

Data Processing and Imaging for Cas A

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Observed Visibility

The observed visibility is

$$\begin{aligned} V_{ij} &= \langle F_i F_j^* \rangle \\ &= g_i g_j^* \int d^2 \hat{\mathbf{n}} d^2 \hat{\mathbf{n}}' A_i(\hat{\mathbf{n}}) A_j^*(\hat{\mathbf{n}}') \langle \varepsilon(\hat{\mathbf{n}}) \varepsilon(\hat{\mathbf{n}}') \rangle e^{2\pi i (\hat{\mathbf{n}} \cdot \mathbf{u}_i - \hat{\mathbf{n}}' \cdot \mathbf{u}_j)} + n_{ij} \\ &= g_i g_j^* \int d^2 \hat{\mathbf{n}} d^2 \hat{\mathbf{n}}' A_i(\hat{\mathbf{n}}) A_j^*(\hat{\mathbf{n}}') I(\hat{\mathbf{n}}) \delta(\hat{\mathbf{n}} - \hat{\mathbf{n}}') e^{2\pi i (\hat{\mathbf{n}} \cdot \mathbf{u}_i - \hat{\mathbf{n}}' \cdot \mathbf{u}_j)} + n_{ij} \\ &= g_i g_j^* \int d^2 \hat{\mathbf{n}} A_i(\hat{\mathbf{n}}) A_j^*(\hat{\mathbf{n}}) I(\hat{\mathbf{n}}) e^{2\pi i \hat{\mathbf{n}} \cdot \mathbf{u}_{ij}} + n_{ij}, \end{aligned} \tag{1}$$

where $\mathbf{u}_{ij} = \mathbf{u}_i - \mathbf{u}_j = (\mathbf{r}_i - \mathbf{r}_j)/\lambda$.

Calibrator

For the calibrator $I(\hat{\mathbf{n}}) = S_c \delta(\hat{\mathbf{n}} - \hat{\mathbf{n}}_0)$,

$$V_{ij}^{\text{calibrator}} = g_i g_j^* A_i(\hat{\mathbf{n}}_0) A_j^*(\hat{\mathbf{n}}_0) S_c e^{2\pi i \hat{\mathbf{n}}_0 \cdot \mathbf{u}_{ij}} + n_{ij}, \quad (2)$$

so

$$V'_{ij} = \frac{V_{ij}}{S_c A_i(\hat{\mathbf{n}}_0) A_j^*(\hat{\mathbf{n}}_0) e^{2\pi i \hat{\mathbf{n}}_0 \cdot \mathbf{u}_{ij}}} = g_i g_j^* + n'_{ij}. \quad (3)$$

Calibration

The eigen-decomposition of the matrix V'_{ij} is

$$V'_{ij}x_j = \lambda x_i, \quad (4)$$

that is

$$V'_{ij} = \lambda x_i x_j^*, \quad (5)$$

Compare 3 and 5, we get

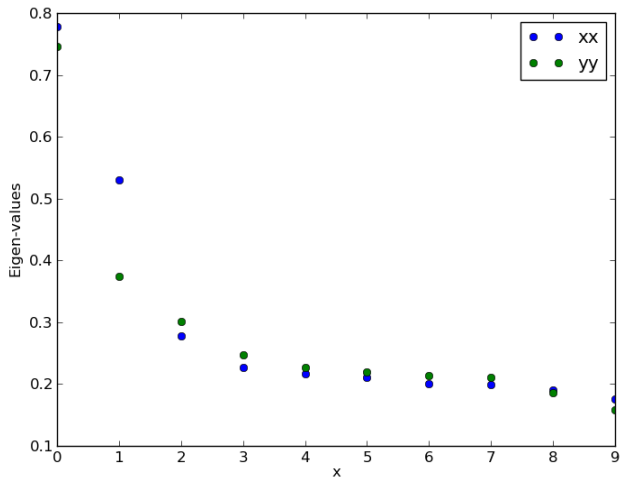
$$g_i = \sqrt{\lambda} x_i. \quad (6)$$

Do the calibration by using the gain in 6,

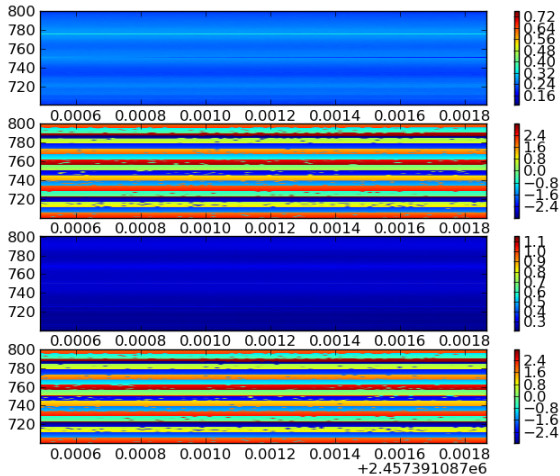
$$V_{ij}^{\text{cal}} = \frac{V_{ij}}{g_i g_j^*} = \int d^2 \hat{n} A_i(\hat{n}) A_j^*(\hat{n}) I(\hat{n}) e^{2\pi i \hat{n} \cdot \mathbf{u}_{ij}}, \quad (7)$$

here I have ignored the noise term.

Eigen-Values

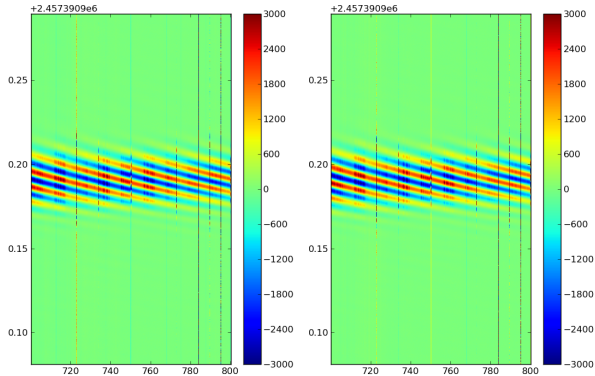


Gain for Antenna 3 — Amplitude and Phase



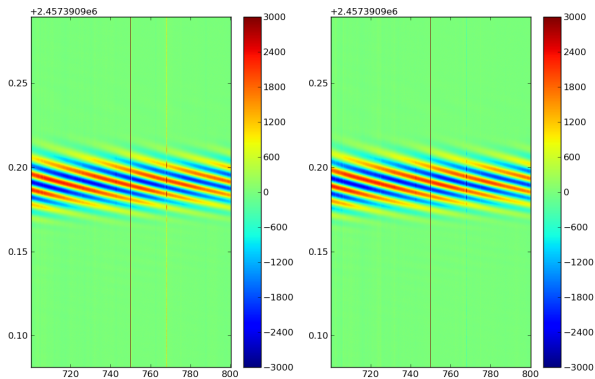
Visibility After Calibration

XX correlation



Visibility After Calibration

YY correlation



Phase to Cas A

Now choose a phase reference point (Cas A) $\hat{s}_0$ such that $\hat{n} = \hat{s}_0 + \boldsymbol{\sigma}$, then the calibrated visibility can be expressed as

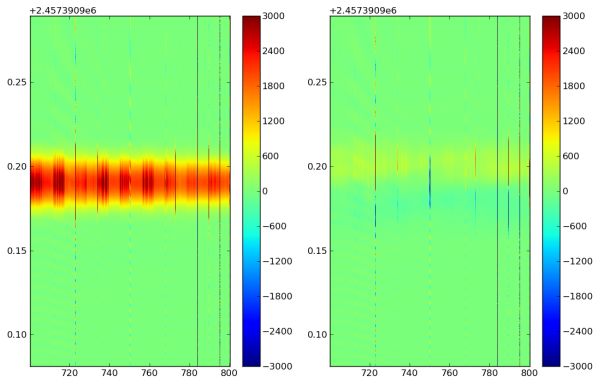
$$V_{ij}^{\text{cal}} = e^{2\pi i \hat{s}_0 \cdot \mathbf{u}_{ij}} \int d^2 \boldsymbol{\sigma} A_i(\boldsymbol{\sigma}) A_j^*(\boldsymbol{\sigma}) I(\boldsymbol{\sigma}) e^{2\pi i \boldsymbol{\sigma} \cdot \mathbf{u}_{ij}}. \quad (8)$$

Phase the calibrated visibility to the phase reference center by divide V_{ij}^{cal} by $e^{2\pi i \hat{s}_0 \cdot \mathbf{u}_{ij}}$,

$$V_{ij} = \int d^2 \boldsymbol{\sigma} A_i(\boldsymbol{\sigma}) A_j^*(\boldsymbol{\sigma}) I(\boldsymbol{\sigma}) e^{2\pi i \boldsymbol{\sigma} \cdot \mathbf{u}_{ij}}. \quad (9)$$

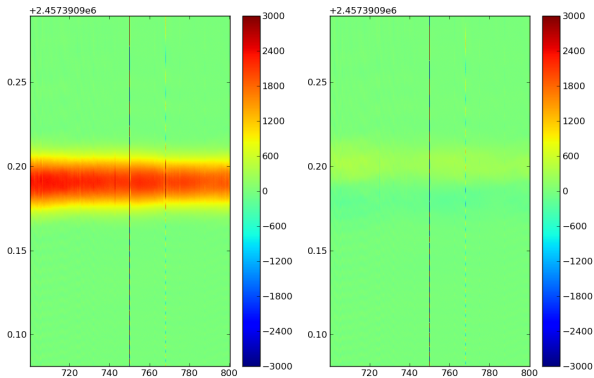
Visibility After Phased to Cas A

XX correlation



Visibility After Phased to Cas A

YY correlation



Imaging

In the flat-sky approximation, and ignore the w -term,

$$V_{ij} = \iint A_i(l, m) A_j^*(l, m) I(l, m) e^{-2\pi i(ul+vm)} dl dm, \quad (10)$$

