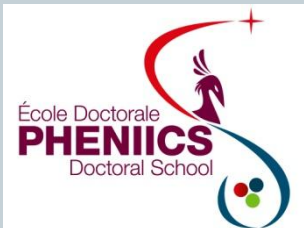


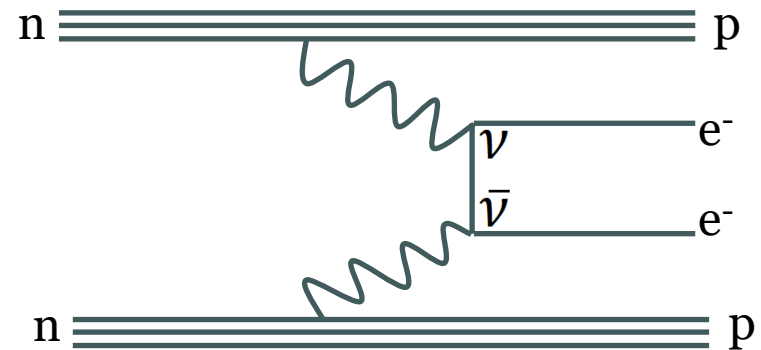
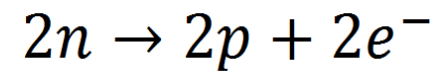
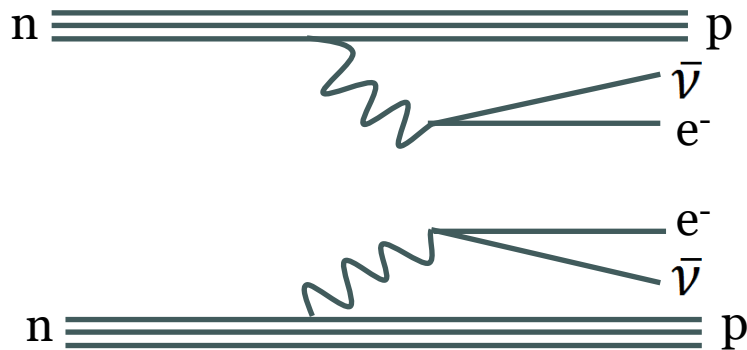
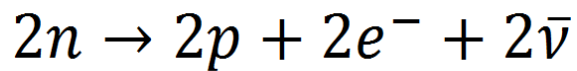
An innovative Neganov-Luke-assisted light detector for the sensitivity enhancement of CUORE experiment



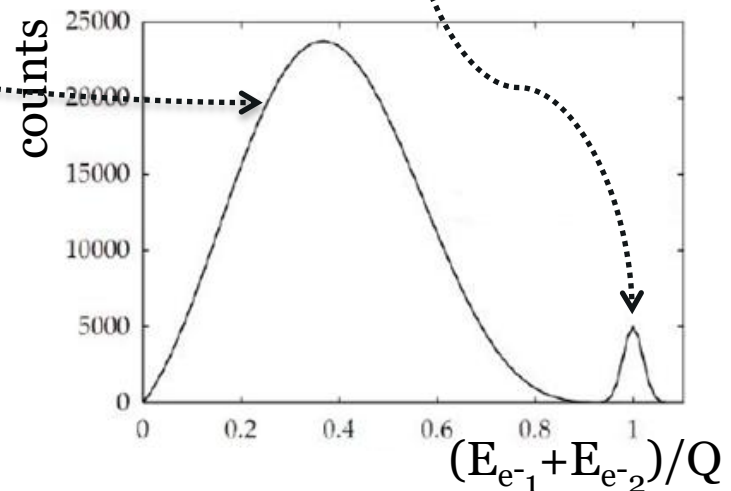
Valentina Novati



What is neutrinoless double beta decay?



Neutrinoless double beta decay is a hypothetical rare nuclear process: its mean life is longer than 10^{26} years, that corresponds to 10^{16} times the age of the Universe



And why do we search for it?

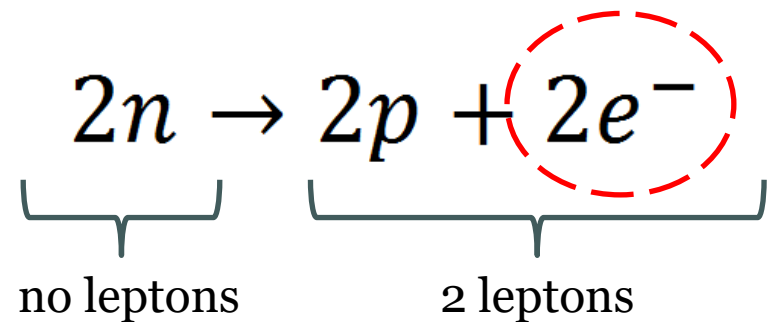
- Majorana nature of neutrino

$$\nu = \bar{\nu}$$

- Neutrino mass

$$T_{1/2} \propto \frac{1}{m_\nu^2}$$

- Lepton number violation

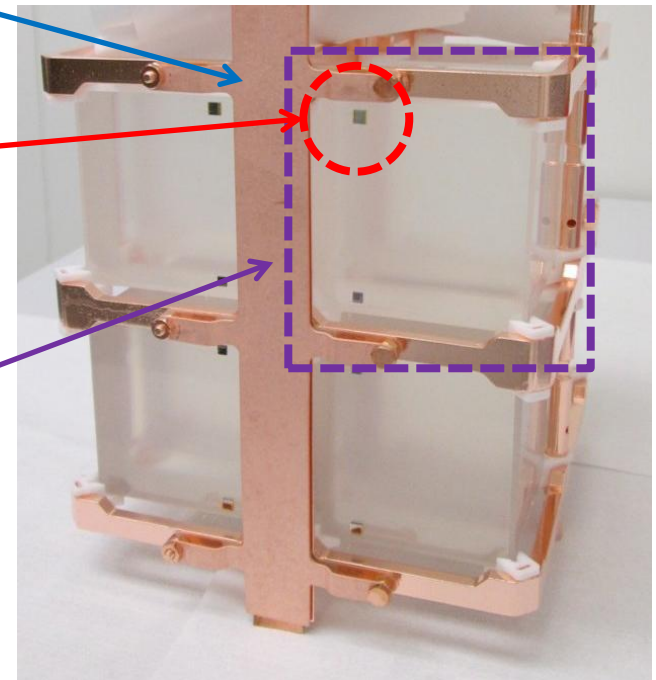
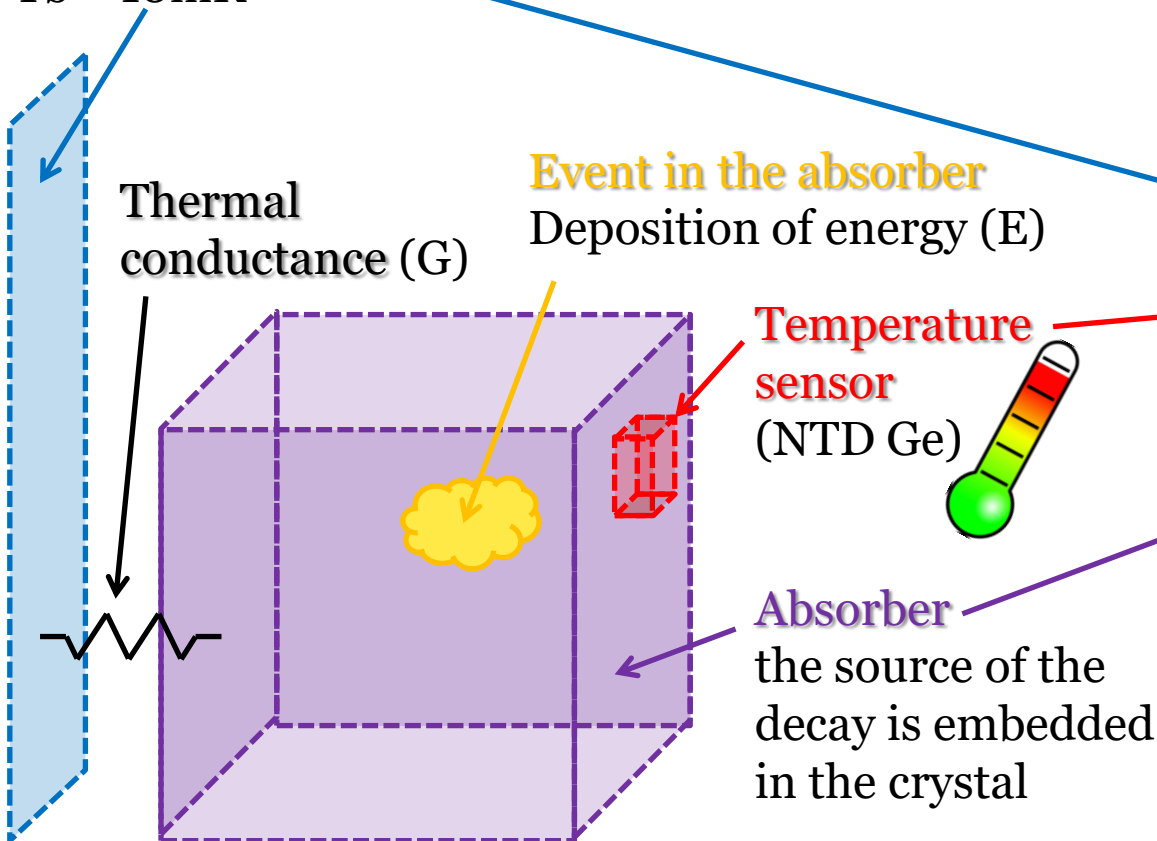


How do we search for it? Bolometers!!

Bolometers work at cryogenic temperatures to be sensitive to temperature variation of the order of μK induced by particles

Thermal bath

$T_b \sim 10\text{mK}$



CUORE experiment

CUORE uses 988 TeO_2 bolometers, where ^{130}Te is the candidate isotope to the neutrinoless double beta decay (with 2527 keV energy)

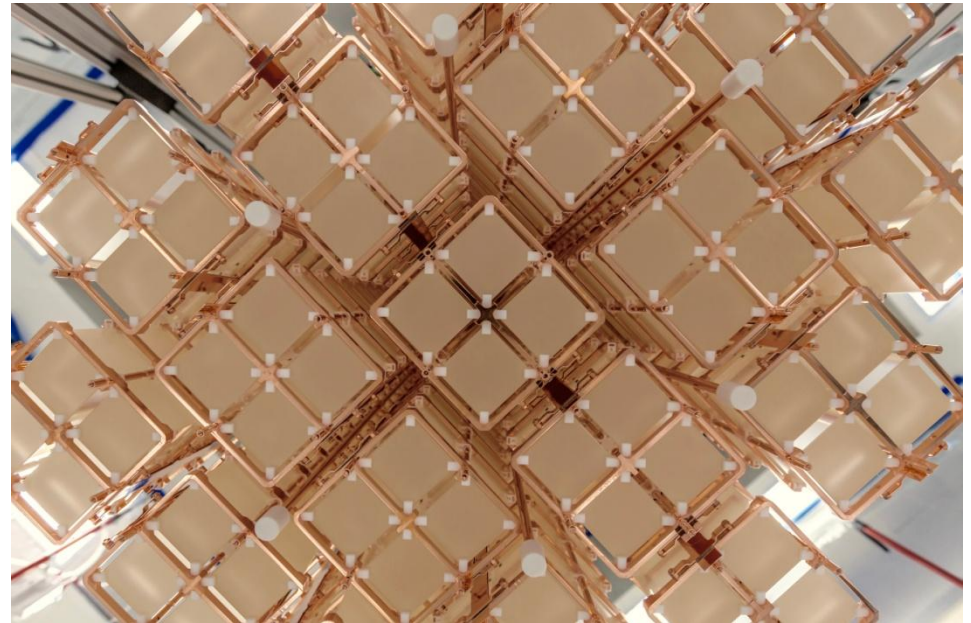
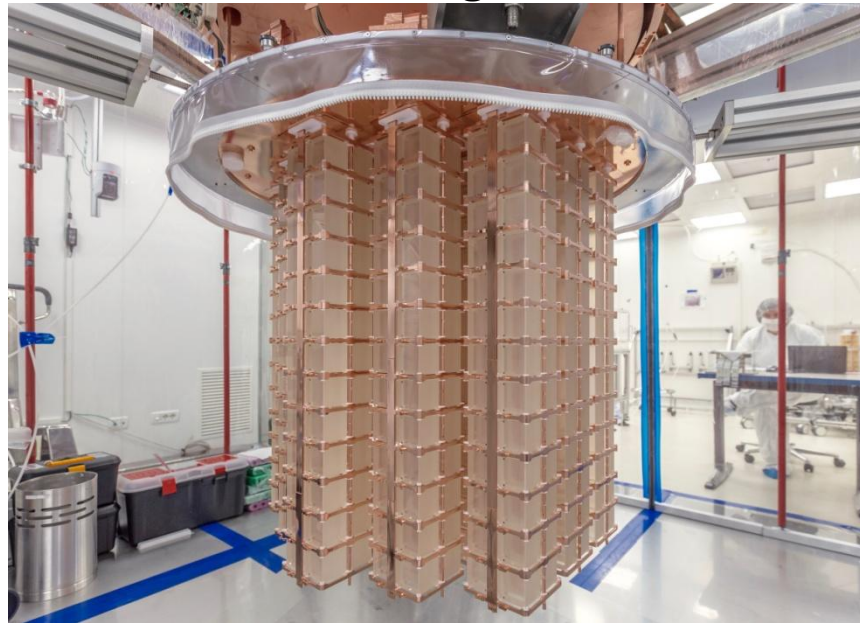
Radiopurity control to minimize the background

Gran Sasso Underground laboratory (LNGS, IT)



Neutrinoless double beta decay is really **rare**

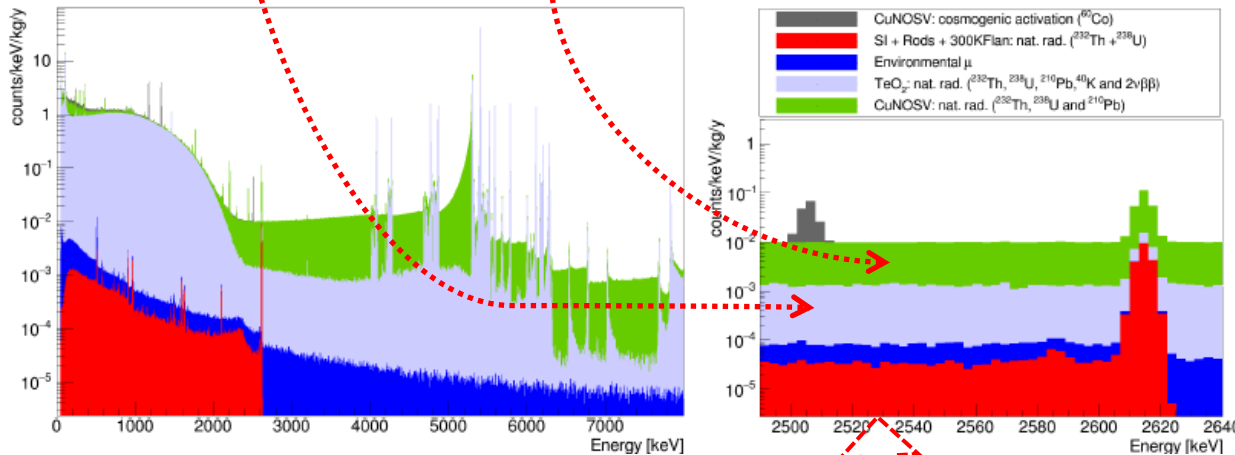
Huge mass:
206 kg of ^{130}Te
and 742 kg of TeO_2



The issue...

CUORE is not a zero-background experiment, the main background is due to **α particles** with degraded energy emitted from the surfaces near the detector.

Background budget of the CUORE experiment



arXiv:1704.08970

32 α events are expected in the neutrinoless-double-beta-decay-energy region in 1 year.

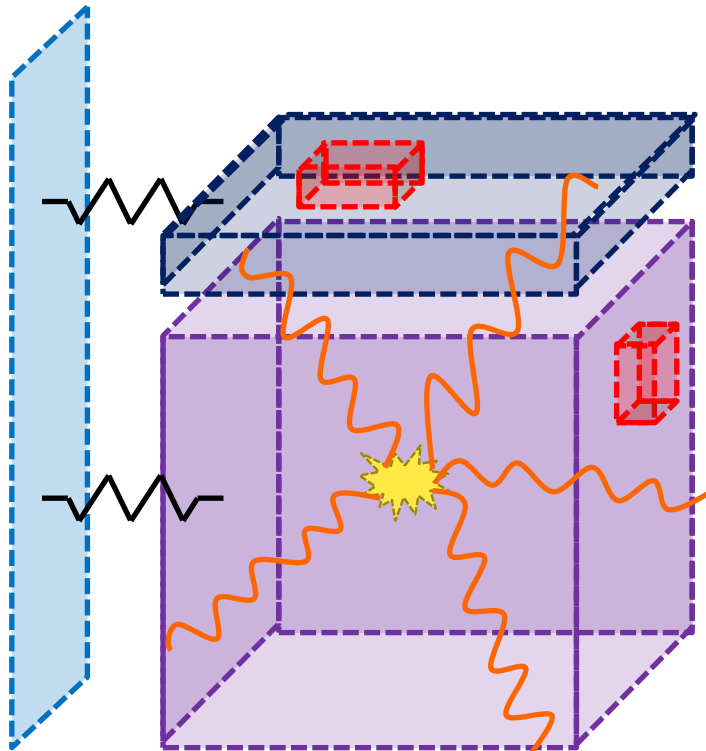
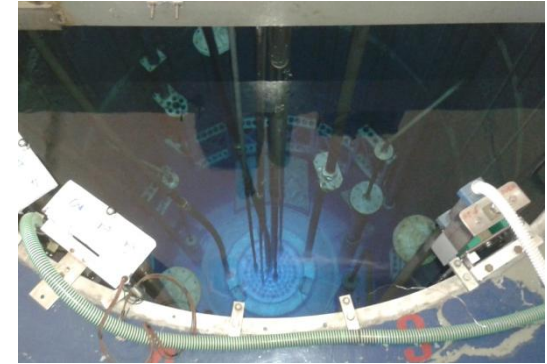
Signal

Our solution: the Cherenkov light tagging

The TeO_2 Cherenkov emission threshold:

- 50 keV for the electrons;
- 400 MeV for α s.

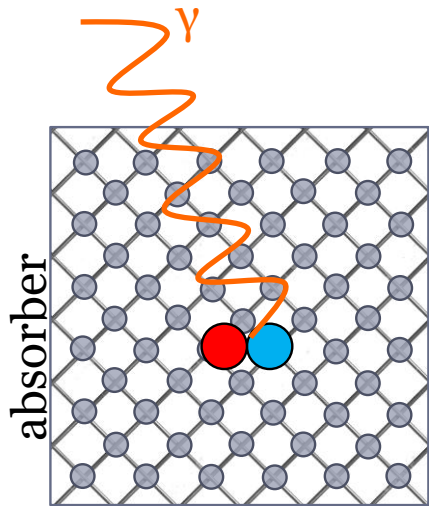
Electrons, produced by neutrinoless-double-beta decay in ^{130}Te , have an energy higher (2527 keV) than the threshold, while surface α s do not.



If the detector is coupled to a **bolometric light detector**, it is possible to measure the Cherenkov light produced by electrons and distinguish their signal from the α background.

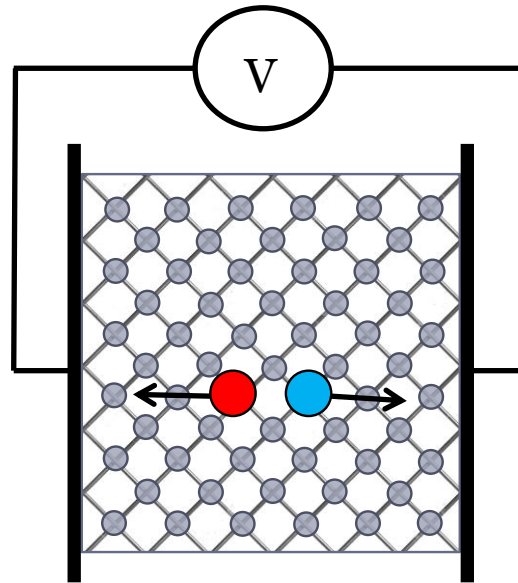
Cherenkov light collected by a light detector corresponds to an energy of around **100 eV**, that would be hidden in the baseline noise of a standard light detector (**~ 50 -100 eV**).

Neganov-Luke effect

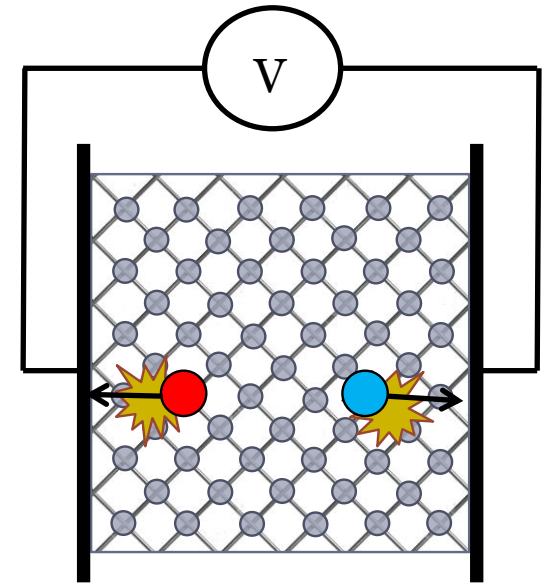


absorber

The absorption of photons produces electron-hole pairs



The electric field *drifts the charges and it prevents their recombination*



Carriers *collide with the lattice during the drift, increasing the temperature*

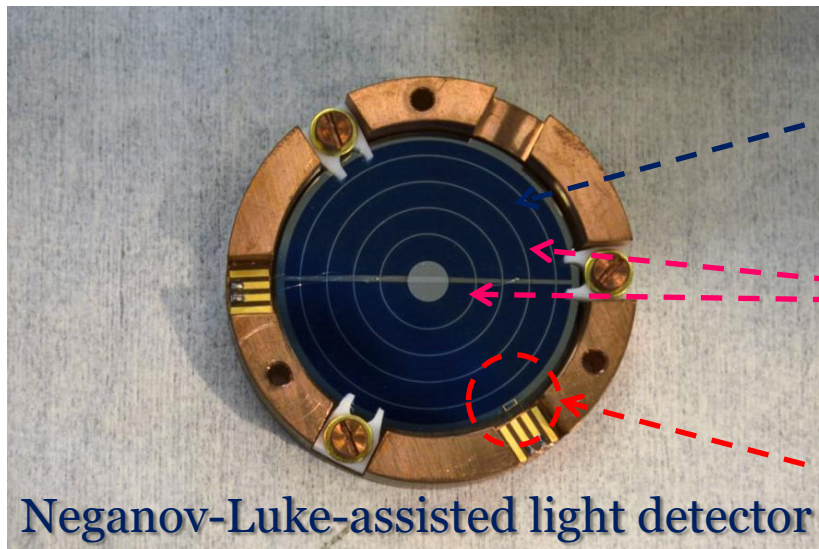
$$E = E_0 \cdot \left(1 + \frac{q \cdot V}{\varepsilon} \right)$$

q electron charge

V voltage applied

ε energy used to produce an electron-hole pair

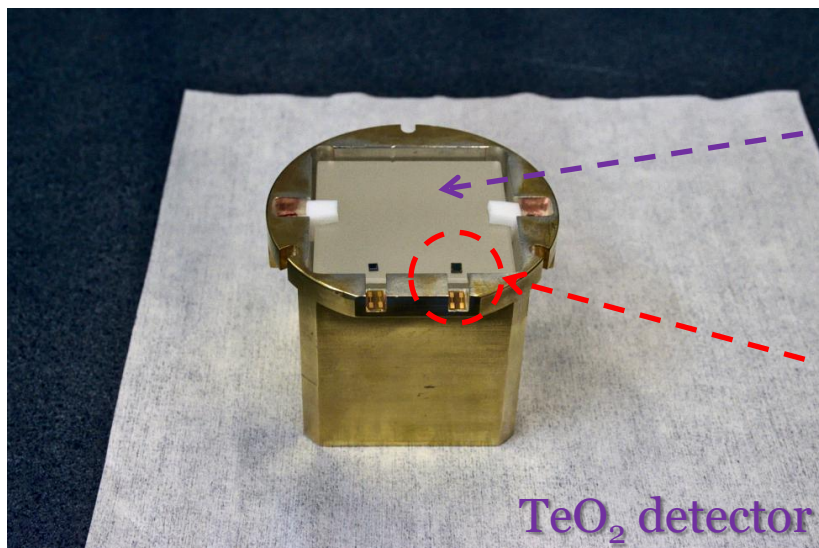
Our hero: Neganov-Luke light detector



Ge wafer
 $\text{\O} = 44 \text{ mm}$, $h = 0.17 \text{ mm}$

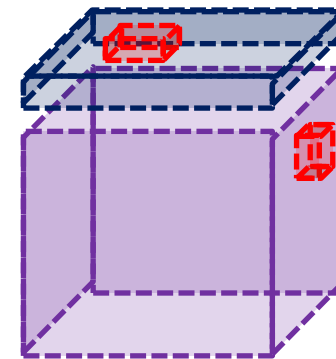
Concentric grids (made of Al)
to apply the electric field for
the Neganov-Luke effect

NTD Ge
mass $\sim 5.5 \text{ mg}$



TeO₂ crystal
volume $5 \times 5 \times 5 \text{ cm}^3$

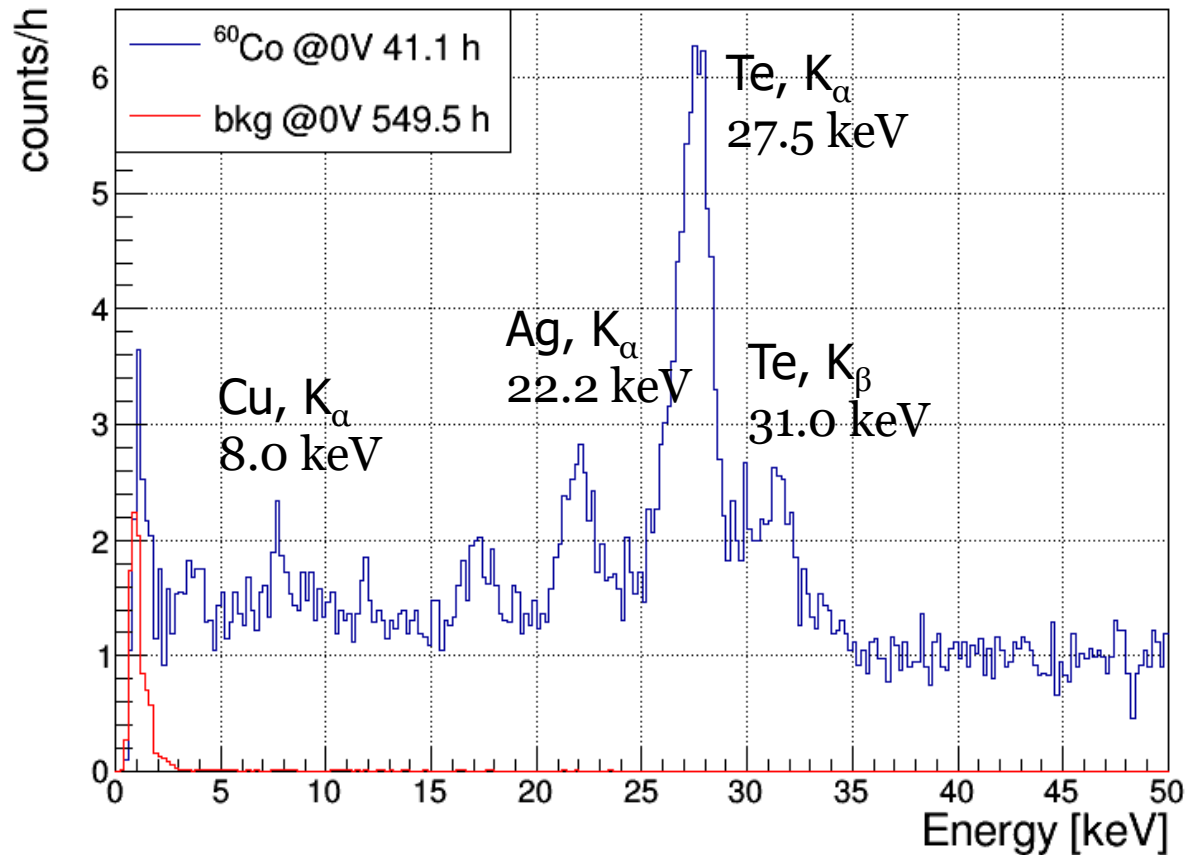
NTD Ge
mass $\sim 48 \text{ mg}$



Smearred alpha source of ^{210}Po

Light detector calibration

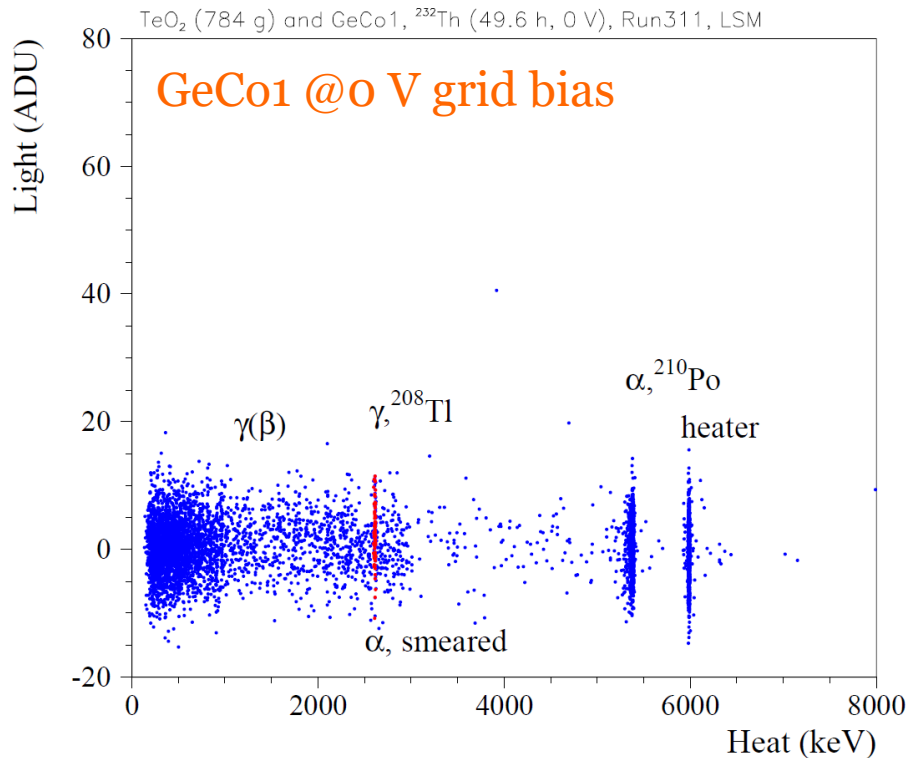
GeCo1 CSNSM, LSM RUN 311 - ra26-rd04



GeCo1 41.1 h (@17 mK) with ^{60}Co source

549.5 h (@17 mK) of **background**

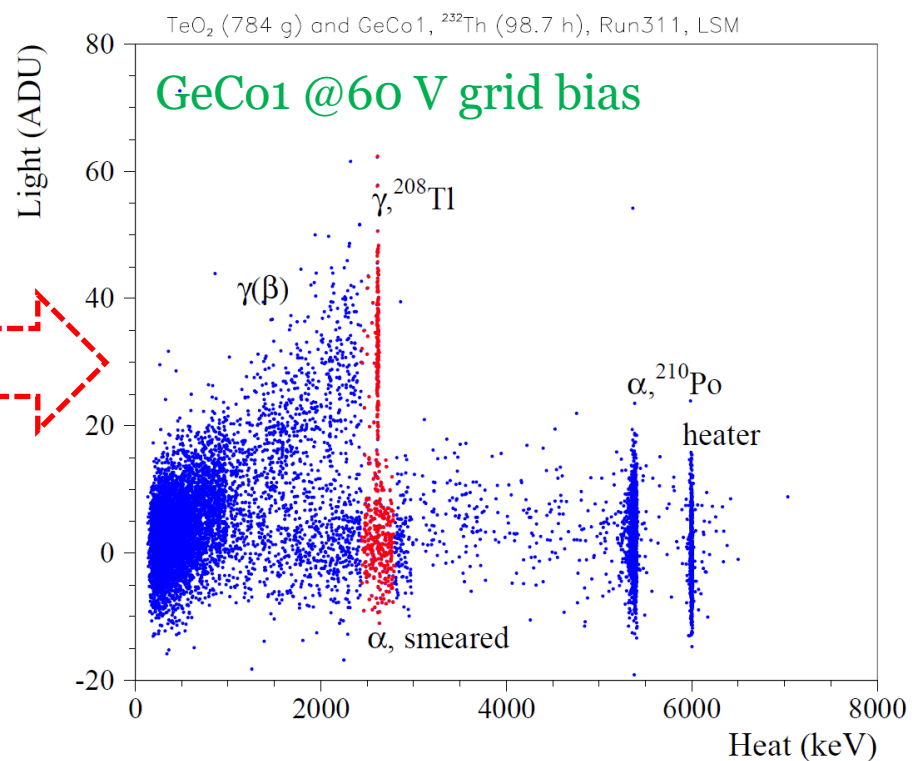
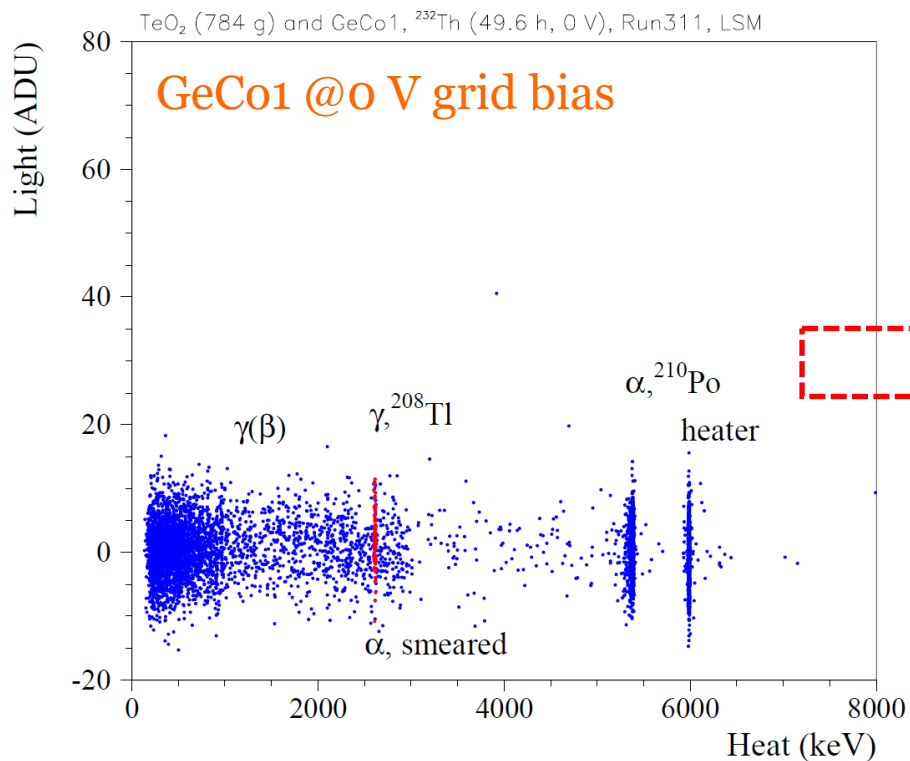
What do we see with a standard light detector



We acquired data without bias on the grids.

baseline-noise RMS = **108 eV**
signal/noise = **0.65**

Let's turn on amplification

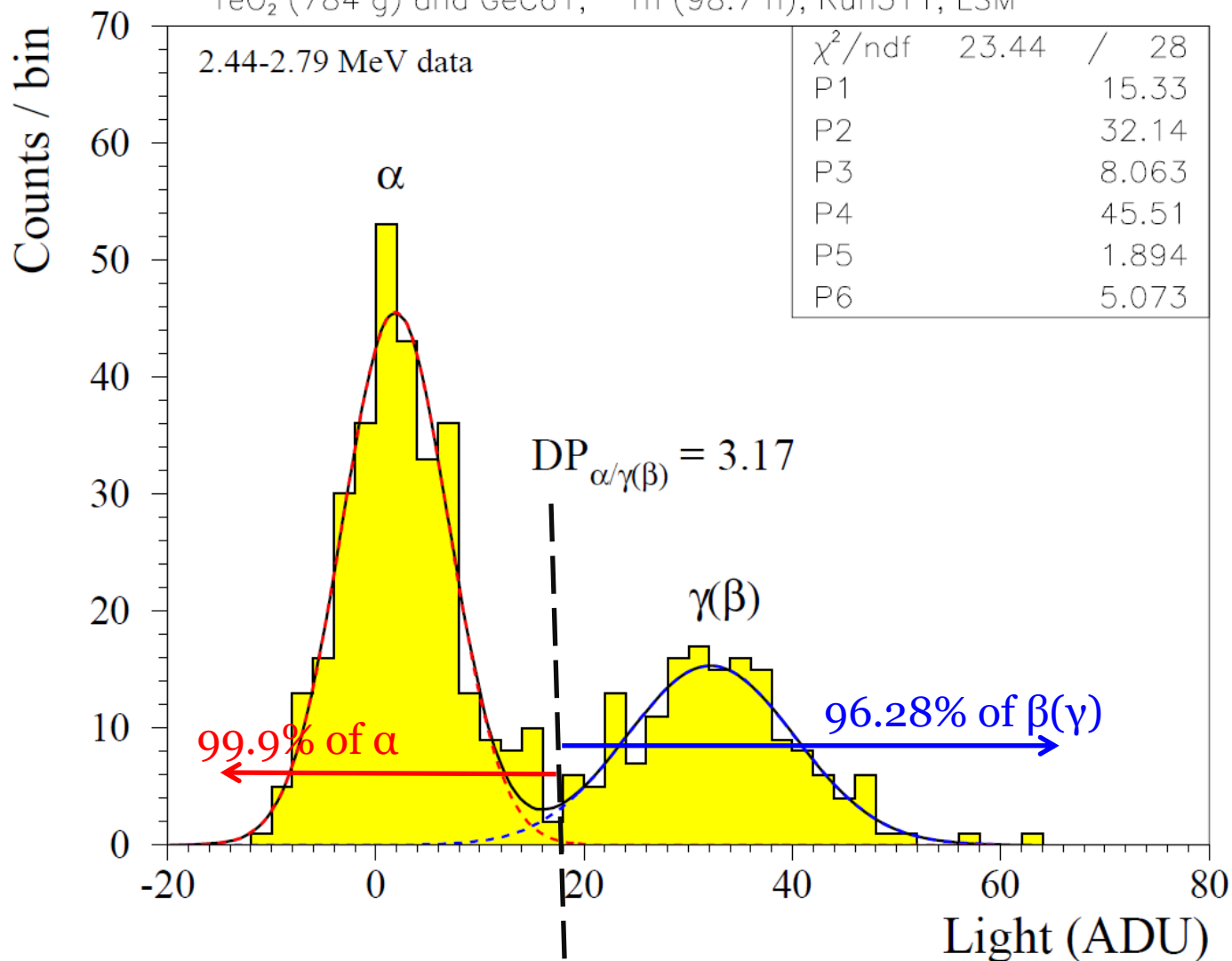


baseline-noise RMS = 108 eV
signal/noise = 0.65

baseline-noise RMS = 9.9 eV
signal/noise = 7.1

$\alpha/\beta(\gamma)$ separation

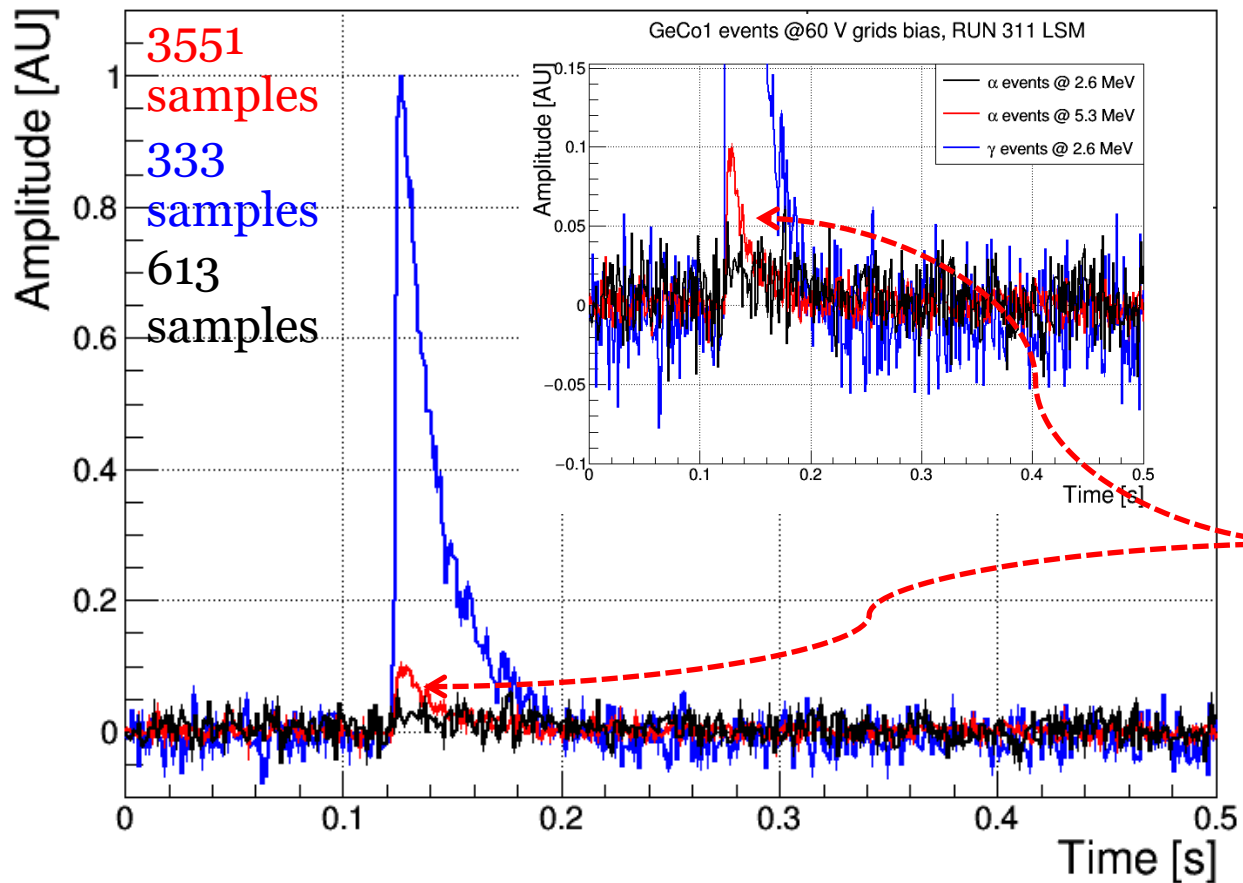
TeO₂ (784 g) and GeCo1, ²³²Th (98.7 h), Run311, LSM



We obtained an acceptance of **96.28%** of $\beta(\gamma)$ events with a rejection of **99.9%** of α particles.

Not only Cherenkov light...

GeCo1 events @60 V grids bias, RUN 311 LSM



Averaging many events to reduce the baseline noise...

Scintillation light of α s emitted by ^{210}Po

Conclusions

- We have been capable of distinguish the β signal from the α background: 96.28% of β events is accepted, rejecting 99.9% of α s.
- The light emitted by TeO_2 crystal is both Cherenkov and scintillation light.

