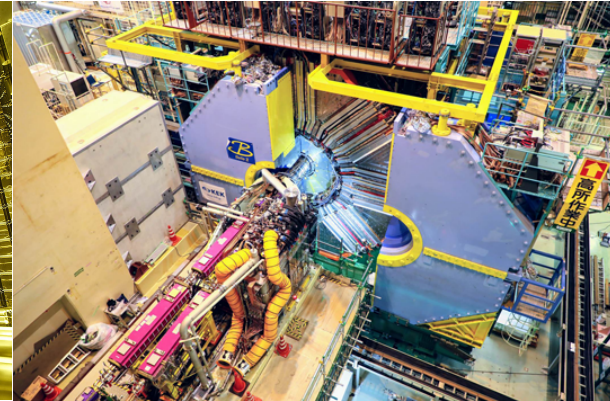
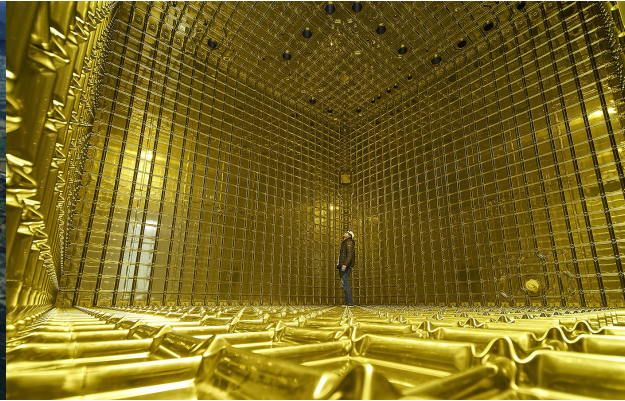
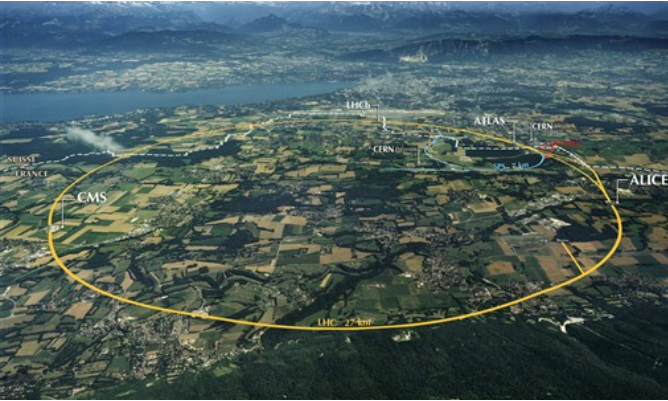


Particle Physics Highlights and Challenges



Beate Heinemann

DESY and Albert-Ludwigs-Universität Freiburg



Outline

(personal selection of material)

CERN-ESU-004
30 September 2019

- Electroweak Physics
- Strong Interactions
- Flavour Physics
- Neutrino Physics
- Dark Matter and Dark Sectors
- Hierarchy Problem

Physics Briefing Book

Input for the European Strategy for Particle Physics Update 2020

Electroweak Physics: Richard Keith Ellis⁴, Beate Heinemann^{2,3} (Conveners)
Jorge de Blas^{4,5}, Maria Cepeda⁶, Christophe Grojean^{2,7}, Fabio Maltoni^{8,9}, Alessandro Nisati¹⁰,
Elisabeth Petit¹¹, Riccardo Rattazzi¹², Wouter Verkerke¹³ (Contributors)

Strong Interactions: Jorgen D'Hondt¹⁴, Krzysztof Redlich¹⁵ (Conveners)
Anton Andronic¹⁶, Ferenc Sikler¹⁷ (Scientific Secretaries)
Nestor Armesto¹⁸, Daniel Boer¹⁹, David d'Enterria²⁰, Tetyana Galatyuk²¹, Thomas Gehrmann²²,
Klaus Kirch²³, Uta Klein²⁴, Jean-Philippe Lansberg²⁵, Gavin P. Salam²⁶, Gunar Schnell²⁷,
Johanna Stachel²⁸, Tanguy Pierog²⁹, Hartmut Wittig³⁰, Urs Wiedemann³¹ (Contributors)

Flavour Physics: Belen Gavela³², Antonio Zoccolli³³ (Conveners)
Sandra Malvezzi³⁴, Ana Teixeira³⁵, Jure Zapun³⁶ (Scientific Secretaries)
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Gudrun Hiller⁴³, Gino Isidori⁴⁴, Yoshikata Kuno⁴⁵, Alberto Lusiani⁴⁶, Yosef Nir⁴⁷,
Marie-Hélène Schune⁴⁸, Marco Sozzi⁴⁹, Stephan Paul⁵⁰, Carlos Pena⁵¹ (Contributors)

Neutrino Physics & Cosmic Messengers: Stan Bentvelzen⁵², Marco Zito^{46,47} (Conveners)
Albert De Roeck⁵³, Thomas Schweitzer⁵⁴ (Scientific Secretaries)
Bonnie Fleming⁵⁵, Francis Halzen⁵⁶, Andreas Haungs⁵⁷, Marek Kowalski⁵⁸, Susanne Merens⁵⁹,
Mauro Mezzetto⁶⁰, Silvia Pascoli⁶¹, Bangalore Sathyaprakash⁶², Nicola Serra⁶³ (Contributors)

Beyond the Standard Model: Gian F. Giudice⁶⁴, Paris Spiliadis^{65,66} (Conveners)
Juan Alcaraz Maestre⁶⁷, Caterina Doglioni⁶⁸, Gaia Lanfranchi^{69,70}, Monica D'Onofrio⁷¹,
Matthew McCullough⁷², Gilad Perez⁷³, Philipp Roloff⁷⁴, Veronica Sanz⁷⁵, Andreas Weiler⁷⁶,
Andrea Wulzer^{4,77,78} (Contributors)

Dark Matter and Dark Sector: Shoji Asai⁷⁹, Marcela Carena⁸⁰ (Conveners)
Babette Dobrich⁸¹, Caterina Doglioni⁸², Joerg Jaeckel⁸³, Gordan Krnjaic⁸⁴, Jocelyn Monroe⁸⁵,
Konstantinos Petrakis⁸⁶, Christoph Weniger⁸⁷ (Scientific Secretaries/Contributors)

Accelerator Science and Technology: Caterina Biscari⁸⁸, Leonid Rivkin⁸⁹ (Conveners)
Philip Barrows⁹⁰, Frank Zimmermann⁹¹ (Scientific Secretaries)
Michael Benedikt⁹², Pierluigi Campana⁹³, Edda Gschwendtner⁹⁴, Erik Jensen⁹⁵, Mike Lamont⁹⁶,
Wim Leemans⁹⁷, Lucio Rossi⁹⁸, Daniel Schulte⁹⁹, Mike Seidel¹⁰⁰, Vladimir Shiltsev¹⁰¹,
Sveinar Stenlund¹⁰², Akira Yamamoto^{103,104} (Contributors)

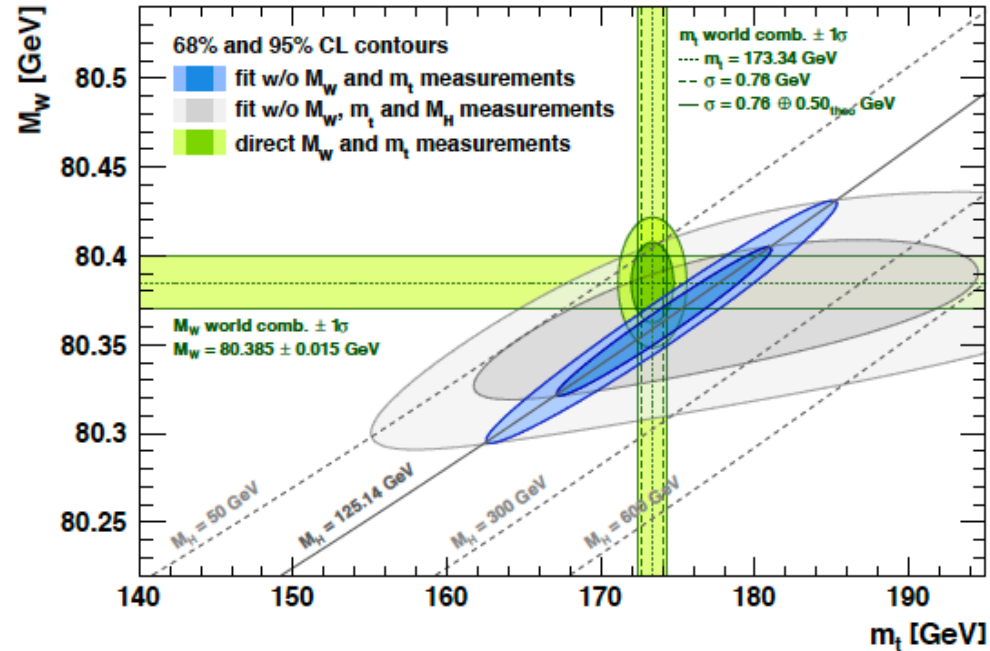
Instrumentation and Computing: Xinchou Lou¹⁰⁵, Brigitte Vachon¹⁰⁶ (Conveners)
Roger Jones¹⁰⁷, Emilia Lesgrands¹⁰⁸ (Scientific Secretaries)
Ian Bird¹⁰⁹, Simone Campana¹¹⁰, Ariella Cattai¹¹¹, Didier Contardo¹¹², Cinzia Da Via¹¹³, Francesco Forti¹¹⁴,
Maria Gironce¹¹⁵, Matthias Kasemann¹¹⁶, Lucie Lüssén¹¹⁷, Felix Sefkow¹¹⁸, Graeme Stewart¹¹⁹ (Contributors)

Editors: Halina Abramowicz⁷¹, Roger Forty²⁰, and the Conveners

Material based largely on [Physics Briefing Book](#) of
European Strategy for Particle Physics Update 2020, released on Oct. 2nd 2019

Electroweak Interactions

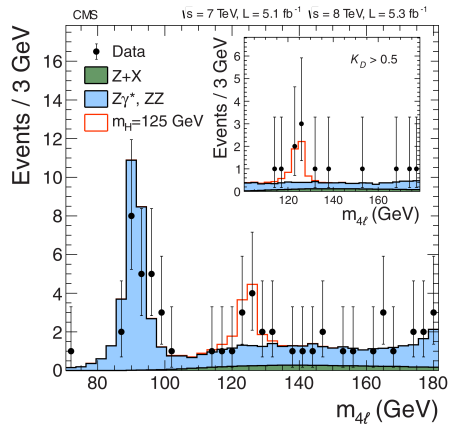
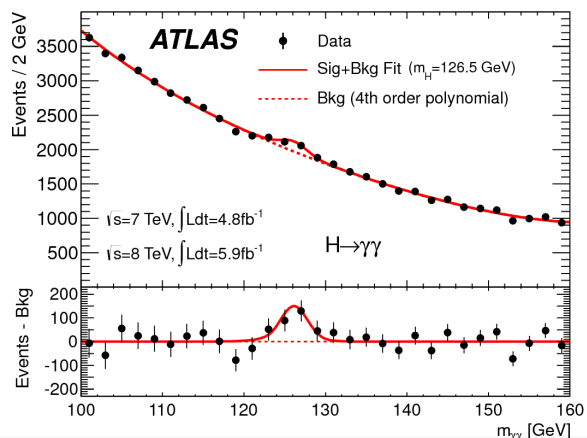
- Discovered mechanism to break electroweak symmetry
 - Higgs mechanism
- Provides technical solution but very unsatisfactory
- Higgs sector contains 15 ad-hoc parameters
 - Only 3 in gauge sector!



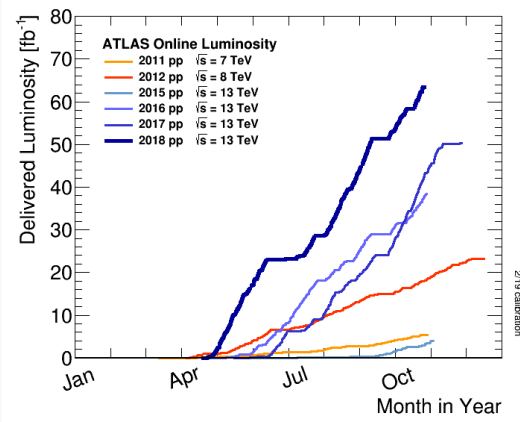
G. Giudice: “Essentially all problems or unsatisfactory aspects of the Standard Model are ultimately related to the structure of Higgs interactions”

Higgs Boson: 2012 vs 2019

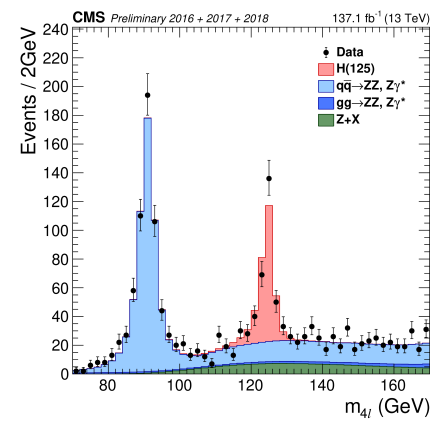
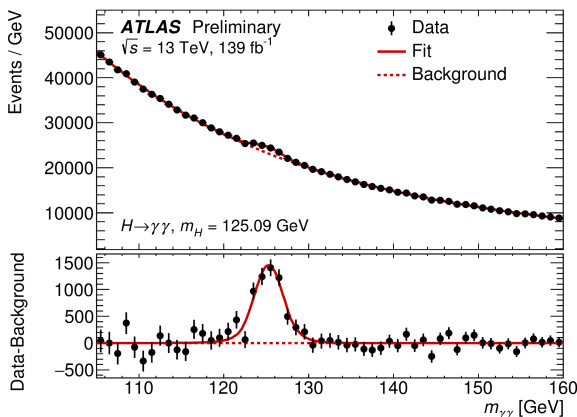
July 2012:
~10 fb⁻¹
@7-8 TeV



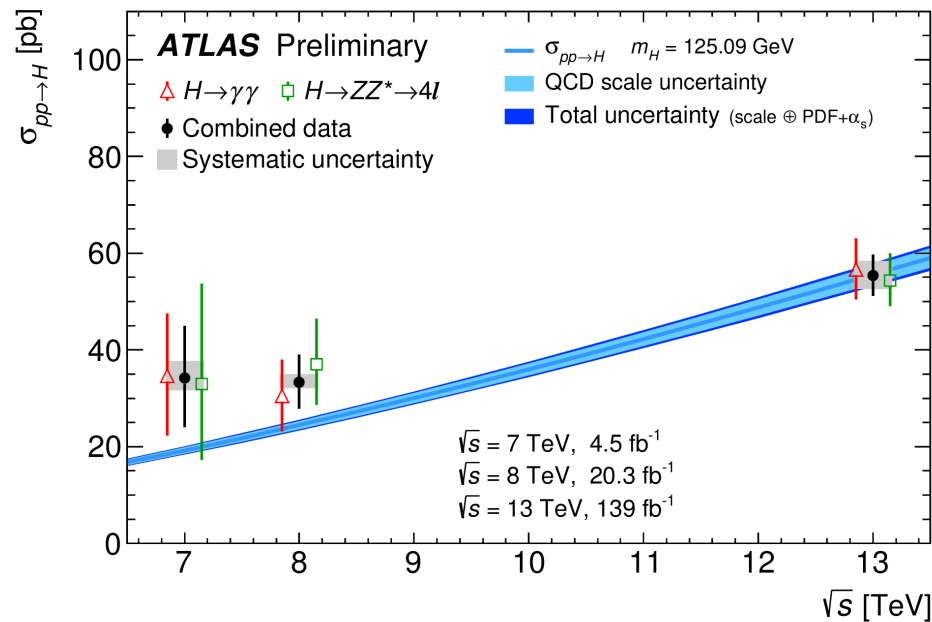
LHC luminosity



Today:
~140 fb⁻¹
@13 TeV

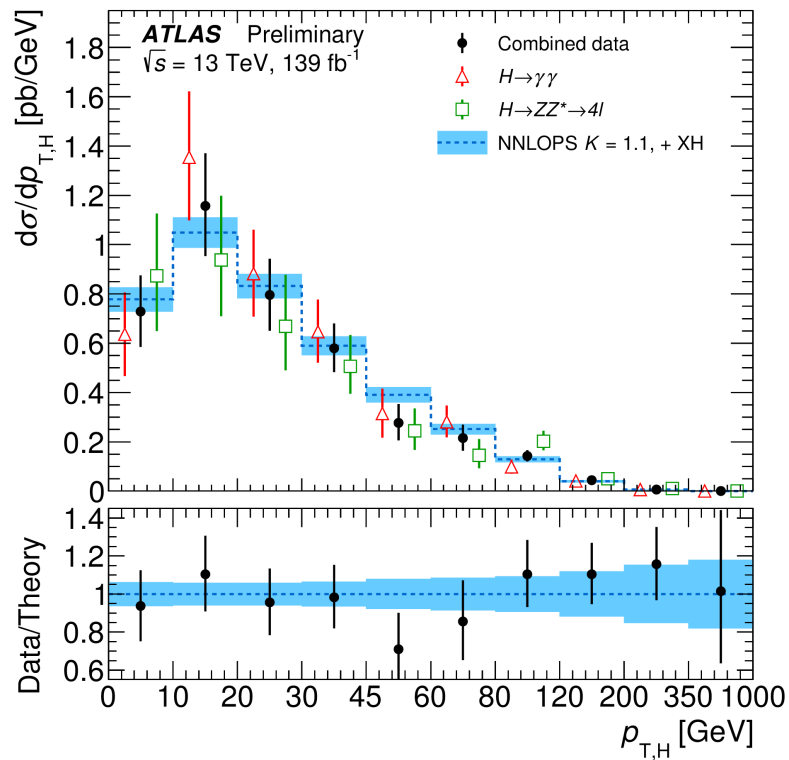


Differential H Cross Section Measurements



$$\sigma_{\text{data}} = 55.4 \pm 4.3 \text{ pb}$$

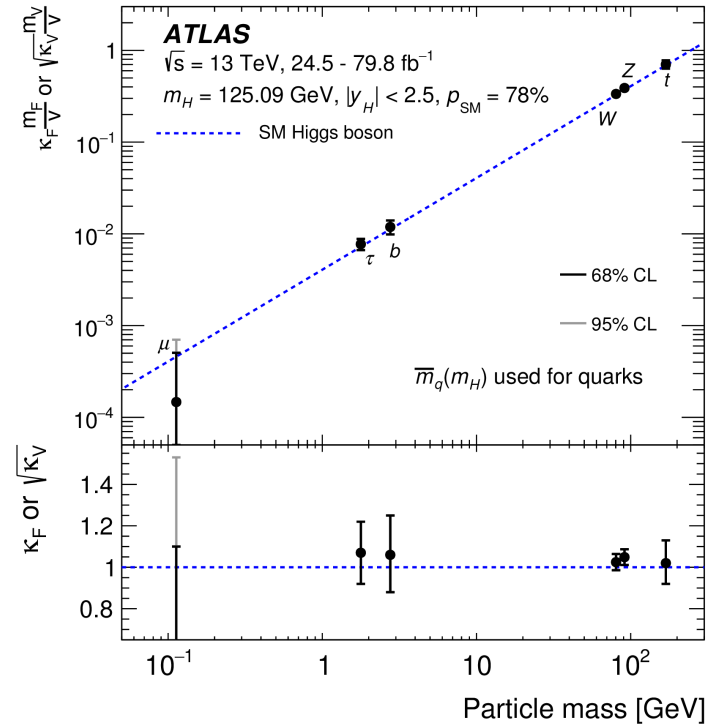
$$\sigma_{\text{SM}} = 55.6 \pm 2.5 \text{ pb (87\% ggF, N}^3\text{LO and NLO EWK)}$$



Higgs Couplings Today

Parameter	Result
κ_Z	1.10 ± 0.08
κ_W	1.05 ± 0.08
κ_b	$1.06^{+0.19}_{-0.18}$
κ_t	$1.02^{+0.11}_{-0.10}$
κ_τ	1.07 ± 0.15
κ_μ	< 1.53 at 95% CL

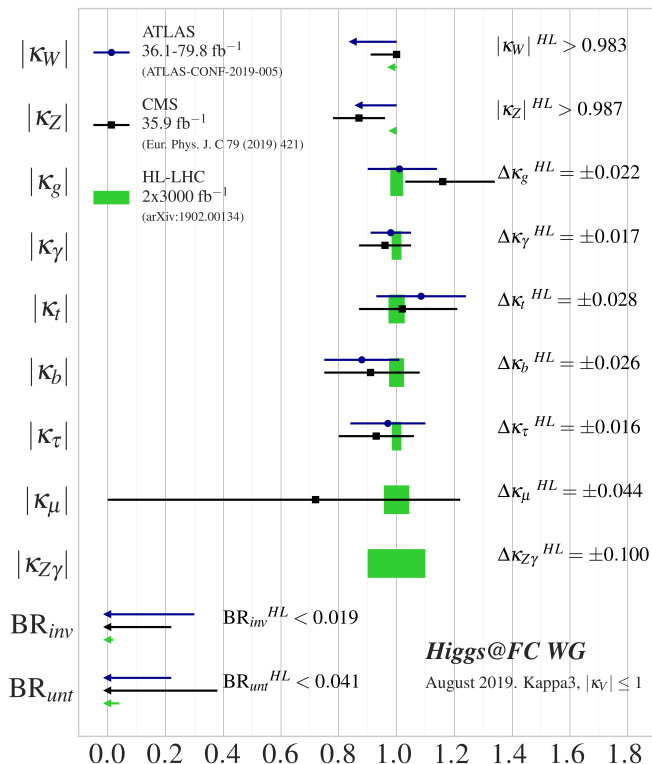
assumes $BR_{BSM}=0$



- Gauge bosons and 3rd gen fermions: $\delta \approx 10\text{-}20\%$

Current Results vs High-Luminosity LHC

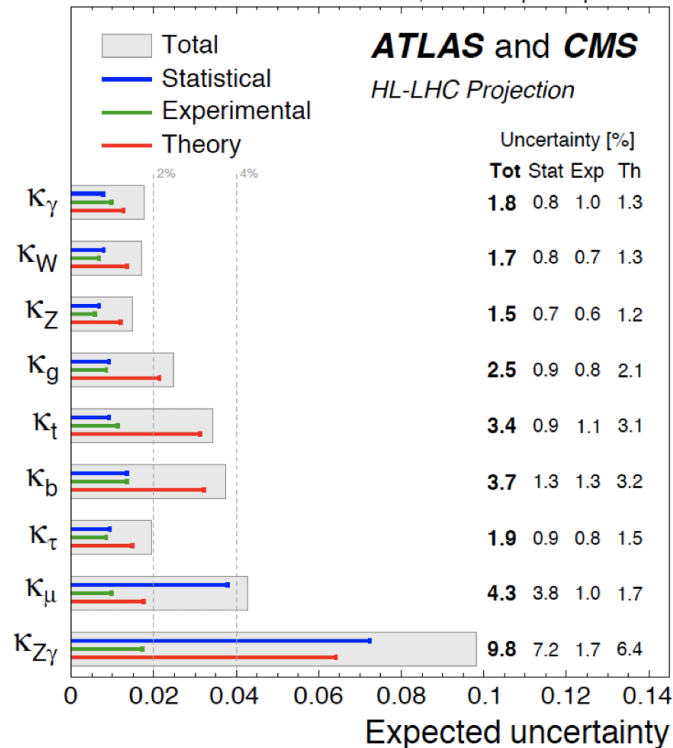
LHC now vs HL-LHC



10%

High Luminosity LHC (HL-LHC)

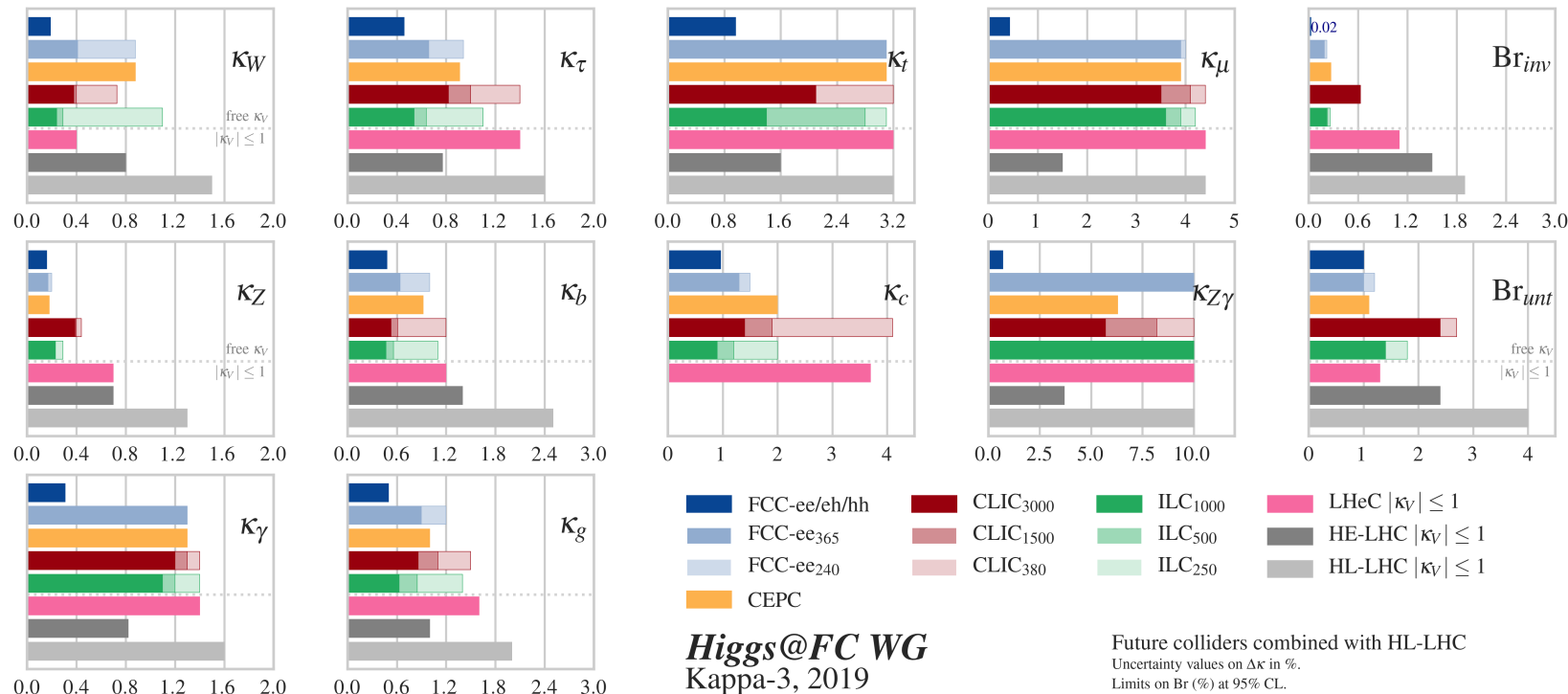
$\sqrt{s} = 14$ TeV, 3000 fb⁻¹ per experiment



10%

Comparison of Future Colliders

arXiv:1905.03764



Requires also significant improvements in theoretical calculations

of “largely” improved H couplings (EFT)

NB: number of seconds/year differs: ILC 1.6×10^7 , FCC-ee & CLIC: 1.2×10^7 , CEPC: 1.3×10^7

		Factor ≥ 2	Factor ≥ 5	Factor ≥ 10	Years from T_0
Initial run	CLIC380	9	6	4	7
	FCC-ee240	10	8	3	9
	CEPC	10	8	3	10
	ILC250	10	7	3	11
2 nd /3 rd Run ee	FCC-ee365	10	8	6	15
	CLIC1500	10	7	7	17
	HE-LHC	1	0	0	20
	ILC500	10	8	6	22
hh	CLIC3000	11	7	7	28
ee,eh & hh	FCC-ee/eh/hh	12	11	10	>50

13 quantities in total

of “largely” improved H couplings (EFT)

NB: number of seconds/year differs: ILC 1.6×10^7 , FCC-ee & CLIC: 1.2×10^7 , CEPC: 1.3×10^7

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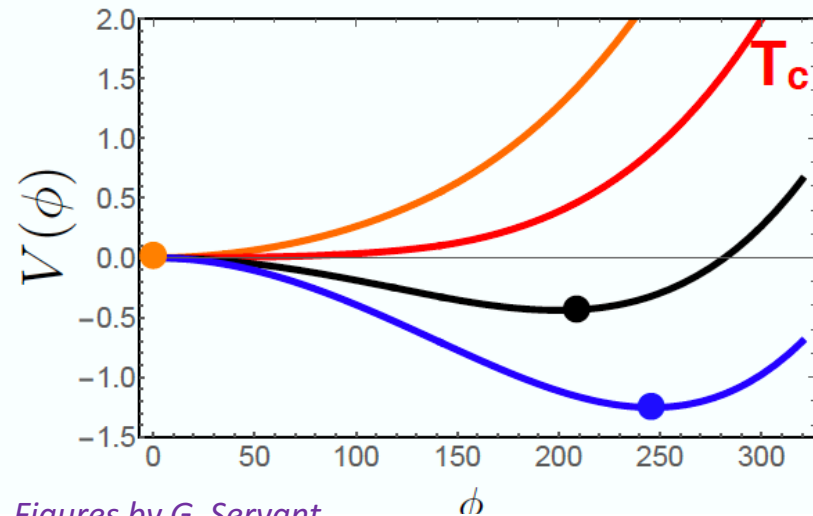
13 quantities in total

About half the couplings improved by factor 5 already in initial run

Electroweak potential

Standard Model

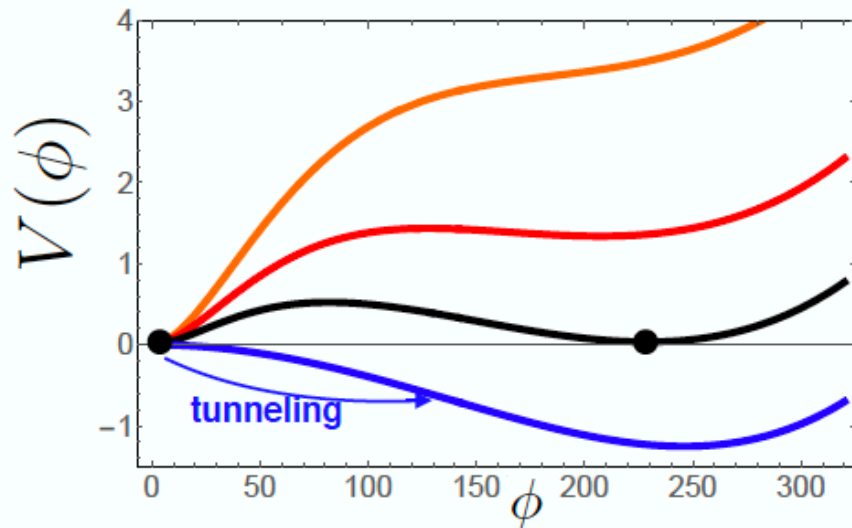
- Electroweak phase transition (EWPT) is a “smooth crossover”
- Electroweak symmetry restored for $T \geq T_c = 130$ GeV



Figures by G. Servant

Alternative idea

- Electroweak phase transition via tunneling: 1st order transition
 - Two phases co-exist
- Electroweak baryogenesis possible if strong 1st order transition $\frac{v(T_{pt})}{T_{pt}} \gtrsim 1.0$

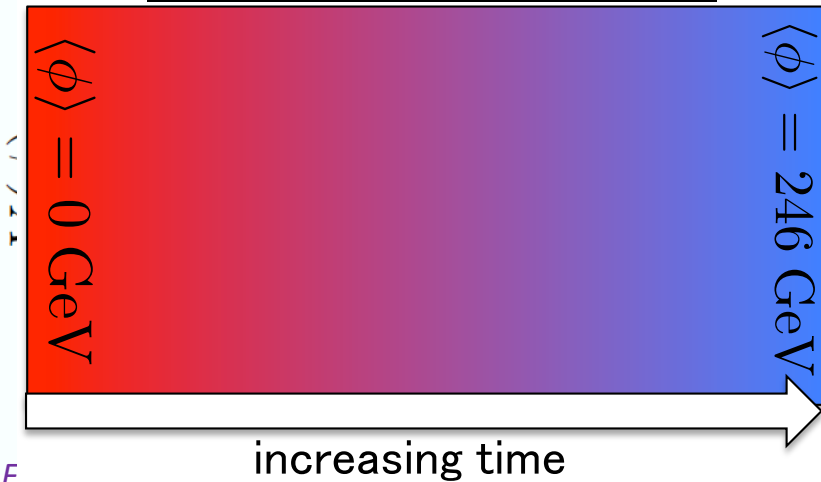


Electroweak potential

Standard Model

- Electroweak phase transition (EWPT) is a “smooth crossover”
- Electroweak symmetry restored for $T \geq T_c = 130 \text{ GeV}$

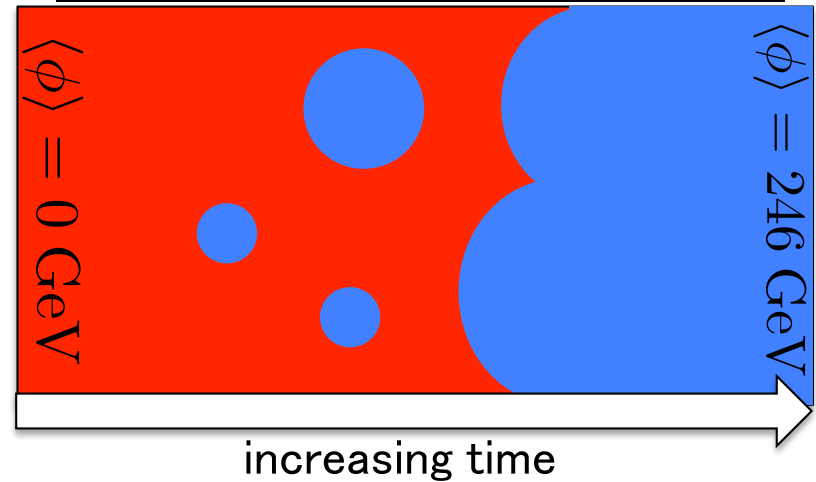
Continuous Crossover



Alternative idea

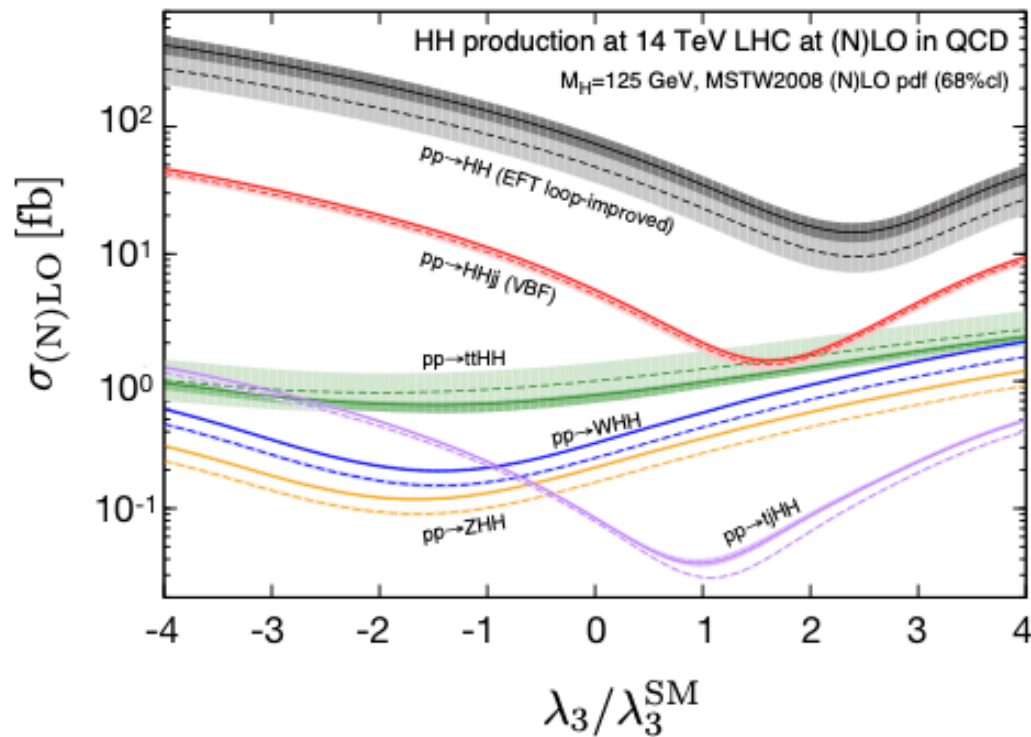
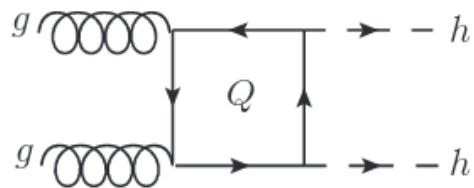
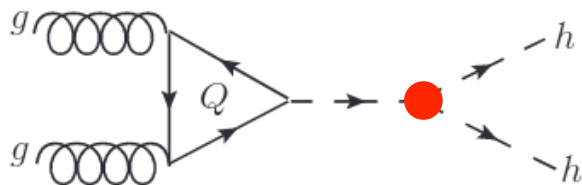
- Electroweak phase transition via tunneling: 1st order transition
 - Two phases co-exist
- Electroweak baryogenesis possible if strong 1st order transition $\frac{v(T_{\text{pt}})}{T_{\text{pt}}} \gtrsim 1.0$

First Order Phase Transition



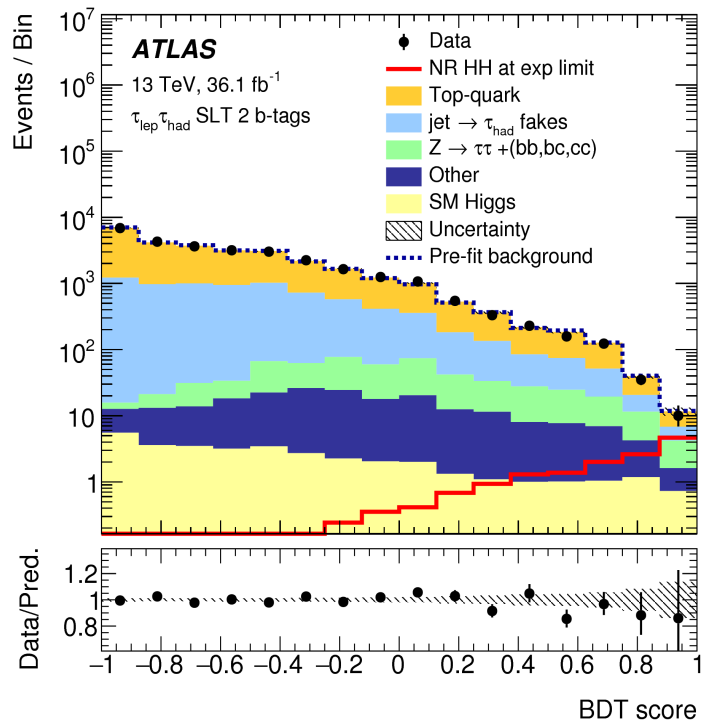
Measuring the Higgs self-coupling: LHC

Hadron collider

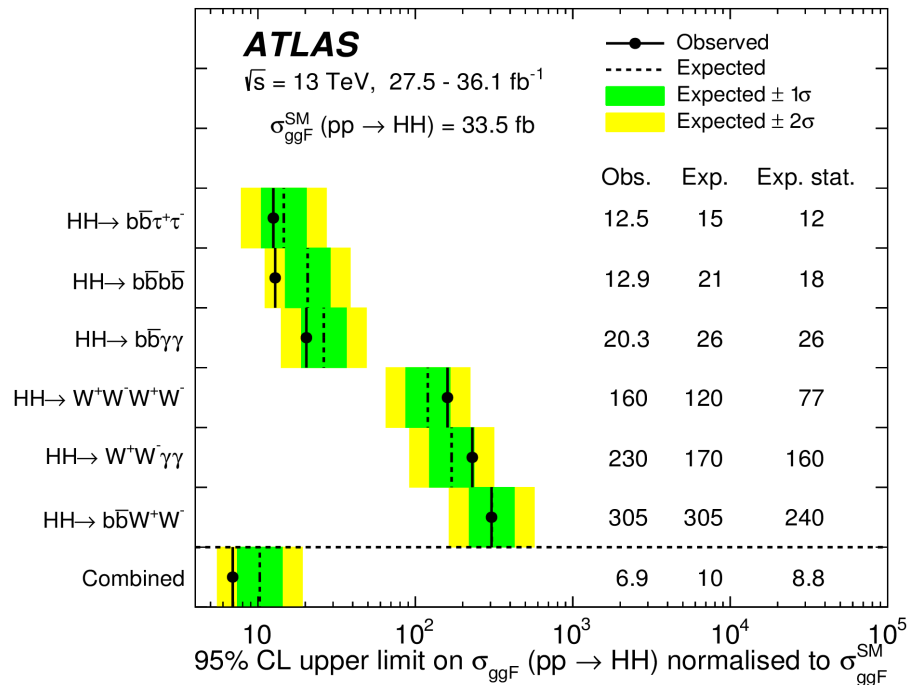


Di-Higgs Production: LHC results

HH → bbττ search result



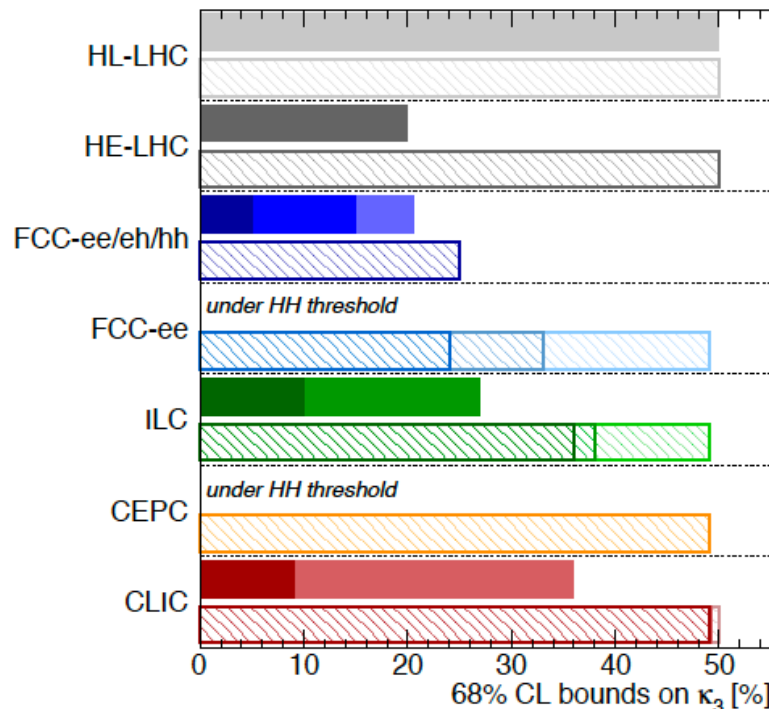
HH analyses combined



=> Upper limit at 95% CL: 6.9 x σ_{SM} => -5 < κ_λ < 12

Sensitivity to κ_λ : future colliders

Higgs@FC WG September 2019

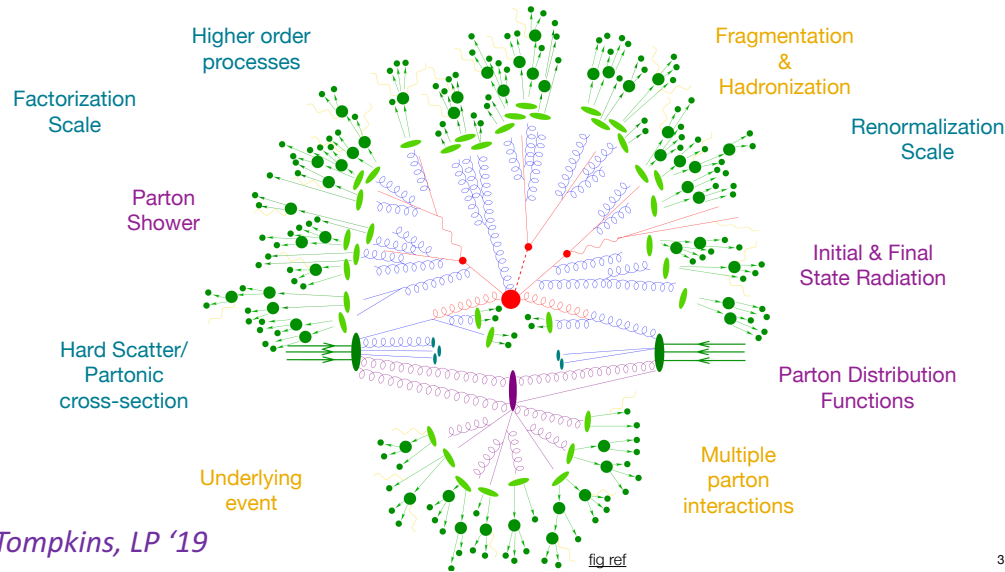


di-Higgs	single-Higgs
HL-LHC 50%	HL-LHC 50%
HE-LHC (10-20)%	HE-LHC 50%
FCC-ee/eh/hh 5%	FCC-ee/eh/hh 25%
LE-FCC 15%	LE-FCC n.a.
FCC-eh ₃₅₀₀ -17+24%	FCC-eh ₃₅₀₀ n.a.
	FCC-ee ₄₀₀ 24%
	FCC-ee ₃₆₅ 33%
	FCC-ee ₂₄₀ 49%
ILC ₁₀₀₀ 10%	ILC ₁₀₀₀ 36%
ILC ₅₀₀ 27%	ILC ₅₀₀ 38%
	ILC ₂₅₀ 49%
	CEPC 49%
CLIC ₃₀₀₀ -7%+11%	CLIC ₃₀₀₀ 49%
CLIC ₁₅₀₀ 36%	CLIC ₁₅₀₀ 49%
	CLIC ₅₀₀ 50%

All future colliders combined with HL-LHC

- HL-LHC: ~50%
 - Already relevant for EWPT!
- Future high-energy colliders: ~5-10%

The Strong Interaction



Conceptional aspects of QCD (closely related to nuclear physics)

- Long-distance phenomena (e.g. confinement)
- Collective behaviour at high temperatures or densities
- Transition of high to low energies (e.g. hadronisation)

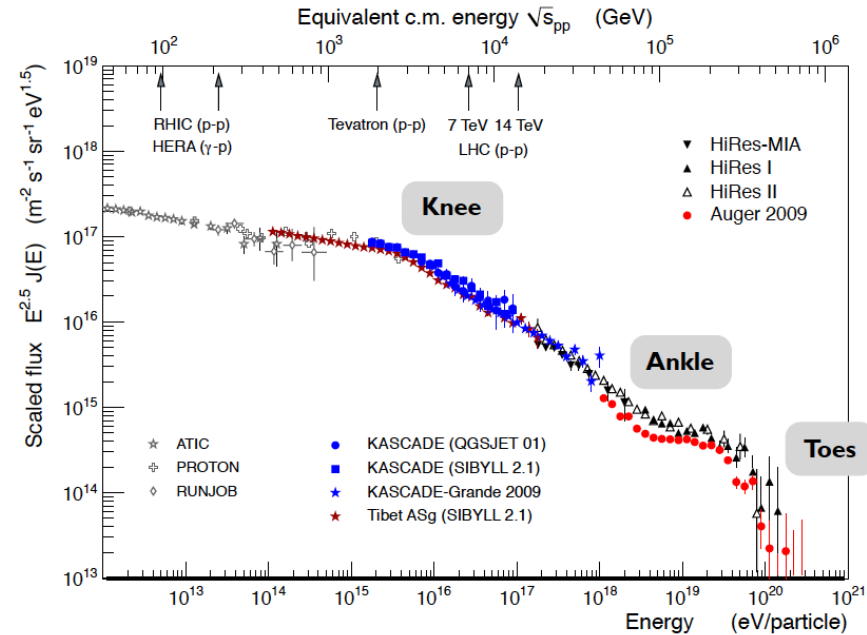
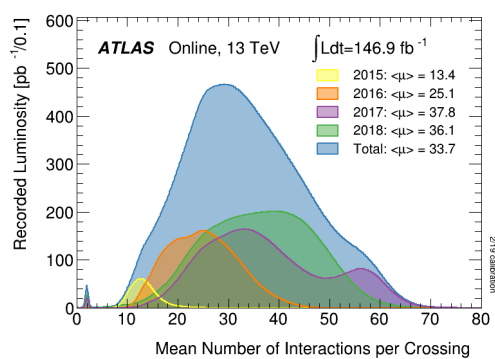
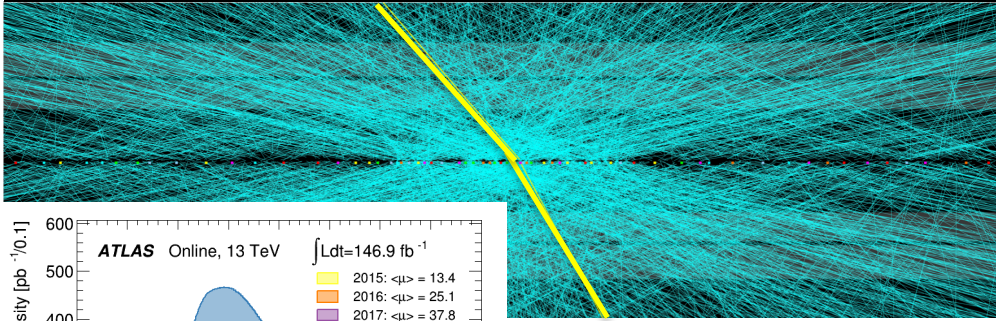
3

Practical aspects of QCD predictions

- **Soft interactions** constitute “noise” in high-pileup environment
- Parton flux in proton determined by **parton distribution functions (PDFs)**
- **Strong coupling constant** governs rate jet production at pp collider
- Precise predictions at hadron colliders require ever **higher order calculations**

Soft QCD processes

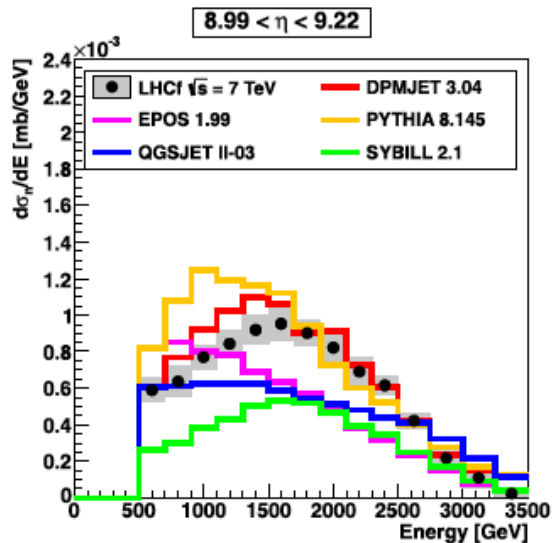
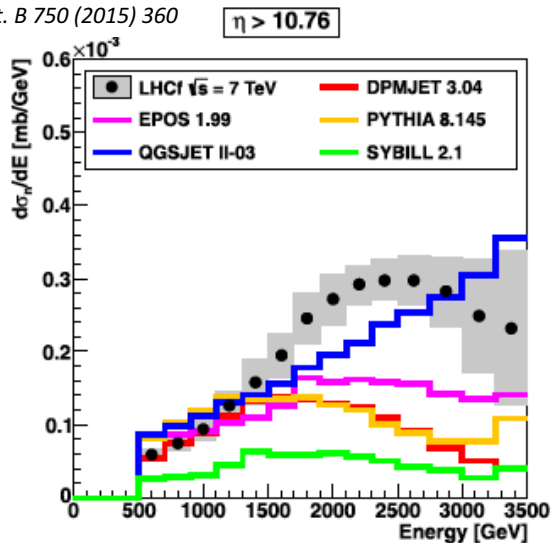
Z->ll event with 65 vertices



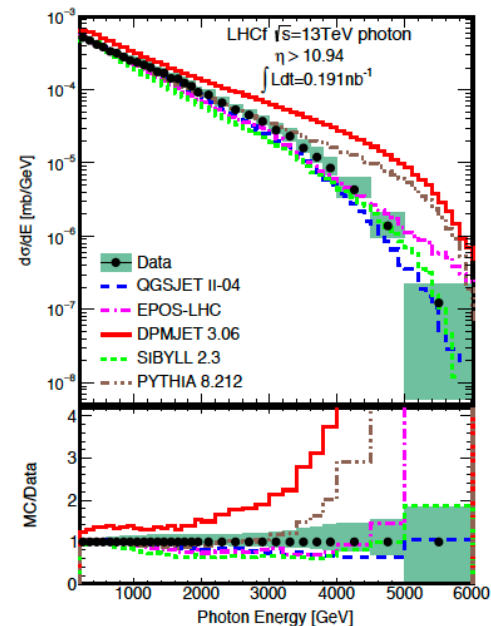
Relevant for understanding of pileup at LHC, for high energy astrophysics and for non-perturbative calculations

Neutral particles: forward n and γ production

Phys. Lett. B 750 (2015) 360

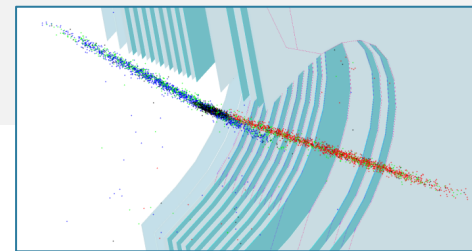


Phys. Lett. B 780 (2018) 233



- LHCf: dedicated experiment measuring production of neutral particles in forward direction
- Data measurements discriminate between different models

Anti-proton production in p-He collisions

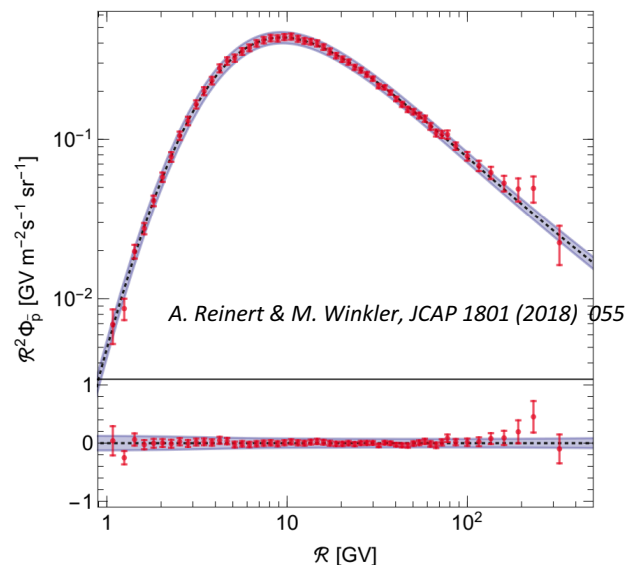


AMS and PAMELA spectrometers measure fraction of antiprotons in space

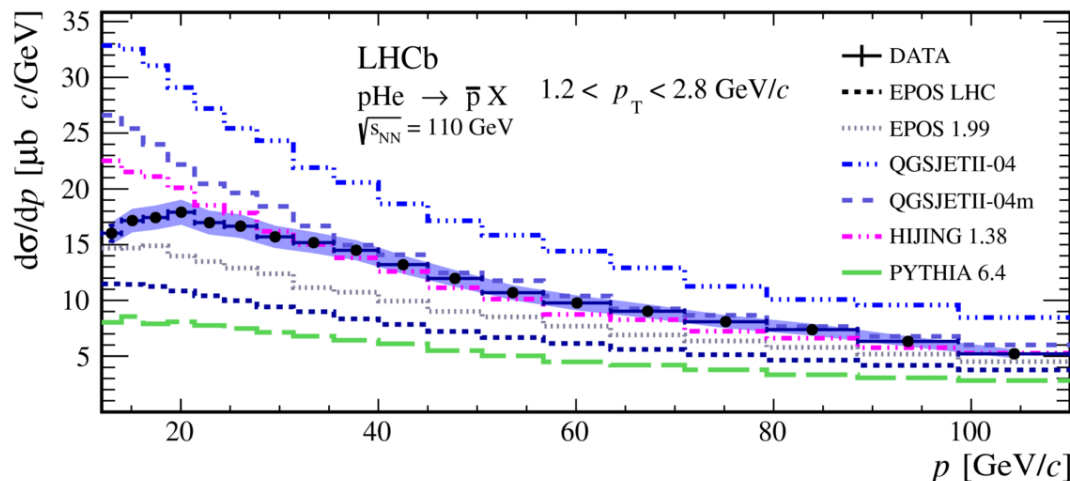
- Indirect probe for dark matter annihilation
- Uncertainties on antiproton cross sections in pp and pHe collisions important

LHCb uses fixed-target mode (p+He gas) to measure antiproton production in pHe collisions

AMS data, M. Aguilar et al., PRL 113 (2014).

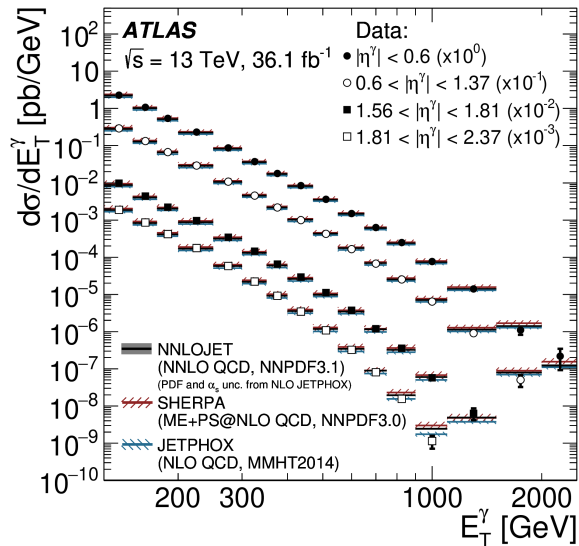


LHCb, arXiv:1808.06127



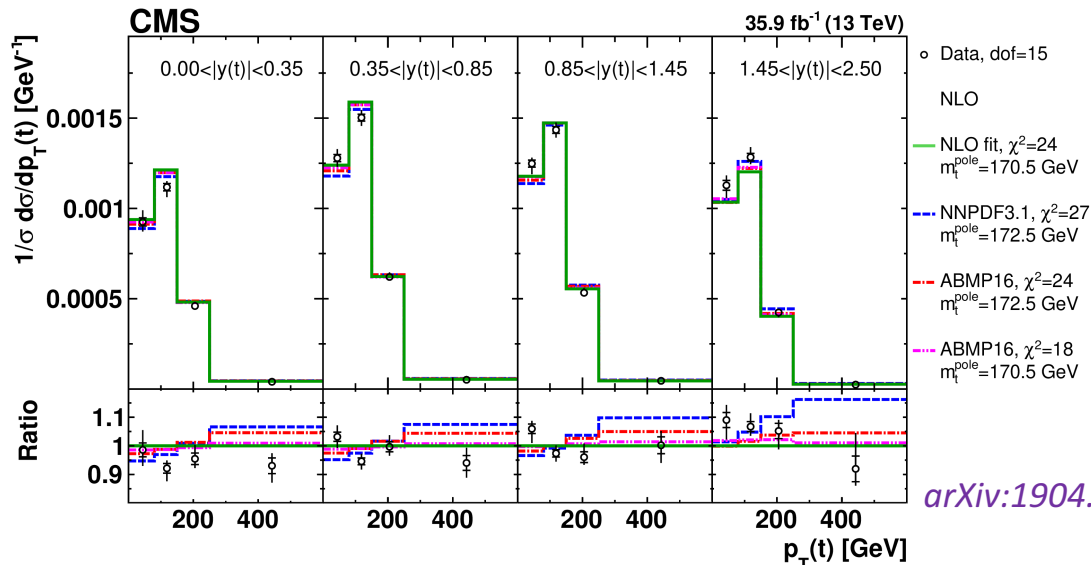
Perturbative QCD at the LHC

$pp \rightarrow \gamma + \text{jets}$



arXiv:1908.02746

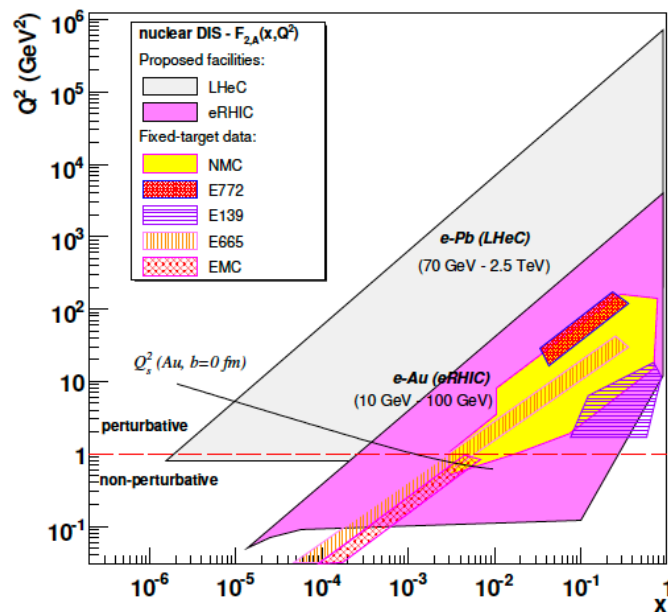
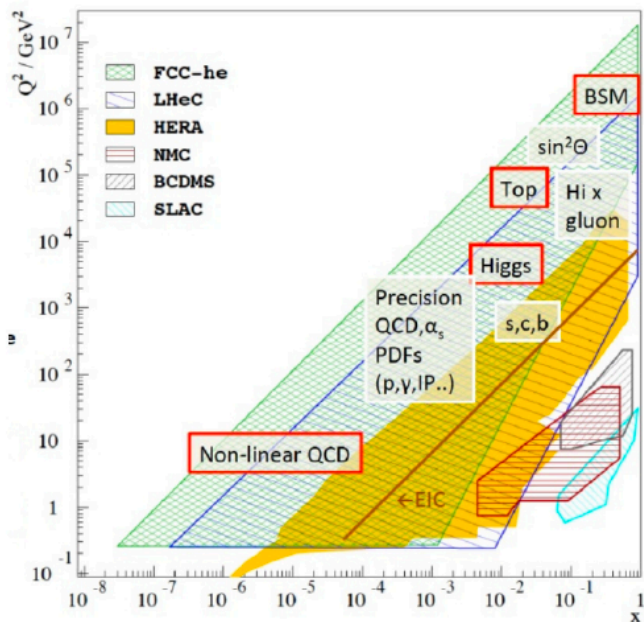
$pp \rightarrow t\bar{t} + X$



arXiv:1904.05237

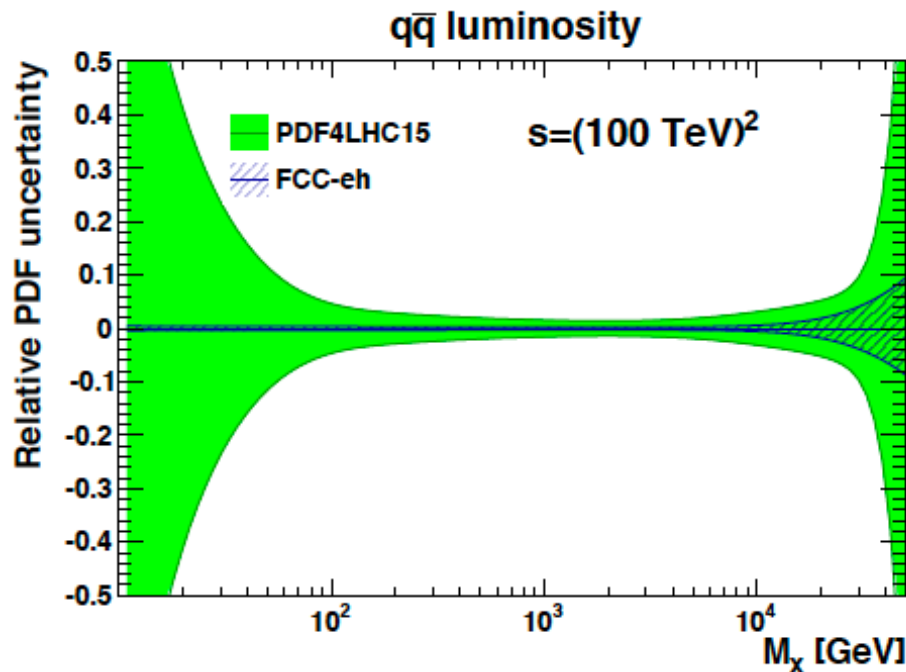
- Multi-differential measurements of QCD production processes
 - Test MC generators and higher order QCD calculations (NNLO+beyond)
 - Determine strong coupling constant α_s and/or parton distribution functions

Parton distribution functions



- Full exploitation of pp data requires precise knowledge of PDFs
 - New ep colliders (LHeC, EIC, FCC-eh) can provide important input

Parton distribution functions

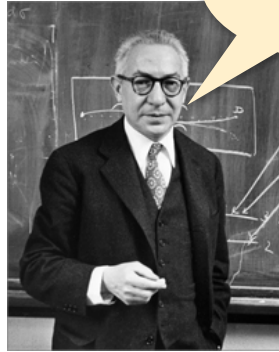


- Full exploitation of pp data requires precise knowledge of PDFs
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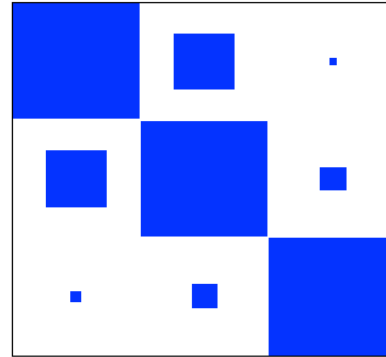
Flavour Physics

I. Raby

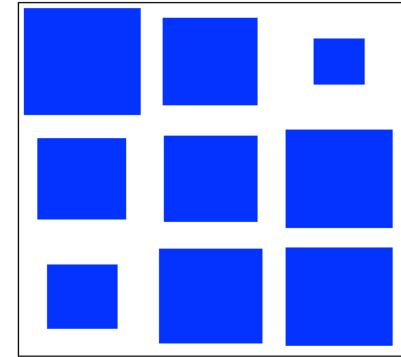
Who ordered that?



Quark mixing



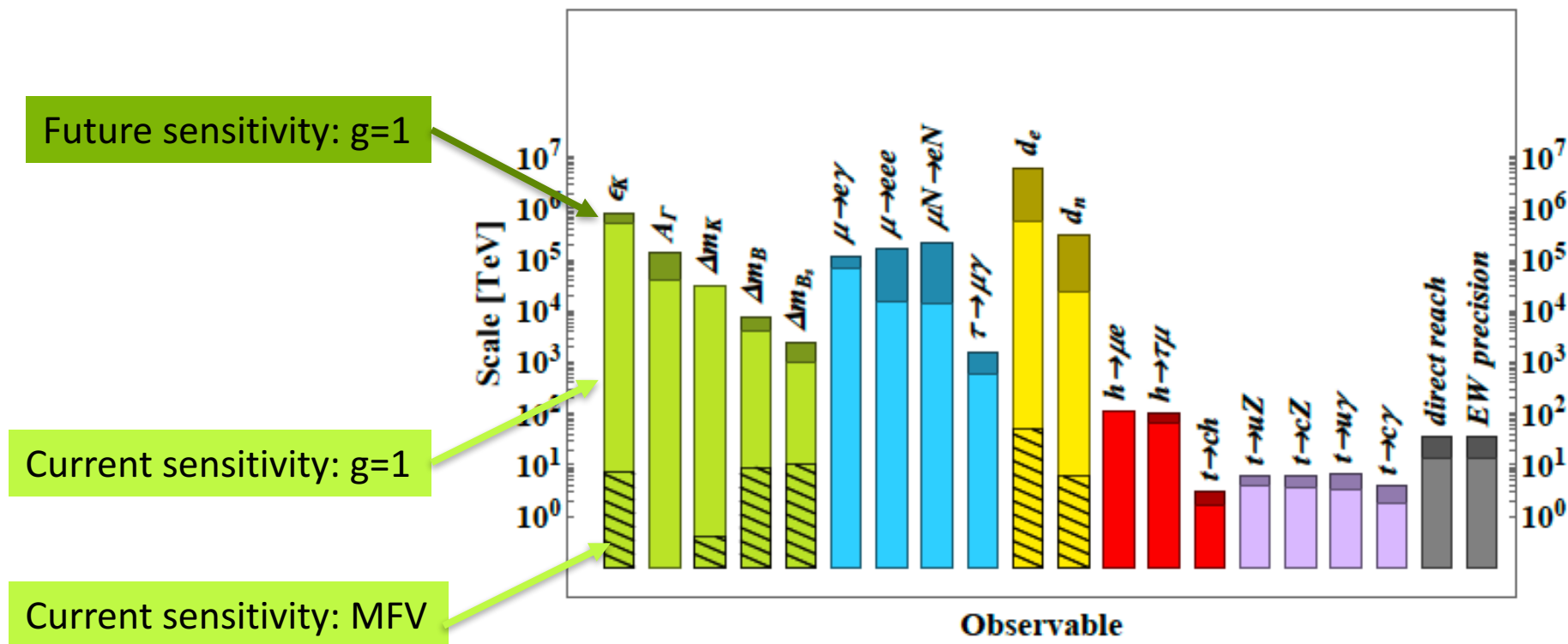
Neutrino mixing



Leptons Quarks	I		
	II		
	III		
	Die Generationen der Materie		
Leptons Quarks	u up	c charm	t top
	d down	s strange	b bottom
	ν_e e- Neutrino	ν_μ μ - Neutrino	ν_τ τ - Neutrino
	e electron	μ muon	τ tau

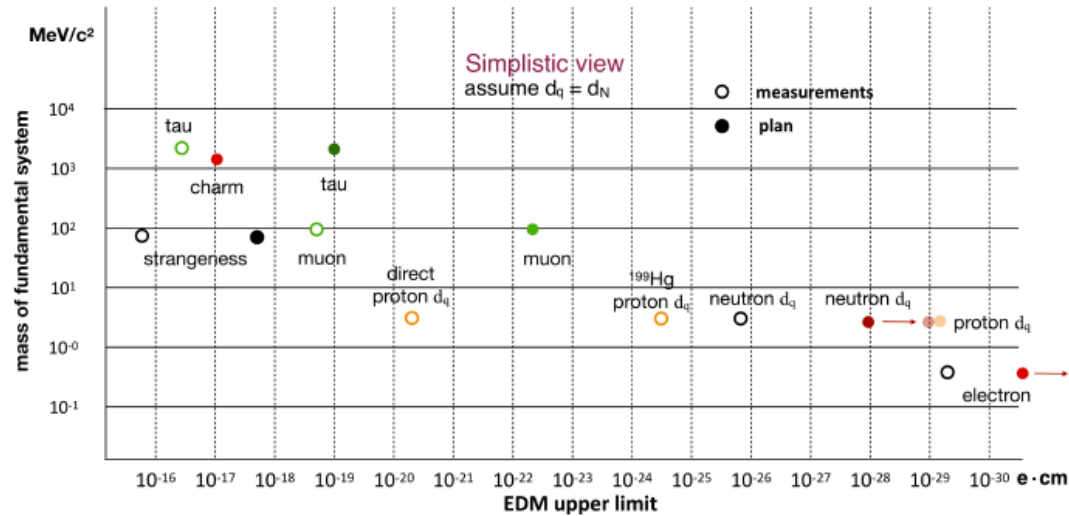
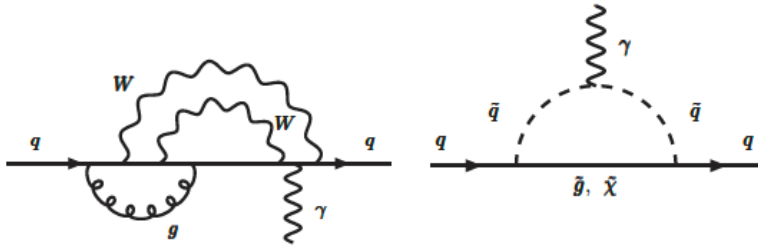
- Flavour puzzle today (adapted from Y. Nir):
 - Why is there so much structure in the quark sector?
 - Why is there no structure in the neutrino sector?
 - Why are there no flavor-changing neutral currents?
- What is source of CP violation explaining lack of anti-matter?
- Flavour is also excellent probe of high-scale physics
 - New physics tends to break accidental symmetries of SM

Flavour: New Physics Sensitivities



Electric Dipole Moment

“Has killed more SUSY models than anything else” (*I. Hinchliffe*)

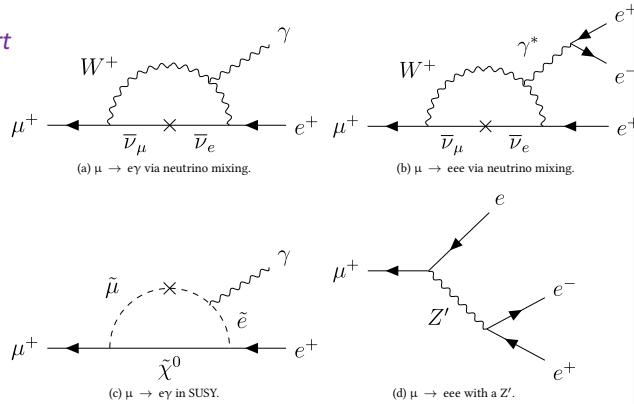


- Current limits: 3.6×10^{-26} for neutron, 1.1×10^{-29} for electron
 - Lepton and quark EDMs are complementary tests of new physics
- Advancements planned in future experiments: factors ~ 10 -1000
- Observation would be clear evidence for new physics

Lepton Flavour Violation: $\mu \rightarrow e$ and $\tau \rightarrow \ell$

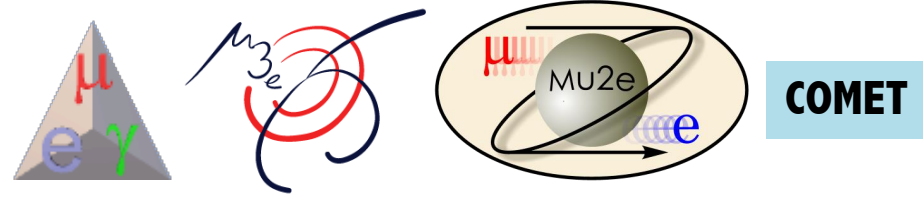
Figures from A-K Perrevoort

SM: $\text{BR} < 10^{-54}$



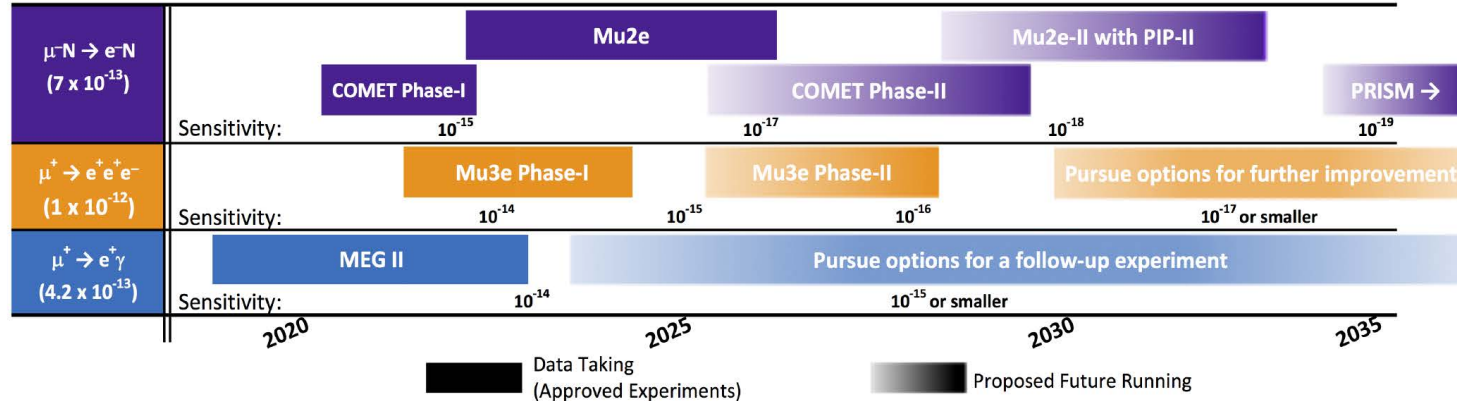
BSM

Several dedicated experiments coming online in near future for $\mu \rightarrow e$ decay or transition in Europe, US and Japan

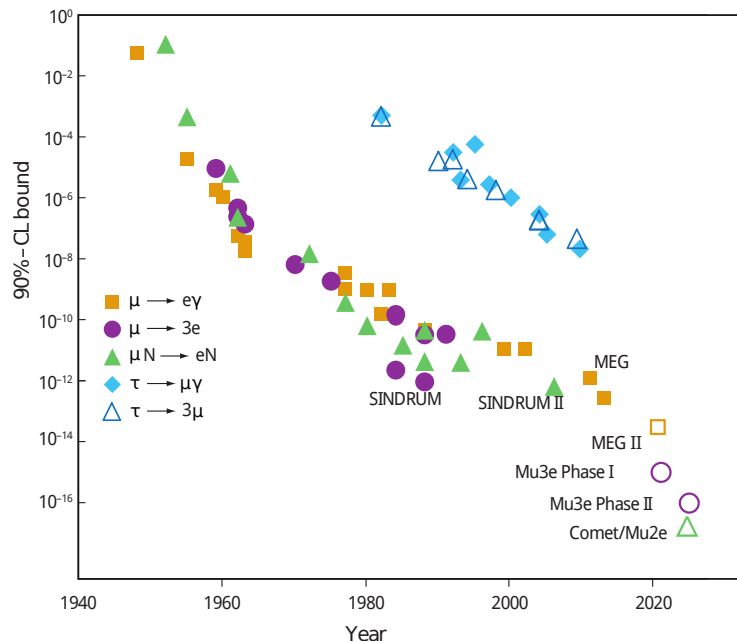


Searches for Charged-Lepton Flavor Violation in Experiments using Intense Muon Beams

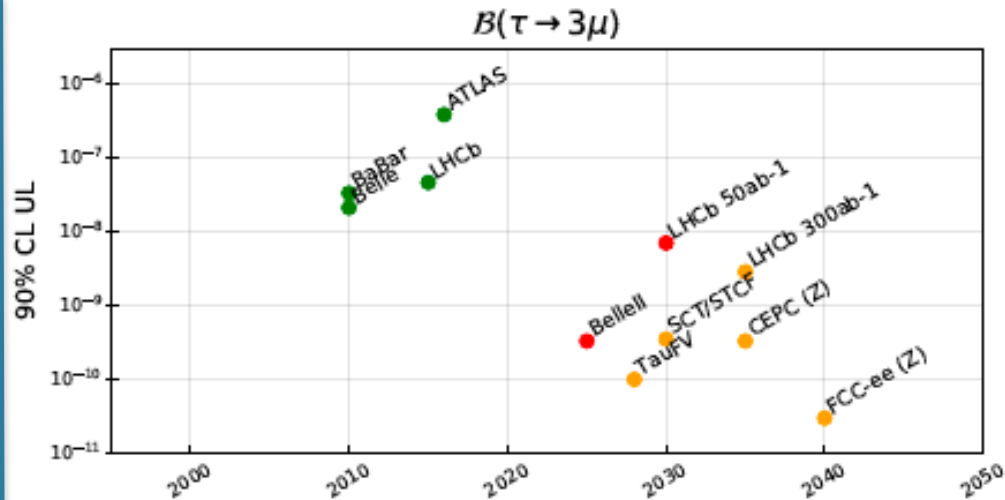
1812.06540



Lepton Flavour Violation: $\mu \rightarrow eX$ and $\tau \rightarrow \ell X$



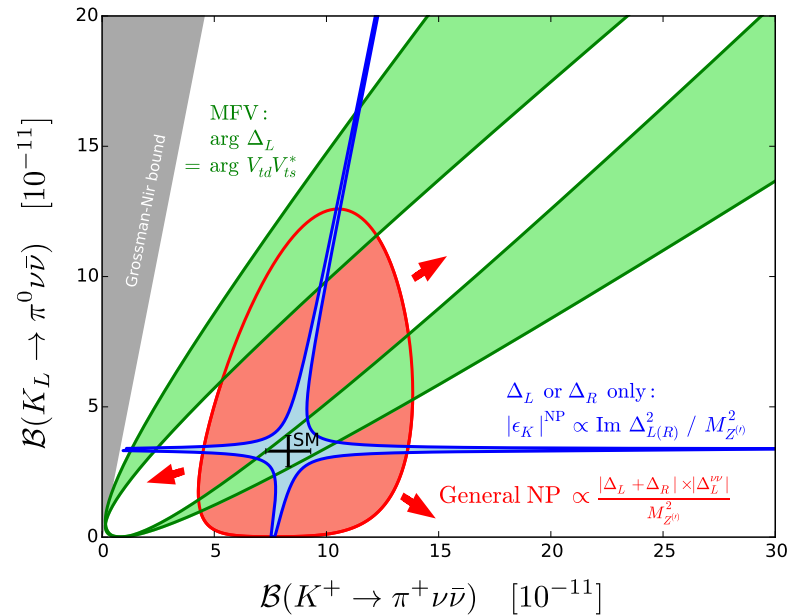
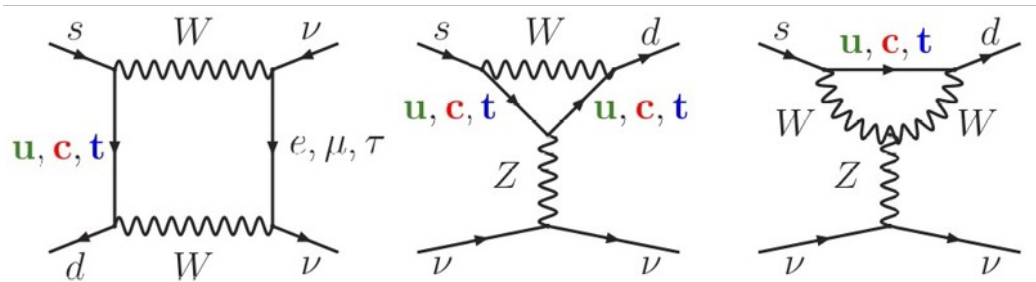
Muons: Expect to improve sensitivity by factor 10^4 !



Taus:

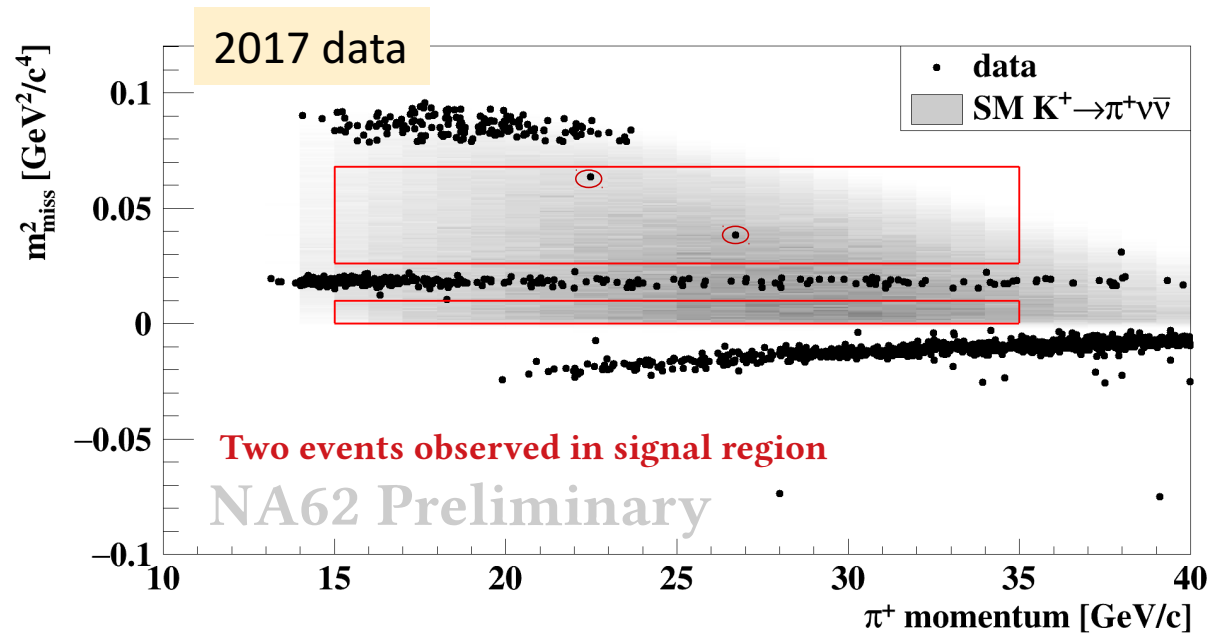
- Next decade (LHCb, Belle-II): $\sim 10^2$
- Beyond next decade (Tera-Z): $\sim 10^3$

Rare Kaon decays



- Charged and neutral kaon decays important probe of new physics
- Major focus: $K^+ \rightarrow \pi^+ \nu \bar{\nu}$ (NA62) and $K_L^0 \rightarrow \pi^0 \nu \bar{\nu}$ (KOTO, KLEVER)

Recent Result by NA62



2016+2017 data

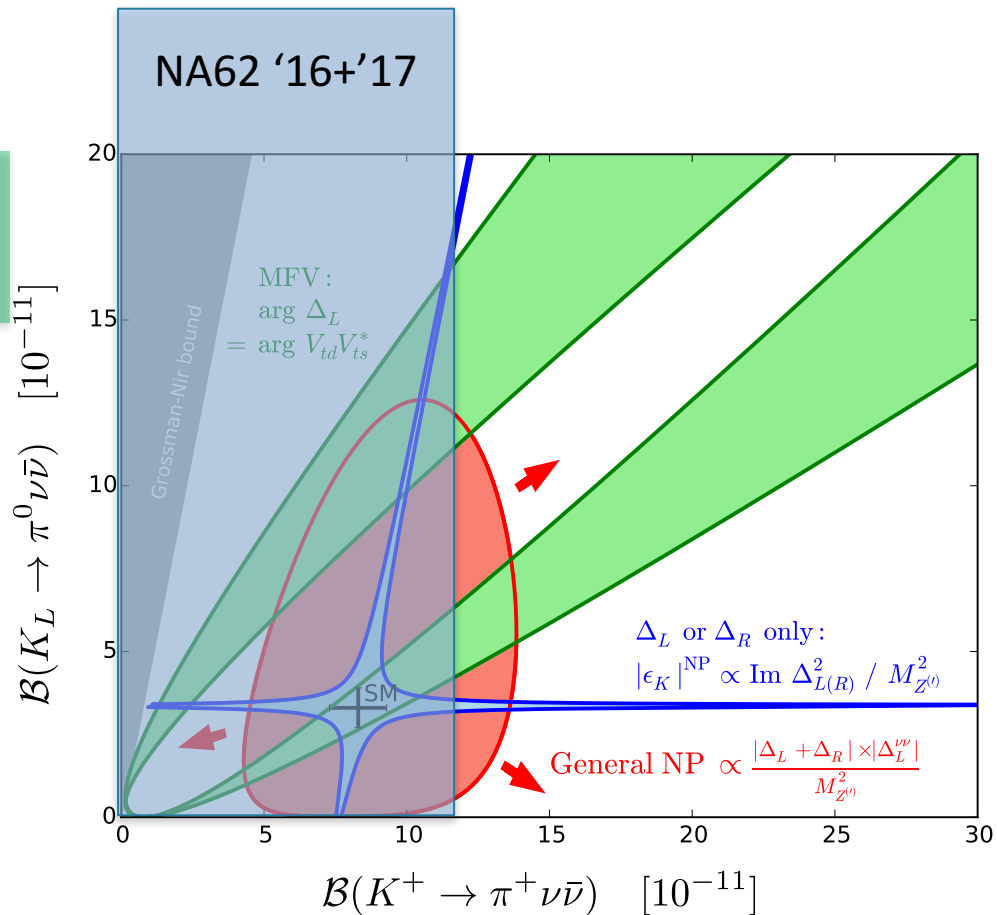
- $N_{\text{data}} = 3$
- $N_{\text{background}} = 1.65 \pm 0.31$

$$BR(K^+ \rightarrow \pi^+ \nu \nu) < 1.85 \times 10^{-10} \text{ @ 90 \% CL}$$

$$BR(K^+ \rightarrow \pi^+ \nu \nu) = 0.47^{+0.72}_{-0.47} \times 10^{-10}$$

Future Prospects

KOTO '15 limit:
 $\text{BR} < 3 \times 10^{-9}$

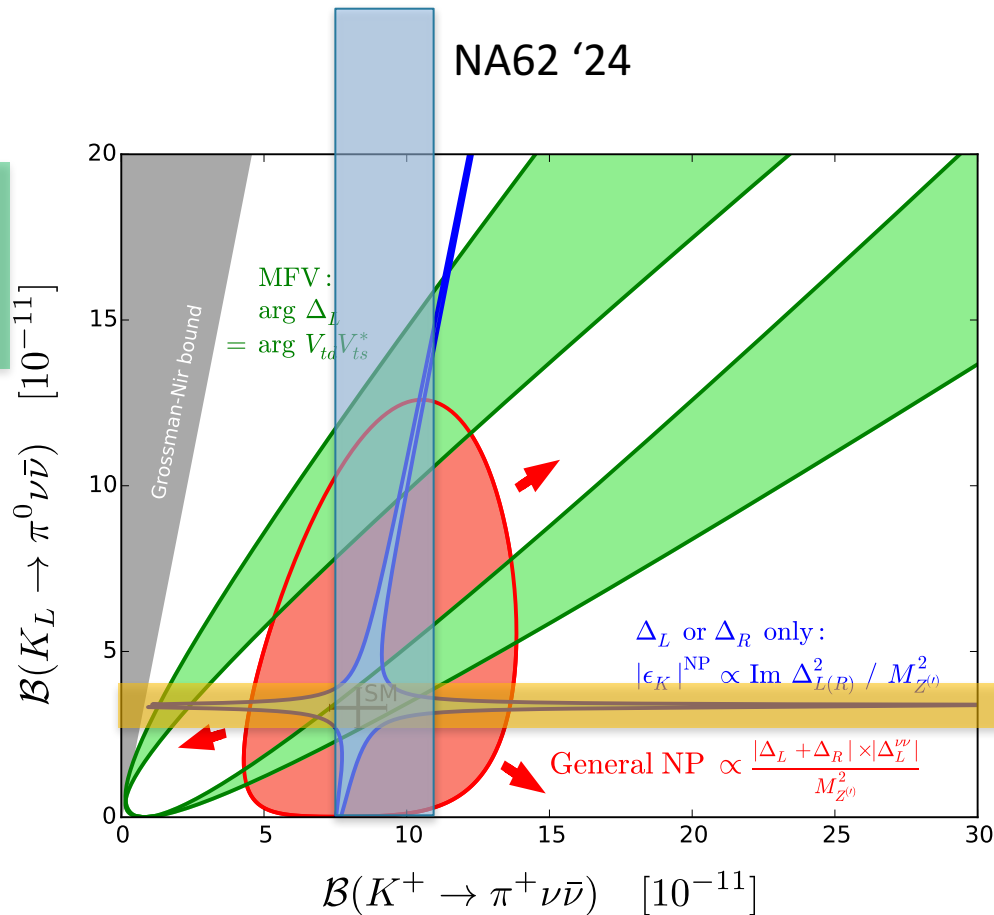
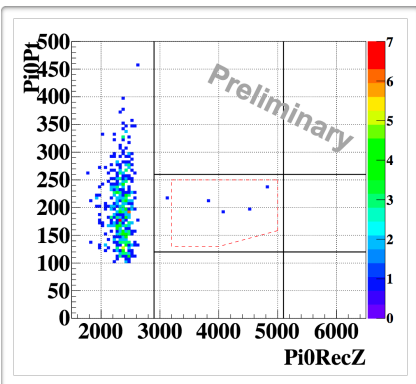


Future Prospects

KOTO '16-'18

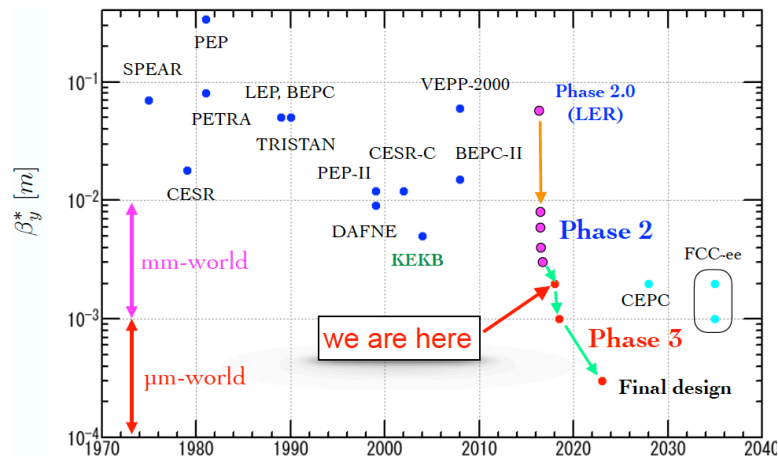
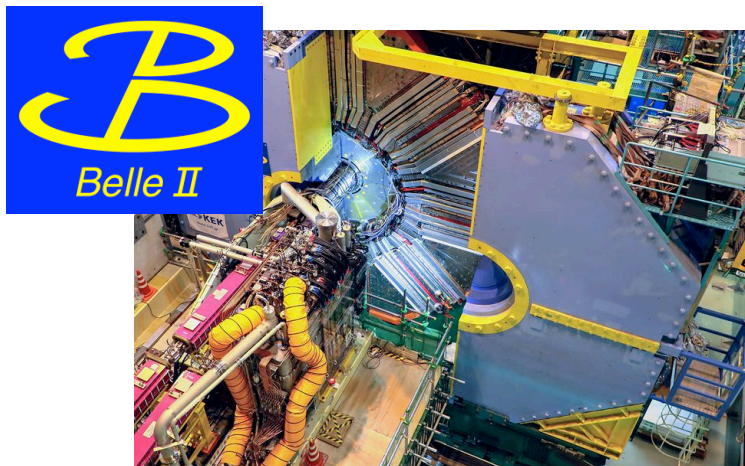
- sensitivity 7×10^{-10}
- 4 events obs.

KOTO

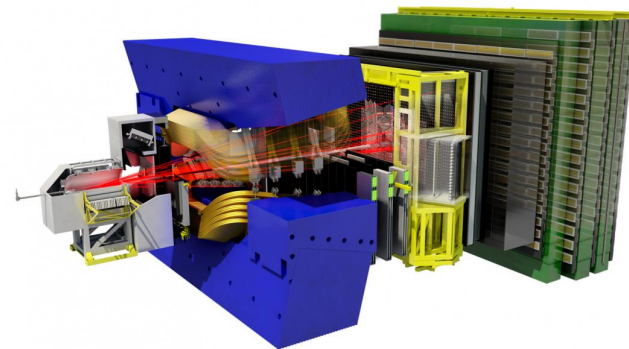


KOTO II, KLEVER

Heavy Flavour



- Major improvements in the next decade expected
 - Super-KEKB and Belle-II
 - Data taking started: towards lumi goal
 - LHCb detector upgrade



Anomalies in flavour physics

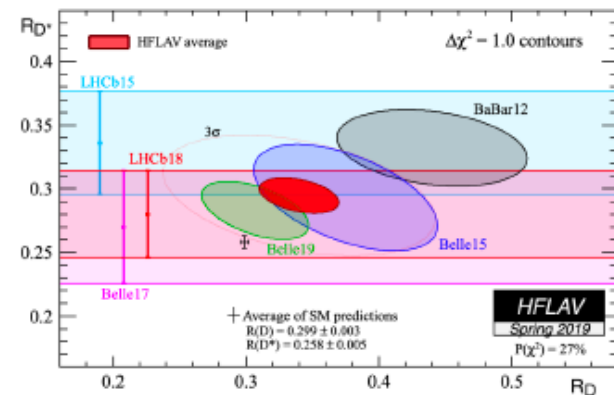
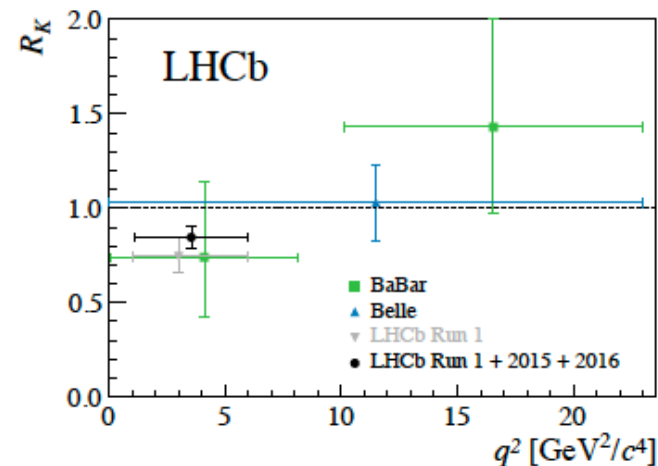
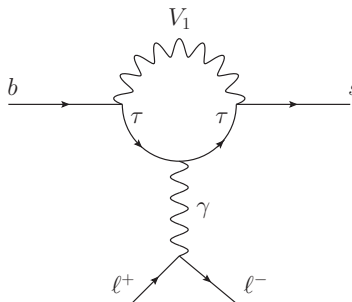
- Lepton flavour violation probes in B-decays

- Define:

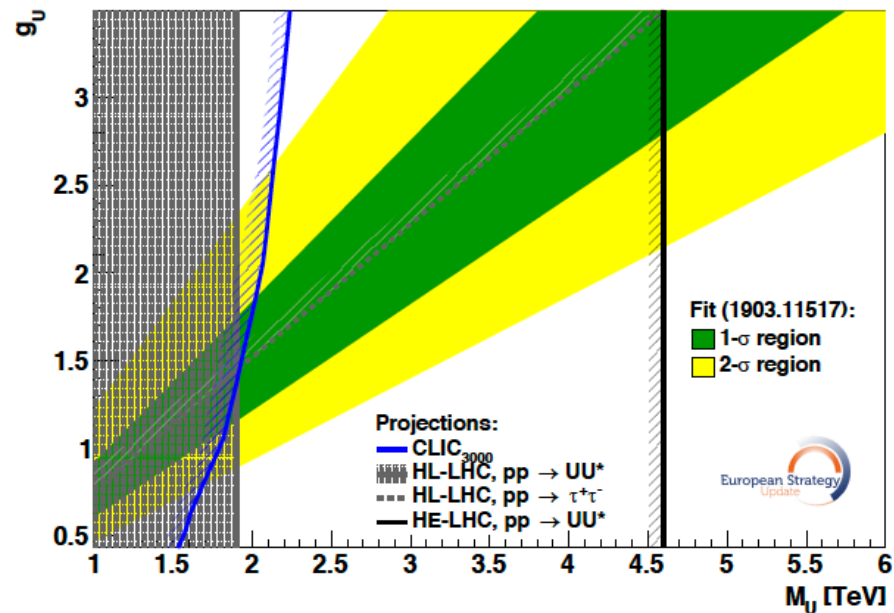
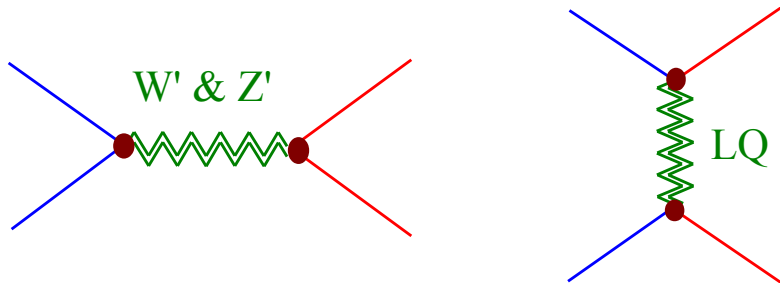
$$R_H = \frac{\int_{q_{\min}^2}^{q_{\max}^2} \frac{d\Gamma[B \rightarrow H\mu^+\mu^-]}{dq^2} dq^2}{\int_{q_{\min}^2}^{q_{\max}^2} \frac{d\Gamma[B \rightarrow He^+e^-]}{dq^2} dq^2}$$

- Several BSM explanations proposed, e.g.

- Leptoquarks
- Z' with LFV couplings



Is flavour anomaly explained by leptoquark or Z' ?

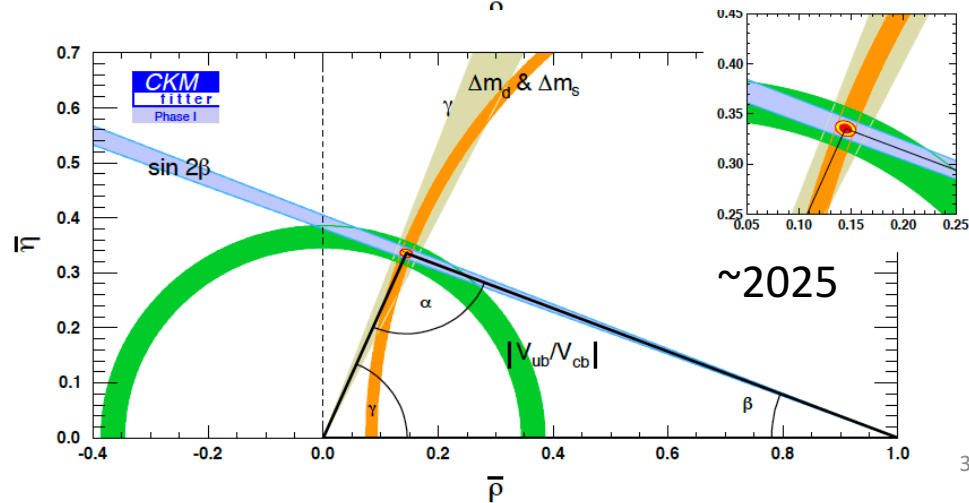
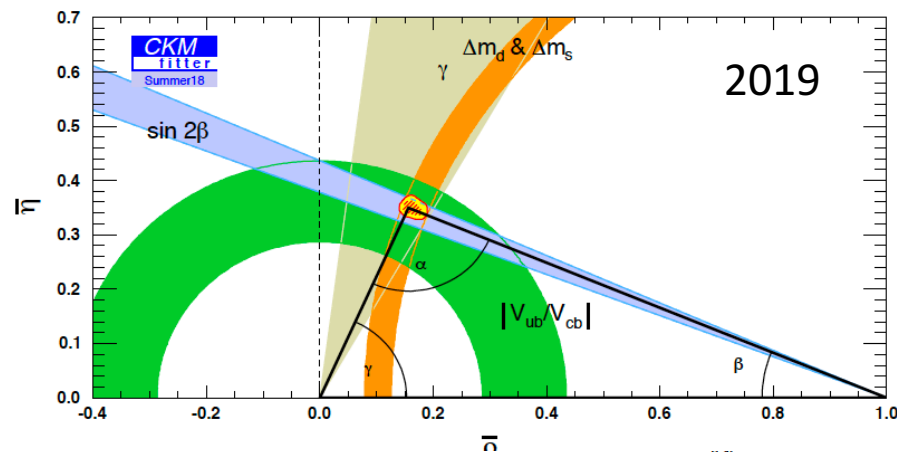


Complementary: direct searches at hadron colliders

CKM Unitarity Triangle

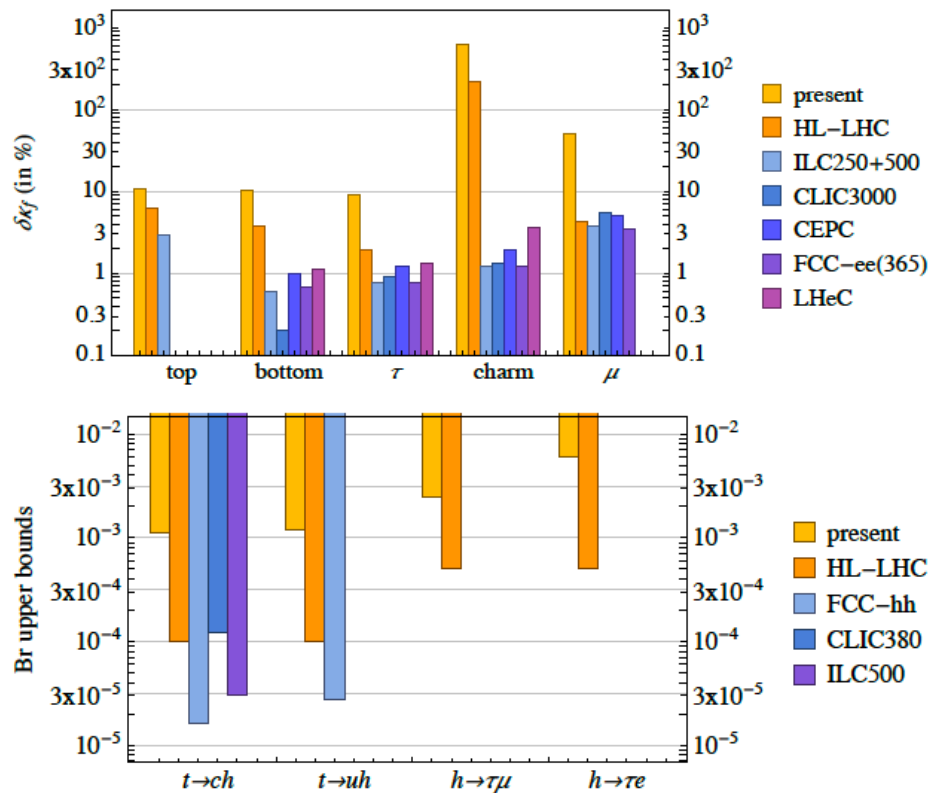
Expect huge improvements in parameters of CKM unitarity triangle by ~2025

- Even higher precision expected by ~2035



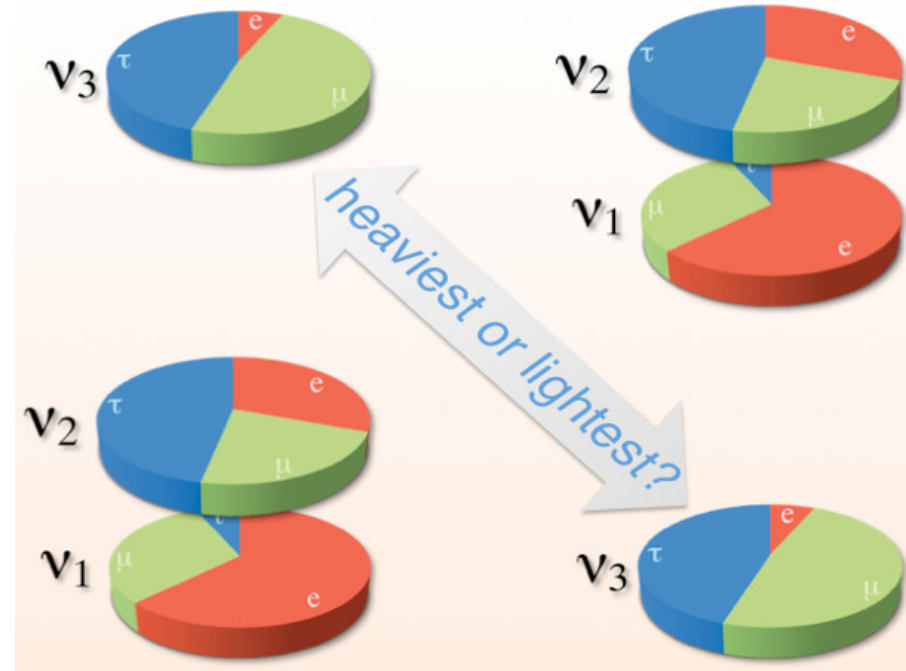
Can Higgs tell us about Flavour?

- Higgs is the only boson that cares about flavour!
 - Could it tell us something?
- Fermion couplings
 - 3rd generation observed
 - Now: 10% precision
 - Future colliders: ~1% precision
 - 2nd generation:
 - Muons: HL-LHC
 - Charm: requires new collider
 - 1st generation: very difficult
 - LFV couplings
 - HL-LHC sensitivity 10^{-3} - 10^{-4}

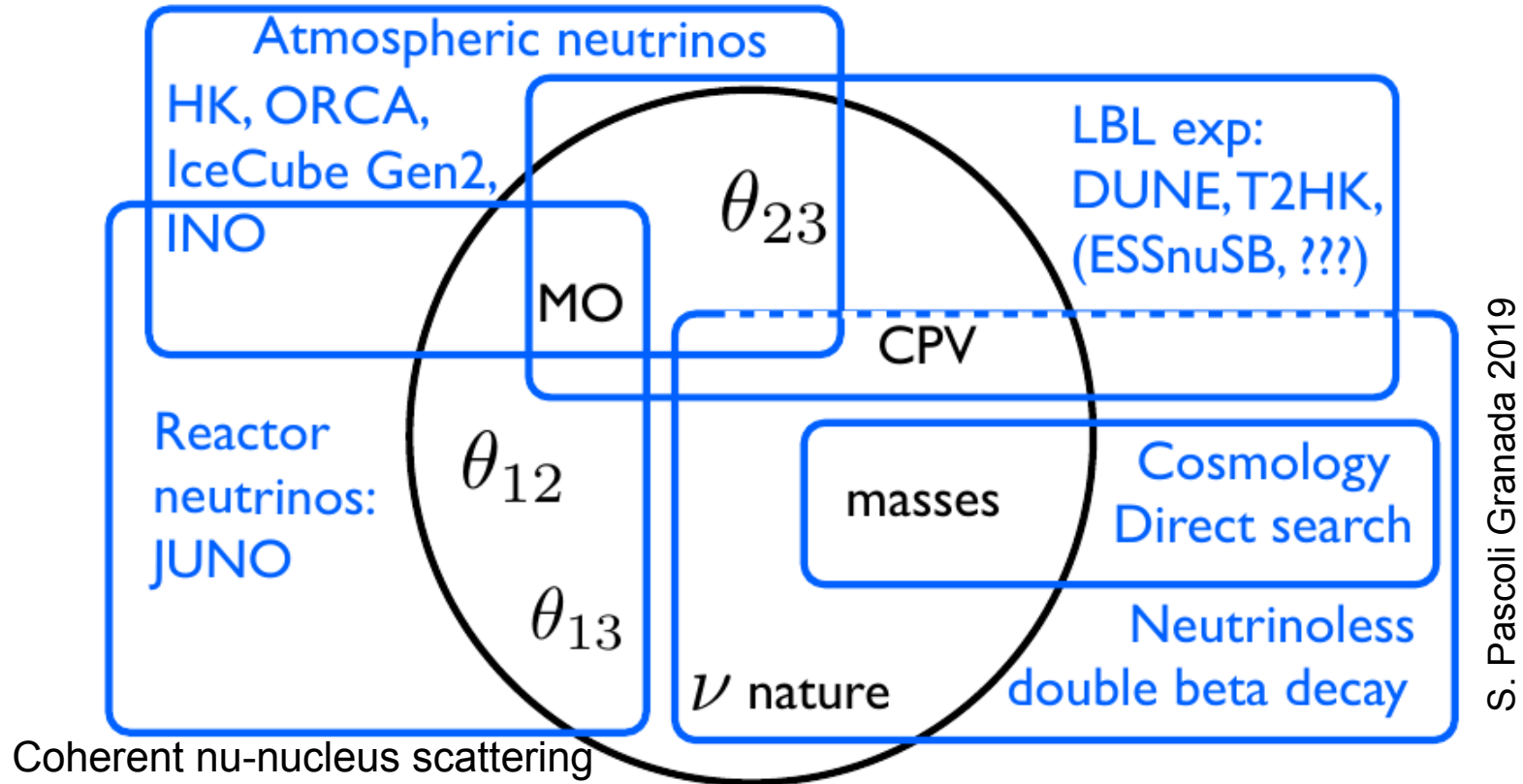


Neutrinos: Many Questions

- What is the mass ordering?
- Is CP violated?
- What is the absolute mass scale?
- Are neutrinos Dirac or Majorana particles?
- Are there heavy neutrinos?
- ...



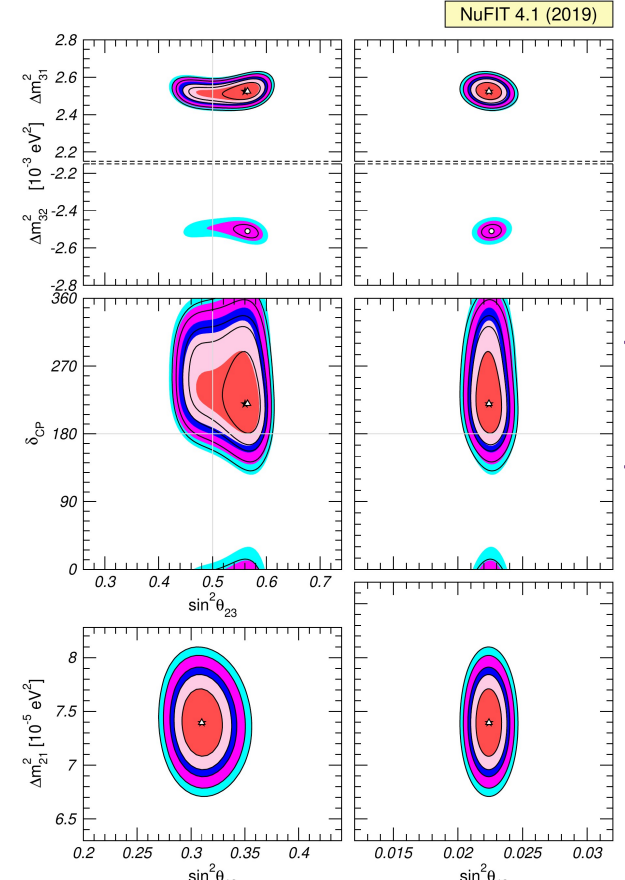
Tackling the neutrino sector experimentally



Fit to World Neutrino data

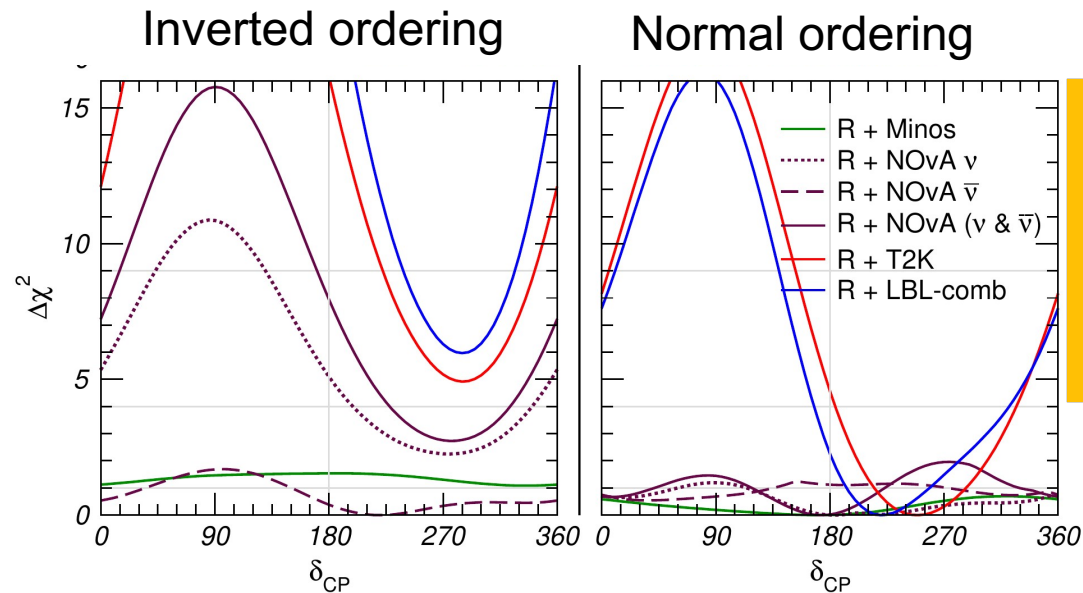
6 free parameters in fit

		Normal Ordering (best fit)		Inverted Ordering ($\Delta\chi^2 = 10.4$)	
		bfp $\pm 1\sigma$	3σ range	bfp $\pm 1\sigma$	3σ range
with SK atmospheric data	$\sin^2 \theta_{12}$	$0.310^{+0.013}_{-0.012}$	$0.275 \rightarrow 0.350$	$0.310^{+0.013}_{-0.012}$	$0.275 \rightarrow 0.350$
	$\theta_{12}/^\circ$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.75}$	$31.61 \rightarrow 36.27$
	$\sin^2 \theta_{23}$	$0.563^{+0.018}_{-0.024}$	$0.433 \rightarrow 0.609$	$0.565^{+0.017}_{-0.022}$	$0.436 \rightarrow 0.610$
	$\theta_{23}/^\circ$	$48.6^{+1.0}_{-1.4}$	$41.1 \rightarrow 51.3$	$48.8^{+1.0}_{-1.2}$	$41.4 \rightarrow 51.3$
	$\sin^2 \theta_{13}$	$0.02237^{+0.00066}_{-0.00065}$	$0.02044 \rightarrow 0.02435$	$0.02259^{+0.00065}_{-0.00065}$	$0.02064 \rightarrow 0.02457$
	$\theta_{13}/^\circ$	$8.60^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.98$	$8.64^{+0.12}_{-0.13}$	$8.26 \rightarrow 9.02$
	$\delta_{CP}/^\circ$	221^{+39}_{-28}	$144 \rightarrow 357$	282^{+23}_{-25}	$205 \rightarrow 348$
	$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$
	$\frac{\Delta m_{3\ell}^2}{10^{-3} \text{ eV}^2}$	$+2.528^{+0.029}_{-0.031}$	$+2.436 \rightarrow +2.618$	$-2.510^{+0.030}_{-0.031}$	$-2.601 \rightarrow -2.419$



I. Esteban et al., see 1811.05487

CP violation and Mass Ordering



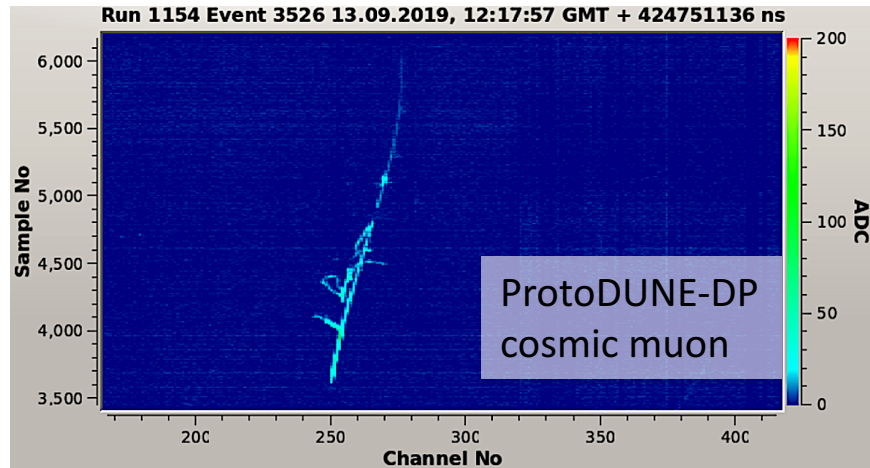
NuFit 4.1 (2019)

- Mass ordering
 - Combined T2K & Nova data prefer normal ordering ($\sim 3\sigma$)
- CP violation:
 - CP violation ($\delta_{CP} \neq 180^\circ$) favoured by T2K data ($\sim 2\sigma$)
 - NOVA data show no preference

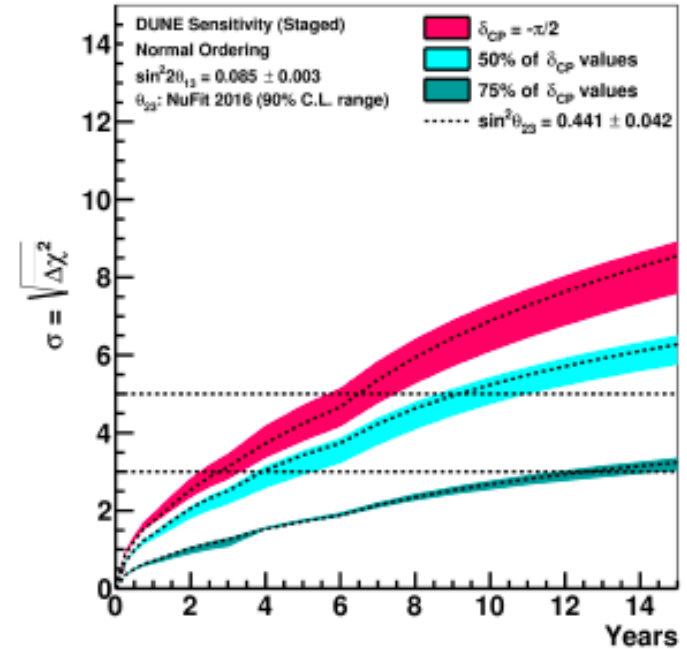
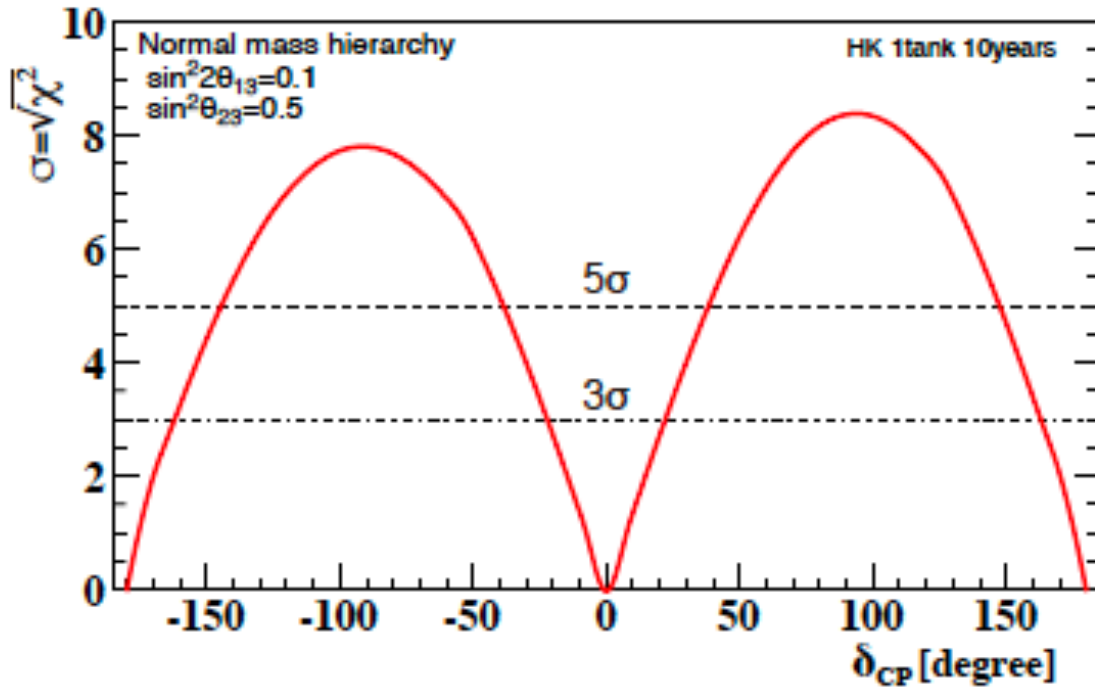
Construction of future Long Baseline Experiments



- CERN neutrino platform plays pivotal role
 - Test & understand protoDUNE single- and dual-phase detectors
 - Critical input for TDRs
- Hyper-K construction starting in 2020

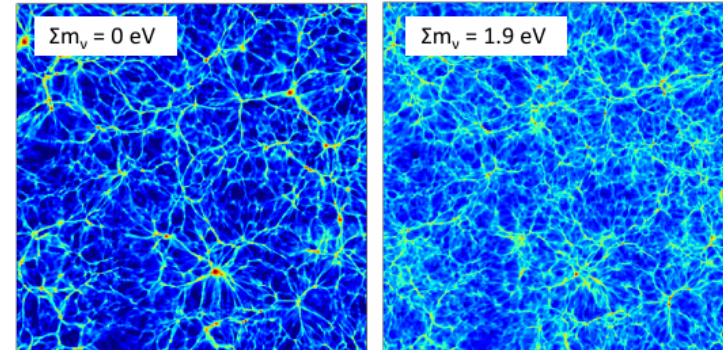
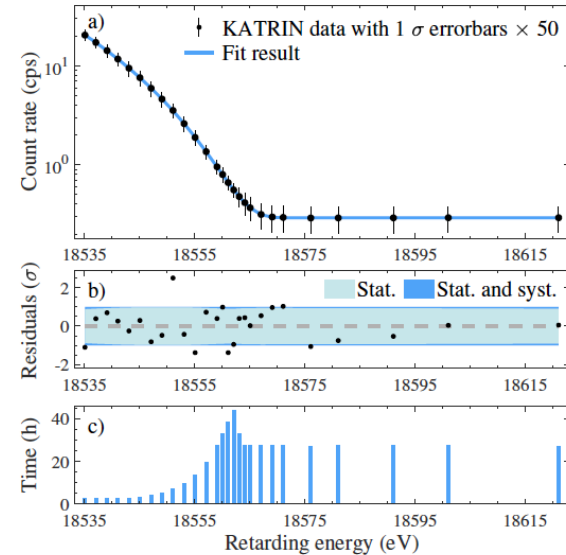
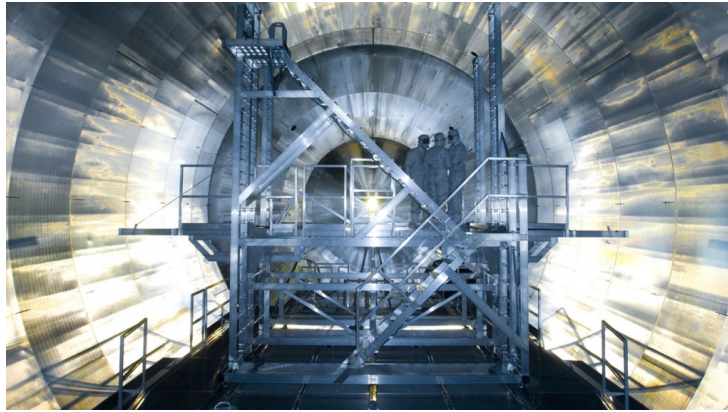


Sensitivity of future experiments



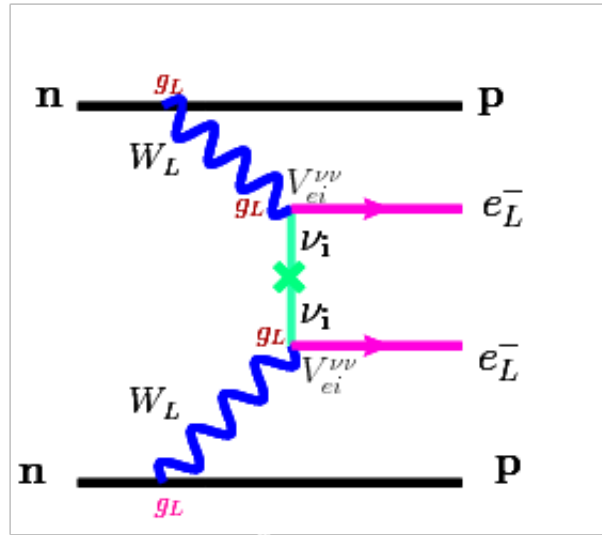
Different technologies of Hyper-K and DUNE (Water Cherenkov vs LAr TPC)
 => somewhat complementary

Absolute Neutrino Mass Scale

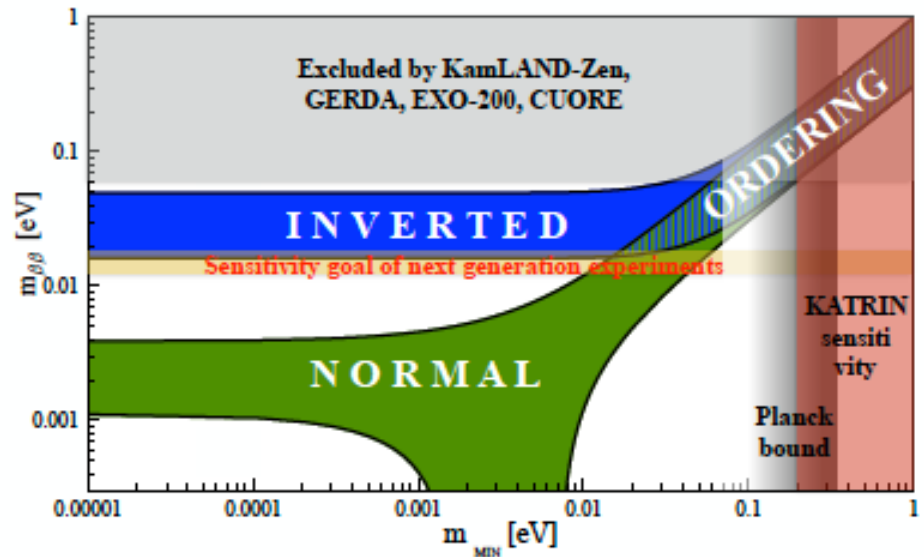
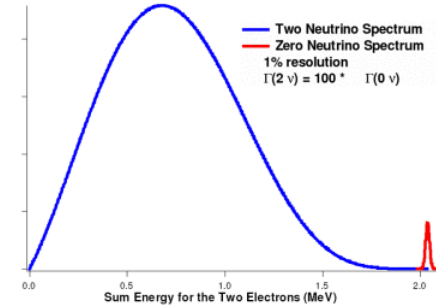


- First results from KATRIN (Sept. 2019)
 - Upper limit: $m_e < 1.1 \text{ eV}$ @ 90% CL
 - Based on 28 days of data taking
 - 5y sensitivity: $m_e < 0.2 \text{ eV}$
- Future cosmology missions (Euclid, DESI) will also have sensitivity

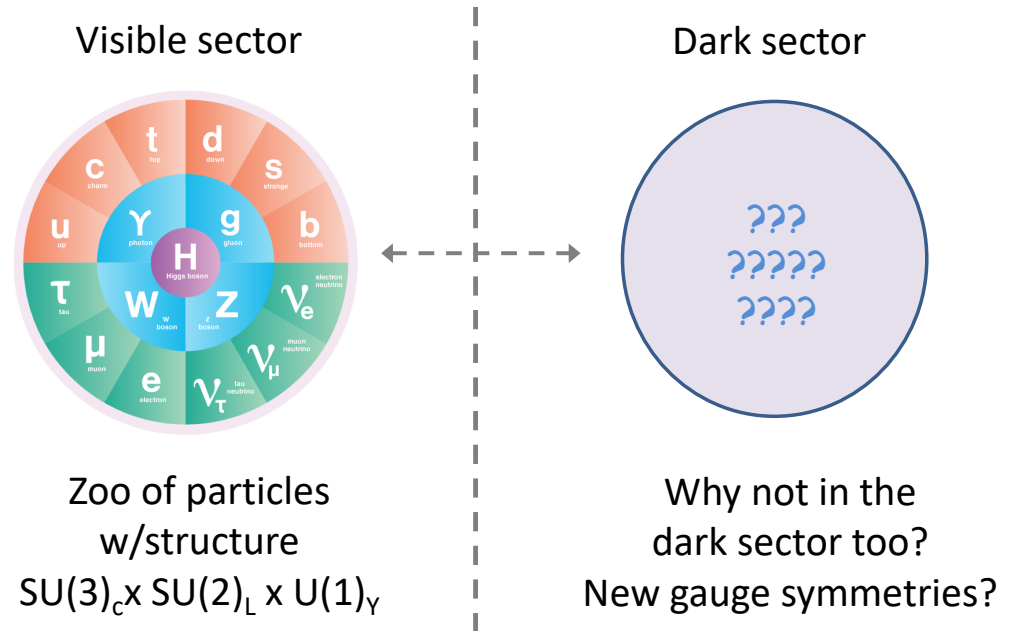
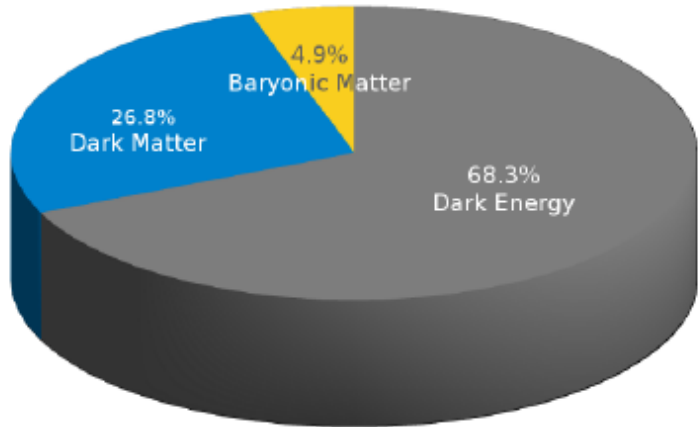
Is the neutrino its own anti-particle?



- Several experiments of next generation planned
- Should probe inverted hierarchy allowed region



Dark Sector

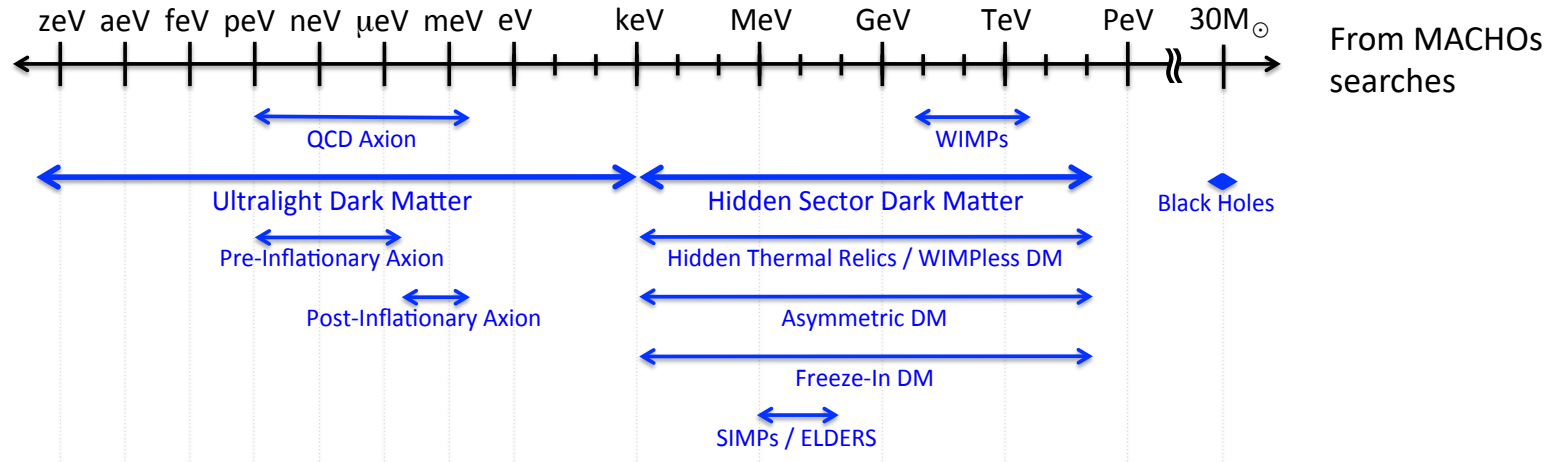


Y. Hochberg, LP '19

Dark Sector

Dark Matter Candidates: Very little clue on mass scales

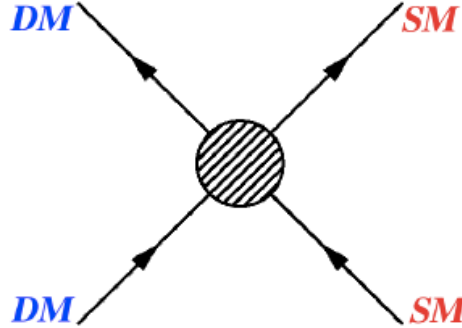
Too small mass
⇒ won't "fit"
in a galaxy!



graphics by M. Carena

How to search for WIMPs?

thermal freeze-out (early Univ.)
indirect detection (now)



production at colliders

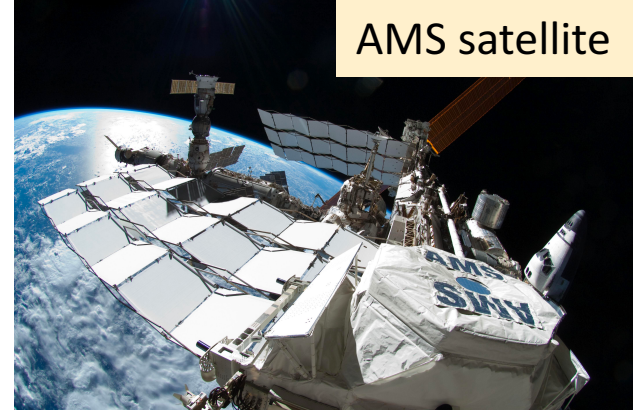
direct detection



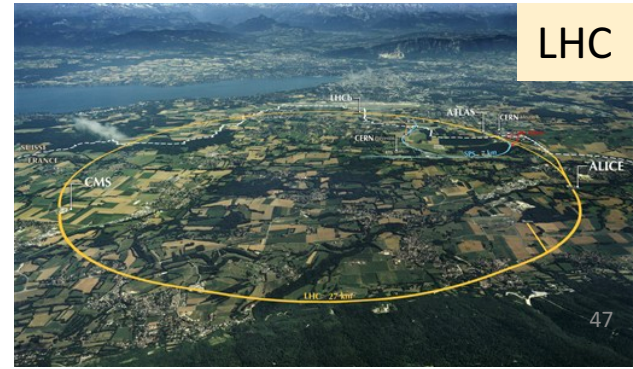
XENON1T in Gran Sasso



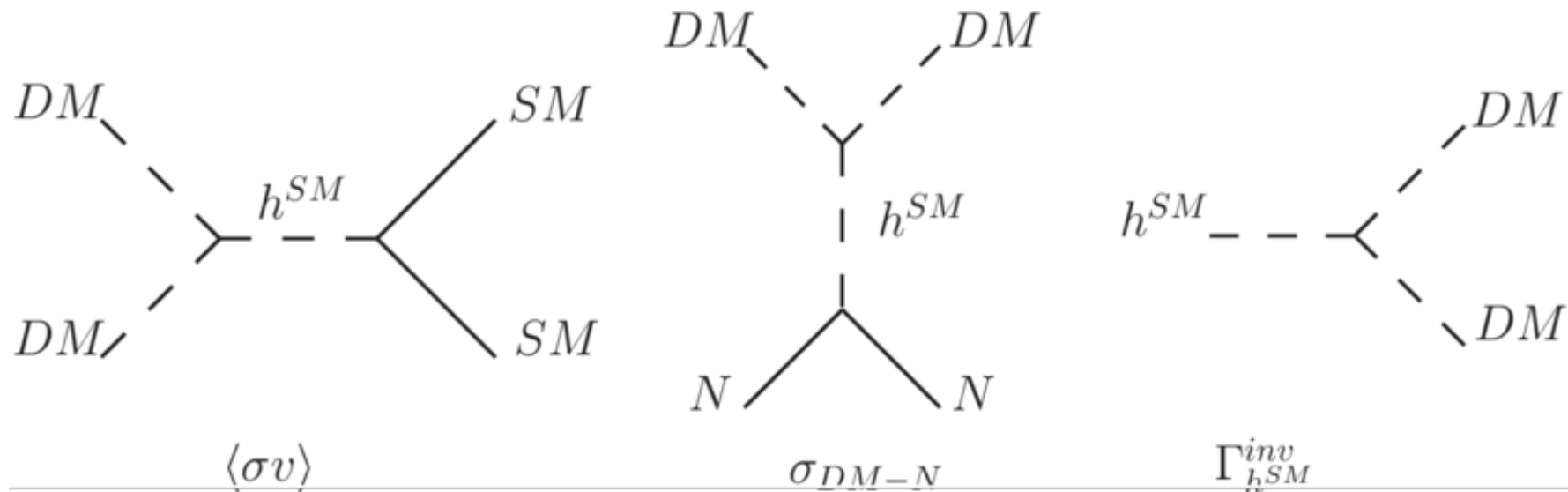
AMS satellite



LHC



Dark Matter processes with Higgs



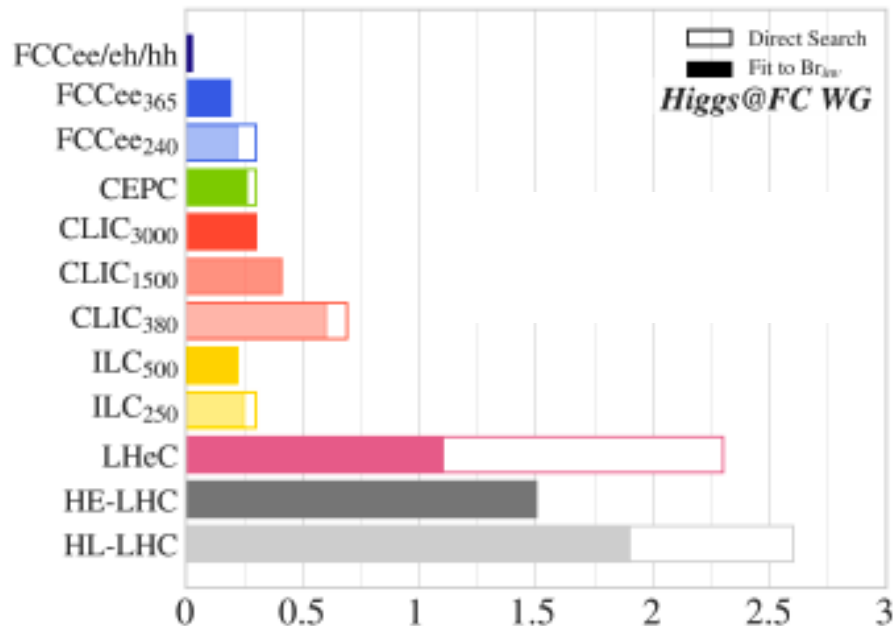
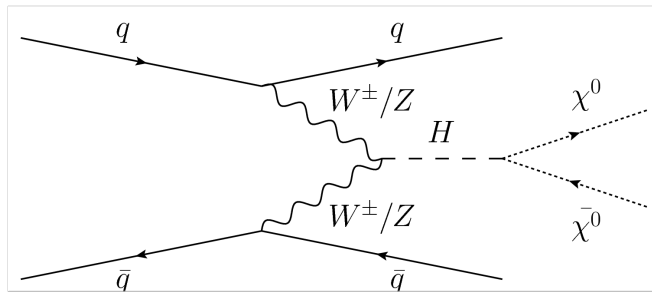
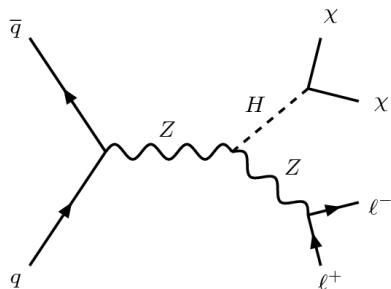
DM annihilation
early Universe &
satellites

DM-N Scattering
(XENON1T etc.)

DM production
early Universe &
LHC

“Invisible” Higgs decays?

- Does dark matter (χ) interact with the Higgs?
 - Higgs can decay to dark matter candidates if $m_H > 2 m_\chi$

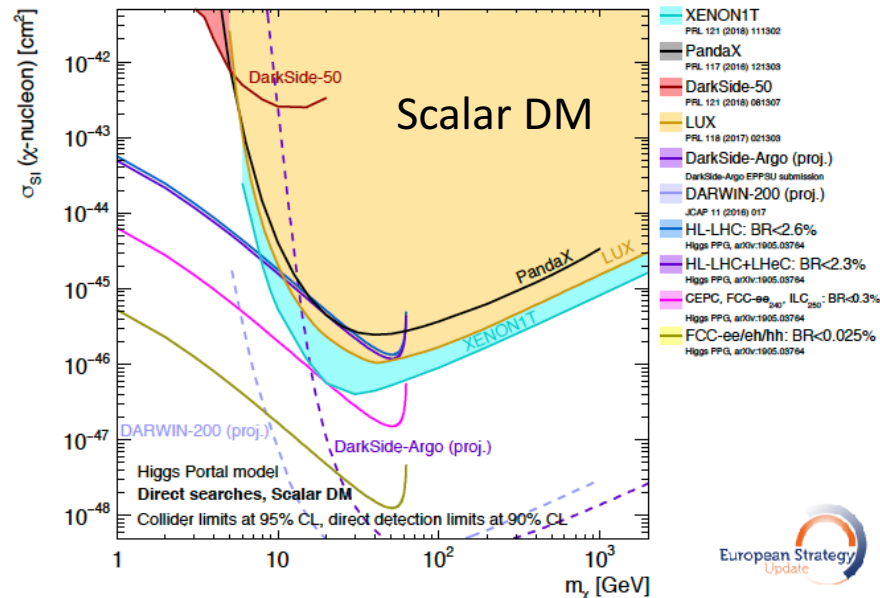


Comparing direct detection and Higgs constraints

Dark Matter can be scalar, vector or fermion:

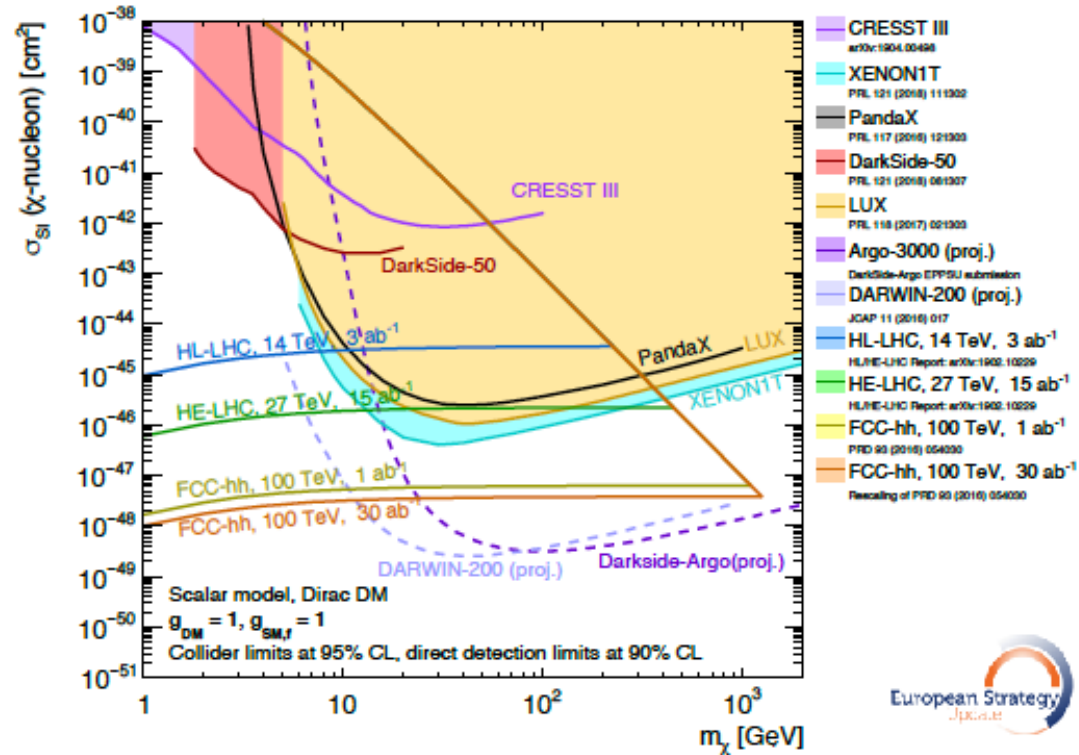
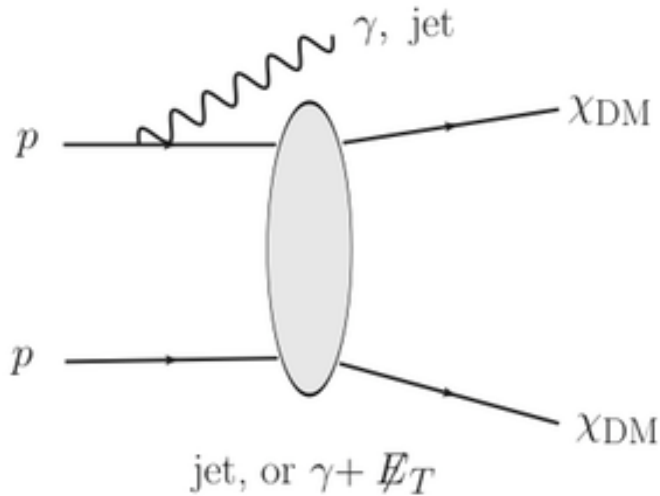
$$\text{BR}_\chi^{\text{inv}} \equiv \frac{\Gamma(H \rightarrow \chi\chi)}{\Gamma_H^{\text{SM}} + \Gamma(H \rightarrow \chi\chi)} = \frac{\sigma_{\chi p}^{\text{SI}}}{\Gamma_H^{\text{SM}}/r_\chi + \sigma_{\chi p}^{\text{SI}}}$$

with $r_\chi = \Gamma(H \rightarrow \chi\chi)/\sigma_{\chi p}^{\text{SI}}$



- Approaches are complementary
 - Higgs more sensitive at low mass, direct detection more sensitive at high mass
- Comparison is mode—dependent
 - This is good: if we see signal we will learn physics from it!

Monojet search



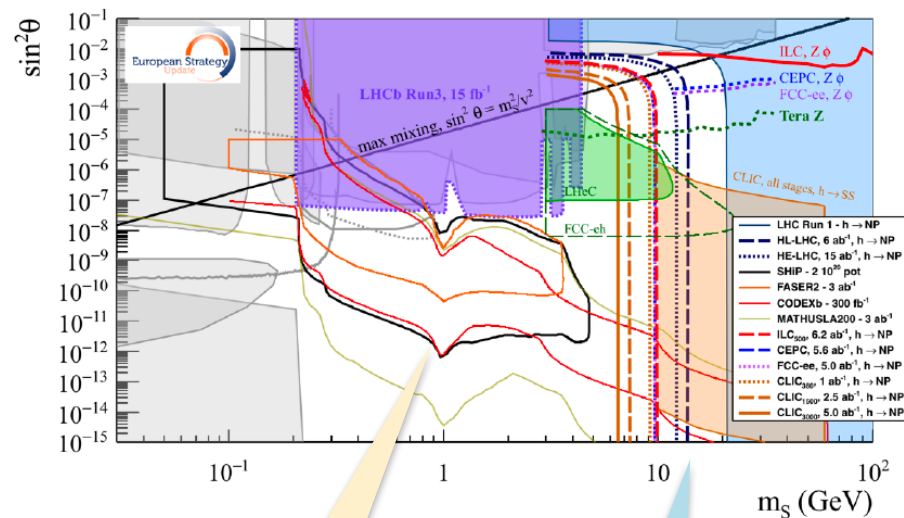
Can fully probe thermal WIMP with FCC-hh

Other Dark(ish) Particles: FIPs

Feebly interacting particles (FIPs)

Portal	Coupling
Vector (Dark Photon, A_μ)	$-\frac{\epsilon}{2\cos\theta_W}F'_{\mu\nu}B^{\mu\nu}$
Scalar (Dark Higgs, S)	$(\mu S + \lambda_{HS}S^2)H^\dagger H$
Fermion (Sterile Neutrino, N)	$y_N L H N$
Pseudo-scalar (Axion, a)	$\frac{a}{f_a}F_{\mu\nu}\tilde{F}^{\mu\nu}, \frac{a}{f_a}G_{i,\mu\nu}\tilde{G}_i^{\mu\nu}, \frac{\partial_\mu a}{f_a}\bar{\psi}\gamma^\mu\gamma^5\psi$

- Scalar portal very interesting!
 - Electroweak phase transition
 - Dark matter
 - Finetuning problem

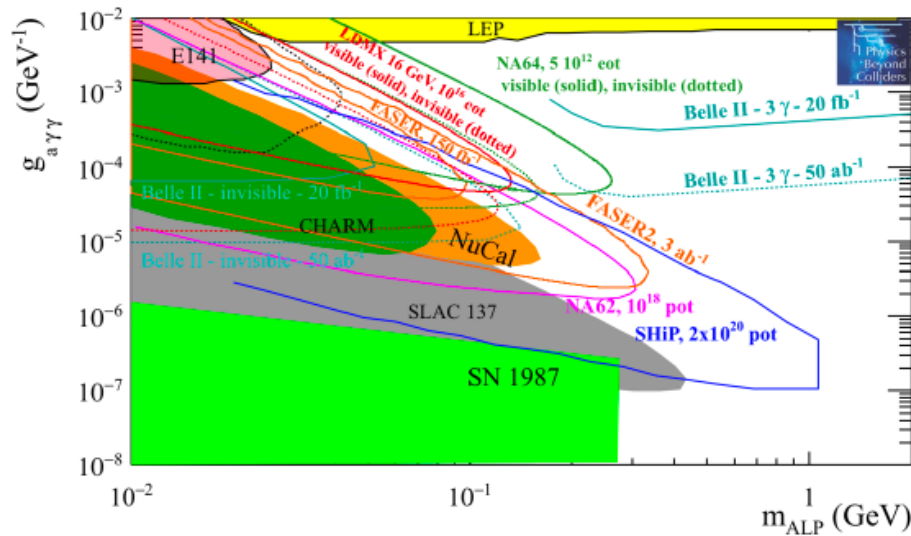
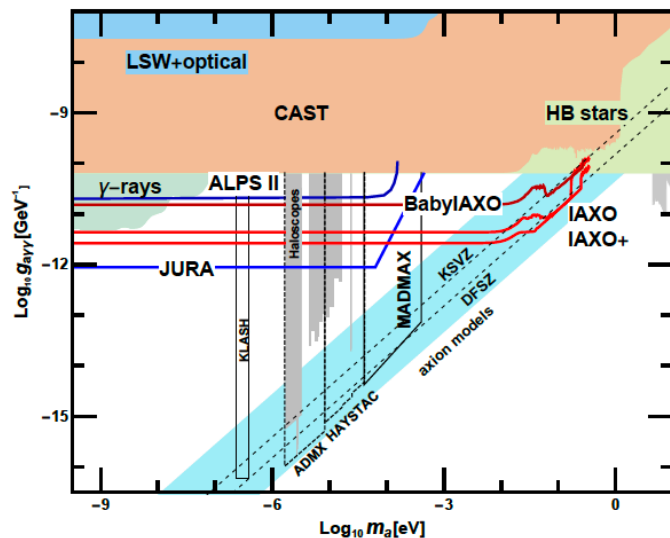


Beam
dump

$H \rightarrow \text{untagged}$

Axion-Like Particles (ALPs)

- Axion originally proposed as solution to strong CP problem (Peccei, Quinn)
 - Could also be relevant for Dark Matter, stellar evolution..
- Many complementary searches for Axion-Like Particles
 - Dedicated experiments (ALPS, CAST, IAXO), beam dump facilities (e.g. SHiP) and colliders



Is our world natural?

- Large quantum corrections to Higgs mass:

$$\Delta m_H^2 = -\frac{|\lambda_f|^2}{8\pi^2} [\Lambda_{UV}^2 + \dots] \sim M_{Pl}^2 \text{ if no new physics with } \Lambda < M_{Pl}$$

- Unnatural that $m_H \ll \Lambda_{NP} \Rightarrow$ “*hierarchy problem*”

- Quantify:

$$\epsilon \equiv \frac{m_H^2}{(\Delta m_H)^2}$$

Method	Dependence	Current Constraint
Direct searches: soft models	$\Delta m_H^2 \sim m_T^2$	$\epsilon \lesssim 1\%$
Direct searches: super-soft models	$\Delta m_H^2 \sim 3y_t^2/(4\pi^2)m_T^2$	$\epsilon \lesssim 10\%$
Direct searches: hyper-soft models	$\Delta m_H^2 \sim 3\lambda_h/(16\pi^2)m_T^2$	$\epsilon \lesssim 100\%$
Higgs couplings	$m_H^2/\Delta m_H^2 \sim \delta g_h/g_h$	$\epsilon \lesssim 10\%$
Oblique parameters (CH models)	$m_H^2/\Delta m_H^2 \sim \delta O \times 3$	$\epsilon \lesssim 30\%$
Oblique parameters (SUSY models)	$m_H^2/\Delta m_H^2 \sim \delta O \times 10^3$	n.a.

Simplicity vs Naturalness

R. Rattazzi

The two Chief Systems

I. The SM is valid up to $\Lambda_{UV} \gg TeV$

- B, L and Flavor: beautifully in accord with observation
- Higgs mass & C.C. hierarchy point beyond naturalness
 - multiverse
 - cosmological relaxation, Nnaturalness, ...
 - failure of EFT ideology (UV/IR connection)

II. Naturalizing New Physics appears at $\Lambda_{UV} \sim 1 TeV$

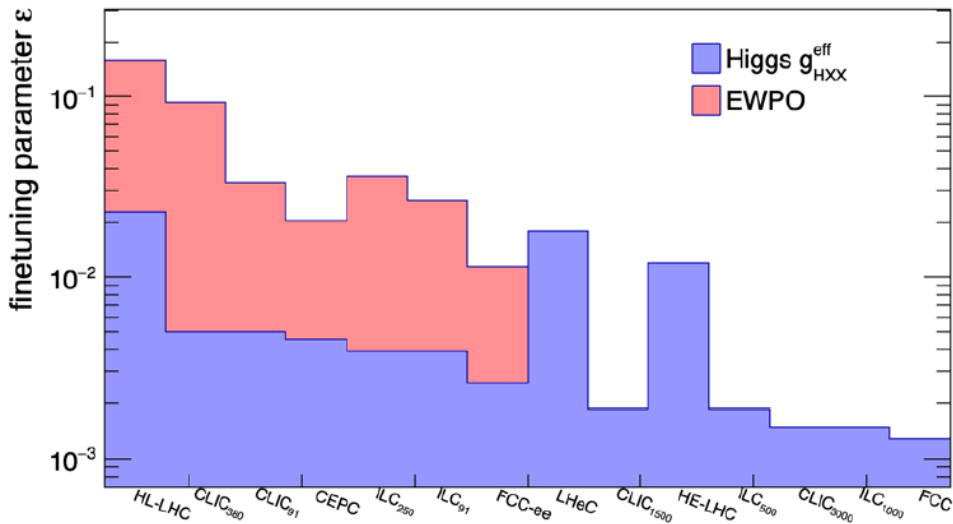
- Constraints on B, L, Flavor & CP met by clever model building

Simplicity

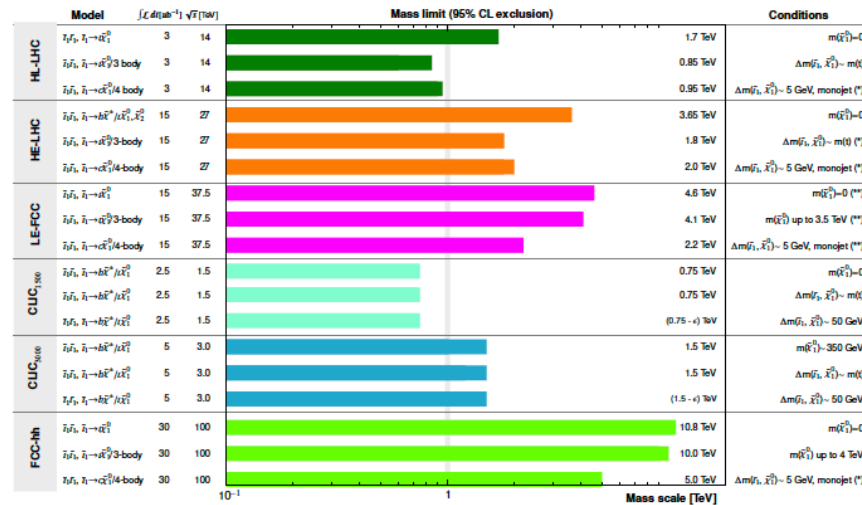


Naturalness

Is our world natural?



All Colliders: Top squark projections (R-parity conserving SUSY, prompt searches)



(*) indicates projection of existing experimental searches

(**) extrapolated from FCC-hh prospects

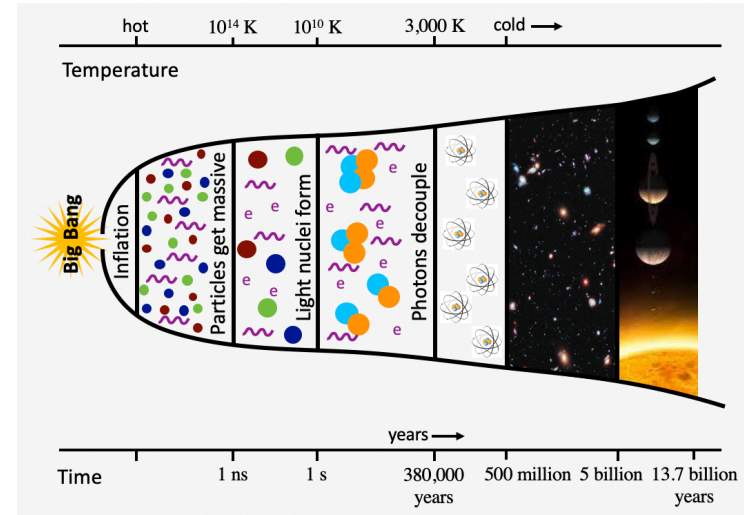
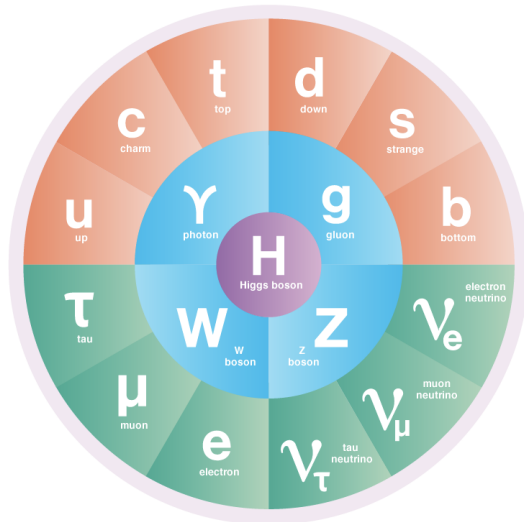
* indicates a possible non-evaluated loss in sensitivity

ILC 500: discovery in all scenarios up to kinematic limit $\sqrt{s}/2$

- Will probe naturalness to levels of 10^{-3} with Higgs couplings
 - Down to 10^{-2} - 10^{-4} with direct searches

Conclusions

- Amazingly exciting world of particle physics
 - So many questions... so many experiments to do...so many challenges!
- Hard to know where the next breakthrough is....
 - But let's break through ... even if it is difficult... and takes a while!



from Selya Ipek

Conclusions

- Amazingly exciting world of particle physics
 - So many questions... so many experiments to do...so many challenges!
- Hard to know where the next breakthrough is....
 - But let's break through ... even if it is difficult... and takes a while!

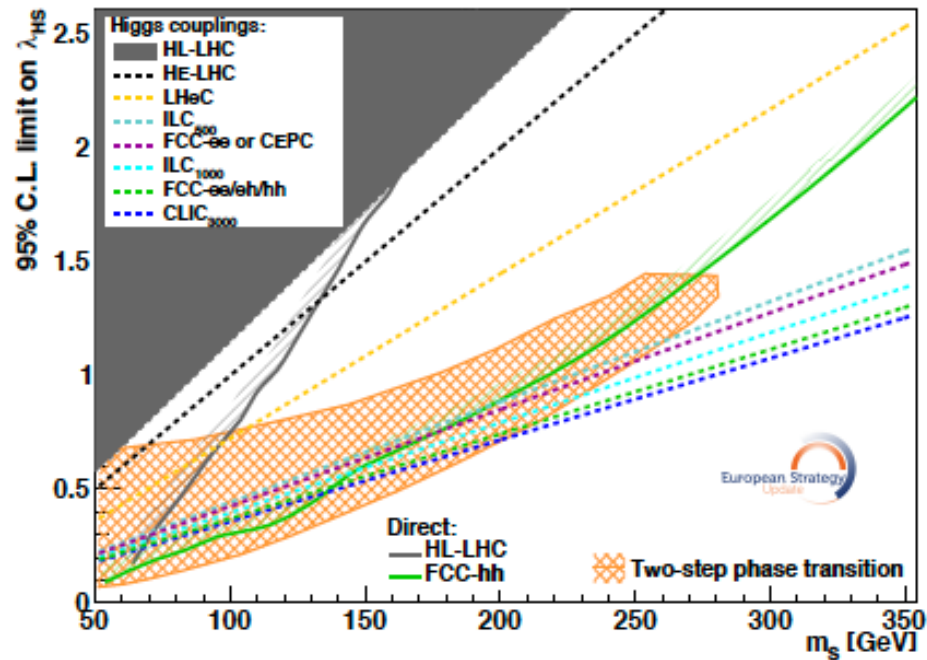
Eliud Kipchoge, Oct. 12th 2019
"No human is limited"

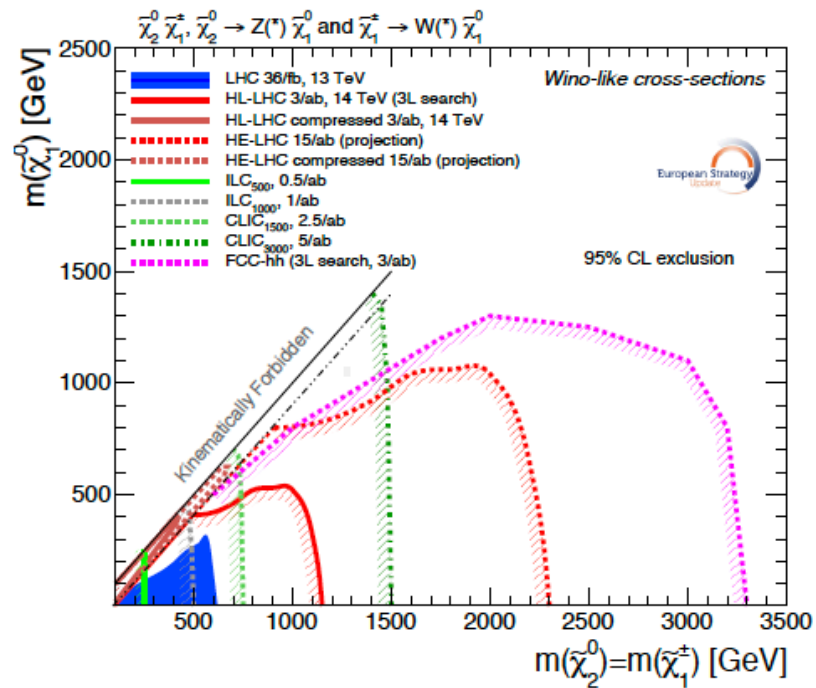


Backup

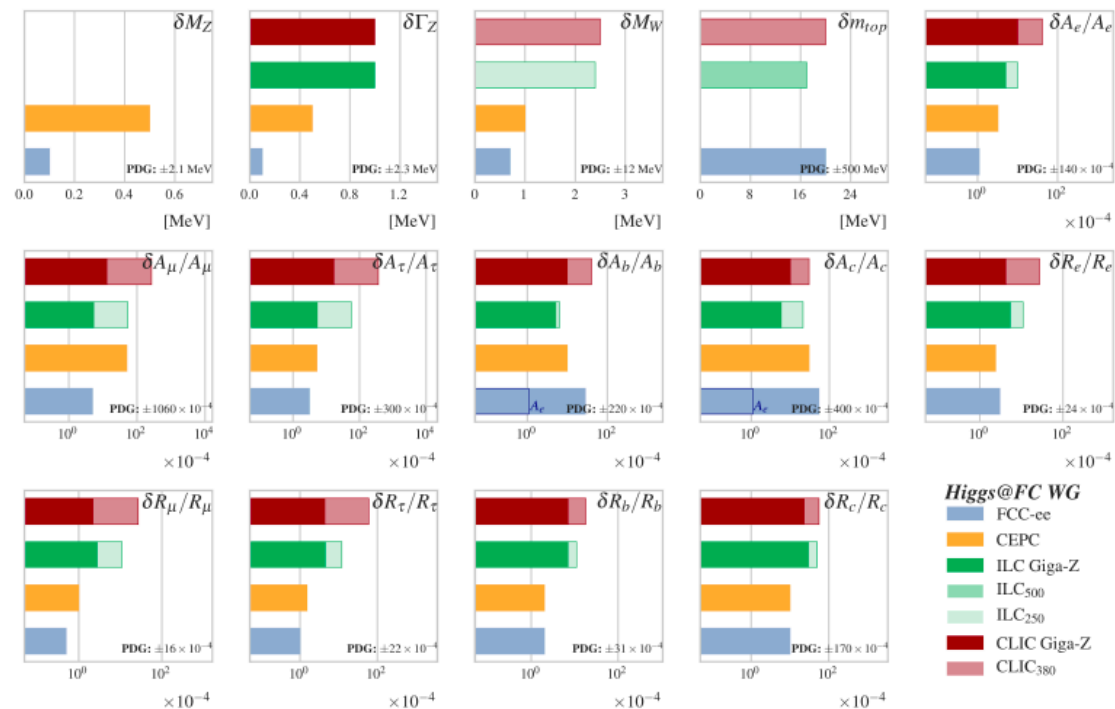
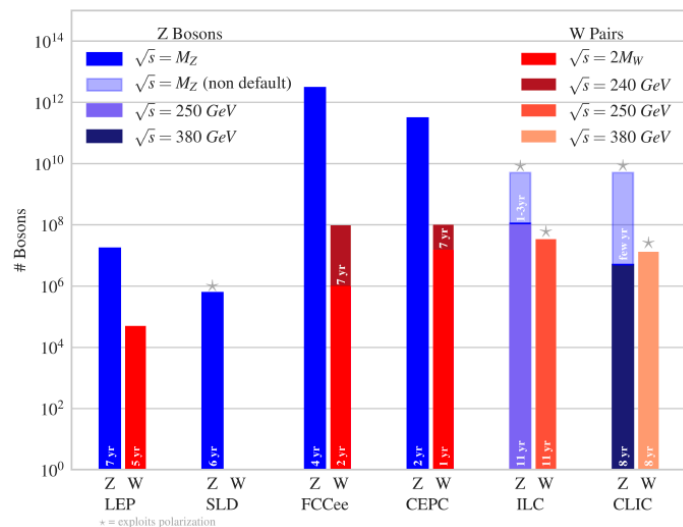
Is there a Singlet?

- Scalar singlet
 - May or may not mix with H boson
 - Mass below or above m_H
- Could address several questions:
 - Order of electroweak phase transition
 - Dark matter
 - ...
- Future experiments will probe very interesting parameter space



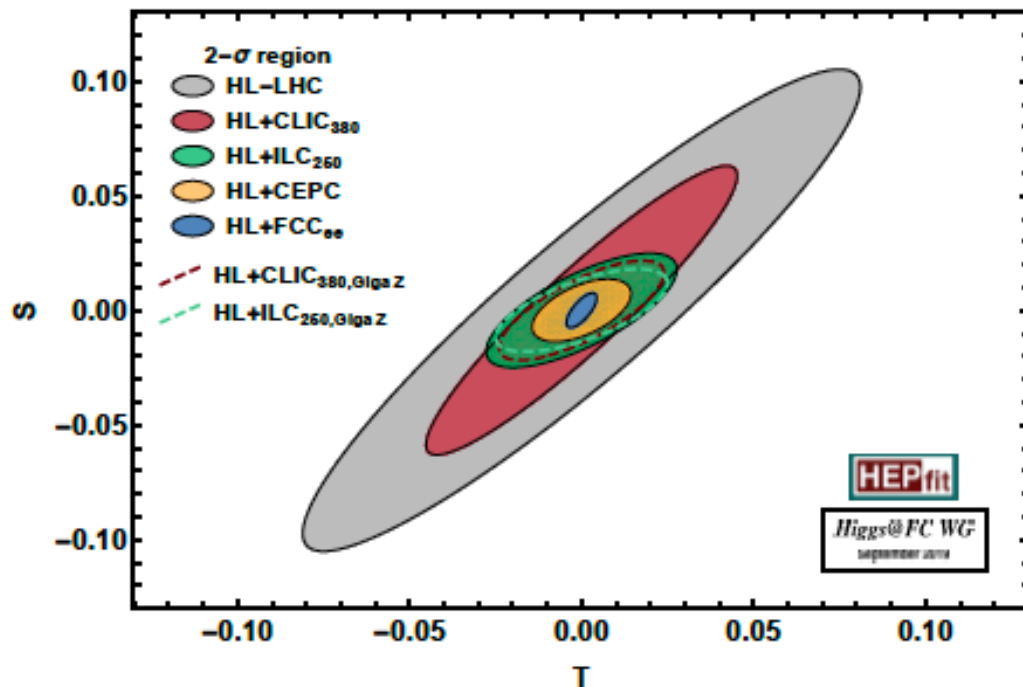
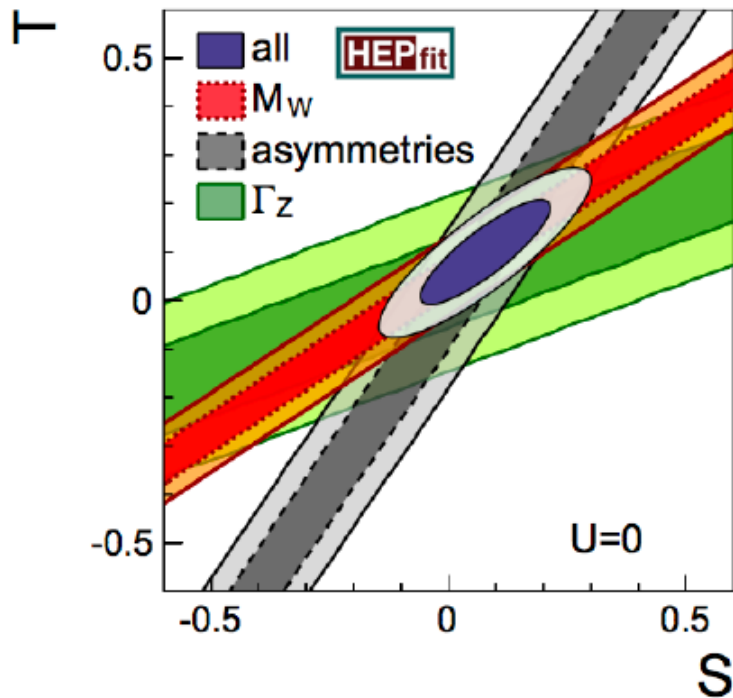


Electroweak Precision Data



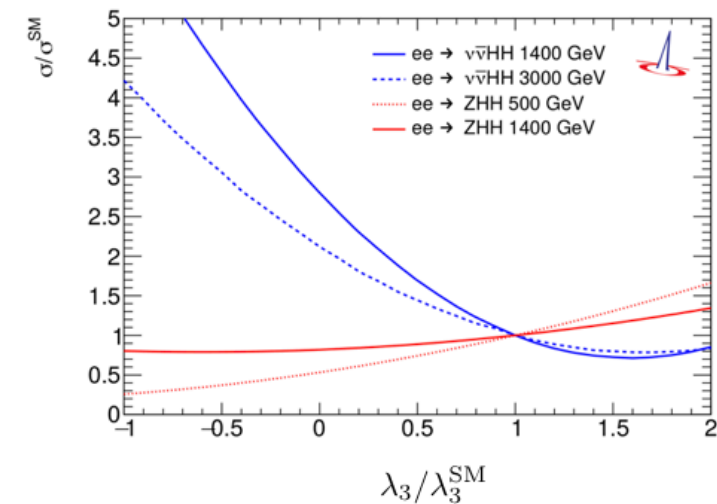
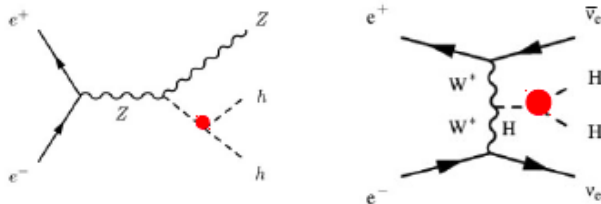
Huge improvements with future e+e- colliders ("Tera-Z")

S and T parameters

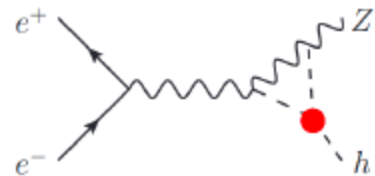


Measurement of Higgs Self-Coupling: Lepton Colliders

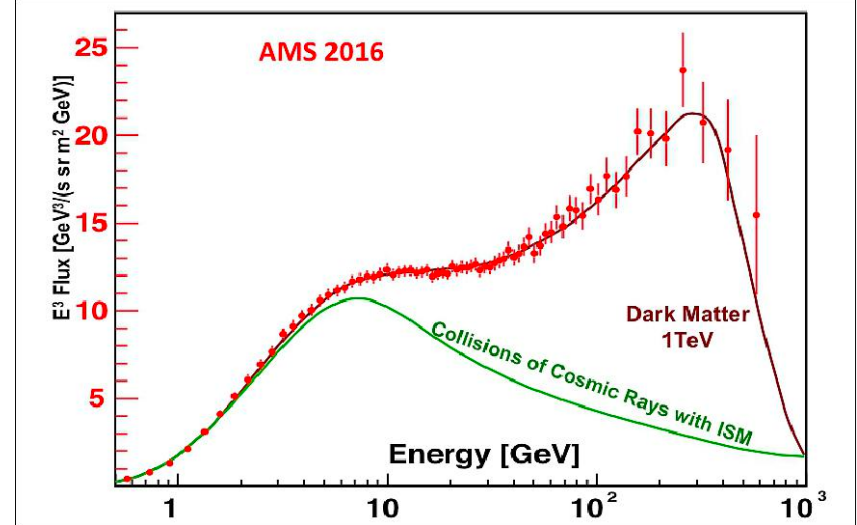
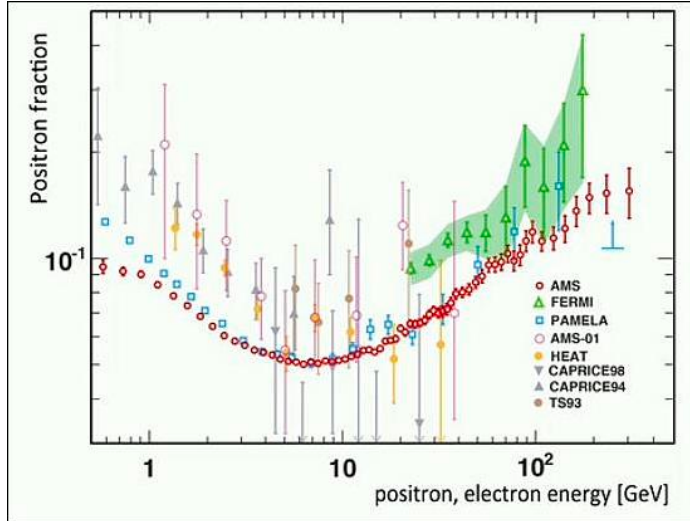
- Di-Higgs processes at lepton colliders
 - ZHH or VBF production complementary



- Single-Higgs production sensitive through loop effects, e.g. for $\kappa_\lambda = 2$:
 - Hadron colliders: $\sim 3\%$
 - Lepton colliders: $\sim 1\%$



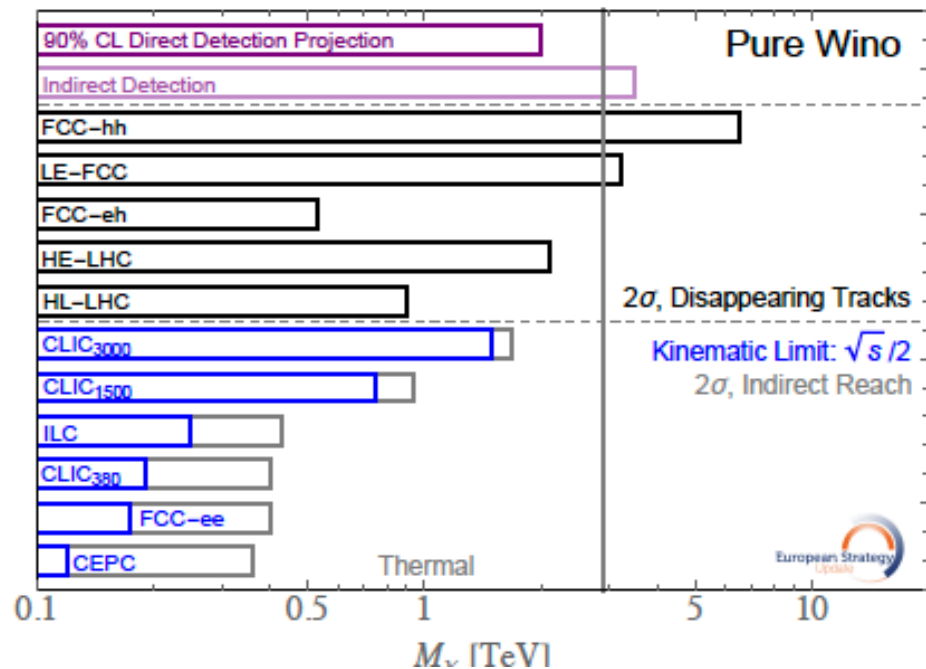
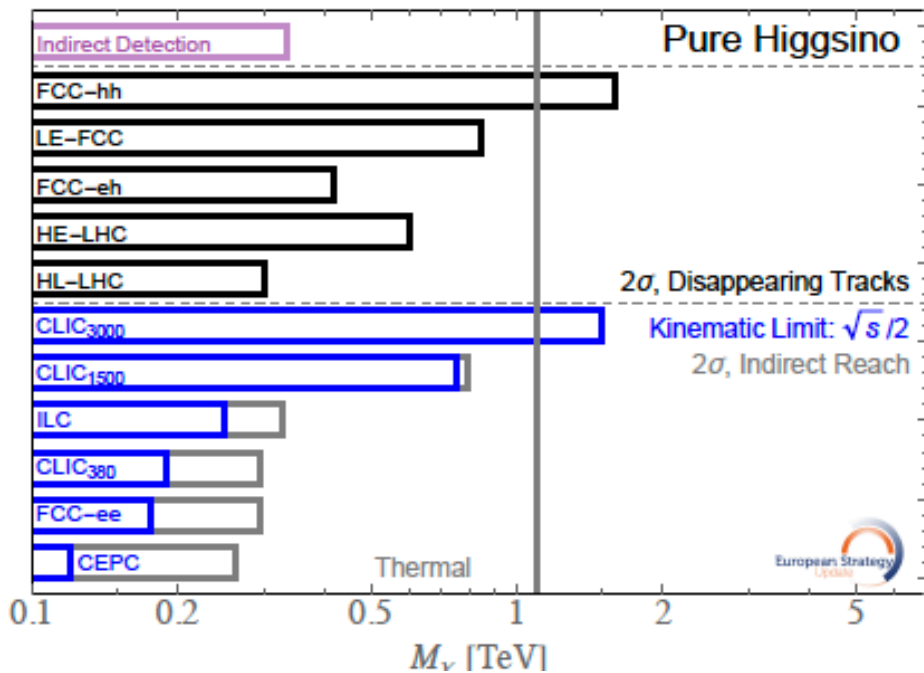
AMS positron flux

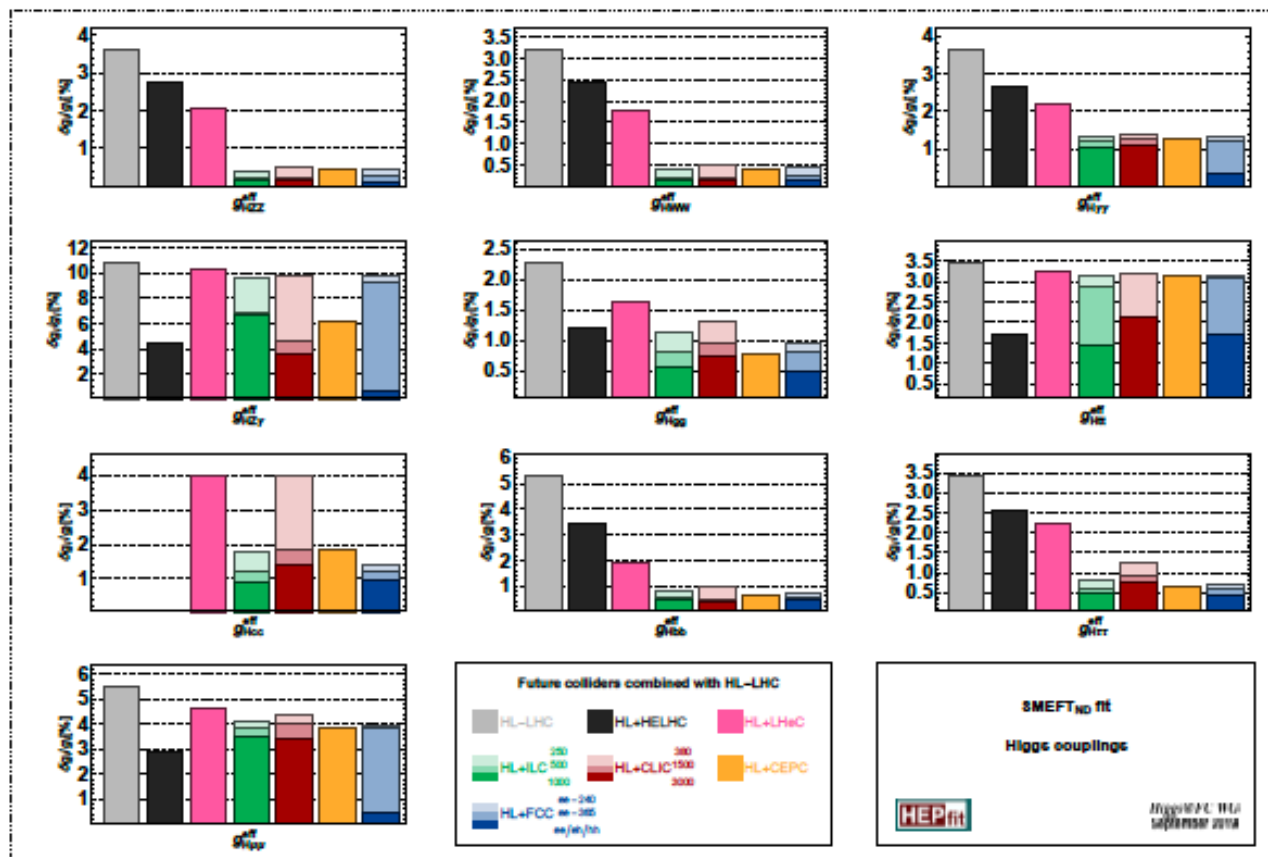


J. Kopp, Phys.Rev. D88 (2013) 076013

- AMS sees increase in positron flux as function of energy
- One possible explanation: DM annihilation to $\tau^+\tau^-$ or $\mu^+\mu^-$ pairs
 - These subsequently decay to electrons
- Another explanation: astrophysical sources, e.g. pulsars

Dark Matter searches compared





Comparing direct detection and Higgs constraints

Dark Matter can be scalar, vector or fermion:

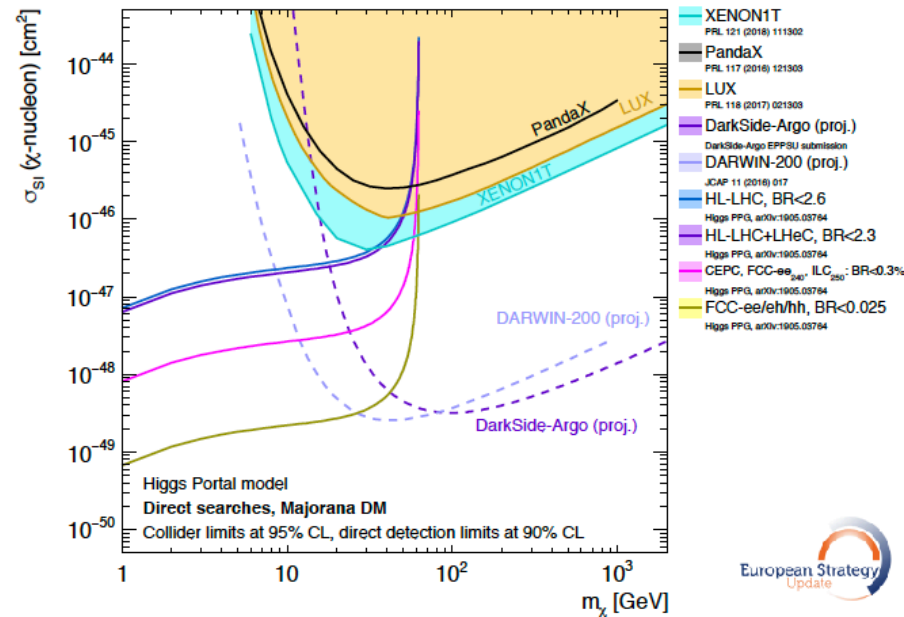
Scalar: $\Delta\mathcal{L}_S = -\frac{1}{2}m_S^2 S^2 - \frac{1}{4}\lambda_S S^4 - \frac{1}{4}\lambda_{hSS}H^\dagger HS^2,$

Vector: $\Delta\mathcal{L}_V = \frac{1}{2}m_V^2 V_\mu V^\mu + \frac{1}{4}\lambda_V (V_\mu V^\mu)^2 + \frac{1}{4}\lambda_{hVV}H^\dagger HV_\mu V^\mu,$

Fermion: $\Delta\mathcal{L}_f = -\frac{1}{2}m_f \bar{f}f - \frac{1}{4}\frac{\lambda_{hff}}{\Lambda}H^\dagger H \bar{f}f + \text{h.c.} . \quad (5)$

$$\text{BR}_\chi^{\text{inv}} \equiv \frac{\Gamma(H \rightarrow \chi\chi)}{\Gamma_H^{\text{SM}} + \Gamma(H \rightarrow \chi\chi)} = \frac{\sigma_{\chi p}^{\text{SI}}}{\Gamma_H^{\text{SM}}/r_\chi + \sigma_{\chi p}^{\text{SI}}}$$

With $r_\chi = \Gamma(H \rightarrow \chi\chi)/\sigma_{\chi p}^{\text{SI}}$



Let there be a singlet

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{2}(\partial_\mu S)(\partial^\mu S) - \frac{m_s^2}{2}S^2 - \frac{a_s}{3}S^3 - \frac{\lambda_s}{4}S^4 - \lambda_{hs}\Phi^\dagger\Phi S^2 - 2a_{hs}\Phi^\dagger\Phi S$$

$$\langle\Phi\rangle = (0, v/\sqrt{2}), \quad \langle S\rangle = v_s$$

Mixing between S and H:

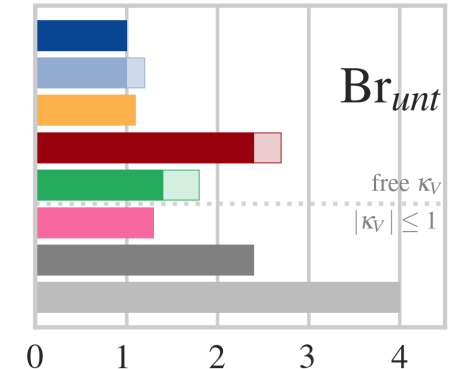
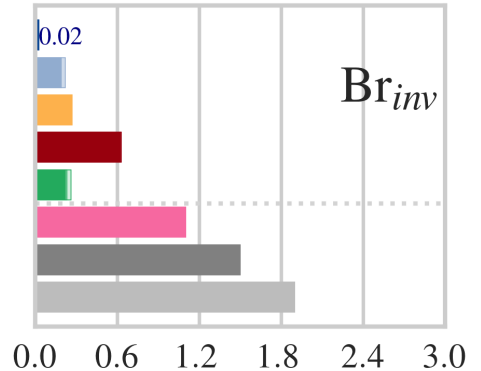
$$\sin 2\theta = \frac{4v(a_{hs} + \lambda_{hs}v_s)}{M_h^2 - M_s^2}$$

Z_2 symmetry: $a_{hs} = v_s = 0$

=> No mixing between H and S

If $M_S < M_H/2 \Rightarrow H \rightarrow SS$ decay possible

- Z_2 symmetry: $S \rightarrow \text{invisibly}$
- $\sin\theta \neq 0$:
 - $S \rightarrow f\bar{f}$ with $y_{f,S} = \sin\theta y_{f,H} \Rightarrow$ direct searches
 - $S \rightarrow \text{undetected}$ is also sensitive in global fit



Let there be a singlet

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{2}(\partial_\mu S)(\partial^\mu S) - \frac{m_s^2}{2}S^2 - \frac{a_s}{3}S^3 - \frac{\lambda_s}{4}S^4 - \lambda_{hs}\Phi^\dagger\Phi S^2 - 2a_{hs}\Phi^\dagger\Phi S \quad \langle\Phi\rangle = (0, v/\sqrt{2}), \quad \langle S\rangle = v_s$$

Mixing between S and H:

$$\sin 2\theta = \frac{4v(a_{hs} + \lambda_{hs}v_s)}{M_h^2 - M_s^2}$$

Z_2 symmetry: $a_{hs} = v_s = 0$

=> No mixing between H and S

If $M_S > M_H/2$ and $\sin\theta \neq 0$

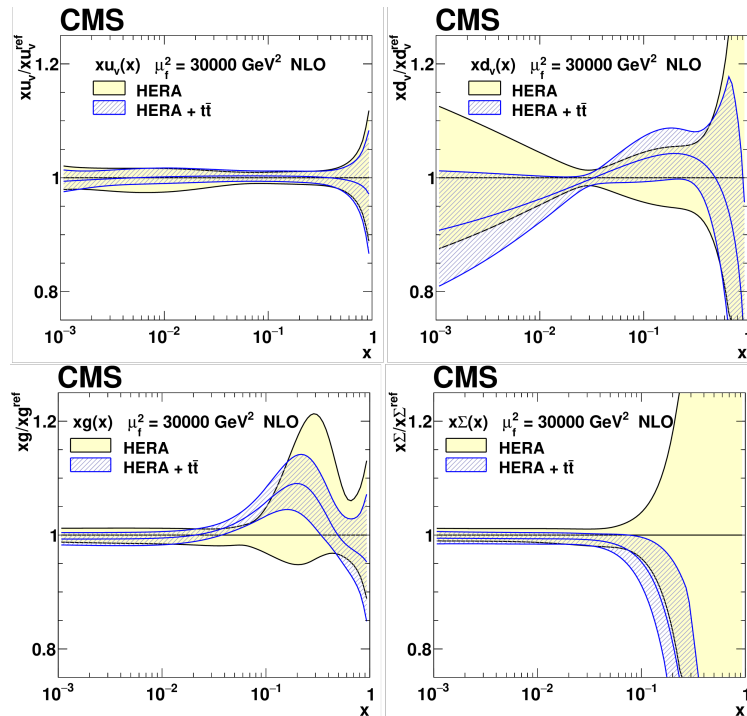
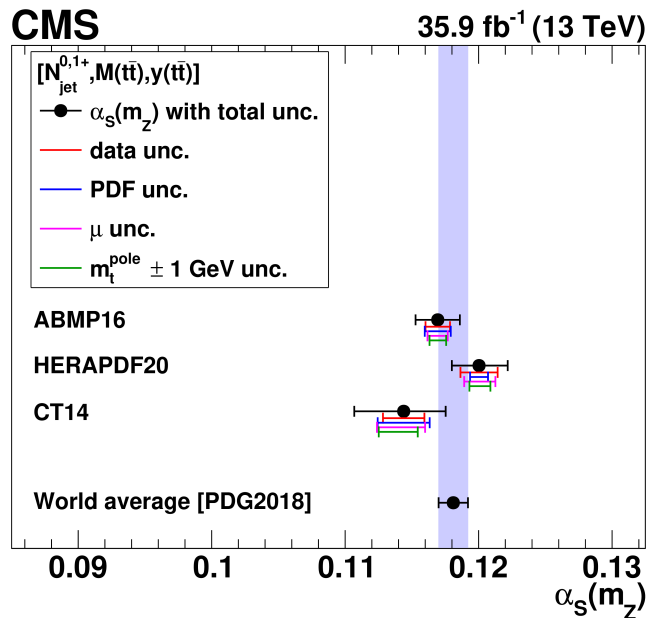
- all H couplings modified => BR unaffected
- Total production cross section modified

$$\mu_i^f \approx \cos^2 \theta.$$

For 1st order ewk phase transition
(Katz, Perelstein, arXiv:1401.1827): $\mu_i^f - 1 \gtrsim 0.6\%$.

	$\Delta\mu$ (%)
now	7.8
HL-LHC	1.5-3
ILC	0.7
FCC-ee	0.5
CLIC	0.24

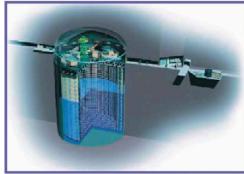
Perturbative QCD at the LHC



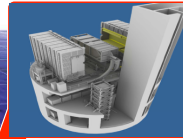
- Important consistency check
 - Actually some tension between $t\bar{t}$ data and other data seen in QCD fits

Long-Baseline Experiments in Japan and USA

T2K



Super-Kamiokande
(ICRR, Univ. Tokyo)



**INGRID +
ND280**

J-PARC Main Ring
(KEK-JAEA, Tokai)



$$E_\nu \simeq 0.7 \text{ GeV},$$

$$\Delta \equiv \frac{1.27 \cdot 0.0025 \text{ eV}^2 \cdot 295 \text{ km}}{0.7 \text{ GeV}} \simeq \frac{\pi}{2}$$

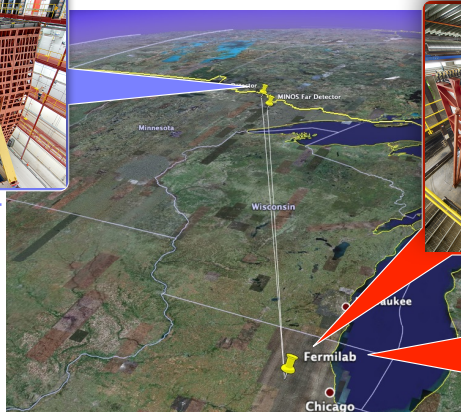
NOvA



NOvA Far Detector



**NOvA
Near
Detector**



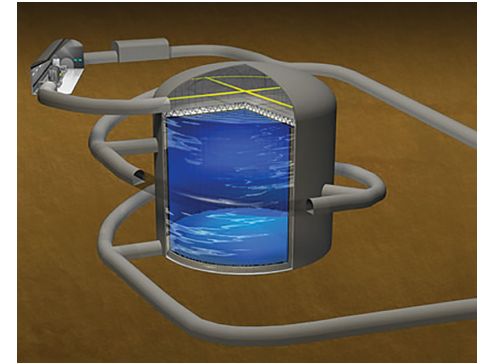
$$E_\nu \simeq 2 \text{ GeV},$$

$$\Delta \equiv \frac{1.27 \cdot 0.0025 \text{ eV}^2 \cdot 810 \text{ km}}{2 \text{ GeV}} \simeq \frac{\pi}{2}$$

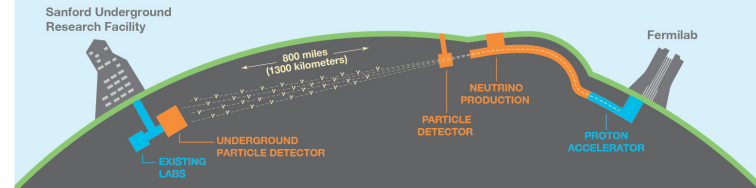
Fermilab Main Injector



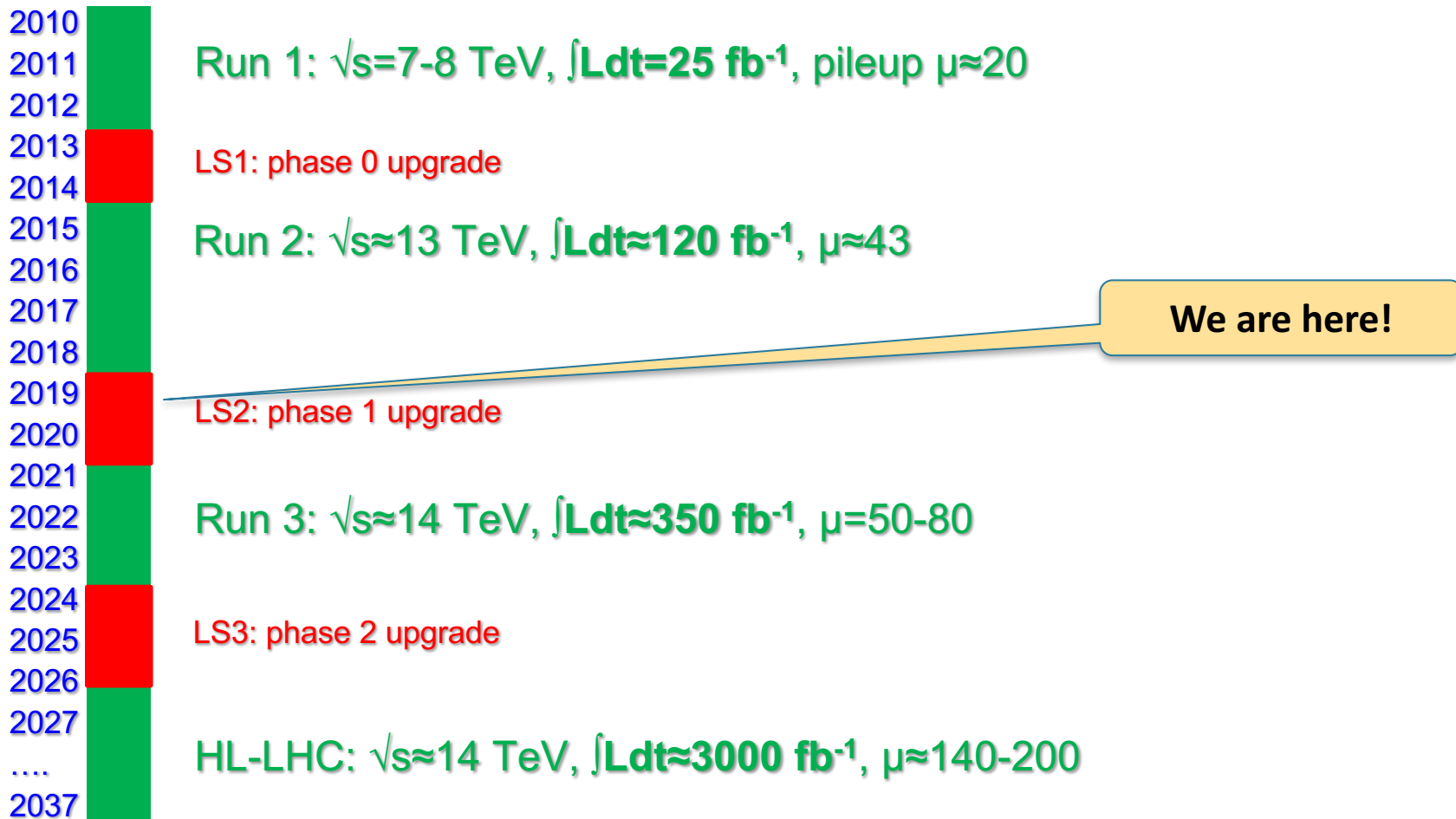
Hyper-Kamiokande



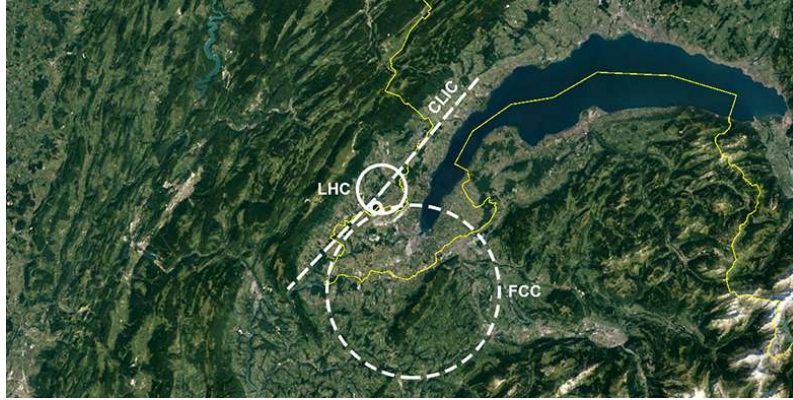
LBNF beamline => DUNE



LHC Roadmap

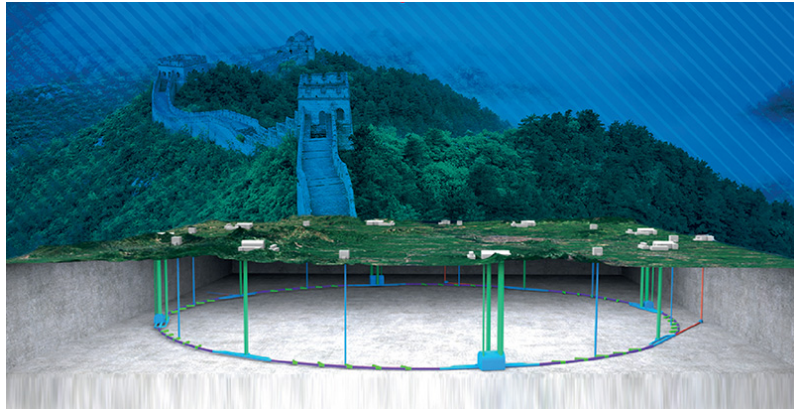


Future Colliders

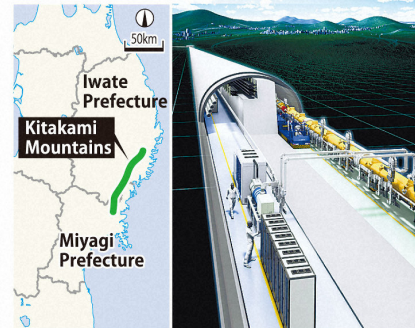


Proposed colliders:

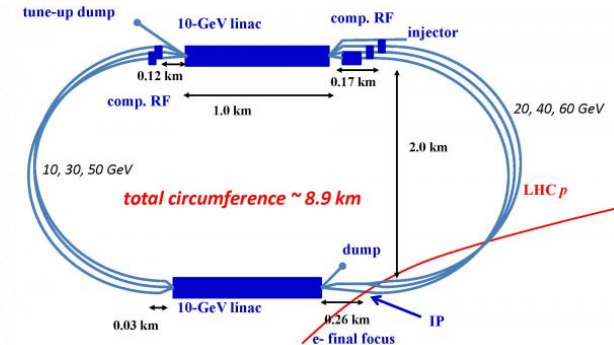
- Linear e^+e^- : ILC, CLIC
- Circular e^+e^- : FCC-ee, CePC
- pp : HE-LHC, FCC-hh, SppC
- ep : LHeC, FCC-eh



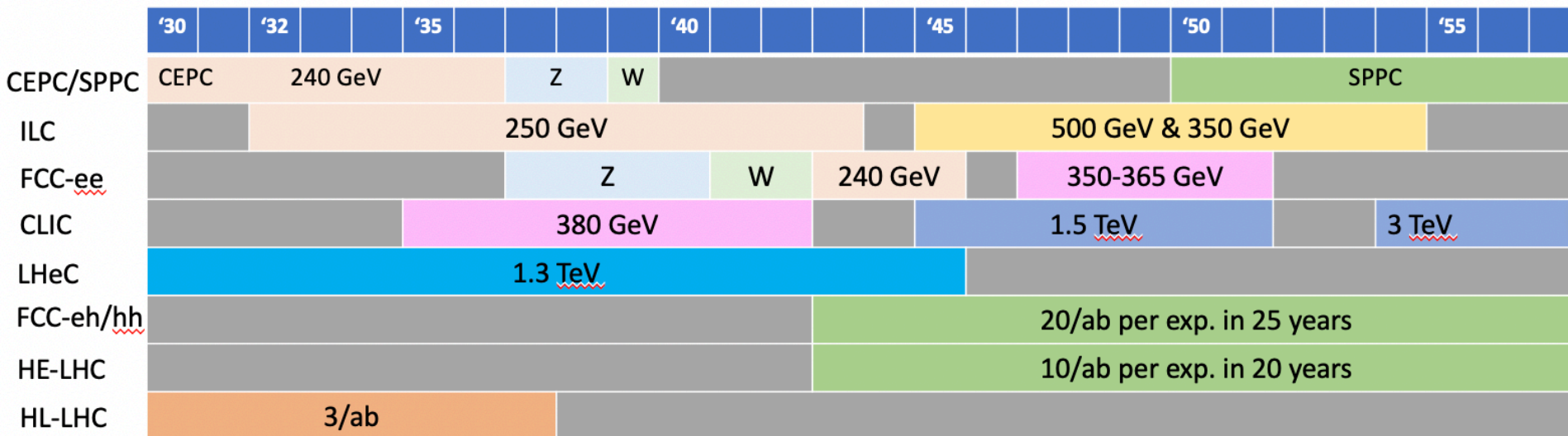
Planned location and artist's rendering of ILC



Artist's rendering provided by the Linear Collider Collaboration

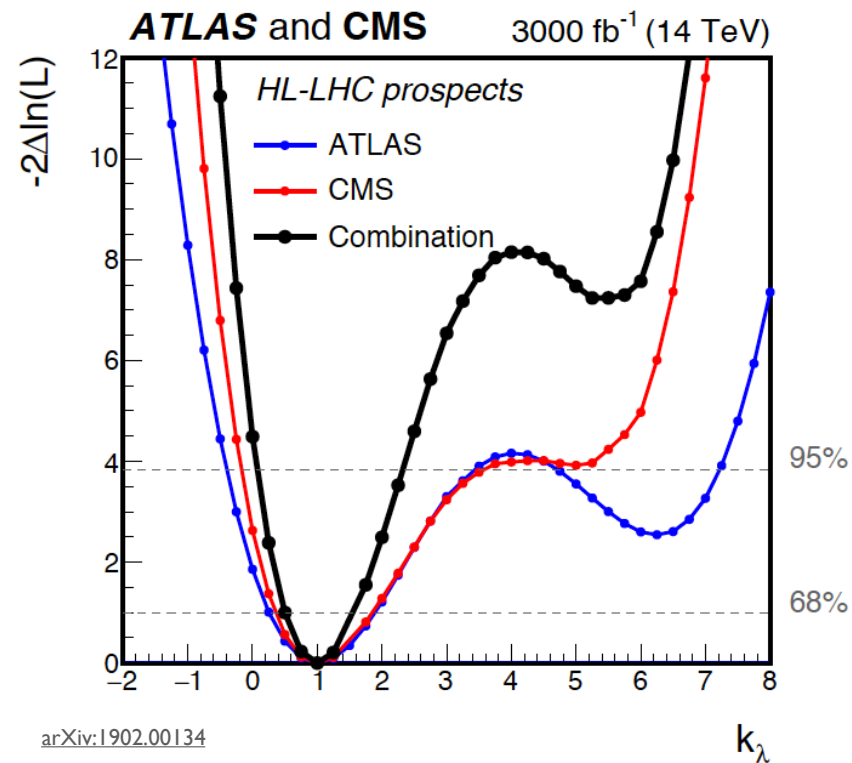


Future Colliders: \sqrt{s} and tentative timescales

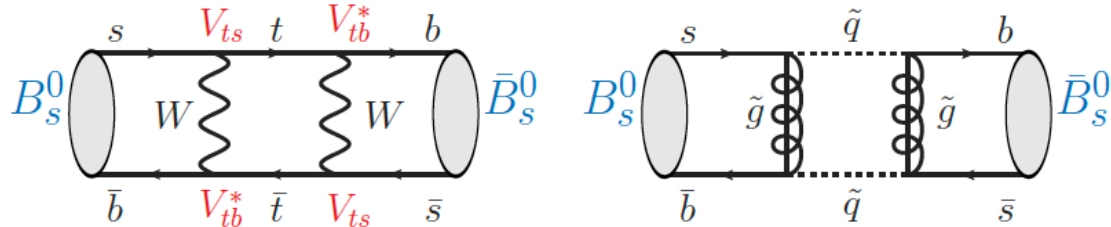
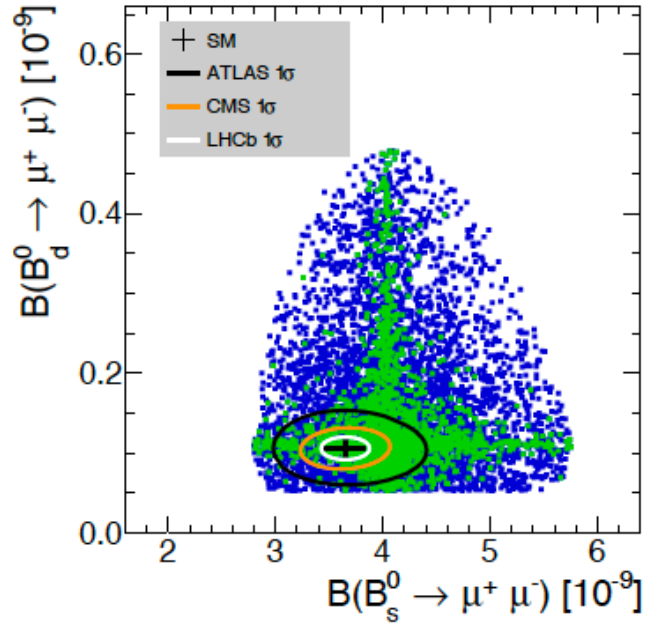


Sensitivity to κ_λ at HL-LHC

- 50% sensitivity at HL-LHC....
- Sensitivity very relevant for probing order of phase transition!



New physics with heavy quarks

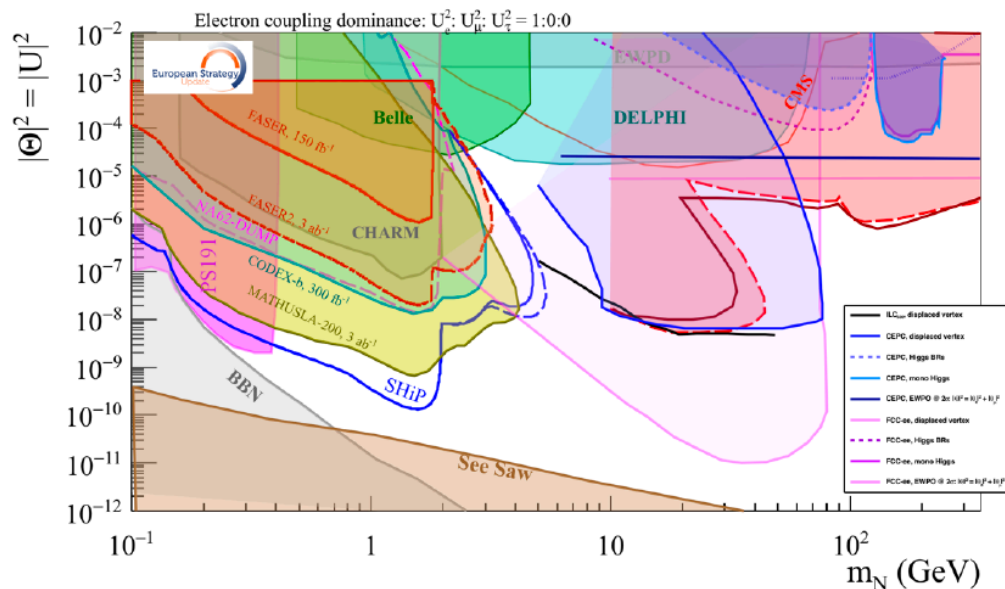


Sensitive to many models of new physics

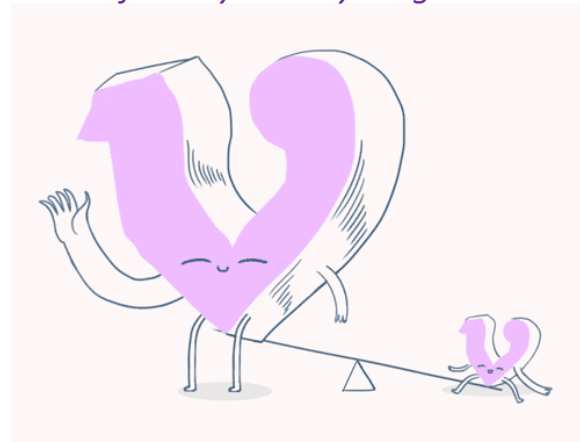
What about right-handed neutrinos

Minkowski (77); Gell-Mann, Ramond, Slansky; Yanagida; Glashow; R. N. M., Senjanovic (1979)

- *Seesaw mechanism* could explain $m_\nu \ll m_e$
- Requires heavy RH neutrino: $M_R/g \gg v$
 - typically at GUT scale but could be lower but could be lower if g very small

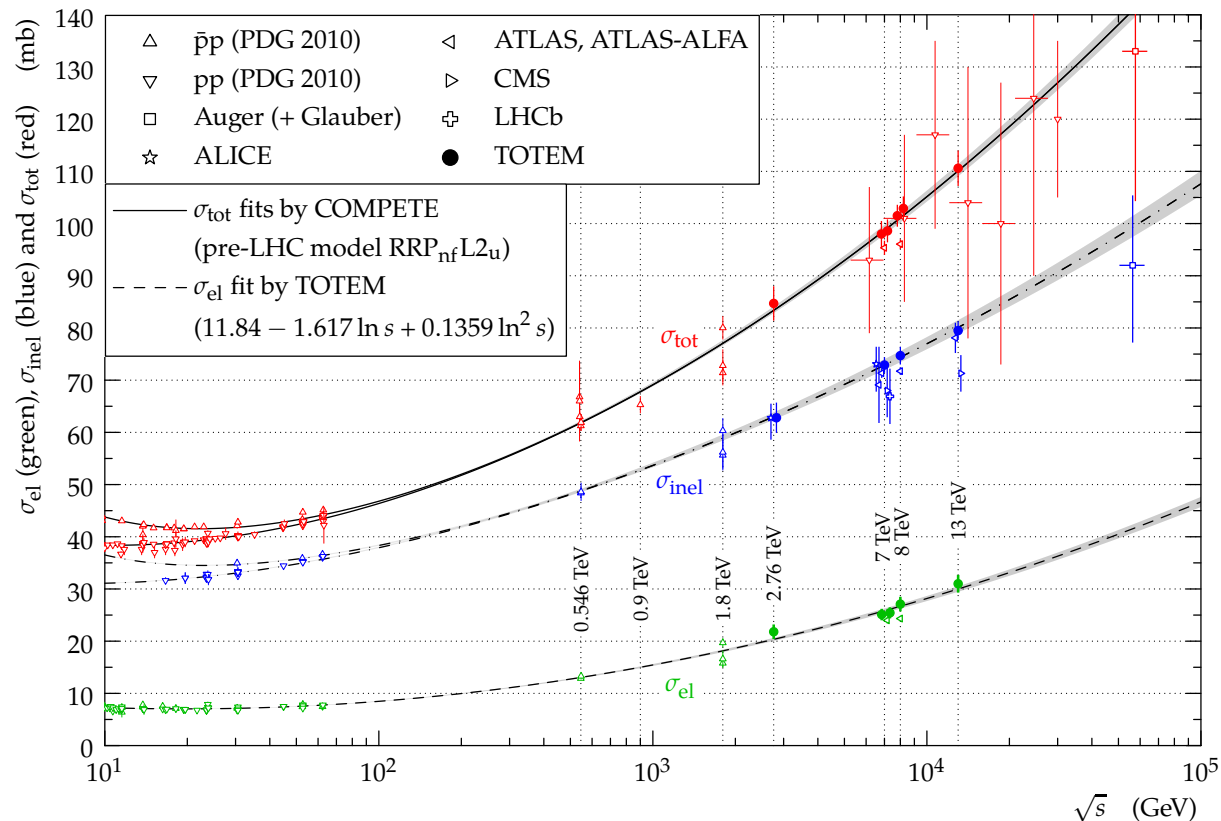


from Symmetry magazine



Probed by e.g. beam dump & collider experiments

Total, Inelastic and Elastic Cross Sections



arXiv:1712.06153
and refs therein

