

# Données Manquantes : 2 exemples

## Séparation de RFI sur SMOS & Video Inpainting

Andrés Almansa

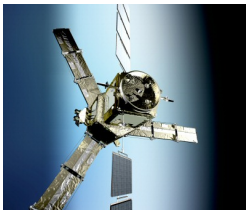


GT SPU "Traitement données spatiales"  
27 septembre, 2014

# SMOS images restoration from L1a data: A sparsity-based variational approach

Andrés Almansa (**Telecom ParisTech**)

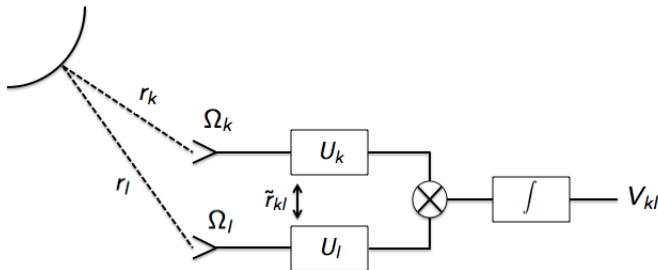
J. Preciozzi, P. Musé (**UdelaR**) S. Durand (**U. Paris Descartes**), A. Khazaal, B. Rougé (**CESBIO**)



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# Interferometry principle

- Measure the phase difference of incident radiation
- Cross-correlation between all pairs of receivers to obtain the *Visibility Function*  $V_{kl}$
- $T_b$  can be obtained indirectly from  $V_{kl}$



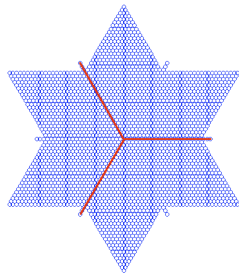
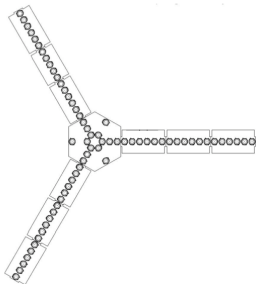
$$V_{k,l} = \frac{1}{\Omega_k \Omega_l} \iint_{\|\xi\| \leq 1} \frac{U_k(\xi) U_l^*(\xi) \tilde{r}_{kl}(t)}{\sqrt{1 - \|\xi\|^2}} (T_b(\xi) - T_r) e^{-i2\pi \mathbf{u}_{kl}^T \xi} d\xi$$

[Corbella et. al. 2004]

# The MIRAS instrument

## Antenna configuration

- Support of  $T_b$  is the unit circle
  - Optimum sampling grid on visibilities is a hexagonal grid
  - Two possible configurations: triangular or Y shaped arrays
  - Frequency coverage is larger for Y-shaped (but does not cover the entire hexagonal domain)



[Camps et al. 1998]

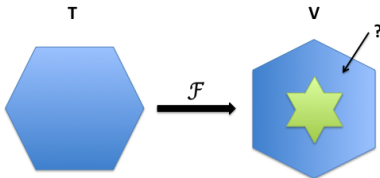
# Recovering $T_b$ from visibilities

$$V_{k,l} = \frac{1}{\Omega_k \Omega_l} \iint_{\|\xi\| \leq 1} \frac{U_k(\xi) U_l^*(\xi) \tilde{r}_{kl}(t)}{\sqrt{1 - \|\xi\|^2}} \overbrace{(T_b(\xi) - T_r)}^T e^{-i2\pi \mathbf{u}_{kl}^T \xi} d\xi$$

- Consider the discrete version of this linear operator, it can be stated by means of matrix  $\mathbf{G}$ :

$$\mathbf{G}T = V$$

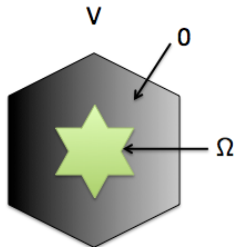
- $\dim(T) > \dim(V)$ : the problem is under constrained



# Generation of L1B: Zero padding regularization

Anterrieu 2004

$$\begin{aligned} \min_T & \|V - \mathbf{G}T\|_2^2 \\ \text{s.t.} & (I - P_\Omega)T = 0 \\ \text{with} & P_\Omega = \mathcal{F}^{-1}Z_\Omega Z_\Omega^* \mathcal{F} \end{aligned}$$



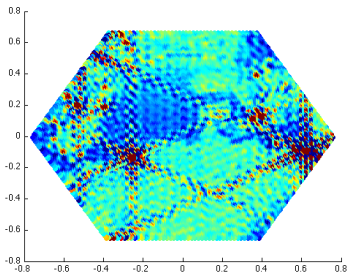
- This problem can be reformulated as:

$$\hat{T} = \arg \min_{\hat{t} \in \Omega} \|V - \overbrace{\mathbf{G}\mathcal{F}^{-1}Z_\Omega}^{\mathbf{A}} \hat{t}\|_2^2$$

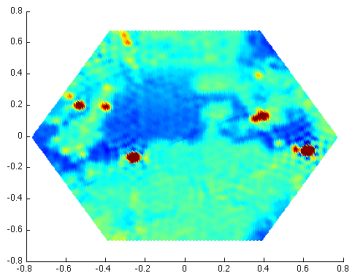
- $\hat{T}$  is SMOS L1B data product,  $V$  is SMOS L1A data product
- $T$  can be simply recovered from  $\hat{T}$  by  $T = \mathcal{F}^{-1}Z_\Omega \hat{T}$

# Zero padding limitations

- **Strong Gibbs effects:** illegal transmitters introduce **outliers**
- Poor spectral extrapolation: limited resolution



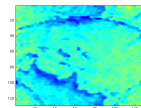
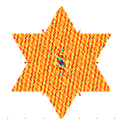
Direct Method:  $T_b = \mathcal{F}^{-1}(\hat{T})$



Regularized Method -  
Blackmann:  $T_b = \mathcal{F}^{-1}(B\hat{T})$

# Objectives of the present work

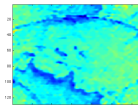
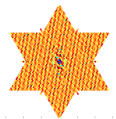
- Recover the brightness temperature **directly** from visibilities



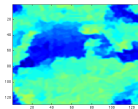
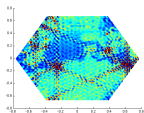


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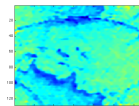
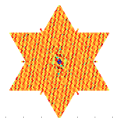


- Remove noise and signal effects generated from illegal emissions (outliers)

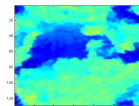
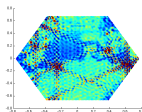


# Objetives of the present work

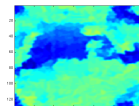
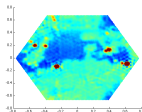
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- Remove noise and signal effects generated from illegal emissions (outliers)



- Extrapolate the image spectrum to minimize Gibbs effects



# Proposed method

## Main idea : Separate 3 sources

$u$  : Original brightness temperature image

$\Rightarrow$  TV semi-norm

$o$  : RFI Outliers

$\Rightarrow$  Sparsity norm ( $\ell^1$  or  $\ell^0$ )

$n$  : Gaussian measurement noise

$\Rightarrow \ell^2$  data fidelity term

## Proposed method

$$\min_{u,o} \left\{ \underbrace{\frac{1}{2} \|\mathbf{G}(o + u) - V\|_2^2}_{E_1(u,o)} + \underbrace{\lambda(\text{TV}_{\mathcal{H}}(u) + \mu S(o))}_{E_2(u,o)} \right\} \quad (1)$$

where

$\mu$  Trade-off between sparsity and regularity

$\lambda$  is chosen to satisfy  $\|\mathbf{G}(o + u) - V\|_2^2 \leq |\Omega|\sigma^2$

- $\text{TV}_{\mathcal{H}}(u)$ : *Total Variation* for  $\mathcal{H}$ -bandlimited images [Moisan 2007]  $\rightarrow$  reduces staircasing effect

[Moisan 2007] *How to discretize the total variation of an image?* Proc. Appl. Math. Mech., 2007

# Proposed Method: Numerical Implementation

Two stage process:

## Stage one

Solve the minimization problem with sparsity term  $S(o) = \|o\|_1$

- the problem is convex
- can be solved iteratively with a Forward-Backward algorithm
- converges to a global minimum

## Stage two

Starting from the previous solution, we solve the same problem with  $S(o) = \|o\|_0$

- the problem is non-convex due to the  $\ell_0$  norm
- for this functional the Forward-Backward algorithm converges to a local minimum [Blumensath and Davies 2005]

## Algorithm

The  $k$ -th iteration starting from seed  $x^0 = (u^0, o^0)$  is

$$\begin{cases} x^{k+1/2} &= x^k - \gamma \nabla E_1(x^k) \\ x^{k+1} &= \text{prox}_{\gamma E_2}(x^{k+1/2}). \end{cases}$$

## Differential operator

$$\nabla E_1(u, o) = (\mathbf{G}^* \mathbf{G}(u + o) - V, \mathbf{G}^* \mathbf{G}(u + o) - V).$$

## Proximal operators

- $\text{prox}_{\gamma E_2}(u, o) = (\text{prox}_{\gamma \lambda TV}(u), \text{prox}_{\gamma \lambda \mu S}(o))$
- $\text{prox}_{\gamma TV}$ : modified version of [Chambolle 2004] with spectral projection
- $\text{prox}_{\gamma \|\cdot\|_1}$ : the *soft-threshold* or *shrinkage* operator
- $\text{prox}_{\gamma \|\cdot\|_0}$ : the *hard-threshold* operator

## Implementation limitations

- $\mathbf{G}^*\mathbf{G}$  is a huge full matrix of size  $16384 \times 16384$
- Explicit multiplication by this matrix on each iteration is impractical.
- Change of basis to Fourier domain:

$$\nabla E_1(u, o) = F^*((\mathbf{GF}^*)^* \mathbf{GF}^* F(u + o) - (\mathbf{GF}^*)^* V)$$

- $\mathbf{FG}^*\mathbf{GF}^*$  is even bigger than  $\mathbf{G}^*\mathbf{G}$  ( $32768 \times 32768$ ) but highly sparse: to keep 99.99% we need 0.008 coefficients.

## Final algorithm

1 Set  $S(\cdot) = \|\cdot\|_1$

2 Initialize  $\lambda$

a Iterate until convergence (FB)

$$\begin{cases} u^{k+1/2} &= u^k - \gamma F^*(\mathbf{FG}^* \mathbf{GF}^* F(u + o) - \mathbf{FG}^* V) \\ o^{k+1/2} &= o^k - \gamma F^*(\mathbf{FG}^* \mathbf{GF}^* F(u + o) - \mathbf{FG}^* V) \\ u^{k+1} &= \text{prox}_{\gamma\lambda \text{TV}_{\mathcal{H}}}(u^{k+1/2}) \\ o^{k+1} &= s_{\gamma\lambda\mu}(o^{k+1/2}). \end{cases}$$

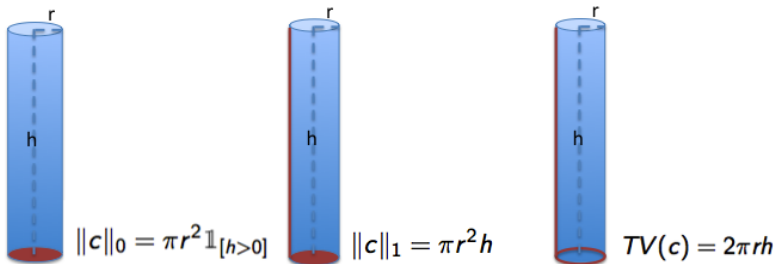
b If  $\|\mathbf{GF}^* F(o + u) - V\|_2^2 - |\Omega|\sigma^2| \leq \epsilon$ , update  $\lambda$  and go to a)

3 Set  $S(\cdot) = \|\cdot\|_0$  and go to 2)



# Numerical implementation: $\mu$ selection

We model an outlier as a cylinder  $c$  of radius  $r$  and height  $h$ :



## $\mu$ selection

- If  $\ell_1$  then  $c$  is an outlier if  $TV(c) \geq \mu \|c\|_1$ , leading to a  $\mu \leq \frac{2}{r}$
- If  $\ell_0$  then  $c$  is an outlier if  $TV(c) \geq \mu \|c\|_0$ , i.e.  $\mu \leq \frac{2h}{r}$

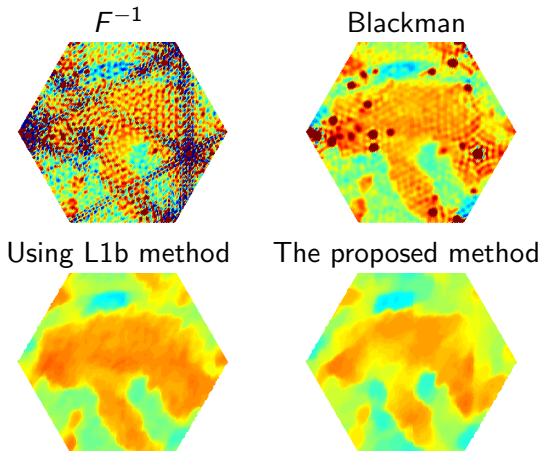
# Experiments on real data



Google Earth view of two of the regions used on the experiments. The left one corresponds to snapshot 996 and the right one to snapshot 1050

# Experiments on real data

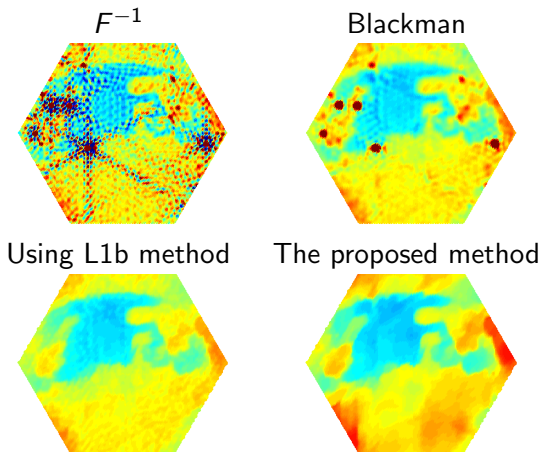
Comparison between previous works and our method.



This snapshot corresponds to Central Europe, with Italy clearly visible. Color scale mapped between 0 and 300 K.

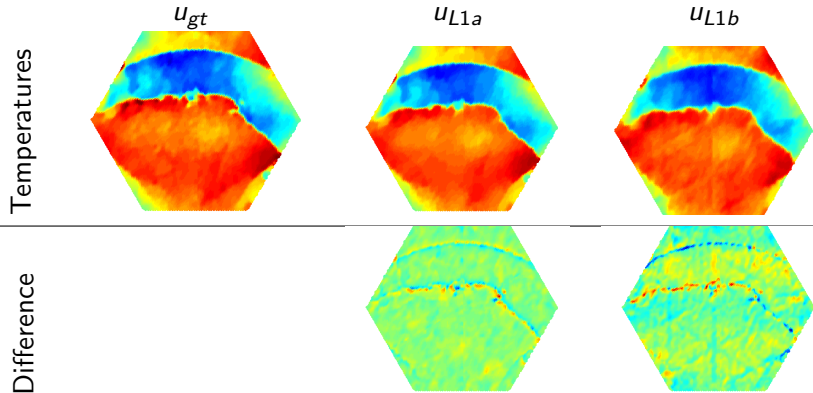
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# Experiments on simulated data



	$L_1$	$L_2$	$L_\infty$
$u_{L1a} - u_{gt}$	3.197598	5.418558	57.679203
$u_{L1b} - u_{gt}$	9.587280	12.994680	87.467700

Results from simulated data. The error is measured over all the image, not only the free of aliasing (FOA) zone.

We propose a variational method to restore images from the L1A SMOS data product.

- The method models the observations as the superposition of three components on the spatial domain:
  - The target brightness temperature map  $u$
  - The outliers image  $o$  due to the illegal emissions
  - A gaussian noise image  $n$
- The method is numerically tractable by a change of basis from spatial to spectrum domain
- The method also extrapolates the spectral domain of  $u$  thanks to the total variation semi-norm
- The method is general enough to be used for other restoration problems

- Better separation if RFI are finely localized [Veterli 2002], [Candes 2013], [Duval 2014]
- More detailed restoration with patch-based (non-local) regularization