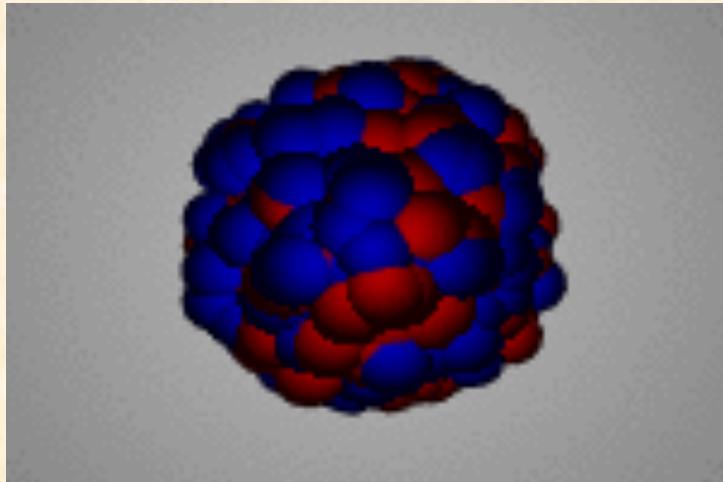




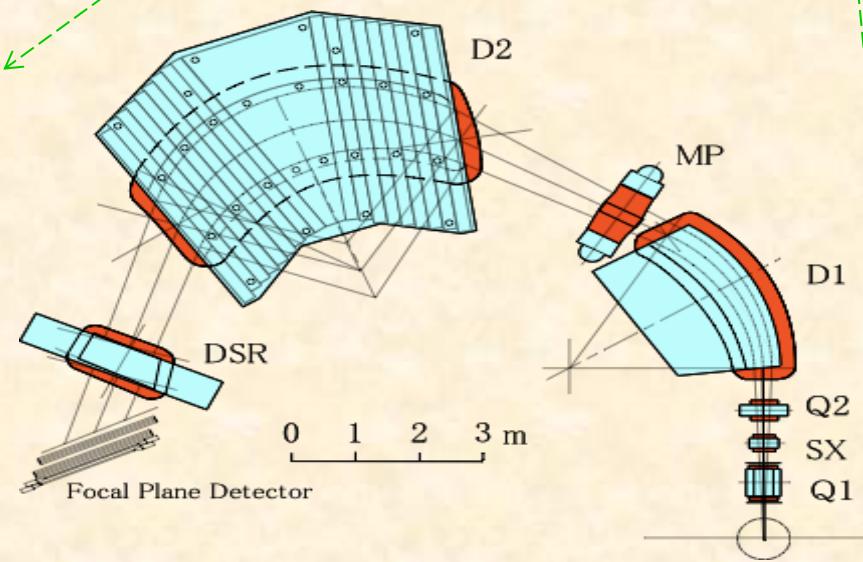
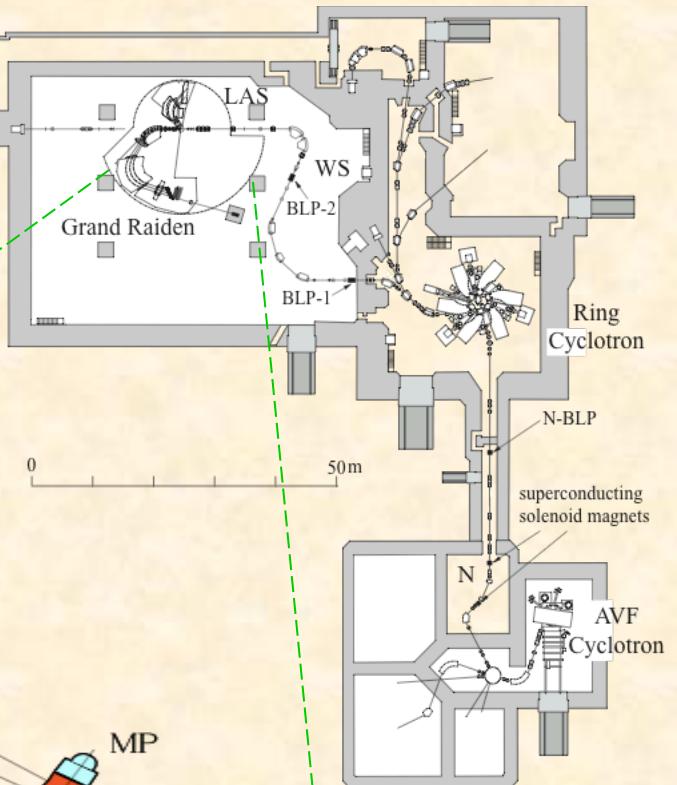
GMR in Stable (and some Unstable) Nuclei: What's New?



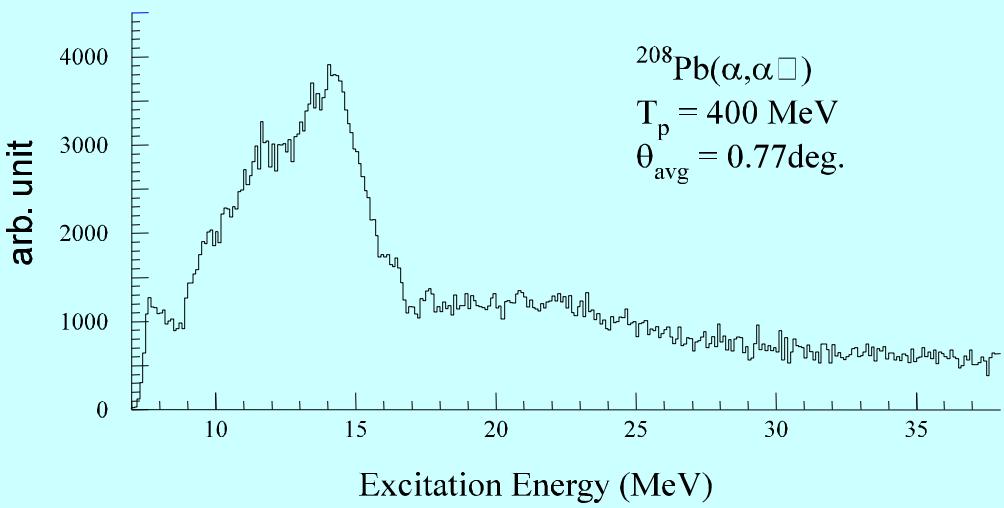
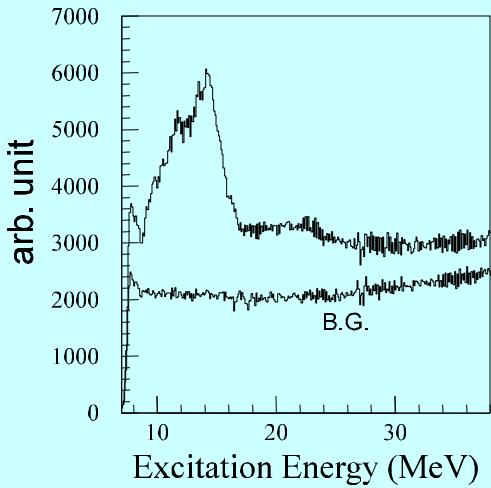
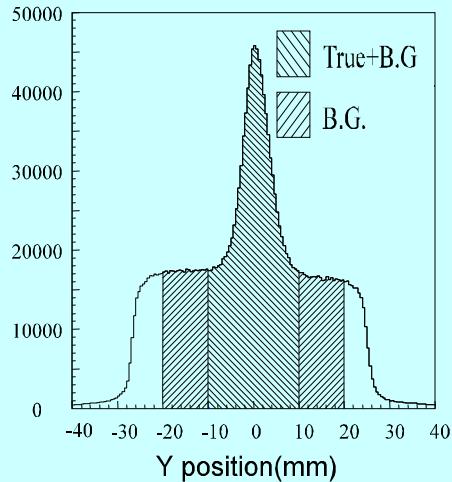
U. Garg
University of Notre Dame

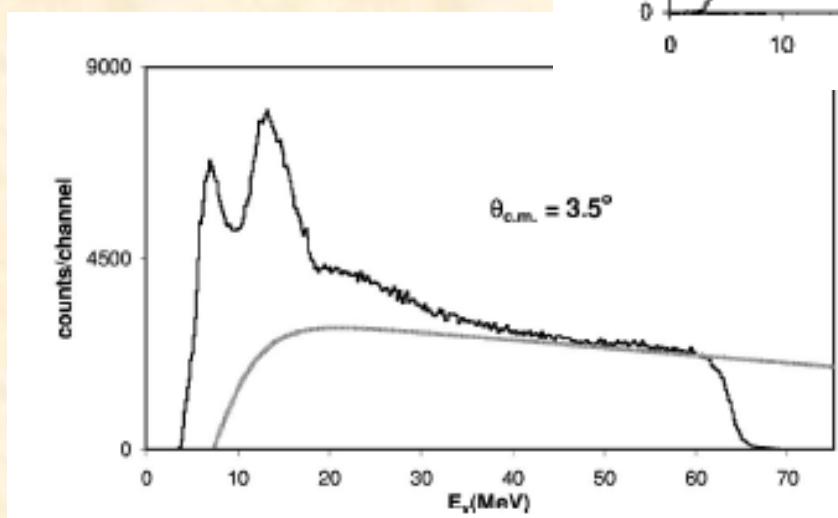
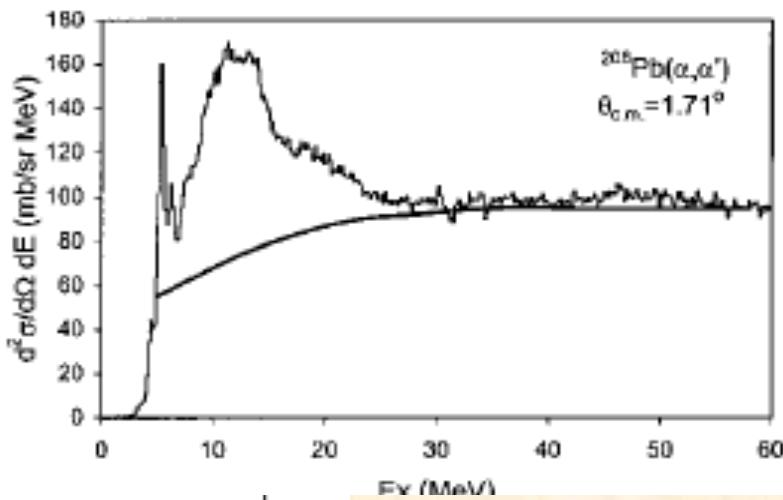
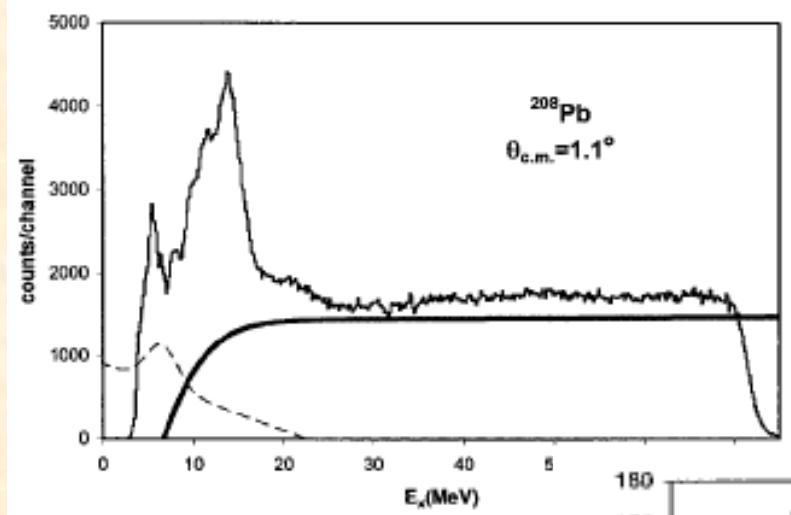
Supported in part by the National Science Foundation

GDR Resanet: GMR Workshop, Orsay, November 25, 2020



386 MeV (α, α')



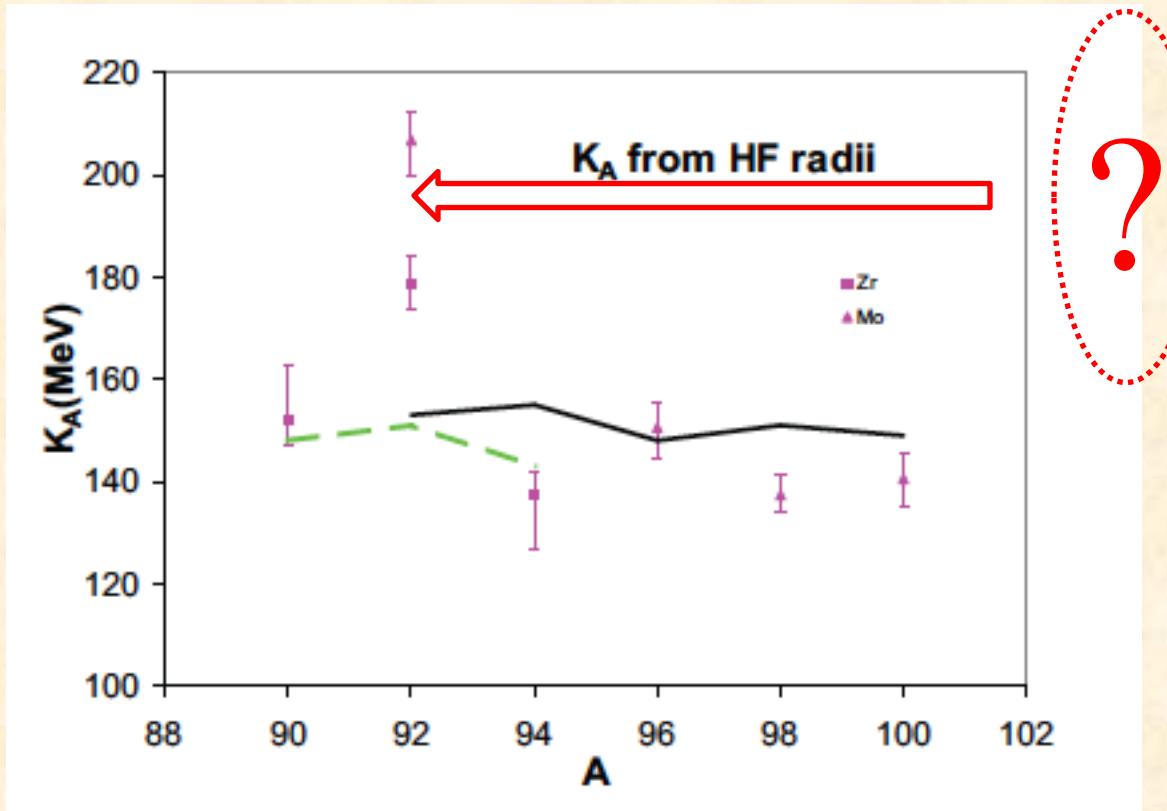


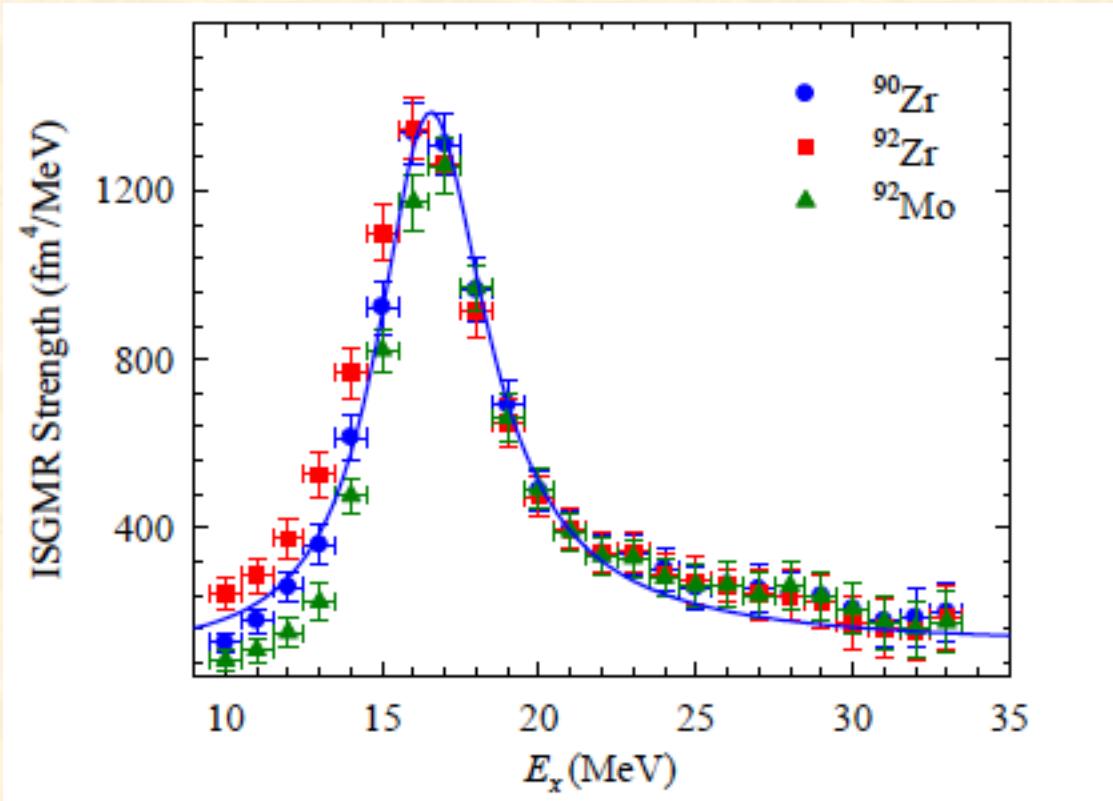


Unexpected characteristics of the isoscalar monopole resonance in the $A \approx 90$ region: Implications for nuclear incompressibility

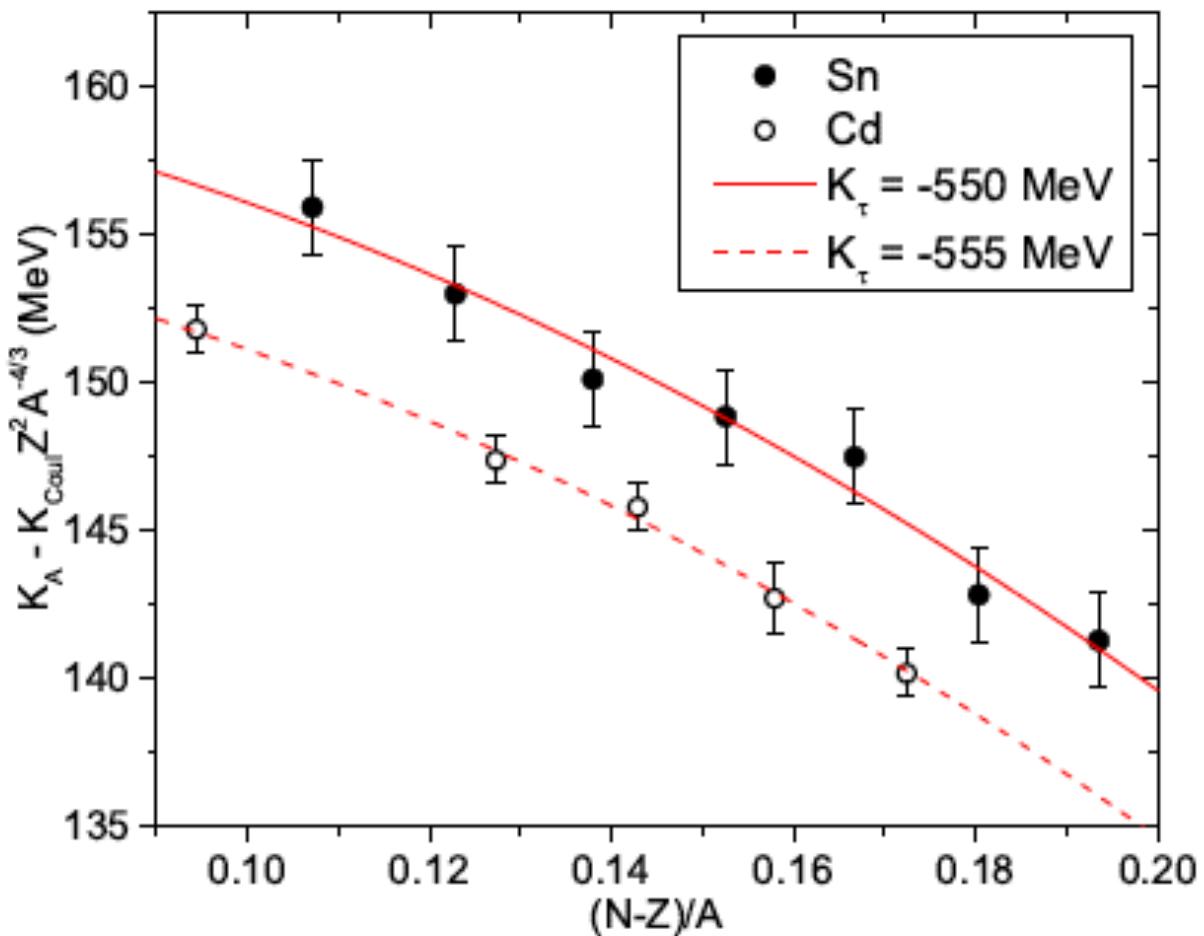
D. H. Youngblood,¹ Y.-W. Lui,¹ Krishichayan,^{1,2} J. Button,¹ M. R. Anders,¹ M. L. Gorelik,³ M. H. Urin,³ and S. Shlomo¹

¹Cyclotron Institute, Texas A&M University, College Station, Texas 77843, USA



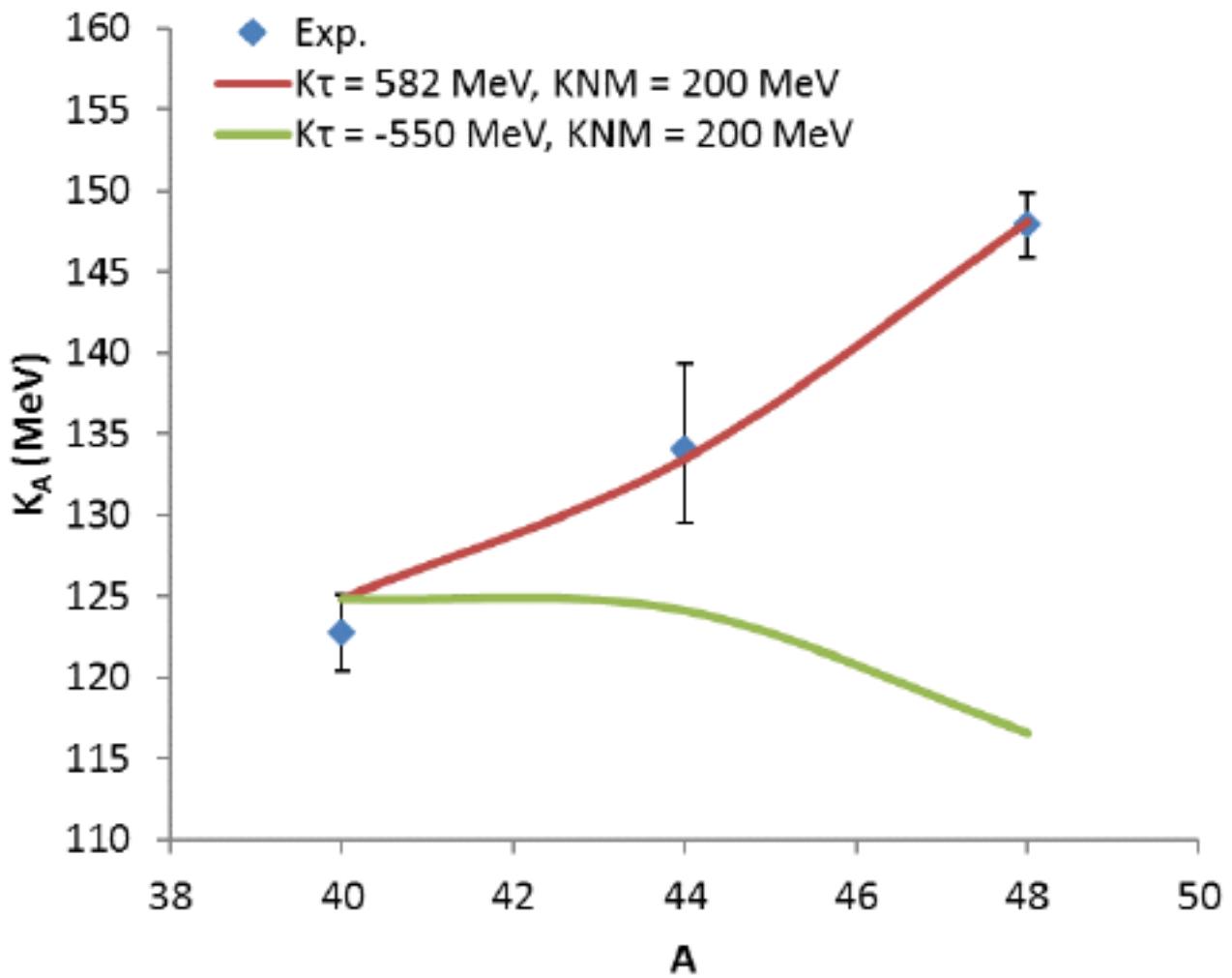


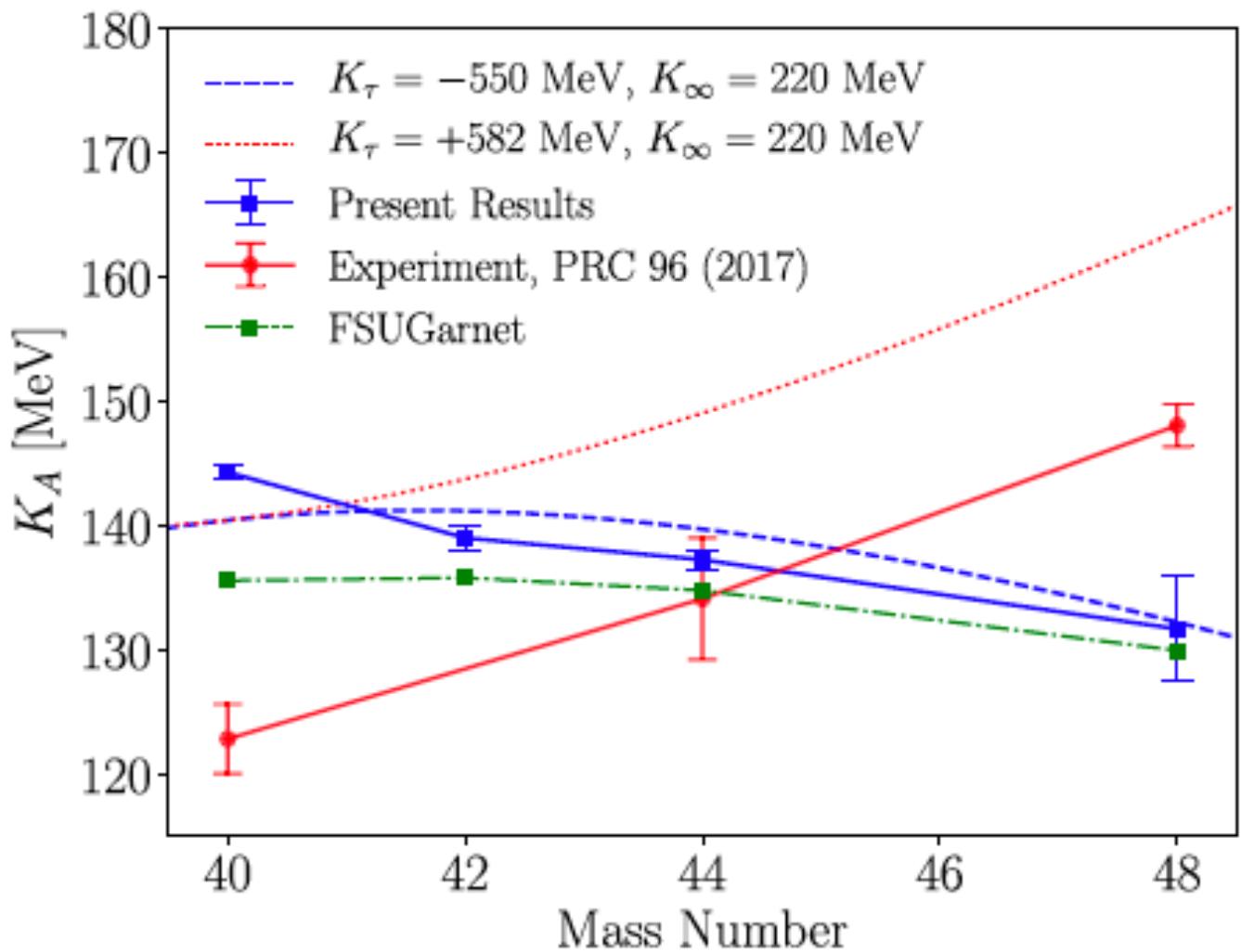
Nucleus	E_m (MeV)	Γ (MeV)
⁹⁰ Zr	16.55 ± 0.08	4.2 ± 0.3
⁹² Zr	16.12 ± 0.04	4.5 ± 0.2
⁹² Mo	16.79 ± 0.11	4.2 ± 0.4



$$K_\tau = -550 \pm 100 \text{ MeV}$$

$$K_\tau = K_{\text{sym}} - 6L - Q_0 L / K_\infty$$

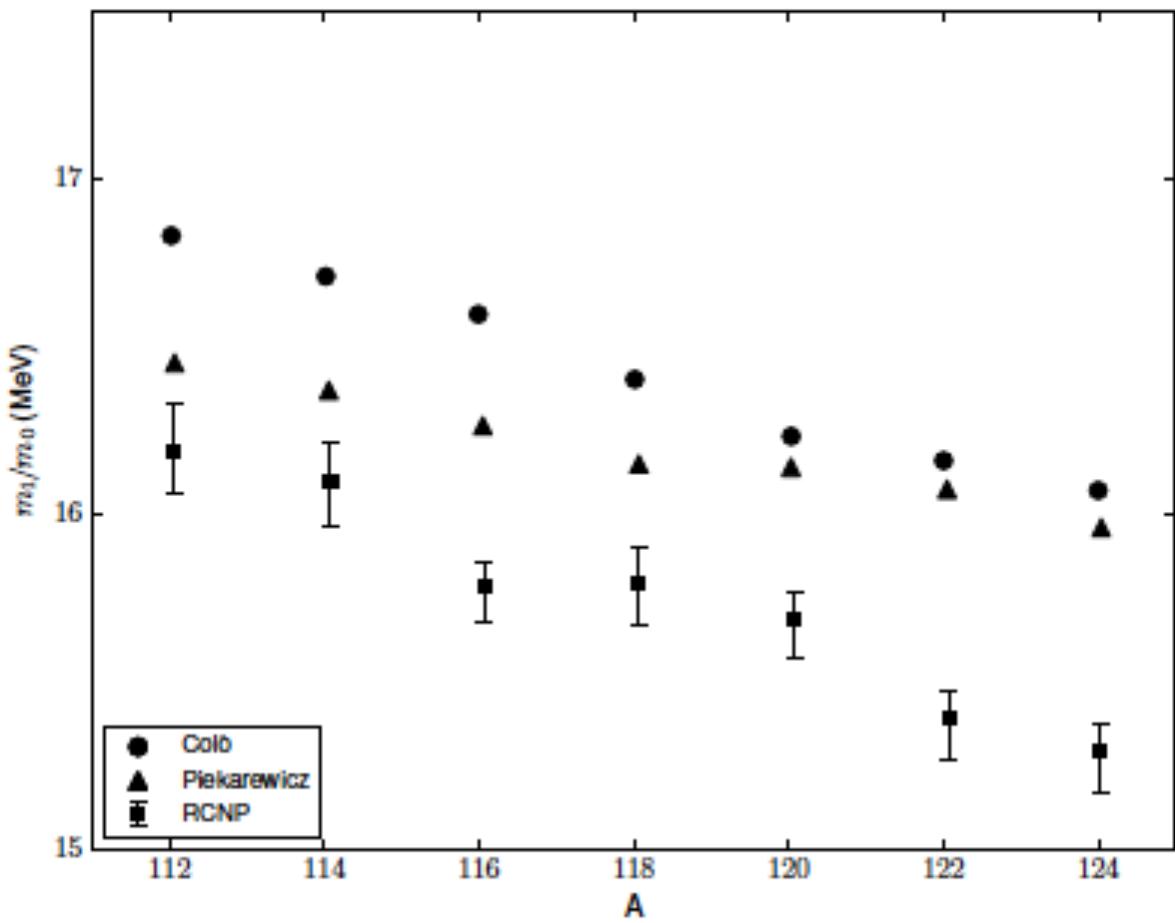






From GMR data on ^{208}Pb and ^{90}Zr
 $K_\infty = 240 \pm 20 \text{ MeV}$

This number is consistent with both GMR and ISGDR data and with non-relativistic and relativistic calculations



Why are tins so “Fluffy”?

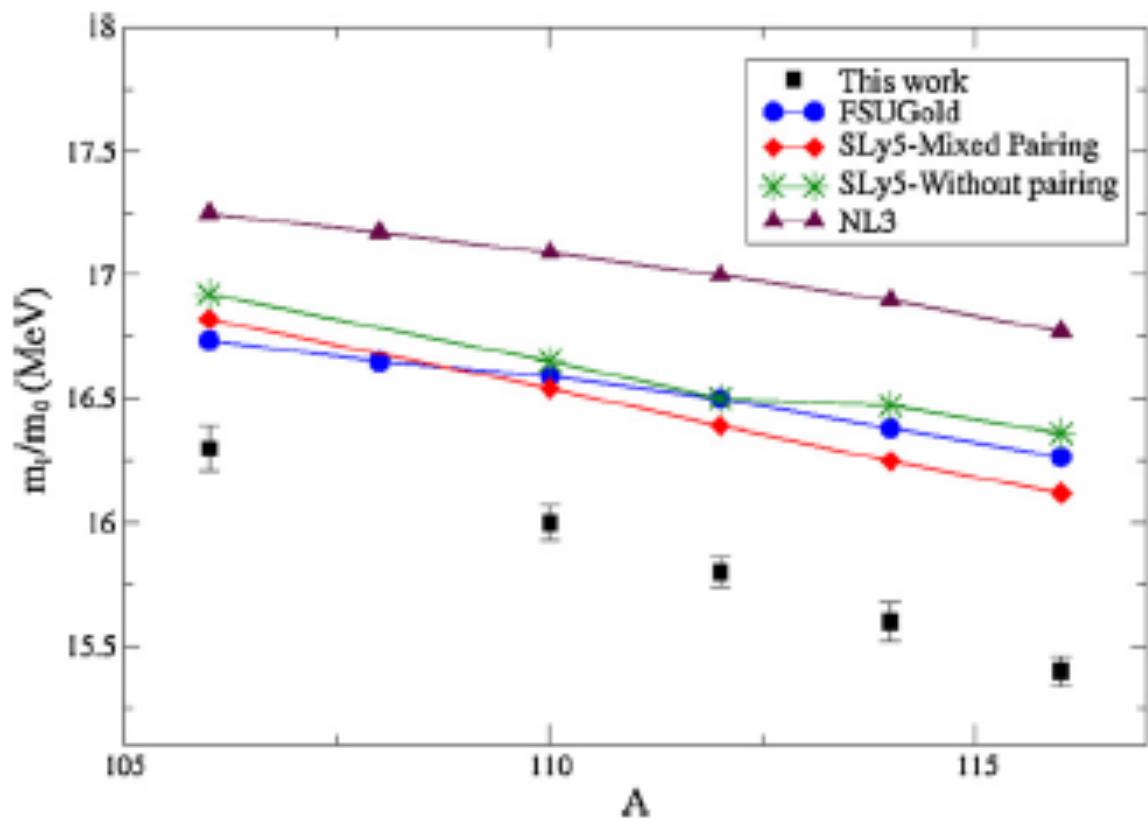
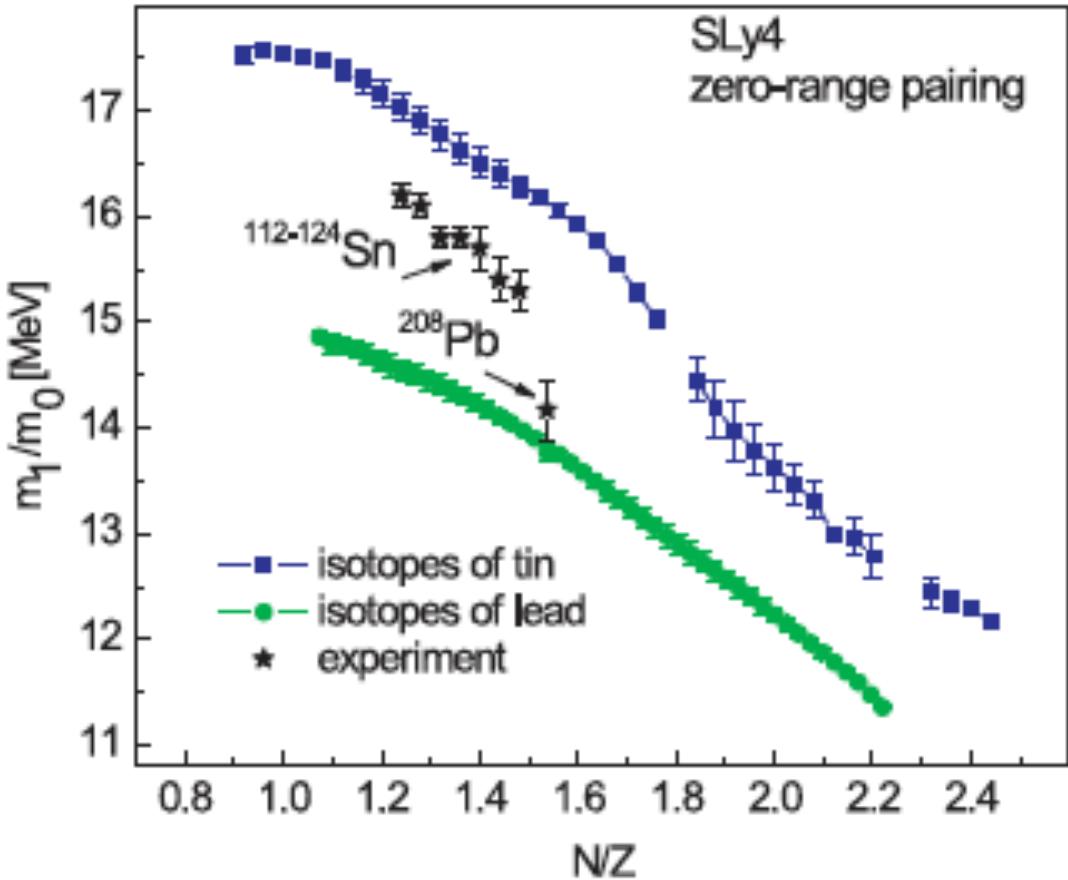
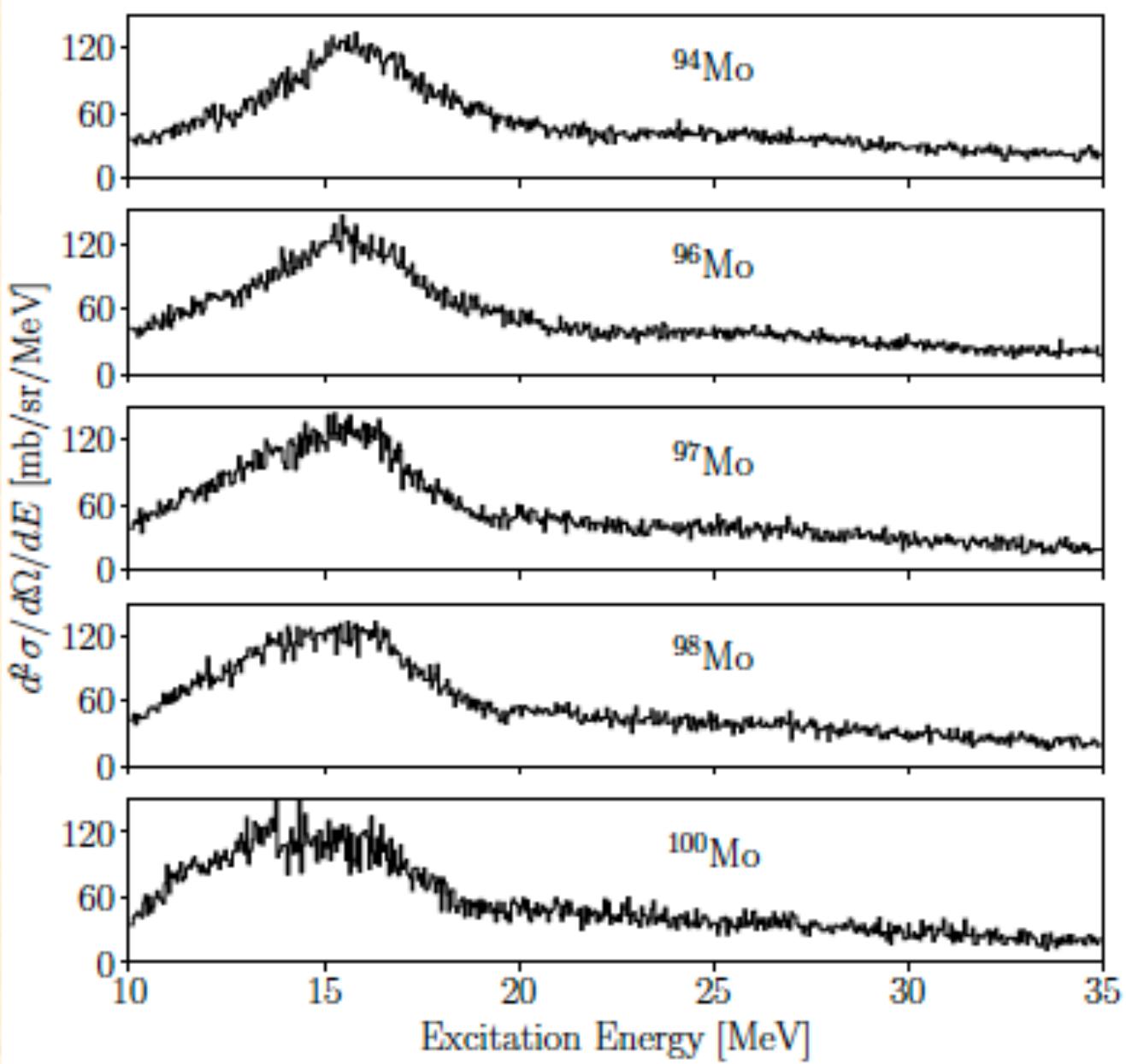
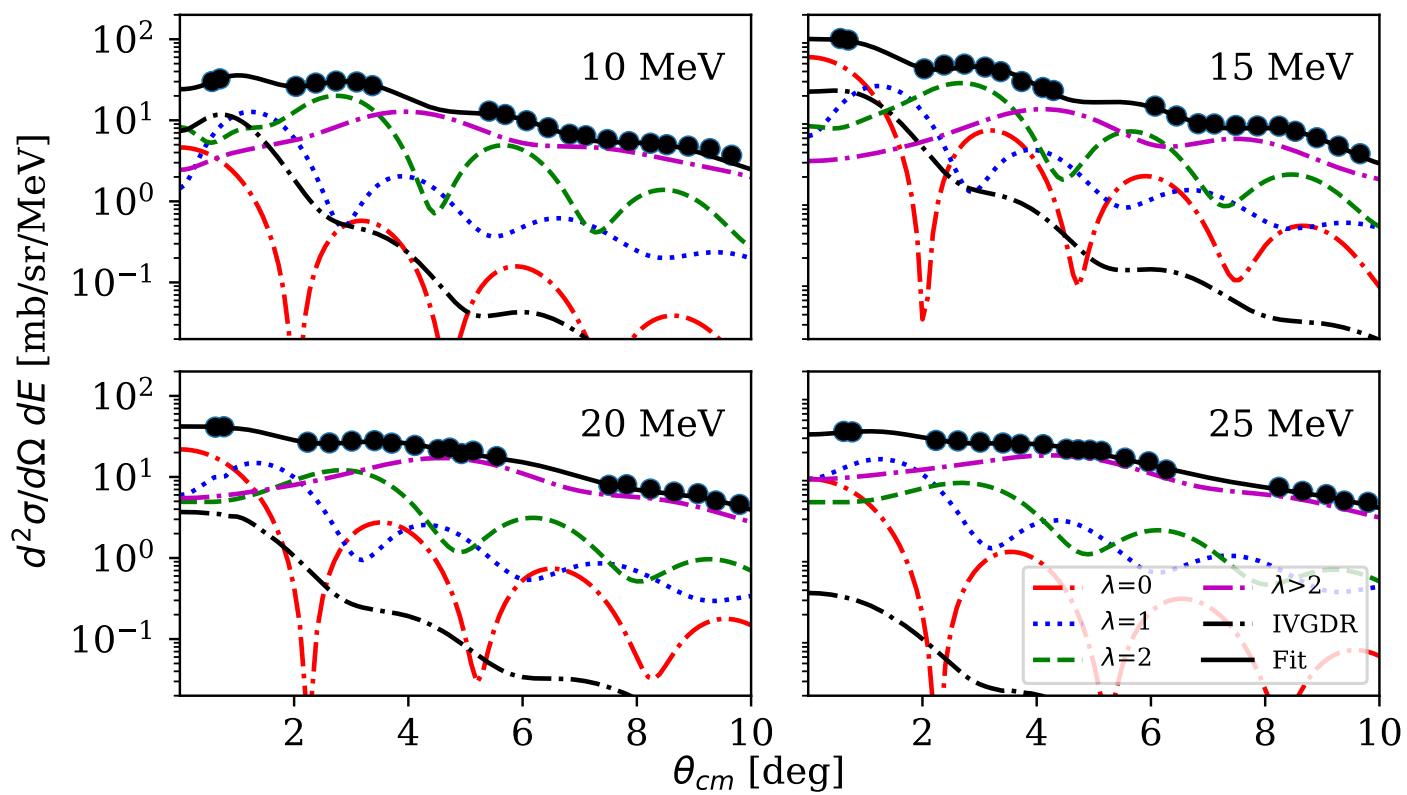


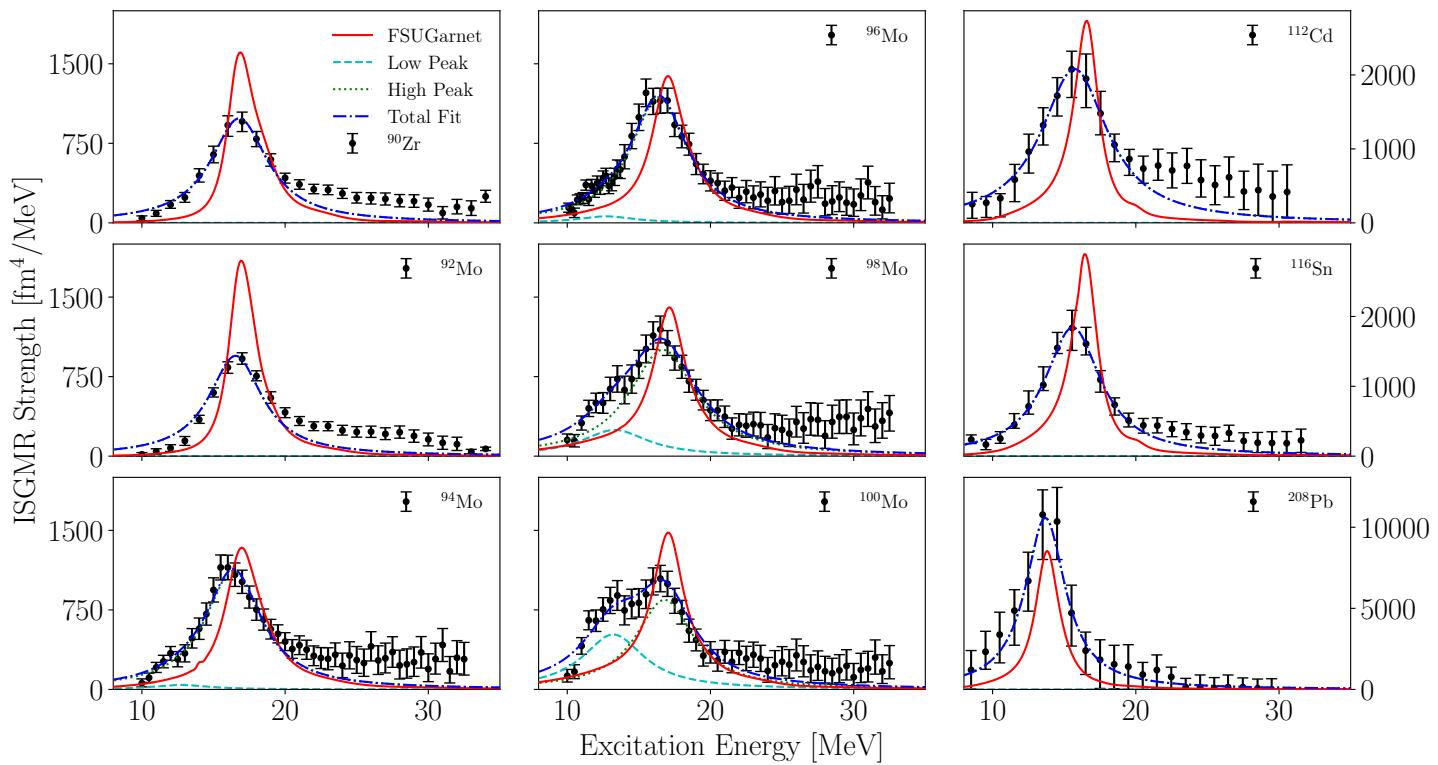
Fig. 3. (Color online.) Systematics of the moment ratio, m_1/m_0 for the ISGMR strength distributions in the Cd isotopes investigated in this work. The experimental results (squares) are compared with relativistic calculations performed using the FSUGold (circles) and NL3 (triangles) effective interactions. Also presented are results from non-relativistic calculations performed using the Sly5 parameter set in the HF-BCS + QRPA formalism with and without the mixed pairing interaction (diamonds and stars, respectively) [36]. The solid lines are to guide the eye.

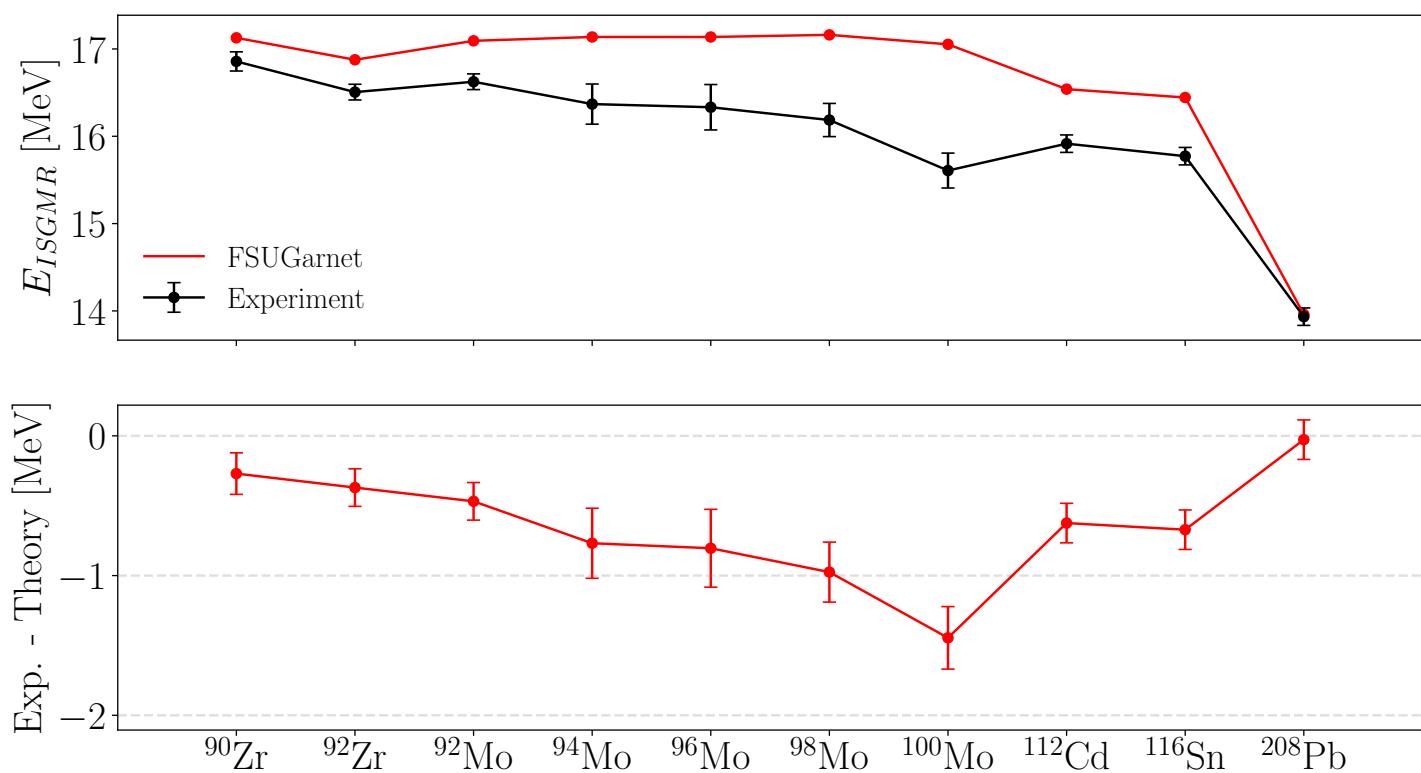


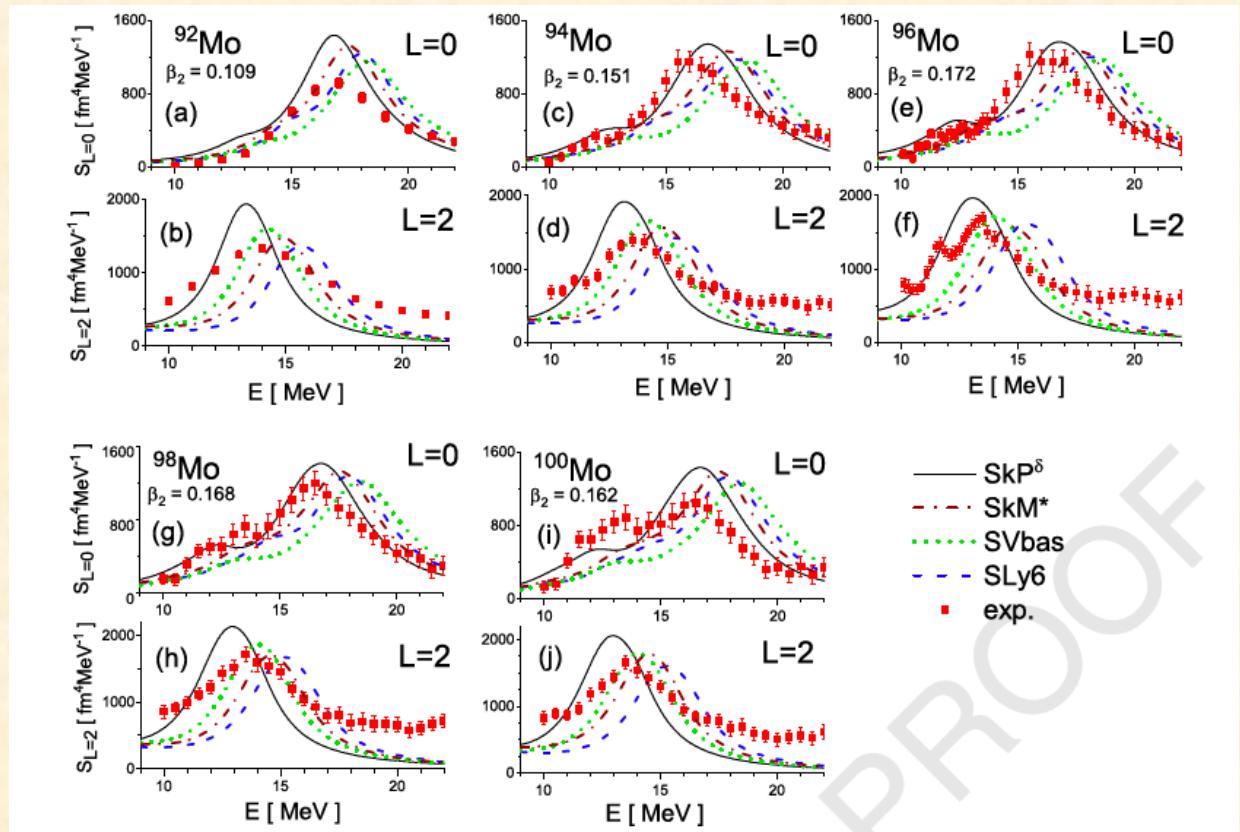




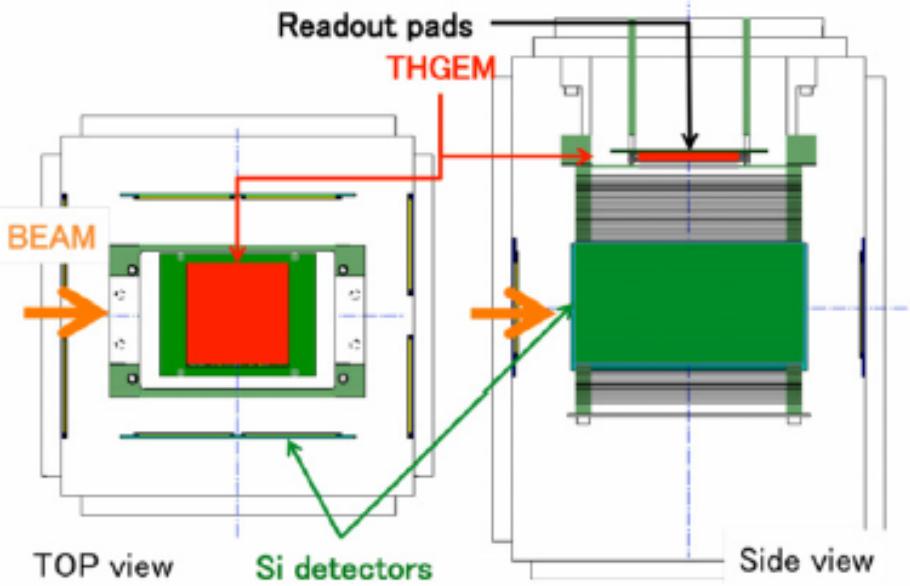
^{94}Mo







the connection between the line shape of the monopole strength ISGMR and the deformation-induced coupling between the ISGMR and the $K = 0$ branch of the ISGQR. The ISGMR is best described by the force SkP^δ , having a low incompressibility $K_\infty = 202 \text{ MeV}$.



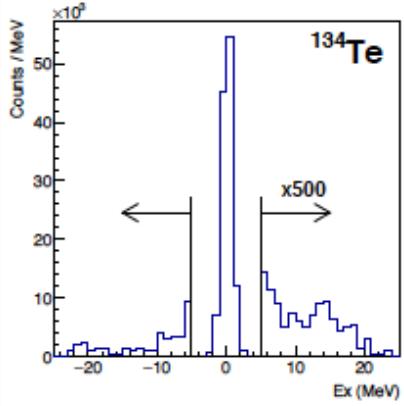
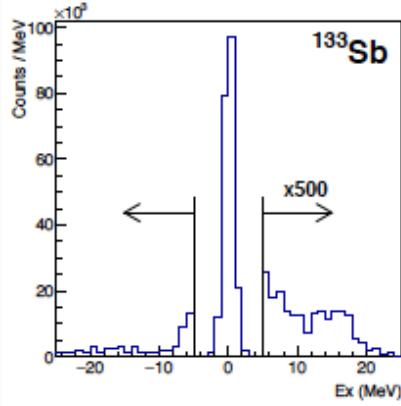
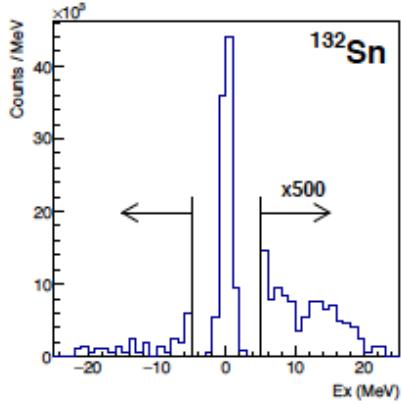
RIBF 113

$^{132}\text{Sn} + ^2\text{H}$
100 MeV/A
>50 kHz ^{132}Sn



Preliminary uncorrected energy spectra for ^{132}Sn , ^{133}Sb , and ^{134}Te .

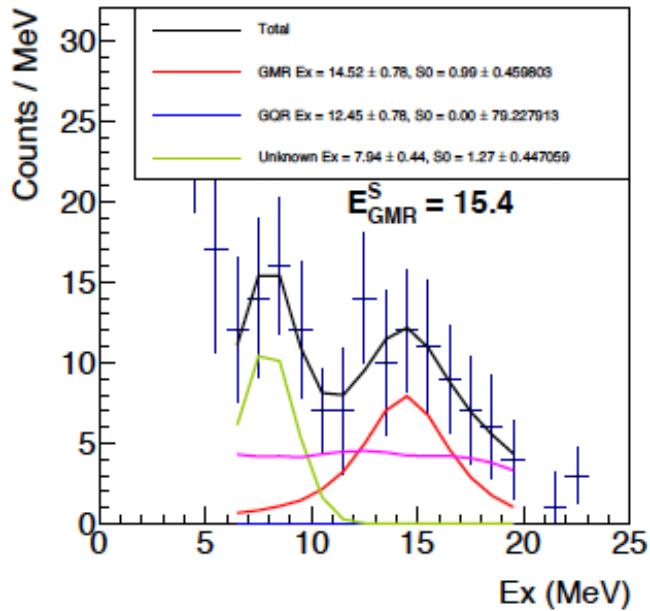
Particle identifications of beam and recoil particles have been done.



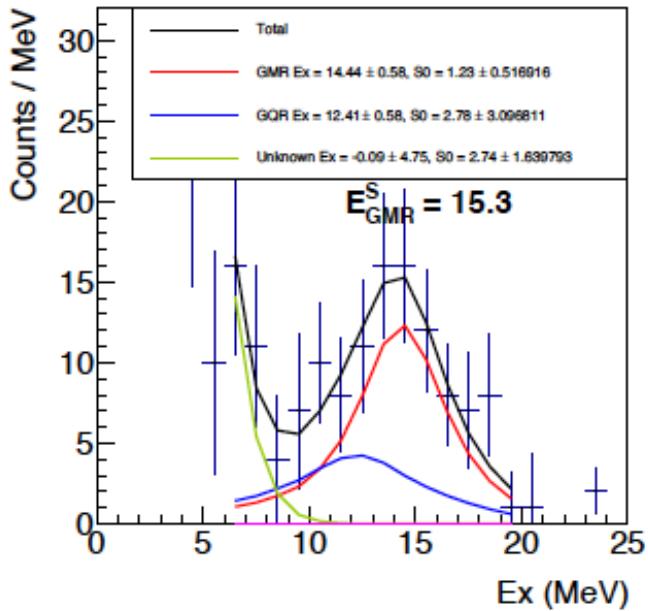
Preliminary results from RIKEN $d(^{132}\text{Sn}, ^{132}\text{Sn}')d$ experiment



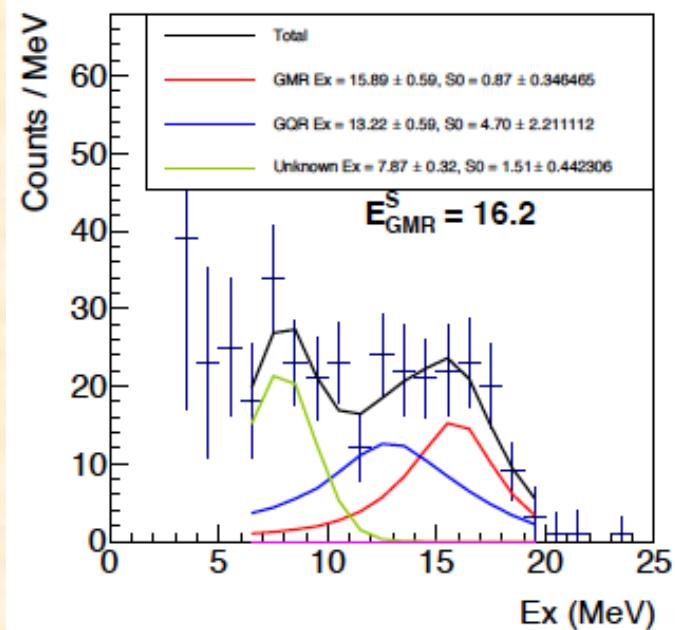
132Sn



134Te

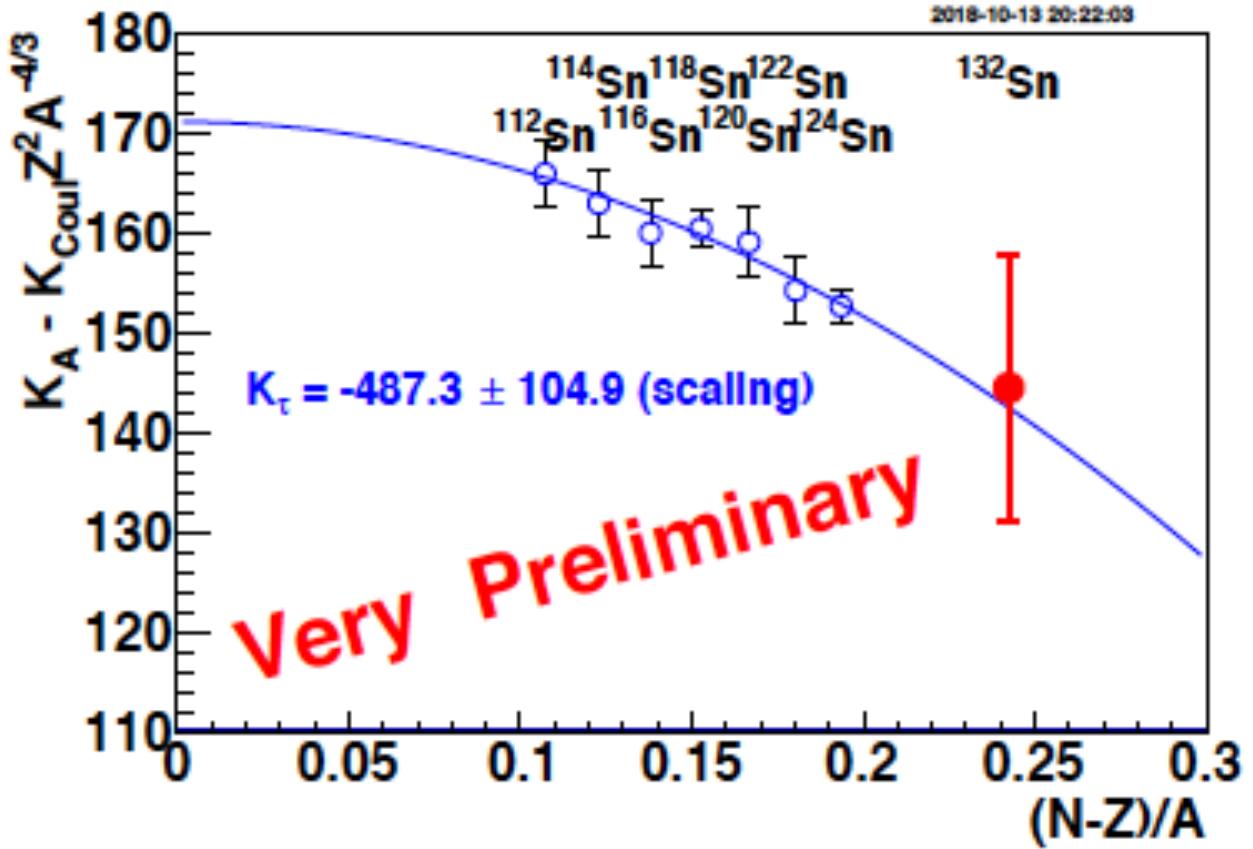


133Sb





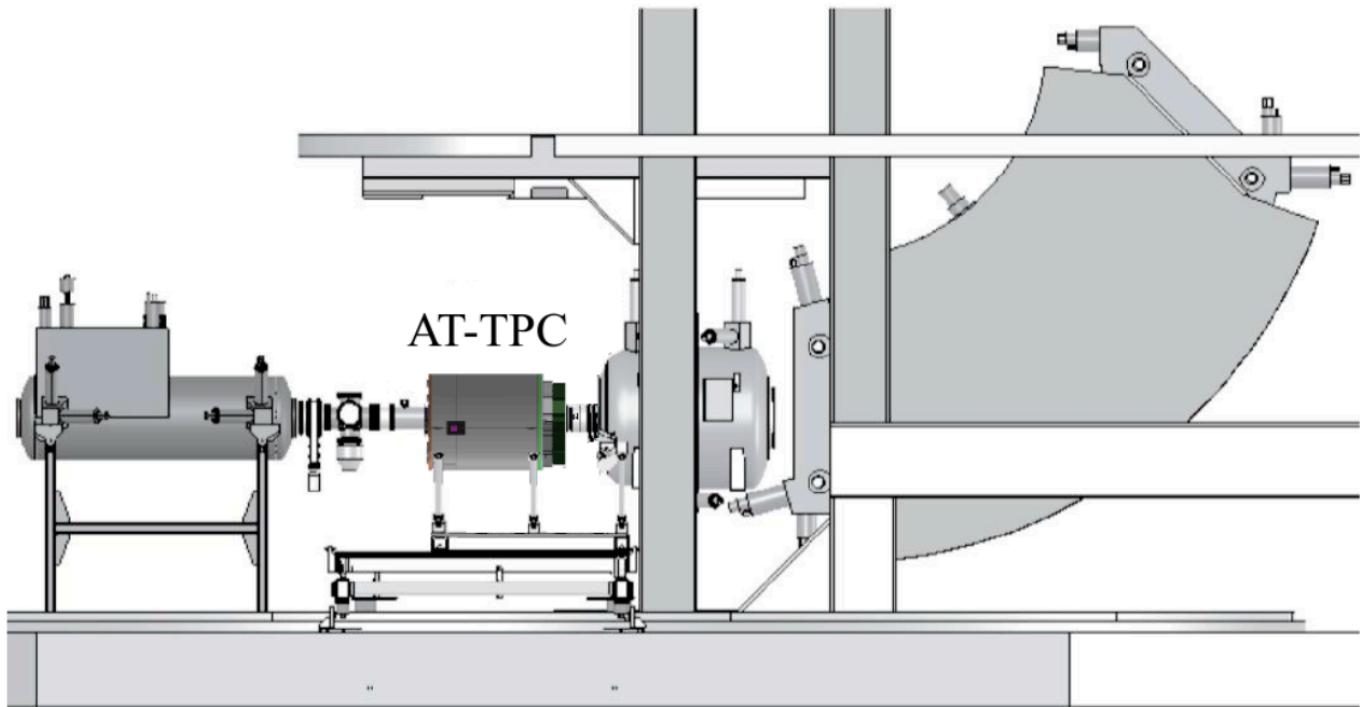
2018-10-13 20:22:03



Preliminary results from RIKEN $d(^{132}\text{Sn}, ^{132}\text{Sn}')d$ experiment



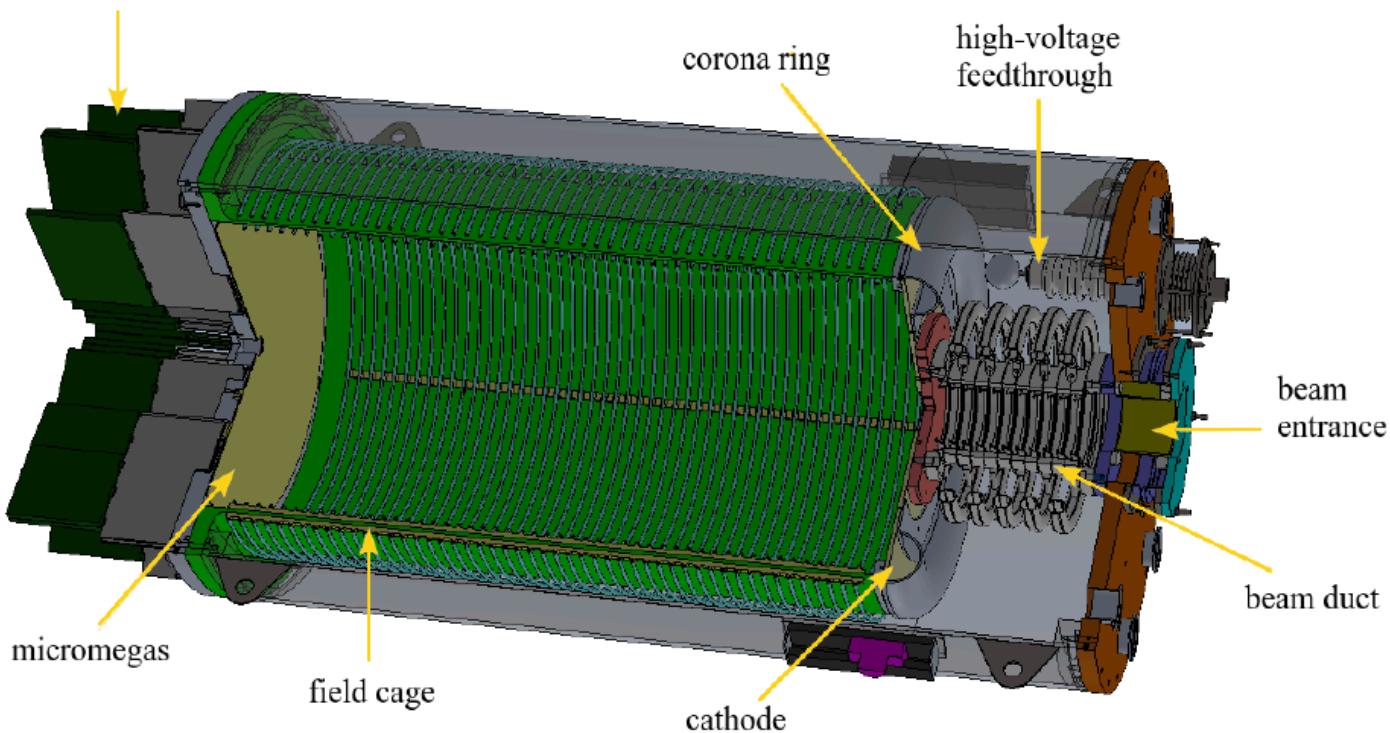
$^4\text{He}(^{70}\text{Ni}, ^{70}\text{Ni}')^4\text{He}$ @ NSCL



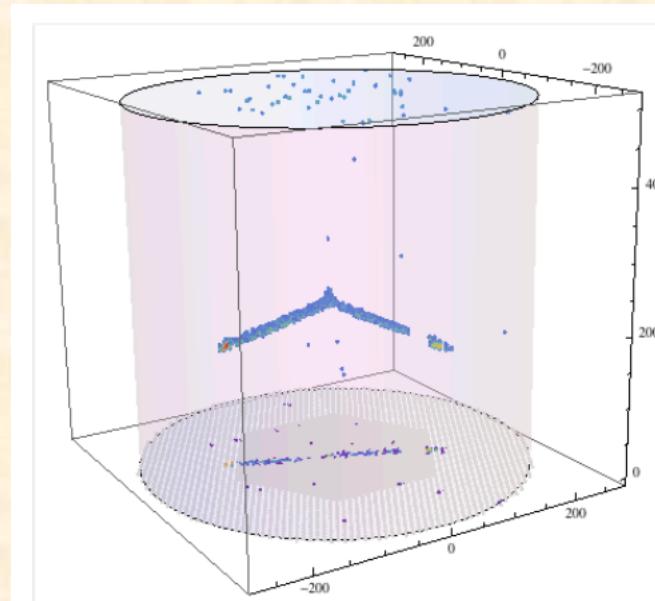
82 MeV/A; ~20k pps

e18207

AsAd boards

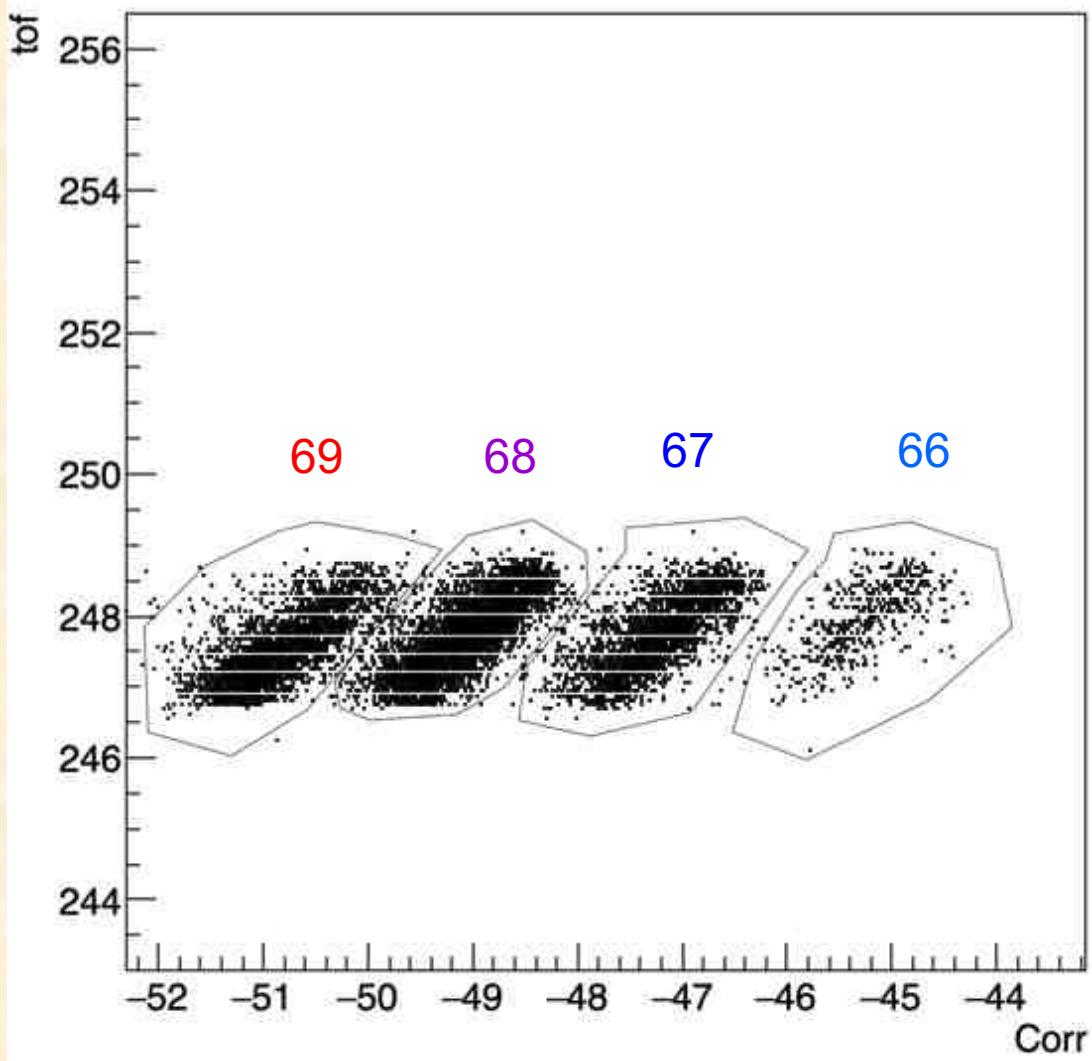


150 torr of
pure ${}^4\text{He}$ gas





tof:Corr





- ◆ We have investigated the ISGMR in the Mo isotopes, via inelastic scattering of 386-MeV α particles at extremely forward angles (including 0°).
- ◆ In the Cd and Sn isotopes, the ISGMR energy was significantly lower than that expected from the accepted value of K_∞ . There has been no satisfactory theoretical explanation of this “fluffiness” of open-shell nuclei.
- ◆ The “fluffiness” appears in the Mo isotopes, beginning with ^{92}Mo , just two nucleons out of the “doubly-closed” nucleus ^{90}Zr .
- ◆ I hope one of the theorists among you will pick up the gauntlet of trying to find a satisfactory explanation of this phenomenon.
- ◆ We have just completed measurements at NSCL on ^{70}Ni using pure 4He gas and trigger with S800. Very clean spectra have been observed with a “peak” at ~17 MeV consistent with GMR.

Need much better statistics, however!



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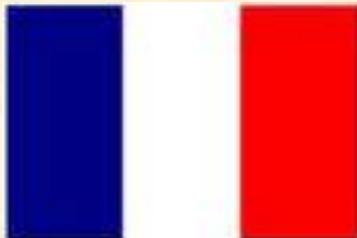




Merci beaucoup

धन्यवाद

Thanks !





The Question Kitten



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