

Rubin-LSST : status and plans

Thibault Guillemin (LAPP)

GDR CoPhy Episode 4
June 1-3 2026
Clermont-Ferrand



The Rubin Observatory

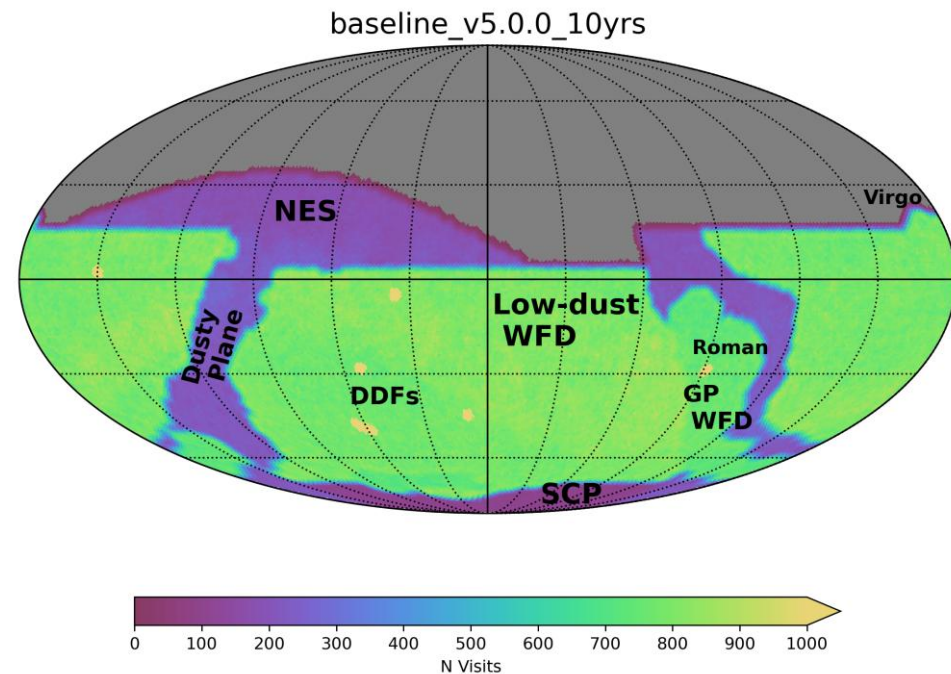
Located on **Cerro Pachón**, in
the Coquimbo region of Chile
Altitude 2647 m



The LSST 10-year survey

The **Wide Fast Deep** philosophy

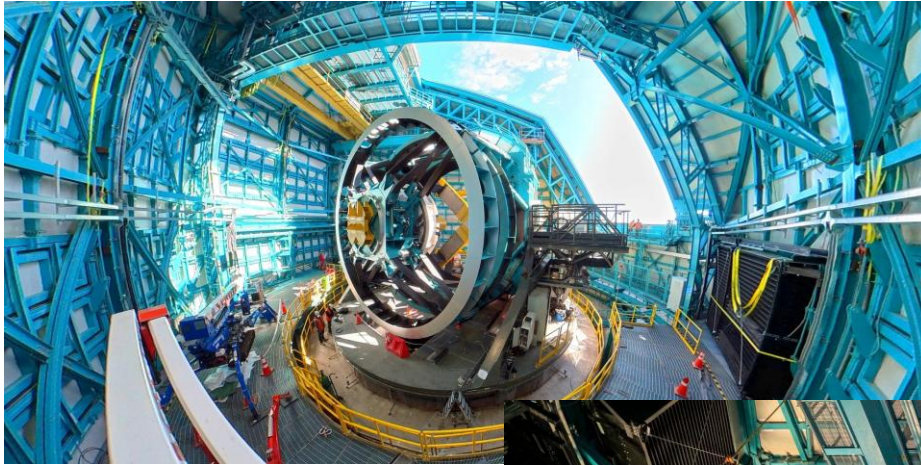
- ❑ Full southern sky $\sim 18,000 \text{ deg}^2$
scanned every 3 nights
30-s exposures $\sim 10 \text{ deg}^2$
- ❑ 6 filter bands (ugrizy) 300-1100 nm
- ❑ 10-year magnitude depths (~ 800 visits)
25.6, 26.9, 26.4, 26.4, 25.6, 24.8
- ❑ Image quality : 0.7 arcsec



Additional programs (<20% of the survey time) :

Deep Drilling Fields, Target-of-Opportunity, Mini-Surveys

Telescope Mount Assembly



2022

Compact steel structure (~300 tons)

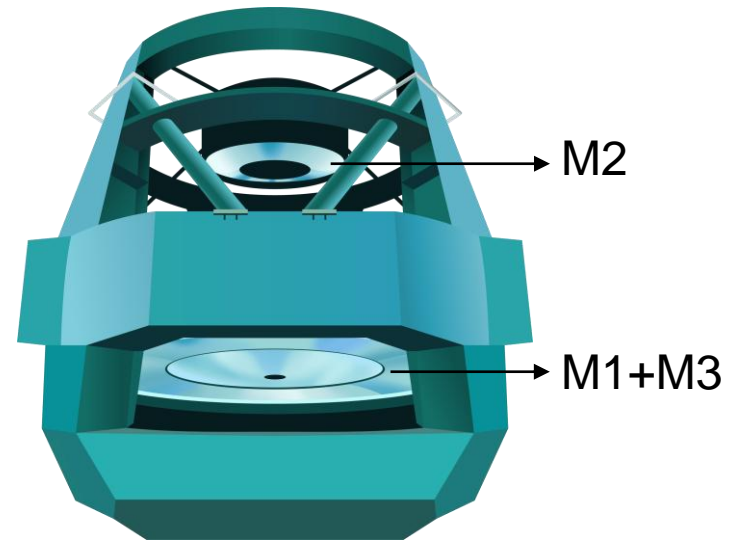
Velocity
~10 °/s (az.)
~5 °/s (el.)

Settling time ~1 s

2025

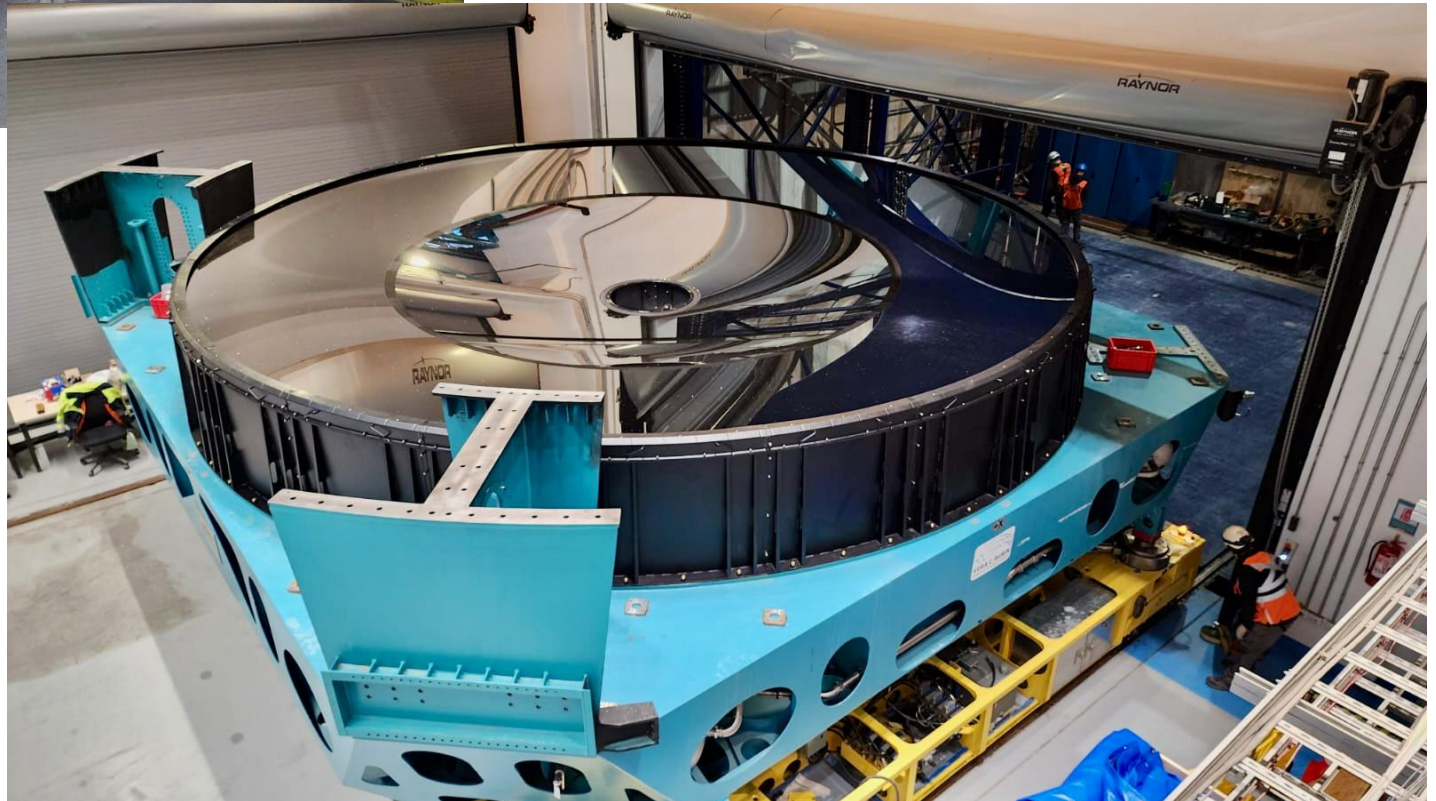


Mirrors

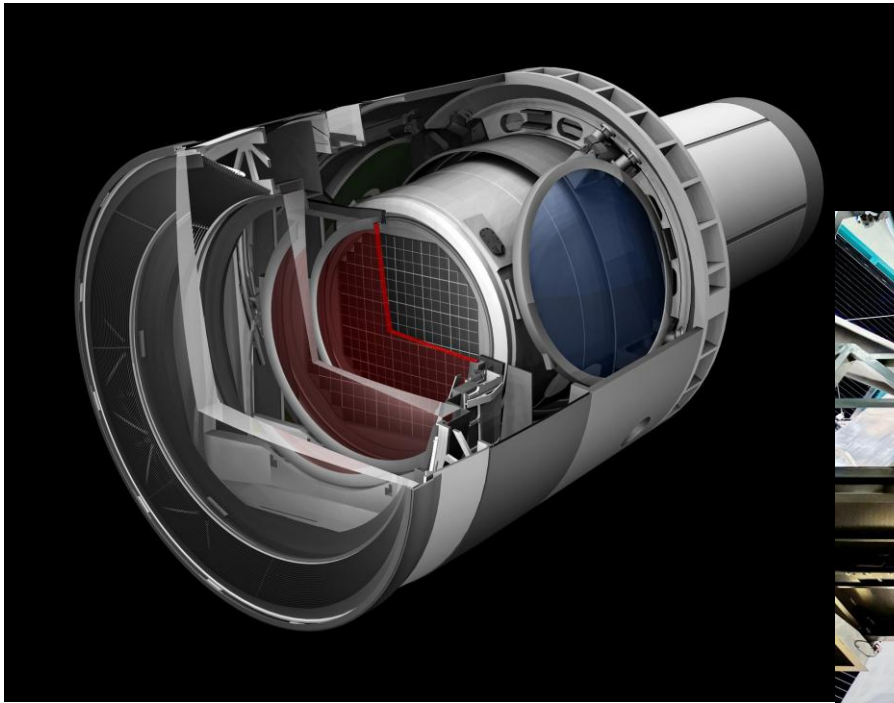


M1 diameter :
8.4 m

Silver



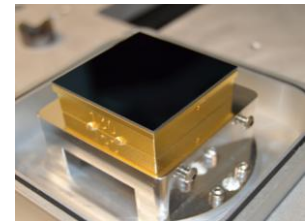
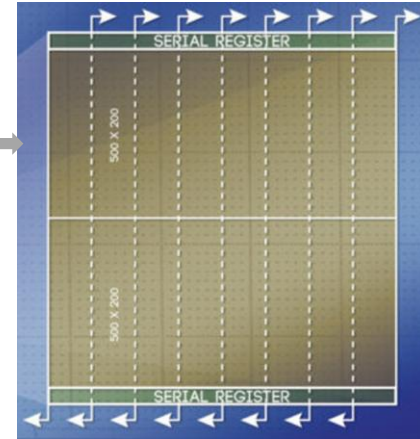
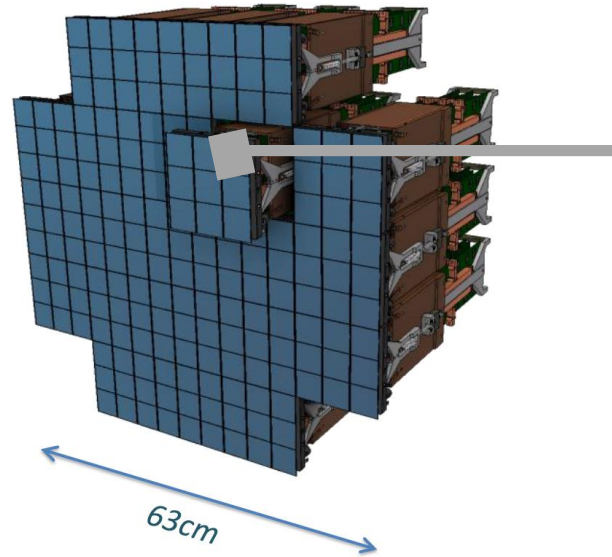
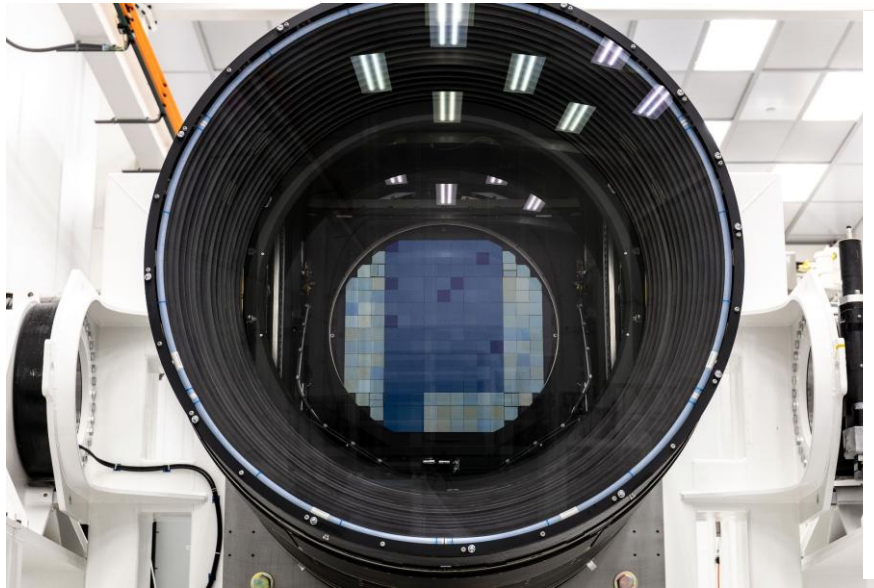
LSSTCam



Diameter 1.6 m
Length 3.7 m
Weight 3 tons



LSSTCam focal plane



CCD ITL

**21 rafts / 1 raft = 9 CCDs / 1 CCD = 16 amp.
189 CCDs / 3024 channels
3.2 Gpixels to calibrate**

Pixel : 10 μ m
0.2 arcsec/pixel
Readout time ~2-3 s

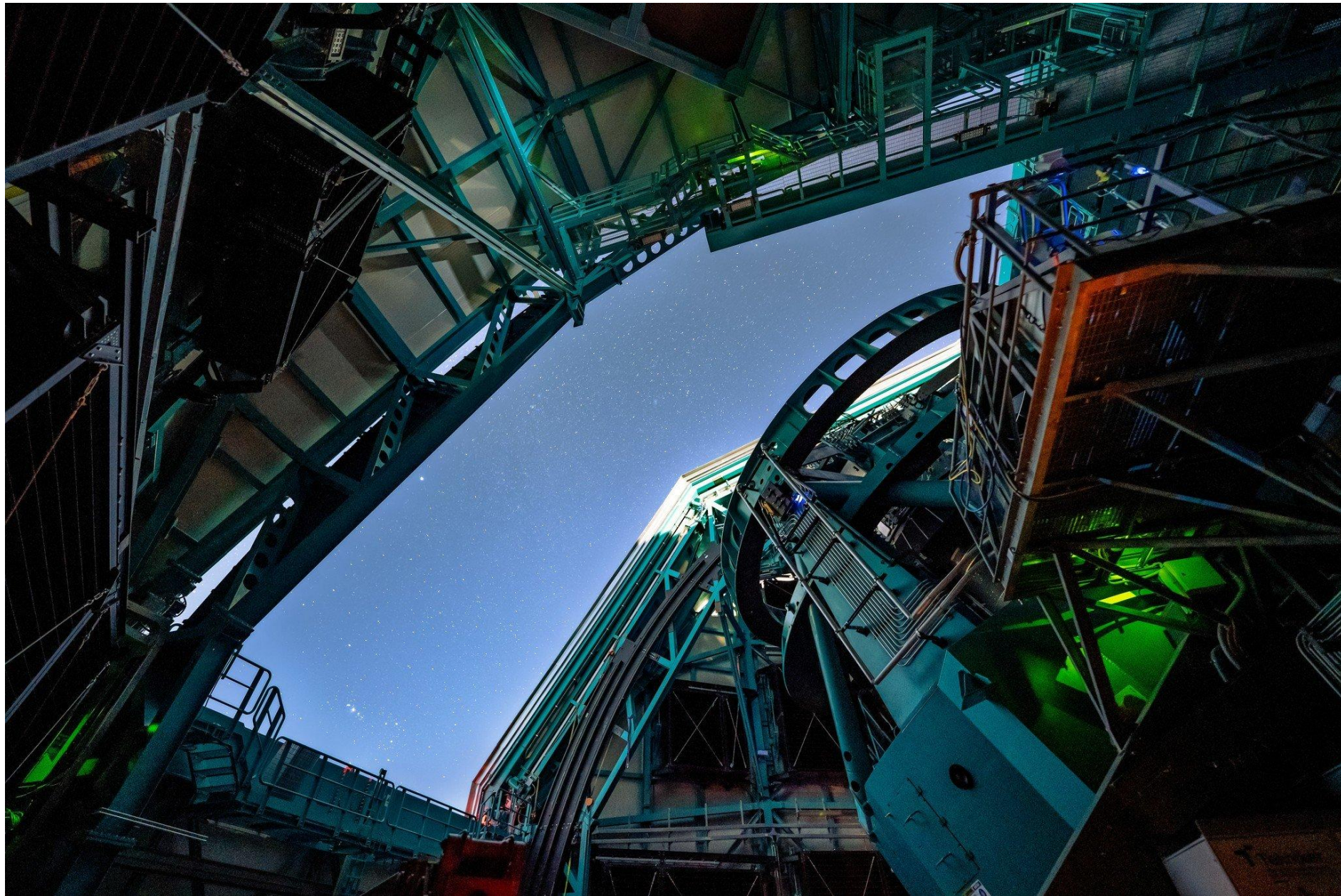
Very complex refrigeration system
Cold system : -40 °C for electronics
Cryostat : -100 °C for CCDs

Data transfer and processing



10 TB/night
~10 million alerts/night
80-hour embargo

Telescope in operation since 1 year! ⁹



Overall organization

Rubin project

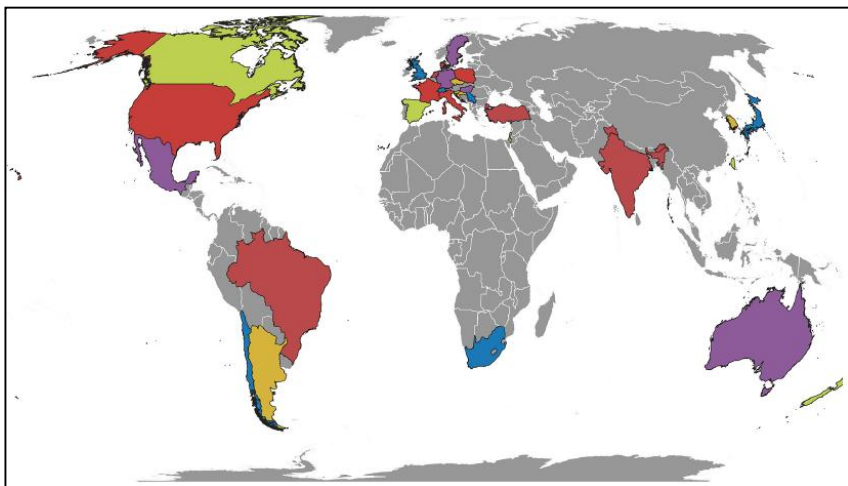
US + Chile (host)

+ France

→ Special case : only non-US major contributor to the construction (camera)

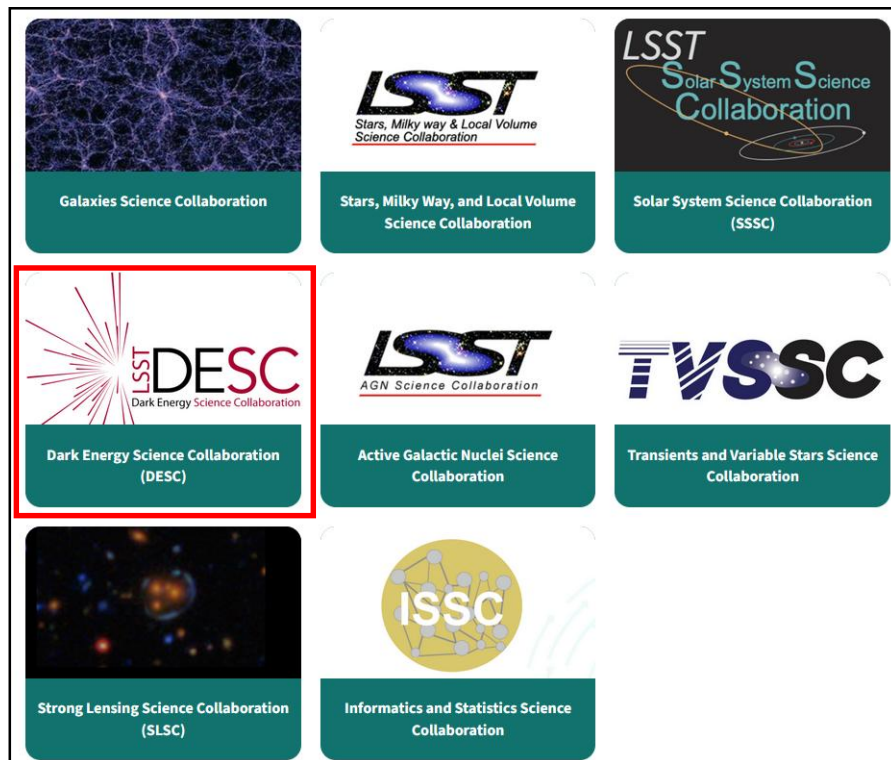
→ Major role in operations : camera and computing (40% of the Data Release Processing)

+ In-kind contributions (to get Rubin data rights)



From 10.5281/zenodo.10393317

Science collaborations



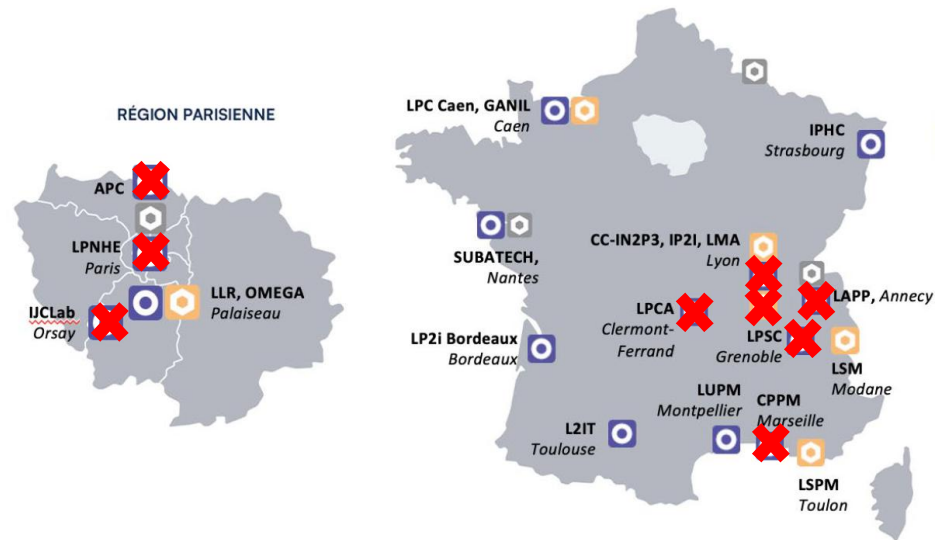
DESC : ~1450 members

LSST-France

~80 researchers at CNRS N&P

~40 IR/IE/T at CNRS N&P

+ now ~20 researchers
at CNRS T&U



CAMERA

COMMISSIONING
PHOTOMETRY

COMPUTING

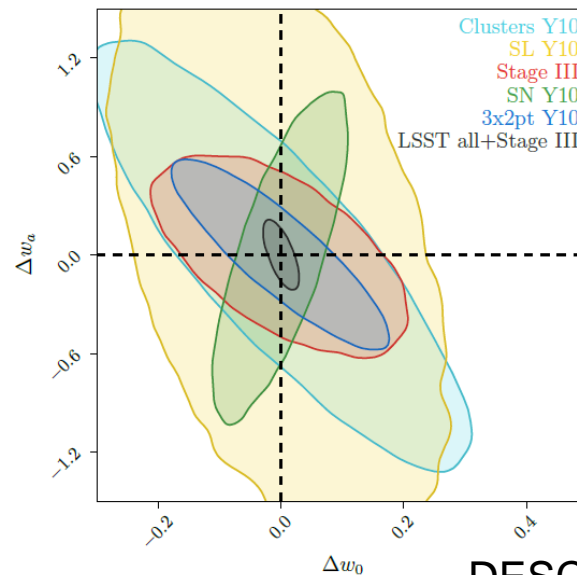
FINK

SCIENCE

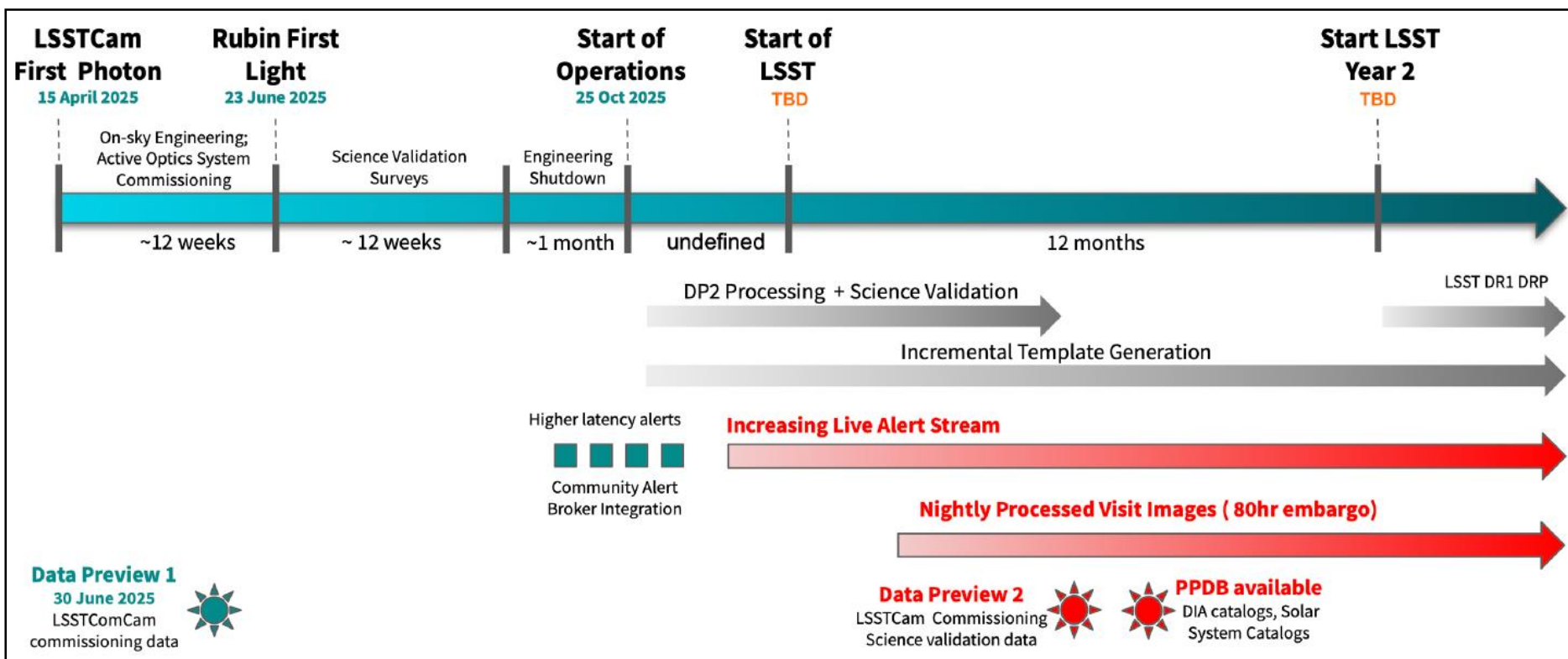
Dark energy probes

- Supernovae
- 3x2-pt
- Galaxy clusters

Preparing
stage-IV
cosmology



Overall planning



- ❑ 2024 : LSSTComCam short campaign
- ❑ 2025 : LSSTCam first photon
- ❑ 2026 : Start of LSST ?

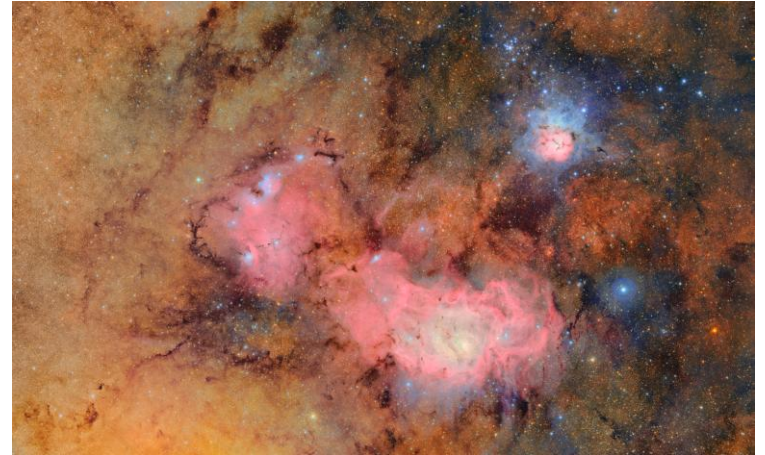
Rubin First Look



Communiqué de presse national

23/06/2025

L'Observatoire Vera C. Rubin dévoile
les premières images du ciel prises
avec la plus grande caméra du monde



<https://rubinobservatory.org/gallery/collections/first-look-gallery>

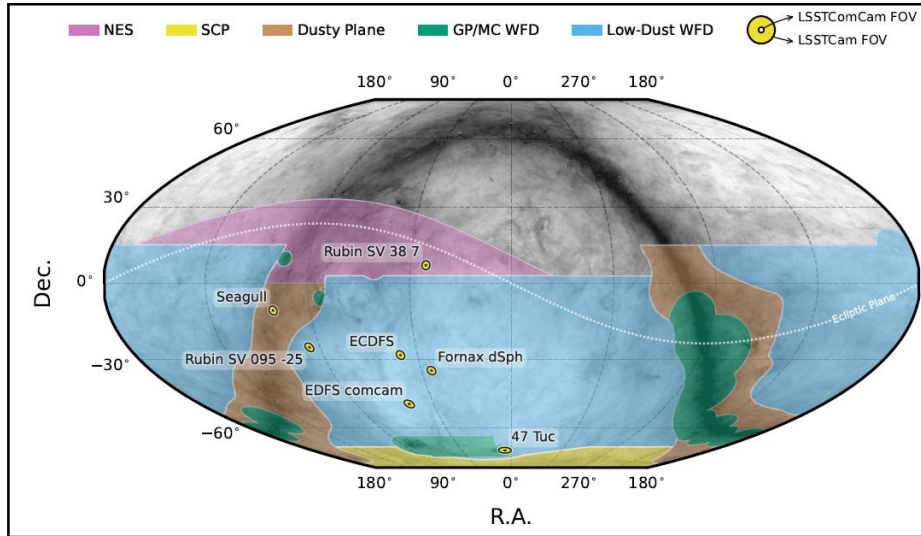
An interlude : Data Preview 1

14

The Vera C. Rubin Observatory Data Preview 1

Paper just published <https://arxiv.org/abs/2603.23786>

(369 authors)

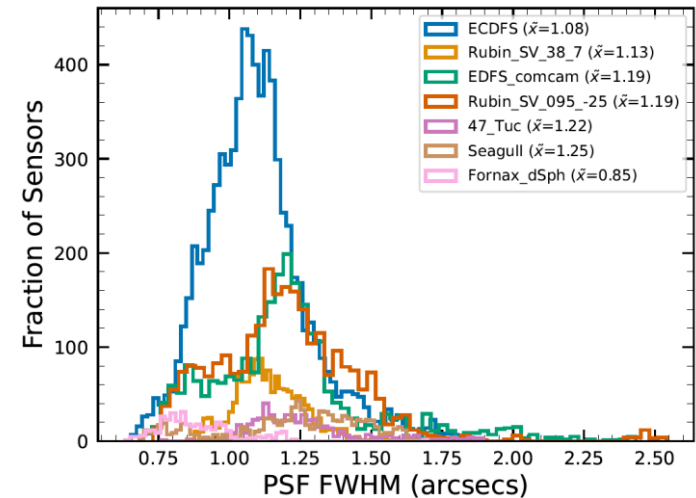
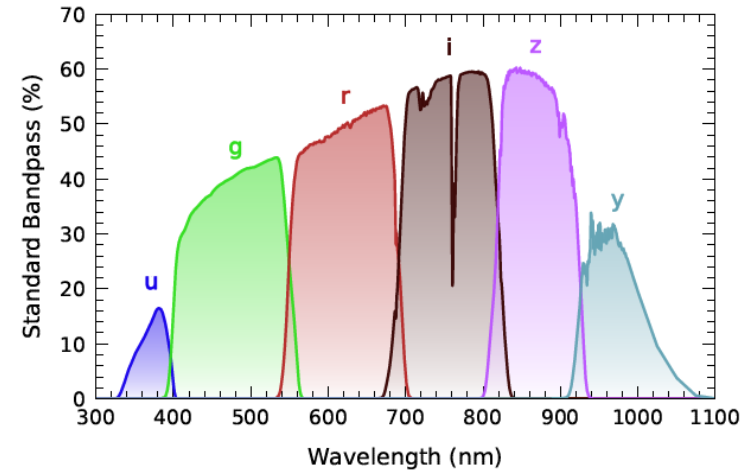


7 fields

- | | |
|------------------|--------------------------------------|
| 47_Tuc | 47 Tucanae Globular Cluster |
| ECDFS | Extended Chandra Deep Field South |
| EDFS_comcam | Rubin SV Euclid Deep Field South |
| Fornax_dSph | Fornax Dwarf Spheroidal Galaxy |
| Rubin_SV_095_-25 | Rubin SV Low Galactic Latitude Field |
| Rubin_SV_38_7 | Rubin SV Low Ecliptic Latitude Field |
| Seagull | Seagull Nebula |

r-band mag depth in range 24.2-26.0

Small-field complete processing exercise



Science with Data Preview 1

Abell 360 for commissioning

Check photometry and photz using red sequence galaxies

Use the cD/BCG to check sky background (over-)subtraction

Check for spurious/missing detections around bright stars

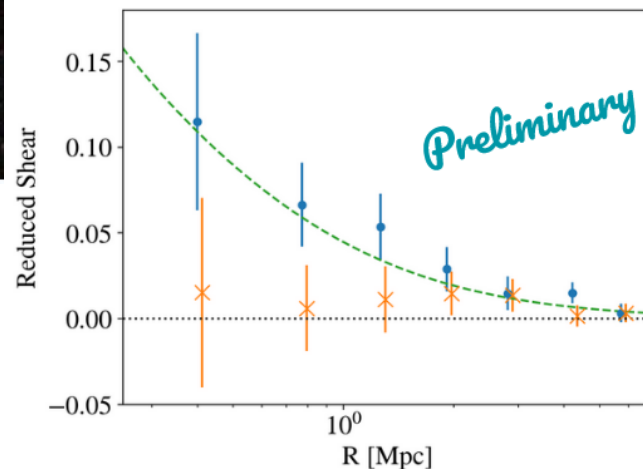
Check PSF modeling in the field

Check deblending

Measure tangential shear signal; check null cross shear signal

A360 shear profiles

Significant detection of tangential shear signal!
Cross-shear consistent with zero!



Slides from C. Combet's talk
at LSST-Europe 2025
Paper in preparation

Data release schedule

Rubin Operations Survey and Data Release Timeline

Event	Date or Date Range	2025	2026	2027	2028
Data Preview 0.1/2/3 (DP0)	2023-06-30				
Rubin First Light (RFL)	2025-06-23	█			
Data Preview 1 (DP1)	2025-06-30	█			
Start of Operations (OPS)	2025-10-25	█			
Rubin First Alerts (RFA)	2026-02-24		█		
PPDB Release (PPDB)	Jun - Sep 2026		█		
Nightly PVIs & Direct-Image Catalogs (NPC)	Jun - Sep 2026		█		
Data Preview 2 (DP2)	Jul - Sep 2026		█		
Start of LSST (SVY)	TBD			🤔	🤔
Data Release 1 (DR1)	TBD			🤔	🤔

J F M A M J J A S O N D J F M A M J J A S O N D J F M A M J J A S O N D J F M A M J J A S O N D

Until recently :

DR1 = LSST Y1 dataset → Planned for **early 2028**

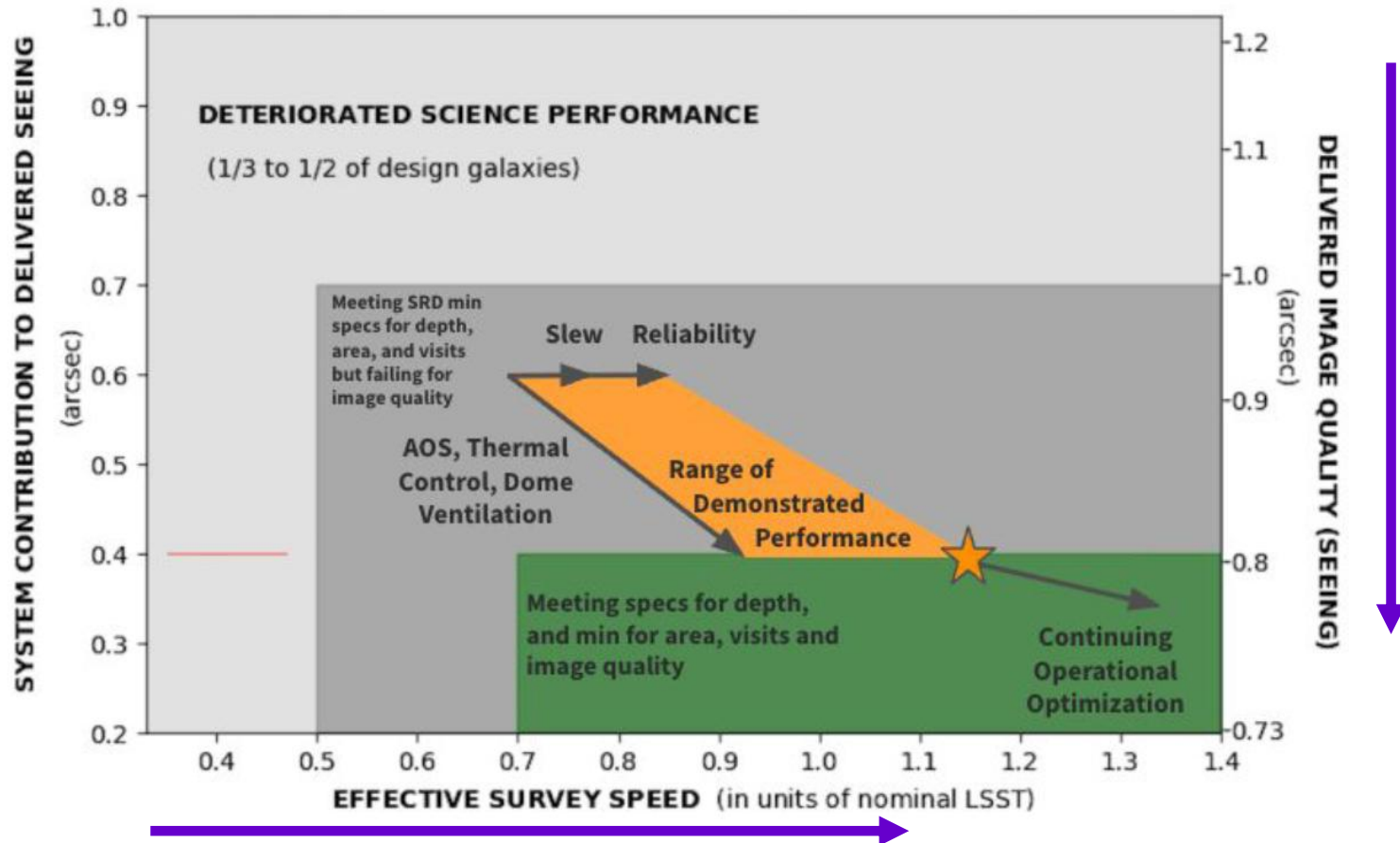
****Assuming a start of the survey at the end of 2025****

Why has the survey not started ?¹⁷

Official statement (early May) :

“We expect to start the LSST in the next months.”

Schematic of the path towards science requirements



Huge work on efficiency, reliability and image quality

Ongoing work

Finalization of the dome construction

Installation of 12 louvers



From J. Neveu

Installation of light wind screens



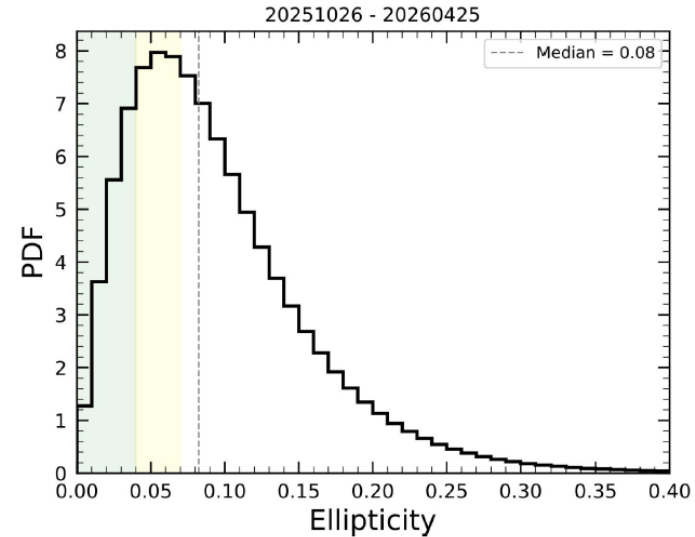
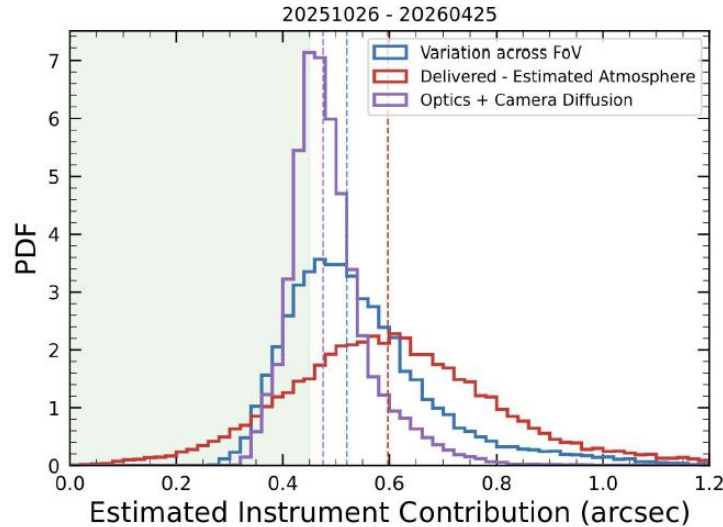
From A. Alexov

+ Work on M1M3 thermal control

Goal : improve dome seeing

From Rubin All Hands
meeting, 14/05/26

Monitoring image quality metrics and PSF ellipticity



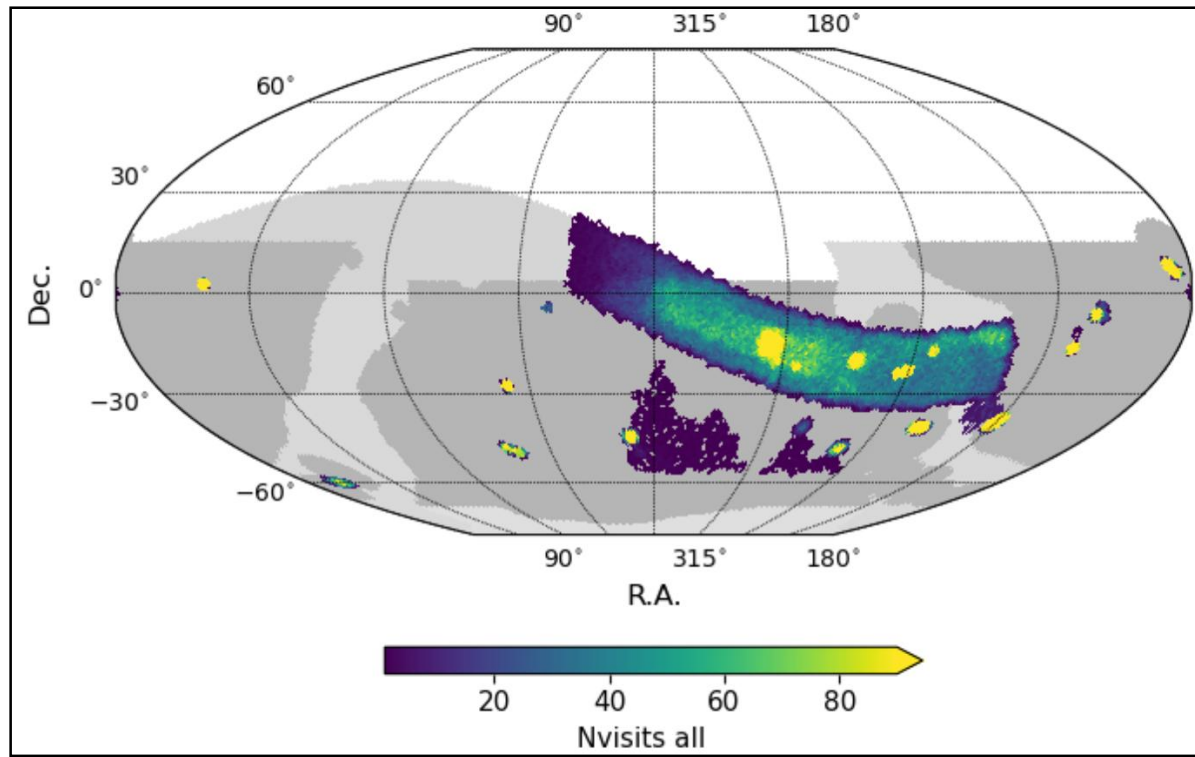
Continuous work :
active optics system, mirror bending
modes, in-dome environment control,
wavefront sensors, etc.

**Image quality and PSF
ellipticity are getting closer
and closer to requirements**

The good news : Data Preview 2²⁰ is coming !

Official statement : “Release planned in July-September”

21,647 science visits in 6 bands



Wide Survey ~750 deg²
Pipelines, templates,
science validation

Deep Survey
Coadds at ~LSST 10-year
depth

More details : <https://rtn-011.lsst.io/>

The good news : Data Preview 2 is coming !

DP2 processing ~3 months

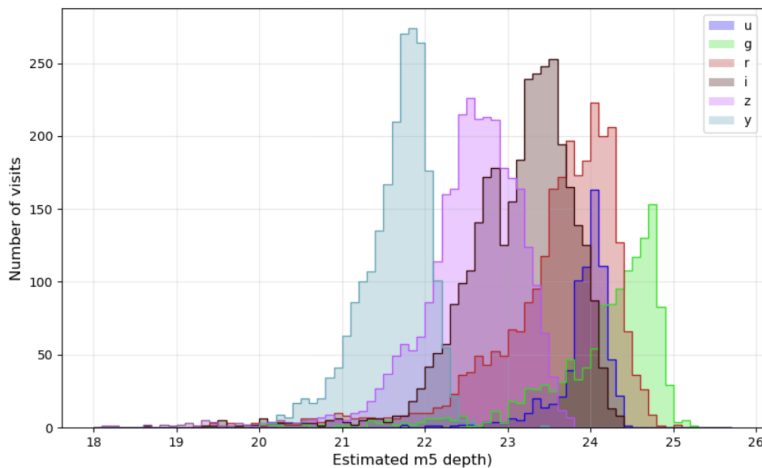
**Reminder : DP2 release
= release to LSST data
right holders
(proprietary period of
2 years)**

Data Product
Raw Images
DRP Processed Visit Images and Source Catalogs
DRP Coadded Images and Object Catalogs
DRP Cell-based Coadded Images and ShearObject Catalog
DRP ForcedSource Catalogs
DRP Difference Images and DIA Catalogs
DRP SSP Catalogs

Photo-z catalog later

Wide Survey

Field	Band						Total N visits
	<i>u</i>	<i>g</i>	<i>r</i>	<i>i</i>	<i>z</i>	<i>y</i>	
750 deg ²	2	4	12	16	11	9	56
	24.4	25.2	25.2	24.8	24.0	23.0	
Median FWHM (arcsec)	1.18	1.26	1.26	1.24	1.17	1.25	



What about alerts ?

**NSF-DOE Vera C. Rubin Observatory
Launches Real-Time Discovery Machine for
Monitoring the Night Sky**

**800.000 science alerts
the night of 24 February**

DDFs first (templates)

Alert-driven science is already happening.

Bellm

Showing results 1 to 50 out of 231

1 2 3 4 5 > >>

ID	Name	Reps	Class	RA	DEC	Obj. Type	Redshift
208084	AT 2026lui	1		12:23:54.129	+08:39:10.09		
207850	AT 2026llp	1		10:00:13.680	+03:40:20.55		
207848	AT 2026lln	1		12:21:27.344	+07:15:12.57		
207828	AT 2026lkw	1		10:05:46.584	+03:00:31.79		
207802	AT 2026lka	1		09:57:15.877	+00:40:00.74		

[Previous]

AT 2026epw: Discovery of a possible nova from Vera C. Rubin Observatory

ATel #17703; *Anrudh Saigundi (UNC), Igor Andreoni (UNC), Akash Anumarpudi (UNC), James Freeburn (UNC), Theophile du Laz (Caltech), Michael Coughlin (UMN), V. Ashley Villar (Harvard), Eric Burns (LSU), Jonathan Carney (UNC), Mansi Kasliwal (Caltech).*

on 27 Feb 2026; 21:00 UT

Notification: Akash Anumarpudi (akasha@unc.edu)

AstroNote 2026-45 Type: Object/s-Data/Analysis

Spectroscopic Observations of AT 2025agpz, a Long-Rising Hydrogen-Rich Transient in Rubin-LSST Commissioning Data

C. R. Angus (QUB), H. Stevance (Oxford), M. Nicholl (QUB), K. W. Smith, S. Smartt (Oxford/QUB), D. R. Young, M. Fulton,...

AstroNote 2026-50 Type: Object/s-Discovery/Classification

Identification of Rubin Supernova Candidates with Superphot+

Kaylee de Soto (Harvard/CfA), Yize Dong (Harvard/CfA), Anya Nugent (Harvard/CfA), S. Karthik Yadavalli (Harvard/CfA), V...

SDSS J095715.86+004000.9

AstroNote 2026-46 Type: Object/s-Data/Analysis

AT 2025agpz / ATLAS25pny / LSST-P-DO-313761043604045880: SLSN candidate in the first Rubin / LSST alerts

I. Pérez-Fourmon (IAC and ULL), F. Poidevin (IAC and ULL), D. Cano-Morales, A.E. Hernández-Díaz, I. Correa-Plasencia, E...

AstroNote 2026-53 Type: Object/s-Data/Analysis

Late-Time Detection of the Ca-rich SN Ia-pec 2024kce in Early LSST Alerts

From Rubin All Hands meeting, 14/05/26

Alert broker : FINK

From J. Peloton

Fink science portal :

<https://lsst.fink-portal.org/>



2026-05-27

Last LSST observing night

29,203

Alerts processed

11,006,068

Total alerts

5,724,965

In catalogs

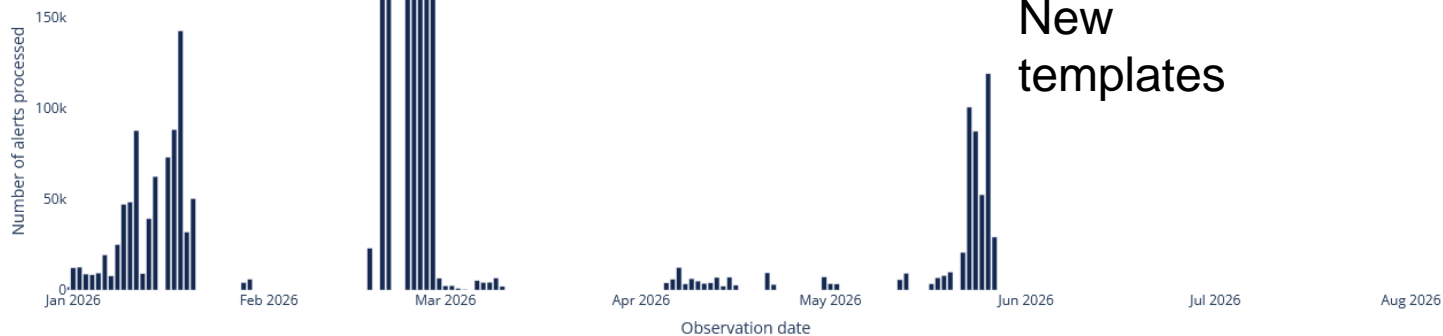
2026



Number of alerts processed

Cumulative

Percentage



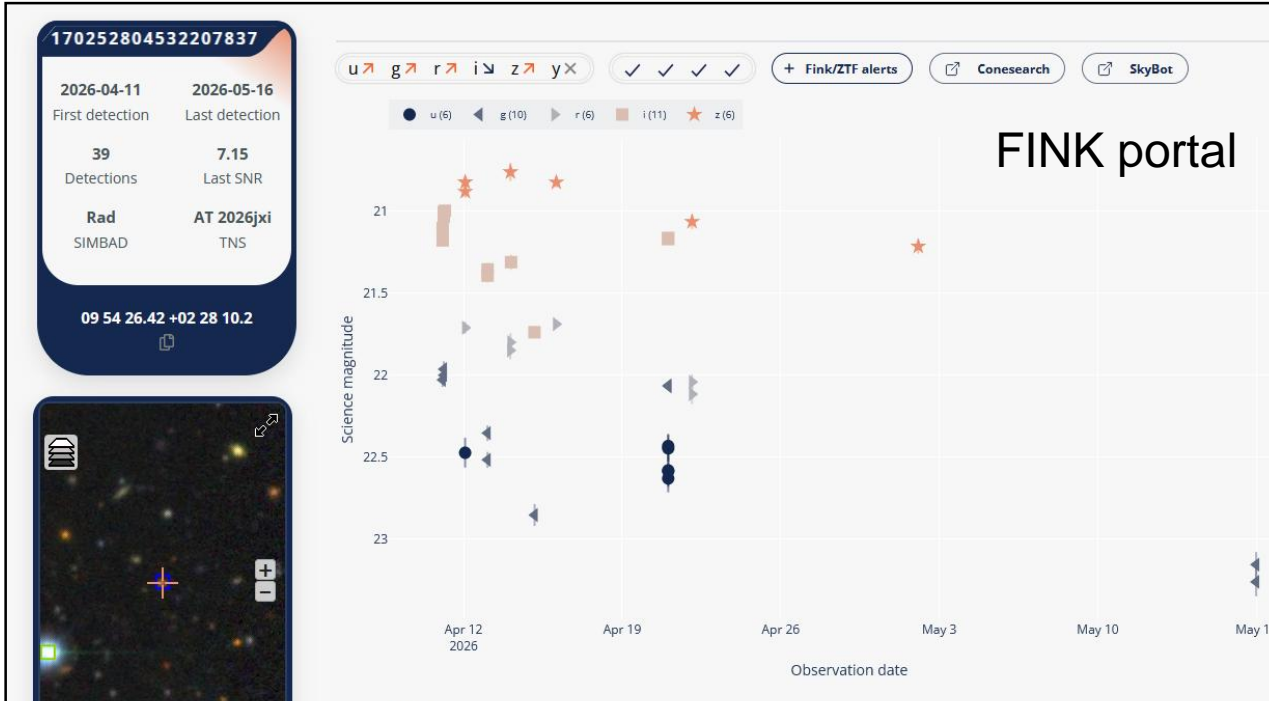
Tests



Real alerts

Alert follow-ups

Thanks to S. Antier for the information



COLIBRI telescope



Sarah Antier

Thu, 04/23/2026 - 17:20

AT 2026jxi: COLIBRI follow-up report

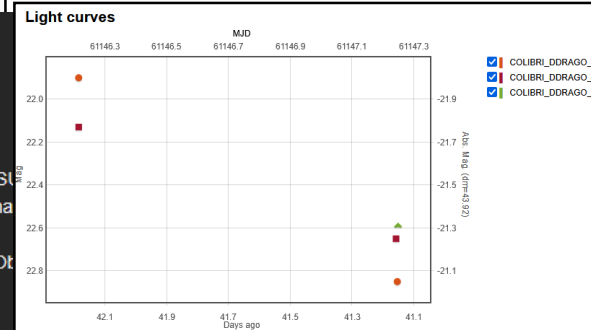
D. Aki (NYUAD), P. Hello (IJCLAB), A. Jacquesson (IJCLAB), R. Hanna (IJCLAB), N. Jasmin (Uljege), L. Chen (TNS) (IJCLAB), S. Antier (IJCLAB), M. Pillas (IAP), M. Coughlin (UMN), A. LeCalloch (UMN), T. DuLaz (Caltech) on behalf of the TNS team

We observed the field of the [2026jxi](#) event with the DDRAGO wide-field imager on the COLIBRI telescope at the Observatorio de San Pedro Mártir, in two consecutive nights on 2026-04-16 and 2026-04-17 in g,r,z filters.

We detected the source with the following magnitudes after subtraction from the template flux reported in <https://www.wis-tns.org/object/2026jxi>

The data were coadded with analysed in STDWeb/STDPipe (Karpov 2025). Comparing our observations (and after performing image subtraction) against Legacy Survey DR10.

However, we observed a re-brightening in the g-band after 6 days post discovery compared to measurements taken by Vera Rubin. This was re-confirmed by further LSST measurement 2 days after. Currently, taken into account the properties of the source in the template, and the light curve of the transient, we cannot determine the class of this transient : a type of AGN ? a type of supernova ?



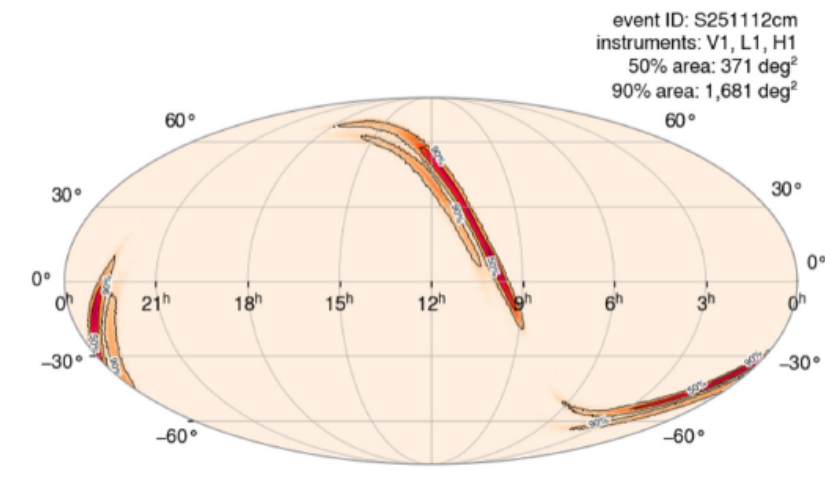
ToO : gravitational waves

Two LVK GW candidates selected for ToO :

- 1) S250725j (July 25 2025)
- 2) S251112cm (November 12 2025)

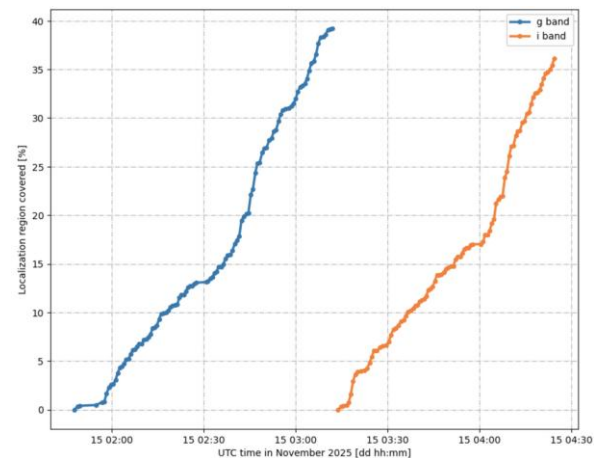
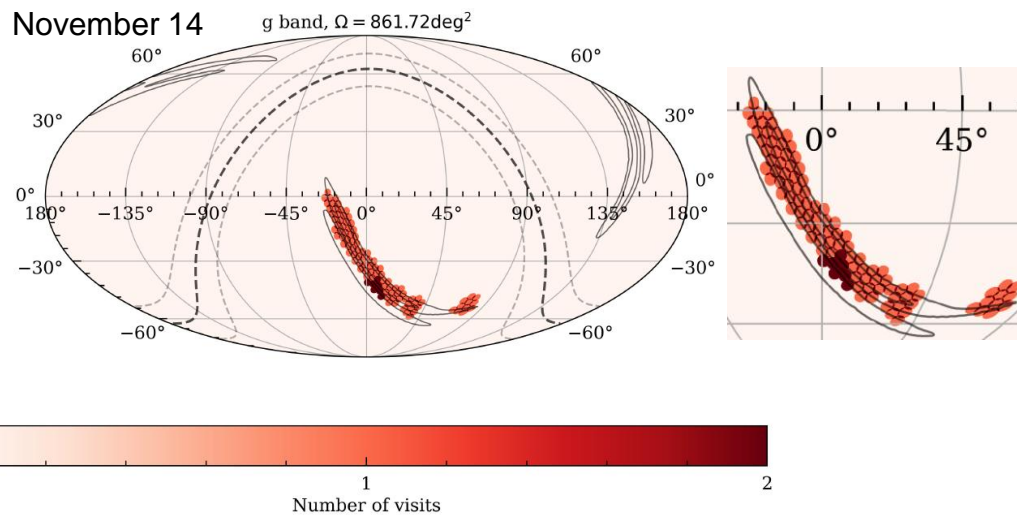
S251112cm discovery

- On 2025-11-12 LVK detected a sub-solar mass merger candidate, **S251112cm**
- **HasSSM > 99%** and **Mchirp** between [0.1, 0.87] Msun
- False Alarm Rate: 1 per 4 yrs
- Localized to **1681 sq. deg** at a distance of **93 +/- 27 Mpc**
- Many teams followed up the event, including GO-TO, BlackGEM, ZTF (GCN 42677; SA+2025) and DECAM



ToO : gravitational waves

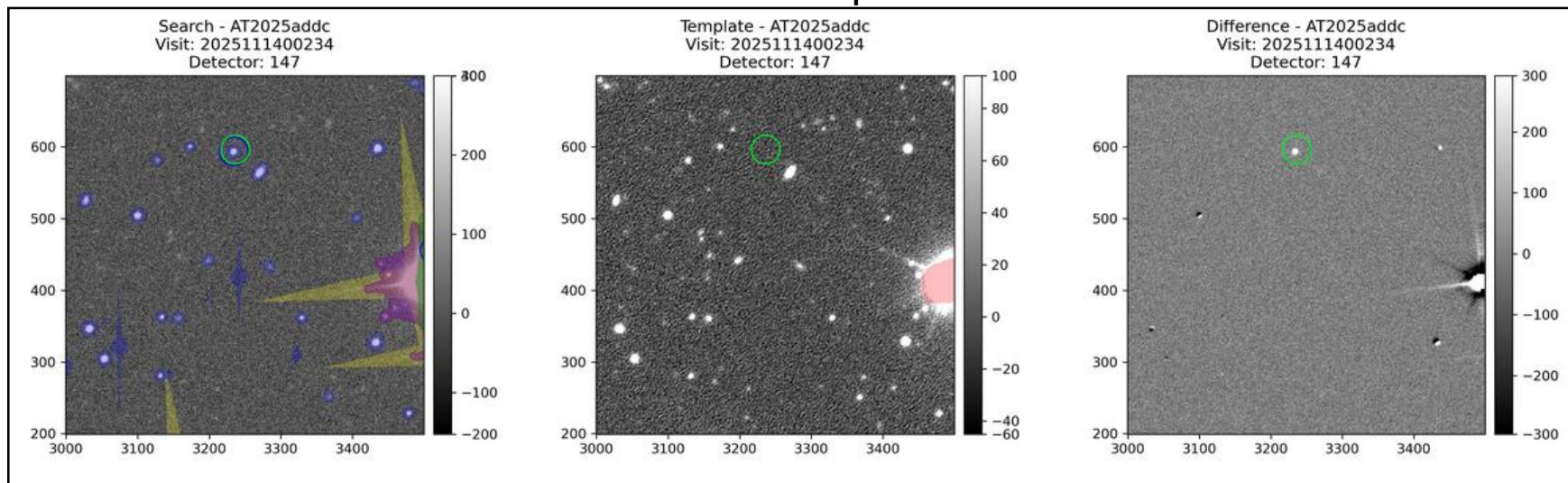
Thanks to S. MacBride
and the Rubin ToO team



~850 deg² in 2 bands in 3 hours
Depths : g 24.5 and i 23.5

Wide Fast Deep in action !

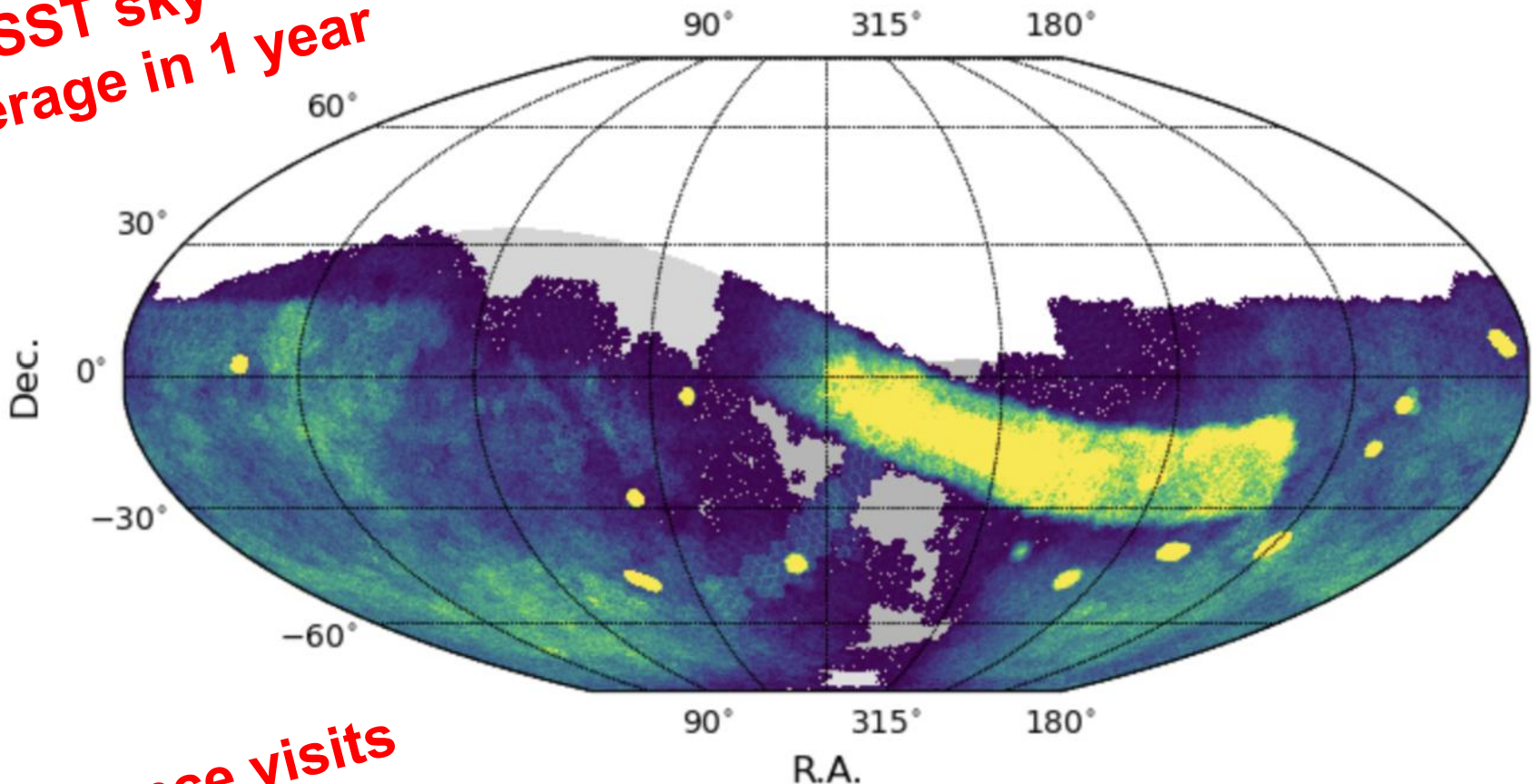
Search for EM counterpart candidates



Conclusion

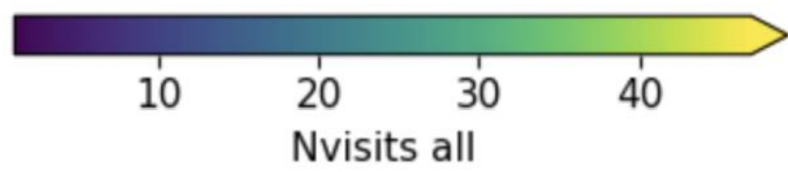
From Rubin All Hands meeting, 14/05/26

PreLSST 20250424-20260429 visits



PreLSST sky coverage in 1 year

>50k science visits



The end

