

Requirements of running at the Z-pole and WW-threshold

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- **Status physics and Collider options**
- **Impact of precision measurements at Z and WW**
- **Calibration Issues**
- **Technical remarks on the machine**
- **Conclusions**

Why running at low energies (Z, WW) at ILC?

Clear distinction needed:

A. For calibration purposes

- Parameter group studies this question: dedicated questions to detector groups
- Status: preliminary answer from ILD (not yet official), not yet from SID
- A few remarks on this later, but not topic of this talk!

B. For physics aspects

- Status of physics case now
- For which cases may running at WW be important?
- For which cases may Z-pole runs be important?

Status physics

- **Current physics case for e⁺e⁻ (as well known):**

- Higgs precision physics
 - Top precision physics
 - Light electroweak particles/DM searches
 - BSM detection in general, complementary to LHC
- } **guaranteed physics!**
- } **well motivated**

- **HEP-e⁺e⁻: ILC, CLIC or CepC or Fcc-ee**

- ILC, $\sqrt{s}=90-500$ (1000) GeV, 163 MW, 31 km, polarized beams, $t \geq 2032$

- CLIC: $\sqrt{s}=500$ GeV, 1.5, 3 TeV, 560 MW, 48 km, polarized beams, $t \geq 2035$ (?)

- CepC: $\sqrt{s}=240$ GeV, 500 MW, 54 km, polarization unclear, $t \geq 2028$

- FCC-ee: $\sqrt{s}=350$ GeV, 500 MW, 100km, no polarization at 350, $t \geq 2035$?

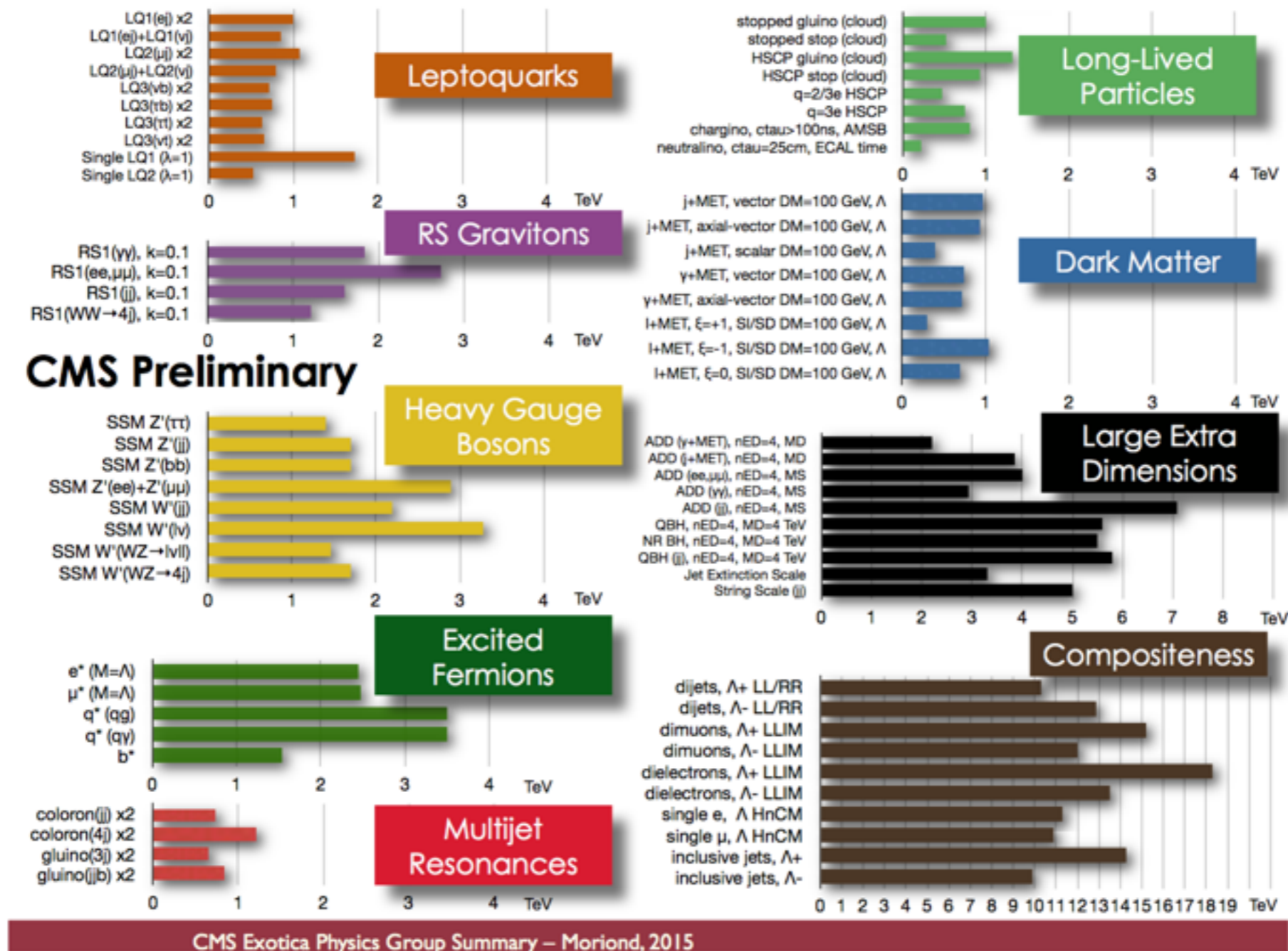
P. Burrows
W. Chou,
EPS15

- *Many ideas, but realistically only one mature project —the ILC—that could start operation before LHC switches off (~2035)*

➤ *very substantial for future development of the field as a whole!*

Current experimental result from LHC (CMS)

- Many model beyond the Standard Model on the market: no clear sign!



- High mass limits: any alternative to get sensitivity to new physics?
- *exploit electroweak precision observables*

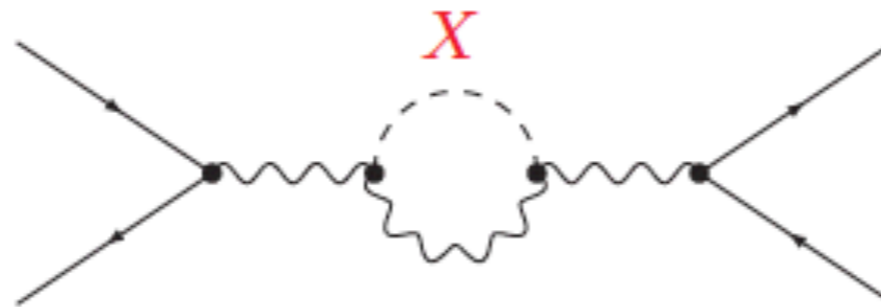
Electroweak precision observables: MW, sin2θeff, MH, (g-2)μ, b physics

- **Comparison of observables with theory**

**Precision data:
MW, sin2θeff, aμ, Mh**

**Theory:
SM, SUSY, Z',...**

Test of theory at quantum level: sensitivity to heavy masses via loops!



SM: limits on MH,

BSM: limits on MX

Very high accuracy of measurements as well as theoretical predictions needed

- only models 'ready so far: SM, SM-like, SUSY
.....still room for improvement from theory
- in the following —**as case study only**— SUSY

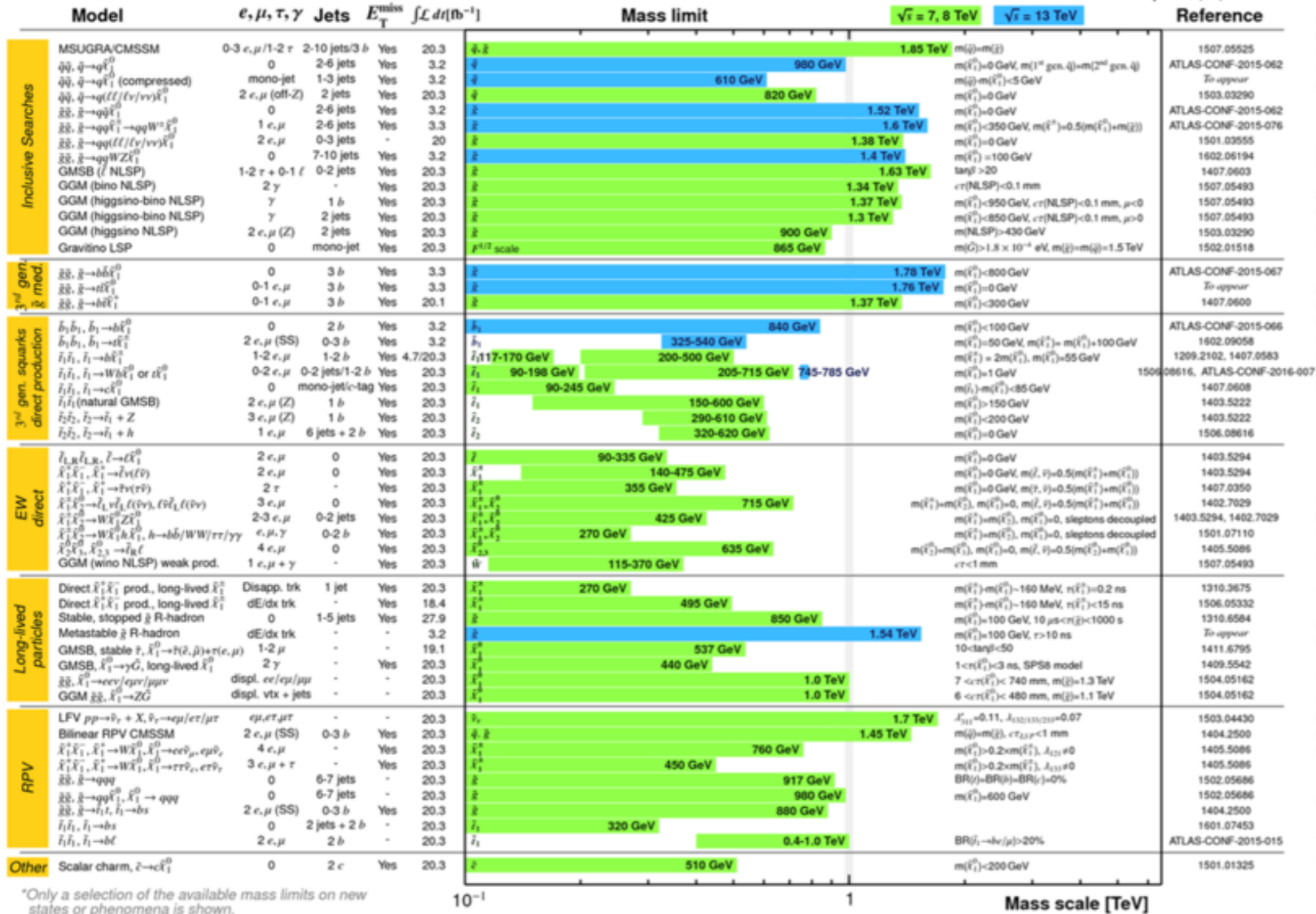
Current experimental result from LHC (ATLAS)

ATLAS SUSY Searches* - 95% CL Lower Limits

Status: March 2016

ATLAS Preliminary

$\sqrt{s} = 7, 8, 13$ TeV



All simplified models!

➤ So far: rather heavy SUSY mass limitsbut not exclusively

Importance of M_W and mixing angle $\sin^2\theta_{eff}$

A. Theoretical prediction for M_W in terms of M_Z , α , G_μ , Δr :

$$M_W^2 \left(1 - \frac{M_W^2}{M_Z^2}\right) = \frac{\pi \alpha}{\sqrt{2} G_\mu} (1 + \Delta r)$$

loop corrections

Evaluate Δr from μ -decay $\Rightarrow M_W$

One-loop result for M_W in the SM:

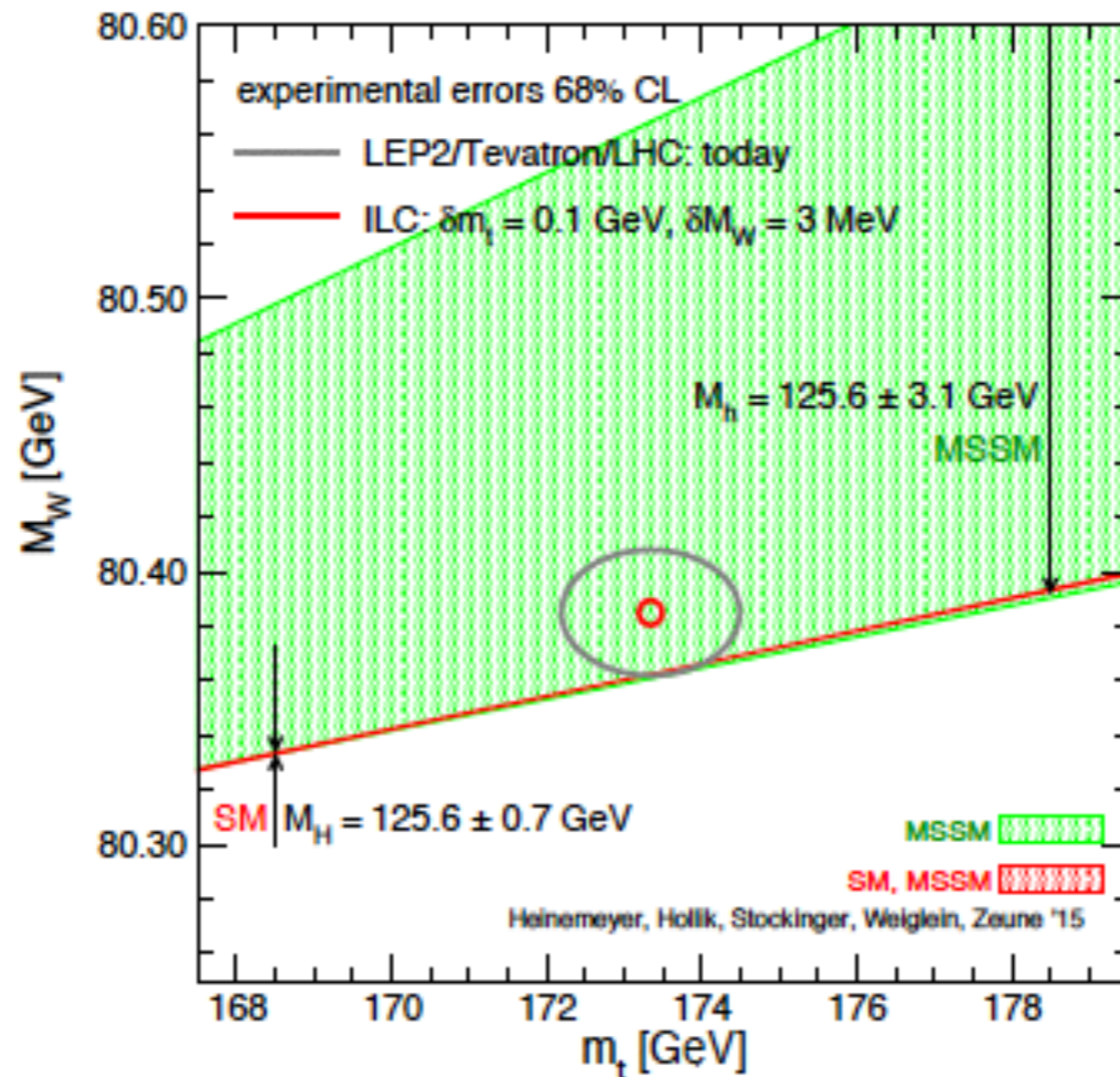
$$\Delta r_{1-loop} = \Delta\alpha - \frac{c_W^2}{s_W^2} \Delta\rho + \Delta r_{rem}(M_H)$$

$\sim \log \frac{M_Z}{m_f}$	$\sim m_t^2$	$\log \frac{M_H}{M_W}$
$\sim 6\%$	$\sim 3.3\%$	$\sim 1\%$

➤ Large impact of m_{top} , M_H and M_Z and their uncertainties!

Which precision in M_W do we need?

- Case study SUSY versus Standard Model
 - Strong correlation of M_W , M_t and M_H :



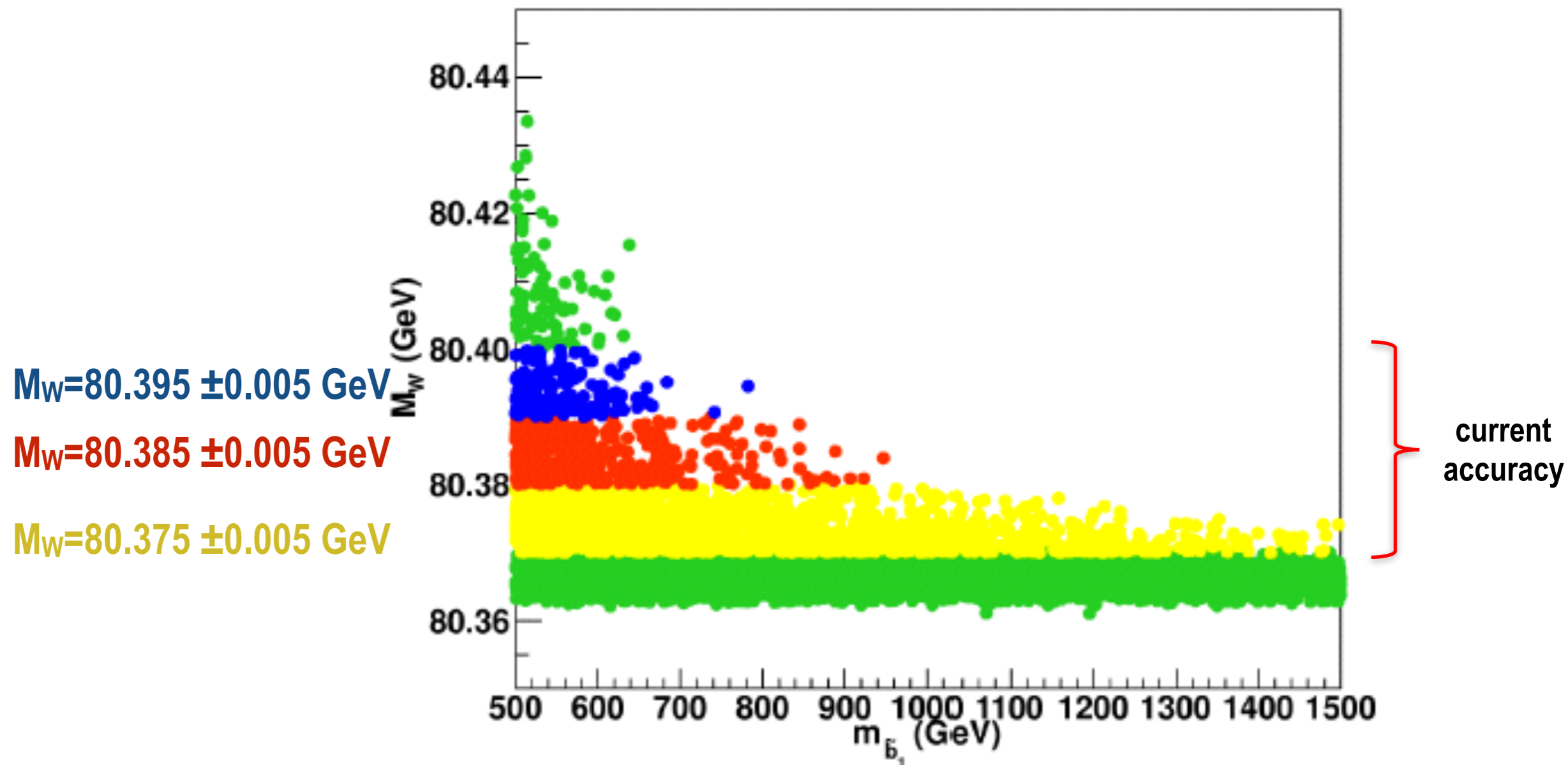
MSSM band: scan

SM and SM-like MSSM

Current experimental world average: $M_W=80.385 \pm 0.015$ GeV, $M_t=173.34 \pm 0.76$ GeV

Which precision in M_W do we need?

- Just as case study SUSY: $m_{\text{stop}1}=400\pm 40$ GeV, other susy masses >500 GeV



➤ *precision for ΔM_W below/about 5 MeV required in order to be sensitive to heavy susy masses!*

The mass of the W boson

- Experimental accuracy:

Today: LEP2, Tevatron: $M_{W\text{exp}}=80.385\pm 0.015$ GeV

ILC: - polarized threshold scan

- kinematic reconstruction of WW

- hadronic mass (single W)

G. Wilson 2013

$\Delta M_{W\text{exp,ILC(FCC-ee)}} < 3$ (1) MeV (from threshold scans) (but without Theo uncertainties!)

- Theoretical accuracies:

intrinsic today: $\Delta M_{W\text{SM,theo}} = 4$ MeV \Rightarrow future: $\Delta M_{W\text{SM,theo}} = 1$ MeV

parametric uncertainties today:

from $\Delta m_t = 0.9$ GeV $\Rightarrow \Delta M_{W\text{mt}} = 5.5$ MeV

from $\Delta \alpha_{\text{had}} = 10^{-4}$ $\Rightarrow \Delta M_{W\alpha_{\text{had}}} = 2$ MeV

from $\Delta m_Z = 2.1$ MeV $\Rightarrow \Delta M_{Wm_Z} = 2.5$ MeV

future:

$\Delta m_t = 0.05$ GeV $\Rightarrow \Delta M_{W\text{mt}} = 0.5$ MeV

$\Delta \alpha_{\text{had}} = 5 \times 10^{-5}$ $\Rightarrow \Delta M_{W\alpha_{\text{had}}} = 1$ MeV

$\Delta m_Z = 1/0.1$ MeV $\Rightarrow \Delta M_{Wm_Z} = 0.2/0.02$ MeV

Impact of the precision of the mixing angle $\sin^2\theta_{\text{eff}}$

B. Effective mixing angle:

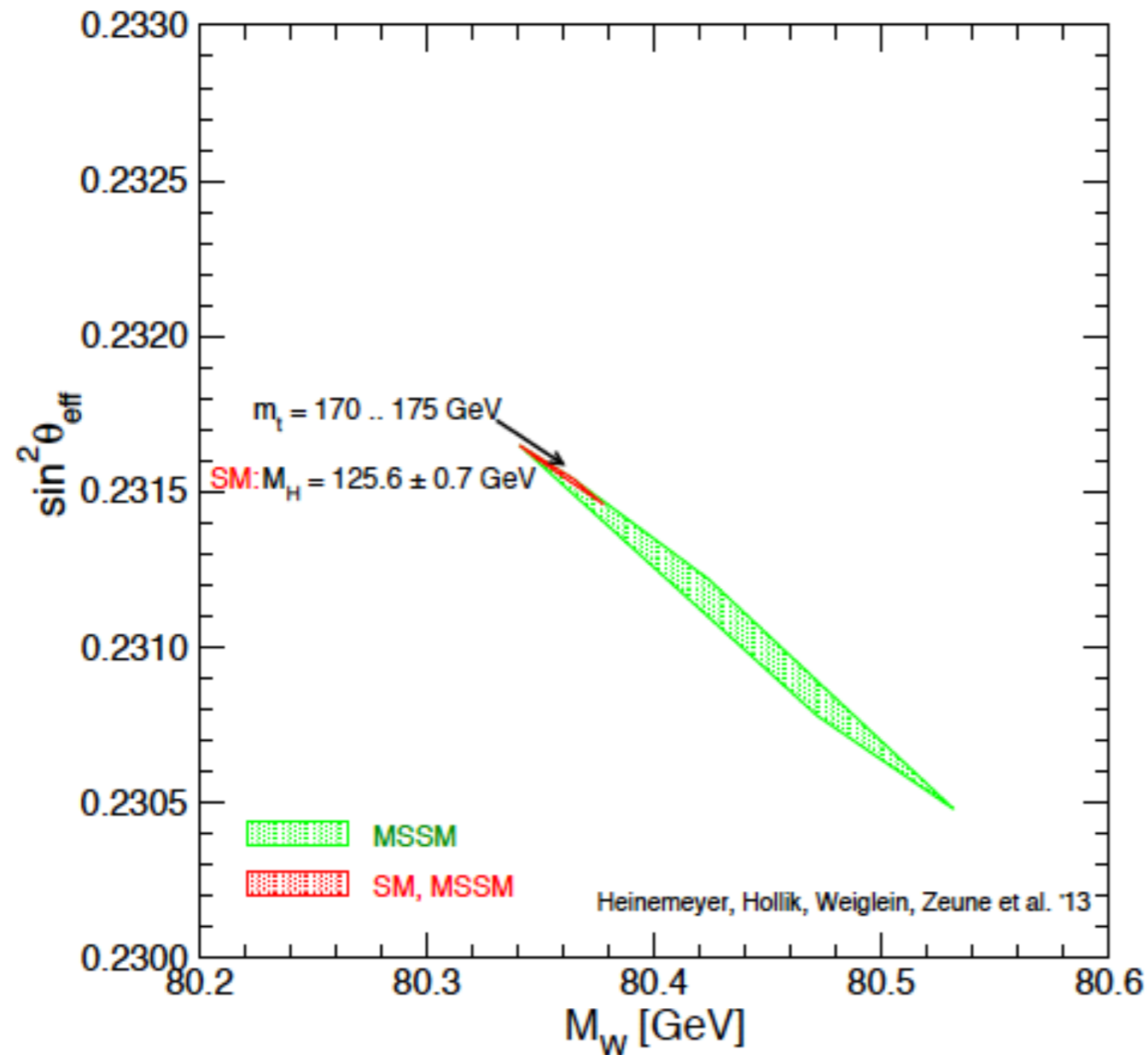
$$\sin^2\theta_{\text{eff}} = \frac{1}{4|Q_f|} \left(1 - \frac{\text{Re } g_V^f}{\text{Re } g_A^f} \right)$$

Higher order contributions:

$$g_V^f \rightarrow g_V^f + \Delta g_V^f, \quad g_A^f \rightarrow g_A^f + \Delta g_A^f$$

also impact of new physics contributions!

Impact of the precision of the mixing angle $\sin^2\theta_{\text{eff}}$

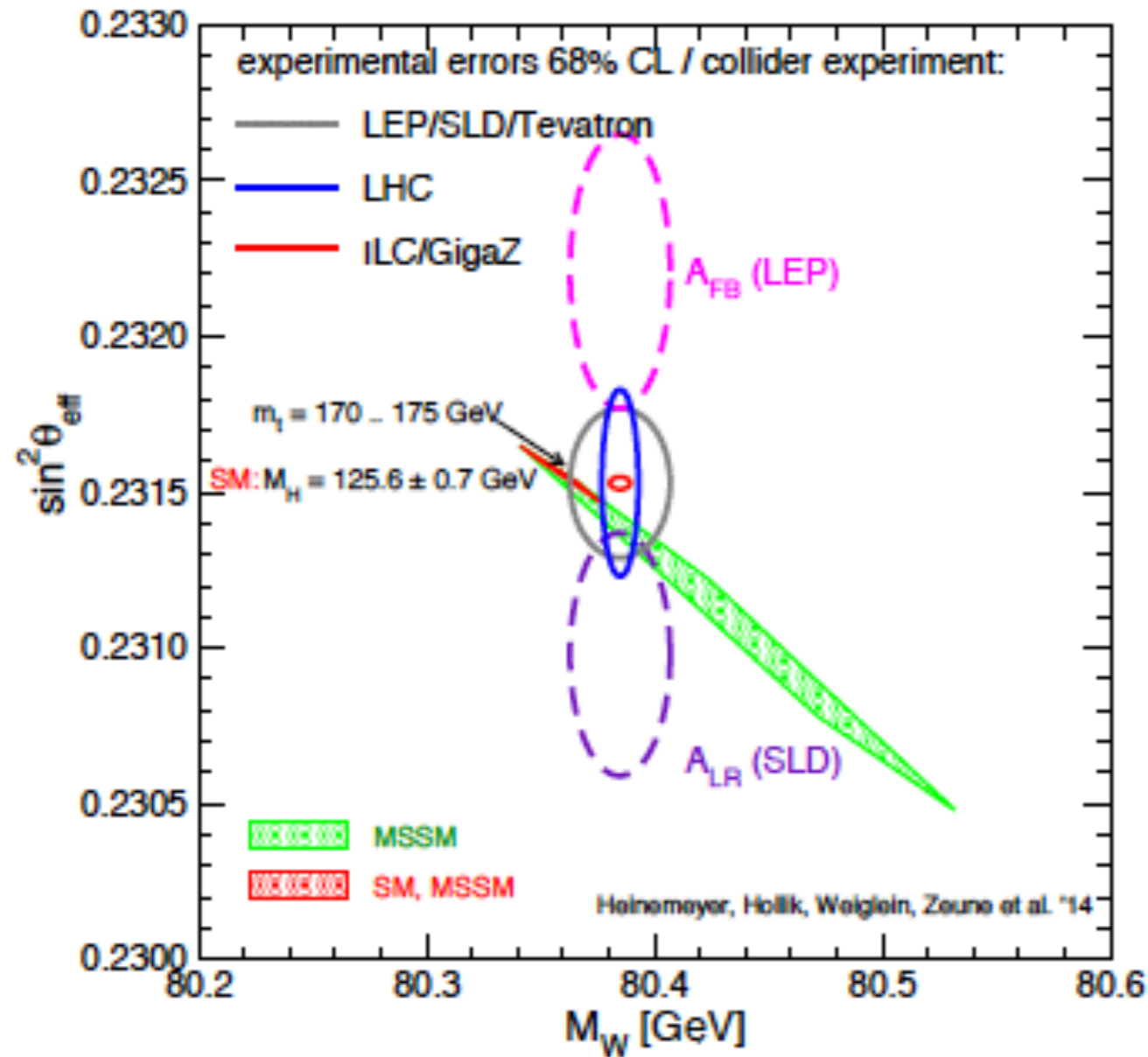


MSSM band: scan over SUSY masses

Overlap:
SM is SUSY-like
SUSY is SM-like

SM band: variation of M_H and M_t

Impact of the precision of the mixing angle $\sin^2\theta_{eff}$



MSSM band:
scan over
SUSY masses

overlap:
SM is MSSM-like
MSSM is SM-like

SM band:
variation of M_H^{SM}

Impact of the precision of the mixing angle $\sin^2\theta_{\text{eff}}$

- Experimental accuracy:**

Today: LEP, SLD: $\sin^2\theta_{\text{eff}}=0.23153 \pm 0.00016$

GigaZ/TeraZ: both beams polarized, Blondel scheme

$\Delta\sin^2\theta_{\text{eff}}^{\text{exp,ILC(FCC-ee)}}=13/3 \times 10^{-6}$ (but without Theo uncertainties!)

- Theoretical accuracies: [10-6]**

intrinsic today: $\Delta\sin^2\theta_{\text{W}}^{\text{SM,theo}}=47 \Rightarrow$ future: $\Delta M_{\text{W}}^{\text{SM,theo}}=15$

parametric uncertainties today:

from $\Delta m_{\text{t}}=0.9\text{GeV} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\text{mt}}=30$

from $\Delta\alpha_{\text{had}}=10^{-4} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\alpha_{\text{had}}}=36$

from $\Delta m_{\text{Z}}=2.1\text{MeV} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\text{mZ}}=14$

future:

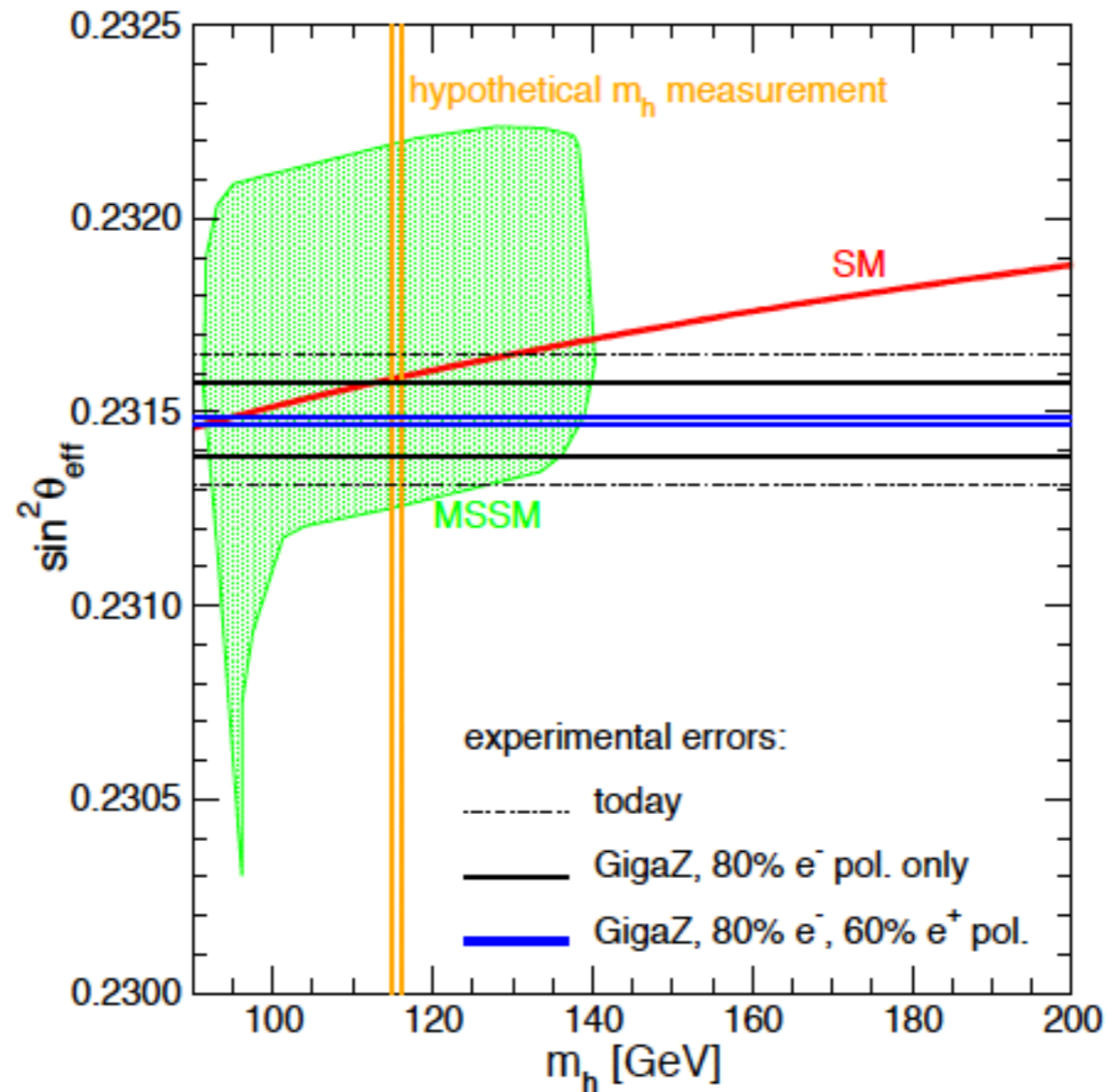
$\Delta m_{\text{t}}=0.05\text{ GeV} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\text{mt}}=2$

$\Delta\alpha_{\text{had}}=5 \times 10^{-5} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\alpha_{\text{had}}}=18$

$\Delta m_{\text{Z}}=1/0.1\text{MeV} \Rightarrow \Delta\sin^2\theta_{\text{eff}}^{\text{mZ}}=6.5/0.7$

Impact of the precision of the mixing angle $\sin^2\theta_{\text{eff}}$

- Last but not least: relevance of both beams polarized:



Heinemeyer, Weiglein 2005

- precision on $\sin^2\theta_{\text{eff}}$ relies strongly on both beams polarized
- crucial to reach sensitivity!

Remark on running at Z/WW for calibration

- **Calibration needs with Z pole data**

- referred to in the TDR, but without specific design
- promised precision can only be guaranteed if matched by alignment and calibration
- not clear whether switching between low and high energy running is critical
- clear question to the detector groups: how often and which lumi is needed and which detector component does really need Z pole running?

*G. Wilson
Talk at DESY
Hamburg, 5/16*

\sqrt{s}	$\sigma(\mu\mu)$ (pb)	$\sigma(q\bar{q})$ (pb)	$\rho_Z(\mu\mu)$	$\rho_Z(q\bar{q})$
91.2	1580	30500	1.0	1.0
250	4.99	50.1	316	609
350	2.57	24.8	614	1230
500	1.30	12.6	1210	2420
1000	0.386	3.64	4080	8370

- Assuming γ scaling of lumi: hadronic events are 440 times higher at Z-pole compared to 500 GeV.
- Calibration that may need about a year at 500 can be done in 1 day at Z-pole,...if well mature designed

Remark on running at Z/WW for calibration

Preliminary intermediate results (thanks to Jenny!)

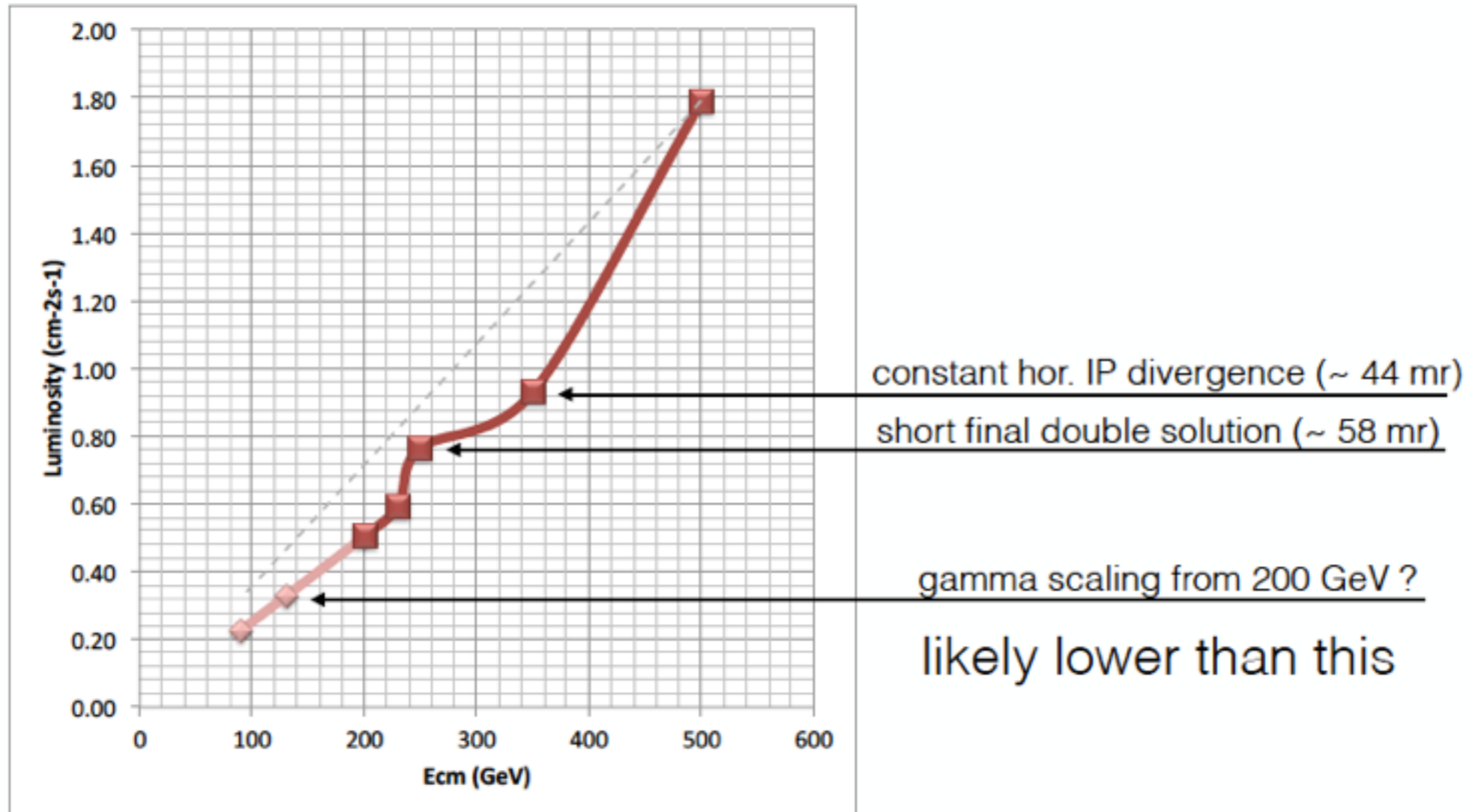
- So far, ILD tracking systems expects frequent Z calibration runs with 10^{32} or 1 pb⁻¹ in few hours
- Calorimetry: absolute scale of ECAL — no Z pole running needed
 - MIP scale calibration requires couple of days Z pole calibration
 - HCAL: MIP scale most efficiently via Z pole
- Jet Energy scale: no dedicated running at Z or WW

However: still many details under work

- no lumi-issues for positron source
- can be delivered with low intensity source

Machine issues: options to run at low energies

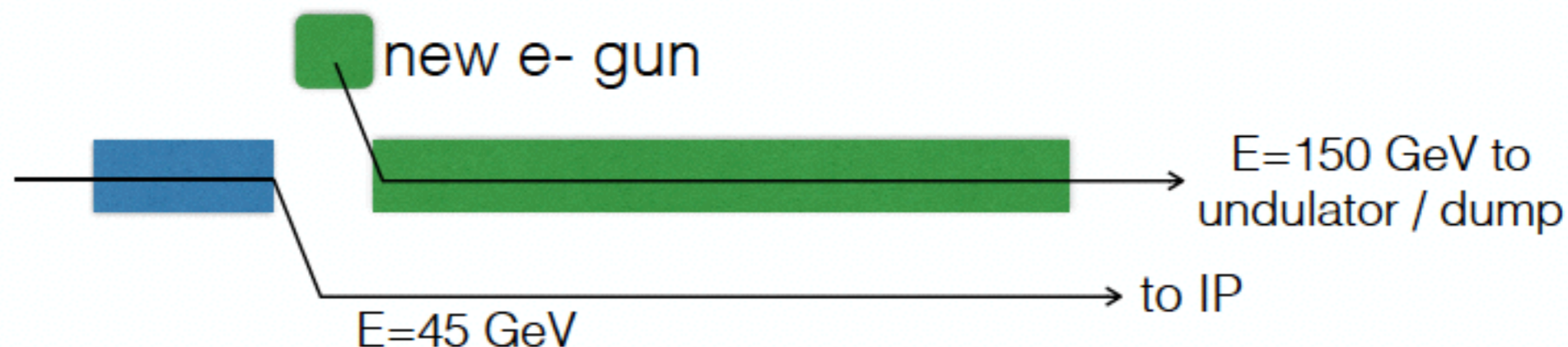
- Luminosity issues: γ scaling may not be appropriate



- also 10 Hz scheme not always applicable if ΔE between pulse too big, e.g. 45 GeV and 150 GeV.....problems with linac dynamics

Machine issues: options to run at low energies

- Further option: split linac !



Major reconfiguration of accelerator

Requires a mini design study

- 3rd beamline in linac tunnel
- additional doglegs, bypasses and possible dumps
- ...

- Z and W running conceptually possible at $2-4\text{\AA}\sim 10^{33}$ and $3-6\text{\AA}\sim 10^{33}$ respectively but much more detailed studies required !

Conclusions

- **Calibration Issues:**

- still under work, some detector parts require Z-pole (10^{32}) (time issues!)
- crucial information how quickly and how often energy changes are required: stay tuned, under work

- **Physics:**

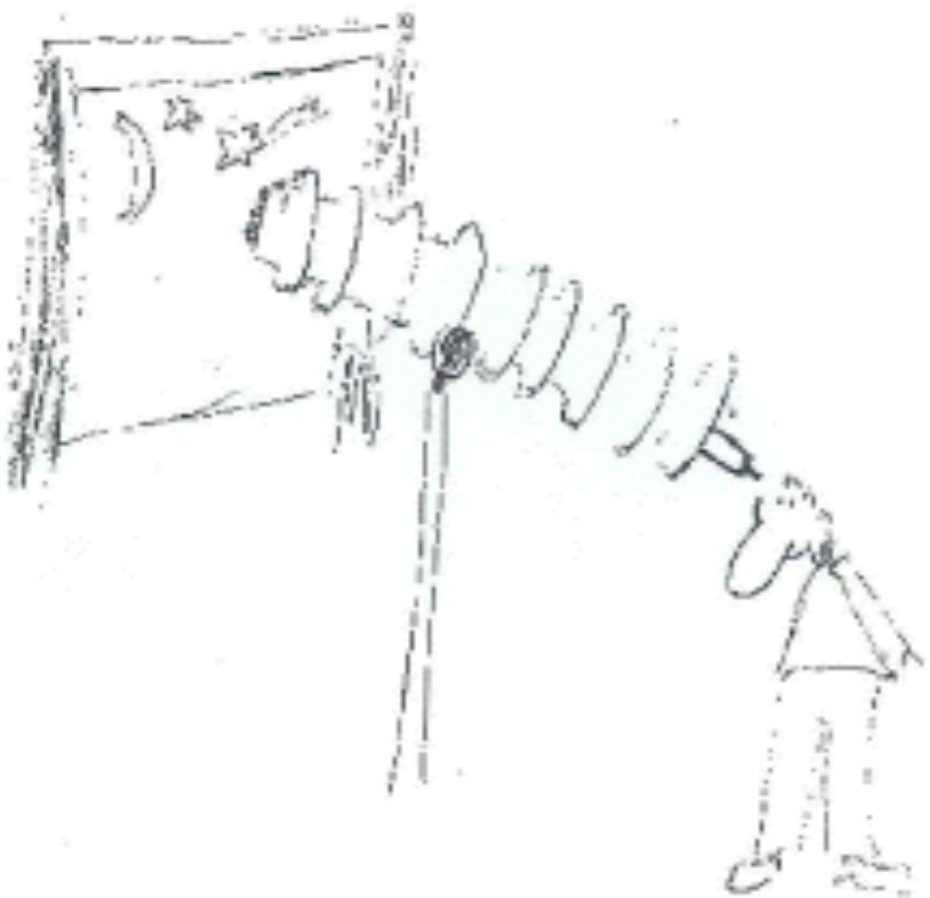
- **Threshold runs at WW needed:** mass precision
- **GigaZ run needed (including polarized beams!):** electroweak precision angle
- FCC-ee option: can not replace ILC@Z,WW

- **Accelerator issues:**

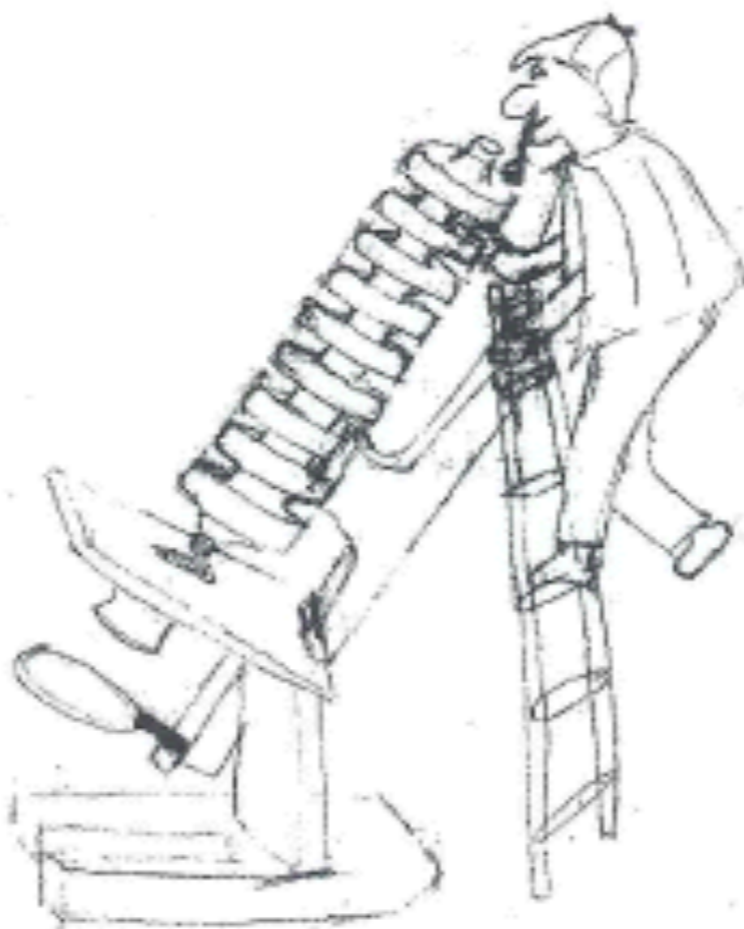
- in principle no lumi issues but design changes required
- Technical solution via split linac
- For lumi upgrade still some more work to do (but also for e-driven source unclear)

ILC only compatible and complementary to LHC if physics runs with 10^{33} at low energy not excluded !

I Love Cavities



Unraveling the past



Revealing present mysteries



Feeding future brains

THE I L C