



# Daya Bay



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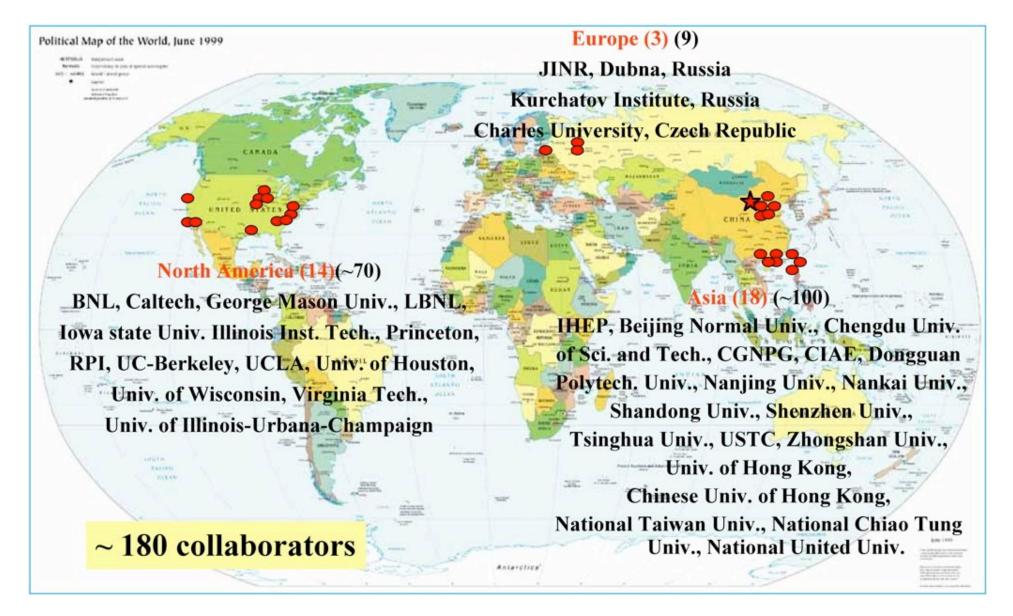
#### and

#### Lawrence Berkeley National Laboratory

For The Daya Bay Collaboration

GDR Neutrino at Universite Victor Segalen Bordeaux 2, France, 25 October, 2007

# Daya Bay Collaboration

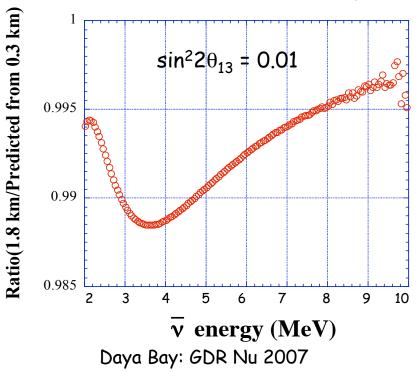




## Daya Bay: Goal And Approach

- Recommendation of the APS Neutrino Study Group:
  - An expeditiously deployed multidetector reactor experiment with sensitivity to  $\overline{\nu}_{e}$  disappearance down to  $\sin^{2} 2\theta_{13} = 0.01$ , an order of magnitude below present limits.
- Determine  $sin^2 2\theta_{13}$  with a sensitivity of  $\leq 0.01$

by measuring deficit in  $\overline{v}_e$  rate and spectral distortion.



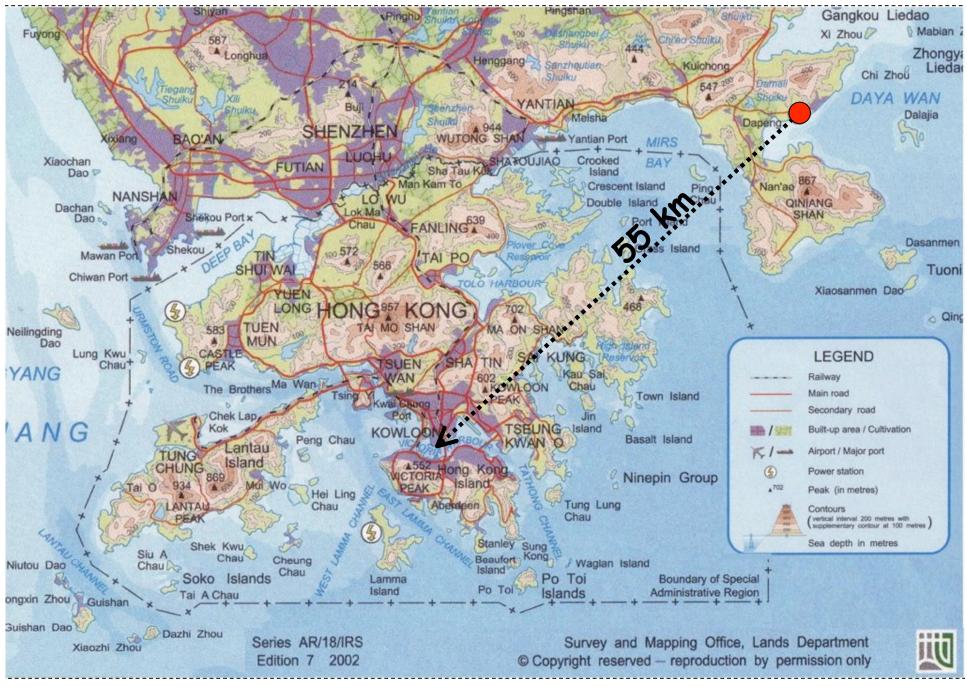
# How To Reach A Precision of 0.01 in Daya Bay?

#### Increase statistics:

- Use powerful nuclear reactors
- Utilize large target mass, hence large detectors
- Long running time

### • Suppress background:

- Go deep underground to gain overburden for reducing cosmogenic background
- Use active shield around the target
- Reduce systematic uncertainties:
  - Reactor-related:
    - Near and far detectors to minimize reactor-related errors
    - Optimize baseline for best sensitivity and smaller residual reactorrelated errors
  - Detector-related:
    - Use "Identical" pairs of detectors to do *relative* measurement
    - Comprehensive program in calibration/monitoring of detectors
    - Interchange near and far detectors (optional)



#### The Daya Bay Nuclear Power Complex

- 12th most powerful in the world (11.6  $GW_{th}$ )
- One of the top five most powerful by  $2011(17.4 \text{ GW}_{\text{th}})$
- Adjacent to mountain, easy to construct tunnels to reach underground labs with sufficient overburden to suppress cosmic rays

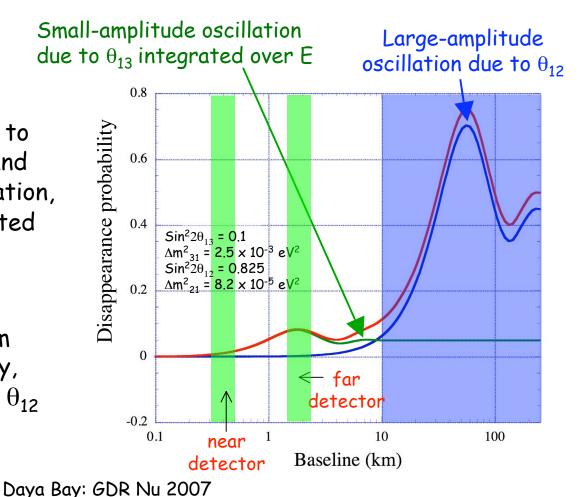


### Where To Place The Detectors ?

• Since reactor  $\overline{v}_e$  are low-energy, it is a disappearance experiment:

$$P(\overline{v}_e \rightarrow x) \approx \frac{\sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E}\right)}{4E} - \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left(\frac{\Delta m_{21}^2 L}{4E}\right)$$

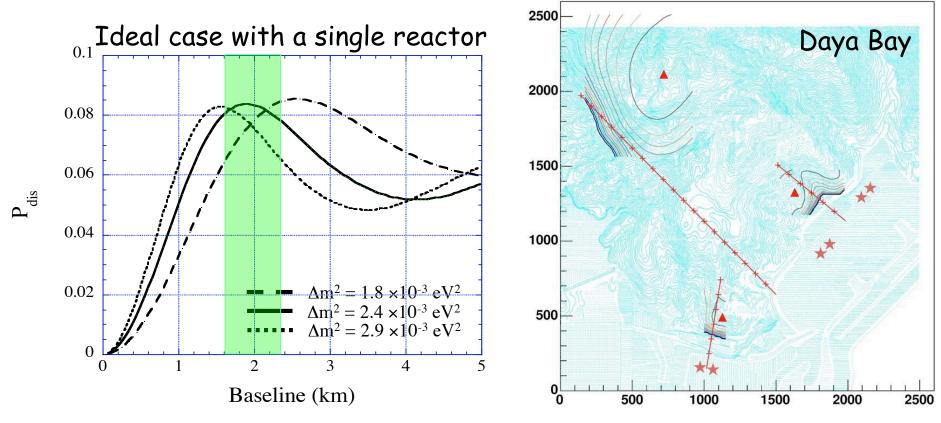
- Place near detector(s) close to reactor(s) to measure flux and spectrum of  $\overline{v_e}$  for normalization, hence reducing reactor-related systematic
- Position a far detector near the first oscillation maximum to get the highest sensitivity, and also be less affected by  $\theta_{12}$

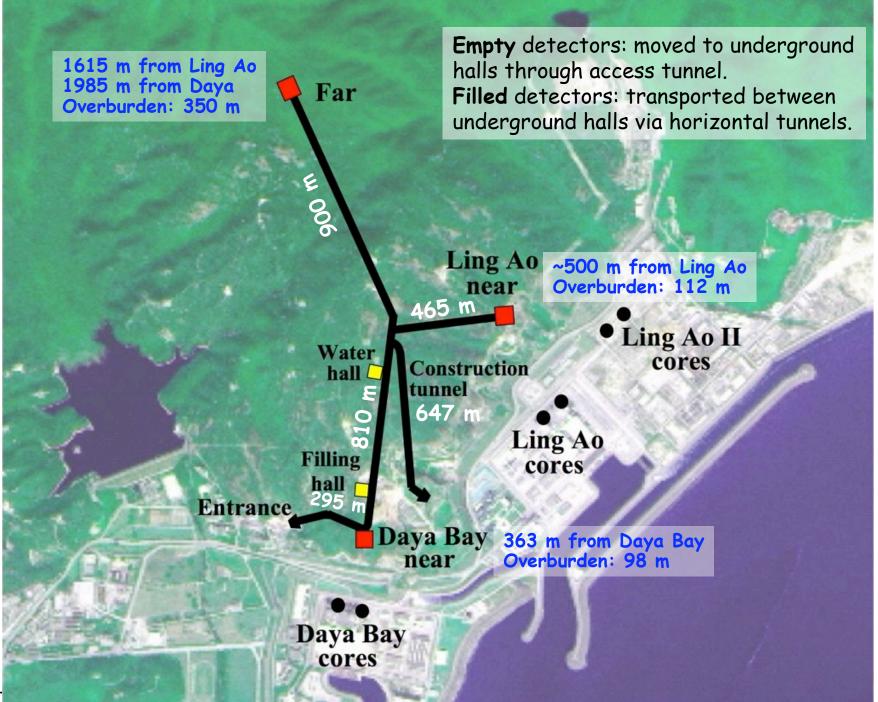


# Baseline optimization and site selection

#### Inputs to the process:

- Flux and energy spectrum of reactor antineutrino
- Systematic uncertainties of reactors and detectors
- Ambient background and uncertainties
- Position-dependent rates and spectra of cosmogenic neutrons and <sup>9</sup>Li

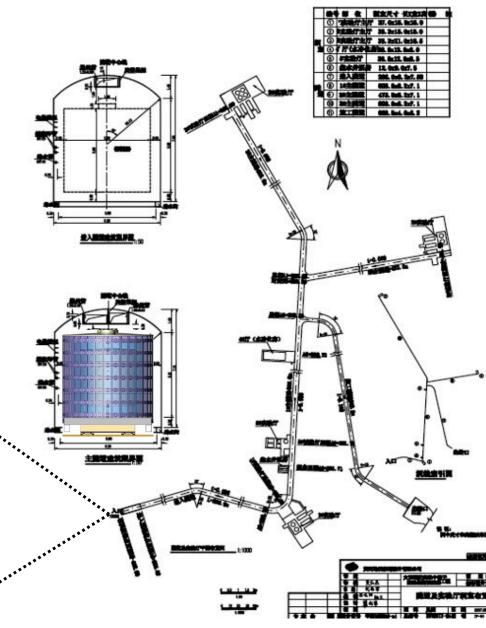




## **Civil Construction**

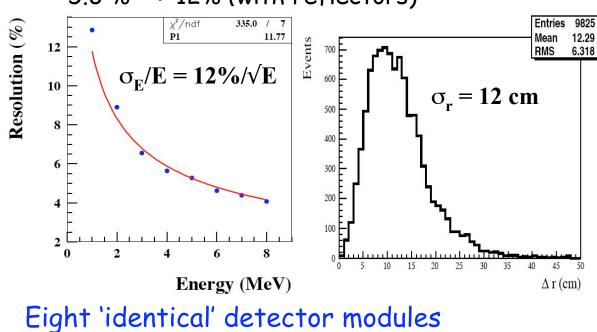
- $\cdot$  Total tunnel length is about 3100 m
- Estimated total construction cost is
  ~\$11 M
- Construction has begun this month
- First experimental hall is available after 13 months
- $\cdot$  Construction time is ~22 months

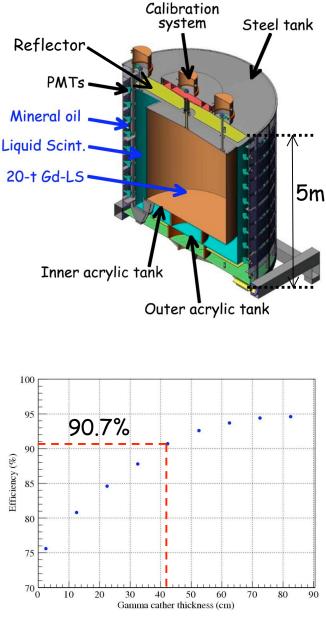




### Antineutrino Detectors

- Three-zone cylindrical detector design
  - Target: 20 T (0.1% Gd-LS), radius = 1.55 m
  - Gamma catcher: 20 T (LS), thickness = 0.42 m
  - Buffer : 40 T (mineral oil) , thickness = 0.48 m
- Low-background 8" PMT: 192
- Reflectors at top and bottom
- Photocathode coverage:
  5.6 % → 12% (with reflectors)



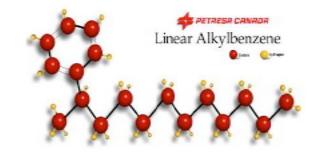


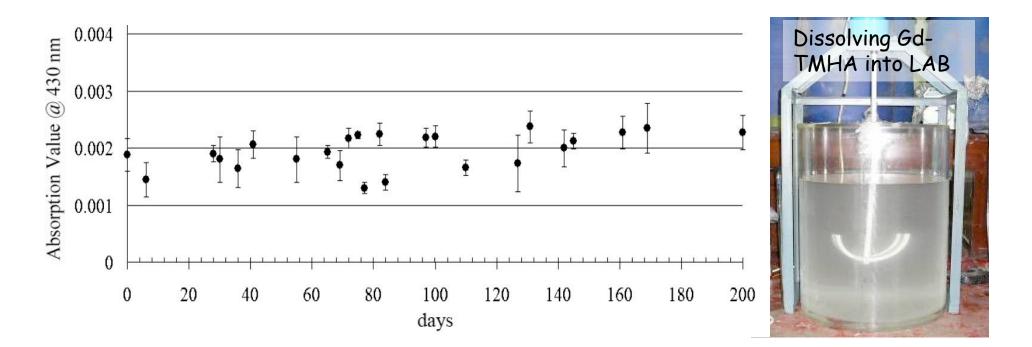
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### **Gd-Liquid Scintillator**

- Use LAB (Linear Alkyl Benzene):
  - high light yield, ~50% of anthracene
  - long attenuation length, > 10 m @ 430 nm
  - high flash point, ~130 °C
  - no EH&S issues
  - compatible with many plastics
  - cheap (raw material for making detergent)

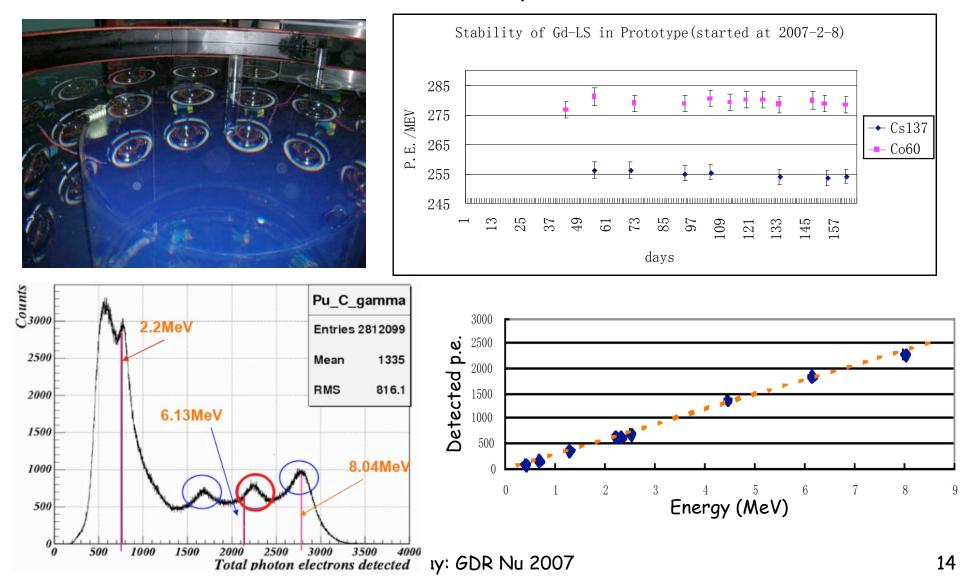




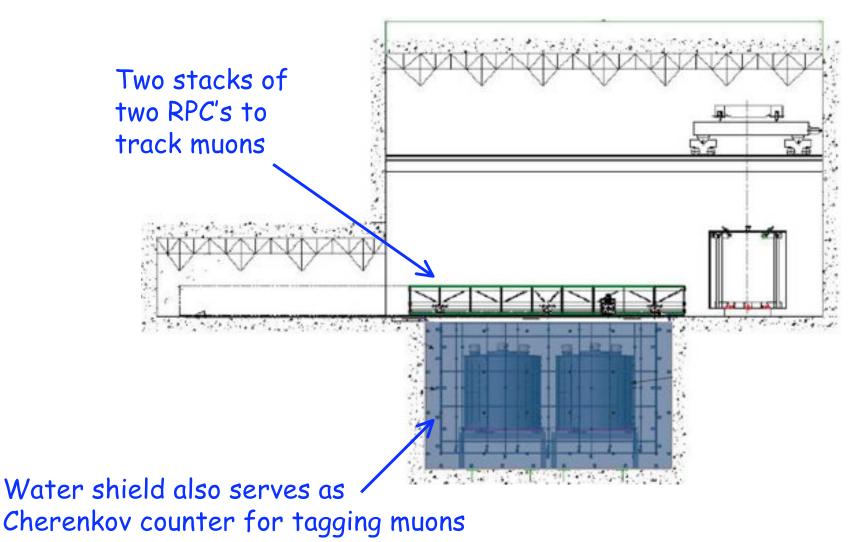


# 2-zone Prototype at IHEP

- 0.5 ton Gd-LS (it was unloaded LS for more than a year)
- 45 8" PMTs with reflectors at the top and bottom



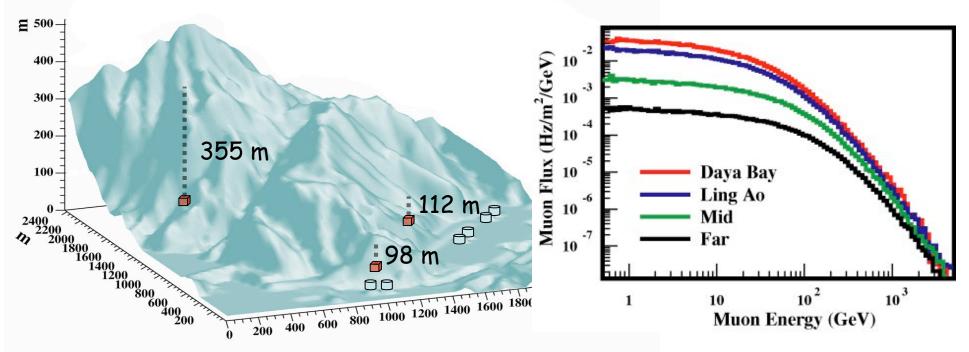
## An Example of the Muon System



 Combined efficiency of Cherenkov and tracker > 99.5% with error measured to better than 0.25%
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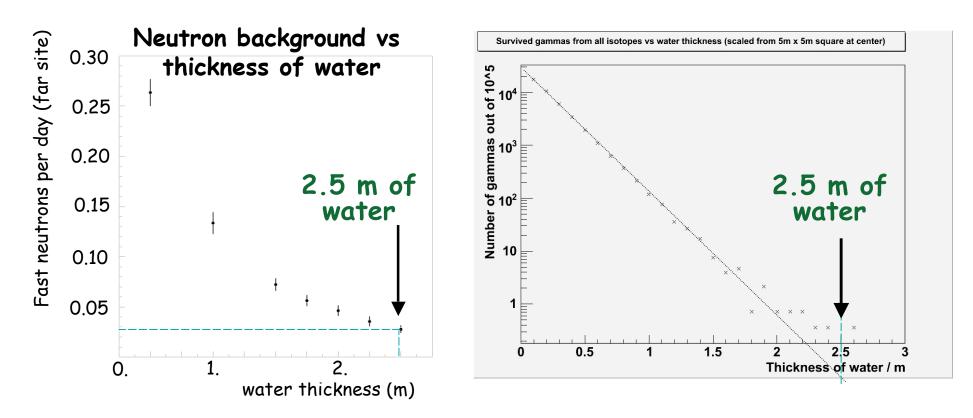
# Cosmic-ray Muon

- Use a modified Gaisser parametrization for cosmic-ray flux at surface
- Apply MUSIC and mountain profile to estimate muon intensity & energy



	DYB	LingAo	Far
Overburden (m)	98	112	355
Muon intensity (Hz/m²)	1.16	0.73	0.041
Mean Energy (GeV)	55	60	138

## Shielding Antineutrino Detectors



 Detector modules enclosed by 2.5 m of water to shield energetic neutrons produced by cosmic-ray muons and gamma-rays from the surrounding rock

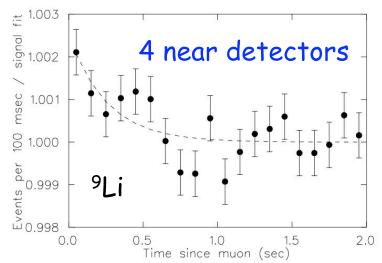
# Background

Uncorrelated background:

Sources: U/Th/K/Rn/neutron

Single gamma rate @ 0.9MeV < 50Hz/module Single neutron rate < 1000/day/module

- Correlated backgrounds: n ∝ E<sub>μ</sub><sup>0.75</sup> Fast Neutrons: double coincidence <sup>8</sup>He/<sup>9</sup>Li: beta-neutron emitting decays
- All these background events can be measured and corrected



Per module	Daya Bay Near	Ling Ao Near	Far Hall	
Baseline (m)	363	481 from Ling Ao	1985 from Daya Bay	
		526 from Ling Ao II	1615 from Ling Ao's	
Overburden (m)	98	112	350	
Radioactivity (Hz)	<50	<50	<50	
Muon rate (Hz)	36	22	1.2	
Antineutrino Signal (events/day)	930	760	90	
Accidental Background/Signal (%)	<0.2	< 0.2	<0.1	
Fast neutron Background/Signal (%)	0.1	0.1	0.1	
<sup>8</sup> He+ <sup>9</sup> Li Background/Signal (%)	0.3	0.2	0.2	

## Systematic Uncertainties

• Reactor-related:

Near detectors really help !

			S	really help :
Number of cores	$\alpha$	$\sigma_{\rho}$ (power)	$\sigma_{\rho}(\text{location})$	$\sigma_{\rho}(\text{total})$
4	0.338	0.035%	0.08%	0.087%
6	0.392	0.097%	0.08%	0.126%

#### • Detector-related:

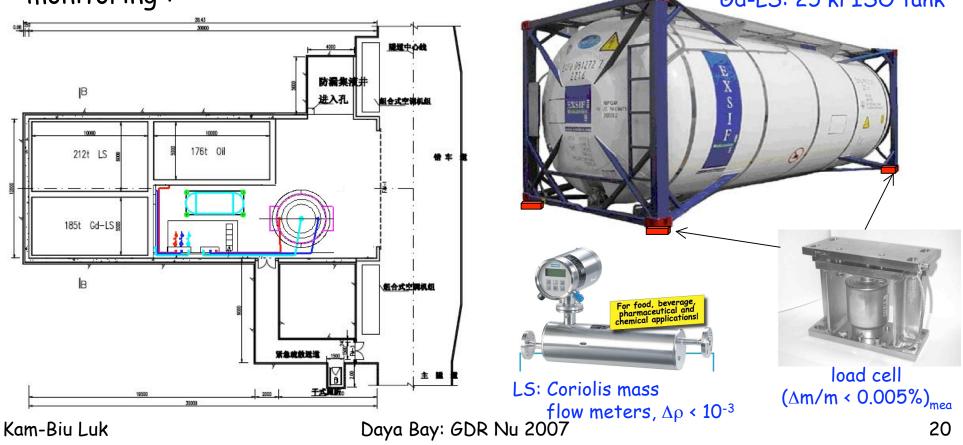
Source of uncertainty		Chooz Daya Bay ( <i>relative</i> )			(relative)
		(absolute)	Baseline	Goal	Goal w/Swapping
# protons	H/C ratio	0.8	0.2	0.1	0
	Mass	-	0.2	0.02	0.006
Detector	Energy cuts	0.8	0.2	0.1	0.1
Efficiency	Position cuts	0.32	0.0	0.0	0.0
	Time cuts	0.4	0.1	0.03	0.03
	H/Gd ratio	1.0	0.1	0.1	0.0
	n multiplicity	0.5	0.05	0.05	0.05
	Trigger	0	0.01	0.01	0.01
	Live time	0	< 0.01	< 0.01	< 0.01
Total detector-related uncertainty		1.7%	0.38%	0.18%	0.12%
			anticipated		
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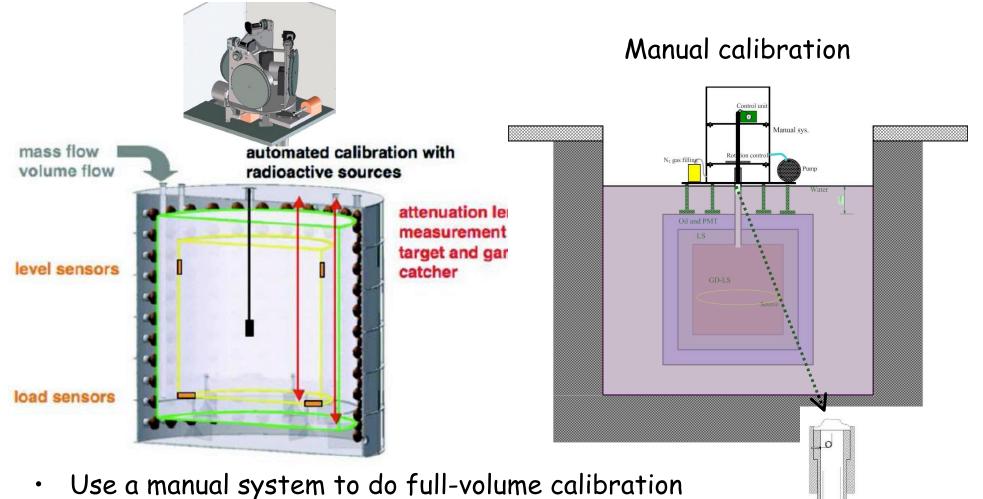
#### **Controlling Target Mass & Composition**

 Use one batch of 400-t pure LAB to prepare the scintillators stored in two underground tanks

• Fill two detectors sequentially at a time underground with controls of load cells (0.008% accuracy), Coriolis mass flow meters (reproducible at 0.1%), and volume flow meters (good to 0.02%), as well as temperature Gd-LS: 25 kl ISO tank

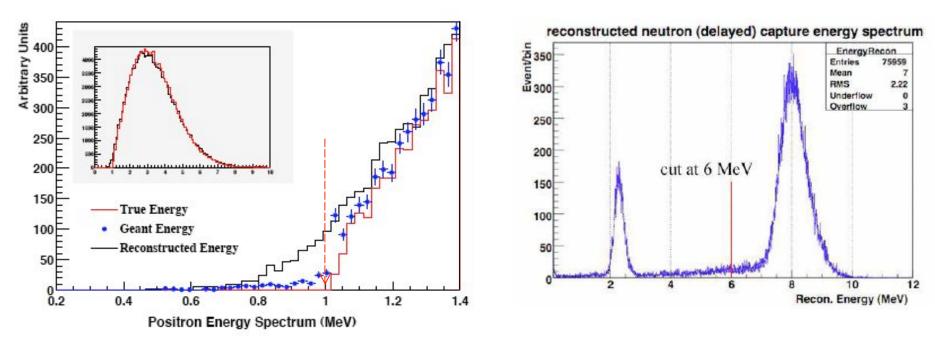


#### **Calibration And Monitoring**



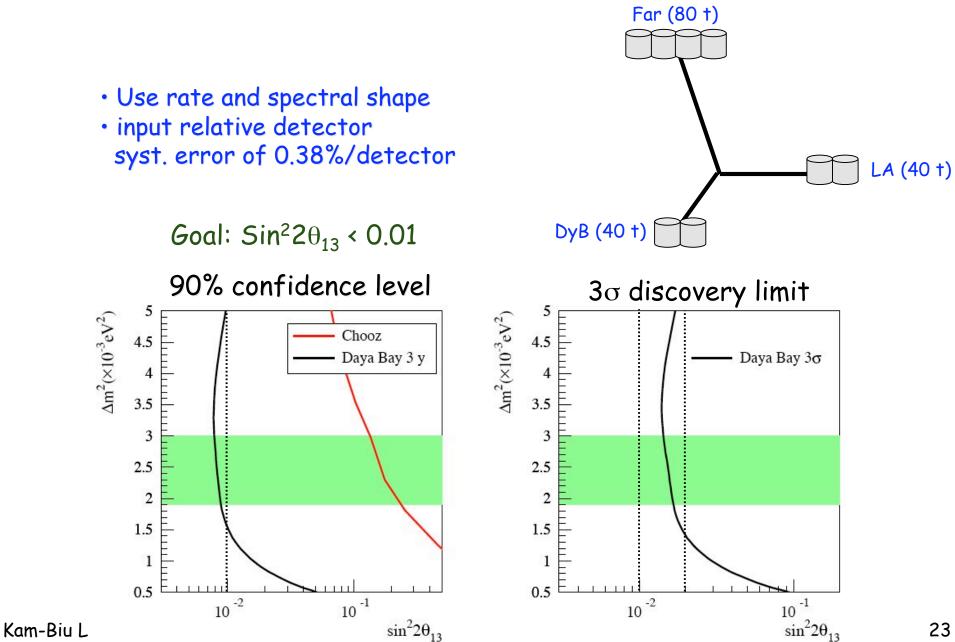
- Monitor level, load and temperature sensors regularly
- Frequent calibration with radioactive sources and LED along vertical axes with an automated calibration system

#### **Calibrating Energy Cuts**



- Routine calibration with •
  - 68Ge: e<sup>+</sup> annihilation yields two 0.511 MeV γ's
  - 60Co: 2.6 MeV in  $\gamma$ 's
  - $^{238}$ Pu- $^{13}$ C ( $\alpha$ -n source):

## Sensitivity of Daya Bay



# Milestones

- US CD-1 review
- Chinese 3-year project funding began to flow
- Start civil construction
- Groundbreaking ceremony

Apr 07 Apr 07 Oct 07 Oct 13,2007



- US CD-2/3a review
- Commission of 2 detectors at Daya Bay Hall
- Begin data taking with 8 detectors in final configuration Oct 10

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Jan 08 Sept 09 Oct 10

# Summary

- Daya Bay will reach a sensitivity of  $\leq 0.01$  for  $\sin^2 2\theta_{13}$
- Detailed design of detectors is nearly complete !!
- Received commitment from Chinese funding agencies
- On track of getting CD2/3a approval in January 2008
- Civil construction has begun:
  - Groundbreaking on Oct 13, 2007
  - Daya Bay Hall beneficial occupancy in 2008 First data with two detectors in Daya Bay Hall in 2009
  - Ling Ao Hall and Far Hall beneficial occupancy in 2009 Data taking with all detectors in 2010
- Daya Bay is sailing forwards

Merci!

# Aberdeen Tunnel Experiment (Hong Kong)

For studying cosmic muons, spallation neutrons, and gamma background

