

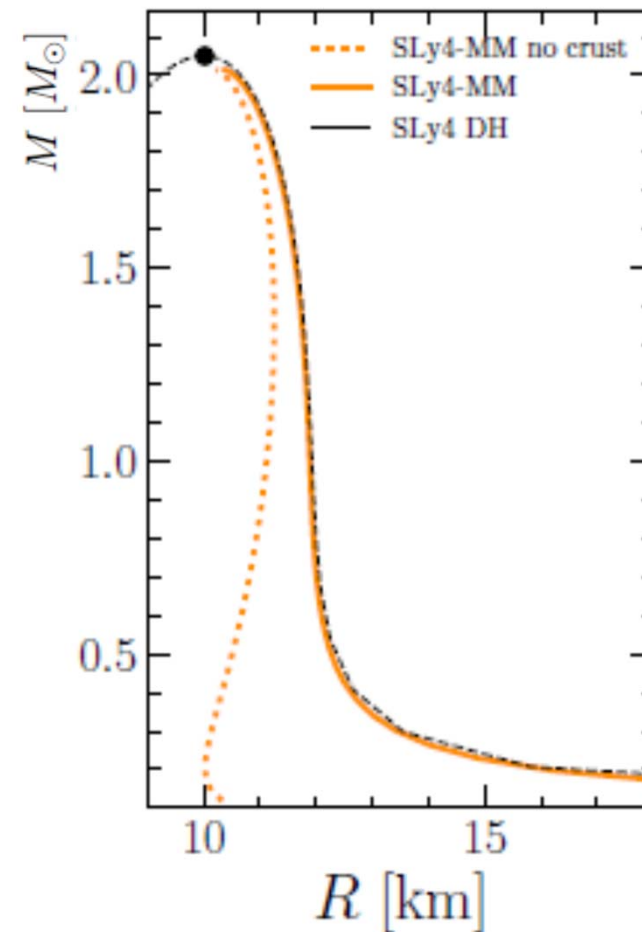
# Nuclear incompressibility and the NS crust

GMR virtual workshop of the GDR RESANET 25-26/11/2020

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# Plan

1. *Incompressibility and other empirical parameters*
2. *Modeling the NS crust*
3. *Influence of incompressibility on crustal observables*



# The EoS empirical parameters

- $e = e_0 = e(\rho, \alpha = 0)$   
 $\rightarrow X_n = \{E_{sat}, K_{sat}, Q_{sat}, Z_{sat}, \dots\}$
- $e = e_{sym} = \frac{1}{2} \frac{\partial^2}{\partial \alpha^2} e(\rho, \alpha = 0)$   
 $\rightarrow X_n = \{E_{sym}, L_{sym}, K_{sym}, Q_{sym}, Z_{sym}, \dots\}$

$$X_n = (3\rho_0)^n \frac{\partial^n e}{\partial \rho^n} (\rho = \rho_0, \alpha = 0)$$

$K_{sat}, K_{sym}$  give the response to a compression mode from  $\rho = \rho_0$

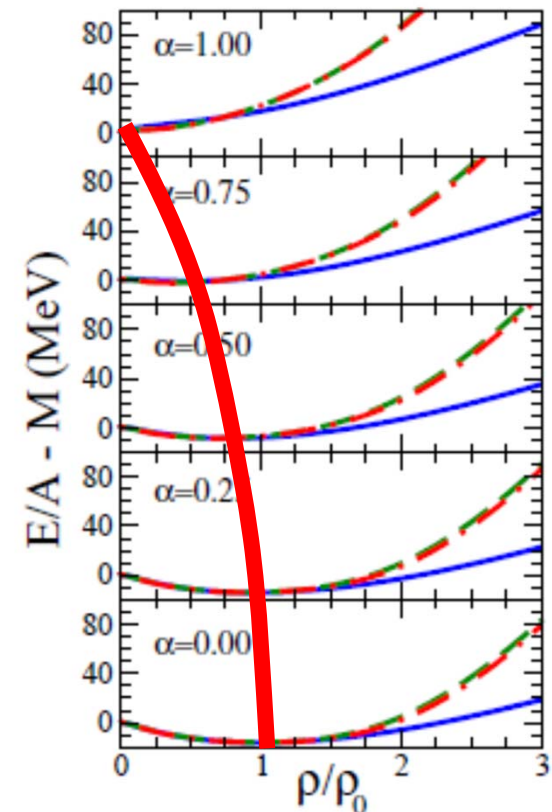
But if  $\rho \neq \rho_0$ , many parameters contribute!

**Example:** saturation density

$$\rho_0(\alpha) = \rho_0 - 3\rho_0 \frac{L_{sym}}{K_{sat}} \alpha^2 + O(\alpha^4)$$

$\Rightarrow$  Incompressibility of asymmetric matter

$$\begin{aligned} K_0(\alpha) &= K_{sat} + K_\tau \alpha^2 + O(\alpha^4) \\ &= K_{sat} + \left( K_{sym} - 6L_{sym} - L_{sym} \frac{Q_{sat}}{K_{sat}} \right) \alpha^2 + O(\alpha^4) \end{aligned}$$



- J. Piekarewicz and M. Centelles, *Phys.Rev.C* 79, 054311 (2009). •

# Constraints from nuclear experiments...

TABLE I. Group A: saturation energy  $E_{\text{sat}}$ , density  $n_{\text{sat}}$ , incompressibility  $K_{\text{sat}}$ , and symmetry energy  $E_{\text{sym}}$  estimated from various analyses of experimental data. See text for more details.

Model	Ref.	$E_{\text{sat}}$ (MeV)	$n_{\text{sat}}$ ( $\text{fm}^{-3}$ )	$K_{\text{sat}}$ (MeV)	$E_{\text{sym}}$ (MeV)
El. scatt.	Wang-99 [55]		0.1607	235 $\pm 15$	
LDM	Myers-66 [56]	-15.677	0.136*	295	28.06
LDM	Royer-08 [57]	-15.5704	0.133*		23.45
LSD	Pomorski-03 [58]	-15.492	0.142*		28.82
DM	Myers-77 [59]	-15.96	0.145*	240	36.8
FRDM	Buchinger-01 [60]		0.157 $\pm 0.004$		
INM	Satpathy-99 [61]	-16.108	0.1620	288 $\pm 20$	
DF-Skyrme	Tondeur-86 [62]		0.158		
DF-Skyrme	Klupfel-09 [63]	-15.91 $\pm 0.06$	0.1610 $\pm 0.0013$	222 $\pm 8$	30.7 $\pm 1.4$
DF-BSK2	Goriely-02 [64]	-15.79	0.1575	234	28.0
DF-BSK24, 28,29	Goriely-15 [65]	-16.045 $\pm 0.005$	0.1575 $\pm 0.0004$	245	30.0
DF-Skyrme	McDonnell-15 [66]	-15.75 $\pm 0.25$	0.160 $\pm 0.005$	220 $\pm 20$	29 $\pm 1$
DF-NLRMF	NL3* [67]	-16.3	0.15	258	38.7
DF-NLRMF	PK [68]	-16.27	0.148	283	37.7
DF-DDRMF	DDME1,2 [69,70]	-16.17 $\pm 0.03$	0.152 $\pm 0.00$	247 $\pm 3$	32.7 $\pm 0.4$
DF-DDRMF	PK [68]	16.27	0.150	262	36.8
Present Estimation		-15.8 $\pm 0.3$	0.155 $\pm 0.005$	230 $\pm 20$	32 $\pm 2$

\*Value determined from  $r_0$ .

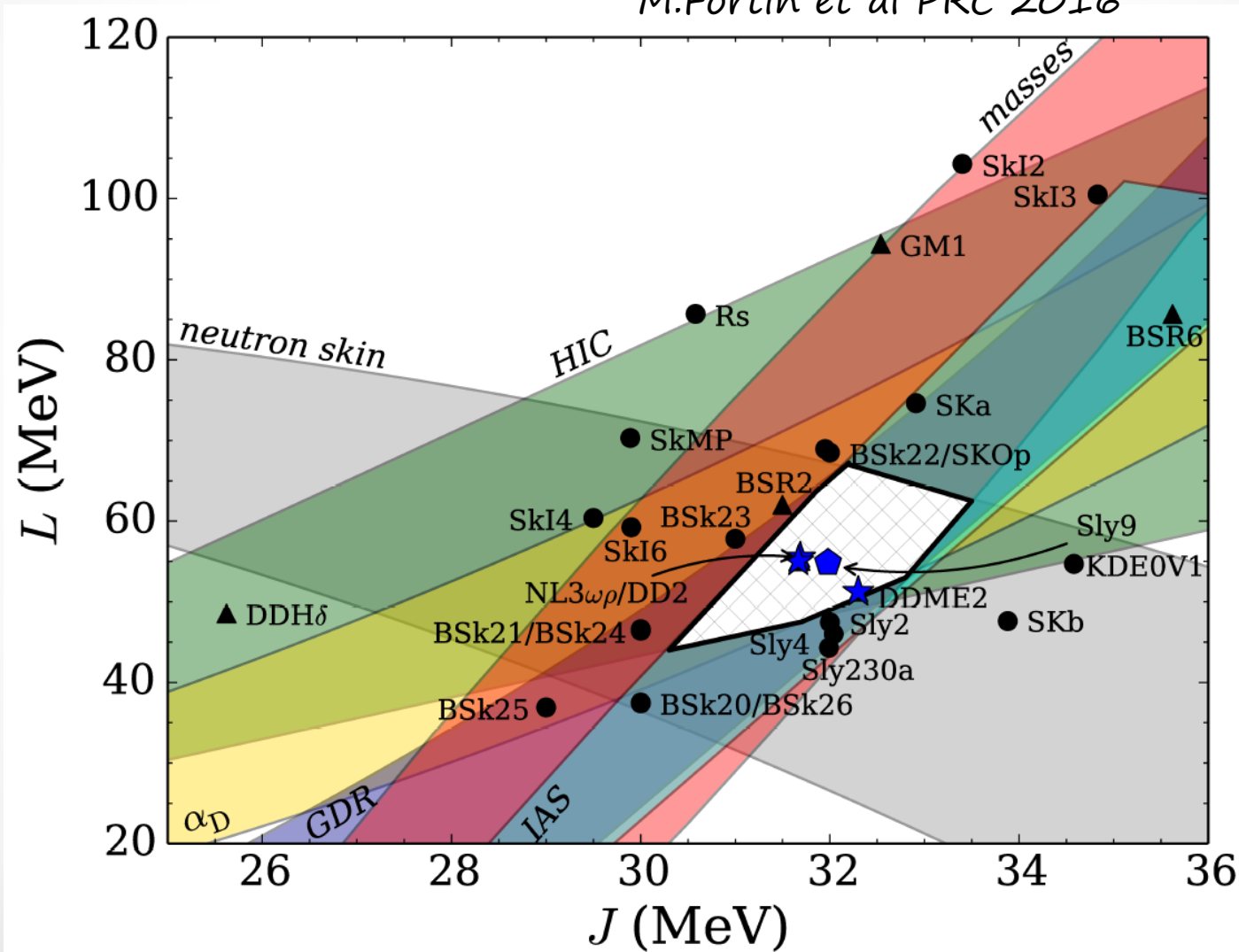
TABLE II. Group B parameters: isoscalar skewness  $Q_{\text{sat}}$ , slope of the symmetry energy  $L_{\text{sym}}$ , and isovector incompressibility  $K_{\text{sym}}$ . The parameter  $K_{\tau}$  is defined as  $K_{\tau} = K_{\text{sym}} - 6L_{\text{sym}} - Q_{\text{sat}}L_{\text{sym}}/K_{\text{sat}}$ . See text for more details.

Model	Ref.	$Q_{\text{sat}}$ (MeV)	$L_{\text{sym}}$ (MeV)	$K_{\text{sym}}$ (MeV)	$K_{\tau}$ (MeV)
DF-Skyrme	Berdichevsky-88 [71]	30	0		
DF-Skyrme	Farine-97 [72]	-700 $\pm 500$			
DF-Skyrme	Alam-14 [31]	-344 $\pm 46$	65 $\pm 14$	-23 $\pm 73$	-322 $\pm 34$
DF-Skyrme	McDonnell-15 [66]		40 $\pm 20$		
DF-NLRMF	NL3* [67]	124	123	106	-690
DF-NLRMF	PK [68]	-25	116	55	-630
DF-DDRMF	DDME1,2 [69,70]	400 $\pm 80$	53 $\pm 3$	-94 $\pm 7$	-500 $\pm 7$
DF-DDRMF	PK [68]	-119	79.5	-50	-491
Correlation	Centelles-09 [73]		70 $\pm 40$		-425 $\pm 175$
DF-RPA	Carbone-10 [74]		60 $\pm 30$		
Correlation	Danielewicz-14 [75]		53 $\pm 20$		
Correlation	Newton-14 [76]		70 $\pm 40$		
Correlation	Lattimer-14 [77]		53 $\pm 20$		
GMR	Sagawa-07 [78]				-500 $\pm 50$
GMR	Patel-14 [79]				-550 $\pm 100$
Present Estimation		300 $\pm 400$	60 $\pm 15$	-100 $\pm 100$	-400 $\pm 100$

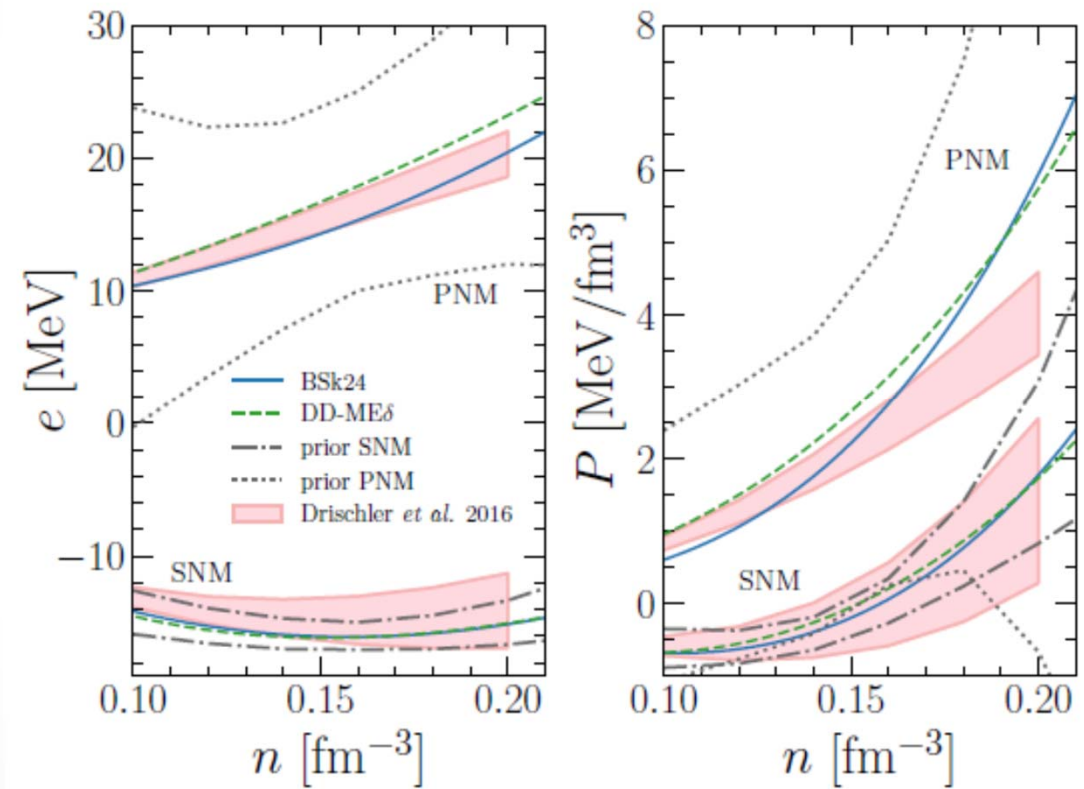
J. Margueron et al PRC 2018

# Constraints from nuclear experiments...

M.Fortin et al PRC 2016

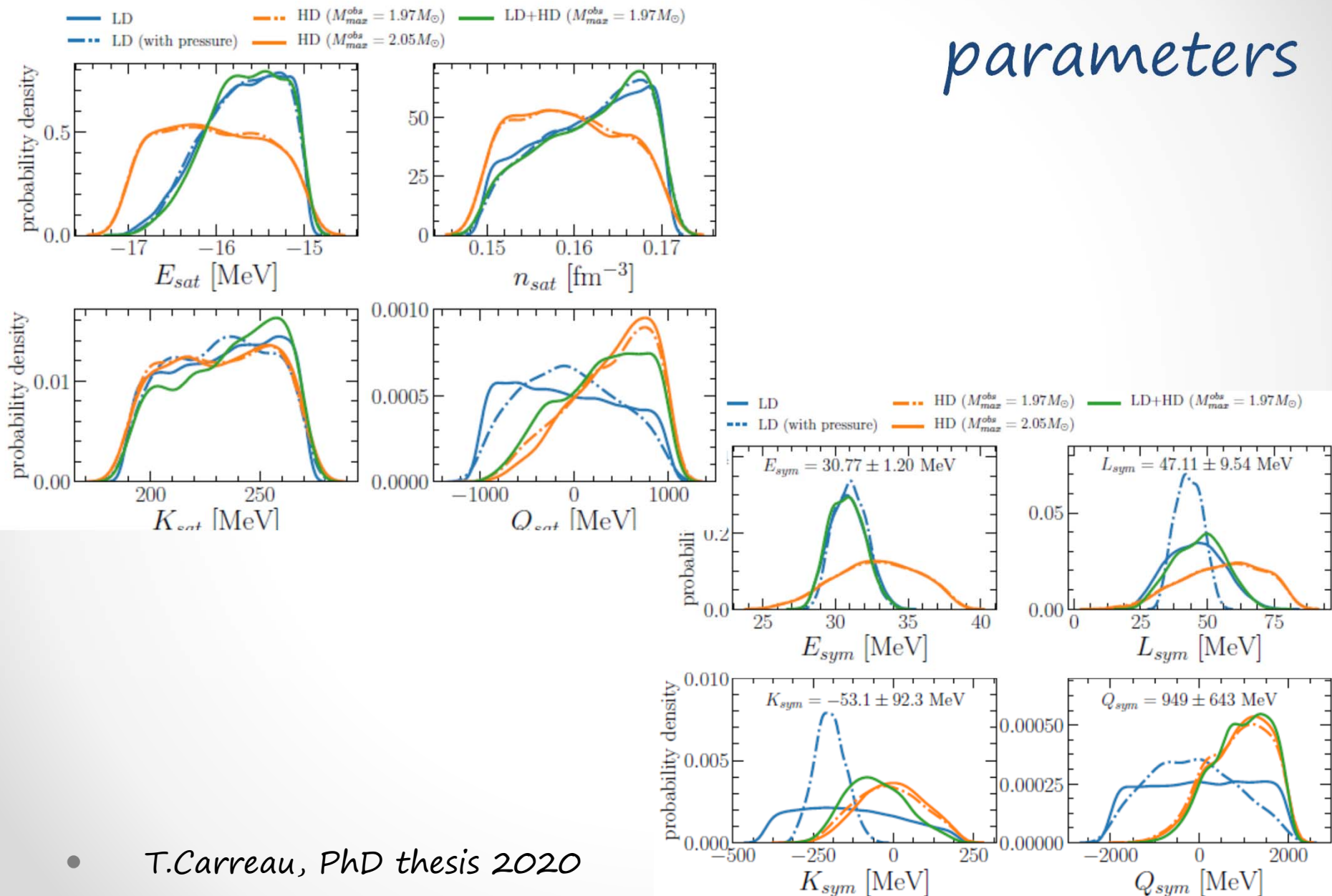


.... and nuclear theory



# Present knowledge of the empirical parameters

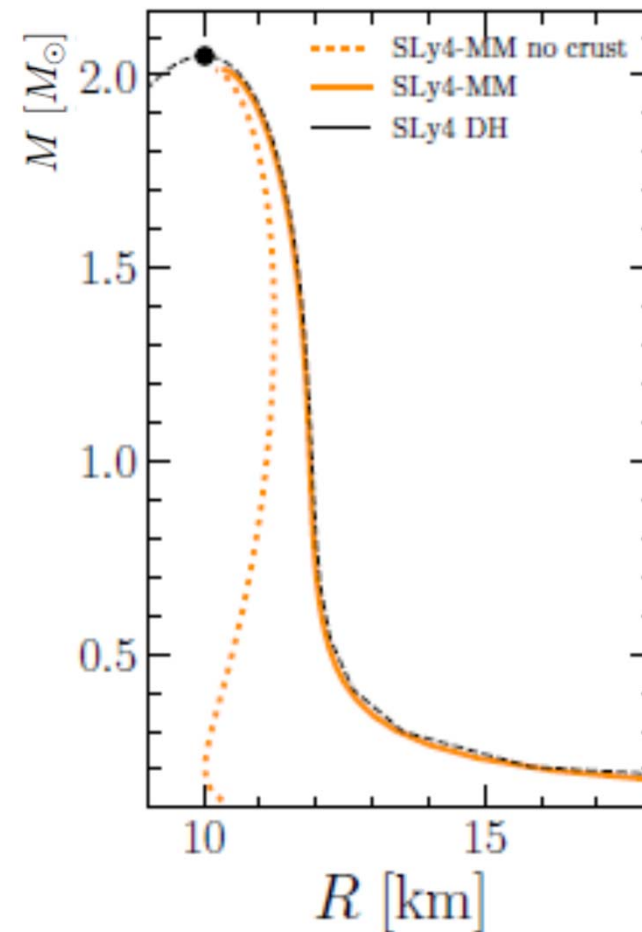
parameters



- T.Carreau, PhD thesis 2020

# Plan

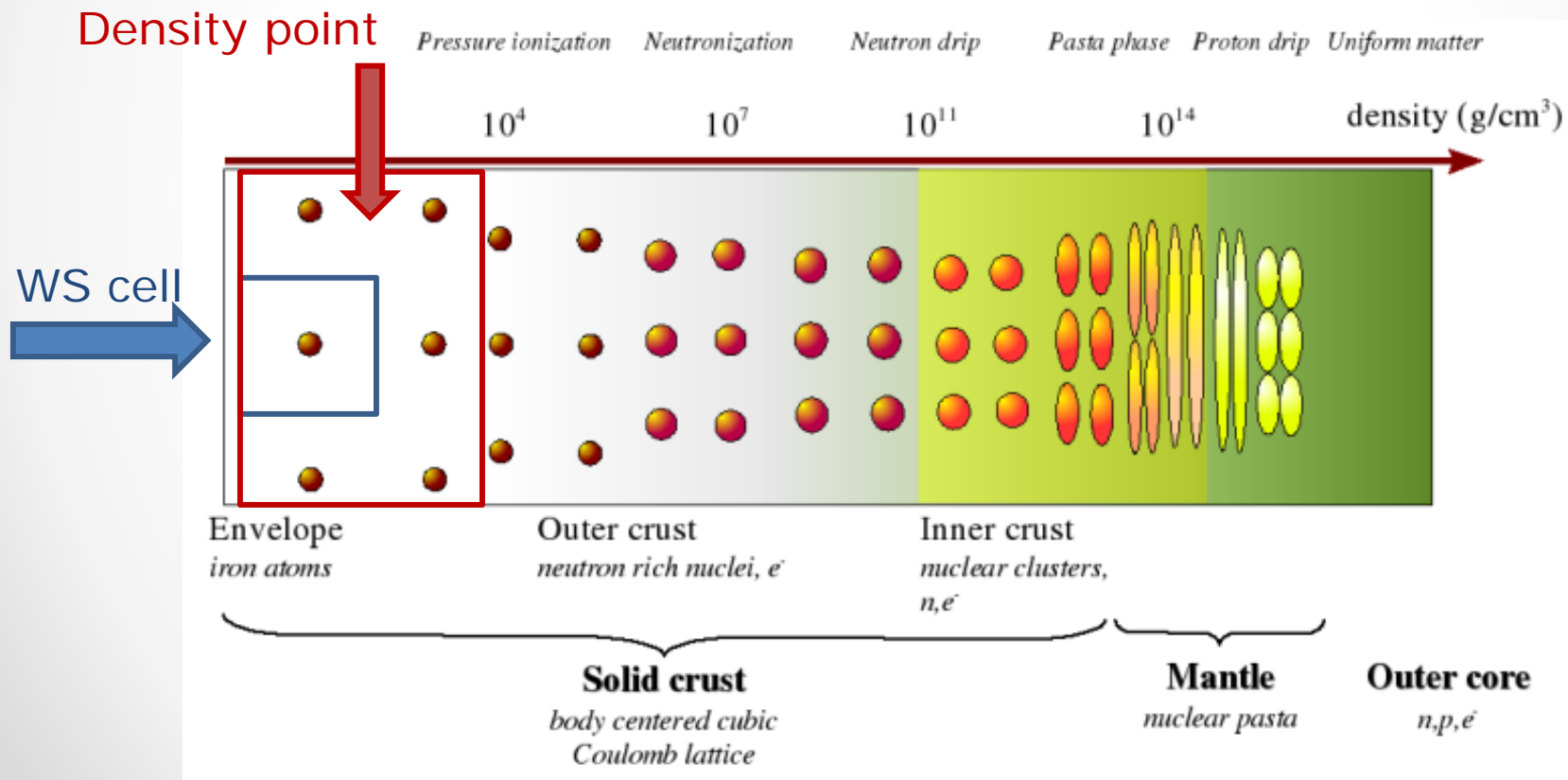
1. *Incompressibility and other empirical parameters*
2. ***Modeling the NS crust***
3. *Influence of incompressibility on the crustal observables*





# The catalyzed NS crust

- Below saturation matter is clusterized
- At T=0: solid state: BCC lattice
- Ground state energy density:  $\varepsilon_{WS}(n_B) = \frac{\sum_i E_{WS}}{\sum_i V_{WS}} = \frac{E_{WS}}{V_{WS}} = \min$



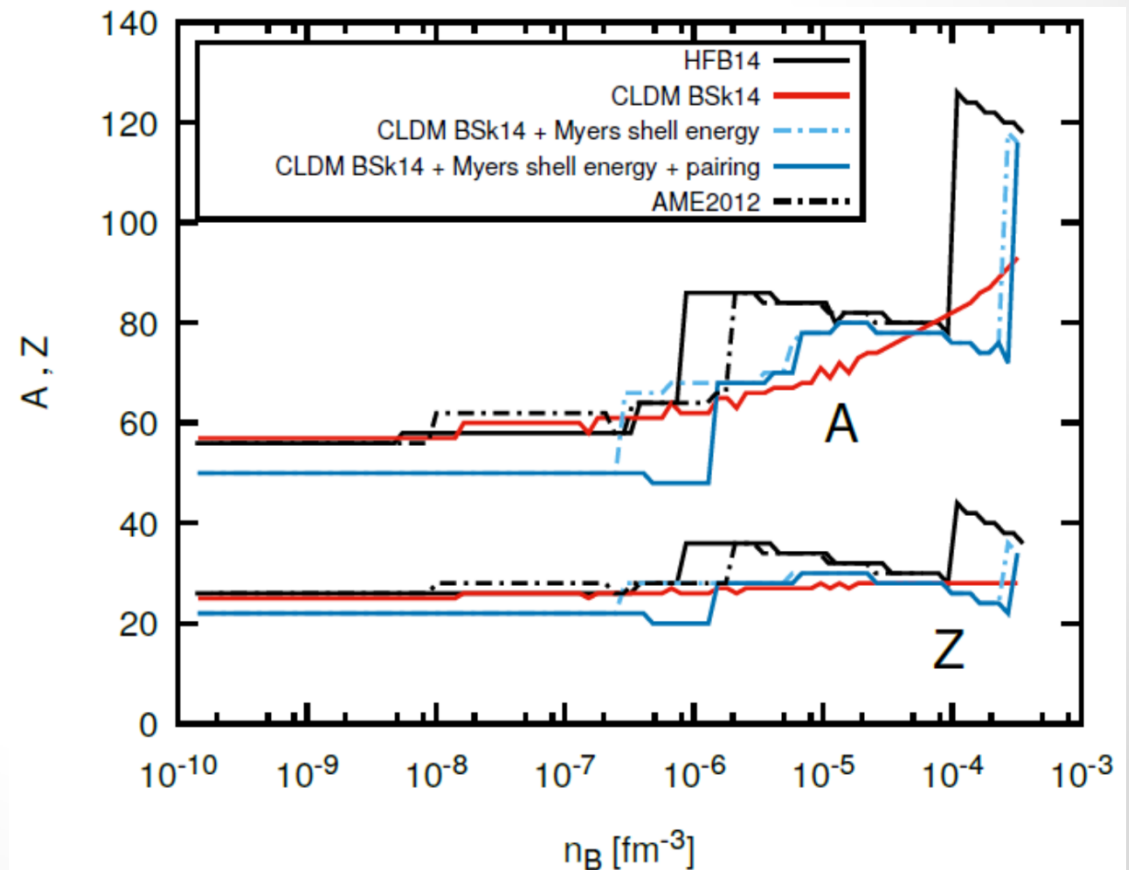
N.Chamel & P.Haensel, living reviews in Relativity 11 (2008)10

# Nuclear modeling & crust composition

$$\varepsilon_{WS}(n_B) = \min$$

1. Choose an energy functional (= set of empirical parameters  $X_n$ )
2. Choose a many-body treatment (HFB, HF, ETF, CLDM...)

=> The inner and outer crust composition depends on the many-body treatment



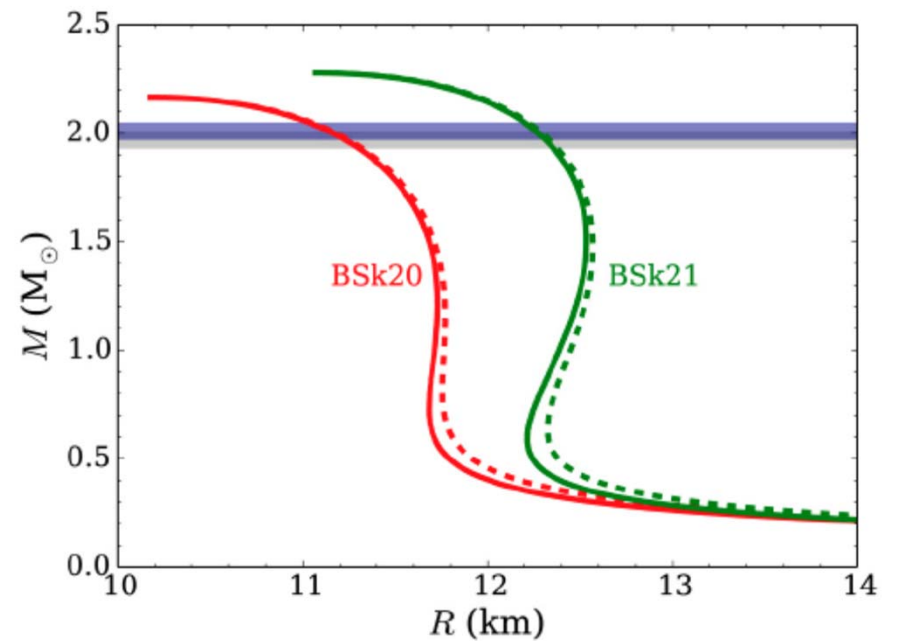
# Nuclear modeling & crust composition

$$\varepsilon_{WS}(n_B) = \min$$

1. Choose an energy functional (= set of empirical parameters  $X_n$ )
2. Choose a many-body treatment (HFB, HF, ETF, CLDM...)

=> The inner and outer crust composition depends on the many-body treatment

=> But the uncertainty on astro observables due to the many body treatment is << than the error due to the EDF !!



M.Fortin et al. PRC 94, 035804 (2016)

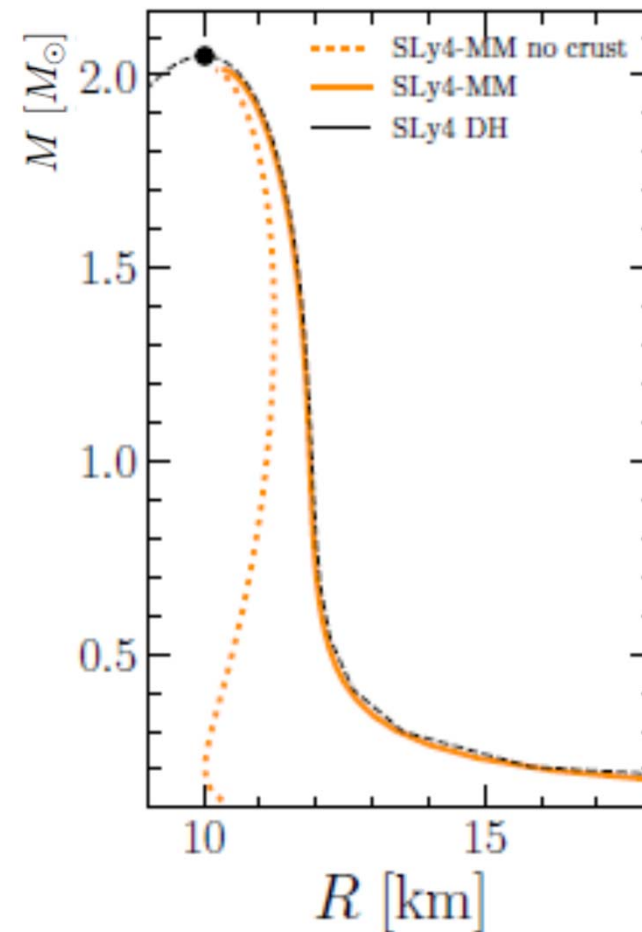
# Nuclear modeling & crust composition

$$\varepsilon_{WS}(n_B) = \min$$

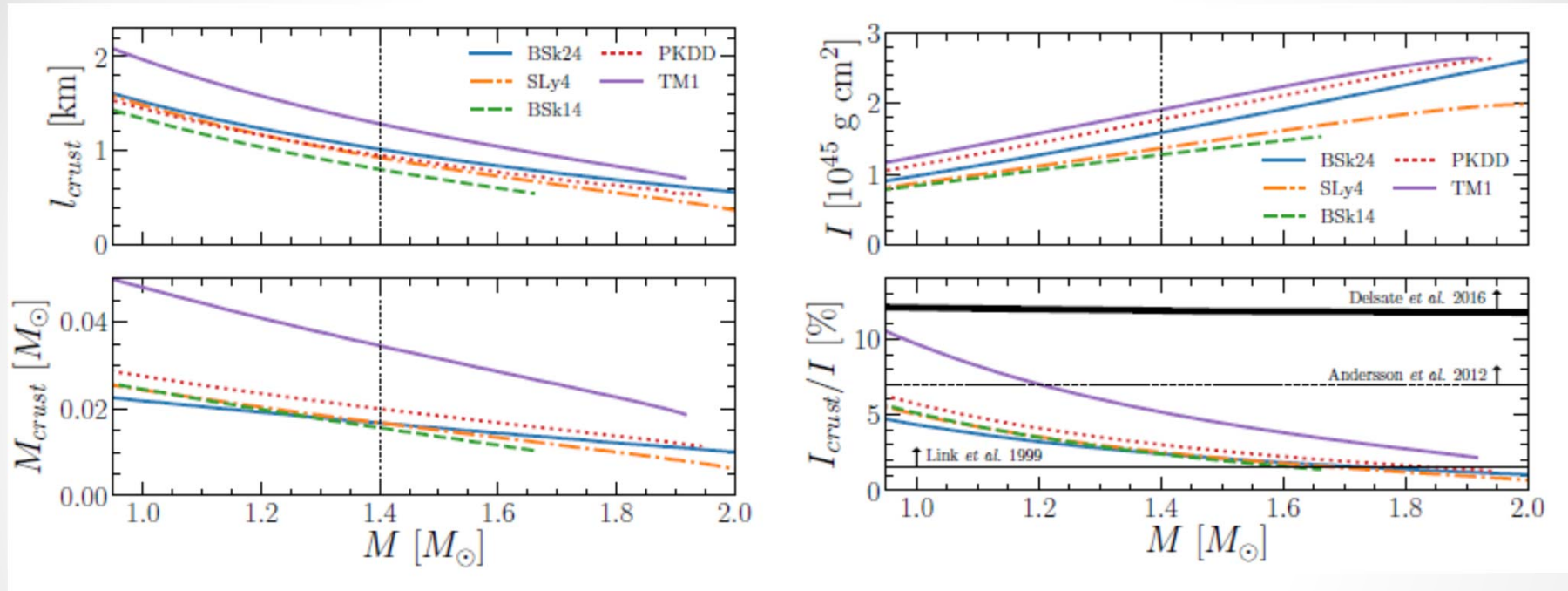
1. Choose an energy functional (= set of empirical parameters  $X_n$ )
  2. Choose a many-body treatment (HFB, HF, ETF, CLDM...)
- **Transition between crust and core:  $\varepsilon_{WS}^{hom.}(n_B) = \varepsilon_{WS}^{inhom.}(n_B)$**
  - **All static observables can be calculated solving the TOV equation of hydrostatic equilibrium:**
  - **in particular the crustal width and moment of inertia (connection to pulsar glitches)**

# Plan

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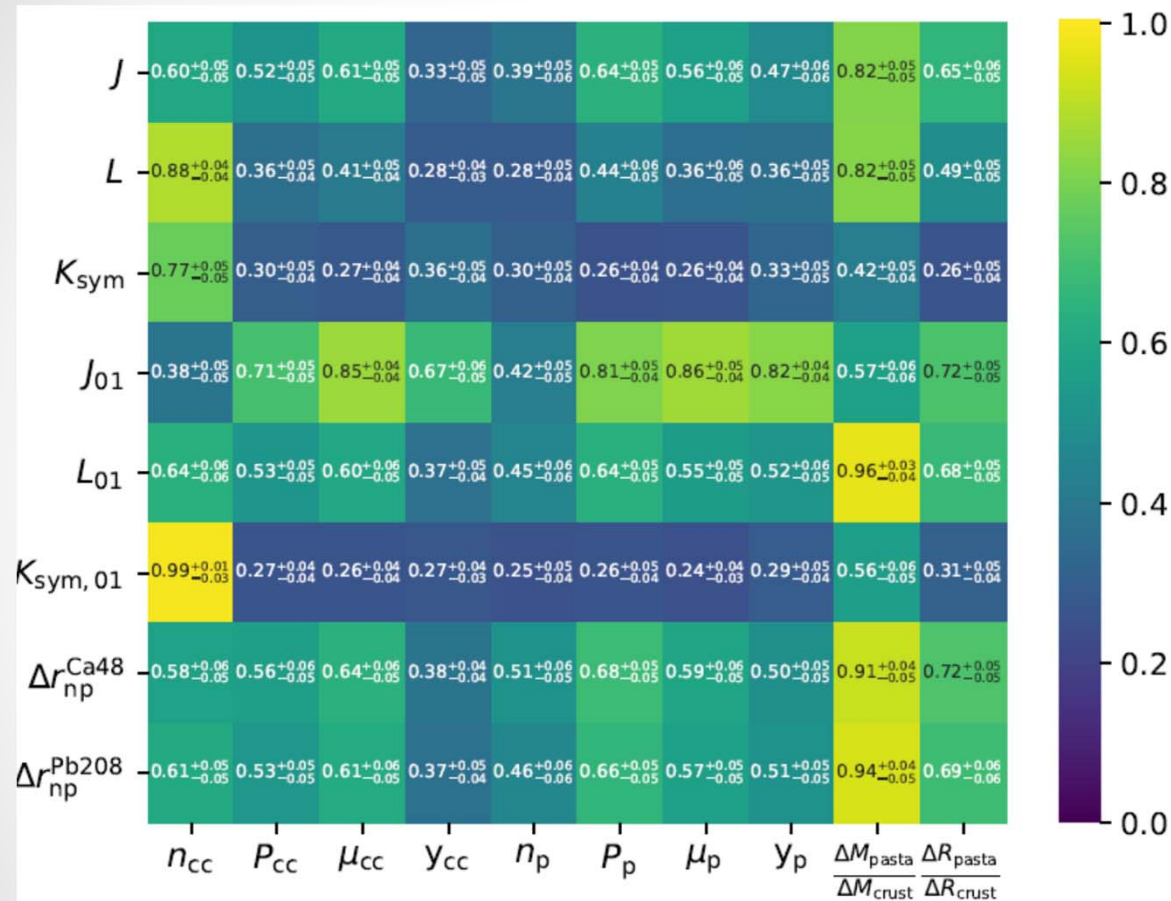
# A chosen set of models



Which EoS parameters play a role?

- T.Carreau, PhD thesis 2020

# Bayesian analysis



- Skyrme interactions and CLDM modeling
- EFT constraint from Drischler et al. included

=>  $J, L (K_{sym})$  are seen to rule crustal physics!

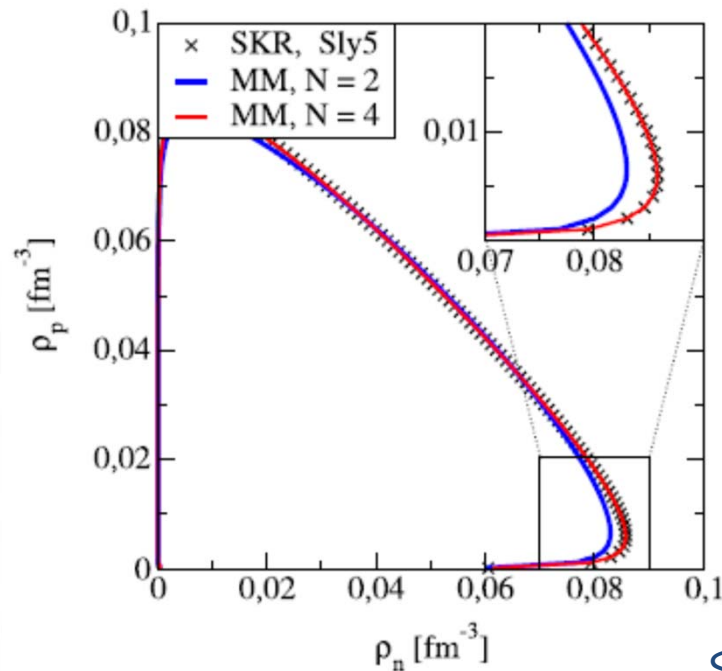
But there are complications...

L. Balliet et al, arXiv:2009.07696

# Beyond J&L: a consistent Bayesian analysis of crust properties

1. The Skyrme functional is not flexible enough

- within Skyrme,  $Q_{\text{sym}}$ ,  $Z_{\text{sym}}$ ,  $Q_{\text{sat}}$ ,  $Z_{\text{sat}}$  are uniquely set from low order parameters



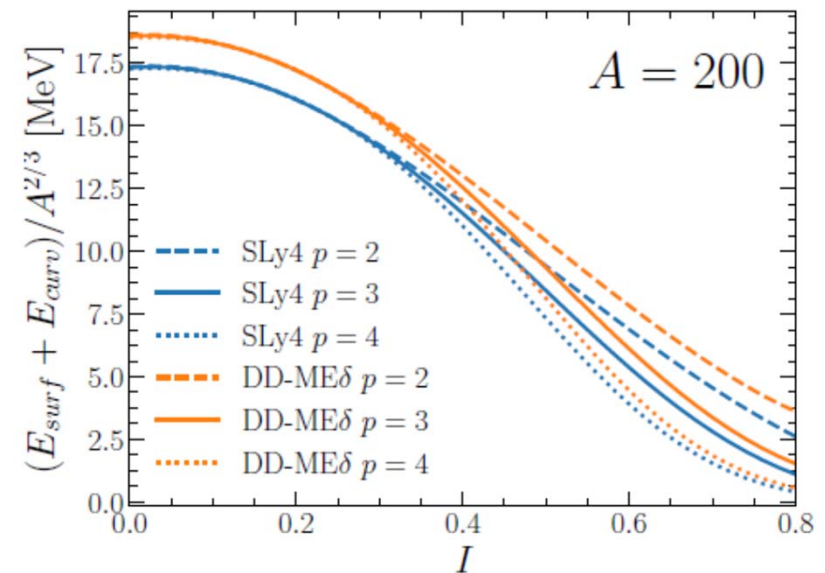
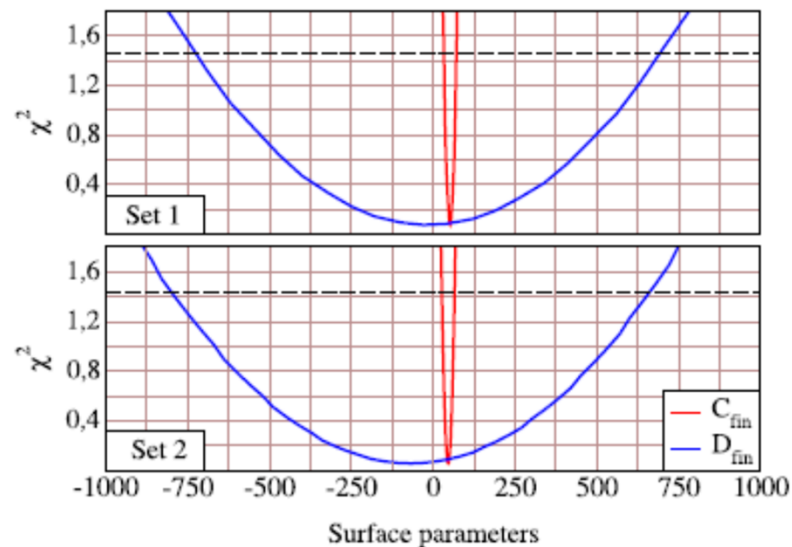
=> the uncertainty on the high order parameters must be included in the analysis

S.Antic et al, Jphys G 46 (2019) 065109



# Beyond J&L: a consistent Bayesian analysis of crust properties

- The energy of inhomogeneous matter depends on the surface tension (= gradient terms) in addition to the EoS



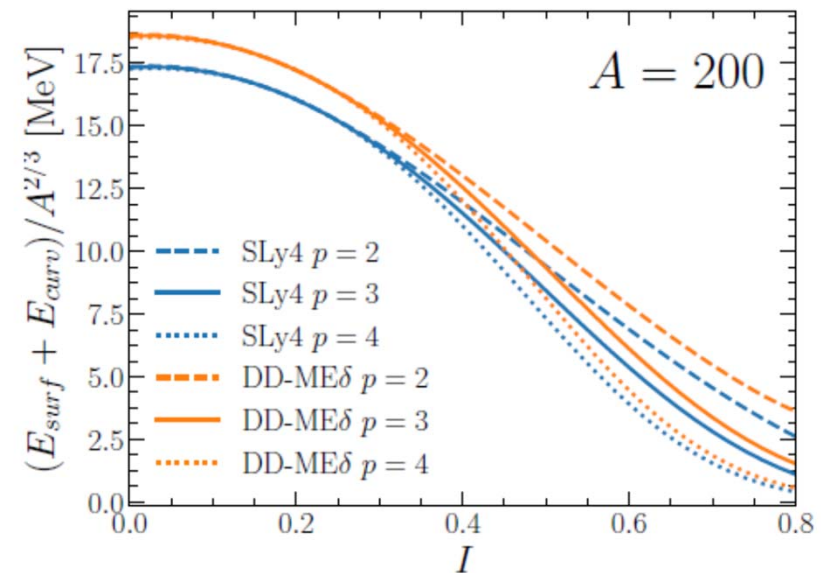
T.Carreau, PhD thesis 2020

S.Antic et al, *Jphys G* 46 (2019) 065109

# Beyond J&L: a consistent Bayesian analysis of crust properties

2. The energy of inhomogeneous matter depends on the surface tension (= gradient terms) in addition to the EoS

- ⇒ The surface tension must be calculated consistently with the bulk (fit on neutron rich nuclei)
- ⇒ The uncertainty at extreme  $I$  must be included in the Bayesian analysis

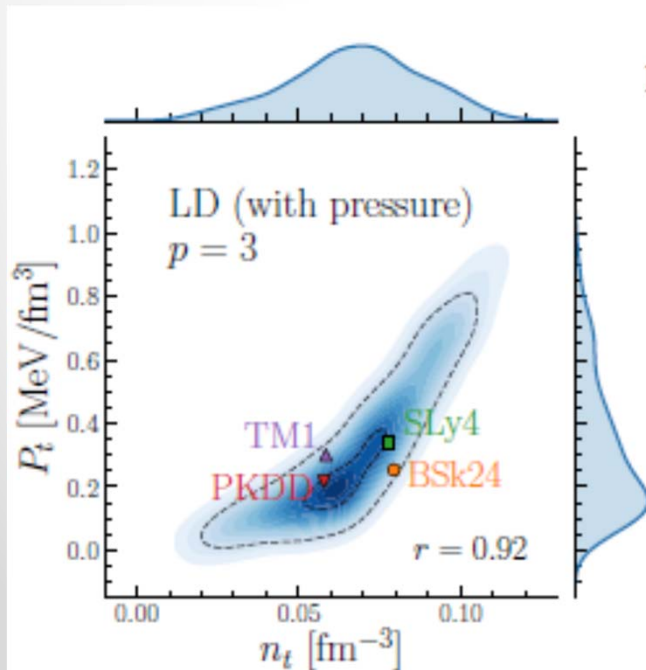


T.Carreau, PhD thesis 2020

# Results

	$n_t$														
LD+HD	-0.11-0.02	0.41	-0.51	0.25	-0.08	-0.26	-0.16	0.43	-0.14	0.11	0.11	0.15	-0.10	-0.08	
LD+HD $p = 3$	-0.10-0.02	0.42	-0.51	0.24	-0.11	-0.29	-0.17	0.46	-0.16	0.00	0.10	0.21	-0.10	-0.13	
LD (w/ $P$ ) $p = 3$	-0.02-0.05	0.12	-0.54	0.05	-0.06	-0.27	0.10	0.59	-0.11	0.00	0.02	0.09	-0.04	-0.06	
LD $p = 3$	-0.08	0.02	0.29	-0.48	0.07	-0.05	-0.20	-0.04	0.47	-0.10	0.00	0.08	0.13	-0.05	-0.11
HD $p = 3$	0.23	0.07	0.30	-0.52	0.25	-0.16	-0.21	-0.06	0.33	-0.13	0.00	-0.23	-0.01	0.17	0.06
Prior $p = 3$	0.17	0.06	0.25	-0.52	0.06	-0.01	-0.22	-0.12	0.32	-0.04	0.00	-0.18	-0.10	0.21	-0.03
	$E_{\text{sat}}$	$n_{\text{sat}}$	$K_{\text{sat}}$	$Q_{\text{sat}}$	$Z_{\text{sat}}$	$E_{\text{sym}}$	$L_{\text{sym}}$	$K_{\text{sym}}$	$Q_{\text{sym}}$	$Z_{\text{sym}}$	$p$	$\sigma_0$	$b_s$	$\sigma_{0c}$	$\beta$

	$P_t$														
LD+HD	-0.08-0.00	0.31	-0.39	0.18	0.06	-0.26	-0.38	0.47	-0.14	0.12	0.07	0.11	-0.01	-0.07	
LD+HD $p = 3$	-0.07-0.01	0.32	-0.39	0.18	0.05	-0.29	-0.41	0.49	-0.16	0.00	0.07	0.03	-0.01	-0.10	
LD (w/ $P$ ) $p = 3$	0.00-0.02	0.11	-0.34	0.02	-0.09	-0.40	0.15	0.74	-0.13	0.00	-0.00	0.08	-0.00	-0.03	
LD $p = 3$	-0.04	0.02	0.19	-0.31	0.06	0.10	-0.19	-0.23	0.46	-0.08	0.00	0.03	-0.05	0.04	-0.06
HD $p = 3$	0.22	0.05	0.27	-0.46	0.19	0.05	-0.16	-0.28	0.39	-0.14	0.00	-0.23	-0.17	0.26	-0.00
Prior $p = 3$	0.15	0.03	0.19	-0.42	0.04	0.15	-0.05	-0.24	0.36	-0.04	0.00	-0.16	-0.20	0.24	-0.06
	$E_{\text{sat}}$	$n_{\text{sat}}$	$K_{\text{sat}}$	$Q_{\text{sat}}$	$Z_{\text{sat}}$	$E_{\text{sym}}$	$L_{\text{sym}}$	$K_{\text{sym}}$	$Q_{\text{sym}}$	$Z_{\text{sym}}$	$p$	$\sigma_0$	$b_s$	$\sigma_{0c}$	$\beta$



The correlations move towards high order parameters....

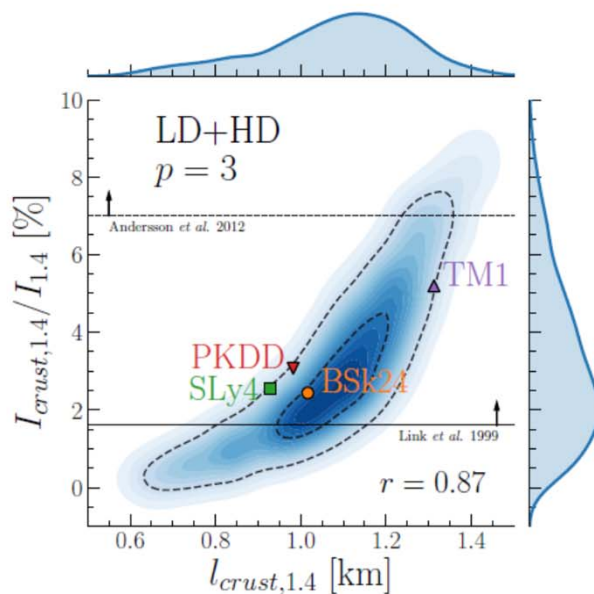
# Results

$l_{crust,1.4}$

LD+HD	-0.03	-0.13	0.37	-0.30	0.22	-0.05	-0.11	0.05	0.52	-0.11	0.10	0.04	0.11	-0.07	-0.04
LD+HD $p = 3$	-0.02	-0.14	0.38	-0.29	0.20	-0.07	-0.14	0.04	0.54	-0.13	0.00	0.02	0.09	-0.06	-0.05
LD (w/ $P$ ) $p = 3$	-0.00	-0.05	0.24	-0.17	0.18	0.00	-0.30	0.08	0.67	-0.17	0.00	0.00	0.02	0.01	-0.04
LD $p = 3$	-0.03	-0.10	0.36	-0.20	0.12	0.02	-0.01	0.04	0.51	-0.09	0.00	0.03	0.03	-0.05	-0.07
HD $p = 3$	0.26	-0.14	0.21	-0.27	0.22	0.35	0.09	0.16	0.35	-0.12	0.00	-0.28	-0.40	0.37	-0.07
Prior $p = 3$	0.22	-0.12	0.23	-0.25	0.13	0.32	0.08	0.11	0.36	-0.09	0.00	-0.24	-0.36	0.34	-0.11
	$E_{sat}$	$n_{sat}$	$K_{sat}$	$Q_{sat}$	$Z_{sat}$	$E_{sym}$	$L_{sym}$	$K_{sym}$	$Q_{sym}$	$Z_{sym}$	$p$	$\sigma_0$	$b^s$	$\sigma_{oc}$	$\beta$

$I_{crust,1.4}/I_{1.4}$

LD+HD	-0.06	-0.05	0.33	-0.35	0.18	0.05	-0.19	-0.28	0.52	-0.14	0.13	0.06	0.12	-0.03	-0.06
LD+HD $p = 3$	-0.05	-0.06	0.34	-0.35	0.18	0.03	-0.22	-0.30	0.54	-0.16	0.00	0.05	0.03	-0.02	-0.08
LD (w/ $P$ ) $p = 3$	-0.00	-0.03	0.23	-0.24	0.10	-0.04	-0.33	-0.02	0.73	-0.24	0.00	0.00	0.05	0.00	-0.04
LD $p = 3$	-0.07	-0.03	0.30	-0.27	0.12	0.16	-0.13	-0.35	0.50	-0.08	0.00	0.07	-0.05	0.00	-0.11
HD $p = 3$	0.24	-0.03	0.29	-0.44	0.20	0.11	-0.03	-0.16	0.44	-0.15	0.00	-0.25	-0.21	0.27	-0.00
Prior $p = 3$	0.19	-0.03	0.25	-0.40	0.14	0.13	-0.01	-0.25	0.42	-0.09	0.00	-0.19	-0.21	0.24	-0.05
	$E_{sat}$	$n_{sat}$	$K_{sat}$	$Q_{sat}$	$Z_{sat}$	$E_{sym}$	$L_{sym}$	$K_{sym}$	$Q_{sym}$	$Z_{sym}$	$p$	$\sigma_0$	$b^s$	$\sigma_{oc}$	$\beta$



The correlations move towards high order parameters....

# Conclusions

- The NS crust properties depend in a complex way on the nuclear energy functional
- High order parameters  $Q_{sat}, Q_{sym}$  play an important role as well as surface properties at high isospin ratio
- How to progress?

# Strategies for a better estimation of crustal properties

- **Theory:** Implementation of low energy constraints from ab-initio theory *R. Somasundaram et al, arXiv:2009.04737*  
*S.Burrello, M.Grasso et al, in progress*

- **Experiment:** looking for observables sensitive to the different unknowns:

- global density variation in asymmetric matter (**GMR in isotopic chains ?**)

$$\Rightarrow K_{\tau} = K_{sym} - 6L_{sym} - L_{sym} \frac{Q_{sat}}{K_{sat}}$$

- isovector density variations at low density (**soft monopole?**)

$$\Rightarrow K_{sym}, Q_{sym}$$

- n-p density gradients (**skin ?**)  $\Rightarrow L_{sym}, D_{fin}$

To be analyzed with the same meta-modeling approach