## Strong EW phase transition from varying Yukawas

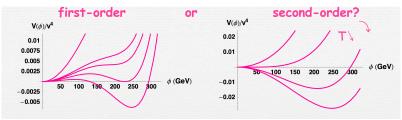
#### lason Baldes



Higgs Hunting - Paris, 1 September 2016

## Electroweak baryogenesis - Requirements

Electroweak baryogenesis — an possible link between the Higgs and cosmology.



## Electroweak baryogenesis requires:

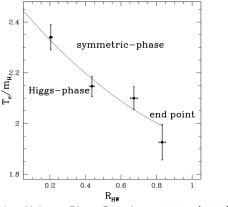
- A strong first order phase transition  $(\phi_c/T_c \gtrsim 1)$
- Sufficient CP violation

#### However in the SM:

- The Higgs mass is too large
- Quark masses are too small

## Electroweak phase transition

Lattice calculations show the SM Higgs mass is too large.



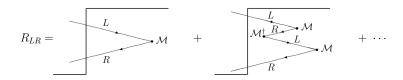
 $R_{HW} \equiv m_H/m_W$ 

Endpoint at:  $m_H \approx 67 \text{ GeV}$ 

- Csikor, Fodor, Heitger, Phys. Rev. Lett. 82, 21 (1999)

The Higgs potential must be modified.

## Baryogenesis from charge transport with SM CP violation



$$\epsilon_{
m CP} \sim rac{1}{M_W^6 T_c^6} \prod_{i>j top u.c.t} (m_i^2 - m_j^2) \prod_{i>j top d.s.b} (m_i^2 - m_j^2) J_{
m CP}$$

- Gavela, Hernandez, Orloff, Pène, Quimbay [hep-ph/9406289],
- Huet, Sather [hep-ph/9404302].

SM quark masses are too small!

## Solutions to the flavour puzzle

#### Yukawa interactions:

$$y_{ij}\overline{f}_L^i\Phi^{(c)}f_R^j$$

#### Possible solutions

- Froggatt-Nielsen
- Composite Higgs
- Randall-Sundrum Scenario

#### Froggatt-Nielsen Yukawas:

$$y_{ij} \sim \left(\frac{\langle \chi \rangle}{\Lambda}\right)^{-q_i+q_j+q_H}$$

Some previous work: Baryogenesis from the Kobayashi-Maskawa phase

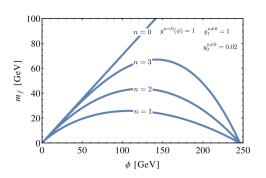
- Berkooz, Nir, Volansky - Phys. Rev. Lett. 93 (2004) 051301

Split fermions baryogenesis from the Kobayashi-Maskawa phase

- Perez, Volansky - Phys. Rev. D 72 (2005) 103522

## Varying Yukawas

Study the strength of the EWPT with varying Yukawas in a <u>model</u> independent way. - IB, Konstandin, Servant (1604.04526)

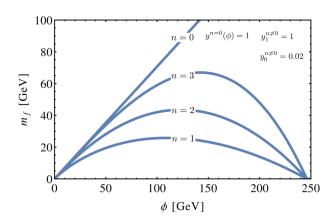


#### Ansatz

$$y(\phi) = \begin{cases} y_1 \left( 1 - \left[ \frac{\phi}{v} \right]^n \right) + y_0 & \text{for } \phi \leq v, \\ y_0 & \text{for } \phi \geq v. \end{cases}$$

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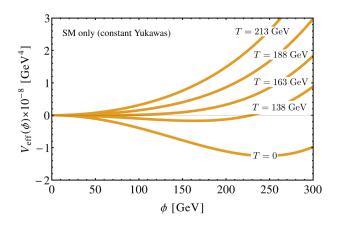
#### **Effective Potential**



## Thermal correction

$$V_{\mathrm{eff}} \supset -\frac{g_*\pi^2}{90} T^4$$

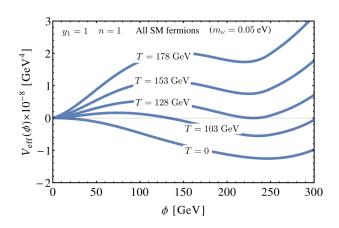
#### Effective Potential - SM case



Second order phase transition  $T_c = 163 \text{ GeV}$ .

$$V_{\mathrm{eff}} = V_{\mathrm{tree}}(\phi) + V_{1}^{0}(\phi) + V_{1}^{T}(\phi, T) + V_{\mathrm{Daisy}}(\phi, T)$$

# Effective Potential - Varying Yukawas



## Strong first order phase transition

$$\phi_c = 230 \text{ GeV}$$

$$\phi_c=$$
 230 GeV,  $T_c=$  128 GeV,  $\phi_c/T_c=$  1.8

$$\phi_c/T_c = 1.8$$

#### Effective Potential - T = 0 terms

$$V_{ ext{eff}} = V_{ ext{tree}}(\phi) + V_1^0(\phi) + V_1^T(\phi, T) + V_{ ext{Daisy}}(\phi, T)$$
 $V_{ ext{tree}}(\phi) = -rac{\mu_\phi^2}{2}\phi^2 + rac{\lambda_\phi}{4}\phi^4$ 

$$V_1^0(\phi) = \sum_i \frac{g_i(-1)^F}{64\pi^2} \left\{ m_i^4(\phi) \left( \text{Log} \left[ \frac{m_i^2(\phi)}{m_i^2(v)} \right] - \frac{3}{2} \right) + 2m_i^2(\phi) m_i^2(v) \right\}$$

Gives a large negative contribution to the  $\phi^4$  term.

- $\bullet$  Can lead to a new minimum between  $\phi=$  0 and  $\phi=$  246 GeV.
- Not an issue for previous  $y_1 = 1$ , n = 1 example.
- Can make phase transition weaker.

# Effective Potential - one-loop $T \neq 0$ correction

$$V_1^T(\phi, T) = \sum_i \frac{g_i(-1)^F T^4}{2\pi^2} \times \int_0^\infty y^2 \text{Log}\left(1 - (-1)^F e^{-\sqrt{y^2 + m_i^2(\phi)/T^2}}\right) dy$$
$$V_f^T(\phi, T) = -\frac{gT^4}{2\pi^2} J_f\left(\frac{m_f(\phi)^2}{T^2}\right)$$

$$J_{f}\left(\frac{m_{f}(\phi)^{2}}{T^{2}}\right) \approx \frac{7\pi^{4}}{360} - \frac{\pi^{2}}{24} \left(\frac{m}{T}\right)^{2} - \frac{1}{32} \left(\frac{m}{T}\right)^{4} \operatorname{Log}\left[\frac{m^{2}}{13.9T^{2}}\right], \quad \text{for } \frac{m^{2}}{T^{2}} \ll 1,$$

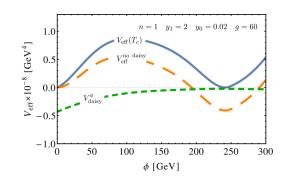
$$\delta V \equiv V_{f}^{T}(\phi, T) - V_{f}^{T}(0, T)$$

$$\approx \frac{gT^{2}\phi^{2}[y(\phi)]^{2}}{96}$$

$$\approx \frac{gT^{2}\phi^{2}[y(\phi)]^{2}}{96}$$

φ [GeV]

## Effective Potential - daisy correction

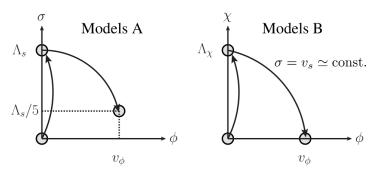




$$V_{\text{Daisy}}^{\phi}(\phi, T) = \frac{T}{12\pi} \left\{ m_{\phi}^{3}(\phi) - \left[ m_{\phi}^{2}(\phi) + \Pi_{\phi}(\phi, T) \right]^{3/2} \right\}$$
$$\Pi_{\phi}(\phi, T) = \left( \frac{3}{16} g_{2}^{2} + \frac{1}{16} g_{Y}^{2} + \frac{\lambda}{2} + \frac{y_{t}^{2}}{4} + \frac{gy(\phi)^{2}}{48} \right) T^{2}$$

## Including the flavon

Flavor Cosmology: Dynamical Yukawas in the Froggatt-Nielsen Mechanism - IB, Konstandin, Servant (1608.03254)

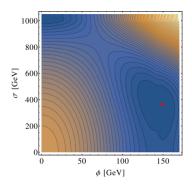


- Have to take into account constraints from flavour physics.
- Flavon dof also affects  $\phi_c/T_c$ .
- Generic prediction: light flavon with mass below the EW scale.

We have implemented this idea in some non-standard Froggatt-Nielsen scenarios.

## Expermental signatures - Model A-2

Couple a flavon to each mass eigenstate - Knapen, Robinson (1507.00009)

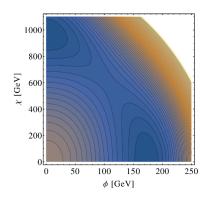


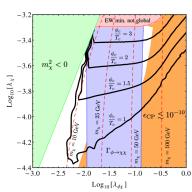
$$\Lambda_s=1$$
 TeV,  $\lambda_s=10^{-5}$  ,  $\lambda_{\phi s}=10^{-3.5}$  ,  $m_\sigma=0.75$  GeV,  $\epsilon_s\equiv v_s/(\sqrt{2}\Lambda_s)=0.12$ 

$$\mathcal{L} \supset \frac{y_b}{\sqrt{2}} \left(\frac{\sigma}{\sqrt{2}\Lambda_c}\right)^2 \phi \overline{b} b \qquad \mathrm{Br}(\phi \to \overline{b} b \sigma) = 1.1\% \left(\frac{0.1}{\epsilon_s}\right)^2 \left(\frac{1~\mathrm{TeV}}{\Lambda_s}\right)^2$$

## Model B-1: $Q_{\rm FN}(X) = -1/2$ - phase transition strength

Here we assume a simple polynomial scalar potential up to dimension four + the Yukawa sector.





$$\Lambda_{\chi}=1$$
 TeV,  $\lambda_{\chi}=10^{-4}$ ,  $\lambda_{\phi\chi}=10^{-2}$ ,  $\emph{m}_{\chi}=14$  GeV

$$\Gamma(\chi o \overline{c}c) \approx 10^{-12} \; {
m GeV} \left( rac{m_\chi}{10 \; {
m GeV}} 
ight) \left( rac{v_\chi^{
m today}}{1 \; {
m GeV}} 
ight)^2 \left( rac{1 \; {
m TeV}}{\Lambda_\chi} 
ight)^4$$

#### Conclusions

The Higgs may have links to cosmology with experimentally accessible signatures.

#### Yukawa variation may allow us to address:

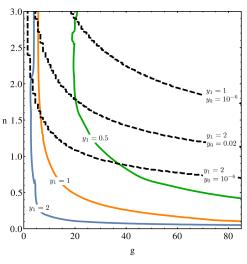
- The lack of a strong first order phase transition in the SM
- The insufficient CP violation for EW baryogenesis
  - Bruggisser, Konstandin, Servant (in preperation)
- The related limits on EDMs (this approach leads to a lack of EDM signals)

This offers additional motivation to consider low scale flavour models and their cosmology.

Other models of flavour are worth looking at too (not just Froggatt-Nielsen). e.g. RS1 - von Harling, Servant (in preperation)

New experimental signatures should then be accessible as we further probe the Higgs potential!

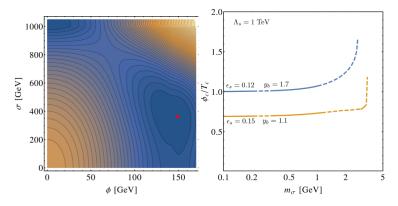
## Strength of the phase transition with varying Yukawas



$$y(\phi) = y_1 \left(1 - \left[\frac{\phi}{v}\right]^n\right) + y_0 \text{ for } \phi \le v$$

## Model A-2: Disentangled hierarchy and mixing mechanism

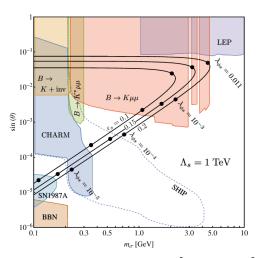
Couple a flavon to each mass eigenstate - Knapen, Robinson (1507.00009)



Here we assume a simple polynomial scalar potential up to dimension four augmented with a  $\sigma$  dependent Yukawa term.

$$\epsilon_s \equiv \frac{\langle \sigma \rangle}{\sqrt{2} \Lambda_s}$$
  $\mathcal{L} \supset \frac{y_b}{\sqrt{2}} \left( \frac{\sigma}{\sqrt{2} \Lambda_s} \right)^2 \phi \overline{b} b$ 

# Constraints on disentangled flavour and hierarchy mechanism

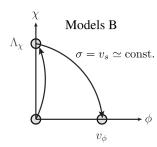


$$\mathrm{Br}(\phi 
ightarrow \overline{b}b\sigma) = 1.1\% \left(rac{0.1}{\epsilon_s}
ight)^2 \left(rac{1~\mathrm{TeV}}{\Lambda_s}
ight)^2$$

## Models B

#### Two FN fields

$$\mathcal{L} = \tilde{y_{ij}} \left( \frac{S}{\Lambda_s} \right)^{\tilde{n}_{ij}} \overline{Q}_i \tilde{\Phi} U_j + y_{ij} \left( \frac{S}{\Lambda_s} \right)^{n_{ij}} \overline{Q}_i \Phi D_j$$
$$+ \tilde{f}_{ij} \left( \frac{X}{\Lambda_{\chi}} \right)^{\tilde{m}_{ij}} \overline{Q}_i \tilde{\Phi} U_j + f_{ij} \left( \frac{X}{\Lambda_{\chi}} \right)^{m_{ij}} \overline{Q}_i \Phi D_j$$



We assume a small VEV for the second FN field today:  $\langle X \rangle \simeq 0$ . The VEV  $\langle S \rangle$  sets the Yukawas today while  $\langle X \rangle$  varies during the EWPT.

#### Model B-1: $Q_{FN}(X) = -1/2$

$$\Lambda_\chi \gtrsim 700 \; {
m GeV} \; (K - \overline{K})$$

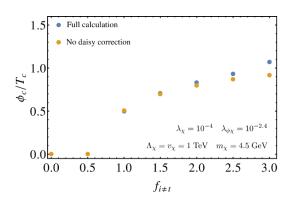
$$\Lambda_\chi \gtrsim 250 \,\, {
m GeV} \,\, (B_s - \overline{B_s})$$

Model B-2: 
$$Q_{FN}(X) = -1$$

$$\Lambda_{\chi} \gtrsim 2.5 \text{ TeV } (K - \overline{K})$$

$$\sqrt{\Lambda_\chi m_\chi} \gtrsim 500 \; {
m GeV} \; (D - \overline{D})$$

## Model B-1: $Q_{\rm FN}(X) = -1/2$ - phase transition strength



$$\mathcal{L}\supset ilde{f}_{ij}\left(rac{X}{\Lambda_{\Upsilon}}
ight)^{ ilde{m}_{ij}}\overline{Q}_{i} ilde{\Phi}U_{j}+f_{ij}\left(rac{X}{\Lambda_{\Upsilon}}
ight)^{m_{ij}}\overline{Q}_{i}\Phi D_{j}+H.c.$$