Advance in particle detectors and new physics

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MPGD2009, Kolymbari, Crete, Greece





Some experiments using Bulk Micromegas









CLAS12G



Be Michigan TPC







T2K Micromegas TPC 3xTPCs, 6 end plates, 72 Micromegas



















50 μm and 25 μm gaps fabricated

Very good energy resolution -11% at 5.9 keV

- 5.5% at 22 keV
- < 1% with Am alpha source

Other advantages

- Flexible structure (cylinder)
- Good uniformity
- Low material
- Low radioactivity

Micromegas micro-bulk in CAST

Greek Institutes in CAST

- NCR "Demokritos", Athens, Greece
- Aristotle University of Thessaloniki, Greece
- University of Patras, Greece
- NTU Athens, Greece







Evolution of Micromegas CAST background



03/11/15

International Axion Observatory (IAXO)

- Towards a new generation axion helioscope
- Conceptual Design Report and Letter of Intent to CERN in preparation.
- **Goal**: 1-1.5 orders of magnitude in sensitivity to g_{ag} better than CAST



Optimised configuration: TOROIDAL with 8 bores 25 m long, 5 m diameter and a peak field of 6 T





Axion search exclusion plots



ILC TPC project - Large International collaboration

G. Aarons et al., arXiv:0709.1893, M. S. Dixit et al., NIMA 518 (2004) 521, M. Kobayashi et al., NIMA581(2007)265,



Momentum resolution=5x10⁻⁵

ILC TPC prototype with Micromegas





Event in DESY test beam



TPC Micromegas advantages

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- 16 z/cm

- Ion suppression .1%
- No ExB effect
- Great resolution $\sim 40 \ \mu m$
- Good energy resolution

Micromegas + micro-pixels

P. Colas et al., NIMA535(2004)506

surface: $1.4 \times 1.6 \text{ cm}^2$ Matrix of 256 x 256 pixel size: 55 x 55 μ m²



Nikhef, Saclay, Bonn collaboration Industrialization by Bonn is going on





Great resolution Single electron counting!!



Gas On Slimmed SIlicon Pixels (GOSSIP) Under study for ATLAS SLHC tracker



InGrid Industrial production assured by Bohn University

2015 successful test beam: large TPC read-out with 160 InGrid detectors (320cm²) Tested in DESY in the ILC TPC test facility



/home/bestbeam/TOS_SR5_quad_EventDisplay/RunData/ForBarbara/run_000143_data_003082_150404_04-45-49.txt

Applications in neutron detection

n-TOF MicroMegas-based neutron transparent flux monitor and profiler

F. Belloni et al., Mod.Phys.Lett. A28 (2013) 1340023



Micromégas Concept for Laser MégaJoule and **ICF** Facilities M. Houry et al., NIM,557(2006)648



neutron tomography



J. Pancin et al., NIMA, 592(2008)104





Petroleum search using neutron detectors *Neutron Logging idea by B. Pontecorvo*





The idea is to build a multilayer Micromegas with thin Bo-10 Goal:

- thermal neutron efficiency >>30%
- much bigger surface detection than He-3 detector

MICROMEGAS PHOTODETECTOR Reflective mode

J. Derre et al, NIM, vol. 449, (2000) p. 314–321



Towards pico-second timing *Collaboration: Saclay-CERN-Princeton aprooved by RD51*

E. Delagnes, G. Fanourakis, E. Ferrer Ribas, I. Giomataris, C. Godinot, D. González Díaz, U. Joshi, Changguo Lu, M. Kebbiri, K.T. McDonald, H. Muller, E. Olivieri, <u>T. Papaevangelou</u>, A. Peyaud, F. Resnati, L. Ropelewsky, R. Veenhof, S. White









Detector prototype



Encouraging results (17 ps) + growing interest!

Second part Spherical detector, light-dark matter search and neutrino physics

Radial TPC with spherical proportional counter read-out

Saclay-Thessaloniki-Saragoza



A Novel large-volume Spherical Detector with Proportional Amplification read-out, I. Giomataris *et al.*, JINST 3:P09007,2008



- Simple and cheap
- Large volume
- single read-out
- Robustness
- Good energy resolution
- Low energy threshold
- Efficient fiducial cut
- Low background capability

NEWS

Light dark matter search using the spherical detector

GOAL: EXPLORE DM MASSES DOWN TO 100 MEV Motivated by:

- Sub-keV energy threshold of the detector
- Versatility of the low-Z target (H, He, Ne)
- Low background capability of this design



University of Saragoza detector

system

Calibration window

HV filter & PreAmplifier

Gas

input

Gas output or air input



Spherical detector propagation





University of Thessaloniki detector

Future project 2m detector will be developed At SNOLAB (G. Gerbier et al.)



University of Tsinghua - HEP detector

NEWS collaborating Institutes and people - Spherical detector project

IRFU Saclay, E. Bougamont1, J. Derre, A. Dastgheibi-Fard, J. Galan, I. Giomataris, G. Gerbier, M. Gros, P. Magnier, J. P. Mols, X.F. Navick, B. Paul, P. Salin, G. Tsiledakis

LSM MODANE, P. Loiza, F. Piquemal, Univ. of Thessaloniki, I. Savvidis

NCR Dimokritos Athens, G. Fanourakis, Univ. of Ioannina, J. D. Vergados

Univ. of Saragoza, P. Iguaz, I. Irastorza, IHEP, Zhimin Wang, Ruiguang Wang

Univ. of Tsighua, C. Tao, Univ. of Shanghai Jiao Tong, K. Giboni, K. Ni



ыоподгариу

I Giomataris et al., JINST 3:P09007,2008., I Giomataris and J.D. Vergados, Nucl.Instrum.Meth.A530:330-358,2004, I. Giomataris and J.D. Vergados, Phys.Lett.B634:23-29,2006.

I. Giomataris et al. Nucl.Phys.Proc.Suppl.150:208-213,2006., S. Aune et al., AIP Conf.Proc.785:110-118,2005.

J. D. Vergados et al., Phys.Rev.D79:113001,2009., E Bougamont et al. arXiv:1010.4132 [physics.ins-det], 2010



Energy resolution ~ 6 % and 9 % for Cu and Cd

If rt ~ 0.0155 ms ==> R = 65 cm 0.014 ms ==> ~70% of signal

New low-energy calibration source *Argon-37*

Home made Ar-37 source: irradiating Ca-40 powder with fast neutrons 7x10⁶neutrons/s Irradiation time 14 days. Ar-37 emits K(2.6 keV) and L(260 eV) X-rays (35 d decay time)





First measurement with Ar-37 source Total rate 40 hz in 250 mbar gas, 8 mm ball 240 eV peak clearly seen A key result for light dark matter search



Search for light dark matter Detector installed at LSM end 2012: 60 cm, Pressure = up to 10 bar Gas targets: Ar, Ne, He, CH4



Internal contamination cleaning Goal: remove Po-210, Pb-210







Ist chemical cleaning of sphere

Conditions :

- Nitric acid (17 %)
- Temperature 10° C
- Cleaning by filling the spherical cavity
- > Washing by pure water
- Drying by hot nitrogen





2nd chemical cleaning of sphere

Conditions :

- Nitric acid (30 %)
- Temperature 30° C
- Cleaning by spray
- Washing by pure water
- Drying by hot nitrogen

Backround evolution of the detector



 β/γ rate evolution





Physic run (light-WIMPs research) Ne + CH₄ (0.7 %) P = 3 bar T = 30.5 jours



Light WIMP search results





Count

29

Shield improvement

During last month intervention (May 2015)

- New platform fabricated and installed to carry detector and shield
 - Chemical cleaning of internal copper shield plates
 - Total lead thickness = 15 cm (from 10cm)
 - Total copper thickness= 7 cm (from 5 cm)
 - Improved anti-radon tent



Summary:

background level among the best experiments Achieved with modest budget and manpower Combined with the low energy threshold and low-Z targets:

Sensitivity for very-light WIMPs of this experiment is out of competition

NEWS-SNOLAB project

Kingston, Saclay, Grenoble, LSM, Thessaloniki..... 2 m detector at 10 bar Pure water shield Funded by Canadian grant of excellence LOI recently approved by SNOLAB committee



Quenching factor measurements

Goal: measure QF down to 500 eV ion energy using the Grenoble MIMAC facility for H, He, Ne, CF4, Ar, Xe at various pressures







Recent investigations with a 15 cm sphere show the capability to measure 500 eV He-4 ions with an estimated QF of about 25% *Saclay, Grenoble, Thessaloniki, Queen's-Kingston*

Additional physics

Neutrino-nucleus coherent elastic scattering

 $v + N \rightarrow v + N \sigma \approx N^2 E^2$, D. Z. Freedman, Phys. Rev.D,9(1389)1974

High cross section but very-low nuclear recoil

Illustration: using the present prototype at 10 m from the reactor, after 1 day

Detector threshold (electrons)	1	2	3	4
Xe	105	32	3	0
Ar	42	24	9	4
Ne	18	12	7	4



Ev [MeV]

A dedicated Supernova detector

Simple and cost effective - Life time >> 1 century Through neutrino-nucleus coherent elastic scattering

Y. Giomataris, J. D. Vergados, Phys.Lett.B634:23-29,2006

Sensitivity for galactic explosion For p=10 Atm, R=2m, D=10 kpc, $U_v = 0.5 \times 10^{53}$ ergs # Number of events (after quenching, $E_{th} = 0.25$ keV) He Ne Ar Kr Xe Xe (with Nuc. F.F) 0.08 1.5 6.7 23.8 68.1 51.8

Idea : A world wide network of several of such dedicated Supernova detectors To be managed by an international scientific consortium and operated by students

0-v ββ Decay

- If 0-*v* decays occur, then: ٠
 - Neutrino mass $\neq 0$
 - Decay rate measures effective mass $\langle m_{\nu} \rangle$
 - Neutrinos are Majorana particles
 - Lepton number is not conserved
- Physics impact is great. ٠

Target >> 1000 Kgr and zero background



- Xenon is relatively safe and easy to enrich
- Natural abundance of 136 Xe is ~ 8%
- EXO and NEXT have 200 kg highly enriched in ¹³⁶Xe
- Low cost
- Pressure variation

High density is desirable to contain event But there is an upper limit! $\rho < 0.55 \text{ g/cm}^3$ Beyond this density, $\delta E/E$ deteriorates rapidly!







Current and future projects

Experiment	Isotope	Method	Resolution (% at $Q_{\beta\beta}$)	Efficiency	Background (10 ⁻³ c/keV/kg/y)	Isotope Mass (kg)	$m_{\scriptscriptstyle etaeta}$ (eV)
GERDA	⁷⁶ Ge	ionization	0.16	0.8	10-1	15-35	100-200
EXO	¹³⁶ Xe	liquid TPC	3.3	0.7	1-0.5	160	100-200
KamLAND- Zen	¹³⁶ Xe	scintillation	9.5	0.8	0.5-0.1	360-1000	90-190
CUORE	¹³⁰ Te	bolometer	0.19	0.9	10-1	206	60-160
NEXT	¹³⁶ Xe	gas TPC	0.7	0.3	0.2-0.06	90-1000	90-190
SNO+	¹⁵⁰ Nd	scintillation	6.5	0.5	10-1	40-500	40-140
SuperNEMO	⁸² Se	tracko-calo	4.0	0.3	0.4-0.06	7-100	200-300

ANEMOS project: Development of a high-pressure spherical detector demonstrator for neutrino less double beta decay search In collaboration with CNBG (F. Piquemal et al.,), CPPM (J. Busto et al.,) The goal is to reach a record low background level $<< 10^{-4}/\text{keV/Kg/y}$ and an energy resolution of .3% We target a sensitivity of m_{bb} <50 meV better than other experiments

ANEMOS strategy

- Demonstrate the ultra-low background capability with a small size 60 cm detector at a pressure of 50 bar
- The 2 m dark matter detector at SNOLAB will demonstrate the scaling (2 m detector) needed for the double beta decay experiment (1 ton scale Xe-136)
- The combination of the two detectors and the good energy resolution to be reached in the R@D phase, will give the proof of principle of an optimized DBD experiment as required for the next generation projects

A main idea is to take advantage of the detector simplicity to severely reduce backgrounds emitted from surrounding materials



ANEMOS demonstrator Identical size to SEDINE (60cm) Made of ultra-pure copper Transported with special castle to avoid Radon contamination and activation Installation inside the current LSM shield (improved)

I. Giomataris

ANEMOS Sensitivity

Simulation model

Sphere diameter: 2 m Xenon gas at 50 bar (1272 Kg) Vessel: Copper PNNL µBq/kg ²³²Th=.034, ²³⁸U=.13 Shield 30 cm copper PNNL

GEANT simulation by J. Galan

- We assume complete suppression of external backgrounds
- Backgrounds are generated by ²³²Th and ²³⁸U contamination of detector and shield

Results Background rate in the region of Q_{bb} (2.46 MeV) 1.54x10⁻⁵/keV/Kg/ year 8.x10⁻⁵/keV/Kg/ year (compared to 2x10⁻³/keV/Kg/ year of running experiments)

Double beta decay experiment with the spherical detector

If additional rejection is required: new idea

Background free double beta decay experiment, I. Giomataris, arXiv:1012.4289

The idea is to detect Cherenkov light emitted by two electrons and then reject background from single electrons (Compton scattering etc..)

Xenon-136 at high pressure of about 25-40 bar is ideal to keep high efficiency for double electrons, Good enough electron path and reduce multiple scattering

A simple read-out is the standard spherical detector signal combined with

CsI photocathode layer deposited at the internal vessel surface, inducing a delayed signal





CONCLUSIONS

- Micromegas detector is widely used in particle physics
- Industrial applications are under development
- A promising low background detector
- Light dark matter search down to 100 MeV