

Classicalization, a Higgsless alternative

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work (in progress) in collaboration with
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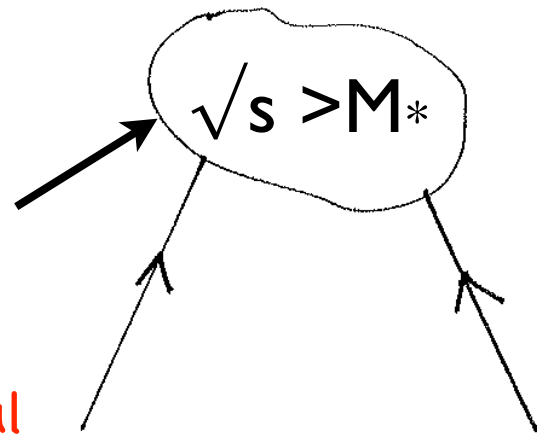
What is classicalization ?

- In QFTs with a **cut off** (like the Higgsless SM) **2→2 scattering amplitudes blow up** at energies near the cut off. Standard (Wilsonian) interpretation => **New Physics** at these scales.
- **Classicalization** is an alternative way of unitarizing these amplitudes in some cases. Inspired by **black hole** physics.

- Take the following simple example of scalar field,

$$\mathcal{L}(\phi) = (\partial_\mu \phi)^2 + \frac{1}{M_*} \phi (\partial_\mu \phi)^2$$

Energy here is like the '**charge**' in electrostatics. It generates a **classical configuration** of ϕ around it.



- **2→2 scattering** gets **suppressed** at high energies. We get instead **2→ Classicalon**

Dvali, Giudice, Gomez
and Kehagias 2010

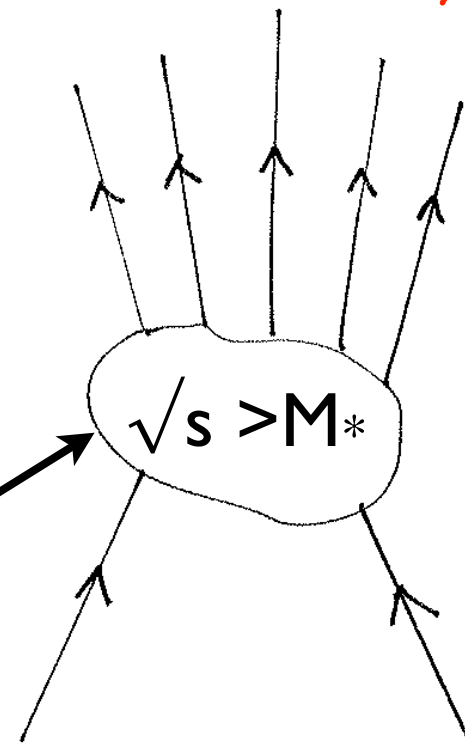
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many quanta
(defines classicality)



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Dvali, Giudice, Gomez
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Unitarizing WW scattering by classicalization

- If there is **no Higgs**, $W_L W_L \rightarrow W_L W_L$ scattering amplitude blows up at around 700 GeV.
- If we add appropriate interactions the **longitudinal Ws and Zs can classicalize. Even more minimal than SM!** For eg. lets add a classicalizing interaction to Higgsless SM,

$$\mathcal{L} = -\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}m_W^2 W_\mu W^\mu + \frac{g^4}{4} (W_\mu W^\mu)^2$$

- With energy size of classicalon increases (like Black Holes). In theories that exhibit classicalization **higher scattering energies does not mean probing shorter distances.**

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A classicalizing interaction which gives classicalons with radii,

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(As $W_L^\mu \rightarrow \frac{\partial^\mu \phi}{m_W}$, ϕ being the 'goldstone')

$$r_* \sim \frac{\sqrt{s}^{1/3}}{v^{4/3}}.$$

the exponent depends on the particular interaction, **=1 for black holes**

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How do we make experimental predictions ?

- The process we can look for in colliders would be:

Weak Boson Fusion \rightarrow Classicalon \rightarrow multiple W/Z

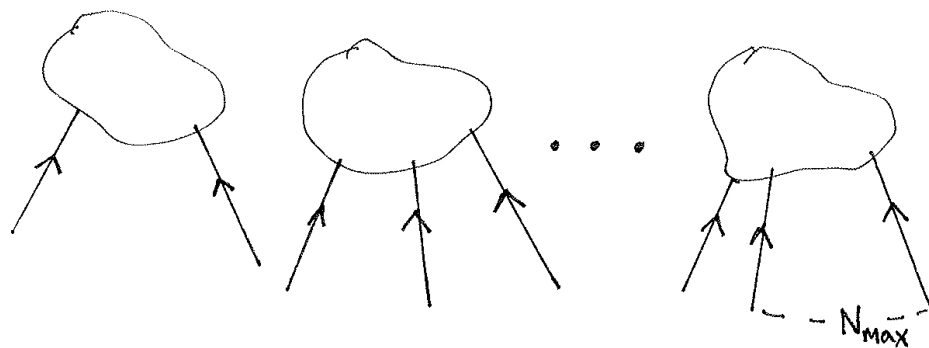
- Cross-section is just the geometric cross-section: $\sigma = \pi r^2$.

unlike black holes, these classicalons would decay only to Ws/Zs

- Experimentally relevant question: What happens **after** we produce a classicalon? What is the **number of particles classicalon decays to** ? What is the **energy distribution** of the decay products?

Classicalons have thermodynamic properties!

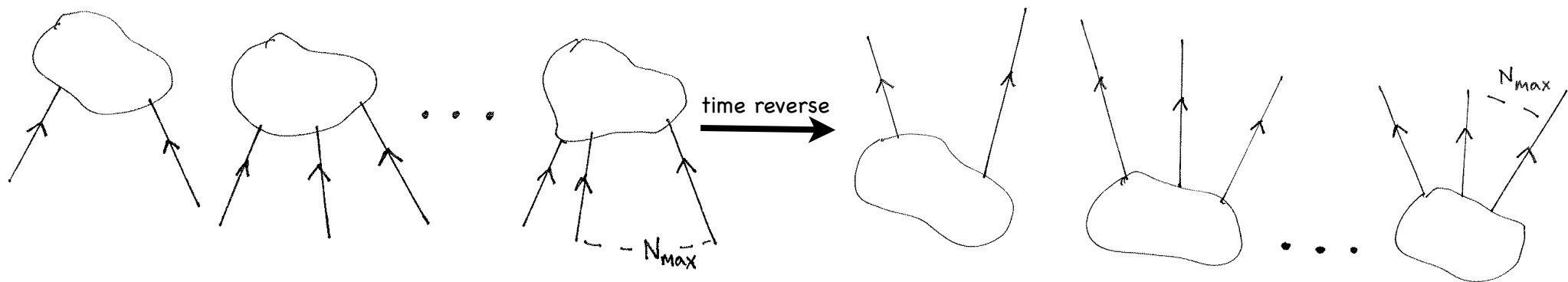
- Every possible way of localizing the required energy in the classicalon radius r^* must form a classicalon, thus **classicalons can be formed from 2,3,4..... N_{\max} particles** (minimum energy of a particle $1/r^* \Rightarrow N_{\max} = \sqrt{s}/(1/r^*) = r^* \sqrt{s}$).



- By **time reversal** symmetry therefore **classicalons can decay into 2,3,4... N_{\max} particles.**
- Classicalons would **rarely decay to 2 particles** as there are **many more ways of decaying to many particles** simply because of combinatorics \Rightarrow **2 \rightarrow 2 scattering suppressed**. We see how **notion of entropy** arises for classicalon.

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Classicalons are Bose Einstein systems

- **Every** possible way of **localizing** the required energy in the classicalon radius r^* must form a classicalon. How do we compute the number of possible ways?
- We must ensure,
 - (1) **Condition 1: energy conservation**, $|\vec{k}_1| + |\vec{k}_2| + \dots + |\vec{k}_N| = \sqrt{s}$
 - (2) **Condition 2**: that all the **energy gets localized inside**, for eg. scattering plane waves would not localize the energy inside r^*
- All possible ways of choosing the momenta \mathbf{k}_i (without any other restriction) respecting energy conservation gives the **Bose-Einstein distribution**,

$$N(\omega)d\omega = \frac{g(\omega)d\omega}{e^{\beta\omega} - 1}$$

- We cannot find the **temperature** ($T = 1/\beta$) and **total number of particles** without knowing the density of states.

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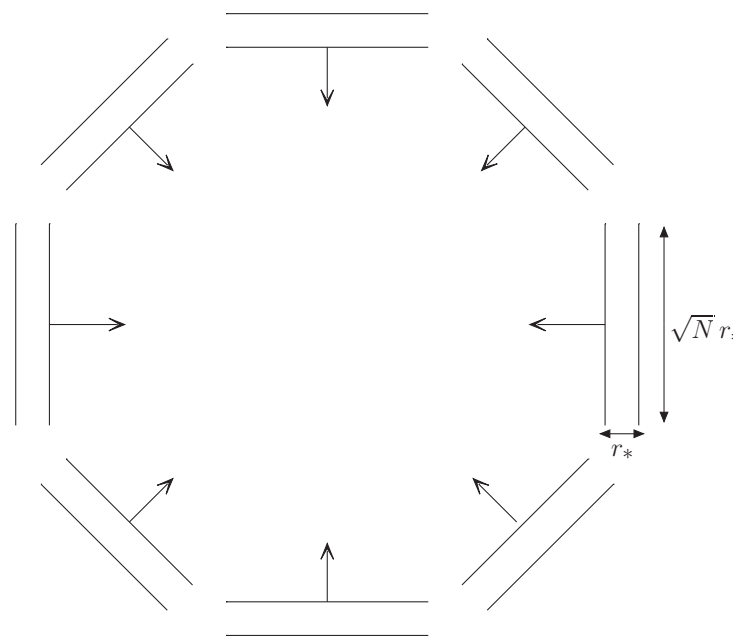
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The Density of states

- In order to localize all the energy in the classicalon radius wavepackets must have size at most r_* in the longitudinal direction.
- The wavepackets can be much longer in the transverse direction, of size $\sqrt{N} r_*$. This leads to a $\Phi \sim 1/r$ field outside the classicalon radius r_* . Most of the energy is still inside the classicalon radius r_* .



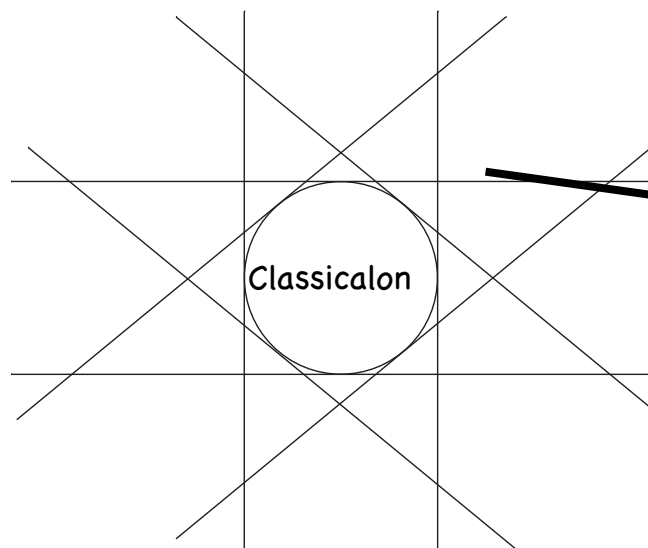
- For such wavepackets density of states function is:

$$g_\omega d\omega = \frac{1}{8\pi^3} (2r_*) d\omega \times L^2 (4\pi\omega^2) = \frac{Nr_*^3 \omega^2}{\pi^2} d\omega$$

- Classicalons like massless Bose-einstein system but with different density of states from blackbody radiation.

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Outside r^* the wavepackets superpose to give **$\Phi \sim 1/r$**

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crucial difference from blackbody radiation

- Classicalons like **massless Bose-einstein system but with different density of states** from blackbody radiation.

Temperature, decay multiplicity and entropy

- With this density of states function and the distribution function we obtain the **temperature, number of decay particles and entropy**.

$$\frac{N_* r_*^3}{\pi^2} \int_{\omega=\pi/r_*} \frac{\omega^2 d\omega}{e^{\beta\omega} - 1} = N_*$$
$$\frac{N_* r_*^3}{\pi^2} \int_{\omega=\pi/r_*} \frac{\omega^3 d\omega}{e^{\beta\omega} - 1} = \sqrt{s},$$

$$\beta \sim r_* \Rightarrow T \sim \frac{1}{r_*}$$

$$N_* \sim \sqrt{s} r_*$$

$$S \sim N_*$$

- For the **black hole** case we recover the well known **proportionality of the entropy with the area**.

$$S \sim N_* \sim \sqrt{s} r_* \sim M_{pl}^2 r_*^2$$

Temperature, decay multiplicity and entropy

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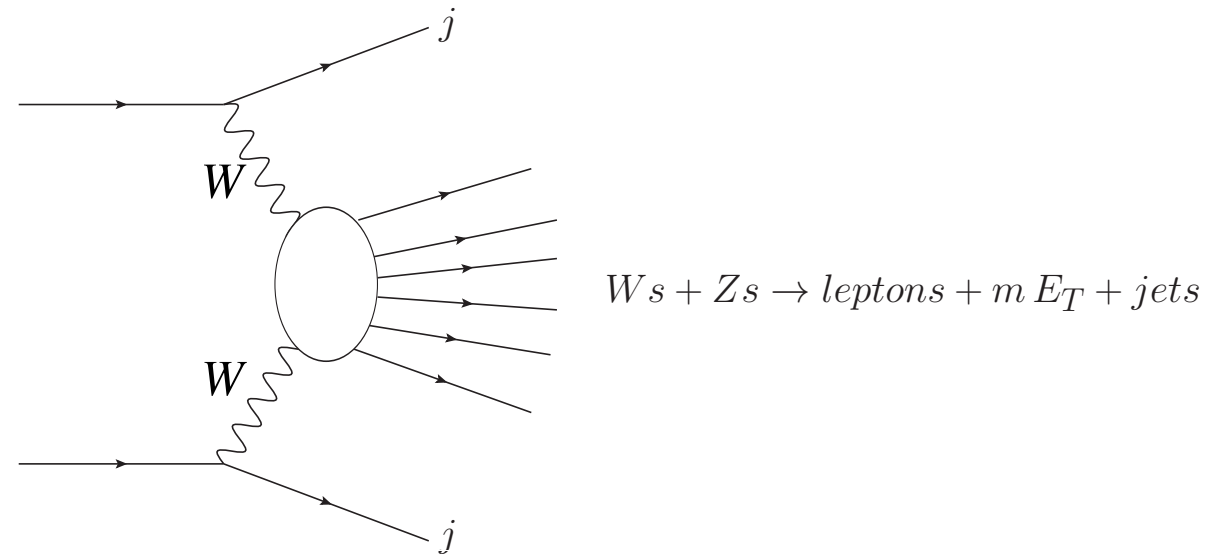
For these classicalons
 $N_* \sim (\sqrt{s})^{(4/3)}$ whereas
 for black holes,
 $N_* \sim (\sqrt{s})^2$.

- For the **black hole** case we recover the well known **proportionality of the entropy with the area**.

$$S \sim N_* \sim \sqrt{s} r_* \sim M_{pl}^2 r_*^2$$

Classicalons at LHC

- The final LHC signature would be **multiple Ws/Zs \rightarrow leptons+jets+mE_T**.



- We use the **effective W/Z approximation** and convolute the geometric cross-section πr_*^2 with the W/Z-boson luminosity function to get the final cross-section :

$$\sigma = \int_{\tau_{min}}^1 d\tau \pi r_*^2(\hat{s})^2 \int_{\tau'}^1 \frac{d\tau'}{\tau'} \int_{\tau'}^1 \frac{dx}{x} f_i(x, q^2) f_j(\tau'/x, q^2) \frac{dL}{d\xi}$$

- We find the **total number of Ws/Zs** a classicalon would decay to by solving for β and N_* the equations,

$$\gamma N_* r_*^3 \int \frac{\omega \sqrt{\omega^2 - m^2} d\omega}{e^{\beta\omega} - 1} = N_*$$

$$\gamma N_* r_*^3 \int \frac{\omega \sqrt{\omega^2 - m^2} d\omega}{e^{\beta\omega} - 1} = \sqrt{s},$$

Results

$\sqrt{\hat{s}_{min}}$ (GeV)	Multiplicity	Total Cross-section (fb)		Signal		Background
		W^+W^+, W^+W^-, W^-W^-	l^+l^+ (fb)	$3l$ (fb)	$3l^+$ (fb)	l^+l^+ (fb)
600	4	57, 56, 14	4.3	4.6	0.5	1.4
740	5	35, 33, 7.8	2.7	4.6	0.4	0.3
870	6	23, 22, 4.8	2.6	4.2	0.4	<0.3
1000	7	16, 14, 3.1	2.2	3.4	0.8	-
1130	8	11, 10, 2.0	1.5	2.4	0.6	-

(preliminary)

(at 14 TeV with $\Lambda = 500$ GeV)

- For $\Lambda = v = 246$ GeV these classicalons can be seen even with 7 TeV LHC energy:
 2 same sign leptons (multiplicity 5): 3.6 fb
 3 same sign leptons (multiplicity 5): 0.5 fb

Results

2 same sign leptons

3 same sign leptons !
hardly any background

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Summary

- Classicalization can **unitarize WW scattering** in the absence of the Higgs.
- Classicalons have **thermodynamic properties** like temperature and entropy and their decay products will have a Bose-Einstein distribution.
- **Prospects** for finding these classicalons are **good at the LHC**. Might even see them in the **7 TeV** run.