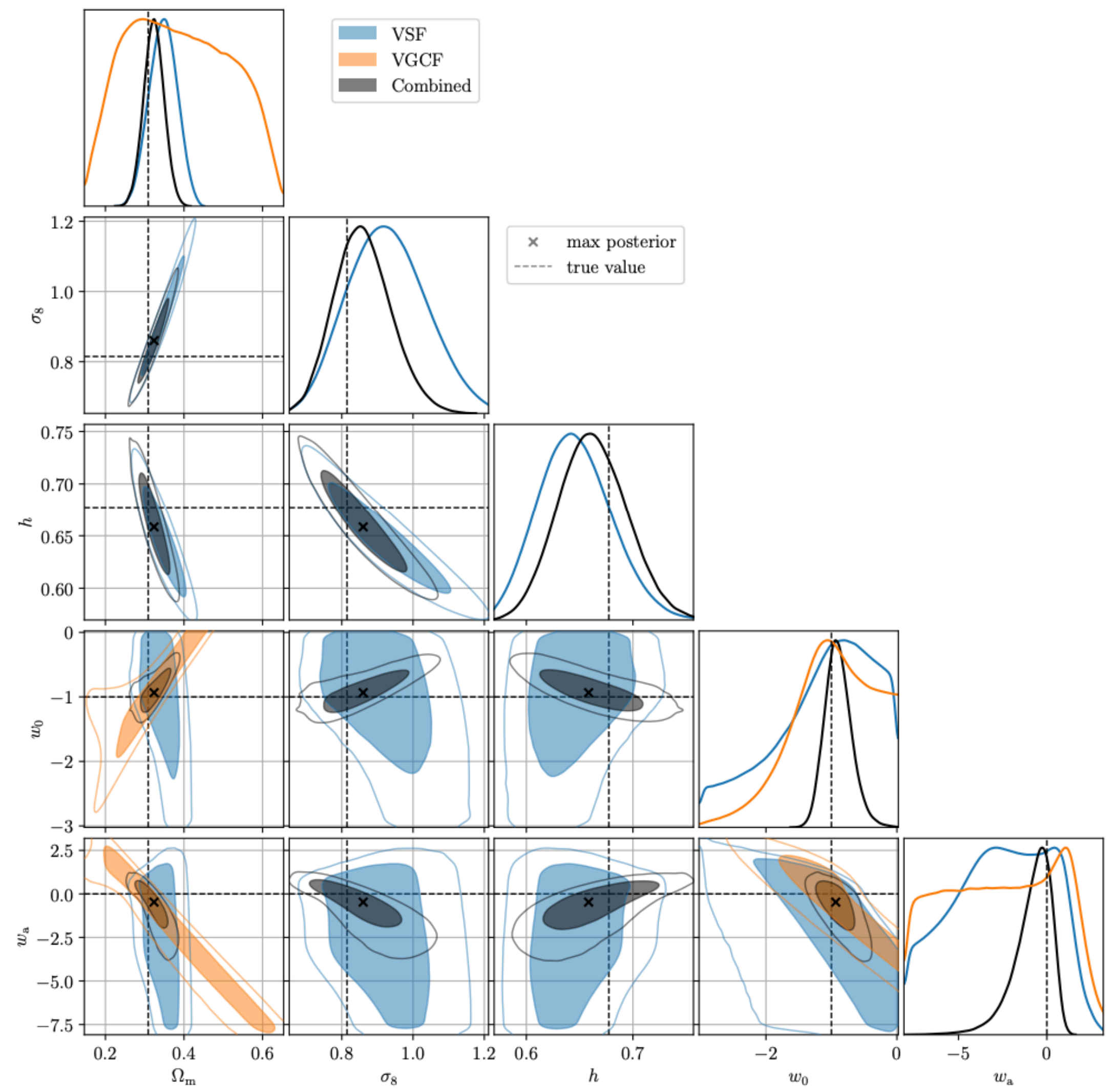
Nadathur et al. 2020



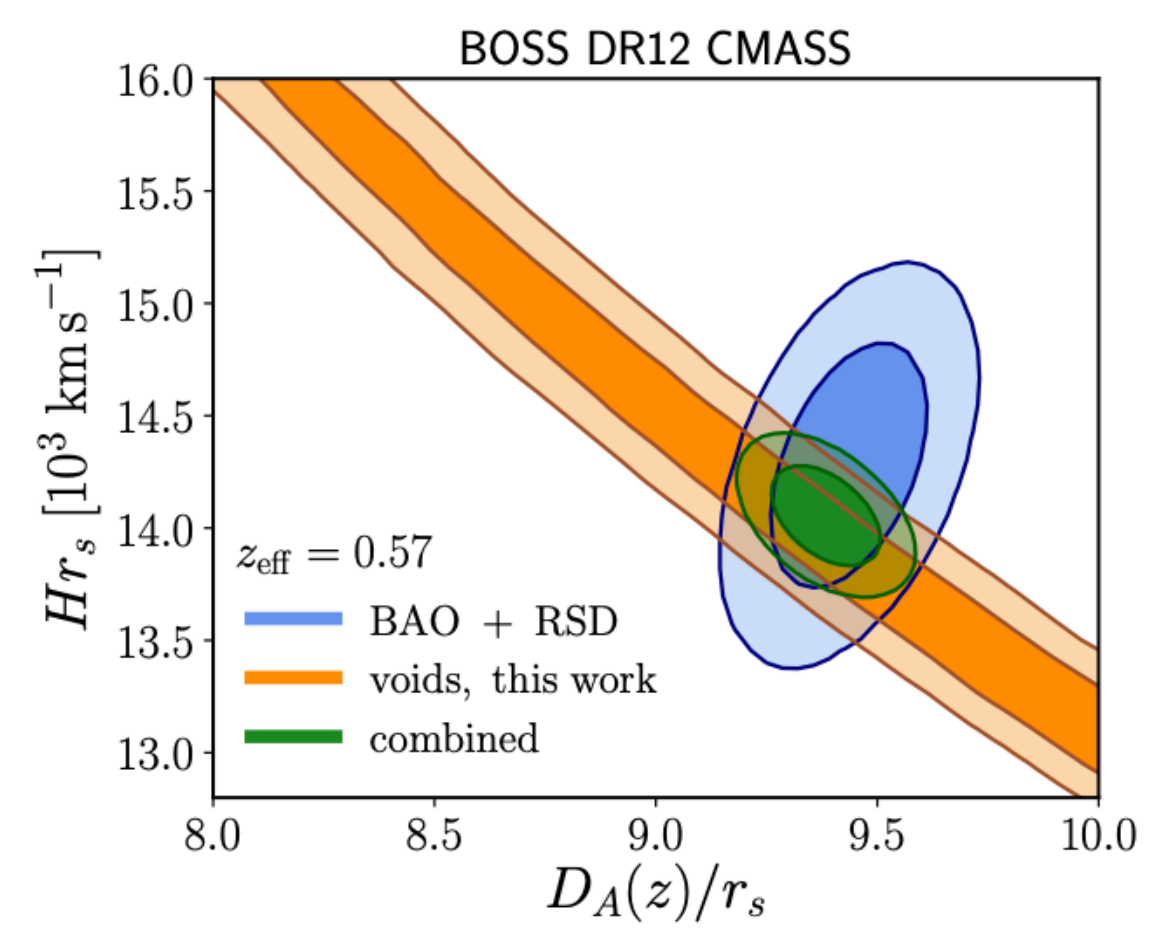
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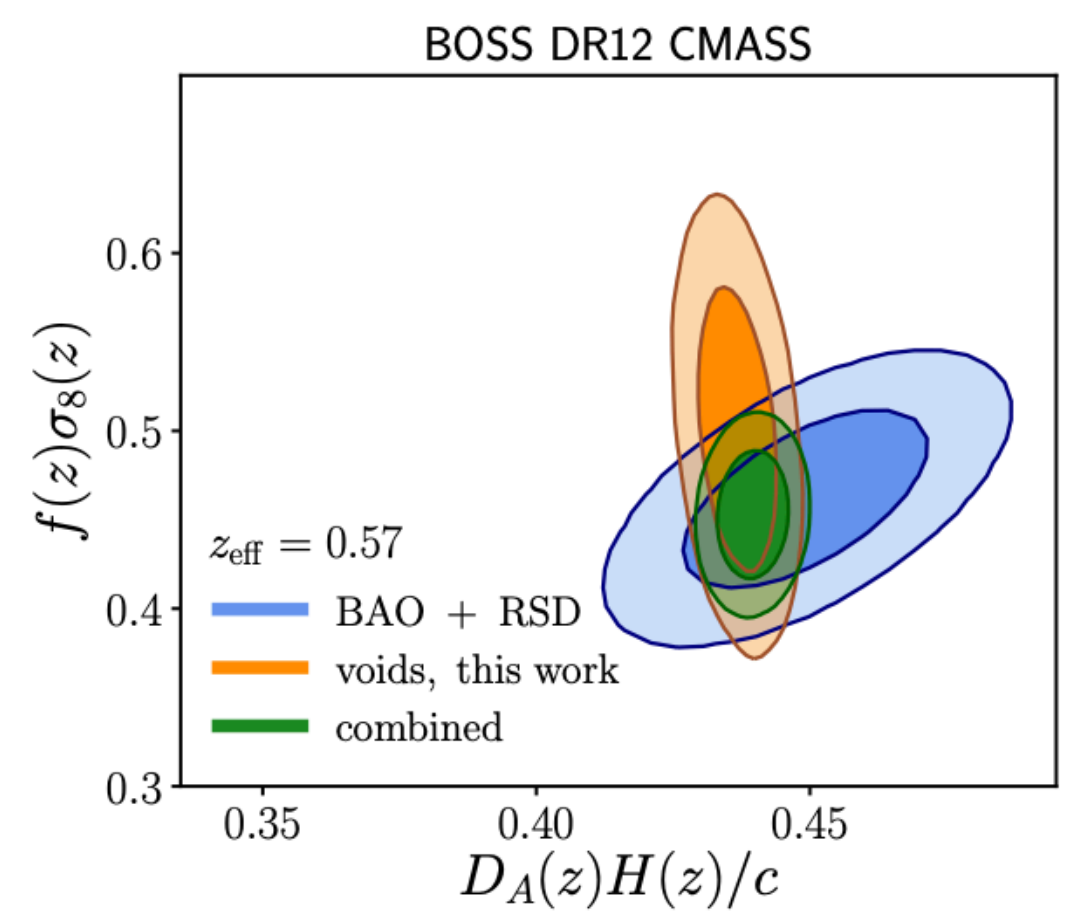
Radinovic et al. 2023



Verza, Degni et al. 2025



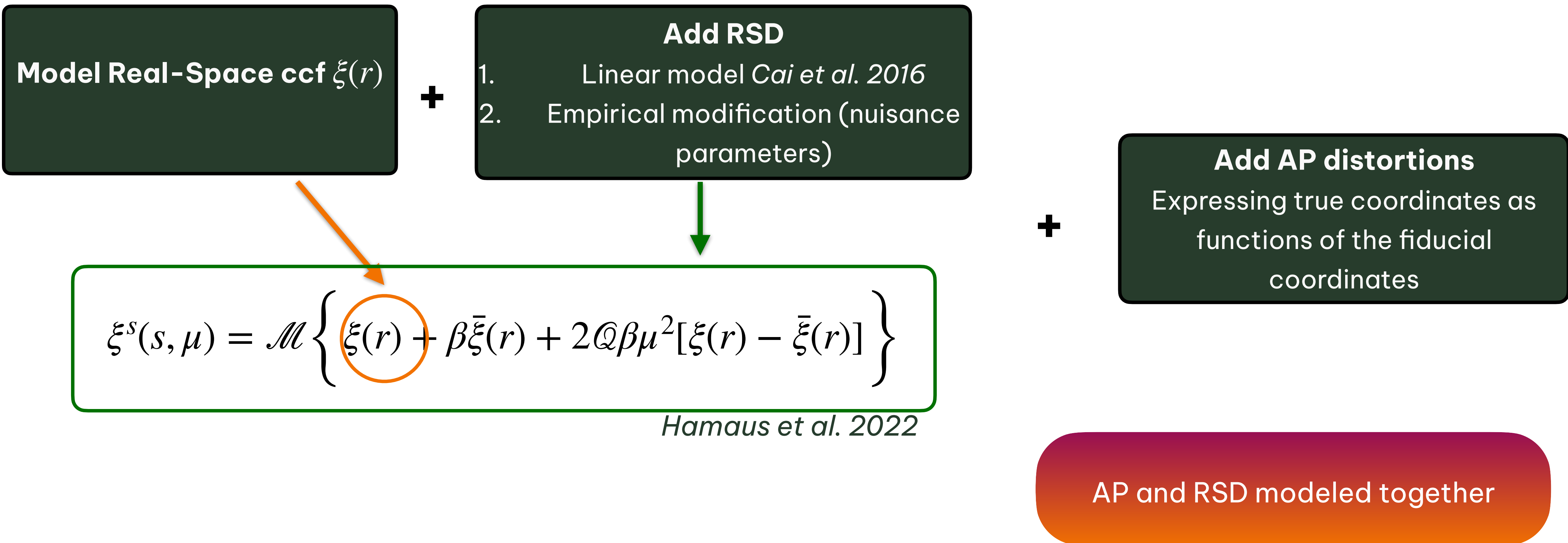
Nadathur et al. 2020



How to deal with RSD in the VGCF?

Standard approach - analytically modeling RSD in voids:

Assume local mass conservation at linear order in density fluctuations with growth rate f



Issues with this approach:

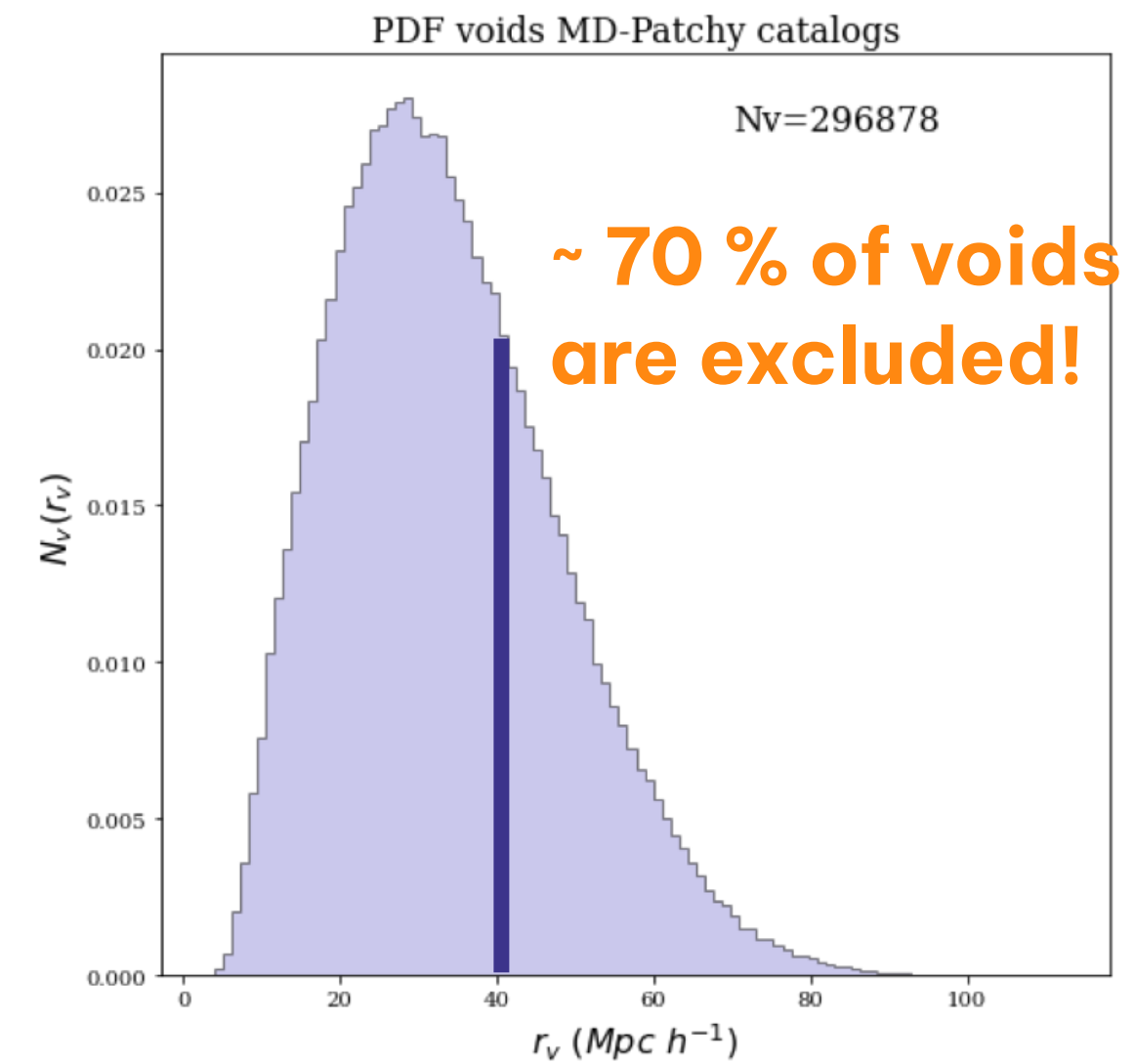
Smaller voids are difficult to model

Difficult to include systematics from the void identification
– see Selection Effect in *Correa et al. 2022*

Solutions and alternative methods :

Eliminating voids at smaller scales

Standard



*Voids radii distribution in
BOSS Patchy mocks*

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In progress

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Dynamical criteria to identify voids

Sartori et al. 2026

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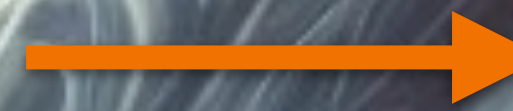
Removing RSD in the void finding process by reconstructing velocities

Nadathur et al. 2019

New approach : THIS WORK !

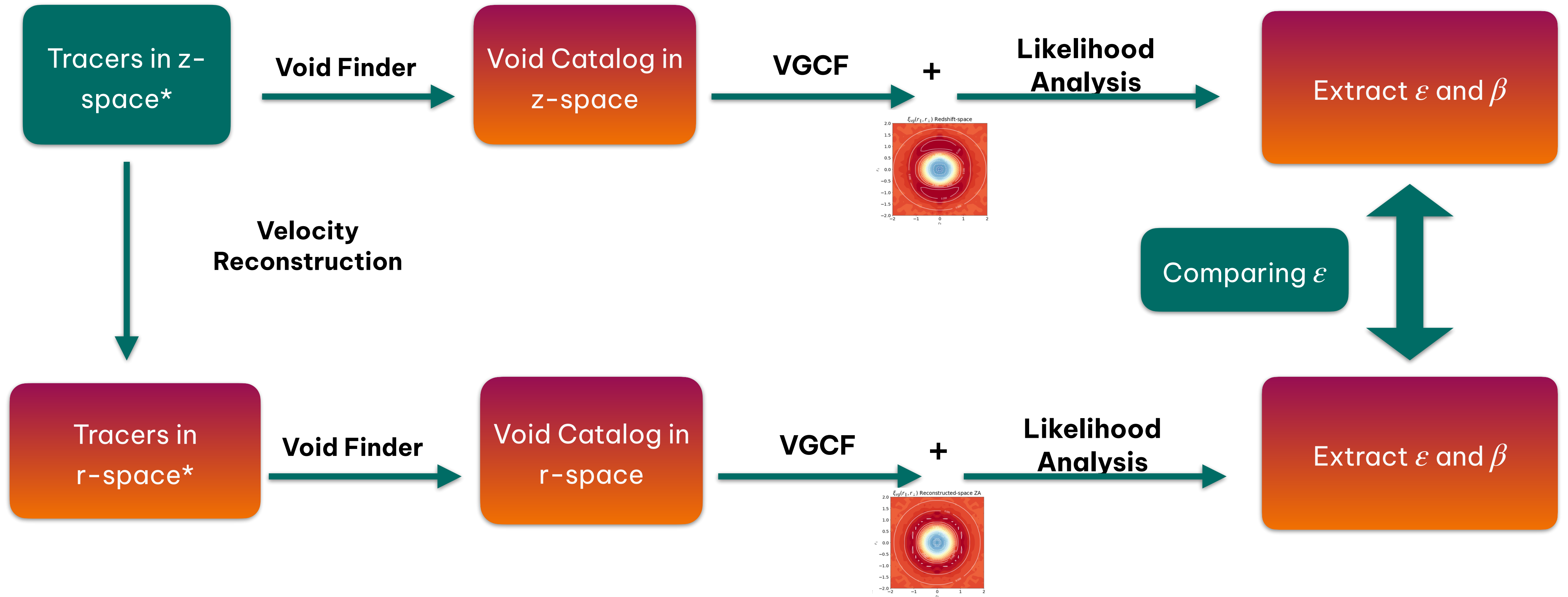
Degni et al. 2026

- Apply a velocity reconstruction technique to eliminate RSD from the position of galaxies creating the **reconstructed space**
- **Perform the AP test in reconstructed space**
- Goal: improve the statistical signal by including all the voids in the sample, reduce the impact of systematics introduced by the void finding process when finding voids in redshift space



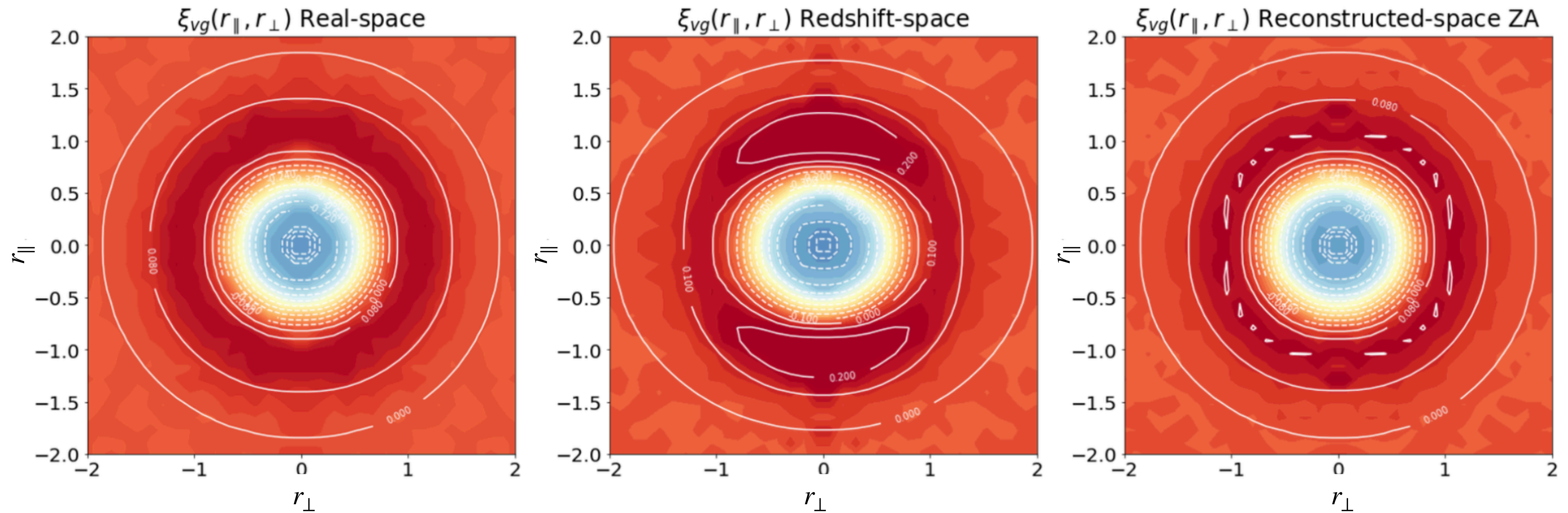
In this work: reconstruction based on Zel'dovich approximation

Strategy

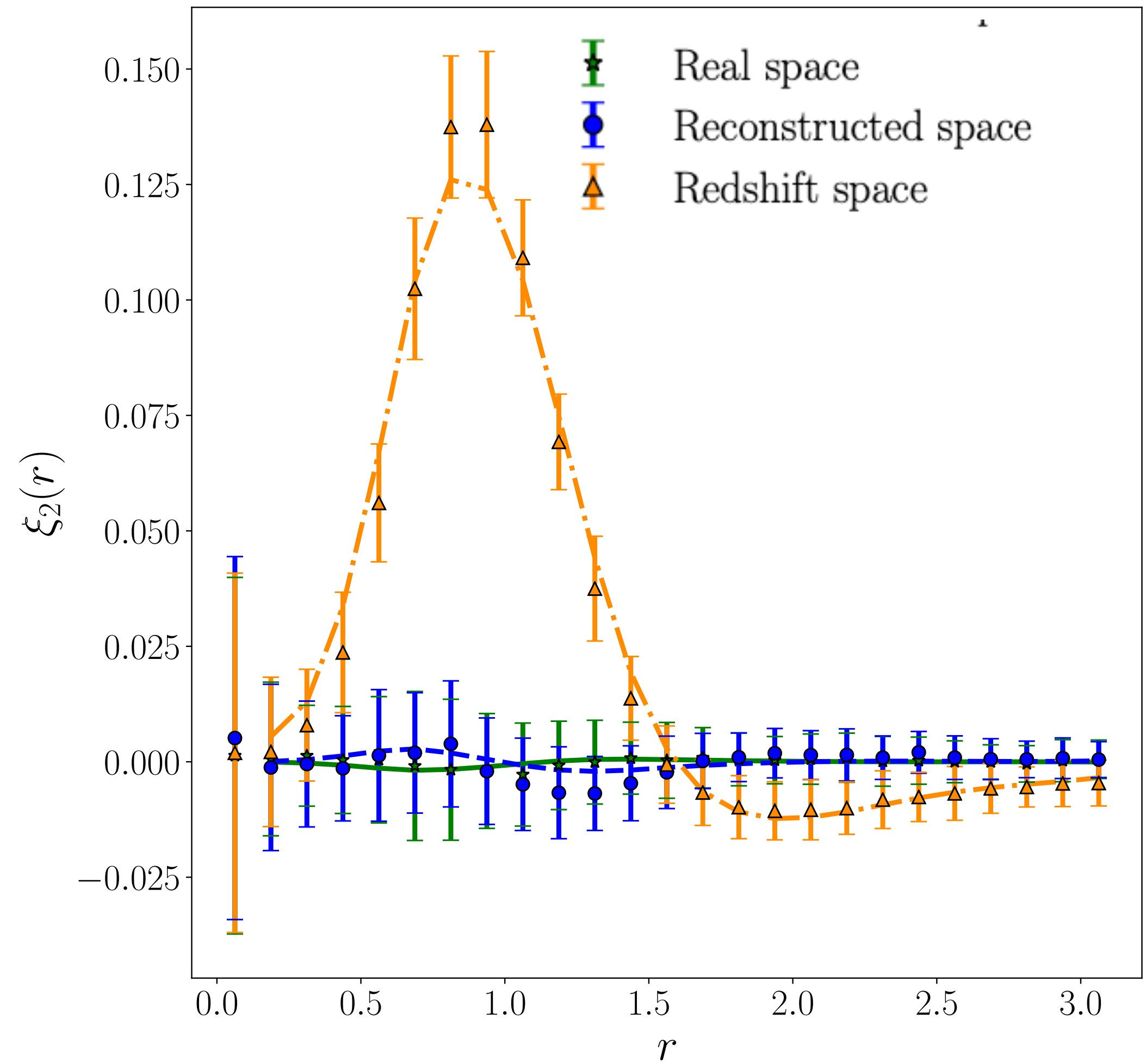
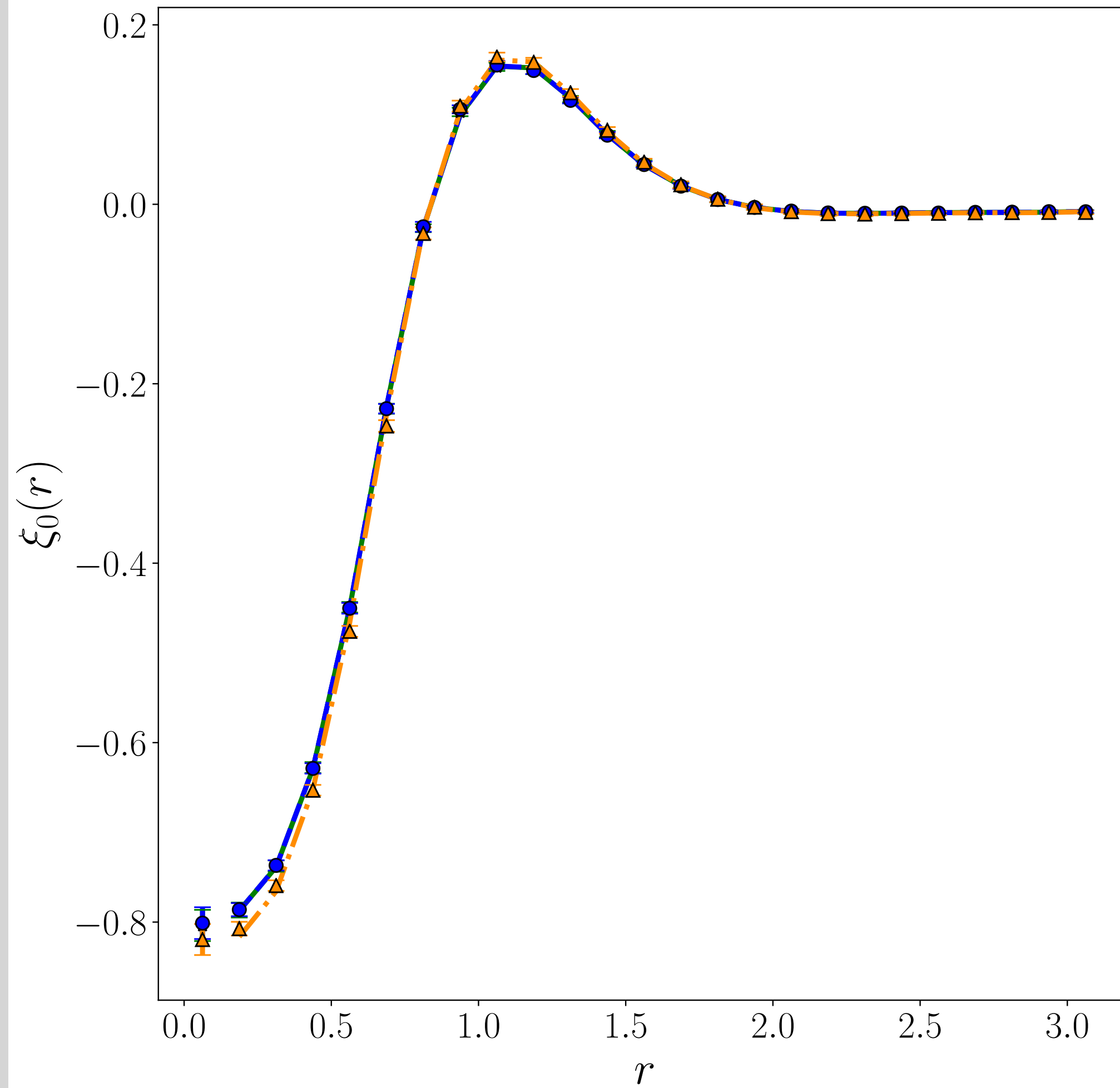


Data: Halos, Cubic Box, $L_{box} = 1000 h^{-1} Mpc$, from Quijote Simulation High Resolution, $\bar{z} = 0.5$

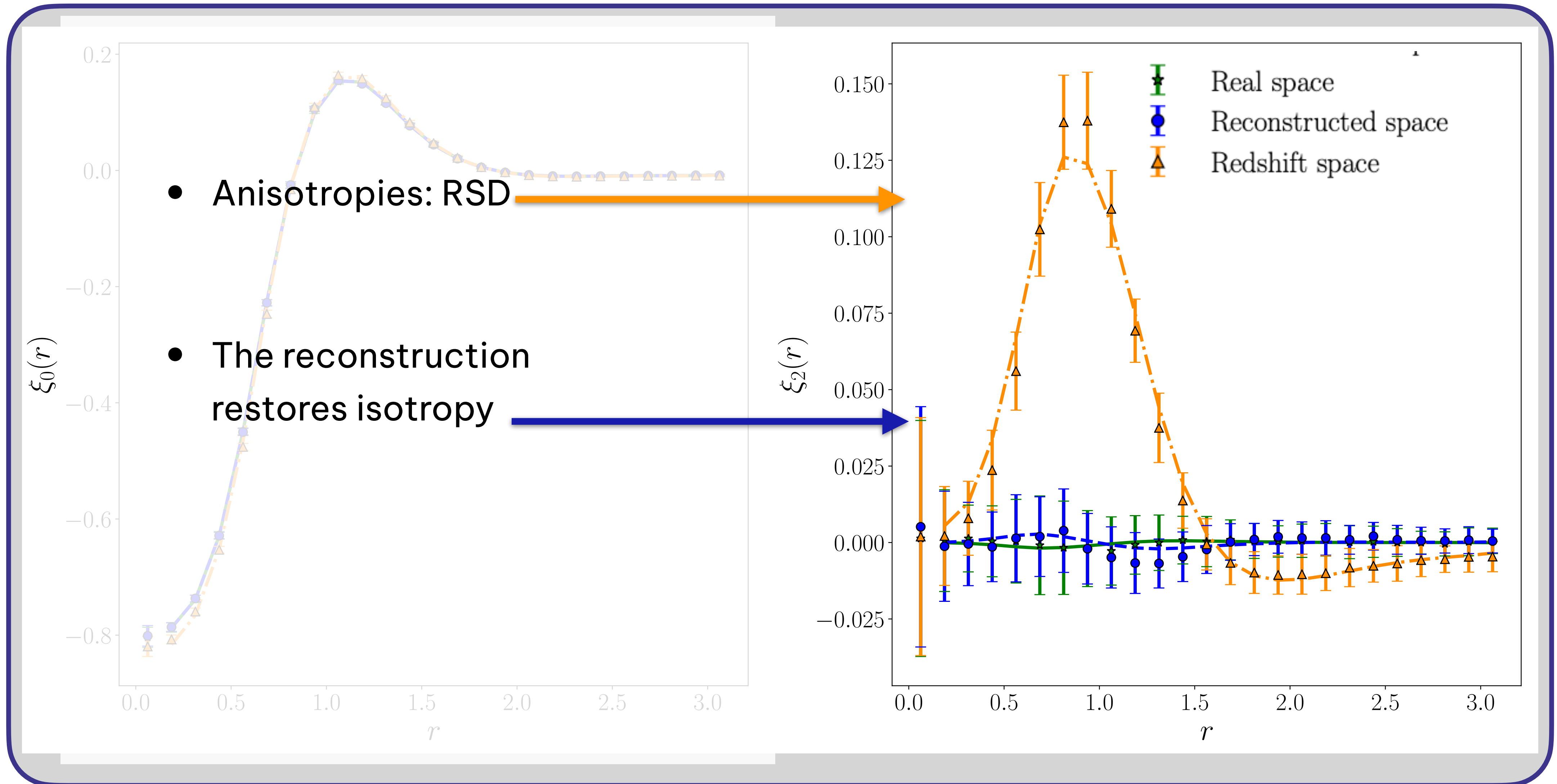
Redshift-space: **anisotropies (only RSD)** Reconstructed-space : **no anisotropies**



Void-galaxy cross-correlation function multipoles

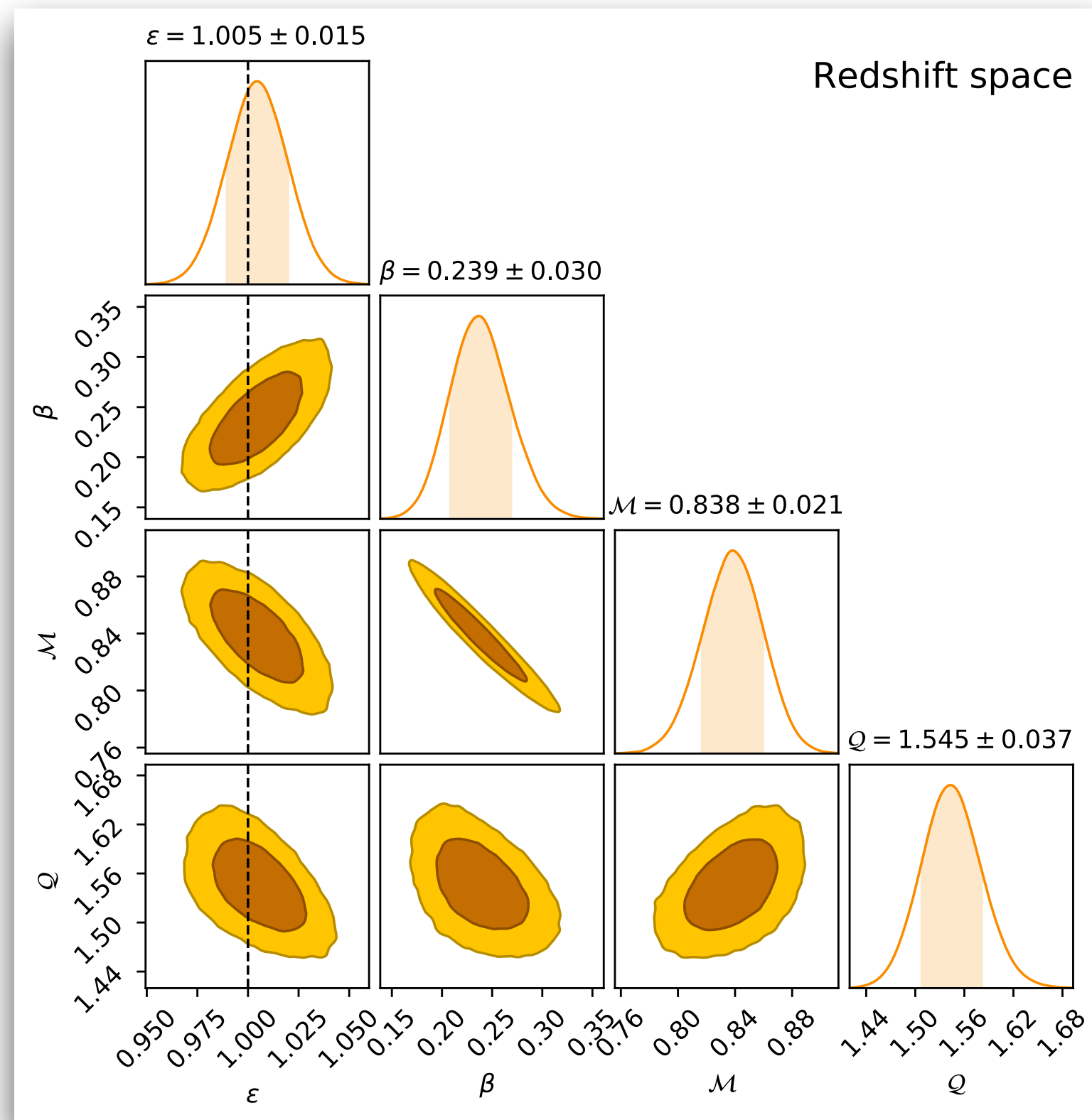


Void-galaxy cross-correlation function multipoles

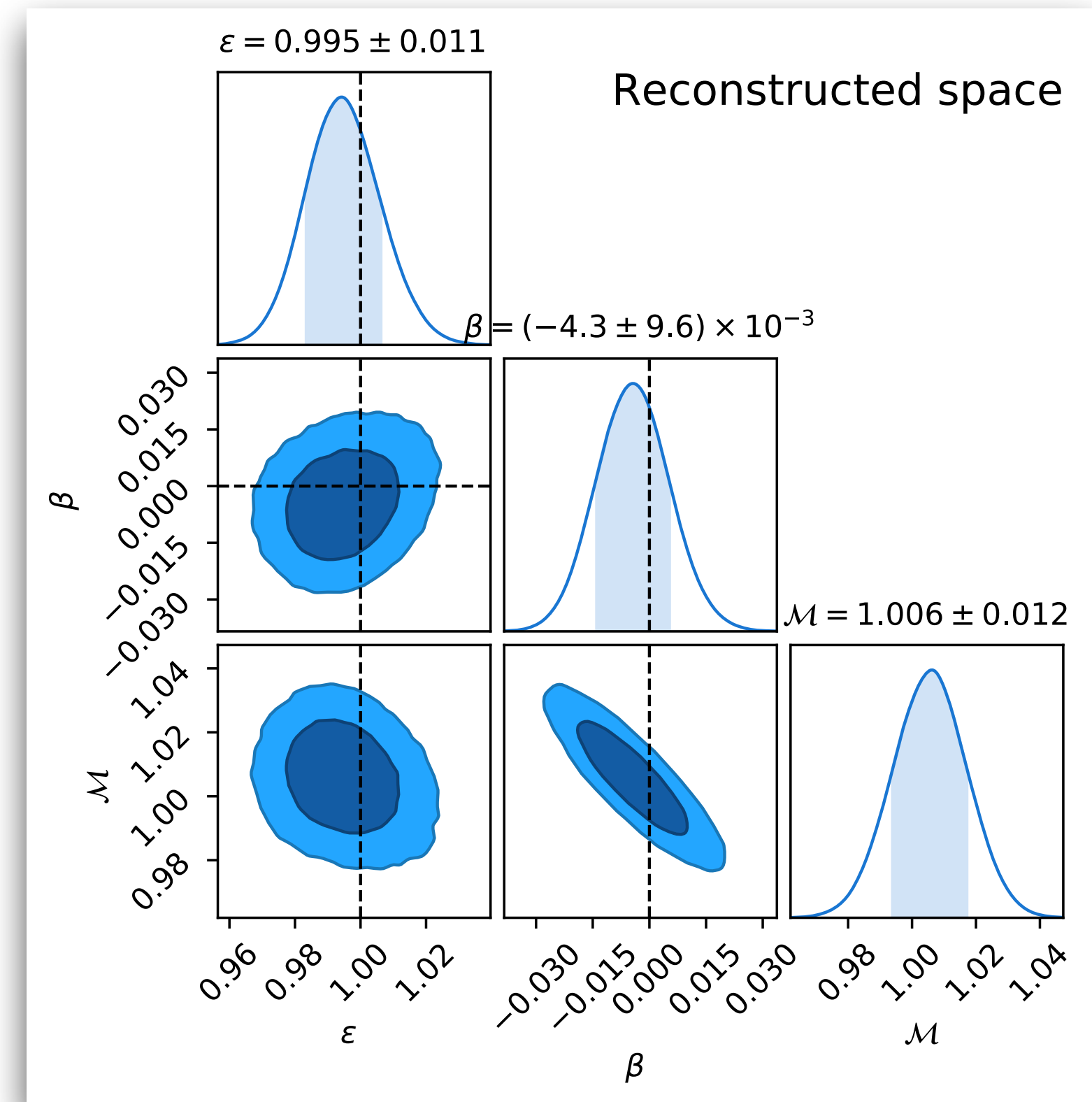


Results

Redshift space – Standard approach

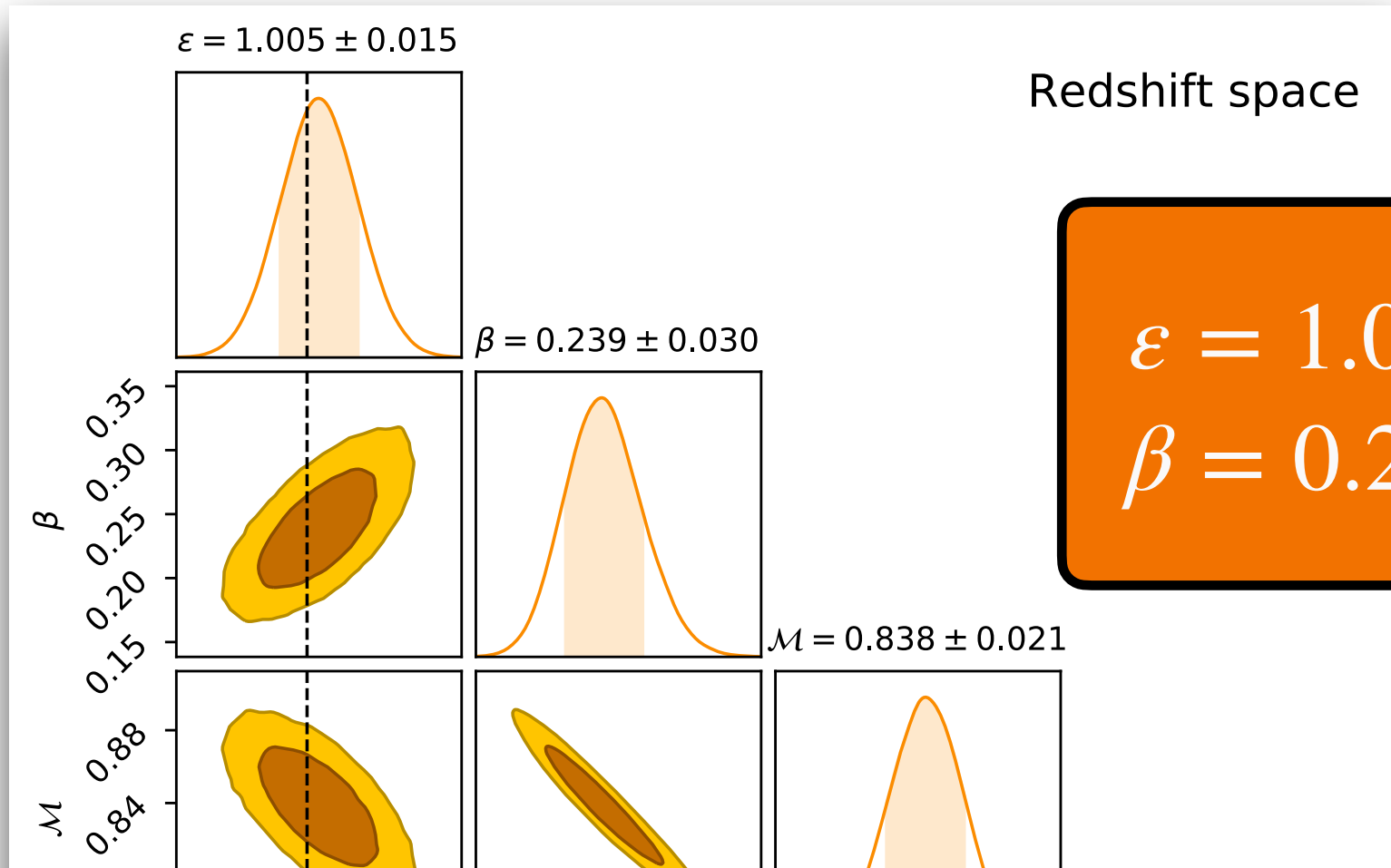


Reconstructed space – New approach



Results

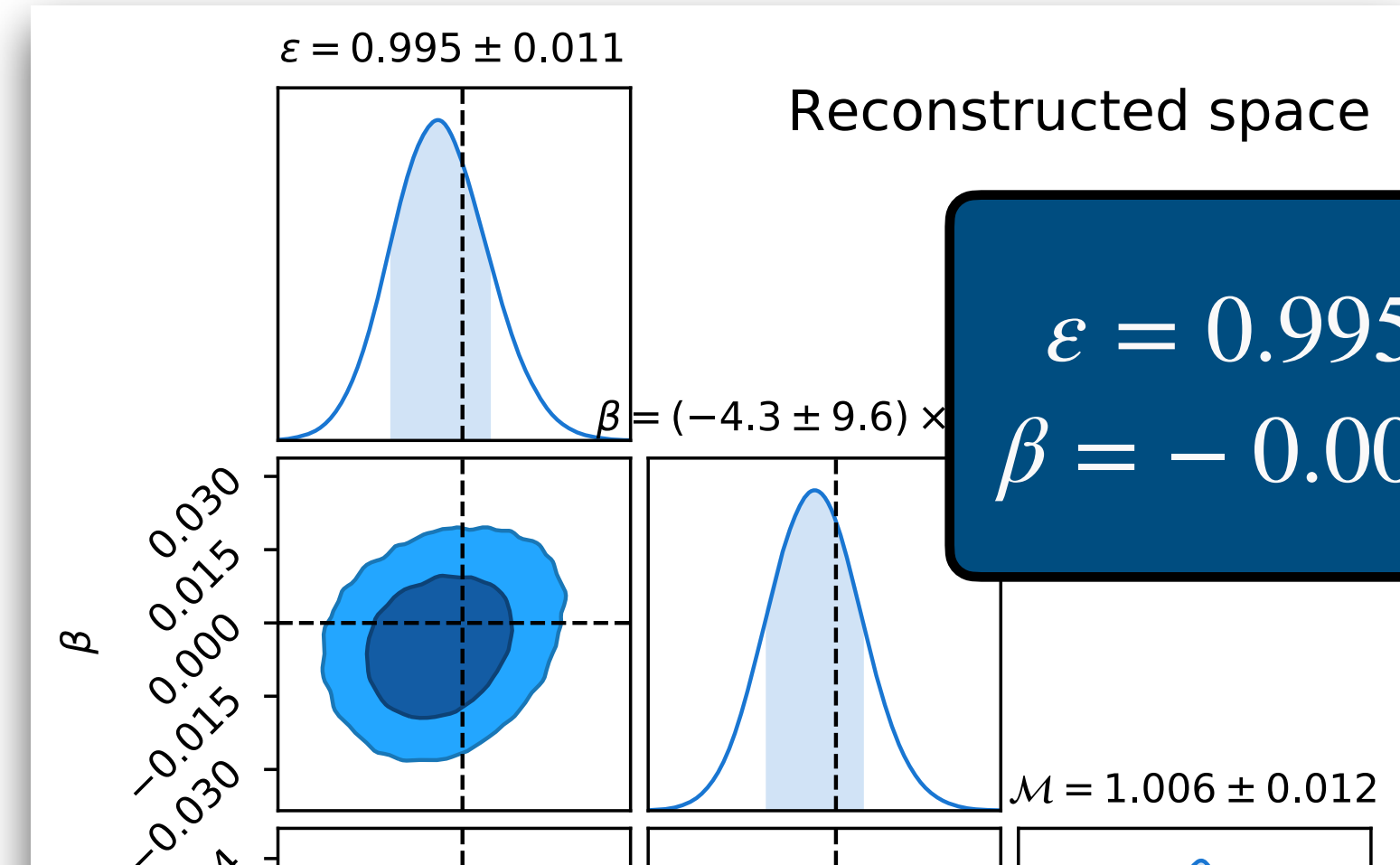
Redshift space – Standard approach



$$\varepsilon = 1.005 \pm 0.015$$
$$\beta = 0.239 \pm 0.030$$

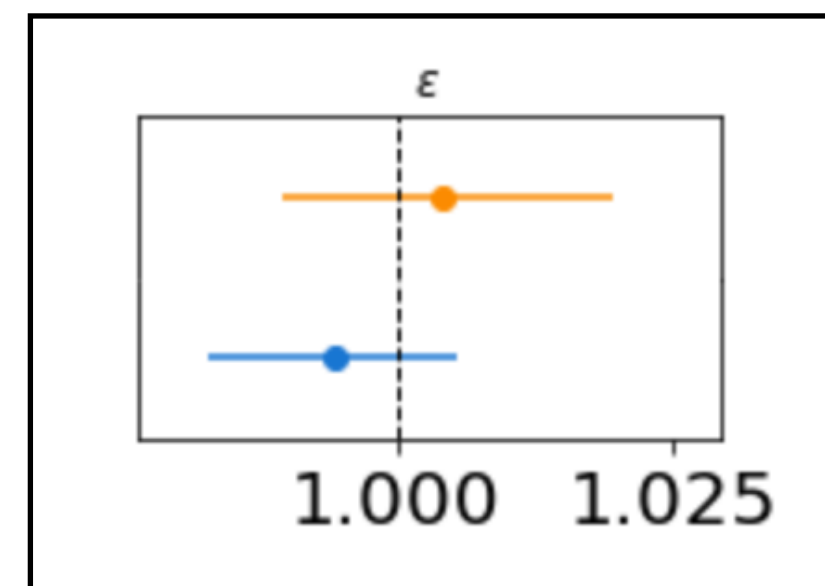
Redshift space
Precision $\sigma_\varepsilon/\varepsilon$:
1.5%

Reconstructed space – New approach



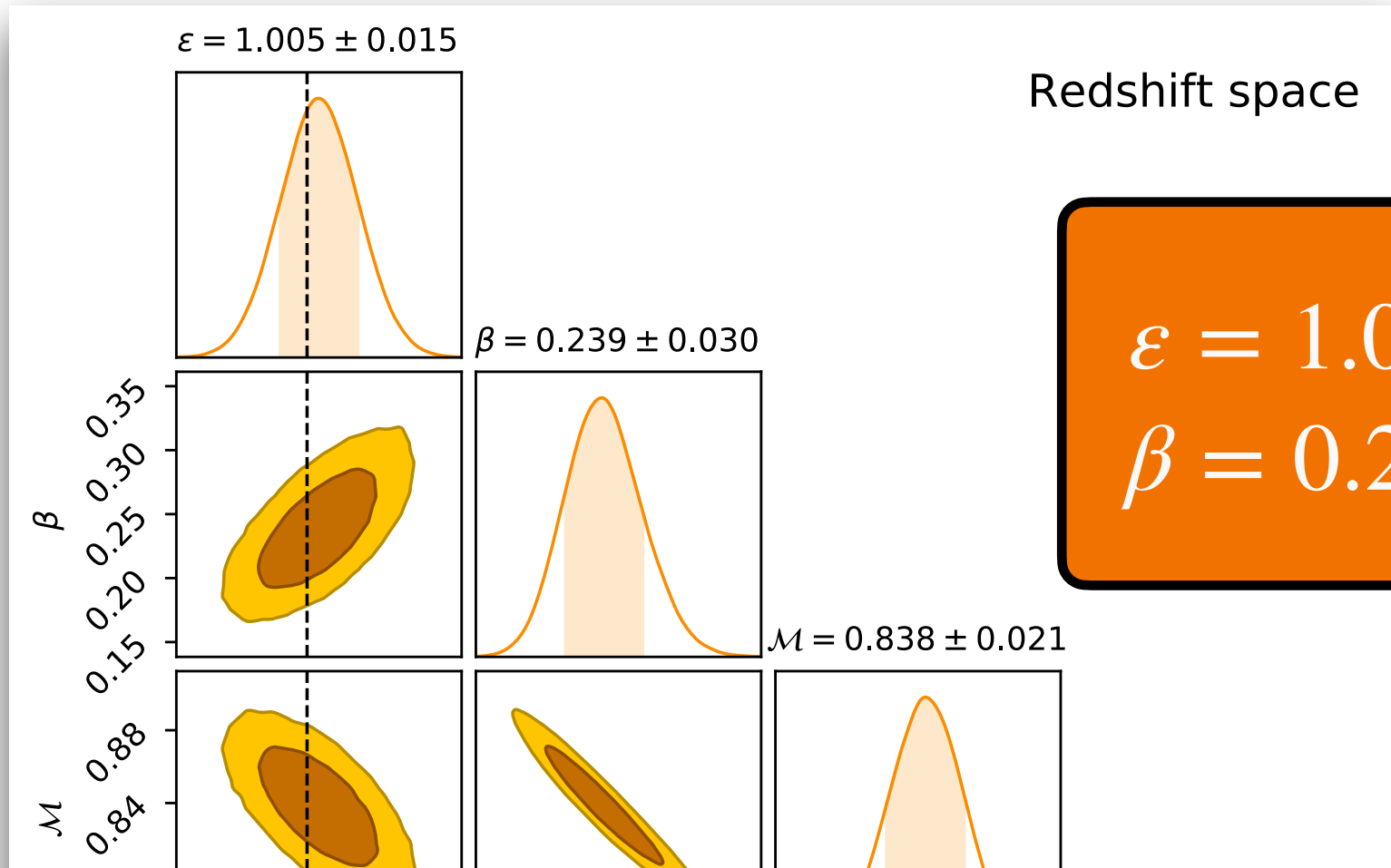
$$\varepsilon = 0.995 \pm 0.011$$
$$\beta = -0.004 \pm 0.010$$

Reconstructed space
Precision $\sigma_\varepsilon/\varepsilon$:
1.1%



Results

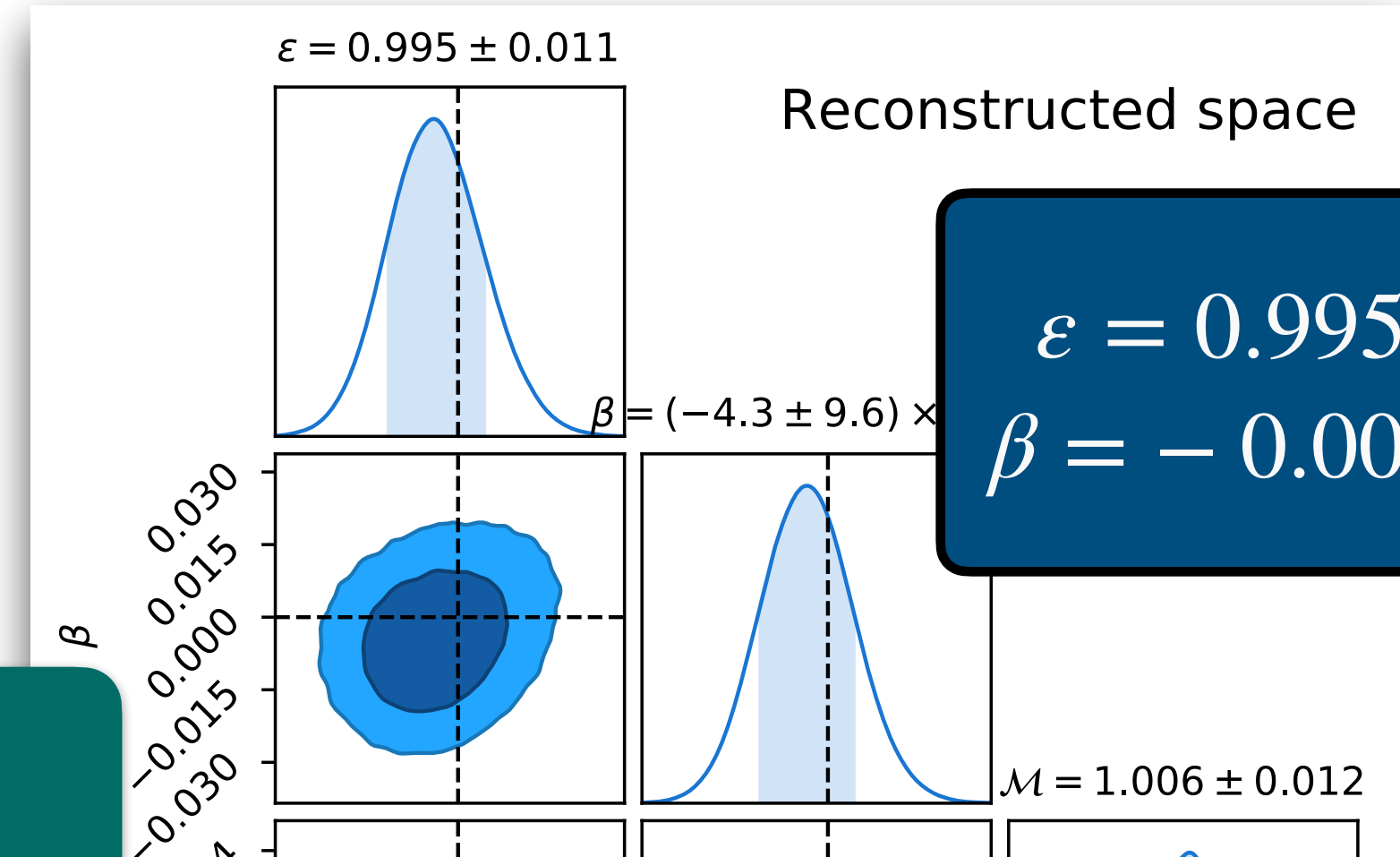
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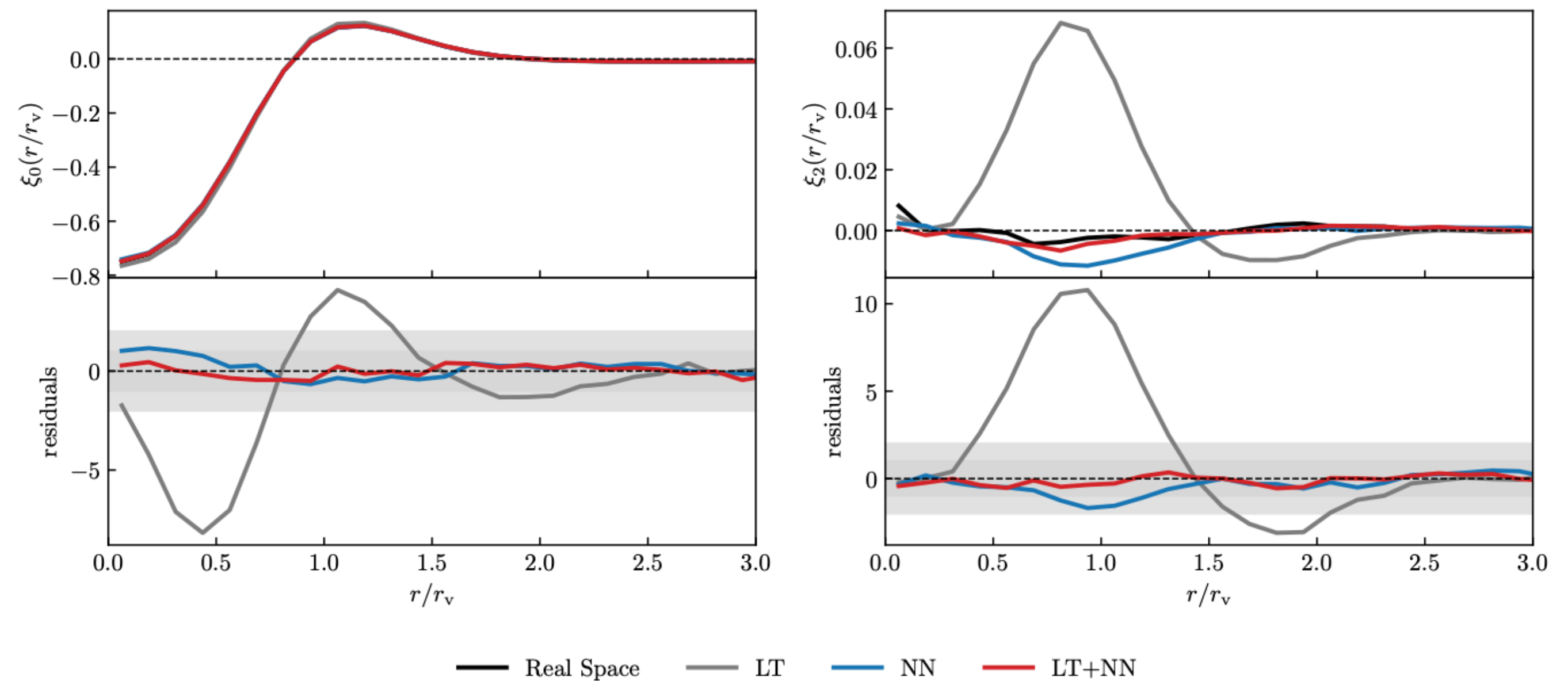
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Reconstructed space
Precision $\sigma_\varepsilon/\varepsilon$:
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Improvement factor
 $I = \sigma_{\text{pre-rec}}/\sigma_{\text{post-rec}} = 1.3$

Qualitative study of the impact of different reconstruction algorithms

- Comparison between three reconstruction methods :
 - LT : Zel'dovich reconstruction - a direct application of linear perturbation theory
 - NN : Neural network reconstruction technique based on Ganeshaiyah Veena et al. 2023
 - LT + NN : hybrid approach in which neural networks is trained on LT reconstructed fields



Maragliano et al. 2025

Conclusions and Future perspectives

- Studied the impact of reconstruction in cosmic voids
- First analysis correlating galaxies and voids in reconstructed space
- Investigating the role of different reconstruction algorithms (*Maragliano et al. 2025*)
- Apply the method to real data with the possibility to increase drastically the statistical signal

Take home message :

Reconstruction help to mitigate the degeneracy between ε and β

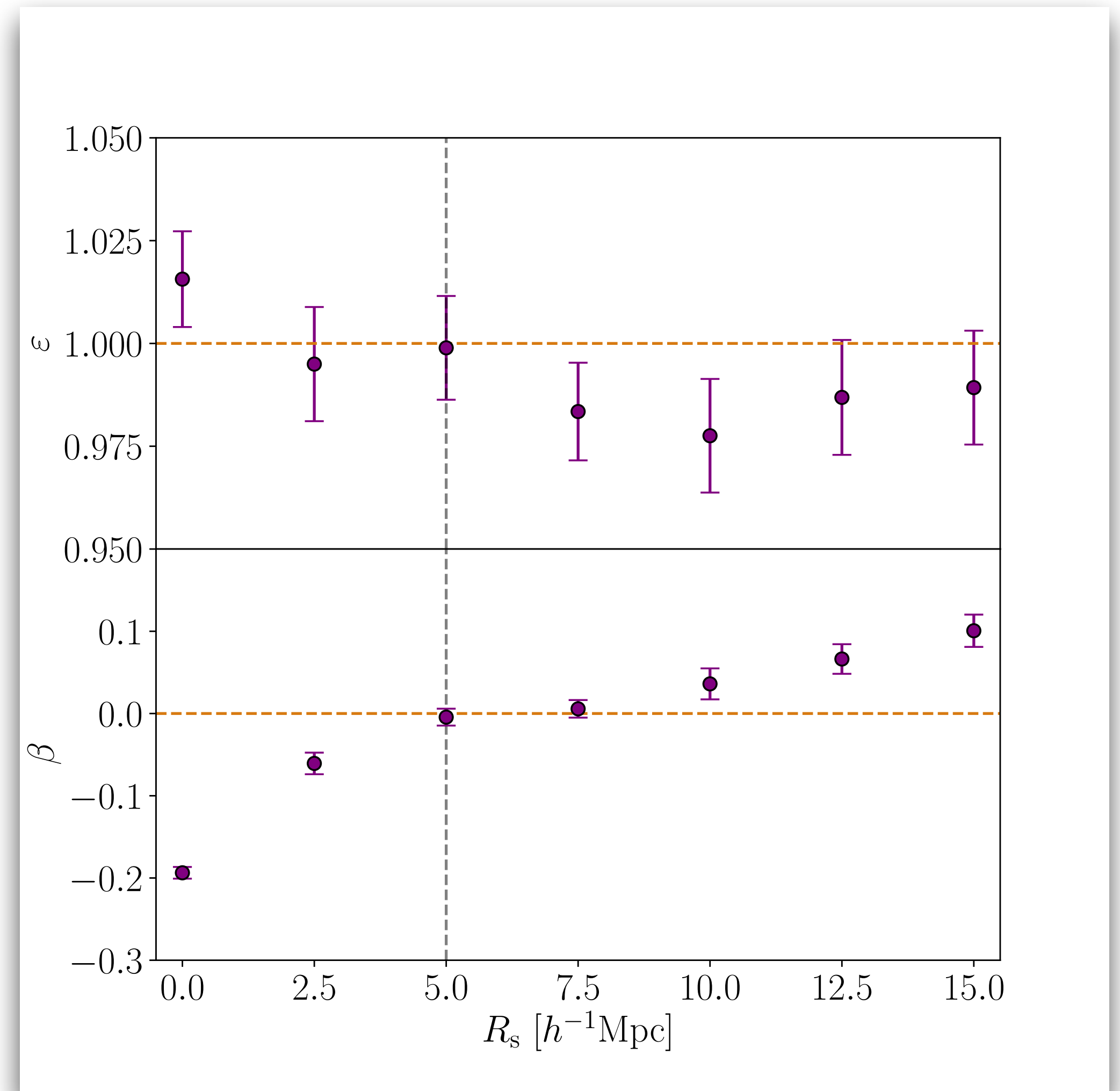
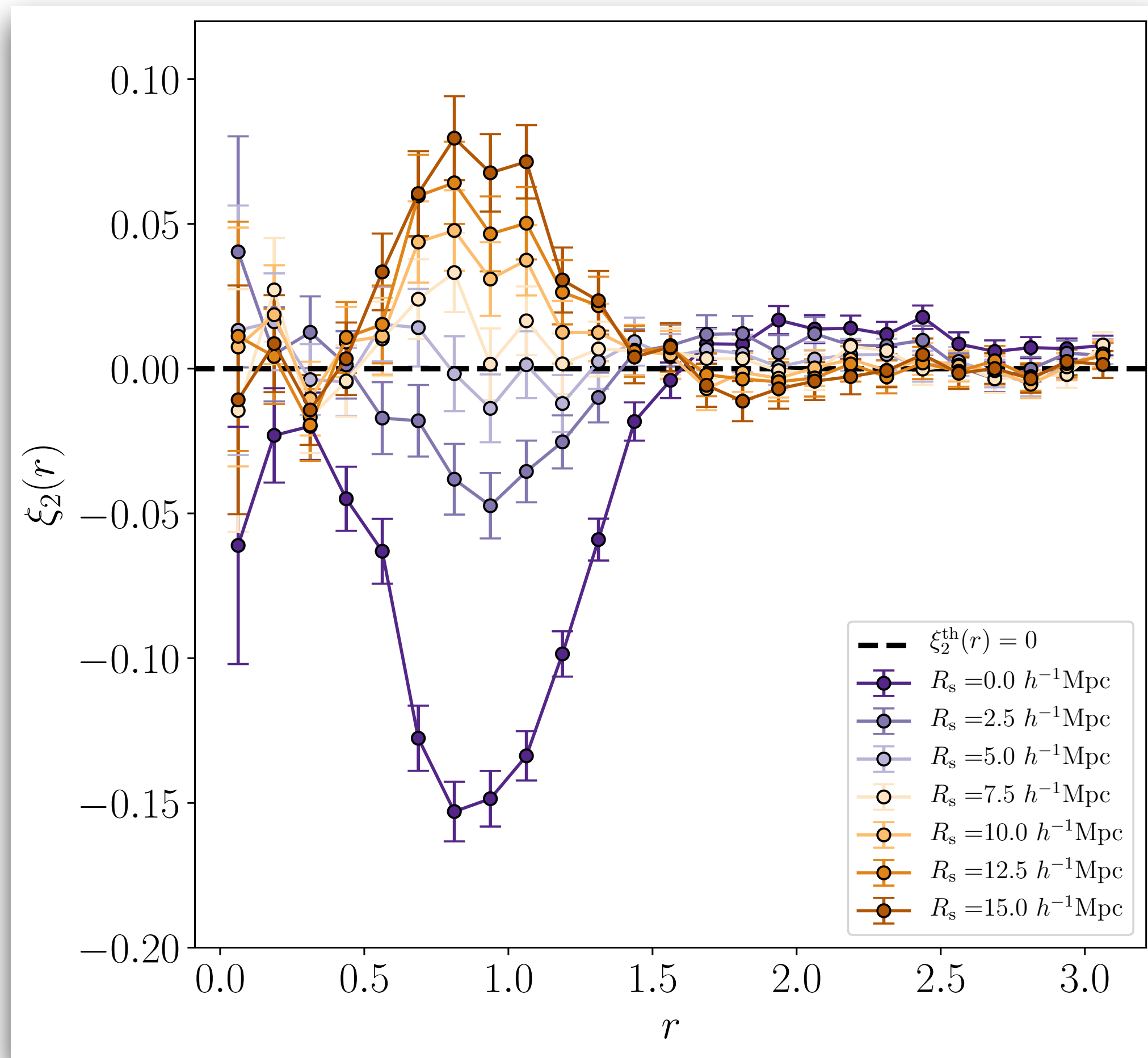
Reconstructed space analysis : more precise in recovering the AP parameter ε



Thank you for the attention :)

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Backup slide: smoothing scale



Backup slide: sensitivity to the fiducial cosmology

