

A light neutralino in hybrid models of supersymmetry breaking

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GDR SUSY

Outline

- 1 Towards hybrid models of Supersymmetry breaking
- 2 The rise of a light neutralino

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SQCD and SUSY breaking

Duality

quarks Q and anti-quarks \tilde{Q}
with N_c colors, N_f flavors
For $N_c < N_f < 3/2N_c$

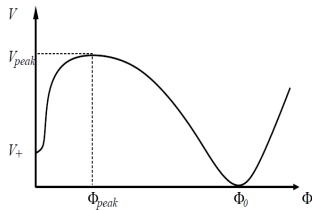


N_f flavors, $N = N_f - N_c$ colors
quarks q , antiquarks \tilde{q} and
meson X

Superpotential

$$W = hqX\tilde{q} - h\mu^2 \text{Tr}X$$

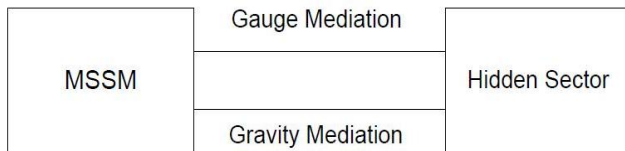
Intriligator, Seiberg and Shih,
hep-th/0602239



Why SUSY is broken

$$F_{X^i} = hq_i^a \tilde{q}_a^j - h\mu^2 \mathbb{1}$$

Mediation of SUSY breaking : General Scheme



Gravity Mediation

- Natural $\mu, B\mu$
- Generates Flavor Changing Neutral Currents
- LSP neutralino

Gauge Mediation

- Have to find a mechanism to generate $\mu, B\mu$
- Flavor Blind process
- LSP gravitino

Having both mediations

Order of magnitude

- Gauge mediation

$$\Lambda_{GM} \sim \frac{\alpha}{4\pi} \frac{F}{M}$$

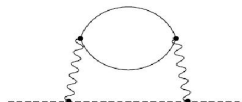
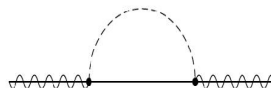
- Gravity mediation

$$m_{3/2} \sim \frac{F}{M_{Pl}}$$

Heavy messengers

$$m_{3/2} \sim (0.01 - 0.1) \Lambda_{GM}$$

$$\rightarrow M \sim 10^{13} - 10^{14} \text{ GeV}$$



Do messengers interact with the GUT sector ?

Model : coupling messengers to the GUT sector

Our Model

Lets consider

- ISS as a hidden sector
- Messengers coupled to the other fields by $\phi(\lambda_X X + \lambda_\Sigma \Sigma)\tilde{\phi}$, with $\Sigma = \text{diag}(2,2,2,-3,-3)v$ that breaks $SU(5)$

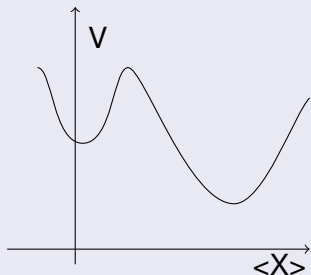
Model bulding characteristics

Why studying this coupling ?

- Gravitational effects : $m_{3/2} \sim (0.1 - 0.01)\Lambda_{GM}$
- One free parameter λ_Σ
- Doublet and triplet messengers have different masses
 $M = \lambda_\Sigma 6vY$

Shape of potentials

Quantum Corrections at one loop for ISS + messengers



We can explicitly compute the Coleman-Weinberg potential and we find

- Metastability for $\lambda_X < 10^{-2}$
- $\langle X \rangle \neq 0$

Mass term for the Higgs

Non-renormalisable term ... but non negligible !

$$W_1 = \lambda_1 \frac{q\tilde{q}}{M_{Pl}} H_u H_d \rightarrow \boxed{\mu}$$

$$W_2 = \lambda_2 \frac{X^2}{M_{Pl}} H_u H_d \rightarrow \boxed{B\mu}$$

UV spectrum

5- $\bar{5}$

$$M_3 = \frac{1}{2} N_m \frac{\alpha_3}{4\pi} \frac{\lambda_X F_X}{\lambda_{\Sigma V}}$$

$$M_2 = -\frac{1}{3} N_m \frac{\alpha_2}{4\pi} \frac{\lambda_X F_X}{\lambda_{\Sigma V}}$$

$$m_Q^2 : m_{U^c}^2 : m_{D^c}^2 : m_L^2 : m_{E^c}^2 \\ \approx 0.79 : 0.70 : 0.68 : 0.14 : 0.08$$

10- $\bar{10}$

$$M_3 = \frac{7}{4} N_m \frac{\alpha_3}{4\pi} \frac{\lambda_X F_X}{\lambda_{\Sigma V}}$$

$$M_2 = 3 N_m \frac{\alpha_2}{4\pi} \frac{\lambda_X F_X}{\lambda_{\Sigma V}}$$

$$m_Q^2 : m_{U^c}^2 : m_{D^c}^2 : m_L^2 : m_{E^c}^2 \\ \approx 8.8 : 5.6 : 5.5 : 3.3 : 0.17$$

Commun feature

$$M_{1,1-loop,GM} \sim \text{Tr}(Y^2/M)$$

$$M_{1,1-loop,GM} \sim \text{Tr}(Y) = 0!$$

$$M_1 \sim m_{3/2}$$

 \neq

Universality for gauginos

$$M_1 : M_2 : M_3 = \alpha_1 : \alpha_2 : \alpha_3$$

How far can we go in the hierarchy?

Limits come from

- Heavy masses for the messengers
- Gravitational effects interesting
- Neutralino LSP

IR spectrum

Constraint by Dark Matter

What are the bounds on $m_{\chi_1^0}$?

- No bound from colliders !
- bound from dark matter constraint (relic density)

What is the lightest neutralino that we can build in this model ?

- $m_{\chi_1^0} \leq 40$ GeV would need R-parity violation
- $m_{\chi_1^0} \sim 40$ GeV with light sleptons
- $m_{\chi_1^0} \geq 40$ GeV but less interesting for the moment...

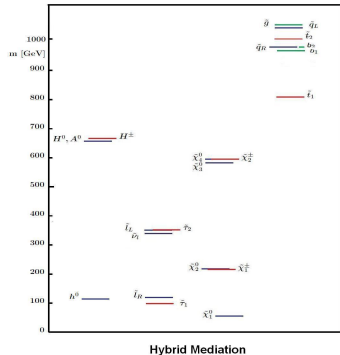
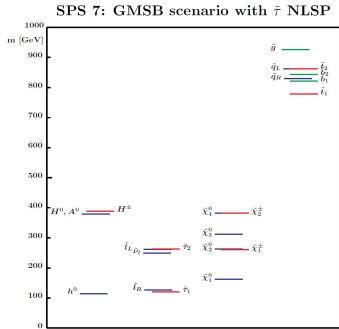
FCNC

Light Sleptons \rightarrow FCNC in the lepton sector

Difficult to avoid because of the very few parameters of the model

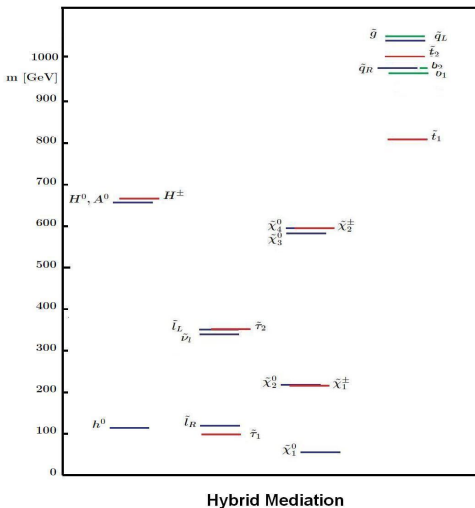
GMSB vs hybrid mediation

From B.C. Allanach et al., hep-ph/0202233



$\Lambda = 40 \text{ TeV}$, $M_{\text{mes}} = 80 \text{ TeV}$, $N_{\text{mes}} = 3$, $\tan\beta = 15$, $\mu > 0$.

GMSB vs hybrid mediation



Conclusions

From the model building point of view

- Explicit models of supersymmetry breaking
- Quite constrained model
- Correct EWSB
- Correct relic abundance
- Still FCNC

Even if gauge mediation, light neutralino as a signature

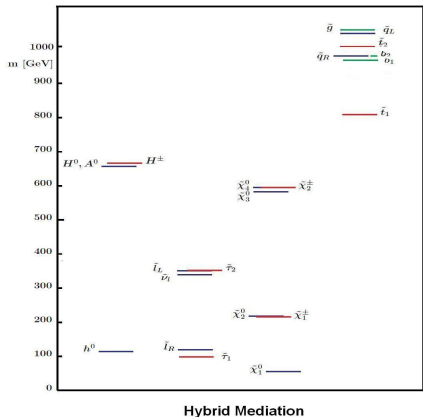
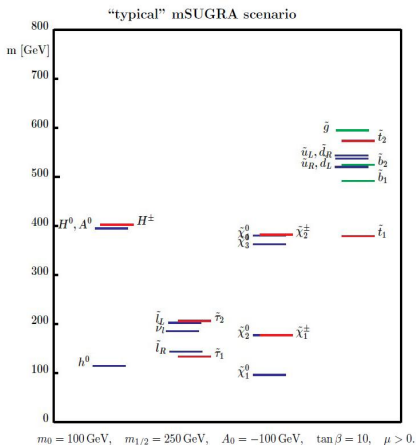
Waiting for the LHC ...

Datas

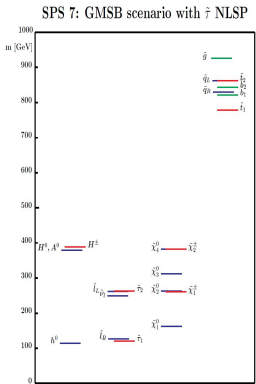
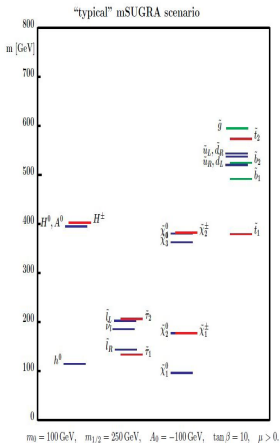
model	1	2	3	3 bis	4	5	6
$N_{(5,5)}$	1	6	0	0	0	1	3
$N_{(10,10)}$	0	0	1	1	4	1	1
M_{GM}	1000	200	300	300	110	220	160
M_1	50	50	50	85	80	85	85
$\tan \beta$	30	24	15	15	9	15	15
$\text{sign}(\mu)$	+	+	+	+	+	+	+
h	114.7	115.0	115.2	115.2	116.5	114.6	114.8
A	779.2	645.4	892.2	892.4	1015	735.8	662.7
H^0	779.2	645.5	892.4	892.6	1015	735.9	662.8
H^\pm	783.3	650.3	895.7	895.9	1018	740.1	667.5
$\tilde{\chi}_1^\pm$	259.4	305.0	560.2	560.3	676.7	408.0	223.9
$\tilde{\chi}_2^\pm$	747.8	636.8	693.9	694.0	970.4	590.4	597.5
$\tilde{\chi}_1^0$	24.5	23.5	23.2	42.9	38.1	43.0	42.9
$\tilde{\chi}_2^0$	259.4	305.0	560.1	560.3	677.1	408.0	223.9
$\tilde{\chi}_3^0$	743.3	629.8	596.9	597.1	691.0	570.8	589.2
$\tilde{\chi}_4^0$	745.7	634.7	693.8	693.9	970.4	590.4	596.3
$ Z_{11} $	0.9982	0.9975	0.9971	0.9971	0.9978	0.9968	0.9969
$ Z_{13} $	0.0599	0.0708	0.0750	0.0755	0.0648	0.0792	0.0772
\tilde{g}	1064	1207	1097	1097	1527	1028	1063
\tilde{t}_1	984.6	927.3	861.7	861.6	1080	795.7	809.5
\tilde{t}_2	1156	1074	1240	1240	1468	1058	1002
\tilde{u}_1, \tilde{c}_1	1195	1087	1135	1135	1361	1006	987.9
\tilde{u}_2, \tilde{c}_2	1240	1115	1327	1327	1555	1118	1043
\tilde{b}_1	1128	1040	1123	1123	1356	995.4	966.2
\tilde{b}_2	1169	1079	1224	1224	1451	1038	987.1
\tilde{d}_1, \tilde{s}_1	1184	1085	1134	1134	1360	1005	987.1
\tilde{d}_2, \tilde{s}_2	1243	1117	1329	1329	1557	1121	1046
$\tilde{\tau}_1$	242.2	99.0	86.3	89.3	87.0	96.7	95.2
$\tilde{\tau}_2$	420.3	289.4	696.2	696.3	753.1	498.6	349.8
$\tilde{e}_1, \tilde{\mu}_1$	294.4	150.6	131.5	133.6	105.4	123.6	117.4
$\tilde{e}_2, \tilde{\mu}_2$	413.4	275.1	699.1	699.2	754.1	500.1	348.5
$\tilde{\nu}_\tau$	396.6	260.5	691.4	691.5	749.0	491.4	337.6
$\tilde{\nu}_e, \tilde{\nu}_\mu$	405.8	263.6	694.8	694.9	750.1	493.9	339.5
$\Omega_{\tilde{\chi}_1^0} h^2$	6.40	0.428	0.279	0.122	0.124	0.118	0.116

mSUGRA vs hybrid mediation

From B.C. Allanach et al., hep-ph/0202233



Comparisons



$\Lambda = 40 \text{ TeV}, M_{\text{mes}} = 80 \text{ TeV}, N_{\text{mes}} = 3, \tan \beta = 15, \mu >$

