

# DARK ENERGY: PRESENT AND FUTURE OBSERVATIONS



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# The *History* of Cosmic Expansion

## Observations

The Universe is Expanding.  
CMB → Hot Big Bang! (inflation)

## Theory & Principles (corroborated by observations)

General Relativity  
Homogeneous & isotropic on large scales

## Friedman Equation: Equation of Motion for Universe

(Dark)  
Matter

$$H^2 \equiv \left( \frac{\dot{R}}{R} \right)^2 = \frac{8\pi G}{3} \rho_M + \frac{\Lambda}{3} - \frac{k}{R^2}$$

$$\Omega_M \equiv \frac{8\pi G}{3H_0^2} \rho_{M,0} \quad \Omega_\Lambda \equiv \frac{\Lambda}{3H_0^2} \quad \Omega_k \equiv -\frac{k}{R_0^2 H_0^2}$$

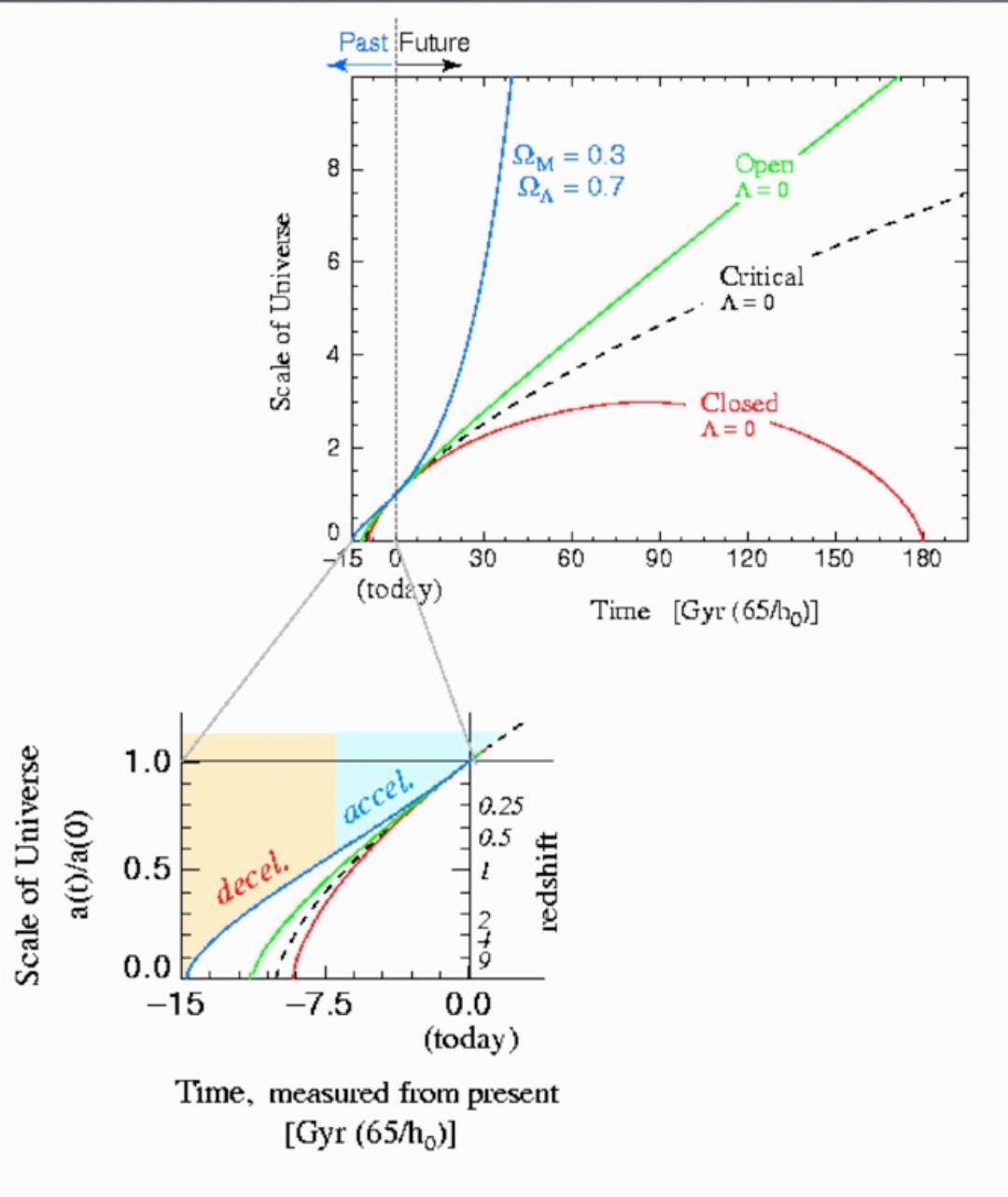
(Dark) Energy  
Einstein's blunder?

$$\Omega_M + \Omega_\Lambda + \Omega_k = -1, 0, +1$$

Geometry  
(Curvature)

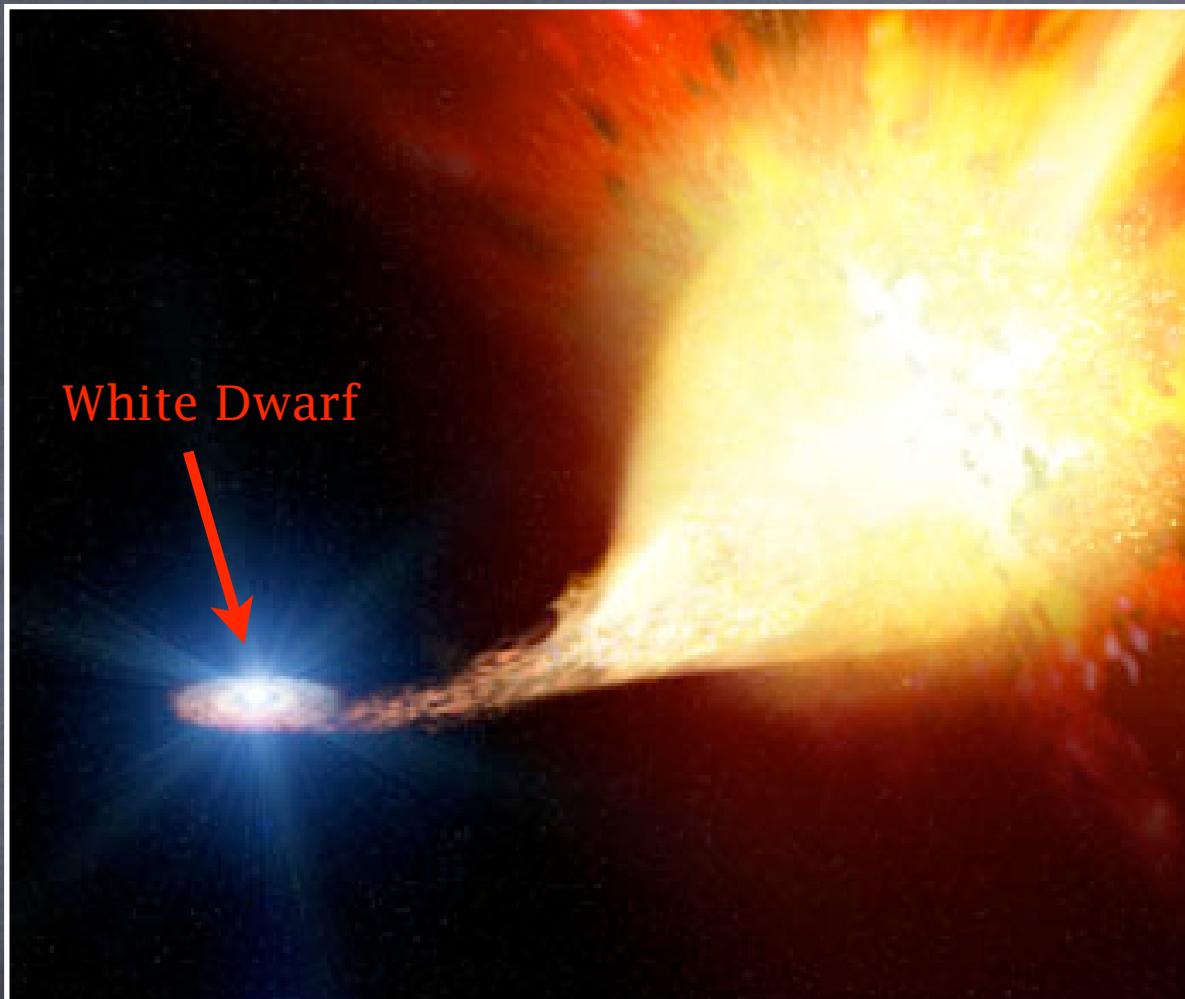
# The Expansion History of the Universe

$$D_L = cH_0^{-1}(1+z)|\Omega_k|^{-1/2}S \left\{ |\Omega_k|^{1/2} \int_0^z dz' [(1+z)^2(1+\Omega_M z) - z(2+z)\Omega_\Lambda]^{-1/2} \right\}$$

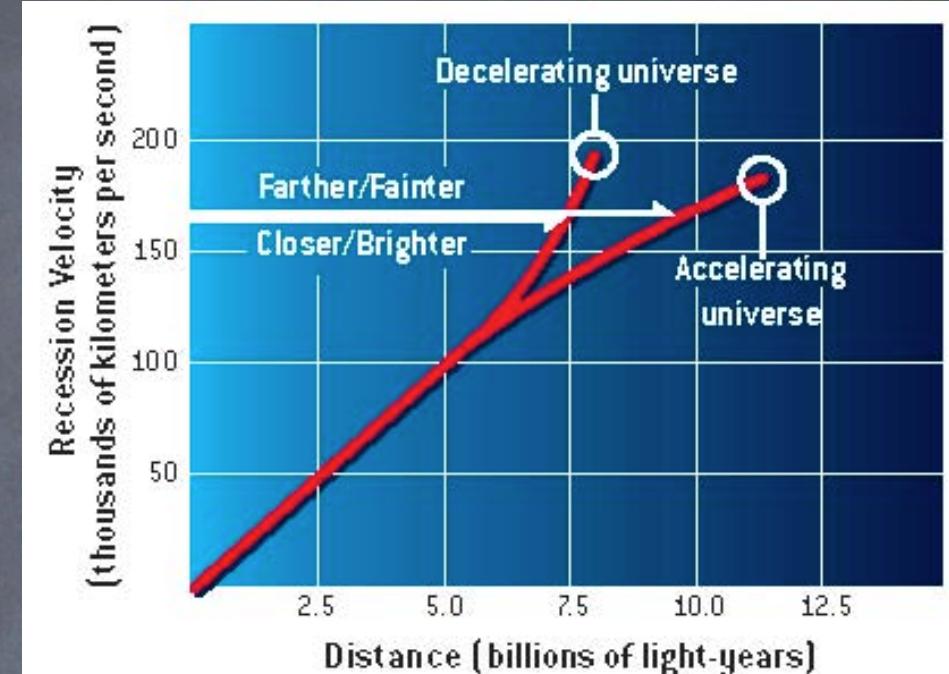
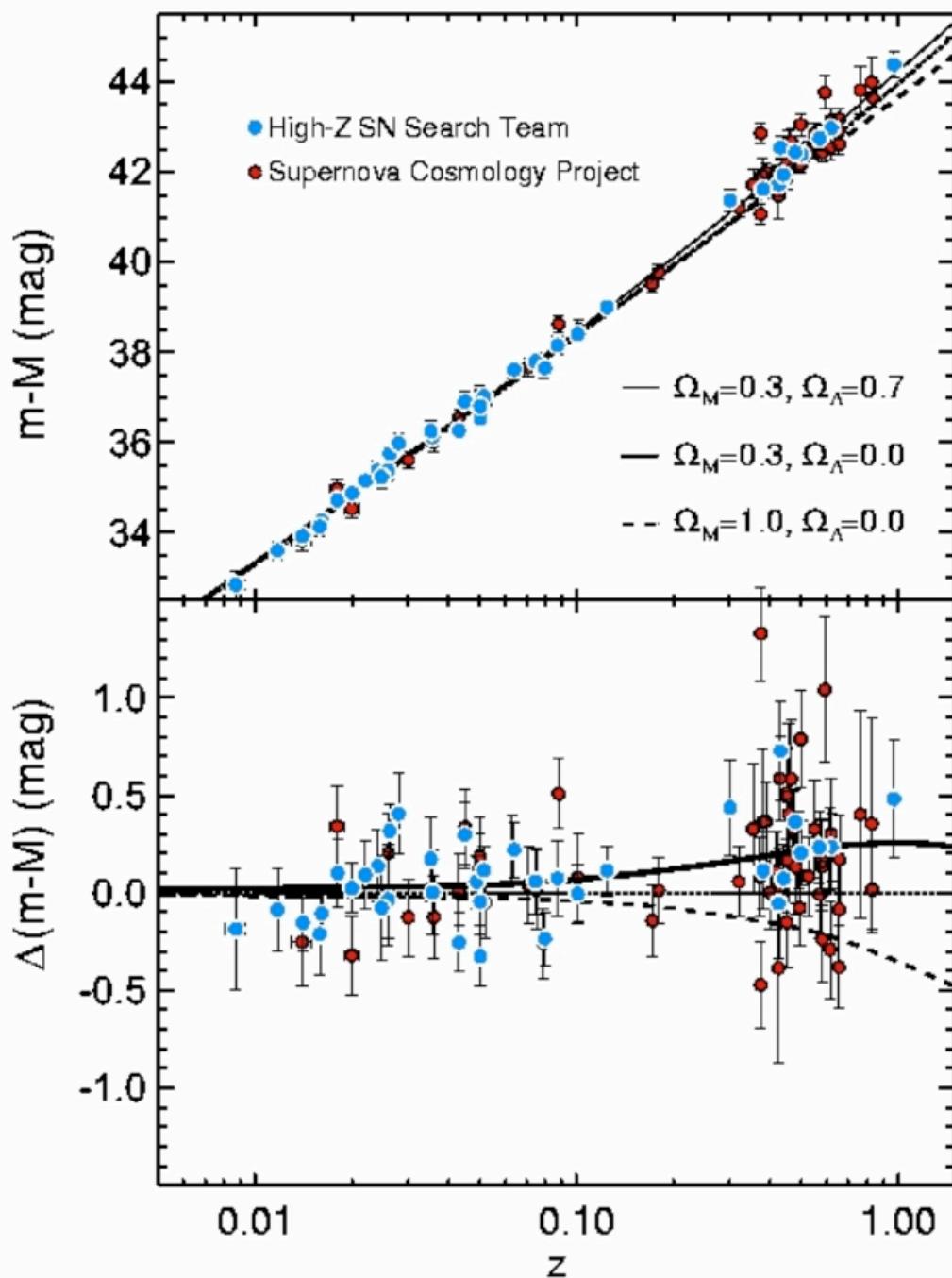


Mass is Destiny

# How (we think) Nature Makes a Type Ia Supernova



- ⦿ Homogeneity:  $1.4 M_{\odot}$ ,  $10^{51}$  ergs
- ⦿ Negligible hydrogen, lots of Intermediate Mass Elements
- ⦿ Mature progenitors
- ⦿ Models (delayed-detonation) good fit to observations

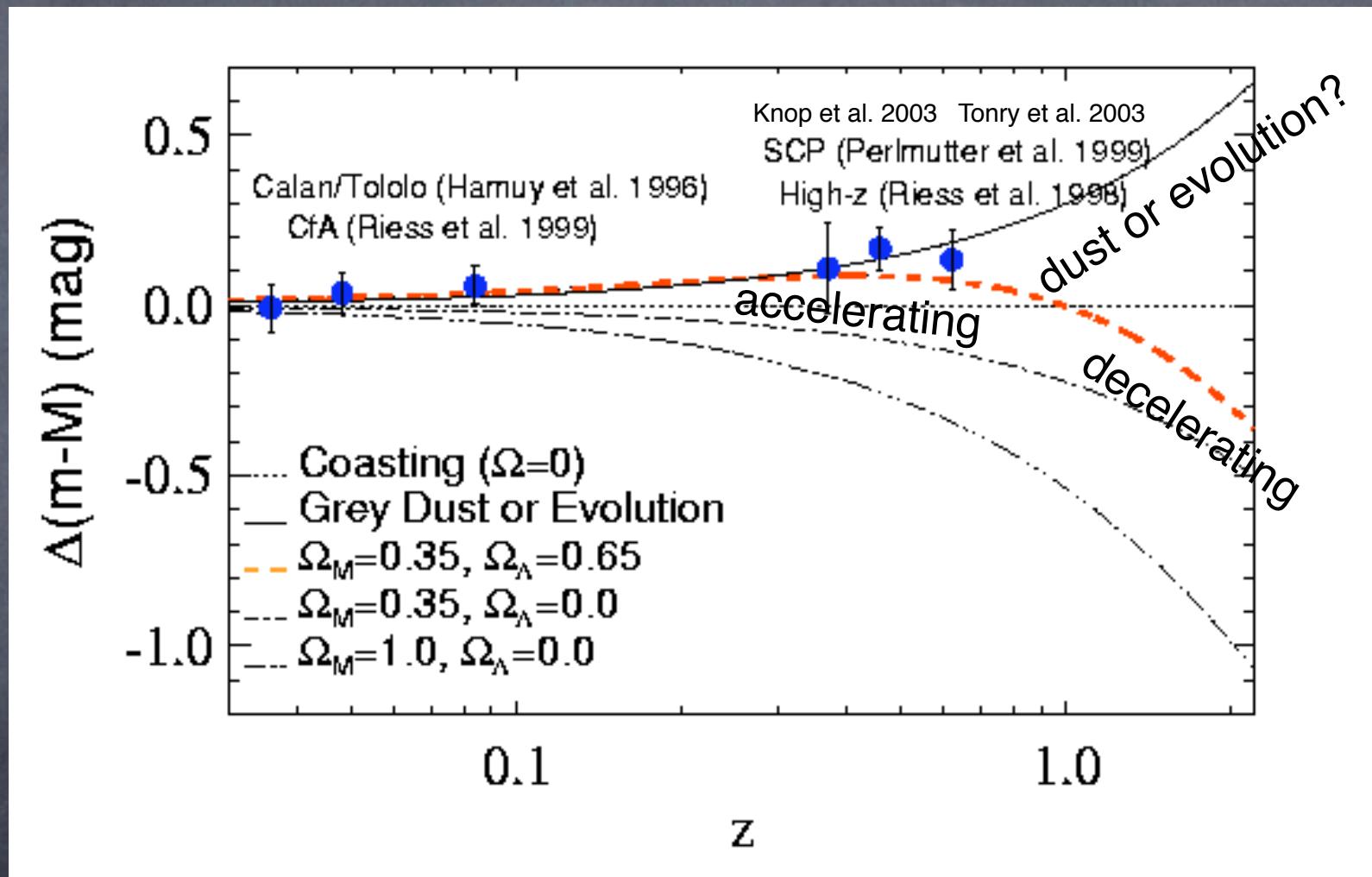
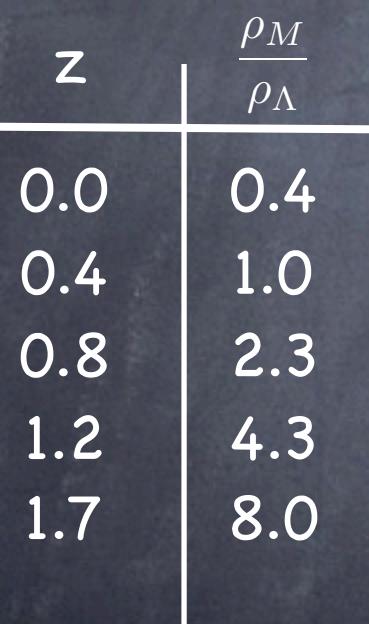


↔ 25% fainter than  
 $\Omega_M = 0.3 \quad \Omega_\Lambda = 0.0$



# Searching for the Epoch of Deceleration

$$\frac{\rho_M}{\rho_\Lambda} = \frac{\rho_{M,0}}{\rho_\Lambda} (1+z)^3$$





## The GOODS ACS Treasury Program...

Type Ia Supernova Discoveries at  $z > 1$  From the *Hubble Space Telescope*: Evidence for Past Deceleration and Constraints on Dark-Energy Evolution<sup>1</sup>

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THE HUBBLE HIGHER- $z$  SUPERNOVA SEARCH:  
SUPERNOVAE TO  $z \approx 1.6$  AND CONSTRAINTS ON TYPE Ia PROGENITOR MODELS <sup>a</sup>

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SWARA RAVINDRANATH<sup>2</sup>, MARIN RICHARDSON<sup>2</sup>, AND EDWARD TAYLOR<sup>15</sup>

Year 1	<b>399 orbits: Deep extragalactic studies 134 orbits: ToO 6 - 8 SNe at <math>1.2 &lt; z &lt; 1.8</math></b>
Year 2	<b>260 orbits w/ Supernova Cosmology Project!</b>
Year 3	<b>360 orbits for supernova studies.</b>
Year 4	<b>parallel data survey (incl. UDF)</b>
Year 5	<b>186 orbits for high-<math>z</math> SNe AND <math>H_0</math>!</b>

18% of HST time

Images

Subtractions

Dist

Ex

E

W

S

N

NE

SE

SW

NW

EW

NS

NE-SW

SW-NE

EW-NS

NS-EW

NE-SW

SW-NE

EW-NS

NS-EW

NE-SW

SW-NE

EW-NS

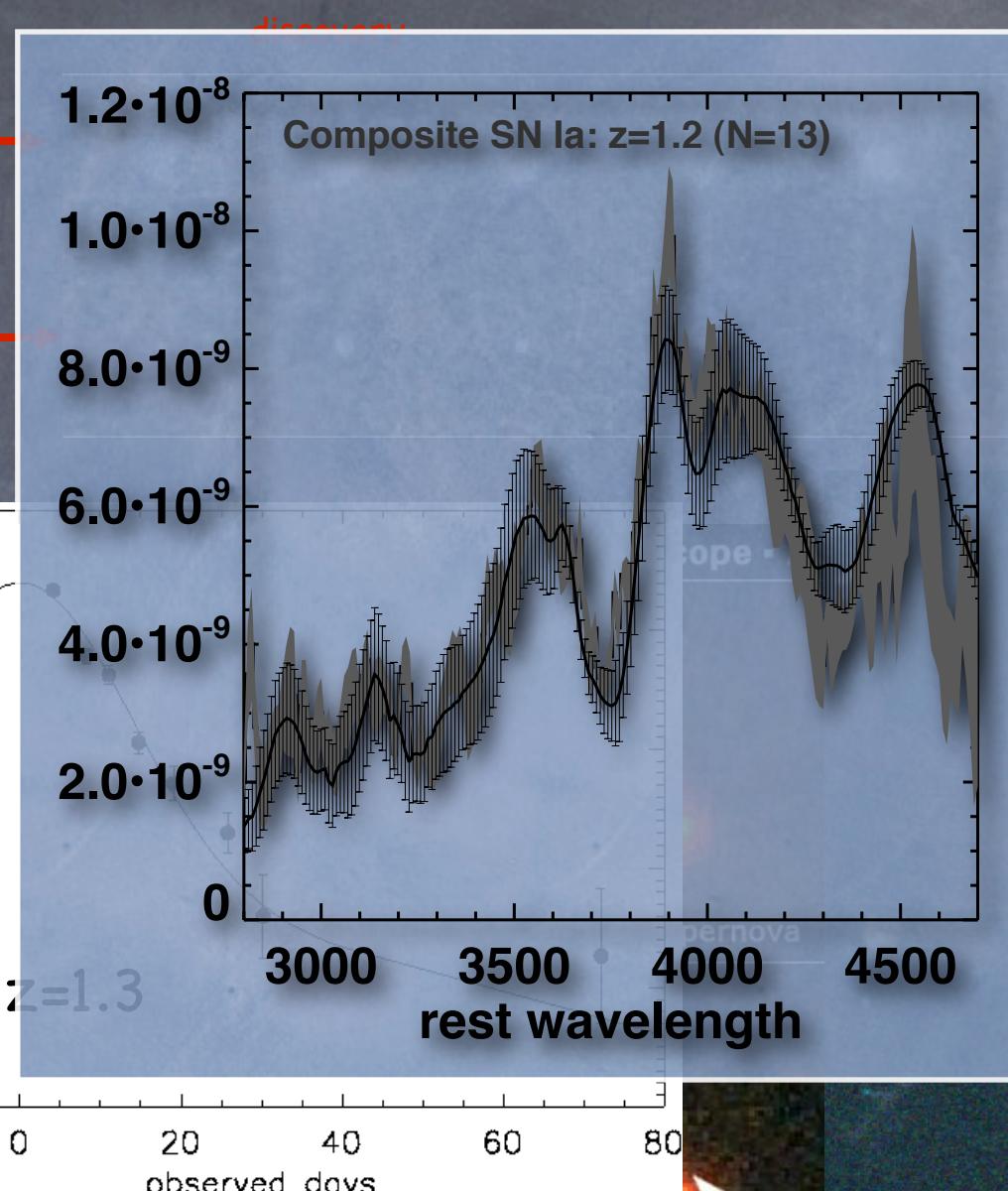
NS-EW

NE-SW

SW-NE

EW-NS

NS-EW

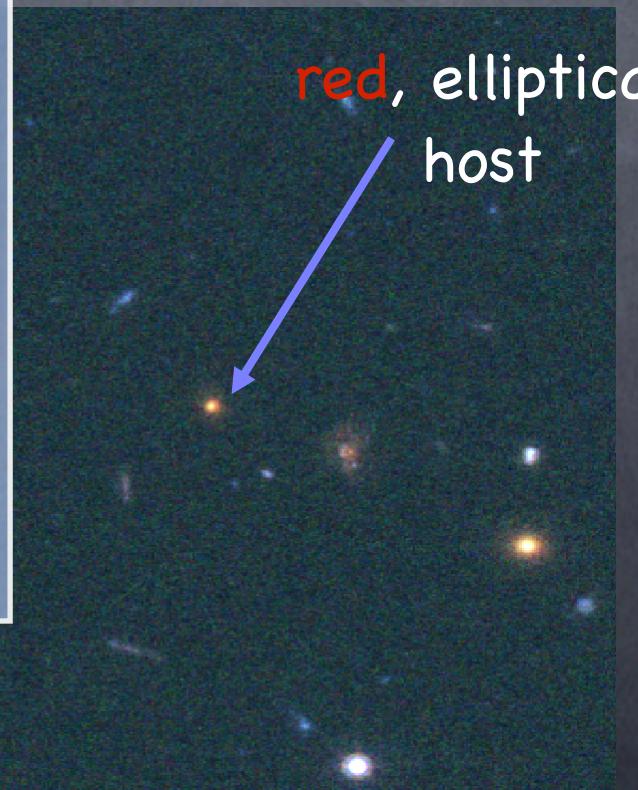


NASA and A. Riess (STScI) & L. Strolger (WKU)

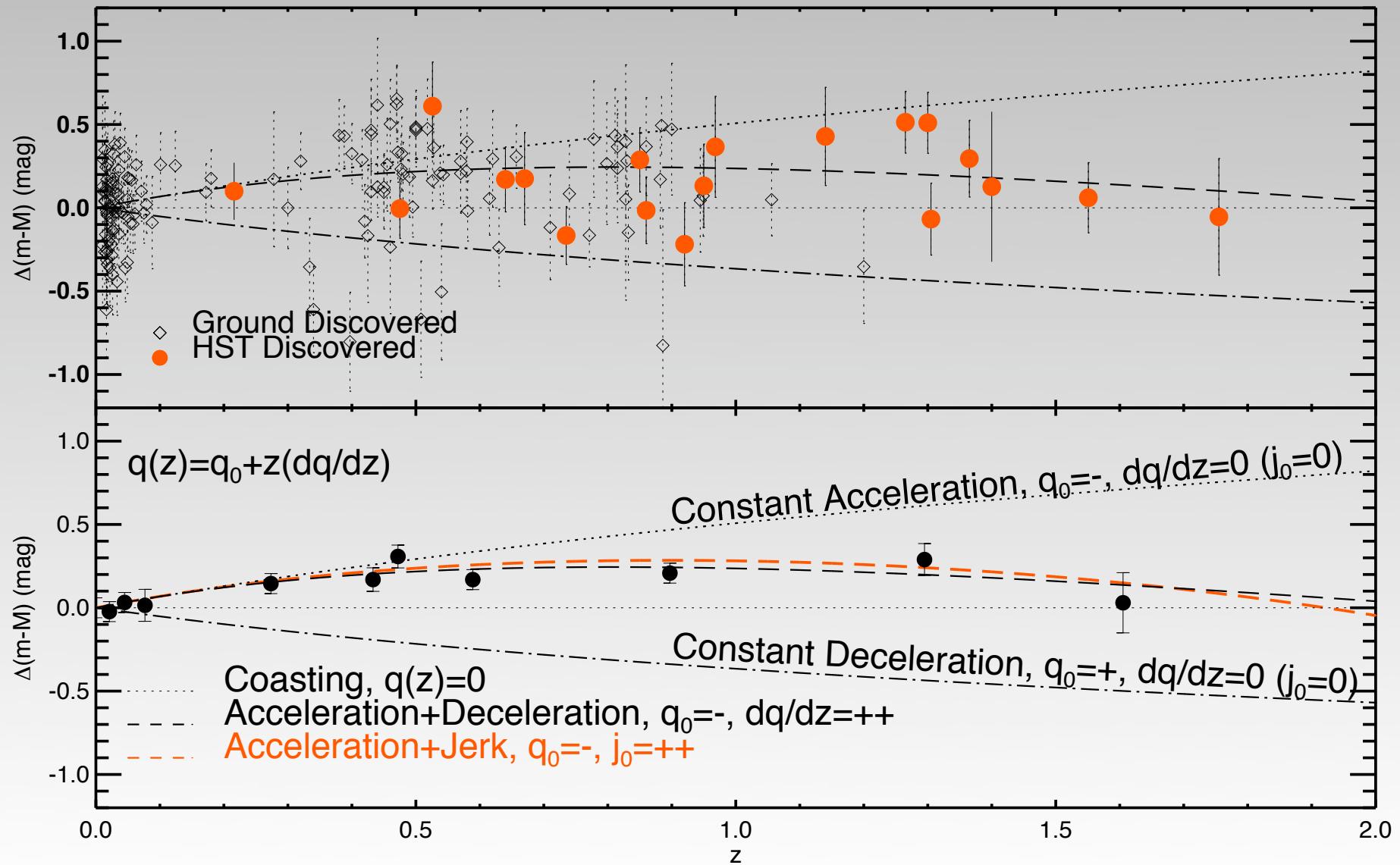
STScI-PRC04-12



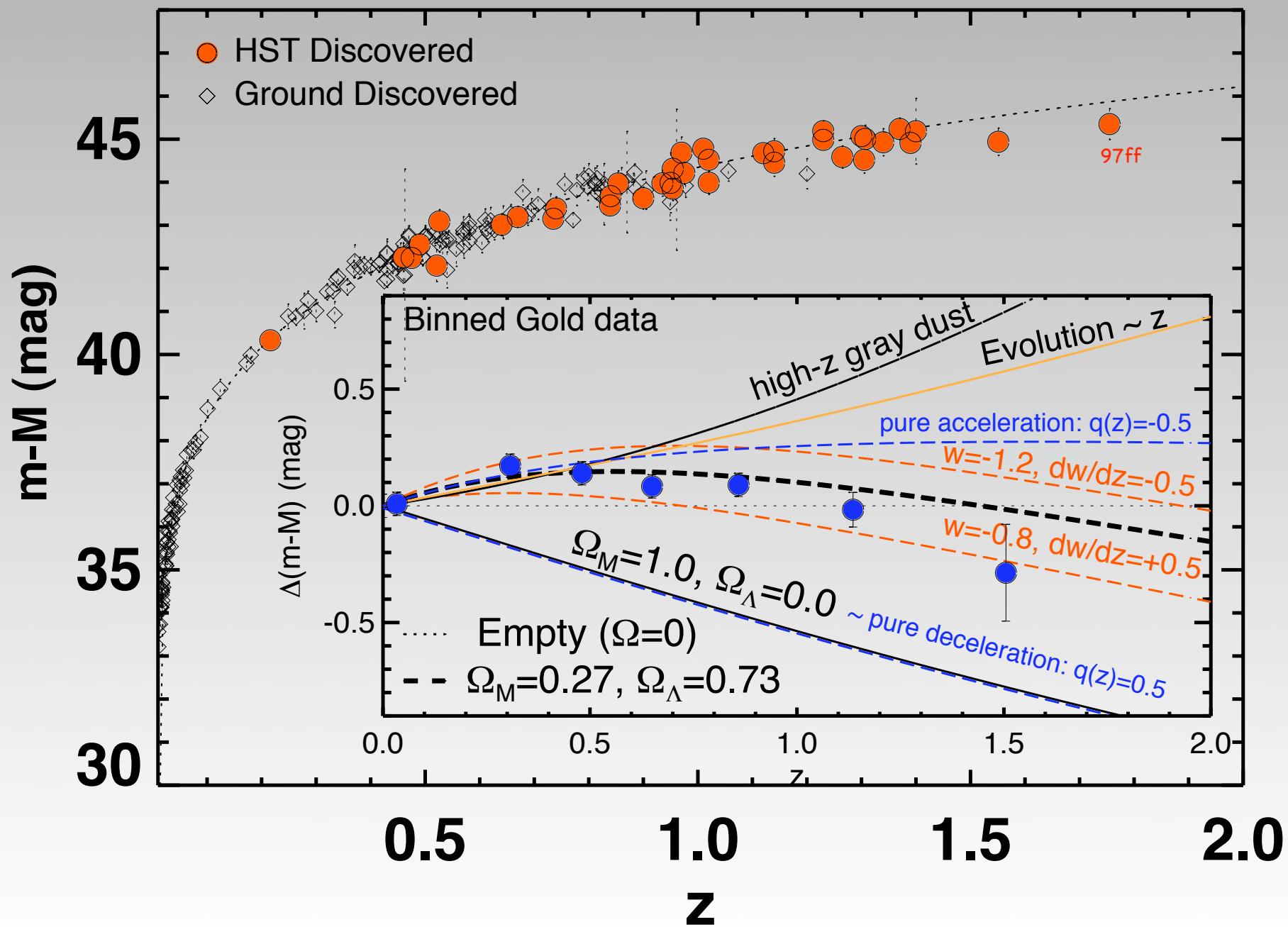
ACS grism spectrum



# A Cosmic Jerk: Deceleration gave way to Acceleration,



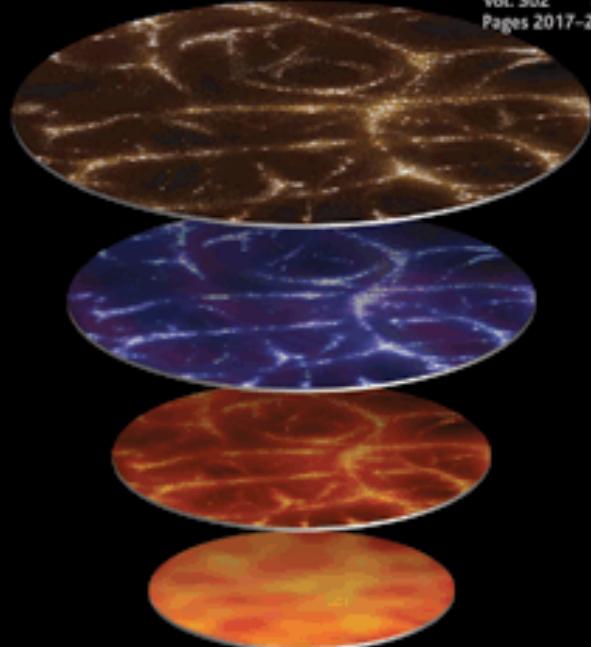
# The New SN Ia Hubble Diagram



# Science

19 December 2003

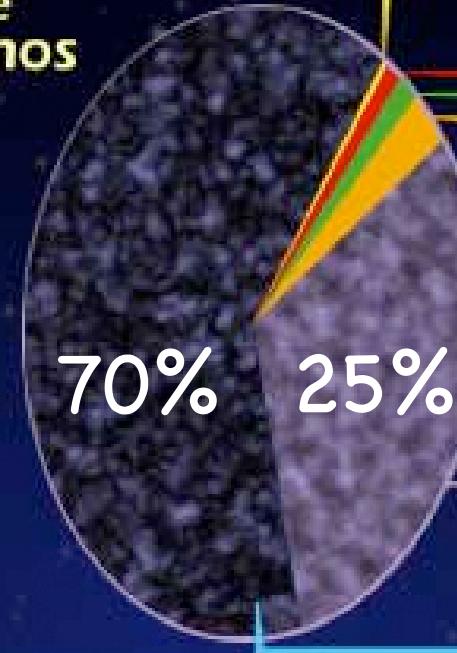
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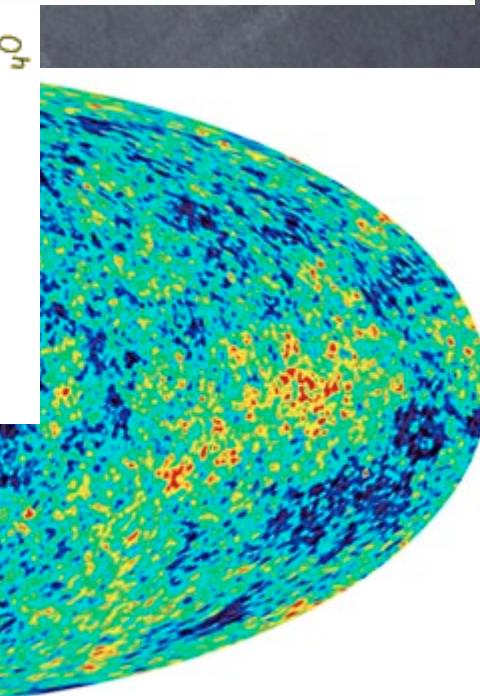
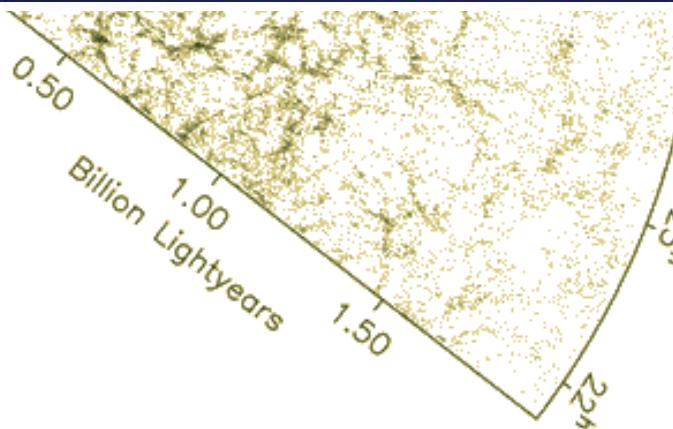
Breakthrough of the Year  
**Cosmic Convergence**

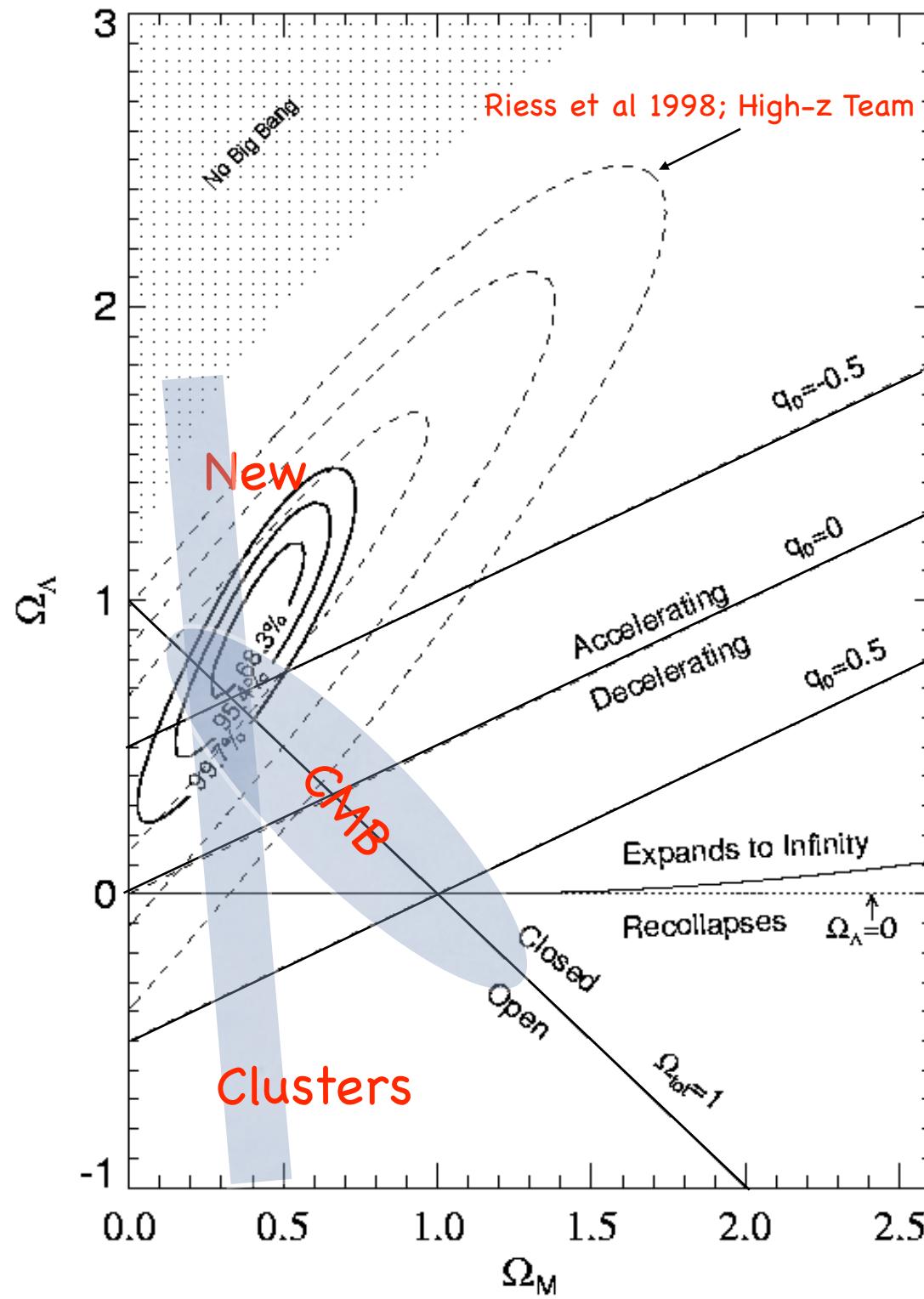
AMERICAN ASSOCIATION FOR THE ADVANCEMENT OF SCIENCE

## Composition of the Cosmos



Heavy elements:	0.03%
Ghostly neutrinos:	0.3%
Stars:	0.5%
Free hydrogen and helium:	4%
Dark matter:	30%
Dark energy:	65%





$\Omega_\Lambda$  only describes  $\rho_{\text{vac}}$

If a perfect fluid (and constant),

then  $U = P\Delta V$  says

$$P_{\text{vac}} = -\rho_{\text{vac}}c^2$$

Equation of state parameter reveals nature of Dark Energy!

$$w \equiv \frac{1}{c^2} \left( \frac{P}{\rho} \right)_{\text{vac}}$$

$$w(z) = w_0 + w'z; \quad w' \equiv \frac{dw}{dz} \quad w(z) = w_0 + \frac{w_a z}{1+z}$$



National Science Foundation  
and the  
National Aeronautics and Space Administration



Dr. Rocky Kolb (Chair)  
Fermi National Accelerator Laboratory

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University of California, Davis

Dr. Gary Bernstein  
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Dr. John Mather  
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Dr. Suzanne Staggs  
Princeton University

Dr. Nicholas Suntzeff  
Texas A&M University

# The Next Step: critical observations

- ⌚ Galaxy Cluster counts (# of large structures)
- ⌚ Baryon Acoustic Oscillations (BAO)
- ⌚ Supernovae
- ⌚ Weak Lensing

And somewhat later:

- ⌚ Integrated Sachs-Wolfe effect
- ⌚ GRBs as standard candles
- ⌚ Gravitational Wave experiments

## The Future... Is NOW!

- ⦿ SNLS, ESSENCE -- Constrains on  $w$  via supernovae  $\langle z \rangle = 0.5$ . SNLS already a 100's!
- ⦿ SDSS -- constraints on  $\Omega_M$ /growth of structure via CL & weak lensing, SN result expected.
- ⦿ WMAP -- BAO, SZ, & critical distance to  $z=1089$
- ⦿ SHOES -- Simultaneous constraints on  $H_0$  and  $w'(w_a)$  via supernovae & cepheids.

A little further down  
the road... < 5 yrs.

- Dark Energy Survey
- Pan-STARRS-4
- ALPACA

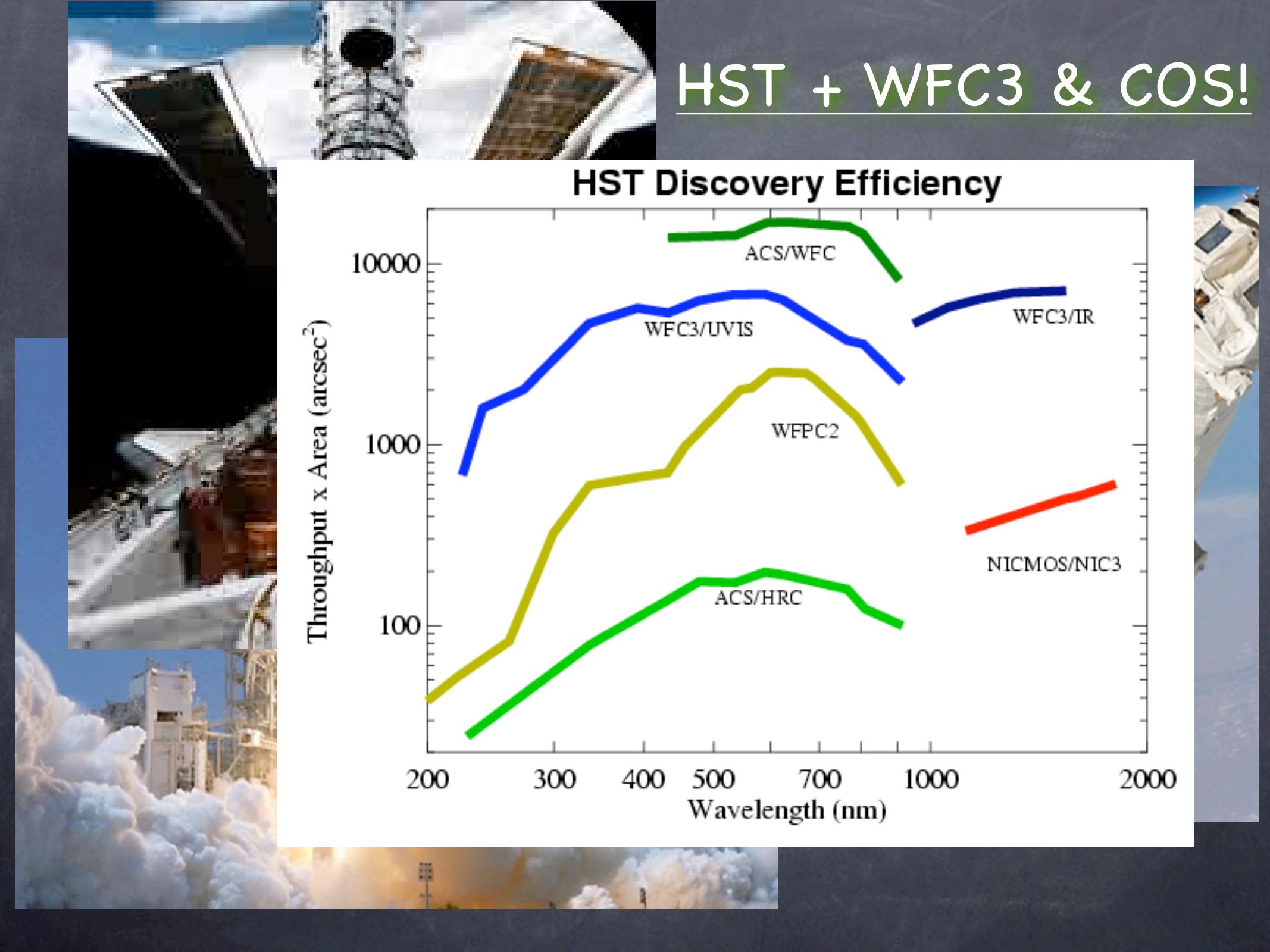
approx.  
1000°'s



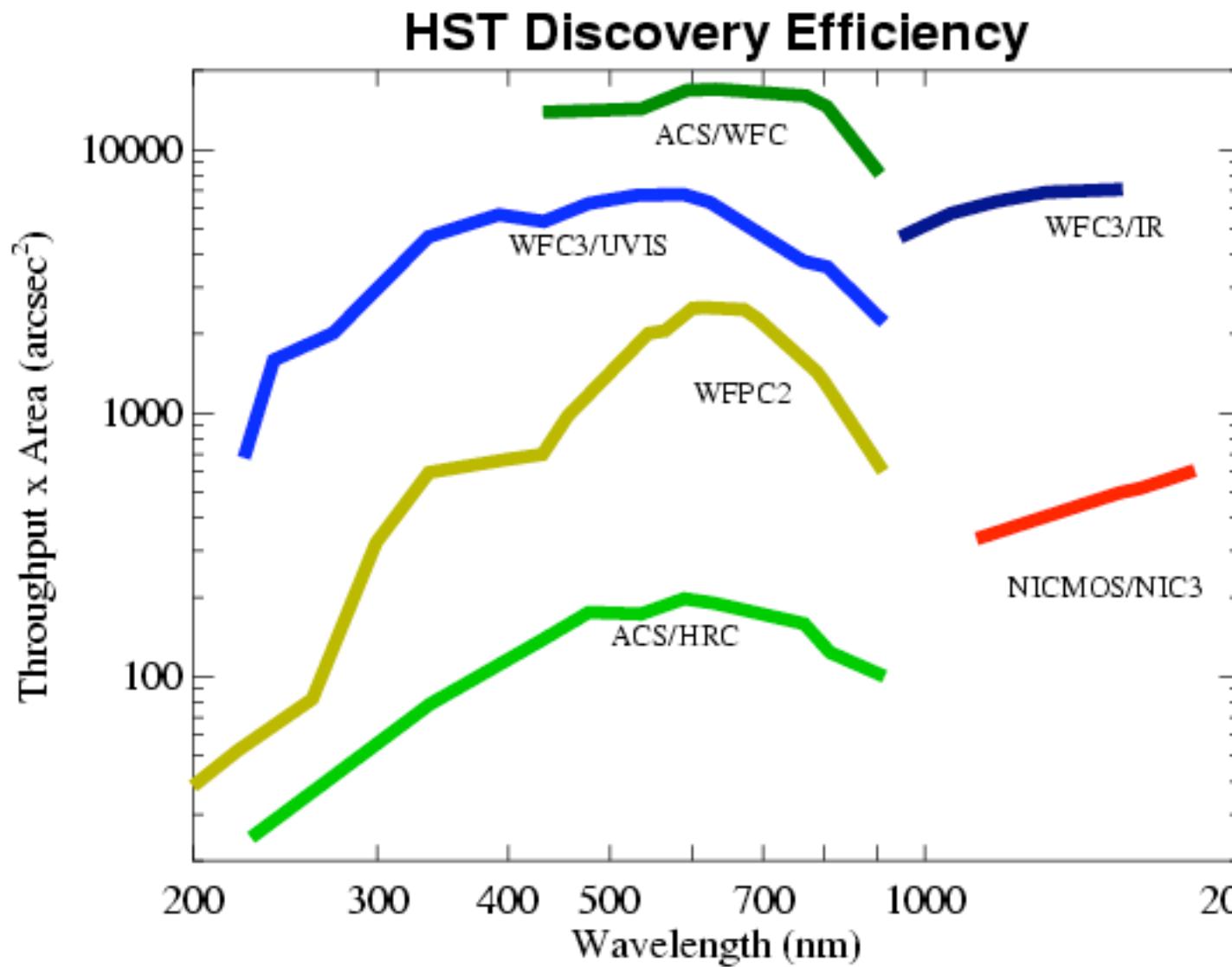
- Cornell-Caltech Atacama Telescope
- Cluster Imaging eXperiment (CIX)

SZ cluster  
detection





HST + WFC3 & COS!



Still further yet...  
approx. 10 yrs.

LSST: 8-m “all sky” surveyor.  
10,000°’s! Hemisphere every week!



GSMT: ultra deep optical & IR.

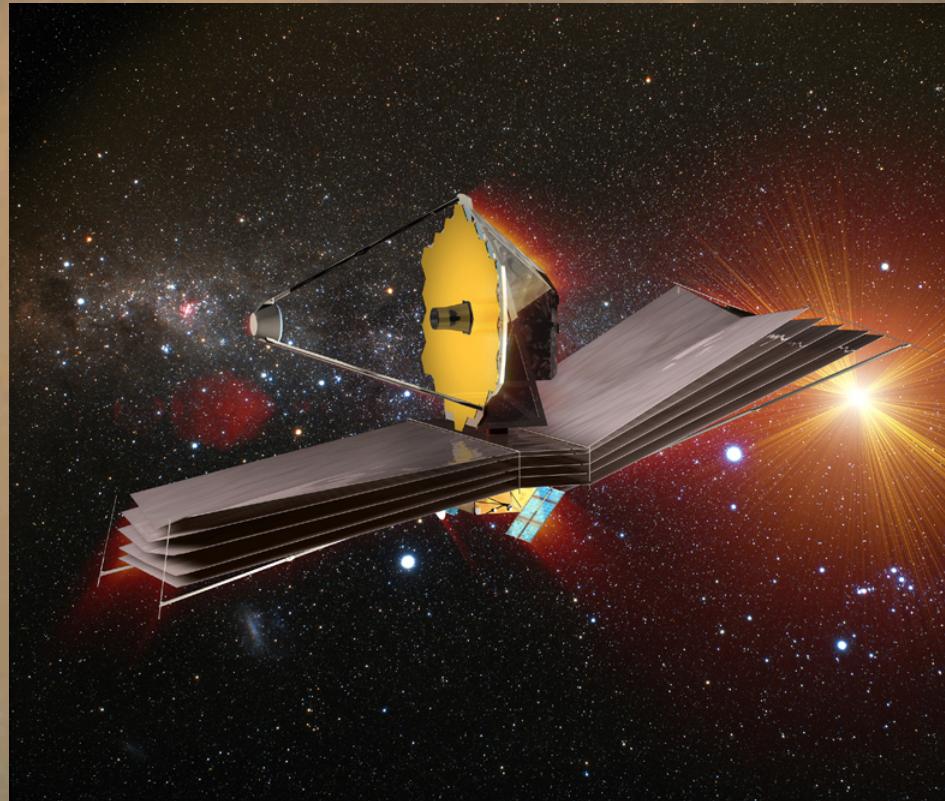
- Giant Magellan Telescope
- Thirty-Meter Telescope

- ⦿ SKA- Next logical step beyond ALMA,  
approx. 2 orders of mag. more sens. than VLA!

The distant, uncertain future...

## space-based projects

- ⦿ Constellation-X: Next logical step from Chandra
- ⦿ JWST

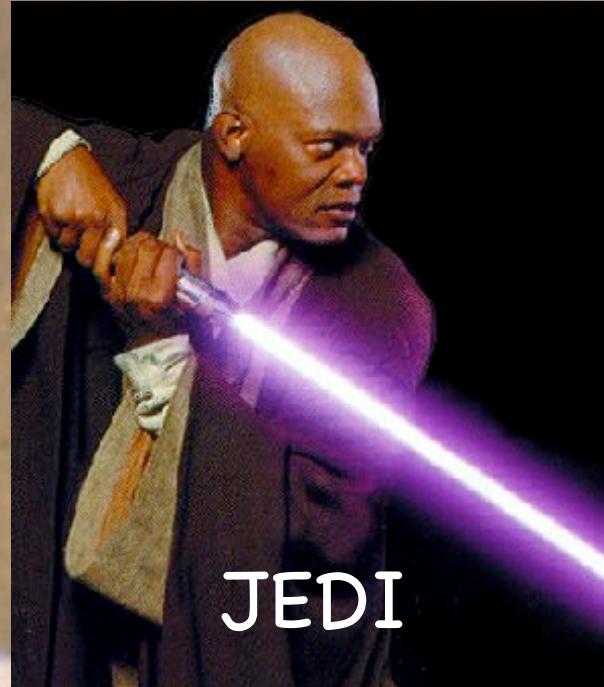
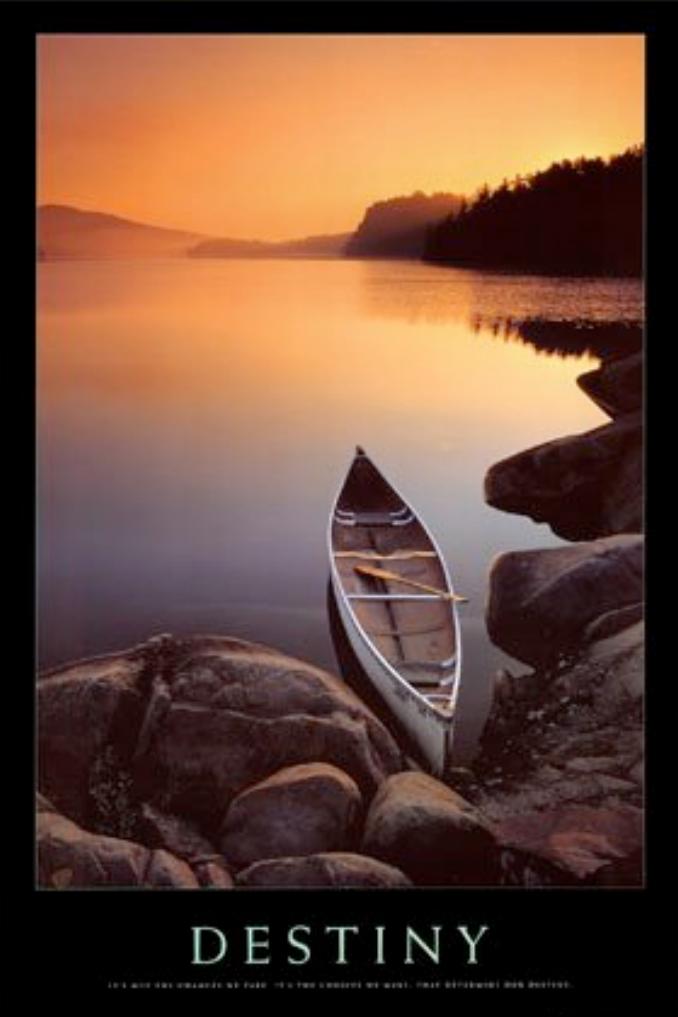


The distant, uncertain future...

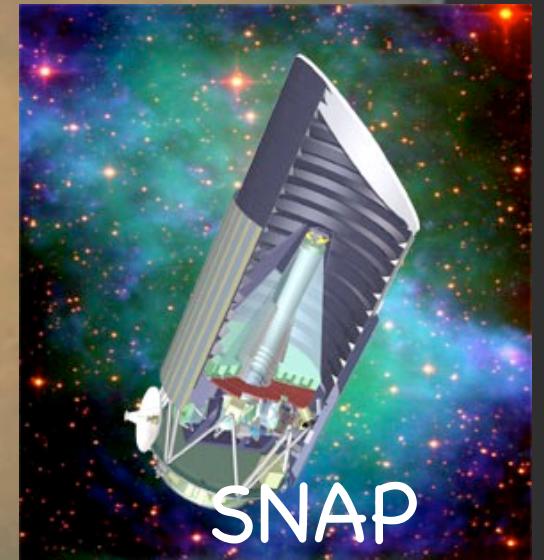
space-based projects



Joint Dark Energy Mission  
(NASA/DOE)



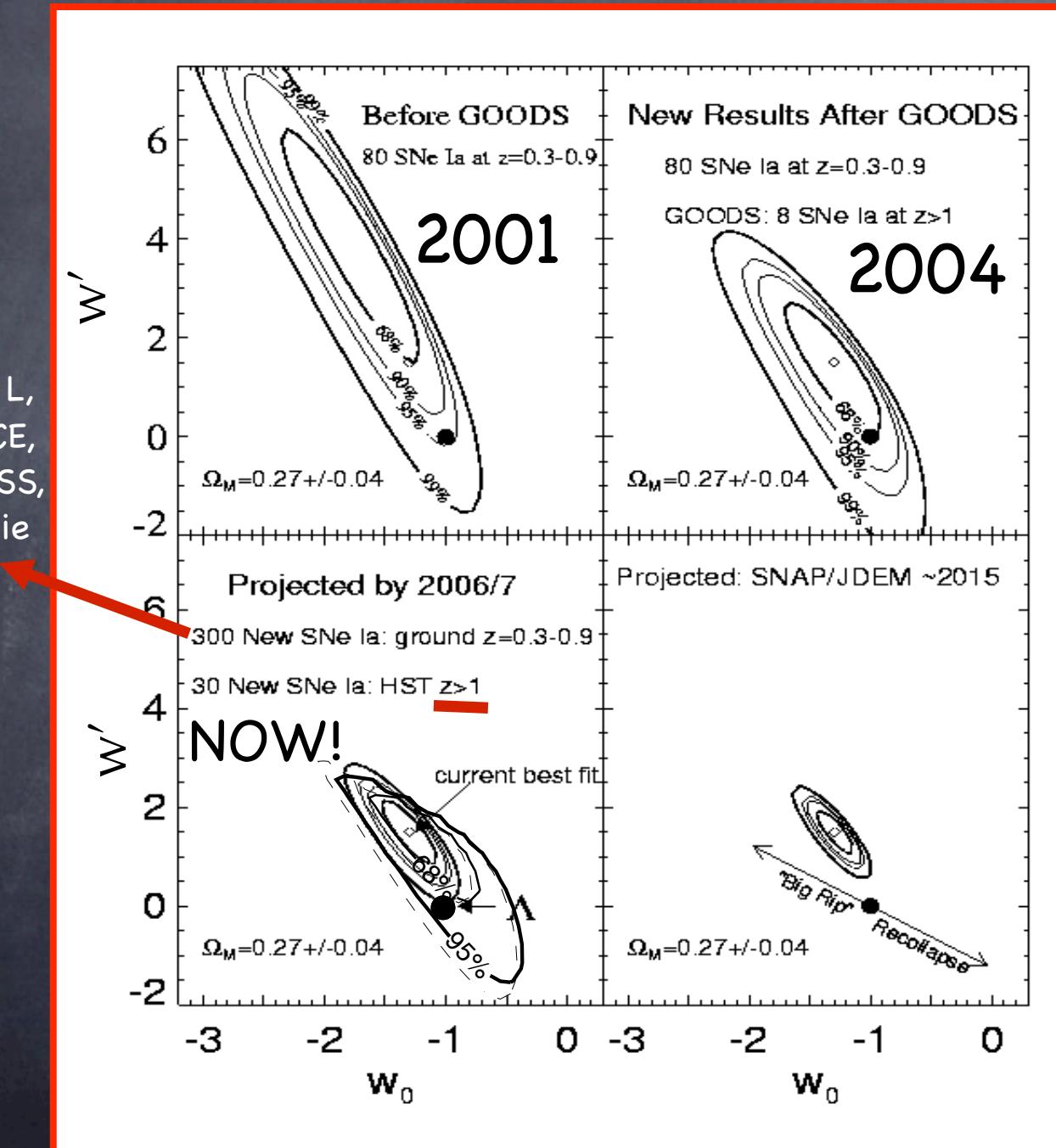
JEDI



SNAP

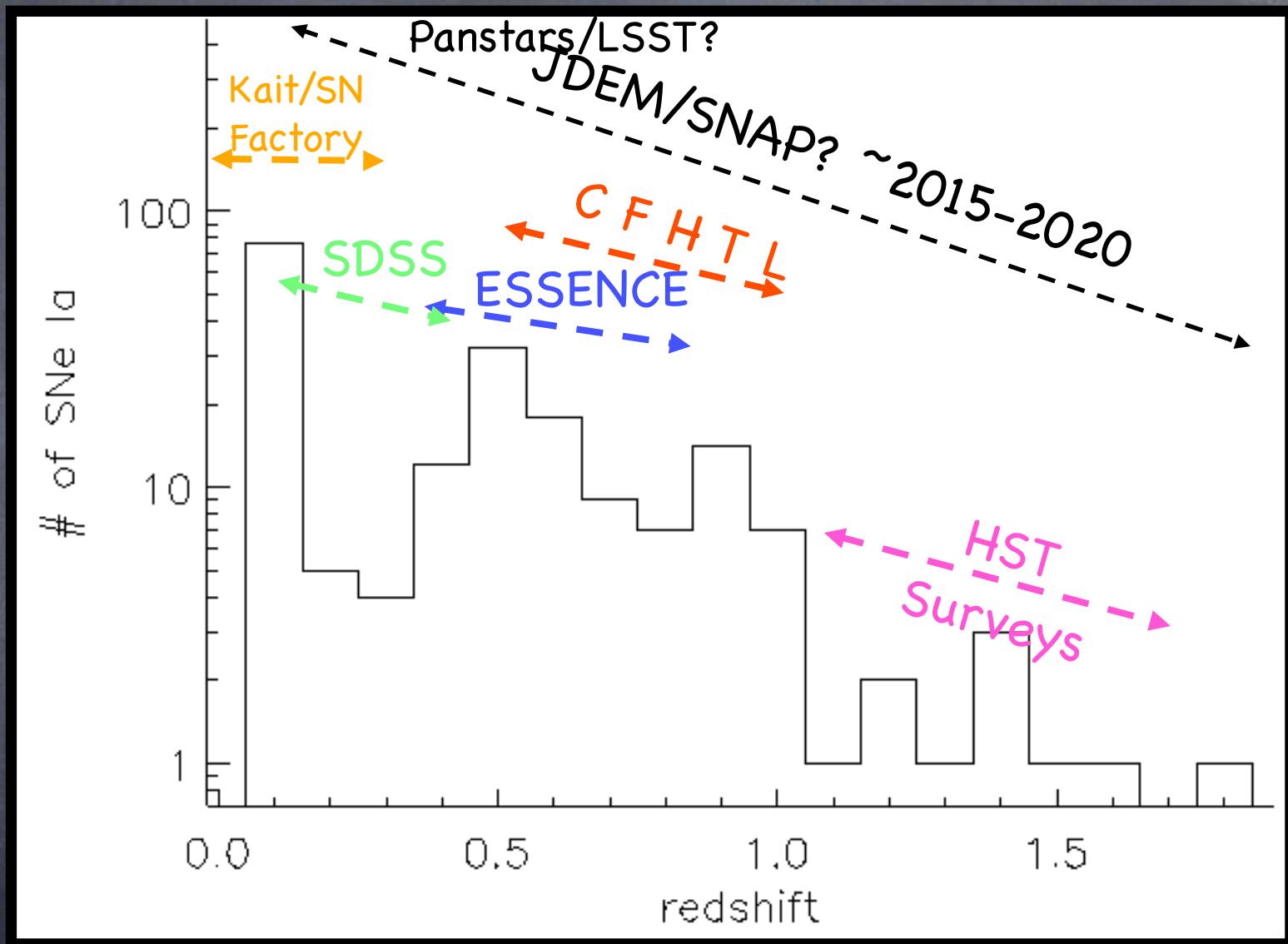
# Closing in on Dark Energy: The Near Future

CFHT L,  
ESSENCE,  
Sloan DSS,  
Carnegie



Rejection of  $\Lambda$ :  
i.e.,  $w_0 \neq -1$  or  $w' \neq 0$   
would be a tremendous  
breakthrough!

# Phase Space of Supernova Dark Energy Surveys



The End