

Dark sector particle searches
and
Directional Detection
with
Gas Time Projection Chambers

Igal Jaeglé

University of Hawai'i at Mānoa

LAL-Orsay, 18 Février 2014

Table of contents

- 1 Personal details
- 2 Introduction
- 3 SuperKEKB, the intensity frontier
- 4 Beam commissioning detectors
 - Directional detection
 - TPC characterisation
 - Micro-TPC design
- 5 Belle/Belle II setups
- 6 Radiative channels
- 7 Higgs-strahlung channels
- 8 Dark matter directional detection

Personal details

Professional Experience

- 2010–present **University of Hawai'i, Honolulu, USA.**
Post-doc in Experimental Physics.
- 2008–2010 **University of Basel, Basel, Switzerland.**
Post-doc in Experimental Physics.
- 2001–2007 **University of Basel, Basel, Switzerland.**
Graduate student in Experimental Physics.
-
-

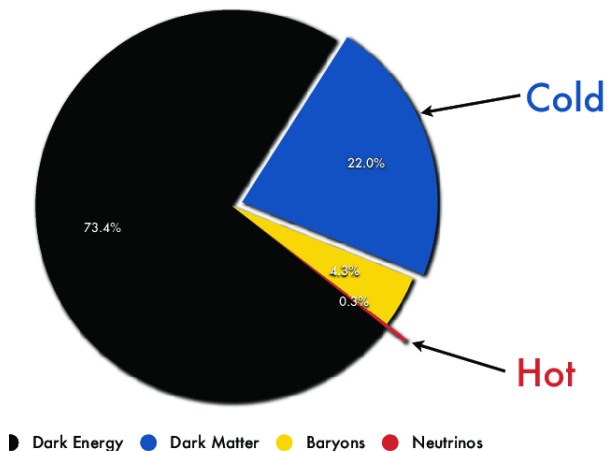
Collaborations

- 2010–present Belle and Belle II (Tsukuba, Japan).
- 2010–present Directional Dark matter Detector - D³ (Honolulu, HI, USA).
- 2010–2013 Dark Matter Time Projection Chamber - DMTPC (Carlsbad, NM, USA).
- 2008–2010 Crystal Ball and Travel Around Photon Spectrometer - A2 (Mainz, Germany).
- 2001–2010 Crystal Barrel and TAPS - CB-ELSA/TAPS (Bonn, Germany).
-
-

Education

- 2001–2007 **University of Basel, Basel, Switzerland.**
Ph.D. in Experimental Physics - "SUMMA CUM LAUDE"
- 1999–2000 **University of Haute-Alsace, Mulhouse, France.**
Diplôme d'Etude Approfondie (DEA) de physique subatomique.
- 1998–1999 **Louis Pasteur University, Strasbourg, France.**
Maîtrise de Chimie-Physique.
-
-

We know there is dark matter / Introduction



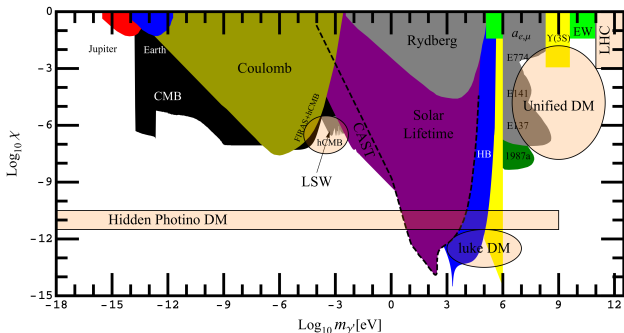
- Can Belle/Belle II and/or a directional detection detector contribute to the dark sector particle searches ?

Dark gauge boson / Introduction

to be distinguished from dark matter or WIMP, **coupling to SM fermions via EM current are:**

- searched since the late 80's, it may or may not exist but
- many theorists think there is a larger dark sector, hence recent strong interest in dark sector models (Unified DM)
- **introduce a vector boson A , and often a dark Higgs h' by a Higgs mechanism**
 - ▶ plot below shows astrophysical and cosmological, constraints and experimental limits: kinetic mixing vs. A boson mass **J. Jaeckel and A. Ringwald - arXiv:1002.0329v1**

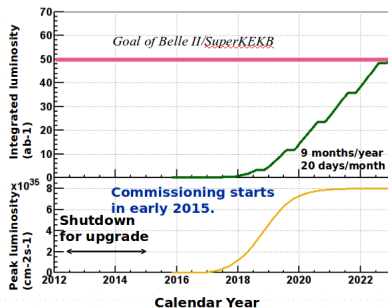
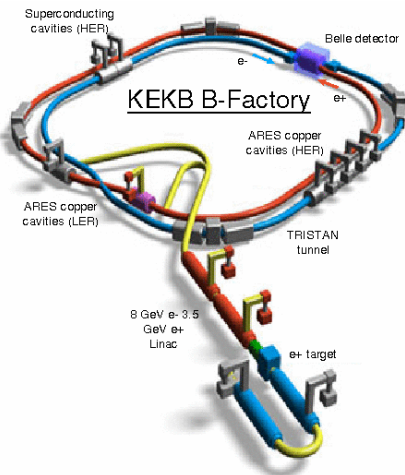
★ kinetic mixing: χ/ϵ



=> BaBar, Belle and Belle II can cover region between few 10's of MeV/c^2 and $10 \text{ GeV}/c^2$

SuperKEKB, the intensity frontier

- Belle/Belle II experiment at KEK/super KEK B-factory in Tsukuba, Japan
- Belle@KEKB $L = 977 \text{ fb}^{-1}$ at $\Upsilon(1S, 2S, 3S, 4S, 5S)$ and continua
- Belle II@superKEKB $L_{\text{projected}} = 40 \text{ ab}^{-1}$ at $\Upsilon(1S, 2S, 3S, 4S, 5S, 6S)$ and continua



$$L = \frac{\gamma_{\pm}}{2\sigma_e} \left(1 + \frac{\sigma_y^*}{\sigma_x^*} \right) \frac{I_{\pm} \xi_{\pm y}}{\beta_y^*} \left(\frac{R_L}{R_y} \right)$$

- nano-beam scheme extremely small β_y^* low emittance
- beam current $\times 2$

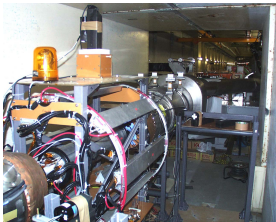
S. Kurokawa and E. Kikutani NIM A 499, 1 (2003)

superKEKB luminosity will generate huge beams-induced background

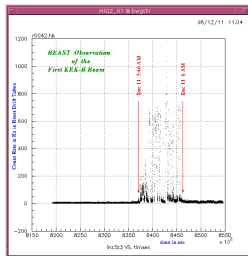
Beam commissioning detectors before Belle

aka BEAST used in 1998 to monitor radiation level and particle rates during KEKB commissioning

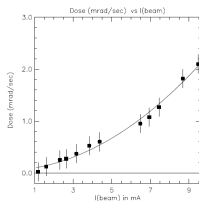
- BEAST in the cave



- first beam



- rate

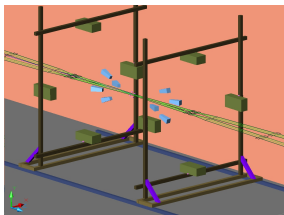


- provided important feedback to accelerator group during commissioning, and ensured background levels acceptable before Belle roll-in
- located at IP composed of PIN diodes, MOSFETs, Drift tubes, CsI and two Silicon Strip Ladders
- but did not prevent synchrotron radiation from damaging first beampipe

Beam commissioning detectors before/during Belle II

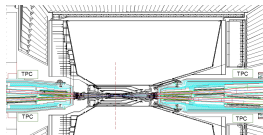
- measure instantaneous and integrated radiation dose at position of Belle II subdetectors
- measurements of luminosity and background levels during beam commissioning
- synchrotron and neutron backgrounds were unexpectedly problematic in Belle, we would like to measure
- details understanding and Monte Carlo simulation validation of beam-induced backgrounds
 - ▶ Touschek: intra-bunch scattering, electron/positron deviates from beam-bunch by Coulomb scattering
 - ▶ Coulomb (beam-gas scattering): electron/positron scattering with residual gas-atoms
 - ▶ Radiative Bhabha process propagate along beam axis and interact with magnet iron
 - ▶ synchrotron radiation: emitted by the beam since $\propto E_e^2$ and B^2 , High Energy Ring beam dominant
 - ▶ also two-photon and beam-beam but not critical background

● phase 1



- non magnetic support structure (Rosen, Hawaii)
- BGO crystals for luminosity monitoring (Wang, NTU)
- Diodes for radiation monitoring (Cinabro, Wayne / Marinas, Bonn)
- few PXD, SVD modules (MPI, KEK)
- CDC prototype (Uno, KEK)

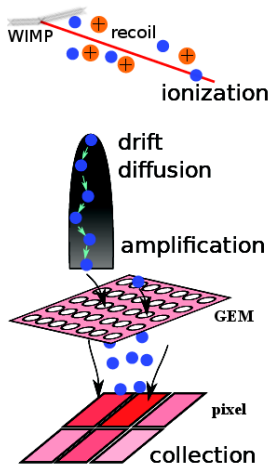
● phase 2



- microTPCs for directional neutron detection, x-rays, tracks (Vahsen, Hawaii)
- thermal neutron detection system (Roney, Victoria)

Gas Time Projection Chambers (TPCs)

Time Projection Chamber as fast neutrons/WIMPs detector filled with gas



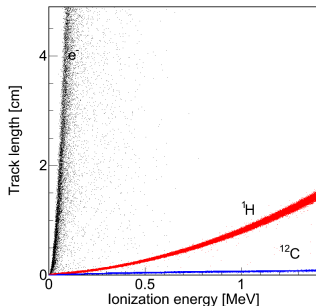
- fast neutrons not detected directly
- but through their scattering product with the gas-nucleus by elastic scattering
$$n + A_{rest} \rightarrow n' + A_{recoil}$$
- nuclear recoil ionizes gas along track
- electric field moves charges
- amplification (by 2 GEMs)
- readout (FE-I4B pixel chips)
 - ▶ 2D charge distribution
 - ▶ + timing information
 - ▶ + known drift velocity \Rightarrow 3D hit information
 - ▶ + known GEMs gain (and QF) \Rightarrow energy

\Rightarrow we reconstruct the nuclear recoil (θ, ϕ, E)

Directional detection concept

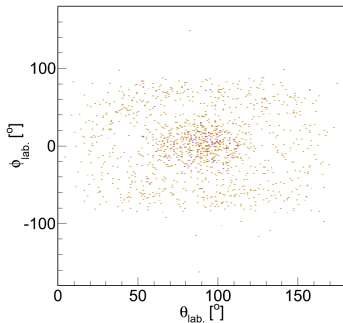
3D-direction and total ionization energy measurements of nuclear recoil allow

- dE/dx topology (iso-butane at 1 atm)



- nuclear recoil “profil”

- ▶ not energy weighted
- ▶ distribution points back to neutron direction (along x-axis)



$$E_{recoil} = E_{neutron}^{incident} \cdot \frac{2m_{recoil} \cdot m_{neutron}}{(m_{recoil} + m_{neutron})^2} \cdot (1 + \cos\psi_{cm})$$

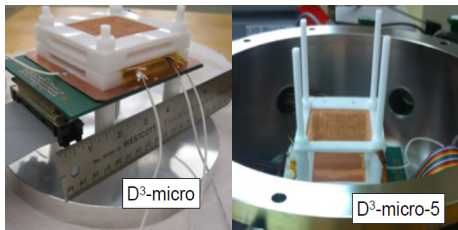
with $\psi_{cm} = 2 \psi_{lab}$, ψ_{lab} opening angle between incident neutron and nuclear recoil

$$(\cos\psi_{lab} = \frac{\vec{n} \cdot \vec{A}}{n \cdot A})$$

pre-prototype-micro-TPC / TPC characterisation

R&D with a pre-prototype TPC: S.E. Vahsen et al. <http://arxiv.org/abs/1110.3401>

- volume from $\sim 0.61 \text{ cm}^3$ to $\sim 2.5 \text{ cm}^3$
- charge amplified by 2 GEMs
- detected by ATLAS Pixel Chip FE-I3
- inside vacuum vessel
NB: white material corresponds to delrin
- vacuum vessel + neutron source

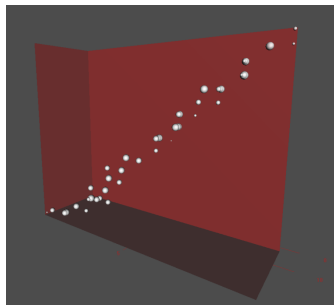


- stable operation for more than a year, large datasets recorded
 - ▶ commissioning w/ Ar:CO₂:70:30; muons, x-rays, alpha-particles (11,12)
 - ▶ detailed calibration & directional neutron detection w/ He:CO₂:70:30 (Fall 12-now)

ATLAS Pixel Chip electronics / TPC characterisation

High gain + ATLAS Pixel Chip electronics allows full 3D tracking

- track range (position resolution)
 - track dE/dx (topology)
 - track total ionization energy (energy resolution)
 - track direction (angular resolution)
- cosmic-ray track: ~ 7 mm and 2 keV measured by micro-D³ prototype



● FE-I3 low-noise and low-threshold

- ▶ chip size 0.84 cm x 0.76 cm
- ▶ pixel size 50 μm x 400 μm
- ▶ 18 column x 160 row
- ▶ 400 ns time range with 16 graduation
- ▶ threshold 3000 e⁻
- ▶ 100k e⁻ charge range with 128 graduation

● FE-I4 low-noise and low-threshold

- ▶ chip size 2 cm x 1.68 cm
- ▶ pixel size 50 μm x 250 μm
- ▶ 80 column x 336 row
- ▶ 1600 ns time range with 64 graduation
- ▶ threshold 1400 e⁻
- ▶ 100k e⁻ charge range with 16 graduation

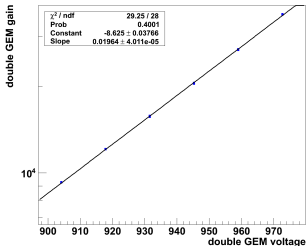
- virtually no noise
- high-single electron efficiency => suitable for low-mass WIMP search

Gain resolution and stability / TPC characterisation

High gain w/o sparking for weeks, measurement w/ pulseheight analyzer

- gain vs voltage

gain vs. GEM voltage

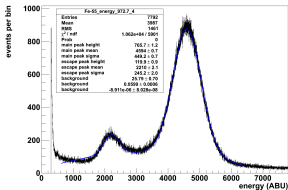


- sufficient gain to achieve single-electron sensitivity if needed
- good gain resolution for MeV-scale signals, adequate even for few-keV signals!

- $^{55}\text{Fe}/x\text{-rays}$ source

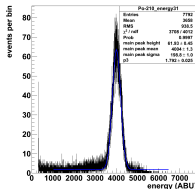
$$\sigma_{\text{gain}} = 11 \% \text{ at } 3 \text{ keV}$$

$$\sigma_{\text{gain}} = 8 \% \text{ at } 5.9 \text{ keV}$$

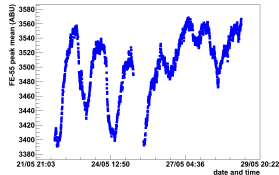


- $^{210}\text{Po}/\alpha$ source

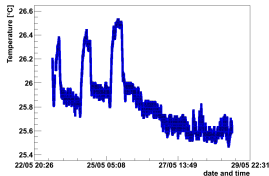
$$\sigma_{\text{gain}} = 5 \% \text{ at } 5 \text{ MeV}$$



- 5.9 keV x-ray peak vs. time

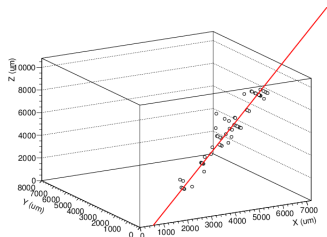
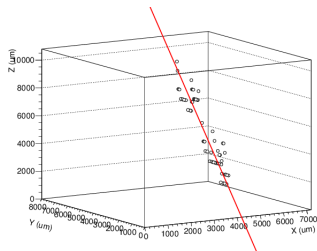
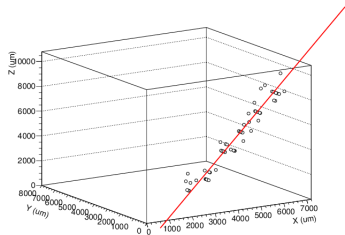


- temperature vs. time



3D point resolution / TPC characterisation

10k events of cosmics recorded with Ar:CO₂:70:30 at 1 atm, use such events to measure detector point resolution ($< 200 \mu\text{m}$)

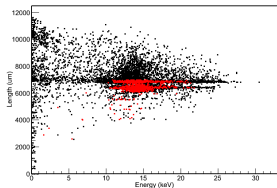


=> based on measured point resolution, expect angular resolution on nuclear recoils ~ 1 degree

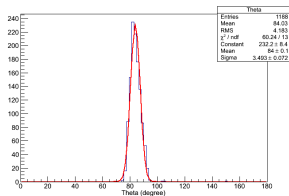
Angular resolution, nuclear recoils / TPC characterisation

^{210}Po alpha-source inside vacuum vessel, He:CO₂:70:30 at 1 atm, plots below show how well we can locate it:

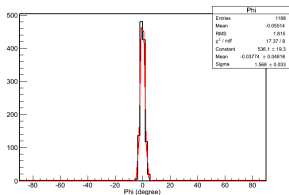
- track length vs. total ionization energy



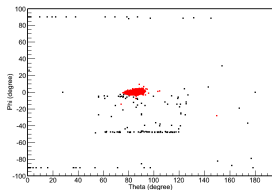
- polar angle distribution



- azimuthal angle distribution



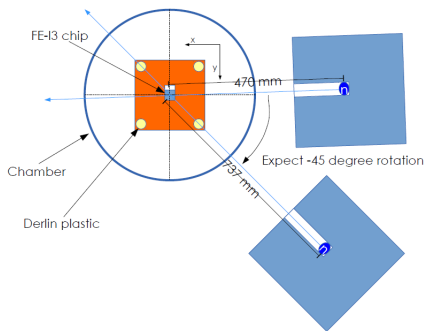
- azimuthal vs. polar angles



- selected events clearly point back to a single source
- No BG after good-track selection
- consistent with $\sigma(\theta, \phi)$ detector $< 1^\circ$

neutron setup / TPC characterisation

Detector Setup Schematic for 5 mm gap

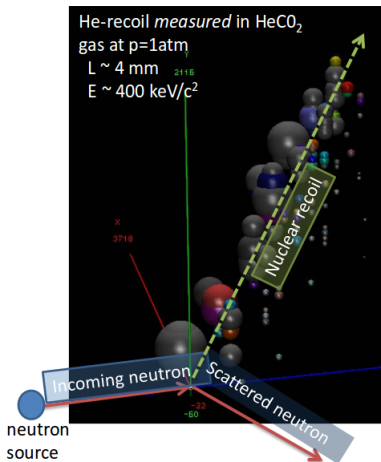


4 settings of the source:

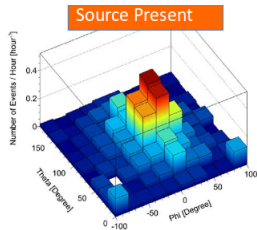
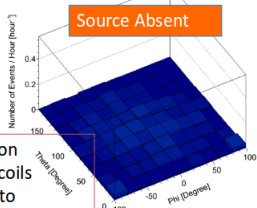
- $\theta = 90$ degrees and $\phi = 0$ degrees
- $\theta = 90$ degrees and $\phi = 45$ degrees
- $\theta = 90$ degrees and $\phi = -45$ degrees
- $\theta = 65$ degrees and $\phi = 45$ degrees

Directional Neutron Detection / TPC characterisation

DCube-micro-5 (He:CO₂:70:30 in 1 atm with FE-I3 board) measurement in LAB. with a pseudo neutron beam emitted by ²⁵²Cf-source



Mean direction of nuclear recoils corresponds to neutron source location

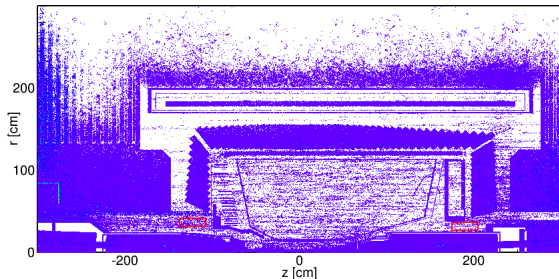


⇒ recoil angular distribution points back to the neutron direction

Micro-TPC conceptual design / Micro-TPC design

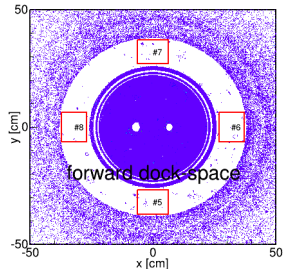
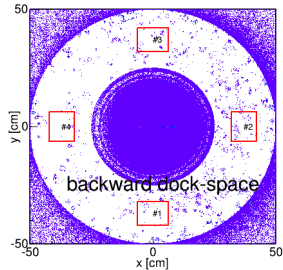
TPCs positioned to disentangle the different sources of neutrons by measuring the angular distribution of the recoil nucleus.

- "phase 2" top view - TPCs located in dock-spaces



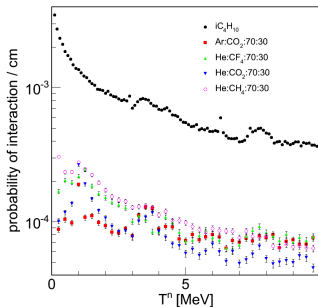
- 8 TPC chambers $10.12 \times 12.7 \times 30 \text{ cm}^3$
- active volume $5 \times 5 \times 25 \text{ cm}^3$
- surface from 0.72×0.8 to $4 \times 3.36 \text{ cm}^2$ can be instrumented
- $E \parallel B$ (0.3 to 0.9 kV/cm [depending of the chip and gas] \parallel 1.5 T)

- "phase 2" R-phi view

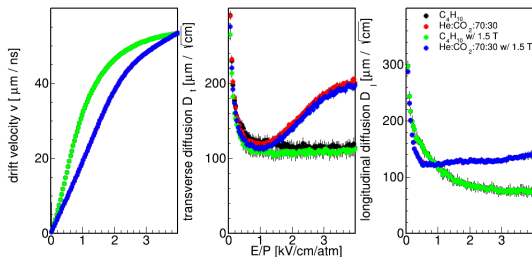


Gas choice / Micro-TPC design

- good neutron detection efficiency
- attachment coefficient low
- gas gain ~ 100
- simulate gas / gas mixture with
 - ▶ iC_4H_{10} (flammable and explosive)
 - ▶ Ar:CO₂
 - ▶ He:CO₂
 - ▶ He:CF₄
 - ▶ He:CH₄ (flammable, but not explosive)



- Gas parameters calculation by MAGBOLTZ and effect of 1.5 T magnetic field



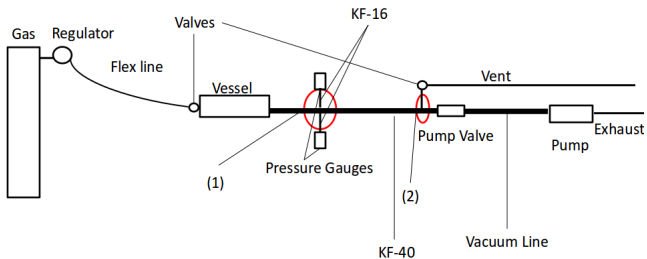
- B-field has negligible effect (for gases with small drift velocity)
- we want $E \parallel B$
- if not $E \parallel B \Rightarrow$ diffusion and drift velocity more complex form

\Rightarrow best trade off between safety, efficiency, gas properties and track length is given to He:CO₂:70:30 at 1 atm

Gas system / Micro-TPC design

- gas
- regulator
- gas in/out
- exhaust
- mechanical overpressure release
 - ▶ gas system designed by Michael Hedges

Design (Flow Chart)



=> all parts ordered

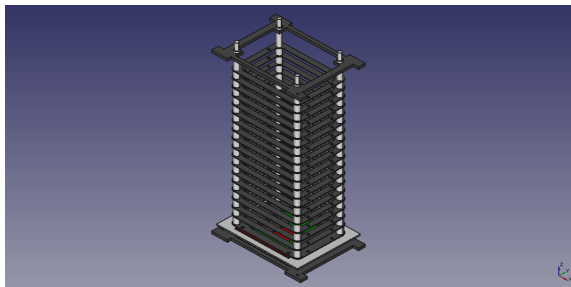
Field cage (FC) / Micro-TPC design

FC encircles sensitive volume and produced the electric field, it is composed of:

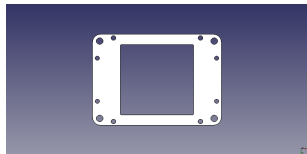
- an anode plane, field-shaping rectangular rings and a cathode
- connected to a resistor chain creating a linearly degrading potential

▶ full structure

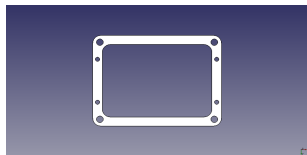
- ★ 7 mm/10 mm gap between each rectangular rings
- ★ 15 cm drift gap
- ★ rings are 1 cm from the vessel wall



▶ cathode



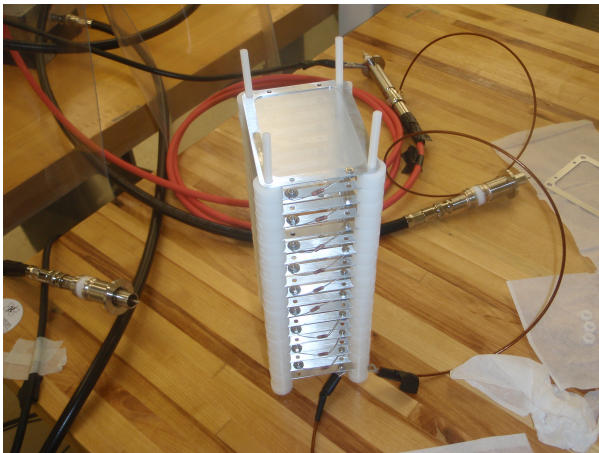
▶ rectangular ring



- anode: rectangular ring at which is glued a mesh

v_{drift} [$\mu\text{m}/\text{ns}$]	D_t [$\mu\text{m}/\sqrt{\text{cm}}$]	E [kV/cm]	V_{anode} [kV]
10	124.3	0.53	$7.95 + V_{GEM_2}^{top}$

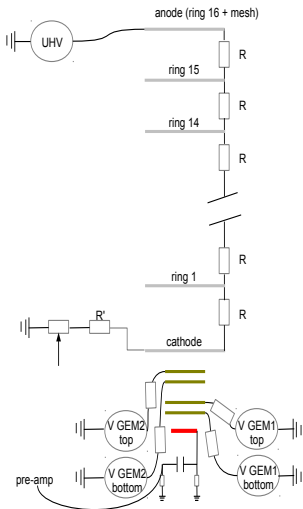
Field cage (FC) / Micro-TPC design



v_{drift} [$\mu\text{m}/\text{ns}$]	D_t [$\mu\text{m}/\sqrt{\text{cm}}$]	E [kV/cm]	V_{anode} [kV]
10	124.3	0.53	$7.95 + V_{GEM_2}^{top}$

HV bias / Micro-TPC design

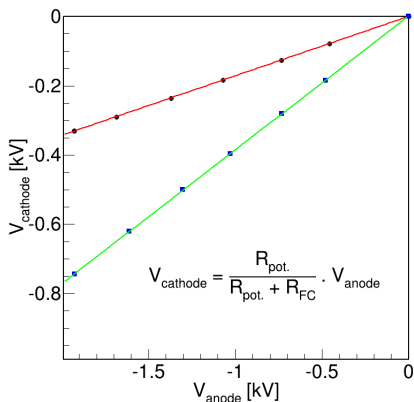
● circuit diagram



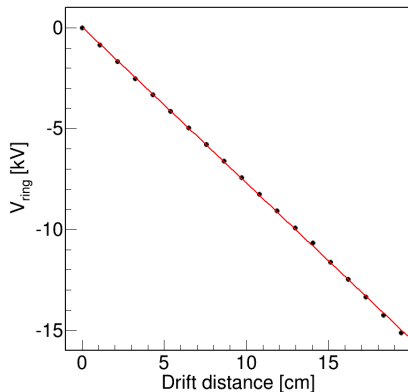
- $R = 4 \text{ M}\Omega \pm 0.1 \% \Rightarrow R^{total} = 64 \text{ M}\Omega$
- GEM HV range between 900 V and 2000 V
- $V_{anode}^{max} = 10.48 \text{ kV}$
- $I = 132.5 \mu\text{A}$
- $R' = 15.16 \text{ M}\Omega$
- $15.16 \leq R' + R_{potentiometer} \leq 45 \text{ M}\Omega$
- $V_{cathode} = V_{GEM2}^{top} + 530 \text{ V/cm} \times 0.06 \text{ cm}$
- $V_{anode} = V_{cathode} + 530 \text{ V/cm} \times 16 \text{ cm}$
- $V_{cathode} = \frac{R' + R_{potentiometer}}{R' + R_{potentiometer} + R} \cdot V_{anode}$
- procedure to set the field:
 - ▶ set first GEM HV
 - ▶ tune **offline** $R_{potentiometer}$ to the desired value
 - ▶ connect potentiometer box cables and switch on UHV
 - ▶ set V_{anode}

HV bias check / Micro-TPC design

- potentiometer range



- field measurement at 15 kV



- $V_{anode} \cdot 0.171 \leq V_{cathode} \leq V_{anode} \cdot 0.386$
- sparking at Feedthrough level above 15 kV
- field homogeneity below 3%, measurement limited by voltmeter precision

Radial displacement (Δr) / Micro-TPC design

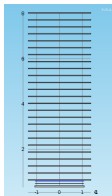
Electric field should be as uniform as possible to minimize distortions of the reconstructed tracks

- $\Delta r(x, y) = \int_{z=0}^{z_{max}/2} \frac{E_r(x, y, z)}{E(x, y, z)} dz$ with

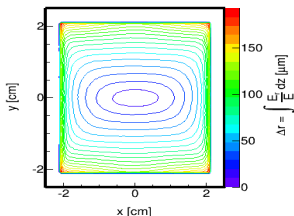
- ▶ z drift distance
- ▶ E_r radial field
- ▶ E field

- Finite Element Method (FEM) used to study the uniformity (in COMSOL)

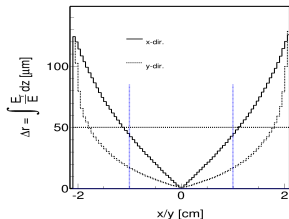
- ▶ field cage in COMSOL



- ▶ Δr vs. x vs. y



- ▶ Δr for $x=0$ and $y=0$, blue dashed lines represent one chip



- point resolution in $r\phi < 125 \mu m$ in the 15 cm drift gap
- point resolution in $z < 250 \mu m$ in the 15 cm drift gap

Gain gas optimization / Micro-TPC design

To achieve the highest possible detection of the primary ionization, GEMs are used to amplify the signal.

- GEMs are gate-less, can operate continuously and have intrinsic ion feedback suppression
- gain has to be adapted to the amount of ionization
- iso-butane at 1 atm
- He:CO₂:70:30 at 1 atm

► SRIM dE/dx

► gain FOM

► SRIM dE/dx

► gain FOM

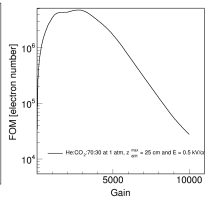
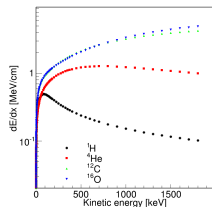
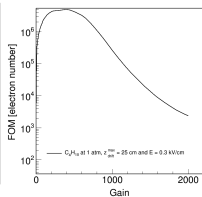
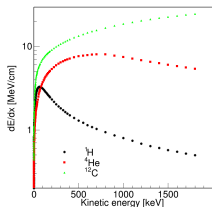


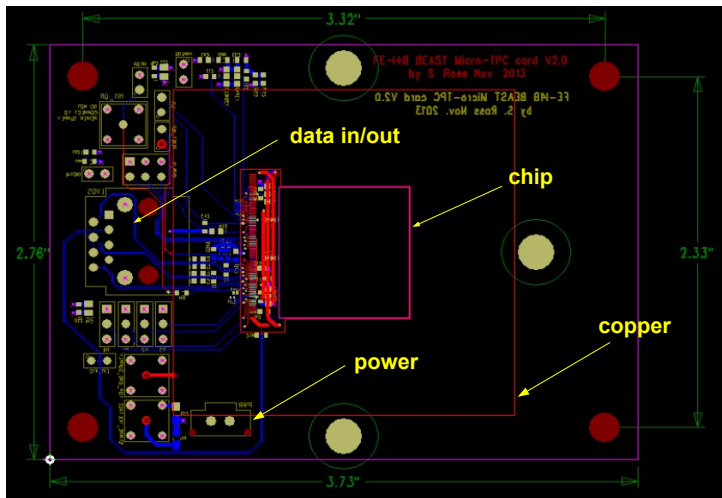
Table: Average electron number per pixel and optimal gain for He:CO₂:70:30 and C₄H₁₀ at 1 atm.

gas	He in He:CO ₂ :70:30	H in C ₄ H ₁₀
electron number per pixel	171	1294
optimal gain	2800	600

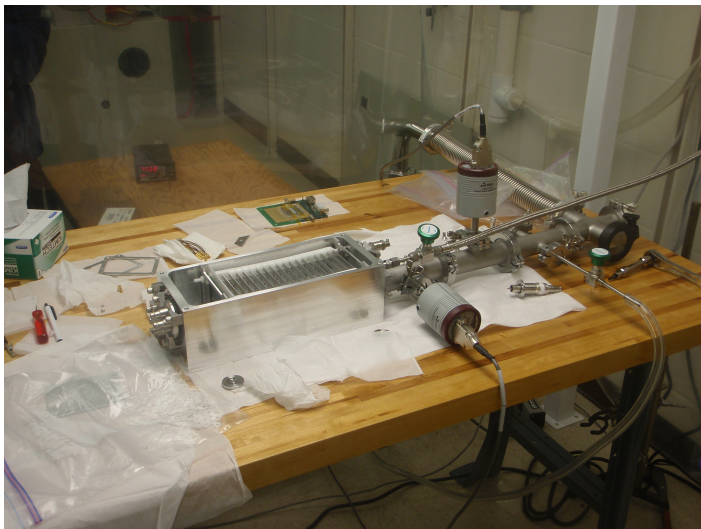
Electronics / Micro-TPC design

We will use FE-I4B with a single chip for the first prototype:

- board



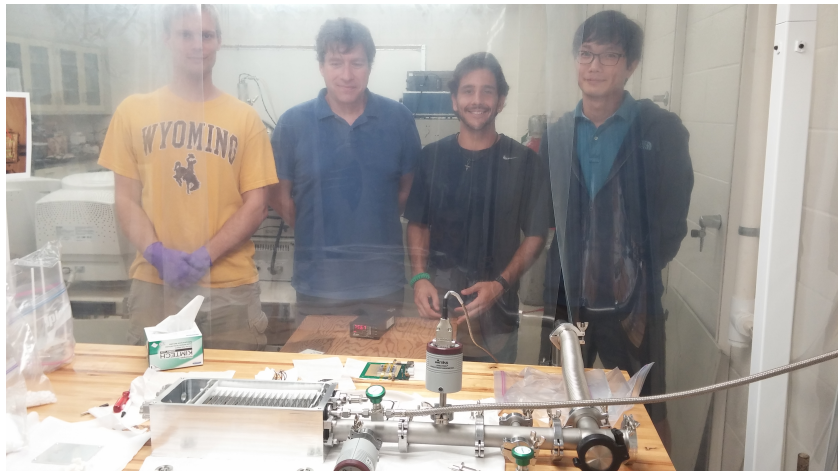
Nice picture I / Micro-TPC design



Nice picture II / Micro-TPC design



Nice picture III / Micro-TPC design

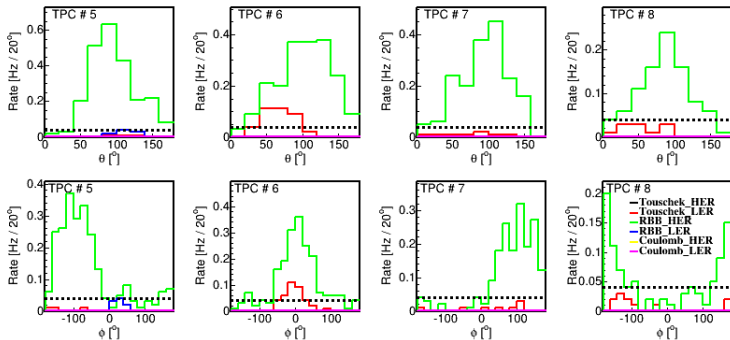
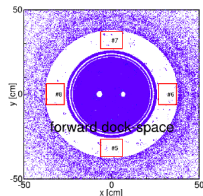


Left to right: Gas System, Speaker, Machine Shop and USBpix-DAQ/LV/HV

Recoil angle rate distribution in forward TPC

MC simulation with 1 chip and He:CO₂:70:30 at 1 atm

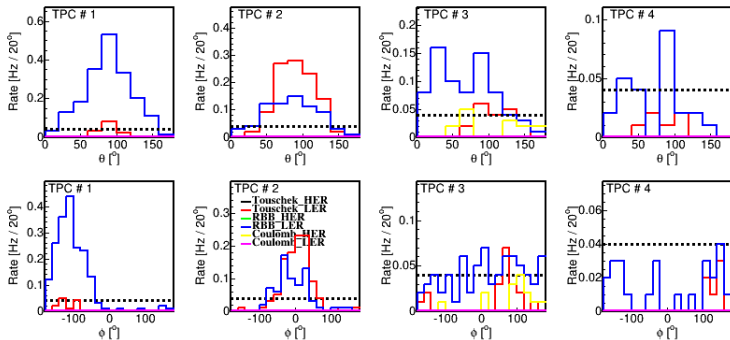
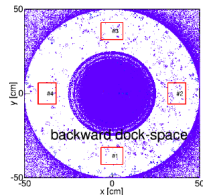
- RBB LER dominates at end of phase2
- with single beams
 - ▶ no RBB
 - ▶ measure Touschek
 - ▶ with vacuum bump Coulomb can be measured



Recoil angle rate distribution in backward TPCs

MC simulation with 1 chip and He:CO₂:70:30 at 1 atm

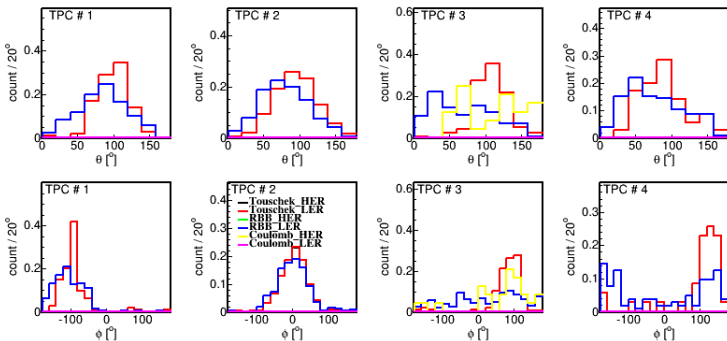
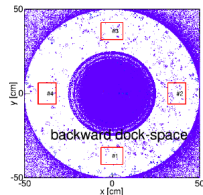
- RBB LER dominates at end of phase2
- with single beams
 - ▶ no RBB
 - ▶ measure Touschek
 - ▶ with vacuum bump Coulomb can be measured



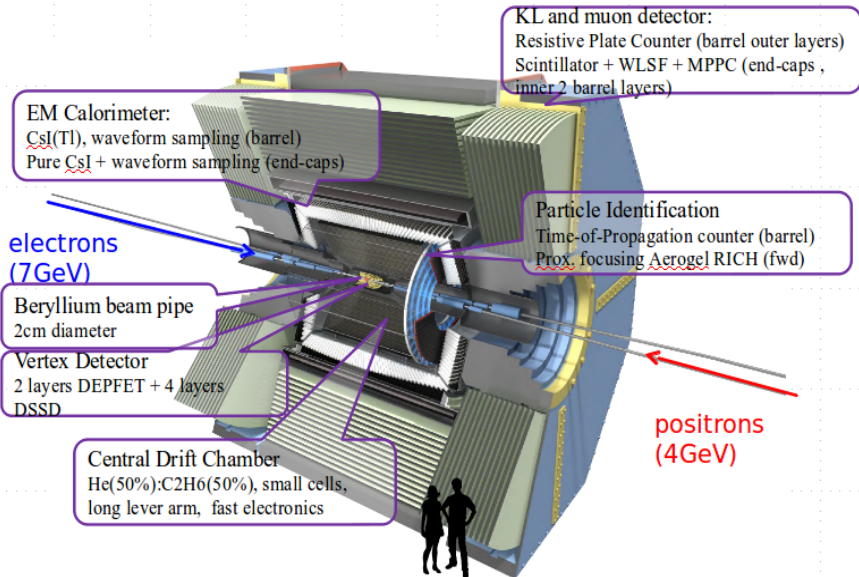
Recoil angle re-normalized distribution in backward TPCs

MC simulation with 1 chip and He:CO₂:70:30 at 1 atm

- re-normalized distributions
- clear difference visible



Belle/Belle II setups

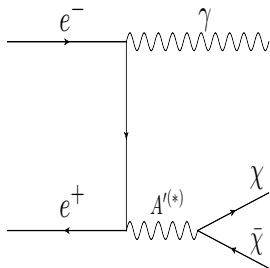


Radiative channels

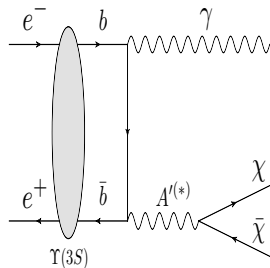
Belle may have or/and Belle II may produce dark matter, χ , dark photon (dark gauge boson), A , and dark Higgs h' as long as the mass of each is low in the following channels:

- Radiative decay:
 - $e^+e^- \rightarrow \gamma X$
 - $e^+e^- \rightarrow \Upsilon(nS) \rightarrow \gamma X$
 - ▶ $X = \chi\bar{\chi}$
 - ▶ $X = A$
 - ▶ $A \rightarrow l^+l^-$ or hadrons or dark matter

★ two body decays



★ upsilon decays



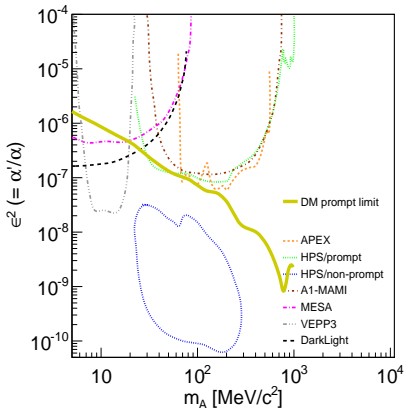
NB: dark photon or dark gauge boson has no normalized name and letter: γ' , A , A' or U

Radiative channels, γA

6 experiments have been approved/commissioned and will cover region between $1 \text{ MeV}/c^2$ and $1 \text{ GeV}/c^2$

- all experiments will look for a prompt decay and $A \rightarrow l^+l^-$ by detecting the leptons
- HPS will also look for a displaced vertex.
- Belle/Belle II could set a limit between $200 \text{ MeV}/c^2$ and $10 \text{ GeV}/c^2$

► predicted sensitivity



- VEPP3/Russia (new setup), $e^+ + p \rightarrow \gamma A$
- APEX/USA-JLAB (new setup), $e^- + \text{nucleus} \rightarrow \gamma A$
- HPS/USA-JLAB, $e^- + \text{nucleus} \rightarrow \gamma A$
- DarkLight/USA-JLAB (new setup), $e^- + \text{H} \rightarrow \gamma A$
- A1-MAMI/Germany, $e^- + \text{nucleus} \rightarrow \gamma A$
- MESA/Germany (new accelerator and setup),
 $e^- + \text{nucleus} \rightarrow \gamma A$

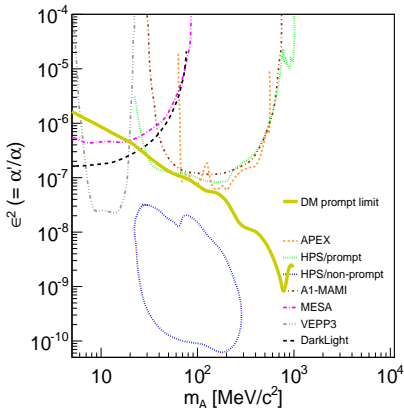
A not necessarily prompt

Radiative channels, γA

6 experiments have been approved/commissioned and will cover region between $1 \text{ MeV}/c^2$ and $1 \text{ GeV}/c^2$

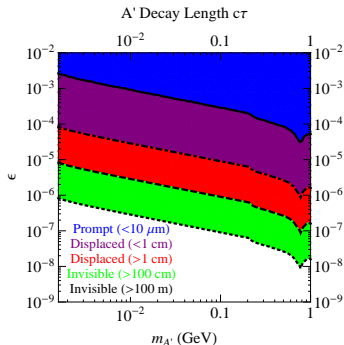
- all experiments will look for a prompt decay and $A \rightarrow l^+ l^-$ by detecting the leptons
- HPS will also look for a displaced vertex.
- Belle/Belle II could set a limit between $200 \text{ MeV}/c^2$ and $10 \text{ GeV}/c^2$

▶ predicted sensitivity



▶ plot shows lifetime of A as a function of its mass m_A and ϵ

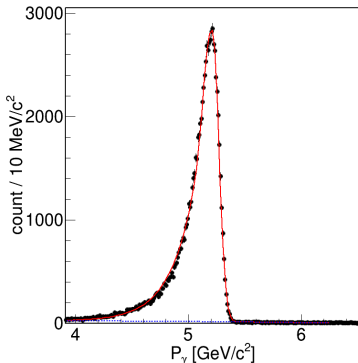
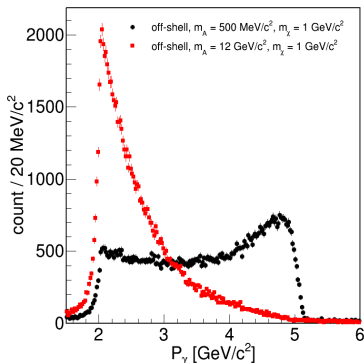
R. Essig et al, arXiv:0903.3941



MC simulation of signal signature / Radiative decay

Dark matter can be produced in two modes:

- off-shell i.e. $e^+e^- \rightarrow \gamma\chi\bar{\chi}$
 - on-shell i.e. $e^+e^- \rightarrow \gamma A, A \rightarrow \chi\bar{\chi}$
 - simulated photon spectra for $W(e^+e^-) = 10 \text{ GeV}/c^2$
- ▶ off-shell case ▶ on-shell case, $m_A = 1 \text{ GeV}/c^2$

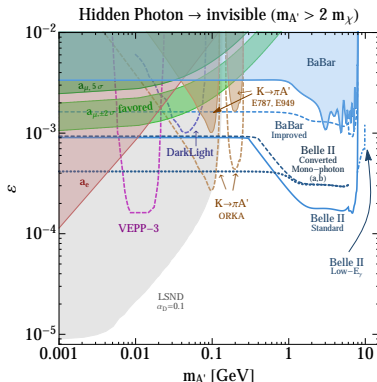


- off-shell signature => broad energy distribution
- on-shell signature => mono-photon, as displaced vertex decaying into leptons

Belle II predicted sensitivity / Radiative decay

Determined by R. Essig et al. [arXiv:1309.5084](https://arxiv.org/abs/1309.5084) for 50 ab^{-1}

- Belle II Standard with prescaled trigger of 100
- Belle II Converted Mono-photon (a,b) no prescaled trigger but $\varepsilon(\gamma \rightarrow e^+e^-) = 5\%$ instead of 1%



At start of Belle II physics run, photon trigger (not prescaled) might be implemented furthermore luminosity much lower than at designed luminosity

Higgs-strahlung channels

Belle may have or/and Belle II may produce dark matter, χ , dark photon (dark gauge boson), A , and dark Higgs h' as long as the mass of each is low in the following channels:

- Higgs-strahlung:

- $e^+e^- \rightarrow Ah'$

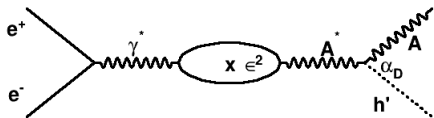
- ▶ if $m(h') > 2m(A)$, $h' \rightarrow AA$

- ▶ if $m(A) < m(h') < 2m(A)$, $h' \rightarrow l^+l^-$ or hadrons

- ▶ if $m(h') < m(A)$, $h' \rightarrow$ dark matter

- ▶ $A \rightarrow l^+l^-$ or hadrons or dark matter

- ▶ Higgs-strahlung decays



- ▶ α_D coupling between the dark Higgs and the dark photon

- ▶ $\alpha' = \frac{g'^2}{4\pi}$ fine structure for the dark photon

- ▶ g' coupling of the dark photon to electrons (or SM)

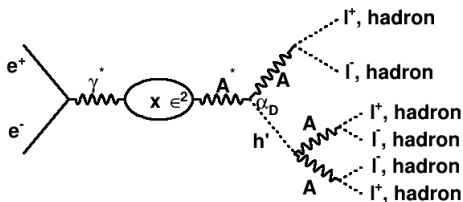
- ▶ $\epsilon = \frac{g'}{e}$ and $\epsilon^2 = \frac{\alpha'}{\alpha}$

- ▶ $\alpha = \frac{e^2}{4\pi} \sim \frac{1}{137}$ the electromagnetic coupling

Search for the dark photon and dark Higgs at Belle

Presented today: A and h' prompt and $m_{h'} > 2m_A$

for $0.1 < m_A < 3.5 \text{ GeV}/c^2$ and $0.2 < m_{h'} < 10.5 \text{ GeV}/c^2$



α_D : dark sector constant

ϵ^2 : kinetic mixing

● channels presented today

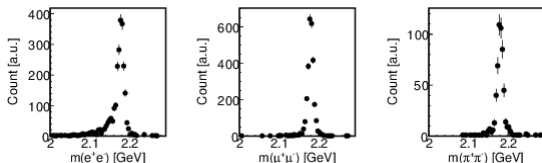
- ▶ $3e^+3e^-, 3\mu^+3\mu^-, 2e^+2e^-\mu^+\mu^-, 2\mu^+2\mu^-e^+e^-$
- ▶ $3\pi^+3\pi^-, 2\pi^+2\pi^-e^+e^-, 2\pi^+2\pi^-\mu^+\mu^-$
- ▶ $2e^+2e^-\pi^+\pi^-, 2\mu^+2\mu^-\pi^+\pi^-, e^+e^-\mu^+\mu^-\pi^+\pi^-$
- ▶ $2e^+2e^-X, 2\mu^+2\mu^-X, e^+e^-\mu^+\mu^-X$

- if A coupling to h' unity
- Higgs-strahlung channel most sensitive to A since QED background low
- than other decays e.g.: $e^+e^- \rightarrow A\gamma$ with huge QED background

Analysis strategy / Higgs-strahlung channels

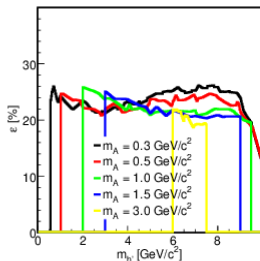
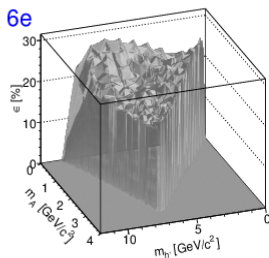
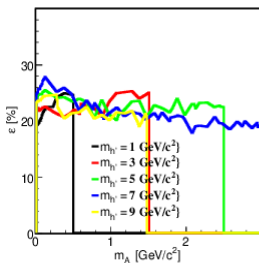
Full reconstruction of exclusive six-lepton/hadron final states from $e^+e^- \rightarrow Ah' \rightarrow AAA$

- final state identification
 - ▶ 6 charged tracks
 - ▶ 3 pairs of opposite charge
- signal reconstruction
 - ▶ impact parameters and χ^2 vertex fit cuts
 - ▶ require energy conservation
 - ▶ calculate invariant mass for each combinations of leptons/hadrons consistent with three distinct $A \rightarrow l^+l^-$ or hadrons
 - ▶ keep combinations with three masses “equal”
 - ▶ plots below show signal Monte Carlo simulation events surviving selection with $m_{h'} = 5 \text{ GeV}/c^2$ and $m_A = 2.19 \text{ GeV}/c^2$



Detection efficiency / Higgs-strahlung channels

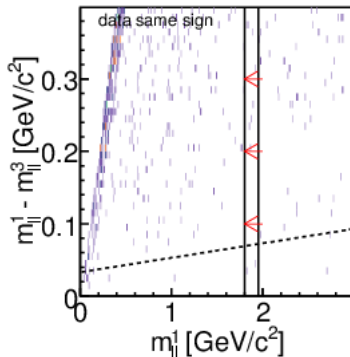
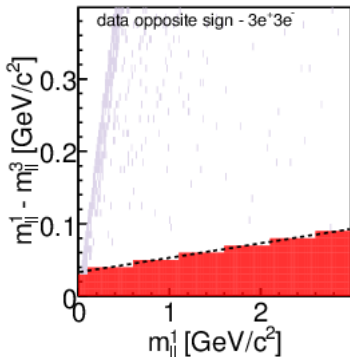
e.g. $e^+e^- \rightarrow Ah' \rightarrow AAA \rightarrow 3e^+3e^-$



Background estimation strategy / Higgs-strahlung channels

Data driven background estimation

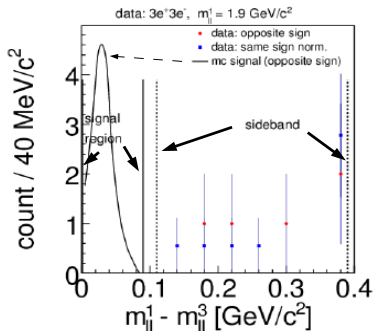
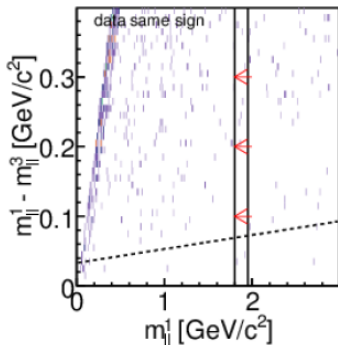
- estimate background using "same sign" events $e^+e^- \rightarrow Ah' \rightarrow A(I^+I^+)A(I^-I^-)$
- order masses of lepton pairs $m_{ll}^1 > m_{ll}^2 > m_{ll}^3$ and plot $m_{ll}^1 - m_{ll}^3$ vs. m_{ll}^1
- select region in m_{ll}^1 and predict background there using same sign
 - ▶ $e^+e^- \rightarrow 3e^+3e^-$



- opposite sign = signal, signal box blinded (red)

Background estimation method / Higgs-strahlung channels

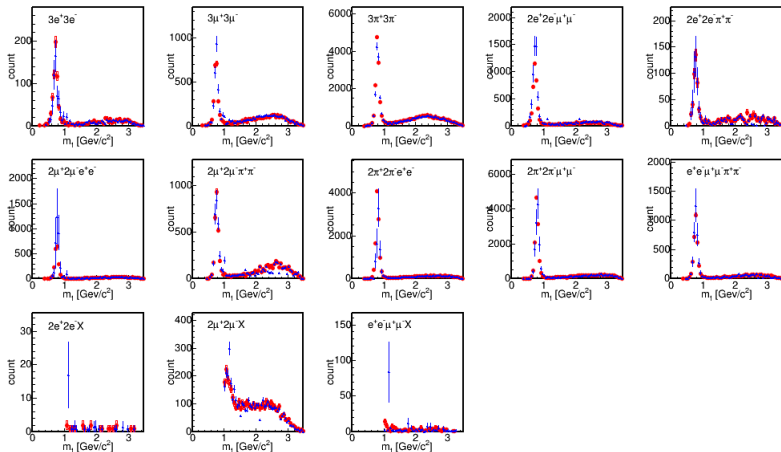
- sideband used to normalize same sign to opposite sign
- background estimated from the number of counts in the signal region of the same sign distributions
 - ▶ projection on $m_{\parallel}^1 - m_{\parallel}^3$ for $m_{\parallel}^1 = 1.9 \text{ GeV}/c^2$



MC simulation test / Higgs-strahlung channels

MC simulation of $e^+e^- \rightarrow \rho^0 \rho^0 l^+ l^-$ or hadrons produced in phase space

- red opposite sign
- blue same sign

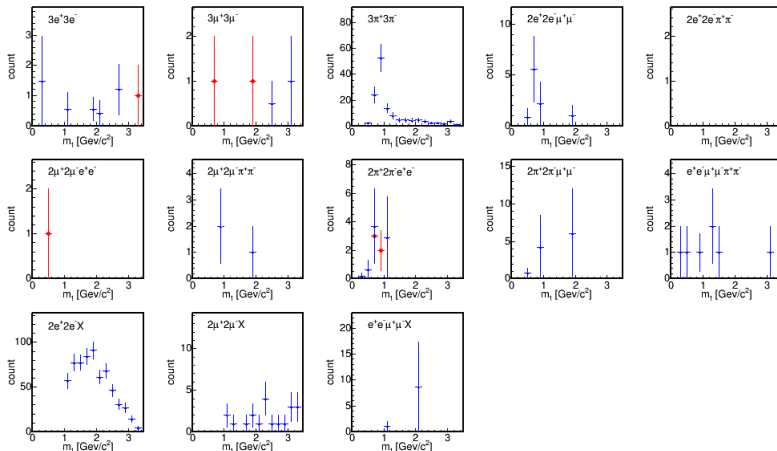


=> background estimation method verified successfully with MC

Background estimation results / Higgs-strahlung channels

Data control sample with the 4 boxes already open:

- red opposite sign
- blue same sign



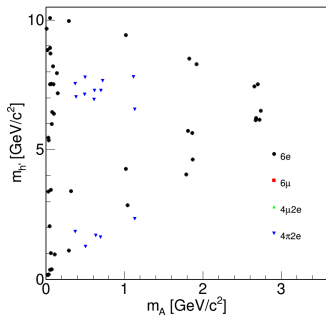
Background estimation results / Higgs-strahlung channels

Comparison between predicted Belle background and Belle number of events measured

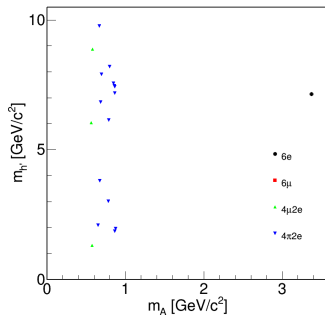
Final state	$6e$	6μ	$4\mu 2e$	$4\pi 2e$
Belle expected	4.75 ± 1.31	1.64 ± 1.12	0	7.4 ± 4
Belle measured	1	0	1	5

=> Number of events measured consistent with background expectation

• same sign (predicted)



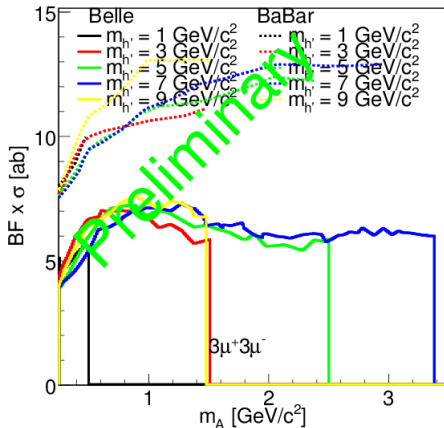
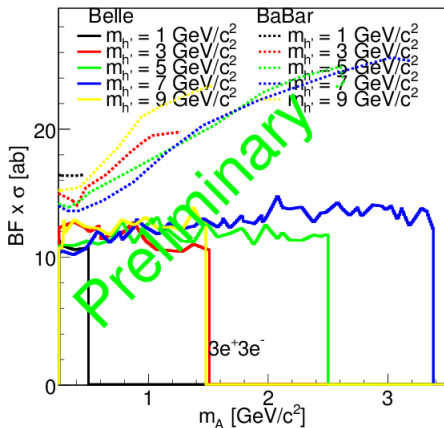
• opposite sign (measured)



background: Standard Model 2γ processes with ρ 's or ω 's in the final state

Belle limits / Higgs-strahlung channels

- Belle preliminary limits for $L = 980 \text{ fb}^{-1}$
- BaBar limits for $L = 520 \text{ fb}^{-1}$ [BaBar Collaboration - arXiv:1202.1313](#)
- $e^+e^- \rightarrow 3e^+3e^-$
- $e^+e^- \rightarrow 3\mu^+3\mu^-$

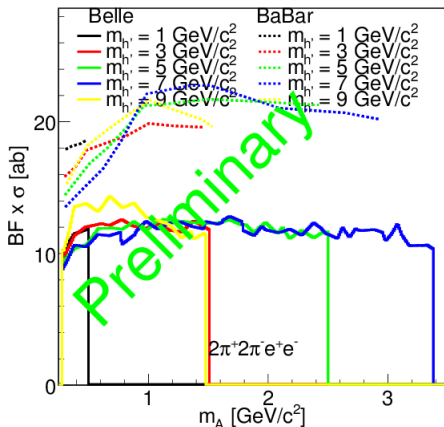


- upper limit (90 % CL) determined by Bayesian inference method with the use of Markov Chain Monte Carlo [A. Caldwell et al., CPC 180 \(2009\) 2197-2209](#)
- Belle limit scales nearly linearly with integrated luminosity

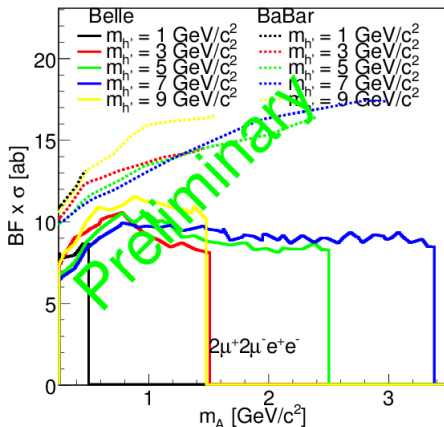
Belle limit / Higgs-strahlung channels

Compare to BaBar limits [BaBar Collaboration - arXiv:1202.1313](#)

● $e^+e^- \rightarrow 2\pi^+2\pi^-e^+e^-$



● $e^+e^- \rightarrow 2\mu^+2\mu^-e^+e^-$

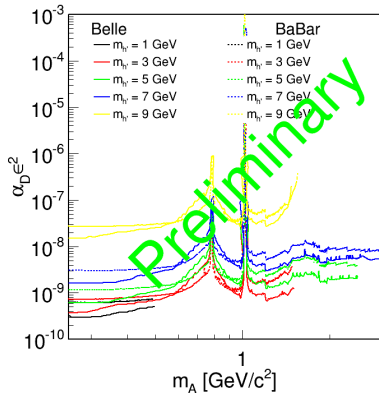


⇒ Belle II will scale nearly linearly with integrated luminosity or with square root of integrated luminosity in presence of background

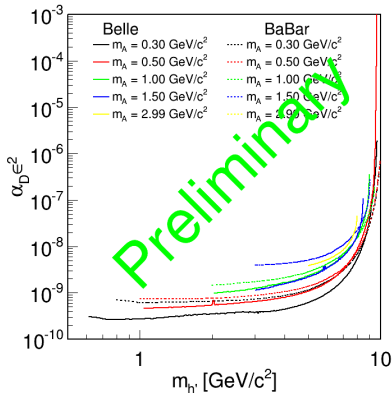
Combined sensitivity to the product of $\alpha_D \epsilon^2$

Belle combined sensitivity compared to BaBar combined limit

- dark photon 90 % CL sensitivity



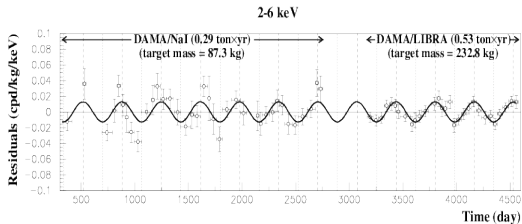
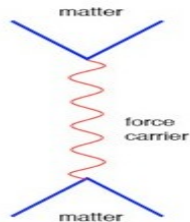
- dark Higgs 90 % CL sensitivity



Branching fractions and couplings versus cross section, dark photon and dark Higgs masses follow B. Batell et al. [arXiv:0903.0363](https://arxiv.org/abs/0903.0363) (2009)

Annual modulation ?

- DM can explain observed anomalies in astrophysical data and dark matter experiments
 - annual modulation due to the Earth's orbit around the Sun as the Sun orbits the galactic center ? DAMA/LIBRA, Eur. Phys. J. C 56: 333-355 (2008)



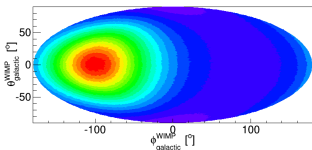
- scattering cross section $\sigma \propto \frac{1}{(q^2 - M_*^2)^2}$ with q : momentum transfer ($q^2 = 2m_{nucleus} E_{recoil}^{nucleus}$) and M_* : carrier particle = A or W, Z
 - $M_* \rightarrow 0$ then $\sigma \propto \frac{1}{q^4}$, long range interaction $\Rightarrow \frac{d\sigma}{dE_{recoil}^{nucleus}} \propto \frac{1}{(m_{nucleus} E_{recoil}^{nucleus})^2}$
 - $M_* \gg q$ then $\sigma \propto \frac{1}{M_*^4}$, contact term interaction
- for $M_{WIMP} = 10 \text{ GeV}/c^2$ and $m_{Na(I)} = 22(129) \text{ GeV}/c^2$, $q^{max} \sim 30(100) \text{ MeV}/c^2$ $\Rightarrow m_A \ll 1 \text{ GeV}/c^2 \ll M_{WIMP}$ and modulation amplitude enhanced

Daily oscillation / Dark matter directional detection

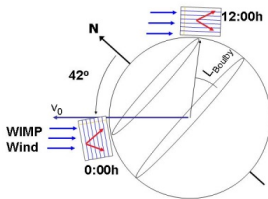
If **annual modulation** observed by DAMA/LIBRA is caused by WIMPs due to:

- Earth's orbit around the Sun as the Sun orbits the galactic centre
- since Earth's rotation a **daily oscillation** in the mean direction of the WIMPs must be observed as well
 - ▶ passing from TPC coordinate to galactic coordinate enforce daily oscillation
 - ▶ at least 10 recoil events to measure WIMP direction
 - => **strong constrain, no known background with this signature**

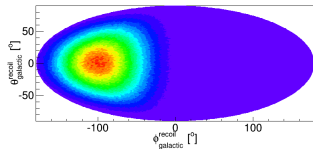
- WIMP flux in galactic coordinate



- daily oscillation scheme



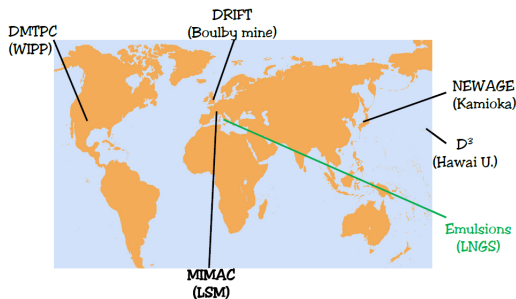
- F recoil angular distribution in galactic coordinate



=> use daily oscillation to clean and identify the WIMP signal

Current projects / Dark matter directional detection

Time Projection Chamber as WIMP detector filled with gas



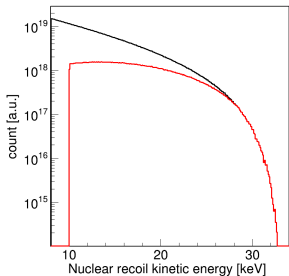
- low target mass, normally need very large gas volume
- but low track reconstruction threshold changes situation for low-mass WIMPs with only 1-10 m³ volume
- several groups attempting directional detection with gas TPC. Advantages of our approach are:
 - ▶ 3D tracking => better DM identification & alpha BG rejection
 - ▶ single electron efficiency => expect very low track-recons. threshold
 - ▶ basically free of noise

Low-pressure operation / Dark matter directional detection

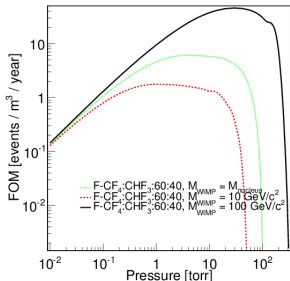
To measure low WIMP mass, eg $10 \text{ GeV}/c^2$:

- gas with target nucleus close to wanted WIMP mass and TPC friendly
=> Fluorine as target and gas CF_4 or gas mixtures $\text{CF}_4:\text{CHF}_3$ or $\text{CF}_4:\text{CS}_2$
- operate at extremely low-pressure, increase track length, $L(\text{at } P) = L(\text{at } P_0) \times \frac{P_0}{P}$,
but diffusion $\sigma(\text{at } P) = \sigma(\text{at } P_0) \times \sqrt{\frac{P_0}{P}}$
- find best trade off between the target mass and track length so that the directional sensitivity is maximized I. Jaegle et al. <http://arxiv.org/abs/1110.3444>

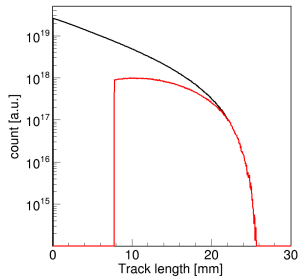
► F recoil kinetic energy
red line: 3D-dir. and 10 keV thres.



► optimal pressure
=> 1.5 torr for $10 \text{ GeV}/c^2$ WIMP



► F track length at 1.5 torr
red line: 3D-dir. and 10 keV thres.

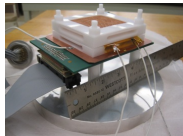


- operation at low-pressure with high gain $\sim 100k$ has been demonstrated in A. Breskin et al., NIM 217, 131 (1983), NIMA 433 476 (1999) and C.K. Shalem et al., NIMA 558 468-474 (2006)

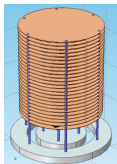
Future plan / Dark matter directional detection

Building this year a 10l volume detector

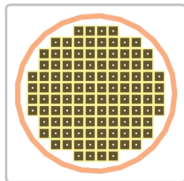
- build and run smoothly



- building this year

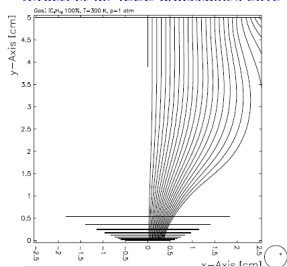


- planned



- to reduce the readout price
 - ▶ larger pixel chips
 - ▶ electrostatic focusing of drift charge
- existing ATLAS DAQ
- negative ion drift or gas w/ slow drift velocity

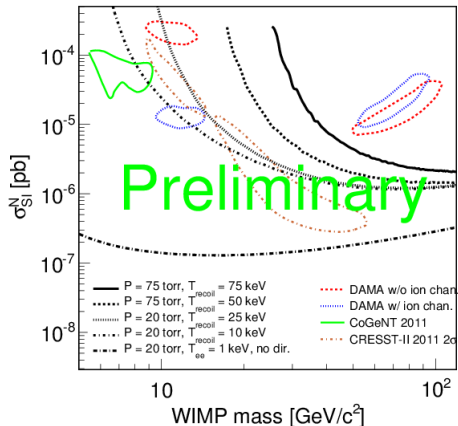
- electrostatic focusing of drift charge
S. Ross et al. *IEEE proceedings* 2012



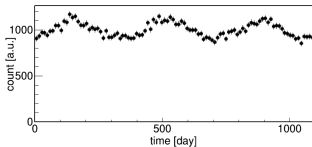
Reach plot for 3 years and 3 m^3

Preliminary reach plot shows advantages of low track energy threshold and possible golden scenario

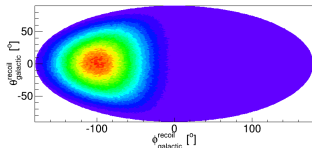
- observe annual modulation
- sub-sample of data for daily oscillation search



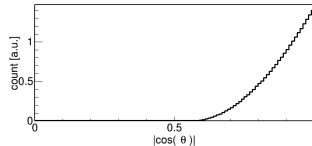
- annual modulation



- F recoil angular distribution in galactic coordinate



- angle between F recoil and WIMP direction



Conclusion

Belle and Belle II can contribute to the dark sector particles searches for the prompt and displaced vertex cases

- radiative decays
 - ▶ with 50 ab^{-1} , Belle II might reach sensitivity to rule out a prompt dark photon and dark Higgs decays into leptons or hadrons
 - ▶ expected background in Belle II is not known
- search in Belle data set for dark photon and dark Higgs in the mass ranges:
 - ▶ $0.1 < m_A < 3.5 \text{ GeV}/c^2$
 - ▶ $0.2 < m_{h'} < 10.5 \text{ GeV}/c^2$
 - ▶ we found that:
 - ★ background is small, implying
 - ★ limit scales nearly linearly with integrated luminosity
 - ★ Belle II expected to scale nearly linearly with integrated luminosity or square root of integrated luminosity

TPC:

- TPC characterisation shows very promising results
- MC simulation
 - ▶ beams-induced background, different background can be disentangled
 - ▶ measurement of WIMPs direction might be possible at low pressure

Thanks for your attention.