

# DUNE and DUNE-Prism concept.

FLC Long talk

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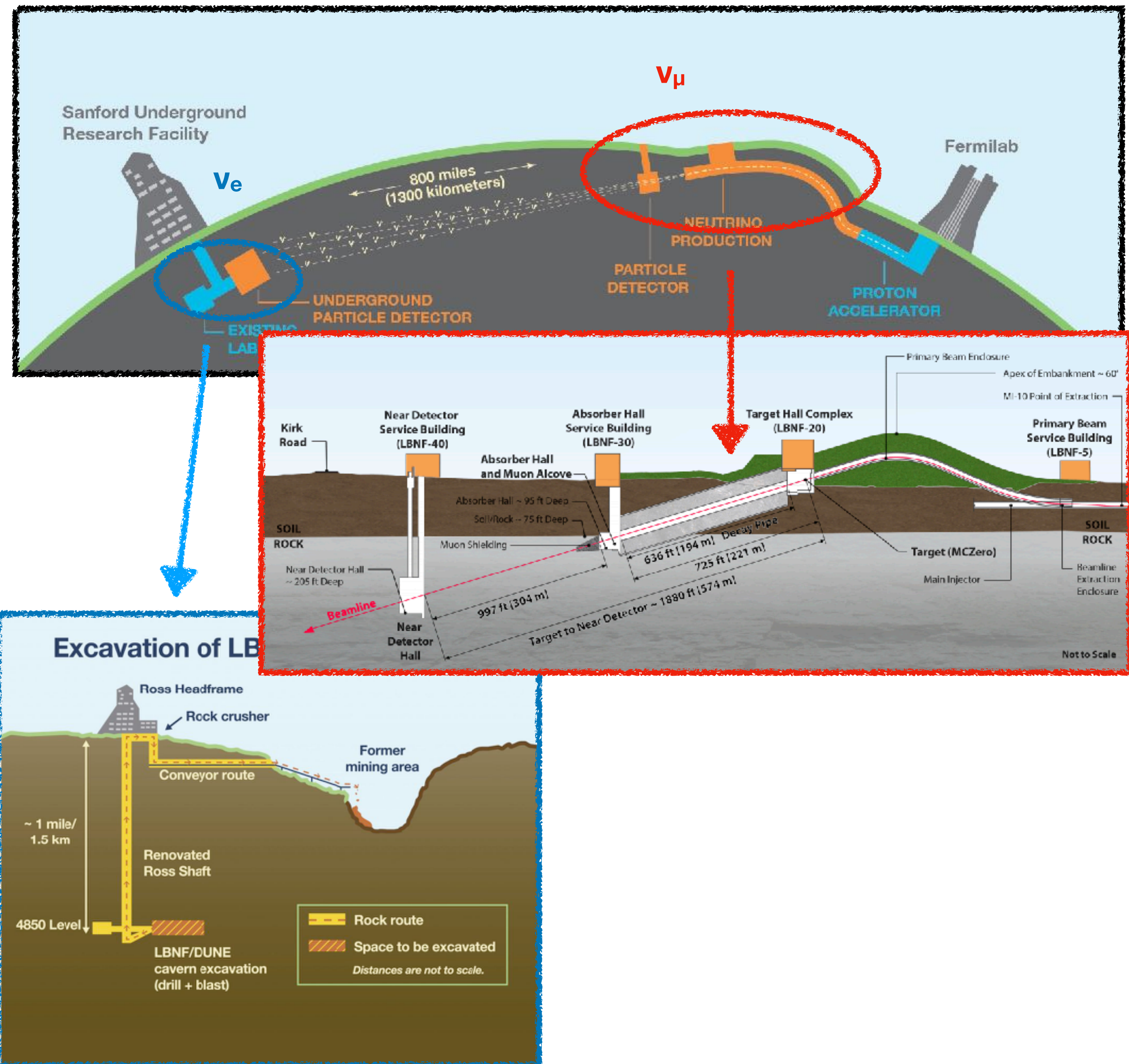
1. The DUNE experiment
2. The Near Detector concept
3. The ND-Gar detector
4. The DUNE Prism concept
5. Outlook and Conclusion



# The DUNE Experiment.

## Deep Underground Neutrino Experiment

- **The next generation neutrino experiment**
  - Long Baseline Neutrino **1300 km**
  - **Wide-band** neutrino beam (GeV range, MW beam)
- Detectors
  - **Near detector** complex at 575 m from neutrino source
  - **Far Detector** made of 4 x 10 kt fiducial mass LAr detector

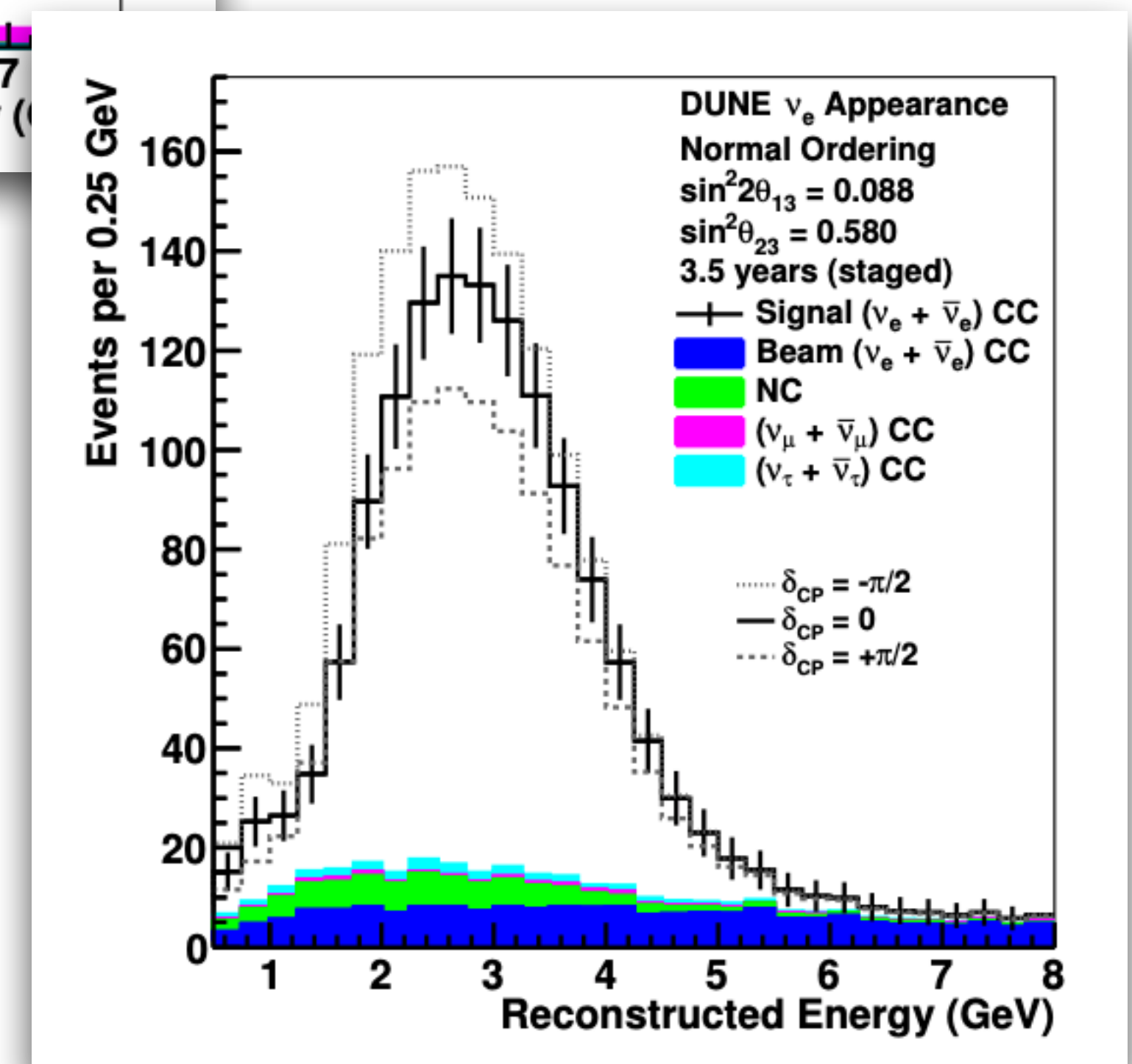
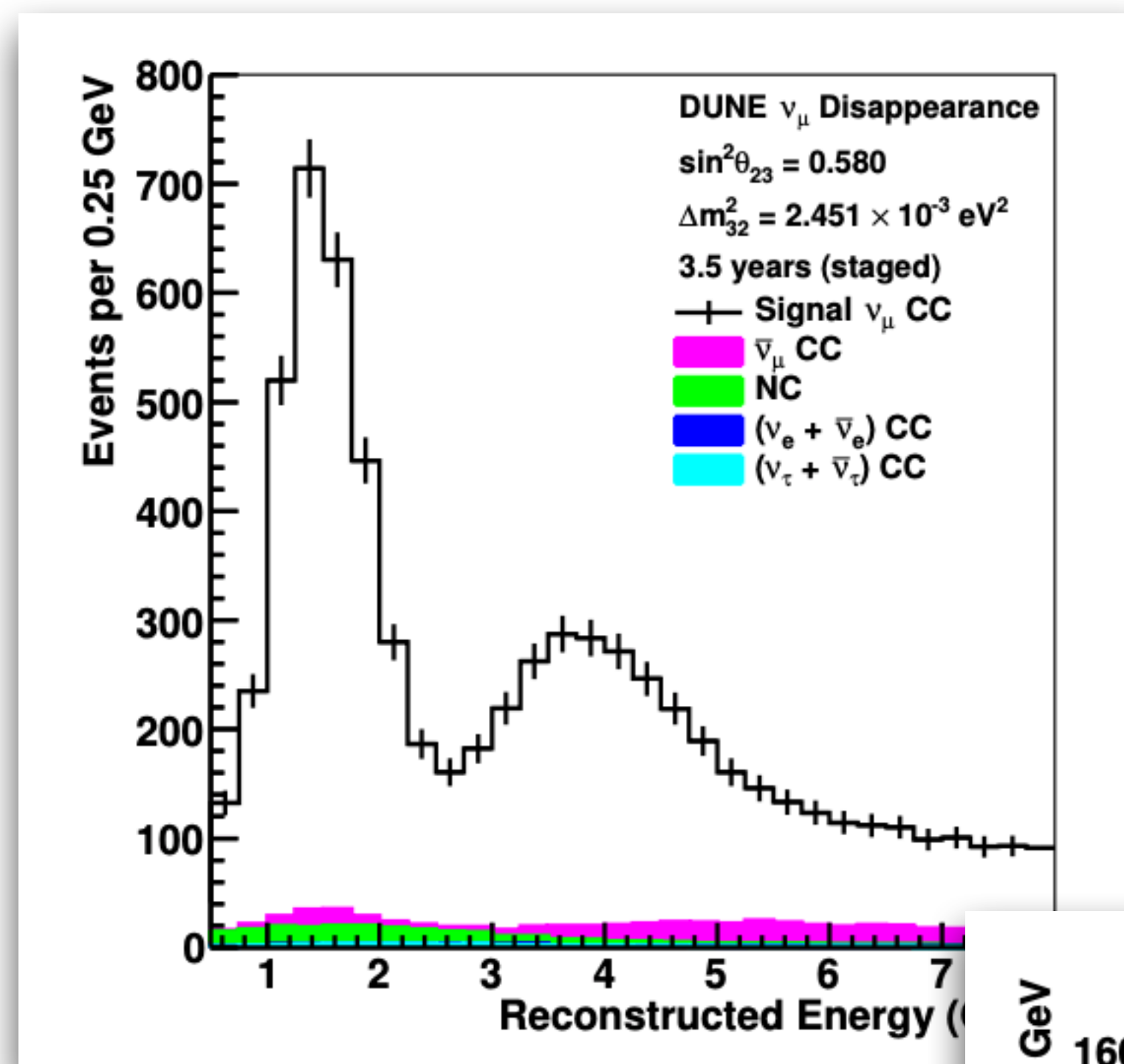




# The DUNE Experiment.

## Deep Underground Neutrino Experiment

- Goals:
  - Neutrino oscillations**  $\implies$  measure  $\nu_\mu$  disappearance and  $\nu_e$  and  $\nu_\tau$  appearance (both FHC and RHC)
  - Measure  $\delta_{CP}$  over 75% of the phase space with a **precision up to  $5\sigma$** , determine mass hierarchy and precise measurement of the mixing angles ( $\theta_{23}$  octant)
  - Beyond the SM physics*: neutrino tridents, DM, sterile neutrinos...
  - Other searches: neutrino supernovae, proton decay...



# How to measure neutrino oscillations?

## Ups and downs of neutrinos

- Want to measure oscillation probability
- However, **detector effects**  $\implies$  need for unfolding / not so easy to cancel systematics
- In reality, this is not easy, need to understand
  - **The neutrino flux**
  - **Cross-section ratios**
  - **Extrapolation near to far**
  - **Detector effects (near and far)**
  - **Relation true to reco neutrino energy**

$$P_{\nu_{\mu} \rightarrow \nu_e}(E_{\nu}) = \frac{\phi_{\nu_e}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{far, no-osc}(E_{\nu})} = \frac{\phi_{\nu_e}^{far}(E_{\nu})}{\phi_{\nu_{\mu}}^{near}(E_{\nu}) * F_{far/near}(E_{\nu})}$$

$$\frac{dN_{\nu}^{det}}{dE_{rec}} = \int \phi_{\nu}^{det}(E_{\nu}) * \sigma_{\nu}^{target}(E_{\nu}) * T_{\nu_{\mu}}^{det}(E_{\nu}, E_{rec}) dE_{\nu}$$

$$\frac{dN_{\nu_e}^{far}}{dE_{\nu}} \bigg/ \frac{dN_{\nu_{\mu}}^{near}}{dE_{\nu}} = P_{\nu_{\mu} \rightarrow \nu_e}(E_{\nu}) * \frac{\sigma_{\nu_e}^{Ar}(E_{\nu})}{\sigma_{\nu_{\mu}}^{Ar}(E_{\nu})} * F_{far/near}(E_{\nu})$$

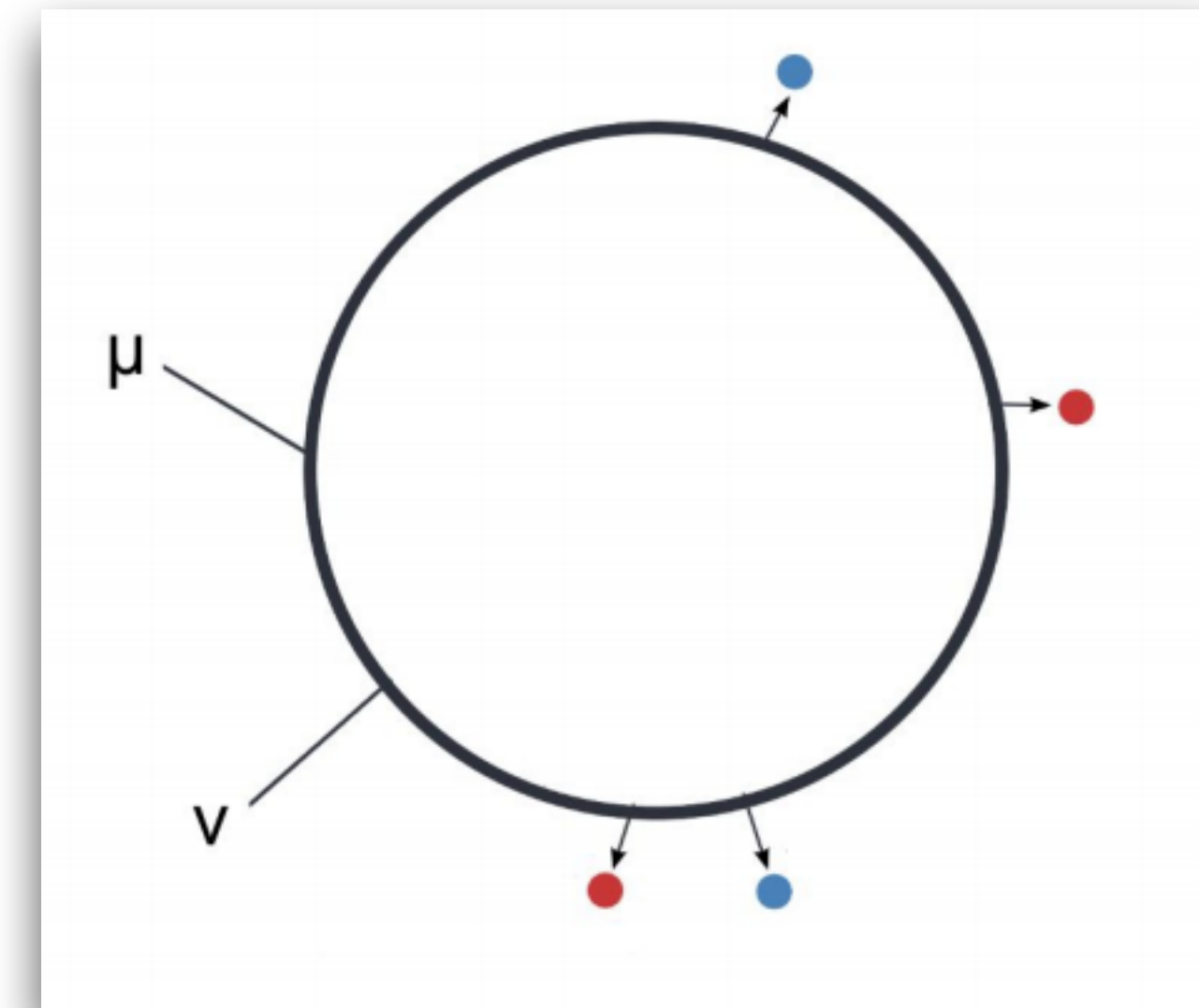
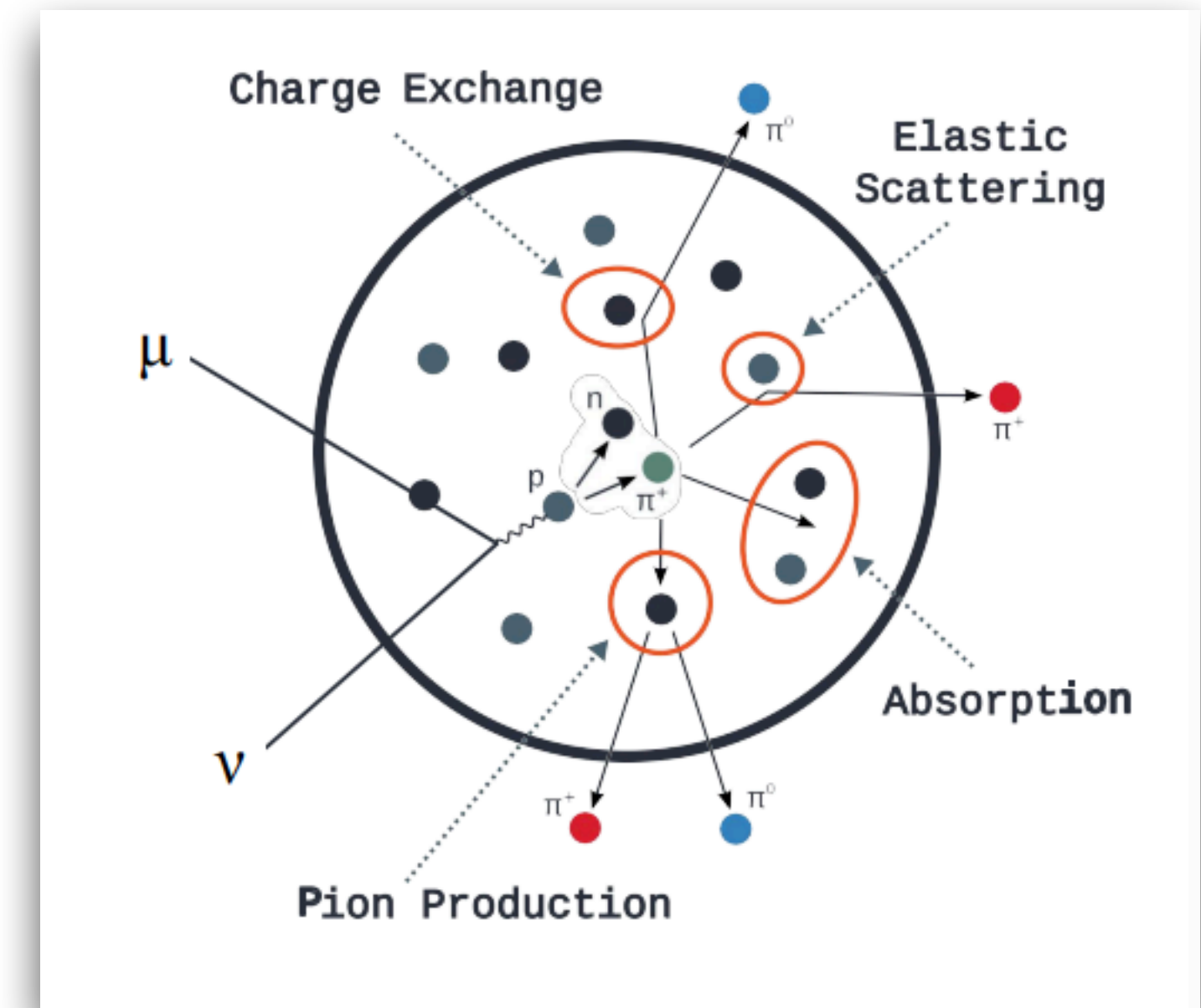
$$\frac{dN_{\nu_e}^{near}}{dE_{\nu}} \bigg/ \frac{dN_{\nu_{\mu}}^{near}}{dE_{\nu}} = \frac{\sigma_{\nu_e}^{Ar}(E_{\nu})}{\sigma_{\nu_{\mu}}^{Ar}(E_{\nu})} * \frac{\phi_{\nu_e}^{near}(E_{\nu})}{\phi_{\nu_{\mu}}^{near}(E_{\nu})}$$

# Example: Impact of cross-section measurement.

What can happen

## Effect on CP sensitivity

- Cross-sections are directly used to reconstruct the neutrino energy
  - **Final state interactions (FSI)** and **nuclear effects** within the nuclei can make different interaction channels with the same topology



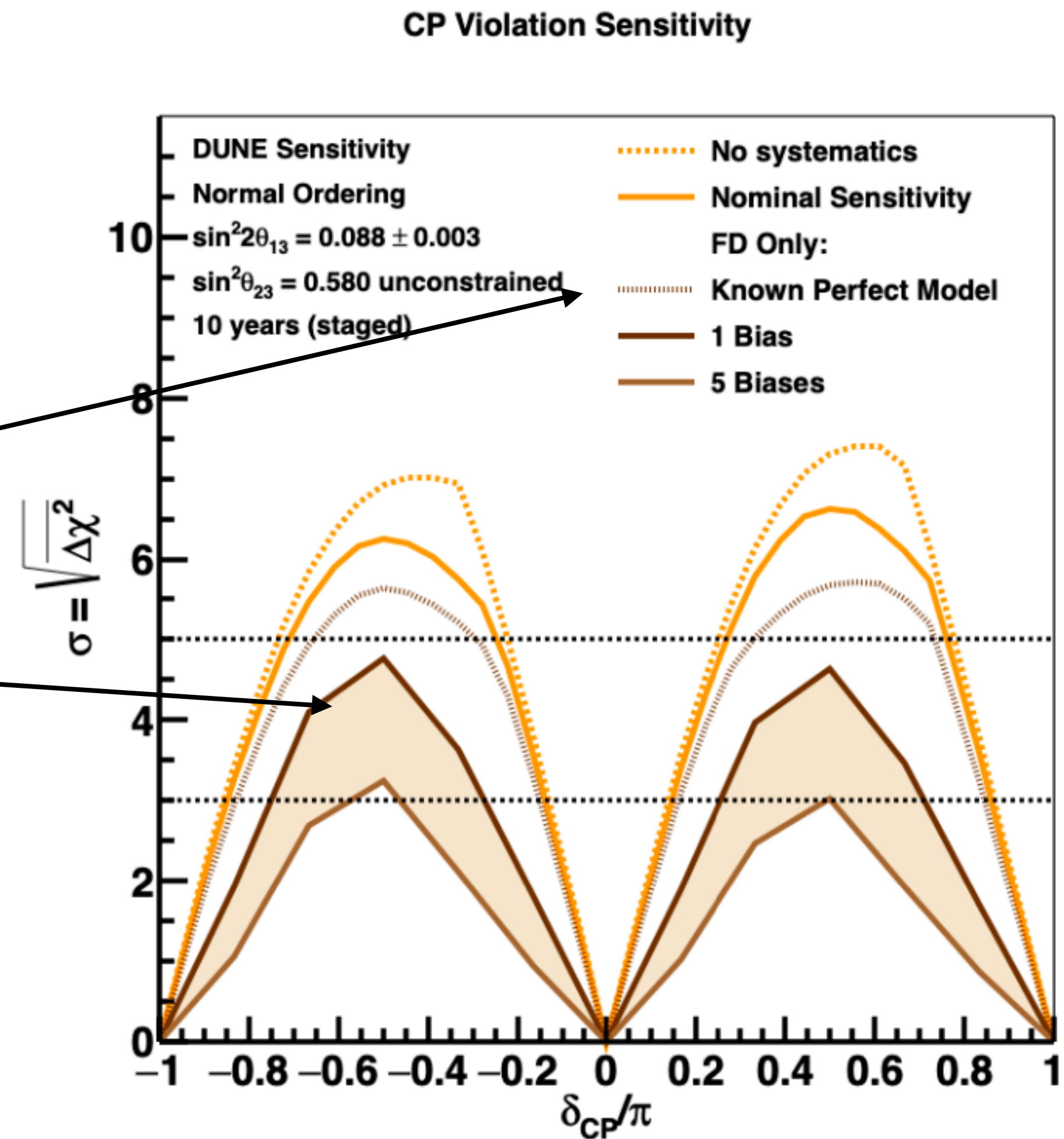
What you observe



# Example: Impact of cross-section measurement.

## Effect on CP sensitivity

- Cross-sections are directly used to reconstruct the neutrino energy
  - **Final state interactions (FSI)** and **nuclear effects** within the nuclei can make different interaction channels with the same topology
- Cross-section systematics are often the **dominant** contribution in the error on neutrino oscillations
  - Sensitive to the **generator model (GENIE)**
  - **Different** but plausible **model**
- Strong possibility of other viable models  $\implies$  **more model bias** will need to be included **reducing the sensitivity** further

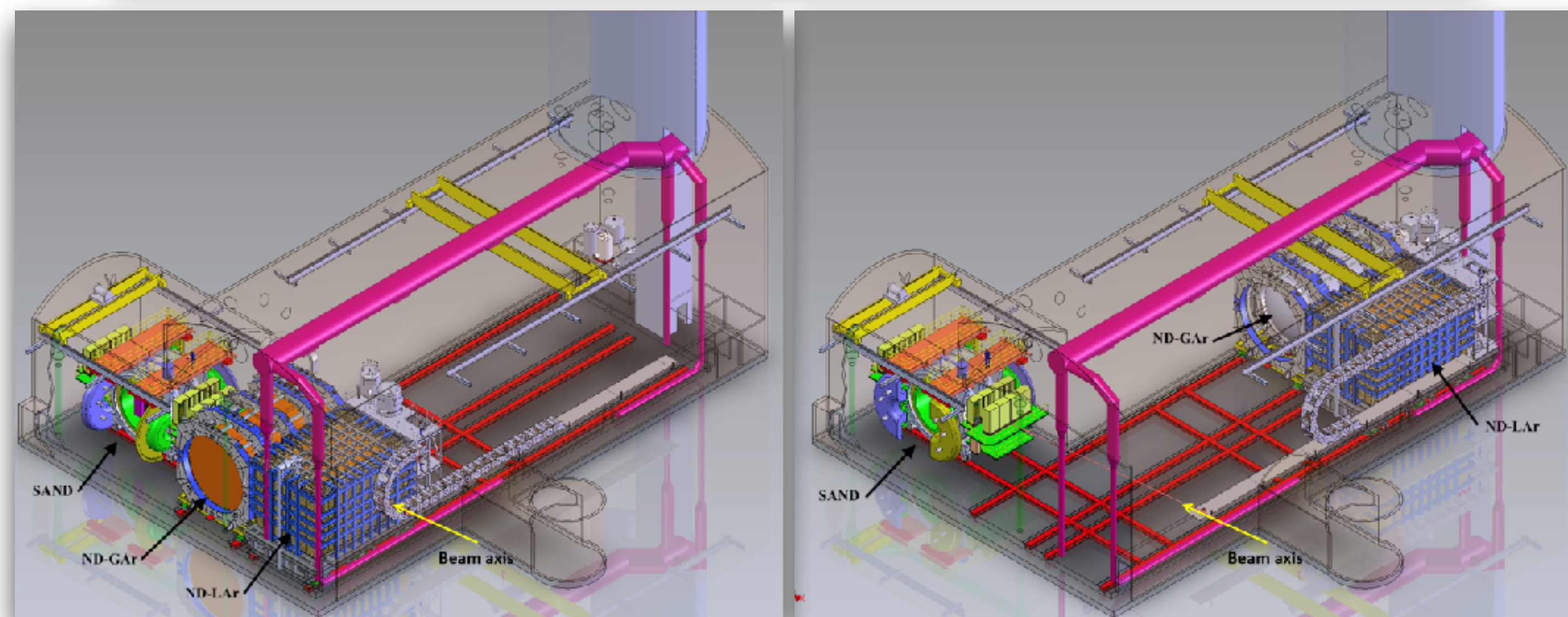
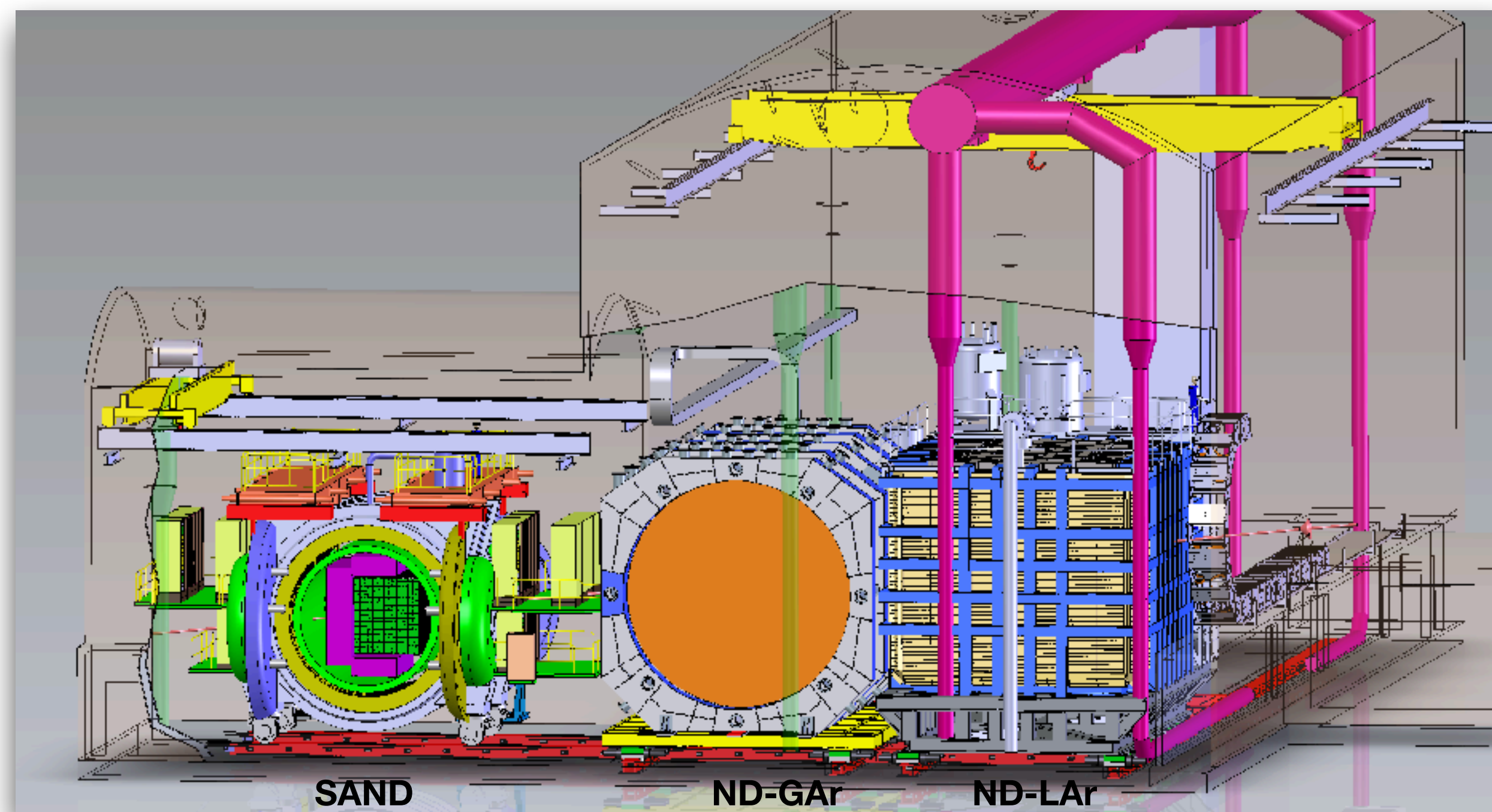




# The DUNE ND Complex.

## A crucial part of DUNE

- **ND-LAr**
  - Highly modular Liquid Argon Time Projection Chamber with pixelated readout (50 t), similar to FD, primary target
- **ND-GAr**
  - High-pressure gas Argon TPC (1 t) surrounded by a high performance ECAL, a magnet and a muon system, muon spectrometer and constrain nuclear modelling on Argon
- **SAND**
  - Highly granular plastic target (8 t) surrounded by trackers and ECAL inside a magnet, on-axis beam spectrum monitor
- **ND-LAr** and **ND-GAr** can move off-axis  $\Rightarrow$  **DUNE-Prism** concept



On-axis

DUNE-Prism



# ND Requirements.

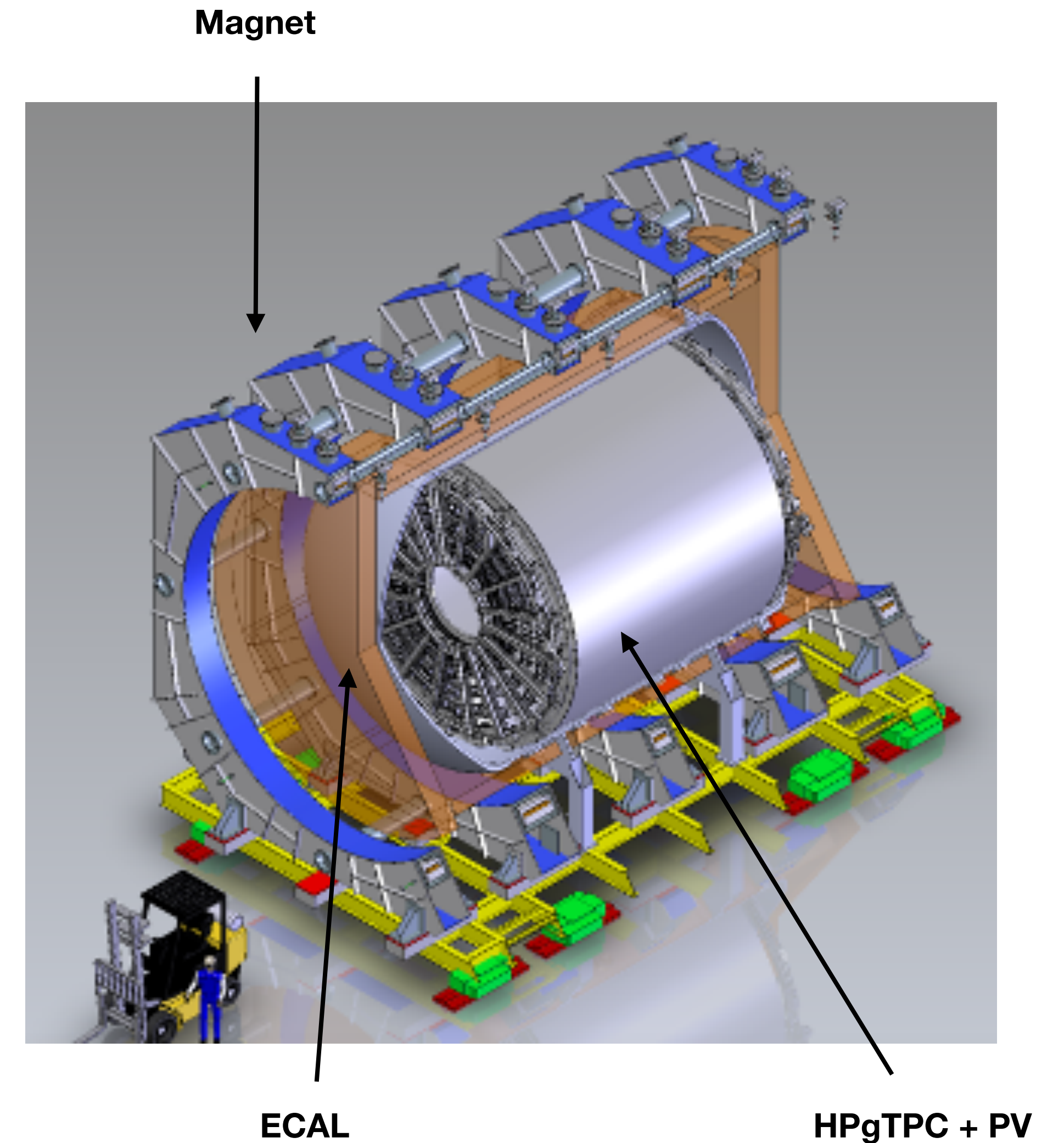
## The goals of the ND

- Predict neutrino flux at the FD
  - ND must be able to **predict observable** at the FD
- Transfer measurements to the FD
  - ND measurements must be **transferable** to the FD to minimise systematics
- Constrain the neutrino cross-section model
  - **Reduces systematics** of the FD response to neutrino energy/flavor
- Measure the neutrino flux
  - Constrain the **flux modelling**
- Obtain data with different neutrino fluxes (**DUNE-Prism**)
  - ND must be able to verify the **robustness of model** predictions with different fluxes
- Monitor the neutrino beam
  - Detect **variations** in the beam flux
- Operate in a **high rate** environment

# The ND-GAr detector.

## The CDR baseline

- **1 ton fiducial mass** Gas Argon TPC at 10 bar inside a pressure vessel
- High-performance **ECAL** using copper as absorber and scintillator tiles/strips with SiPM readout
- A super-conducting **magnet** and a complementary **muon system**
- Acts as a **spectrometer** for muons exiting the ND-LAr
- **Lower target energy threshold** ( $\sim$ MeV compared to  $\sim$ 30 MeV)
  - improve understanding of nuclear models to reduce systematics
- Key for some **channels** that are **hard to distinguish** in the ND/FD-LAr, such as multi-pi final states

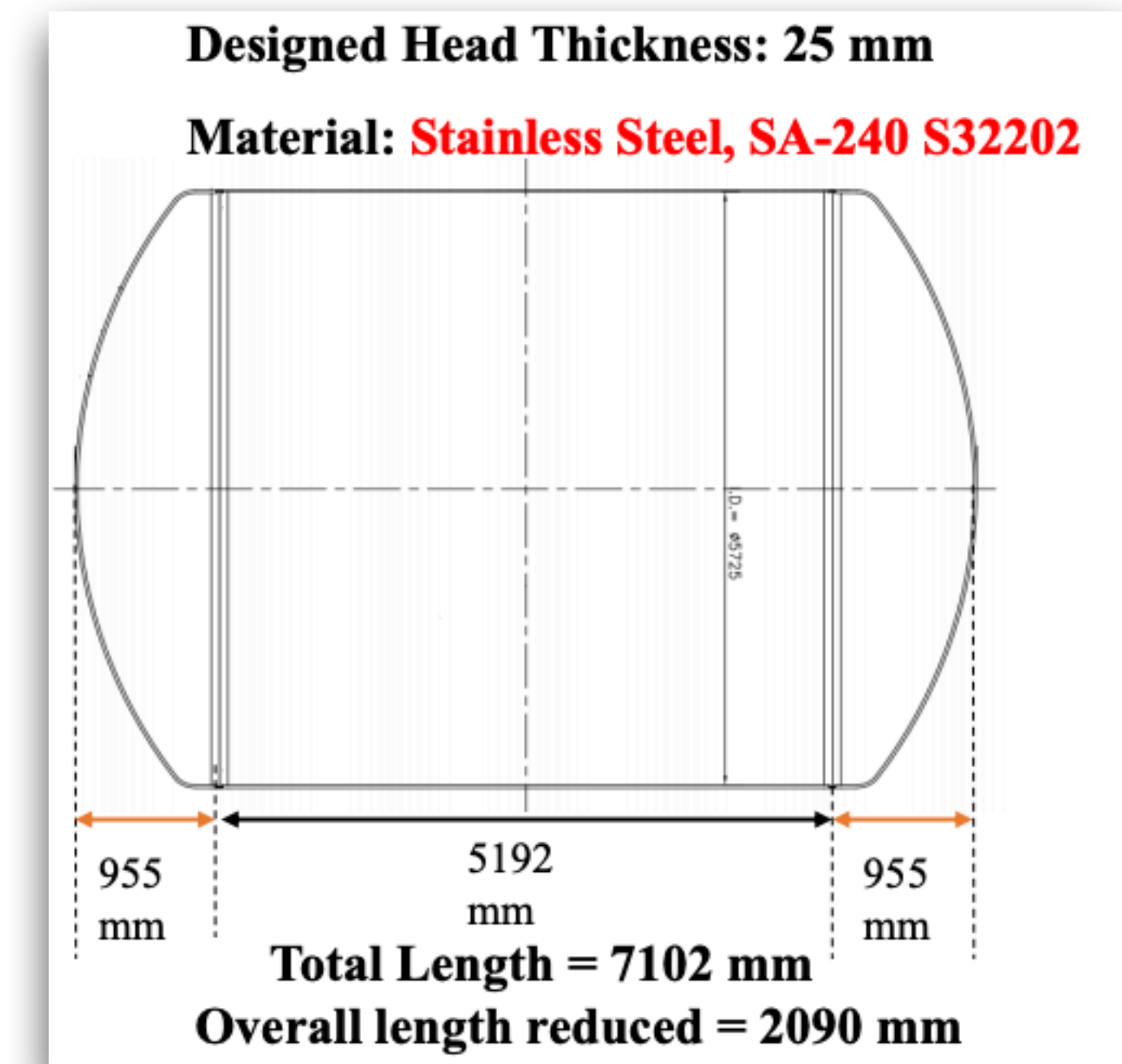
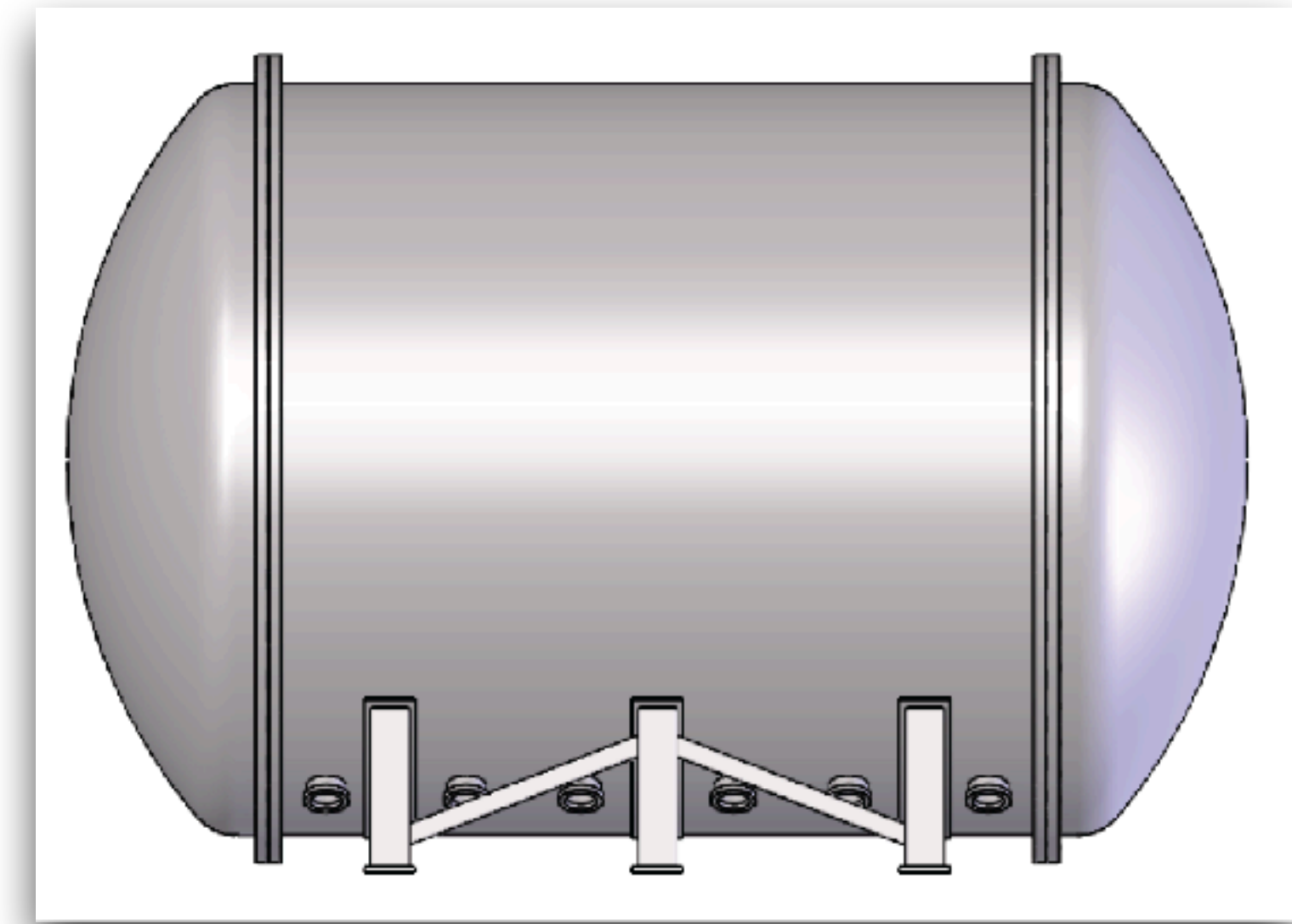




# The ND-GAr detector.

## Fast evolution of the concept

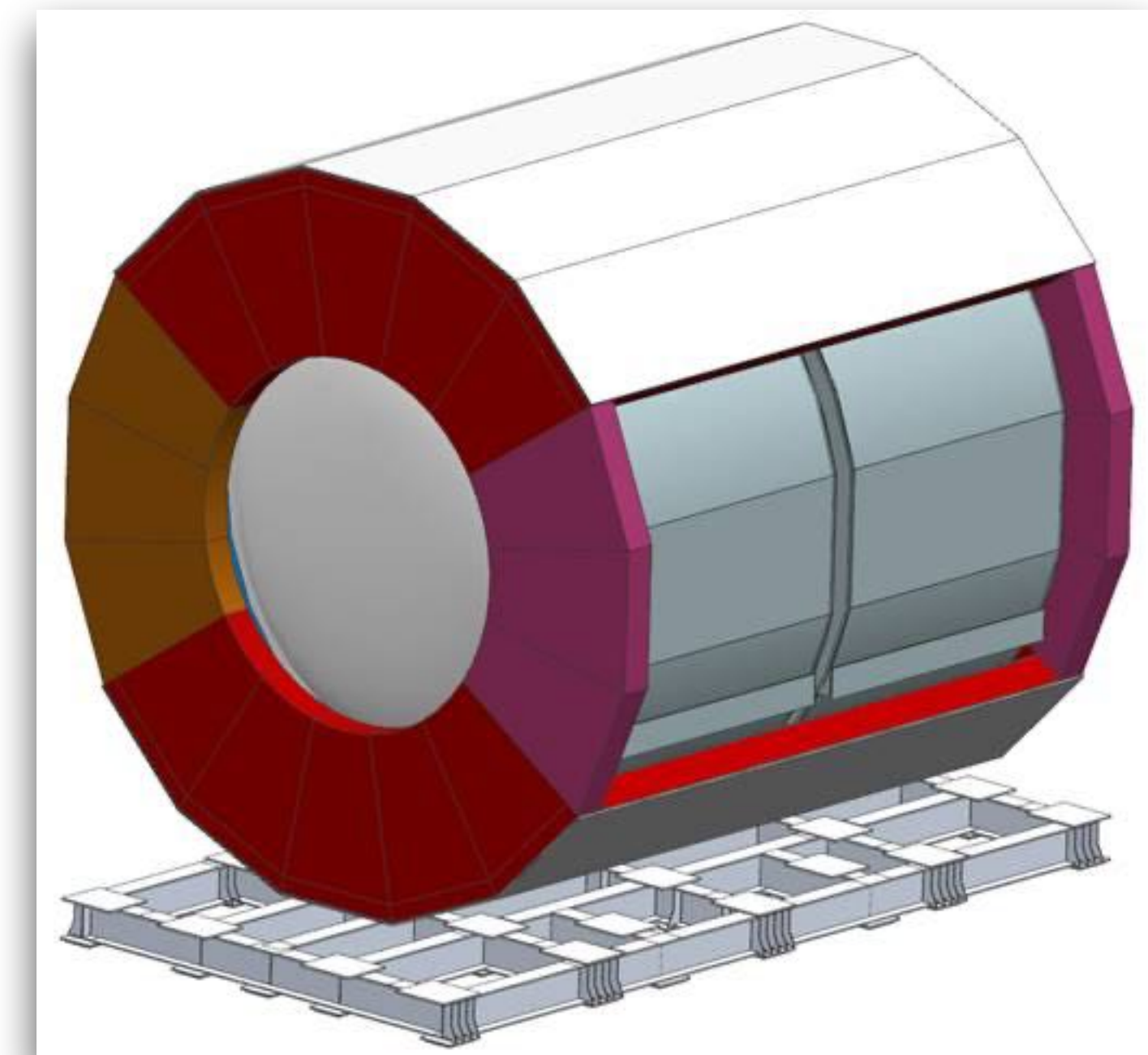
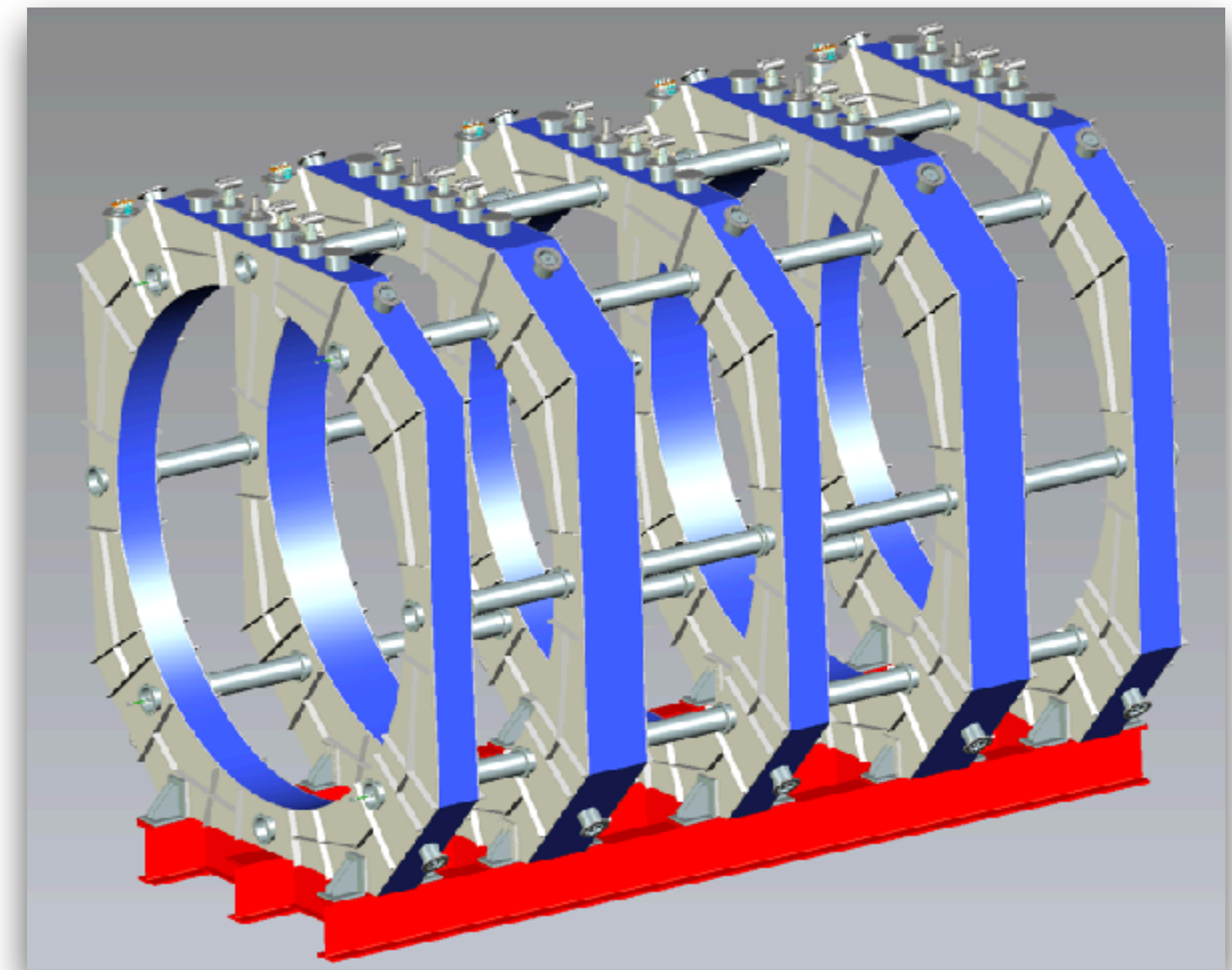
- Engineering design of the pressure vessel is evolving fast
  - Understand **mechanical constrains** due to pressure, welding etc...
- Barrel thickness designed to be  $\sim 0.5 X_0$  ( $\sim 4.4$  cm Al)
  - Due to mechanical constrains  $\Rightarrow$  pressure vessel head need to be  $\sim 1 X_0$  (if using Al)
  - Stainless steel being investigated  $\Rightarrow$  would reduce overall ND-GAr length by  $\sim 2$ m with a thickness of 25 mm ( $\sim 1.4 X_0$ )
- This has an impact on the overall ND-GAr design especially on the **ECAL**
  - ECAL endcaps need to be **inside** the pressure vessel



# The ND-GAr detector.

## Fast evolution of the concept

- Engineering design of the ND-GAr magnet has changed a lot in the past few months
  - Allowance in the overall detector radial size reduced
  - Impact of the stray B-field on SAND (forces on the SAND yoke)
  - Cost
- Original design using 3 to 5 **Helmholtz-coils**
- New design investigated using a **solenoid with a partial return yoke (SPY)**
  - Reduces **material** in front of the ND-GAr
  - Reduces **overall size** in diameter
  - Reduces **stored energy, stray field and cost**
  - Can include a **muon system** inside the yoke

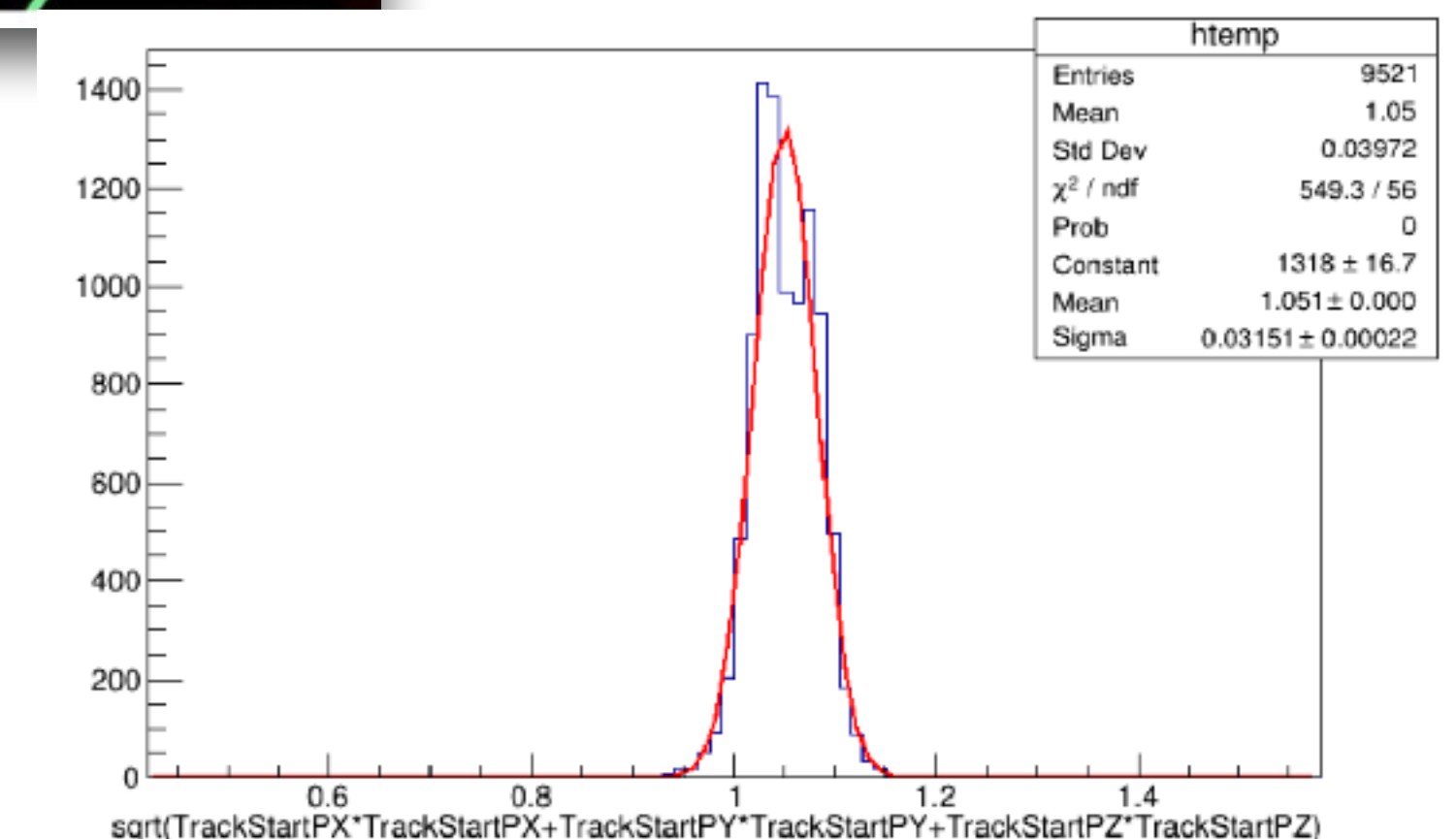
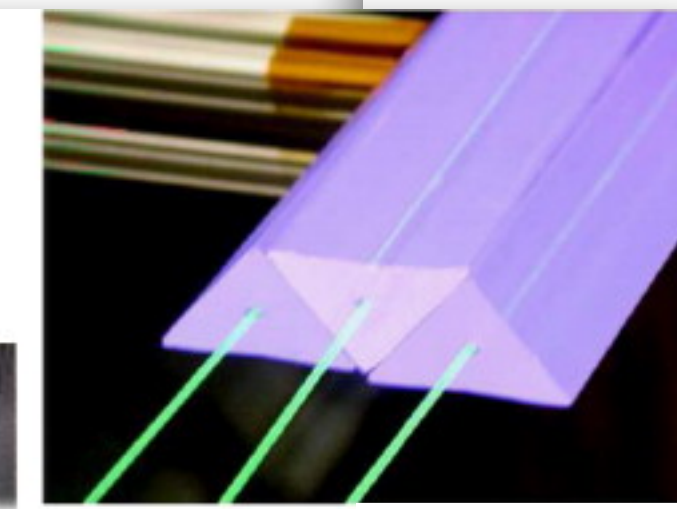
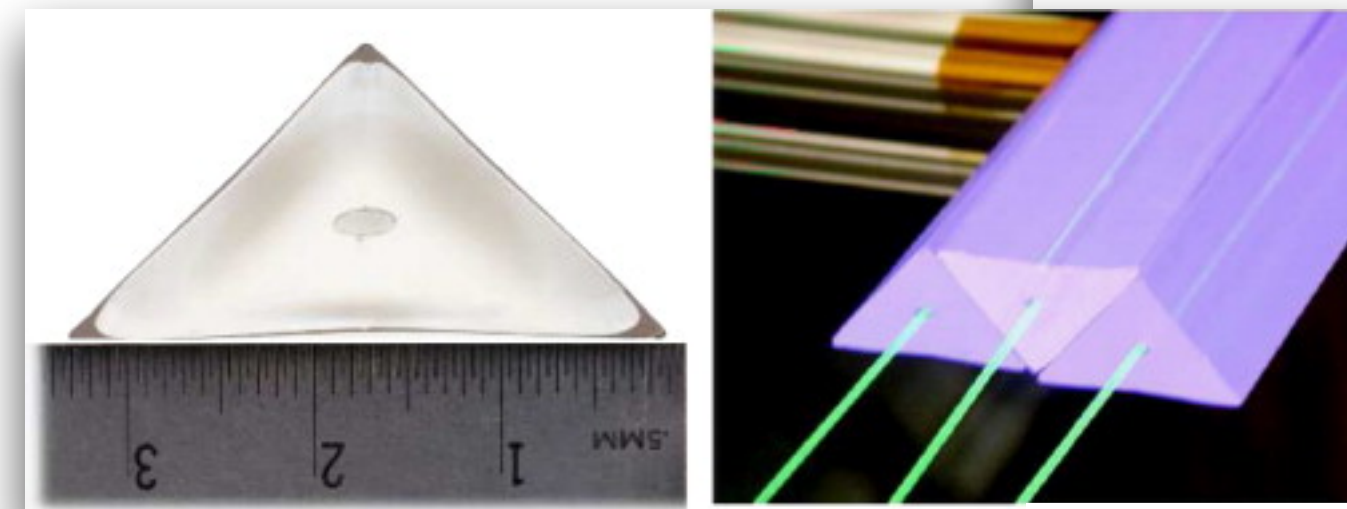
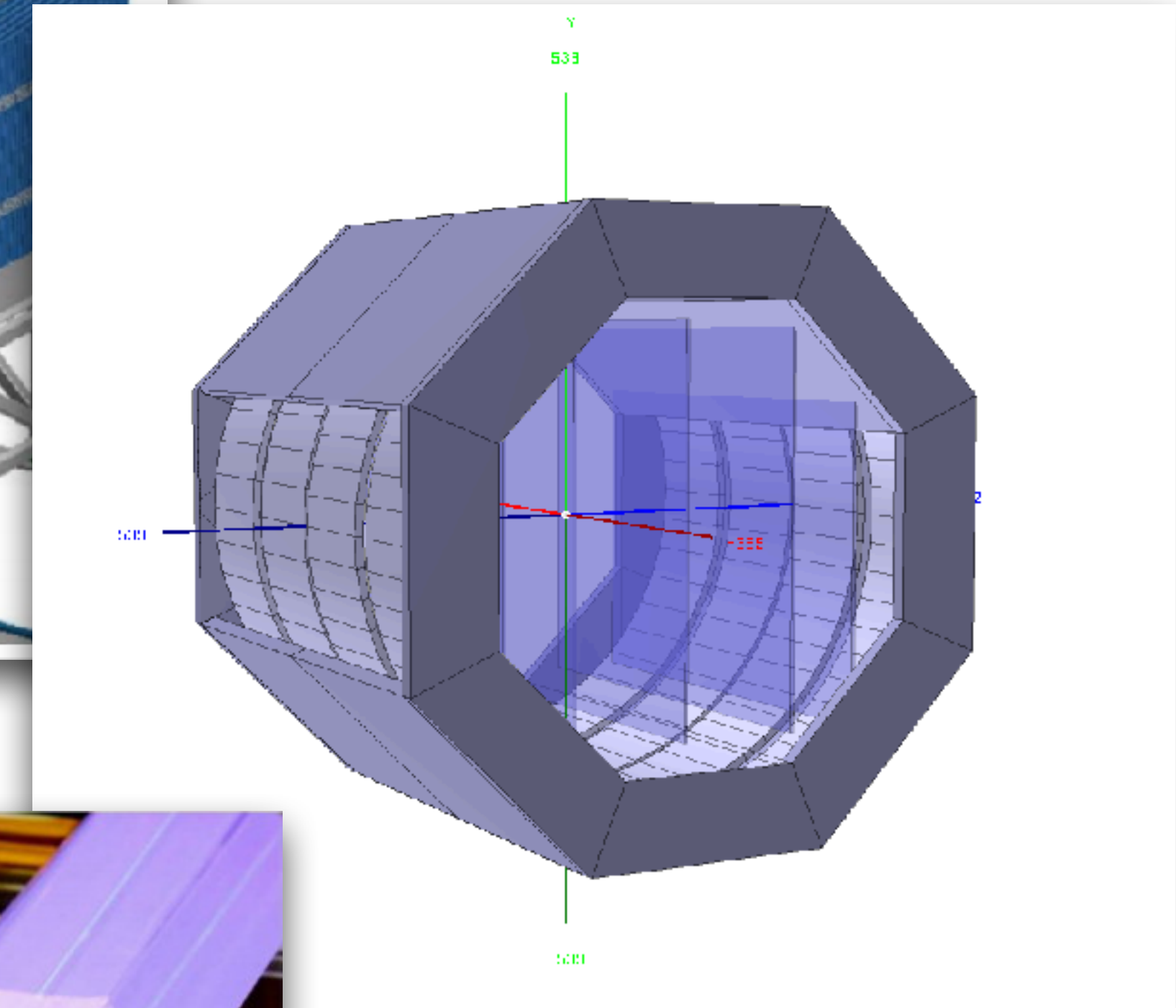
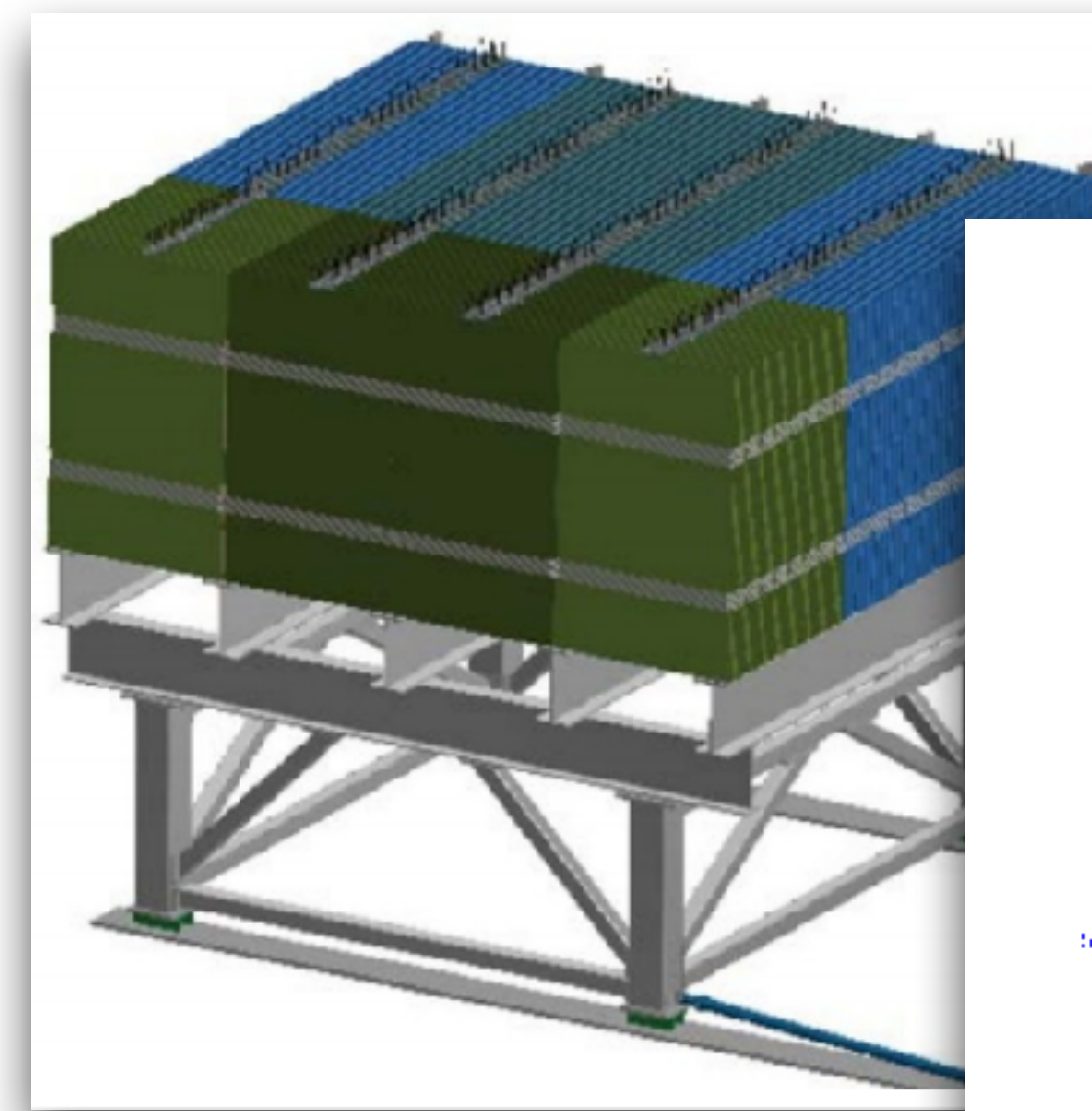




# The ND-GAr detector.

## Fast evolution of the concept

- The full ND-GAr will **not** be available on day-one
- Needs to be able to measure muons exiting from the ND-LAr
- Needs to meet the DUNE **physics requirements** (3+ years running)
  - **$3\sigma$  CP violation sensitivity** at  $\delta = \pi/2$
  - **Mass-ordering** determination
- **Alternatives** are being investigated
  - Temporary muon spectrometer system (**TMS**) using magnetised steel and scintillator planes
  - ND-GAr with the **SPY** magnet system + 5 scintillator tracker planes (Minerva-like)

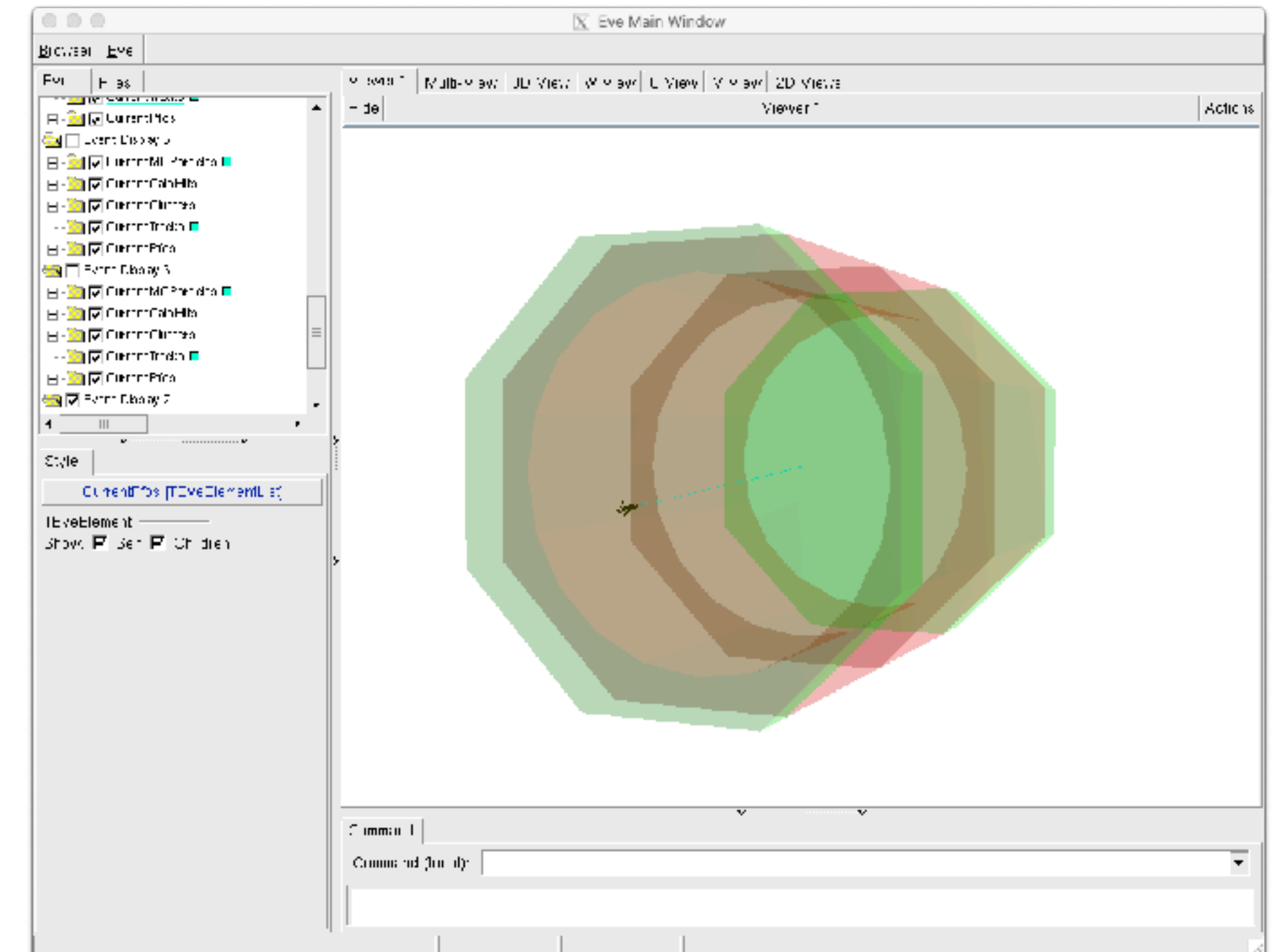


# The ND-GAr detector.

## Integration of Pandora

- Software framework is evolving a lot also
- Full simulation and reconstruction chain being **harmonised** within the ND groups
  - Agreements on **generator** (GENIE) and **simulation** wrapper (edep-sim)
- **PandoraPFA** algorithms are used to reconstruct the neutrino event in the FD
  - Why not also using it for the **ND**?
- ND-LAr and ND-GAr groups are interested in implementing Pandora into their reconstruction chain
  - ND-GAr  $\Rightarrow$  PandoraPFA is now **implemented** due to large similarities with ILD (using for now LCContent algorithms)
  - Pandora group is interested in providing support into the implementation of PandoraPFA for the ND complex

```
MSG
Begin processing the 1st record. run: 1 subRun: 0 event: 1 at 09-Jul-2020 13:11:2
MSG-i PandoraInterface - produce: PandoraInterface:pandora@BeginModule 09-Jul-2
MSG
Running Algorithm: Alg0001, CaloHitPreparation
Running Algorithm: Alg0002, EventPreparation
Running Algorithm: Alg0003, ClusteringParent
--> Running Algorithm: Alg0004, ConeClustering
--> Running Algorithm: Alg0005, TopologicalAssociationParent
-----> Running Algorithm: Alg0006, LoopingTracks
-----> Running Algorithm: Alg0007, BrokenTracks
-----> Running Algorithm: Alg0008, ShowerMipMerging
-----> Running Algorithm: Alg0009, ShowerMipMerging2
-----> Running Algorithm: Alg0010, BackscatteredTracks
-----> Running Algorithm: Alg0011, BackscatteredTracks2
-----> Running Algorithm: Alg0012, ShowerMipMerging3
-----> Running Algorithm: Alg0013, ShowerMipMerging4
-----> Running Algorithm: Alg0014, ProximityBasedMerging
-----> Running Algorithm: Alg0015, TrackClusterAssociation
-----> Running Algorithm: Alg0016, ConeBasedMerging
-----> Running Algorithm: Alg0017, TrackClusterAssociation
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-----> Running Algorithm: Alg0019, TrackClusterAssociation
-----> Running Algorithm: Alg0020, SoftClusterMerging
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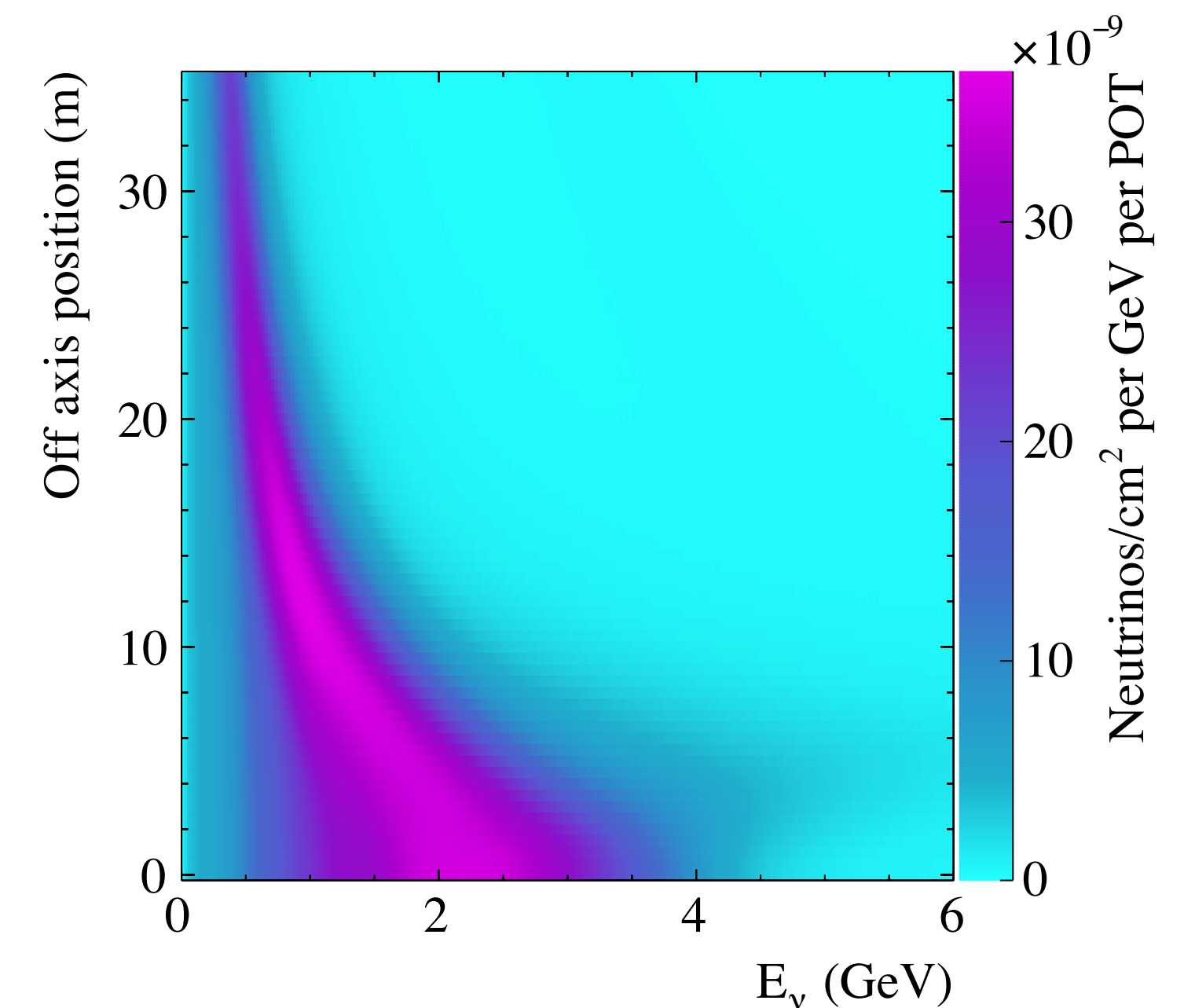
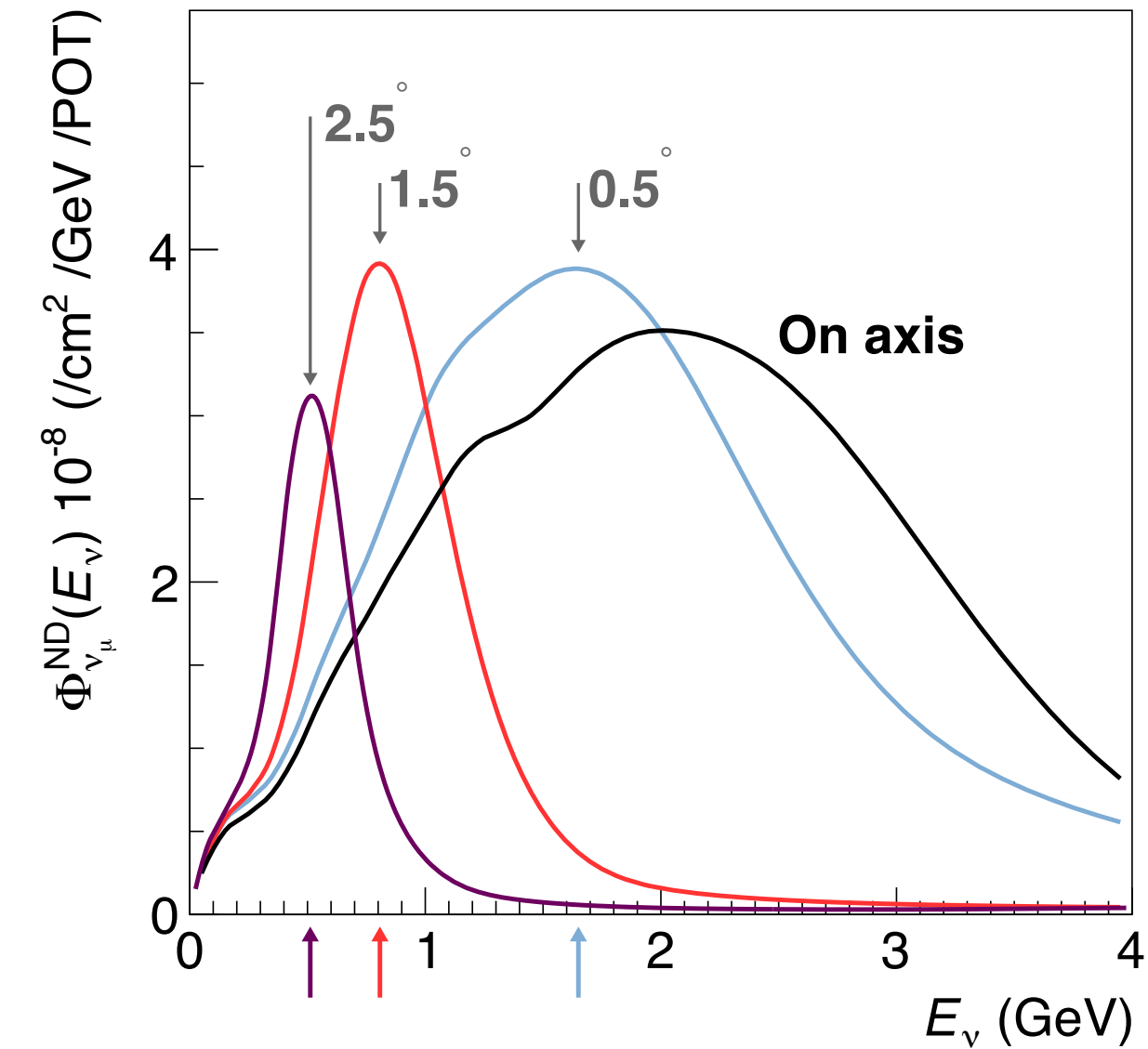




# The DUNE Prism concept.

## Motivation

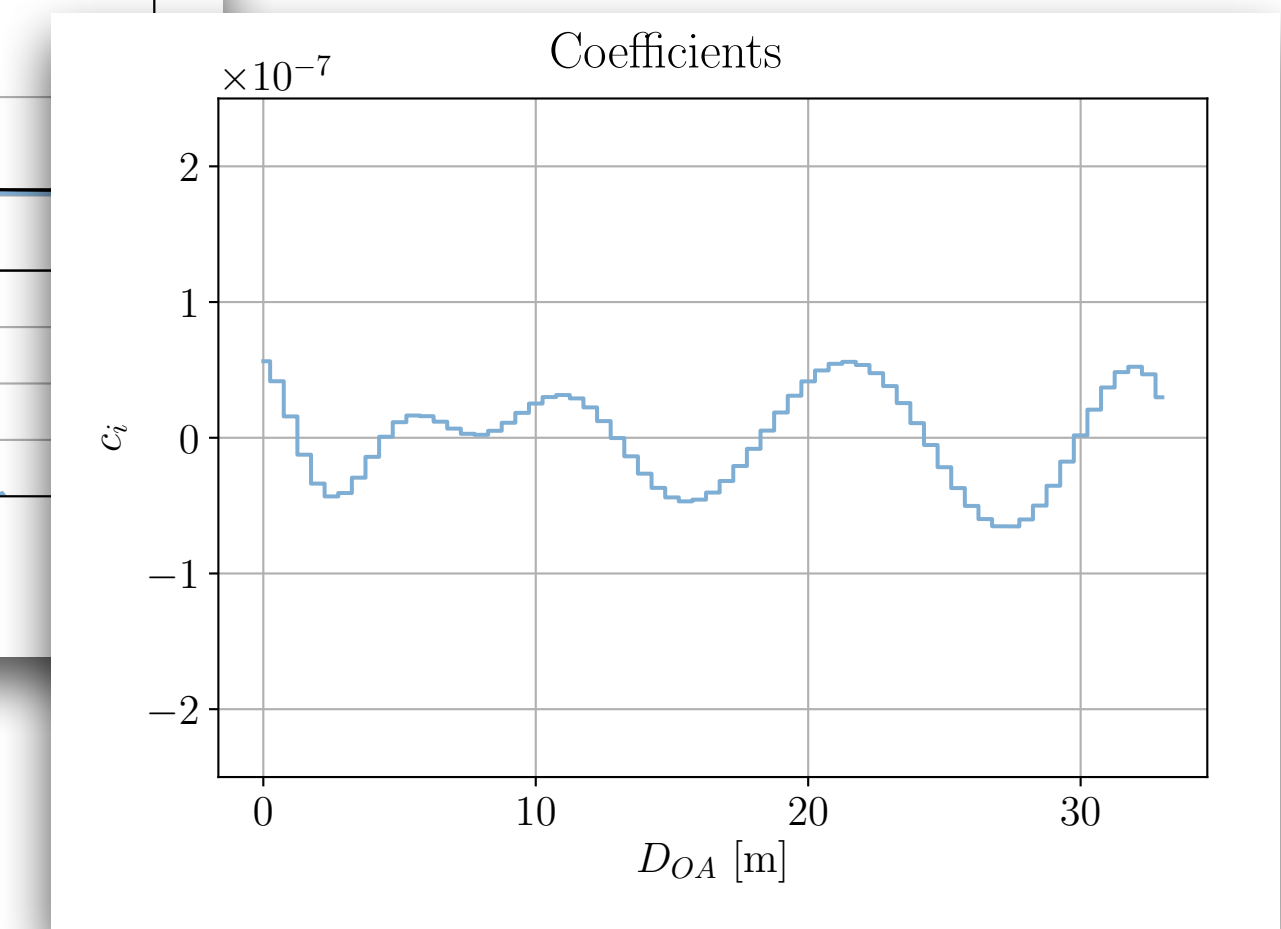
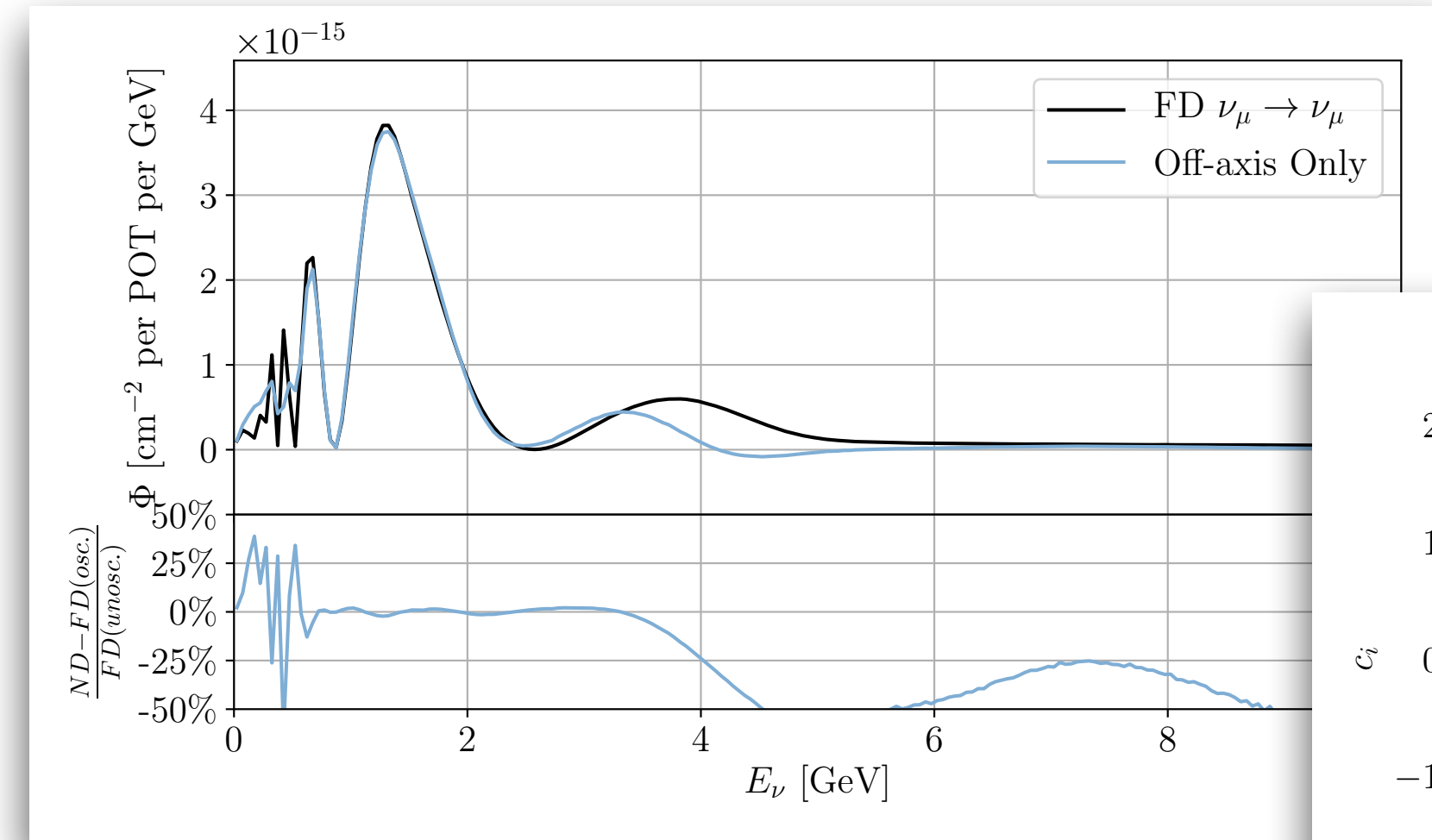
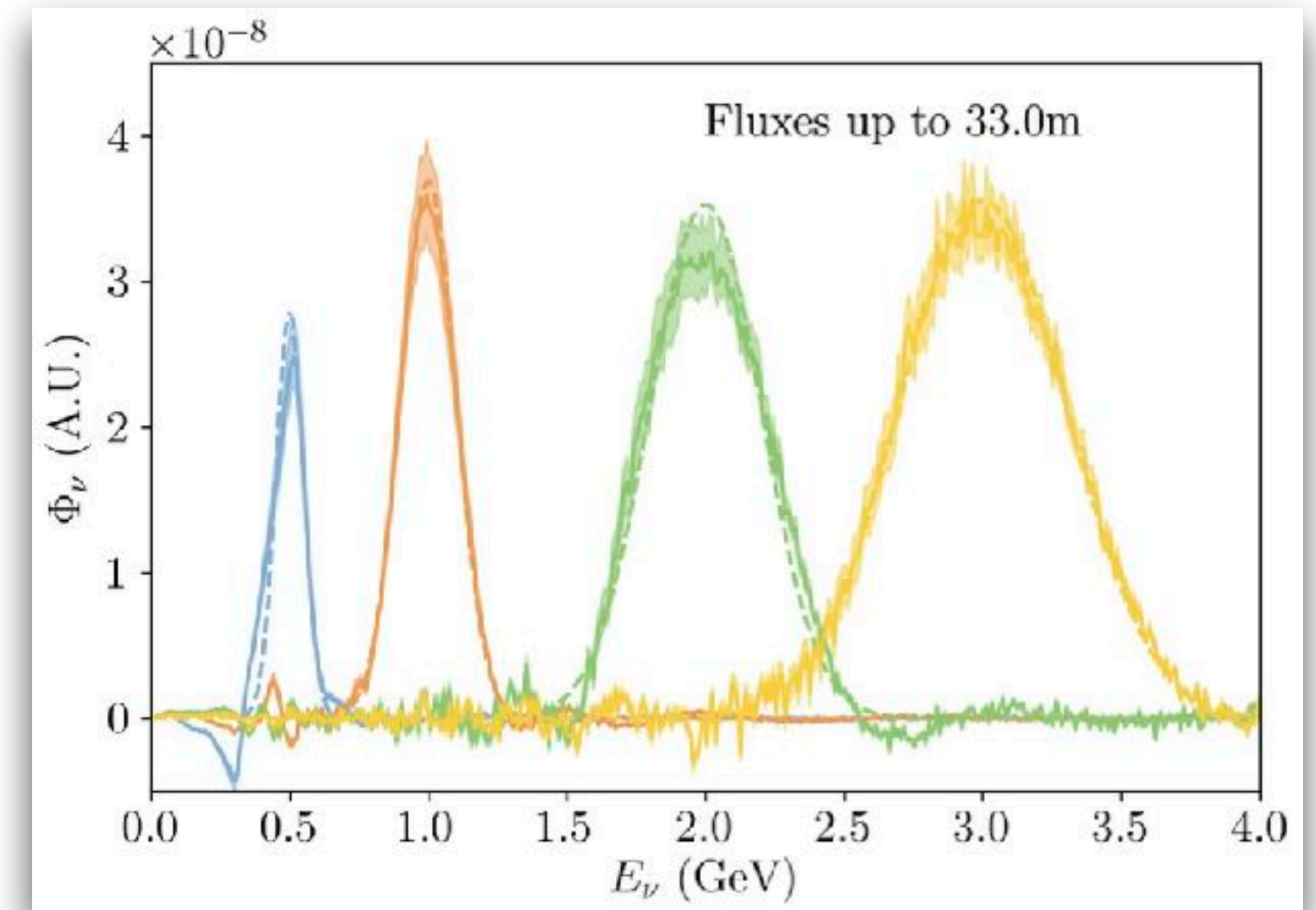
- Neutrino energy depends on the interaction model
- Constraining neutrino interactions uncertainty is difficult due to the lack of a complete model (**unknown** biases that can be difficult to estimate)
- DUNE neutrino beam
  - Peak energy decrease as function of the angle relative to the beam  $\Rightarrow$  generally used in **narrow-band** experiment (T2K, NOvA)
- **DUNE-Prism** concept exploits this
  - Measure various **off-axis** positions
  - Provide an **additional degree of freedom** to constrain systematics uncertainties
- Allow for **data-driven determination** of the relation  $E_{\text{true}} - E_{\text{reco}}$  (less sensitive to neutrino interaction model)



# The DUNE Prism concept.

## Flux matching

- Off-axis positions constitute a **set of fluxes** peaking at different energies across the DUNE neutrino energies
- Using these fluxes, one can mock-up a nearly Gaussian spectra using **linear combination** of the energy spectra
- Going even further, one can directly **construct** the oscillated energy spectrum at the FD for any oscillation parameter
  - This minimises the ND/FD flux differences and associated systematics
  - Gives a set of coefficients that can be applied to any ND observable to get the **FD prediction**
  - In the limit that the model flux is perfect, this gives a **nearly model independent** measurement





# Conclusion.

## A lot of progress

- The DUNE experiment is the next generation neutrino experiment that will probe the neutrino sector to great details
- The DUNE Near Detector complex is crucial to achieve DUNE's physics goals
- The ND-GAr is a necessary complementary detector to the ND-LAr
- Its conceptual design is evolving fast
- The ND group is going toward a unified software framework
- The DUNE-Prism concept is a fundamental part of the DUNE ND program, providing datasets that will enable us to understand to great details neutrino oscillations
- The DUNE ND CDR is currently under review and hopefully will be soon available publicly



**Backup Slides.**

