

# KATRIN: from Neutrino Mass to Dark Matter

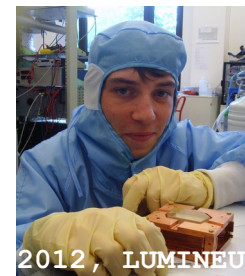
Séminaire du PHE, IJCLab

Thibaut Houdy

11<sup>th</sup> of October, 2021

# Who am I?

- **New comer to the IJCLab** as maître de conférence.
- Joining the DUNE experiment and **Fabien Cavalier's team**.
- 2014-2017: **PhD between APC/CEA** about solar and sterile neutrinos with the SOX project and the Borexino experiment.
- 2018-2021: **Postdoc in Susanne Mertens' group** in Max Planck Institute for Physics, Munich about KATRIN.
- Mostly working on neutrino/dark matter physics with an experimental taste.
- During my studies on Orsay's campus: internship @ CSNSM, lecture @ IPNO, Englert's conference on Higgs @ LAL ...
- Looking forward to collaborating with you all!



—► **KATRIN: from Neutrino Mass to Dark Matter**

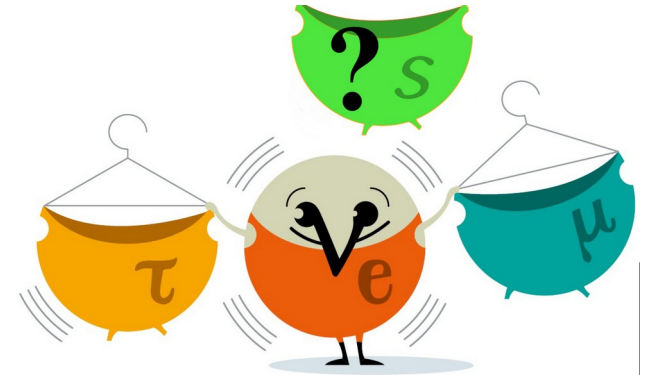
# This talk is about:

What is the mass  
of the neutrino?



What is dark matter?

Do sterile  
neutrinos  
exist?





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What is the mass  
of the neutrino?



*Beyond the  
Standard Model*



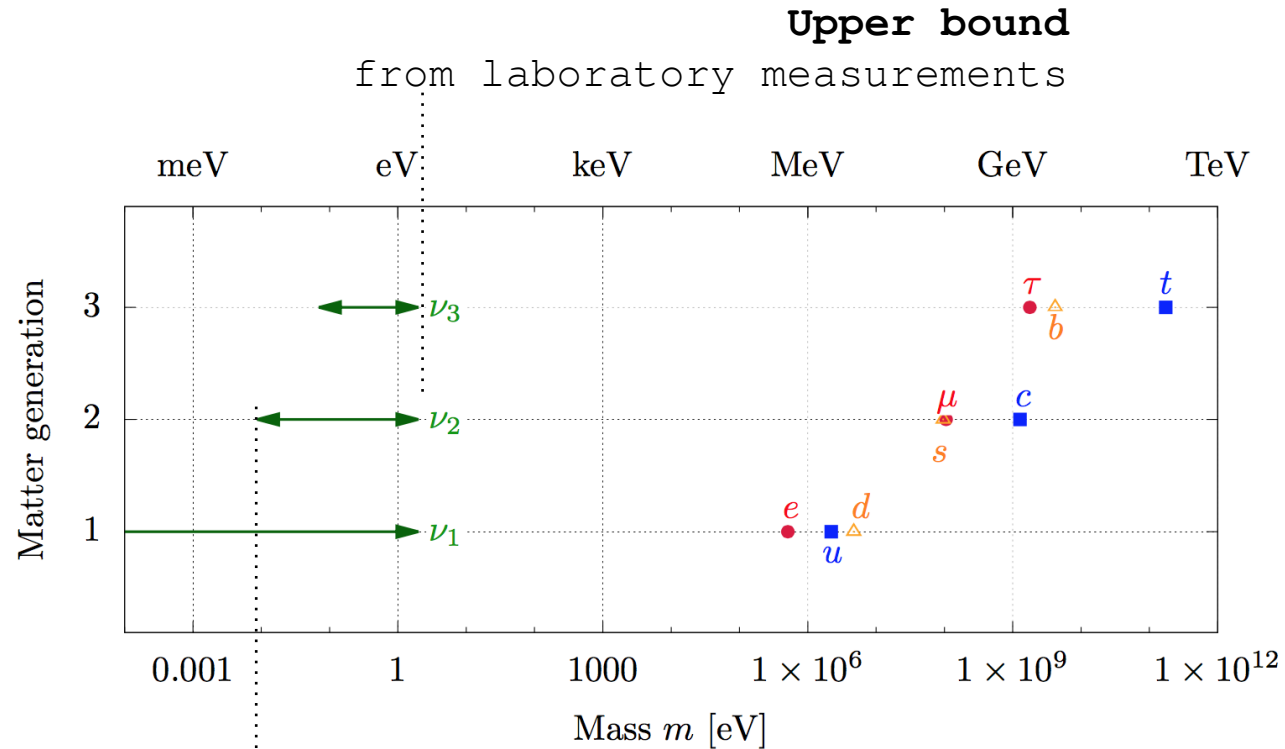
What is dark matter?

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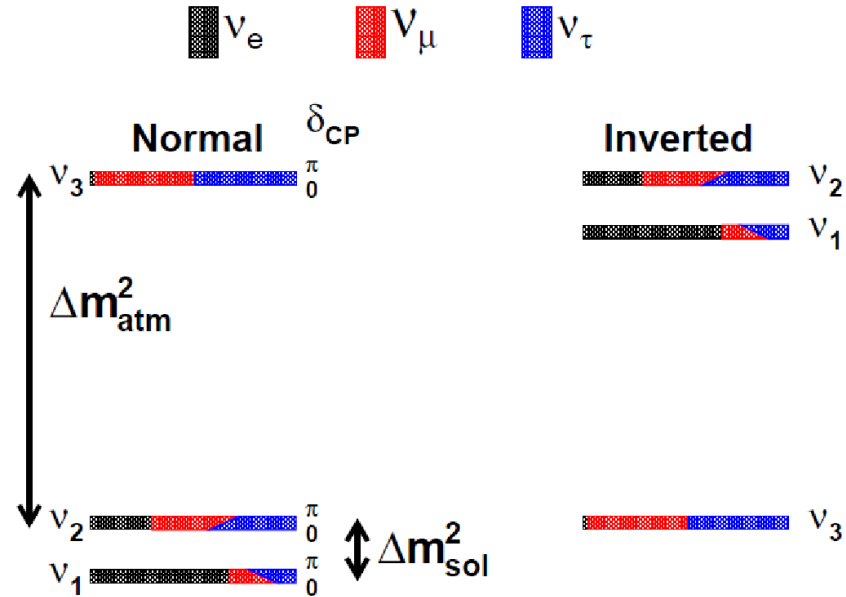
# Neutrino mass status



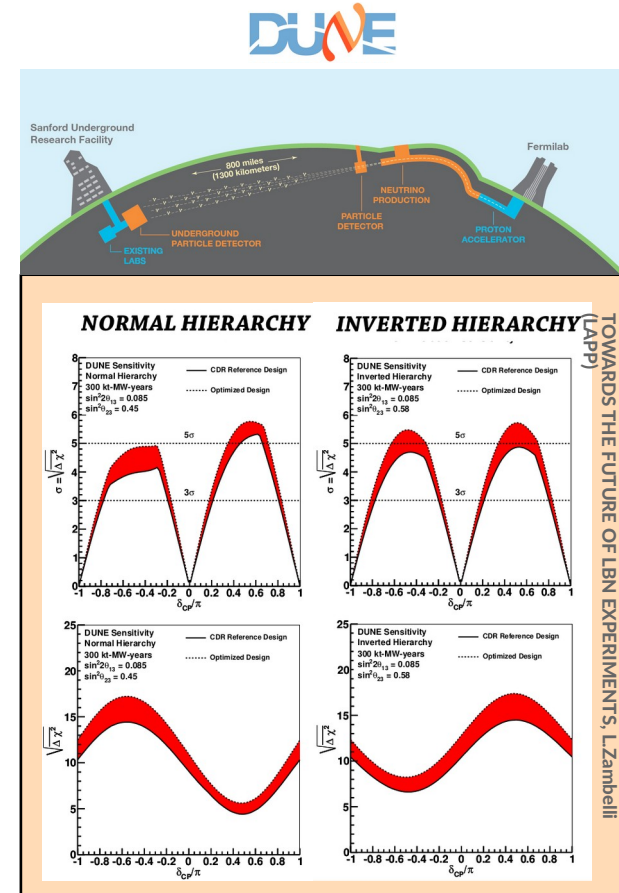
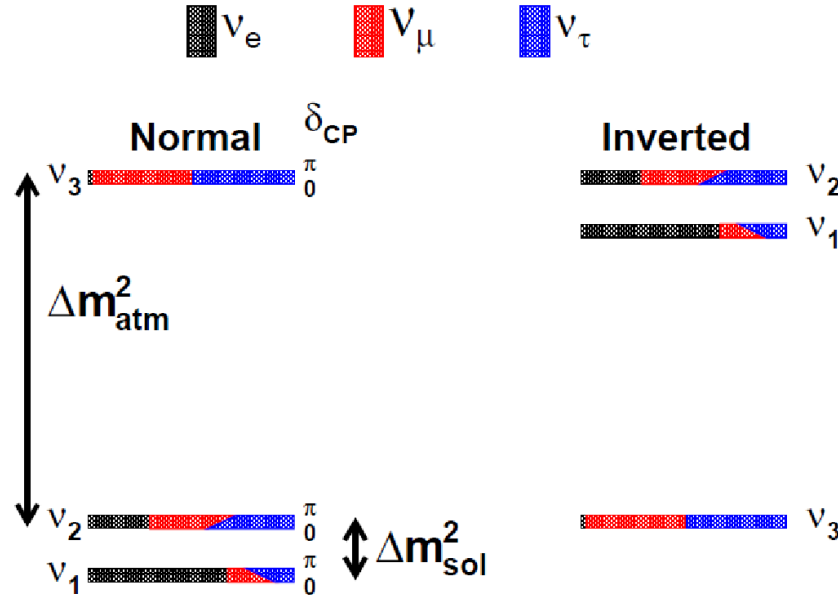
**Lower bound**

from oscillation experiments

# Neutrino mass hierarchy



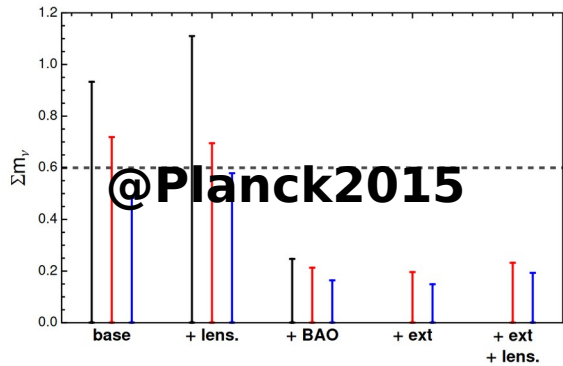
# Neutrino mass hierarchy



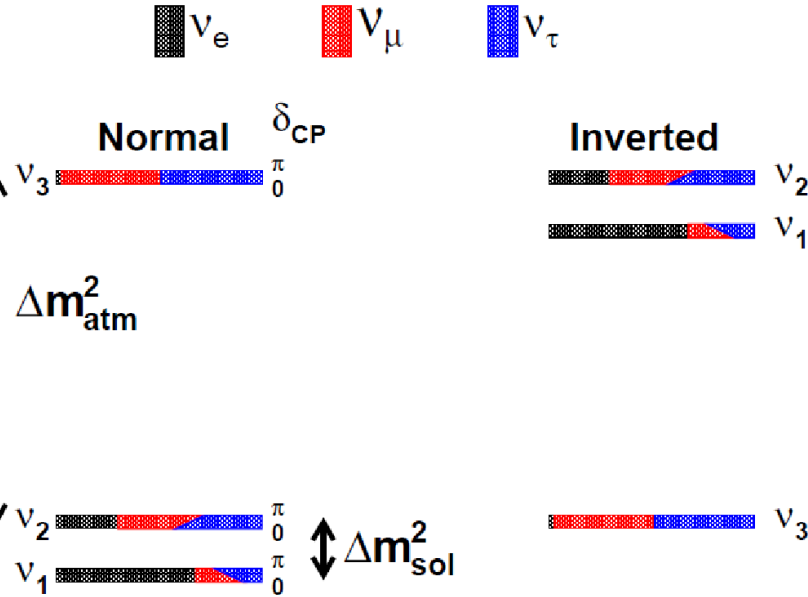


# Neutrino mass measurements

- **Cosmology**  
**model dependent**  
 $\Sigma m_\nu < 120 \text{ meV}$   
 Ex : Planck, eBOSS



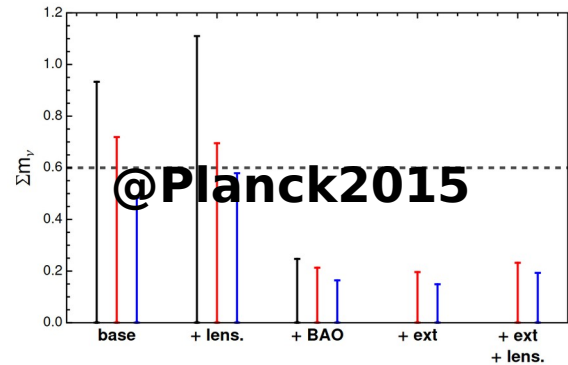
M.Lattanzi 2016 JPConf.Ser.718 032008



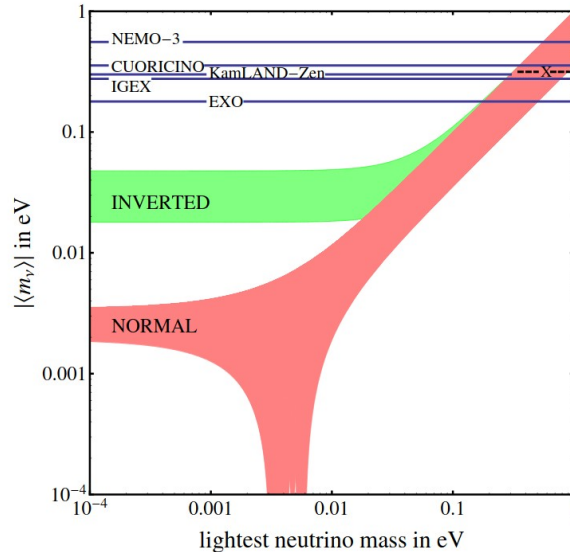
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- **$0\nu\beta\beta$**   
laboratory based  
 $m_{\beta\beta} < 100 \text{ meV}$   
Ex : CUPID, LEGEND, etc



M.Lattanzi 2016 JPConf.Ser.718 032008



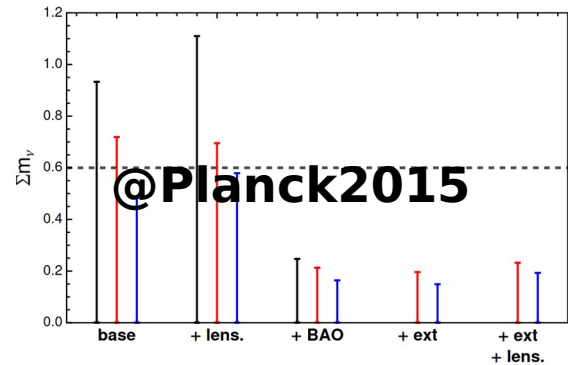
J. Barea et al. PRL 109, 042501 - (2012)

# Neutrino mass measurements

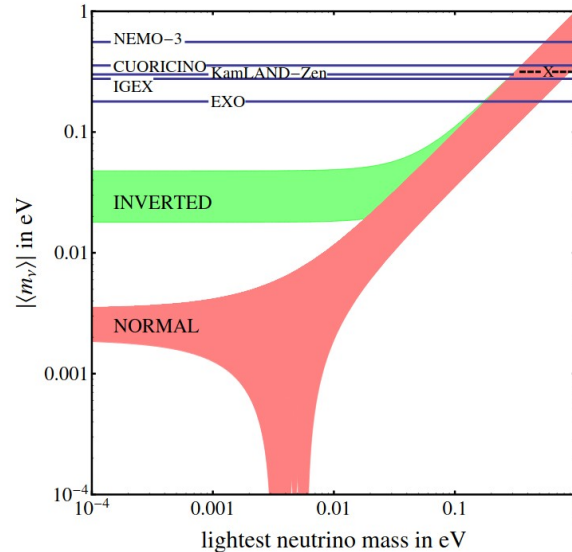
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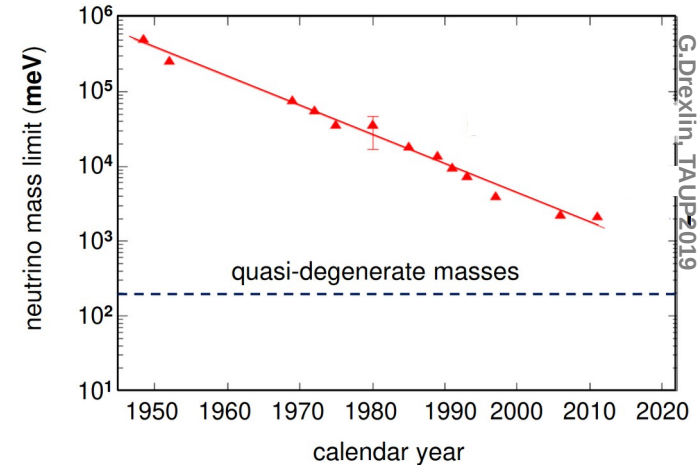
- **Kinematics of  $\beta$ -decay**  
**laboratory based**  
 $m_{\nu e} = 50\text{-}200 \text{ meV}$   
 Ex : ECHO, KATRIN, Project8



M.Lattanzi 2016 JPConf.Ser.718 032008

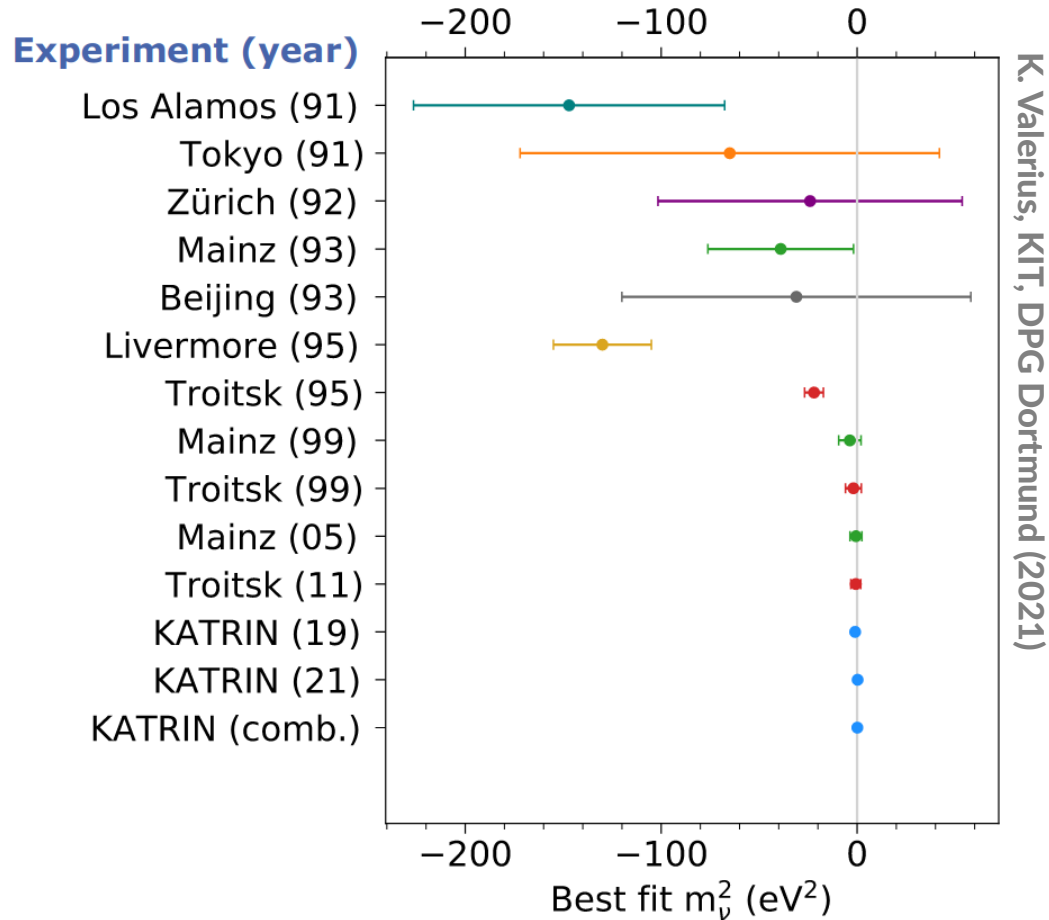


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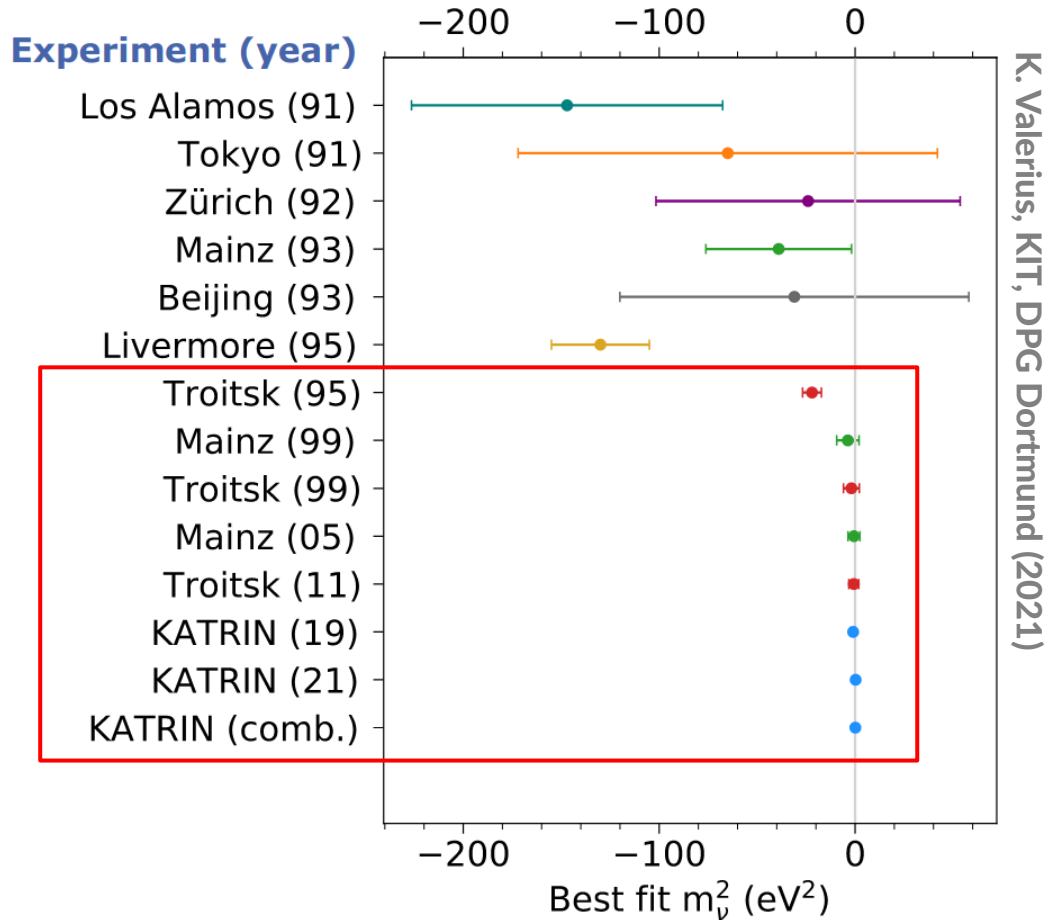




# Determining $m_\nu$ from $\beta$ -decay



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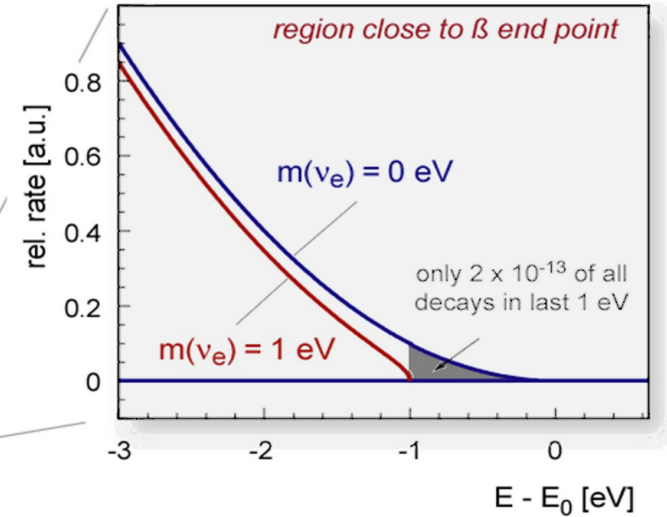
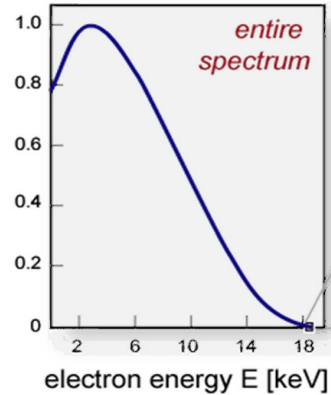
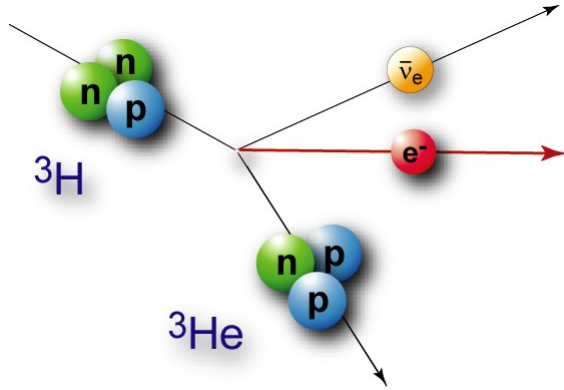


- Troitsk
- Mainz
- KATRIN

→ MAC-E filters with tritium source

# Determining $m_\nu$ from $\beta$ -decay: KATRIN

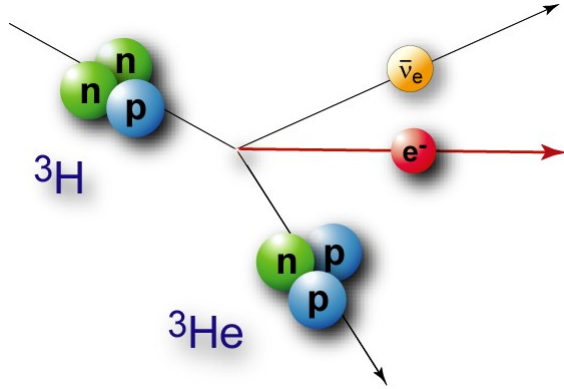
## General Idea



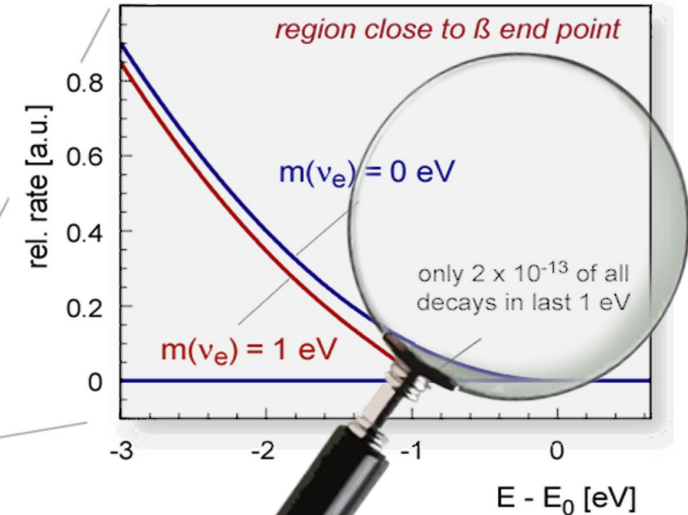
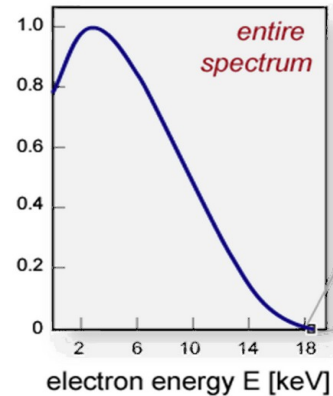


# Determining $m_\nu$ from $\beta$ -decay: KATRIN

## General Idea



- Ultra-strong  $\beta$ -source  $10^{11}$  decays/s
- Low background level  $< 0.1$  cps
- Excellent energy resolution  $\sim 1$  eV
- Precise understanding of spectrum



# Karlsruhe TRitium Neutrino Experiment : KATRIN



- Experimental site : Karlsruhe Institute of Technology (KIT)
- International Collaboration (150 members)
- Design sensitivity: 0.2 eV (90% CL)

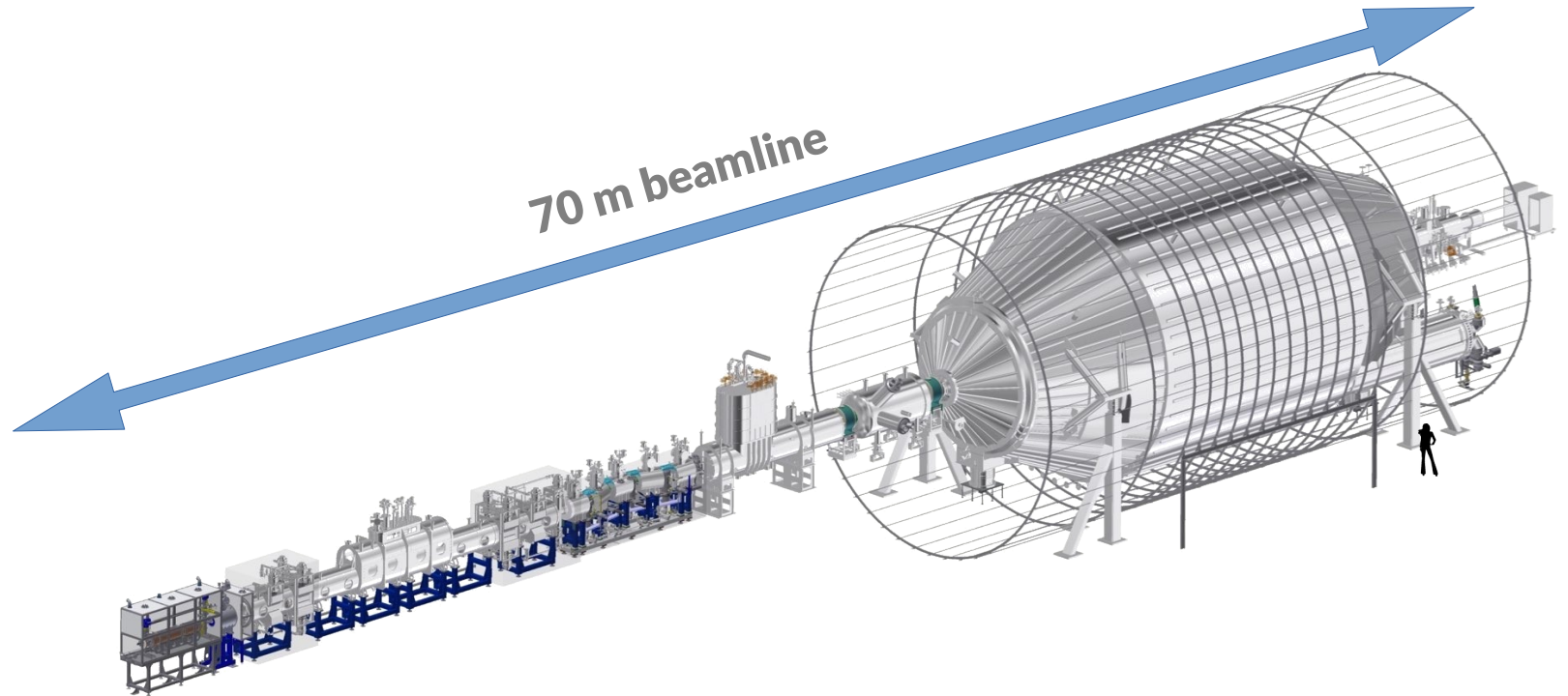


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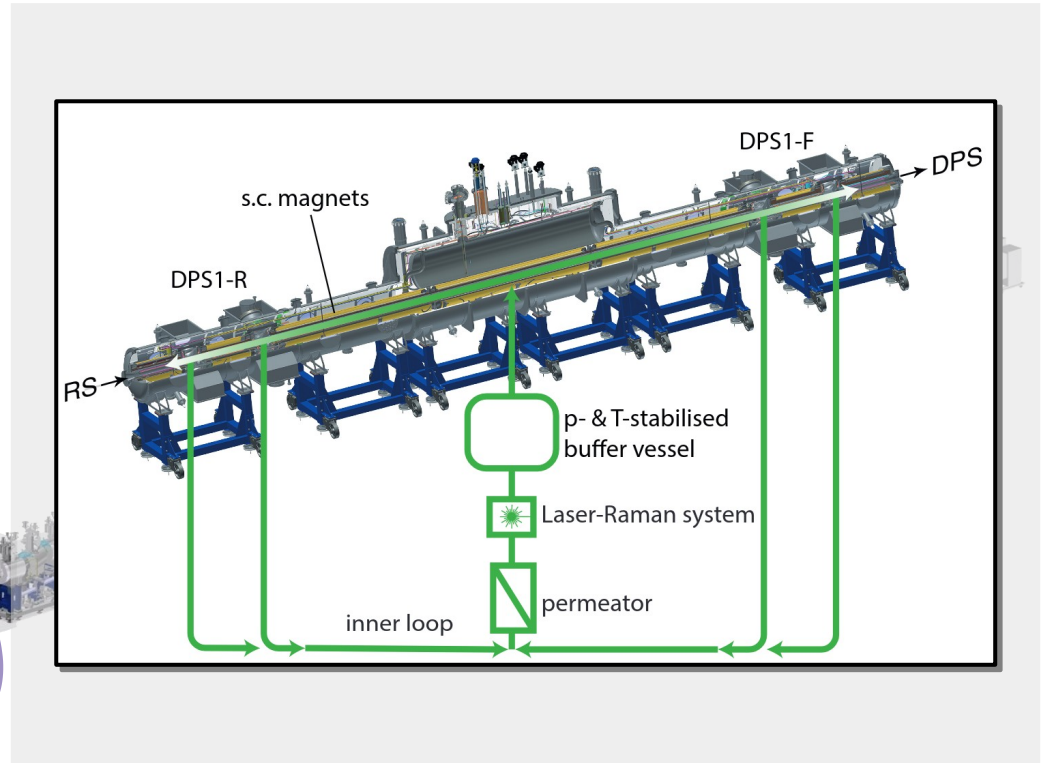
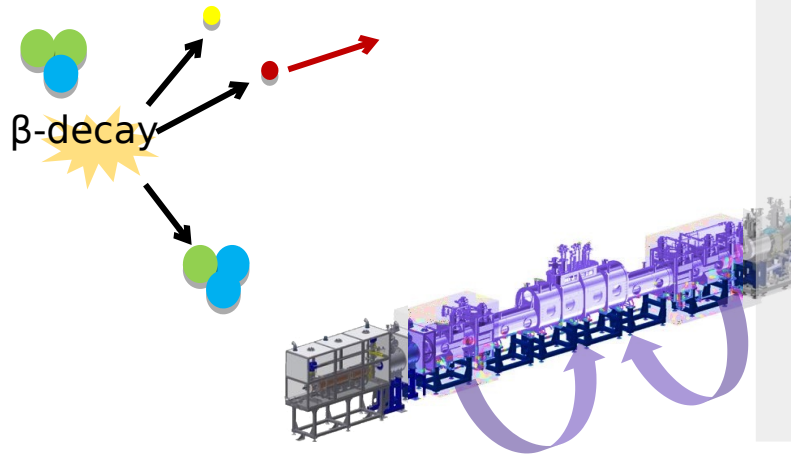
# KATRIN Working Principle



# KATRIN Working Principle

## Windowless gaseous tritium source

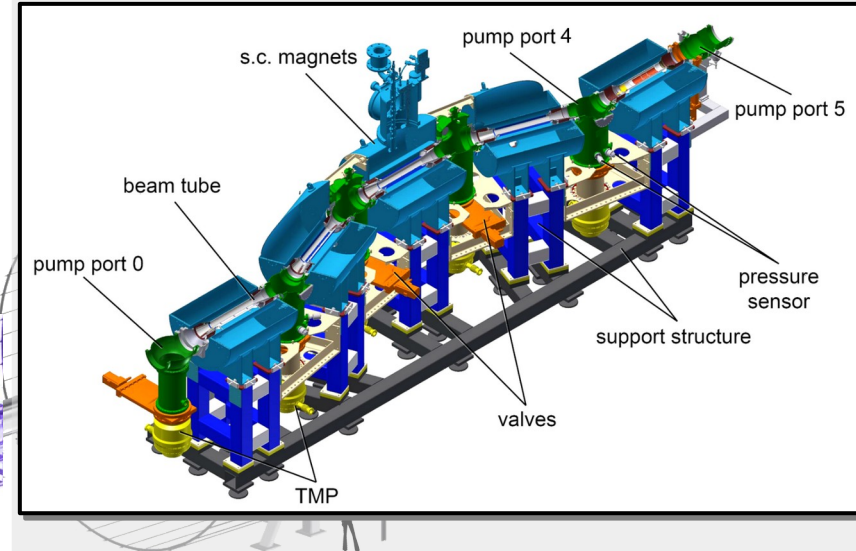
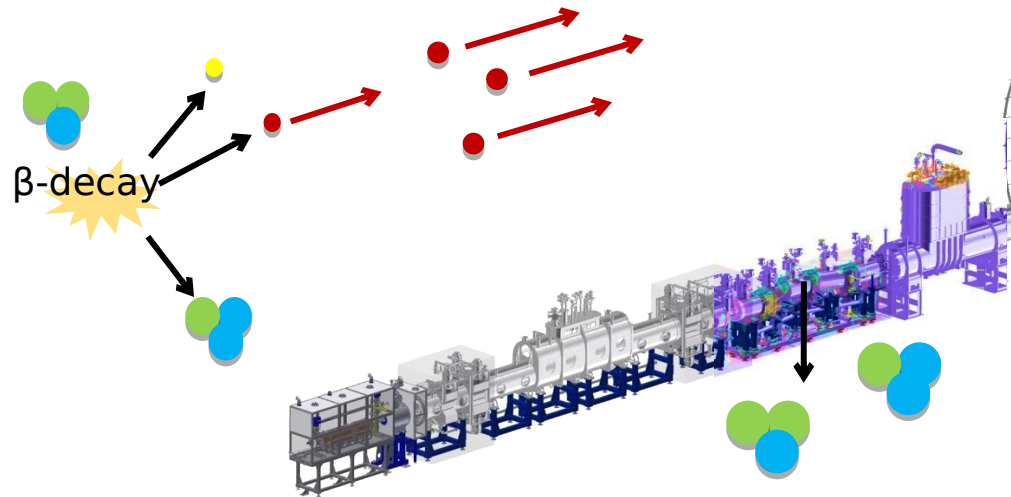
- molecular tritium in closed loop system
- $10^{11}$  decays/s



# KATRIN Working Principle

## Transport section

- magnetic guidance of electrons (@ 4 T)
- tritium flow reduction by  $> 10^{14}$  + tritium ion removal

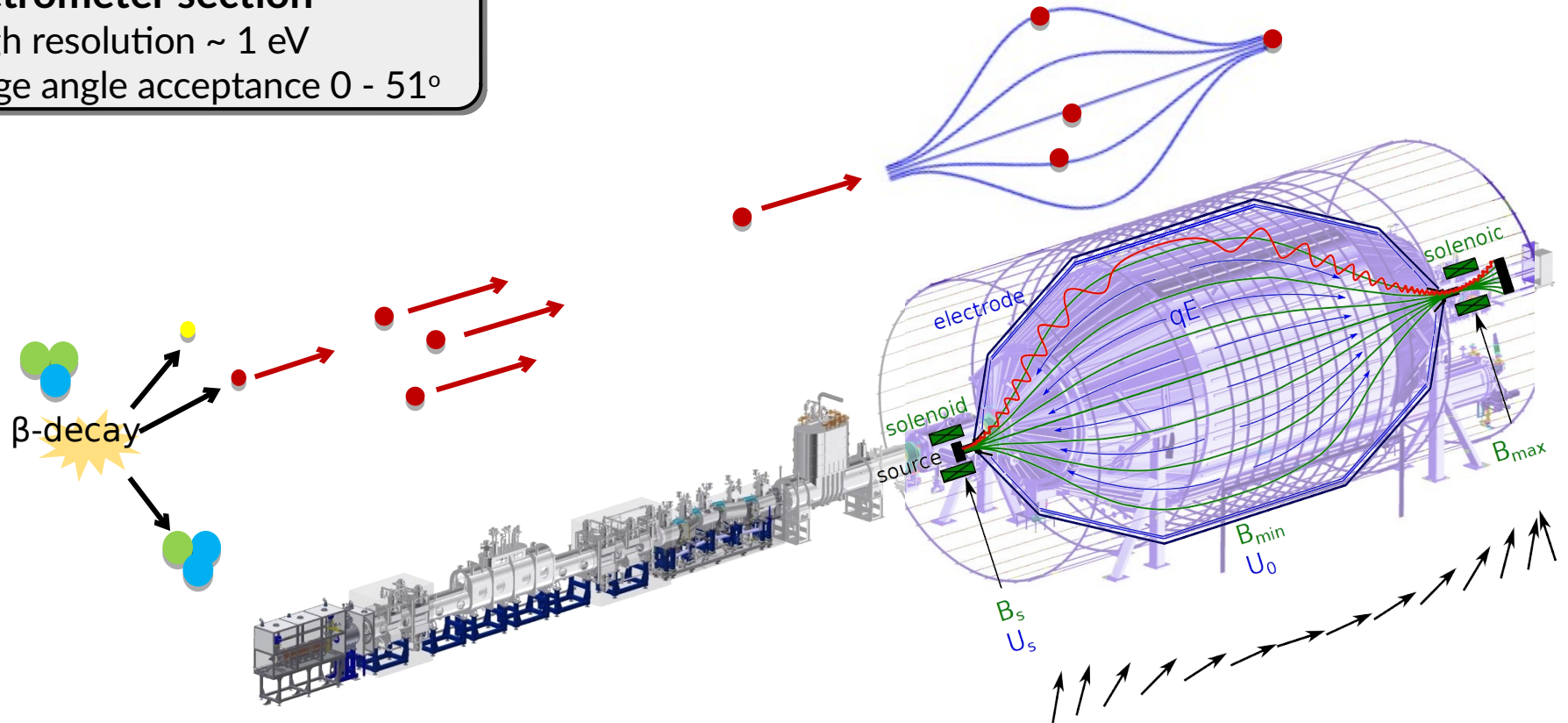




# KATRIN Working Principle

## Spectrometer section

- high resolution  $\sim 1$  eV
- large angle acceptance  $0 - 51^\circ$

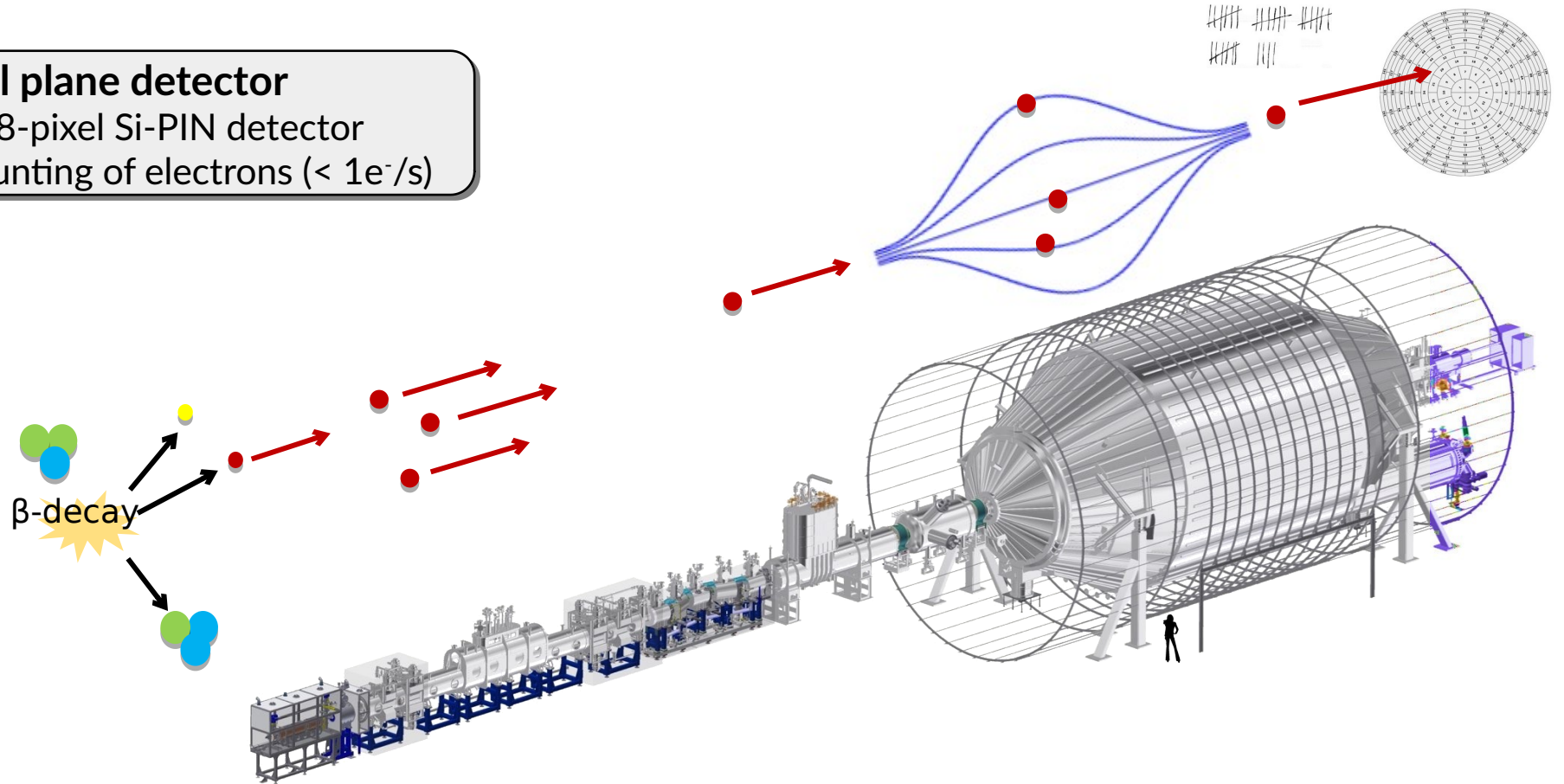




# KATRIN Working Principle

## Focal plane detector

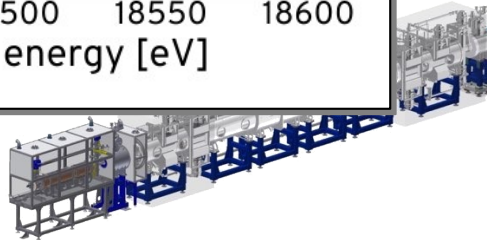
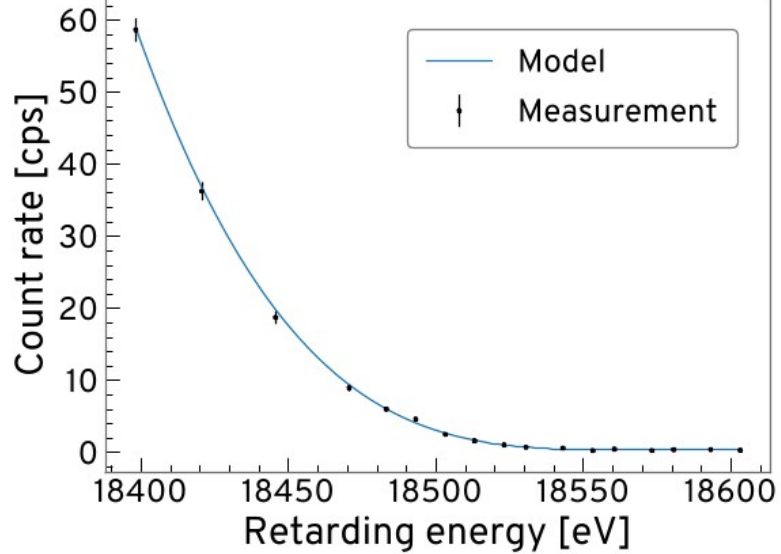
- 148-pixel Si-PIN detector
- counting of electrons ( $< 1e^-/s$ )



# KATRIN Working Principle

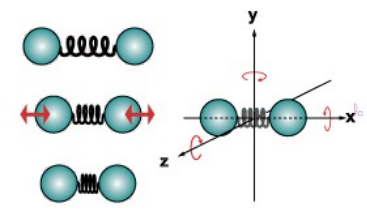
## Focal plane detector

- 148 pixel Si PIN detector



# Systematic uncertainties

## Molecular Final States



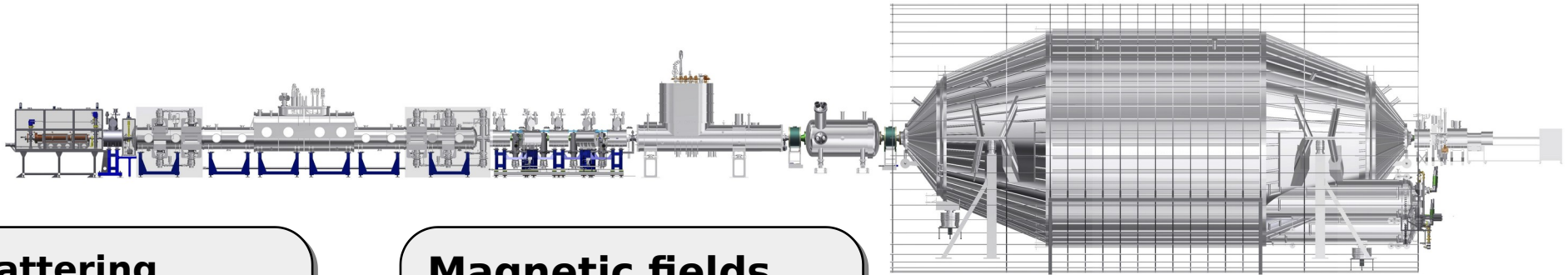
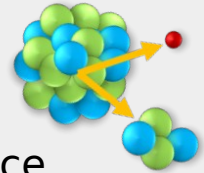
## Activity fluctuations

Challenge:  $< 0.1 \%$



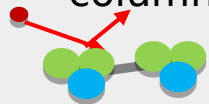
## Background:

- time correlation
- retarding potential dependence
- Challenge:  $< 0.1$  cps



## Scattering

- energy loss
- column density



## Magnetic fields

- source
- spectrometer
- detector



## Data combination

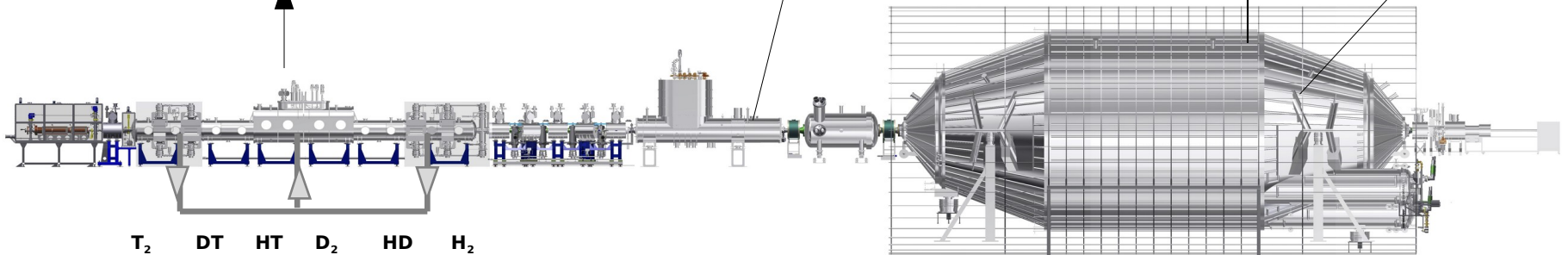
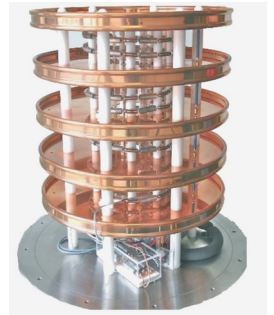
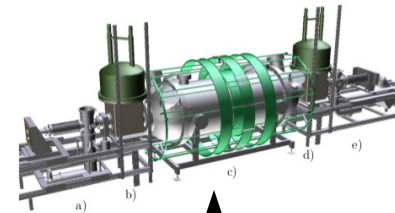
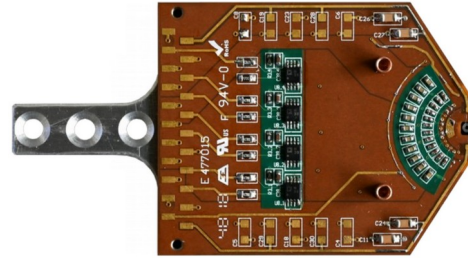
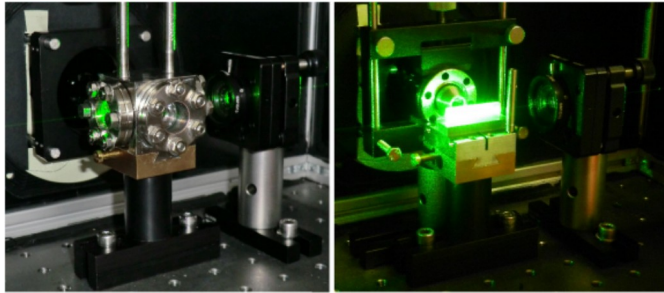


# Key Monitoring Devices

**Laser Raman system:** monitoring of tritium purity and gas composition at the 0.1% level

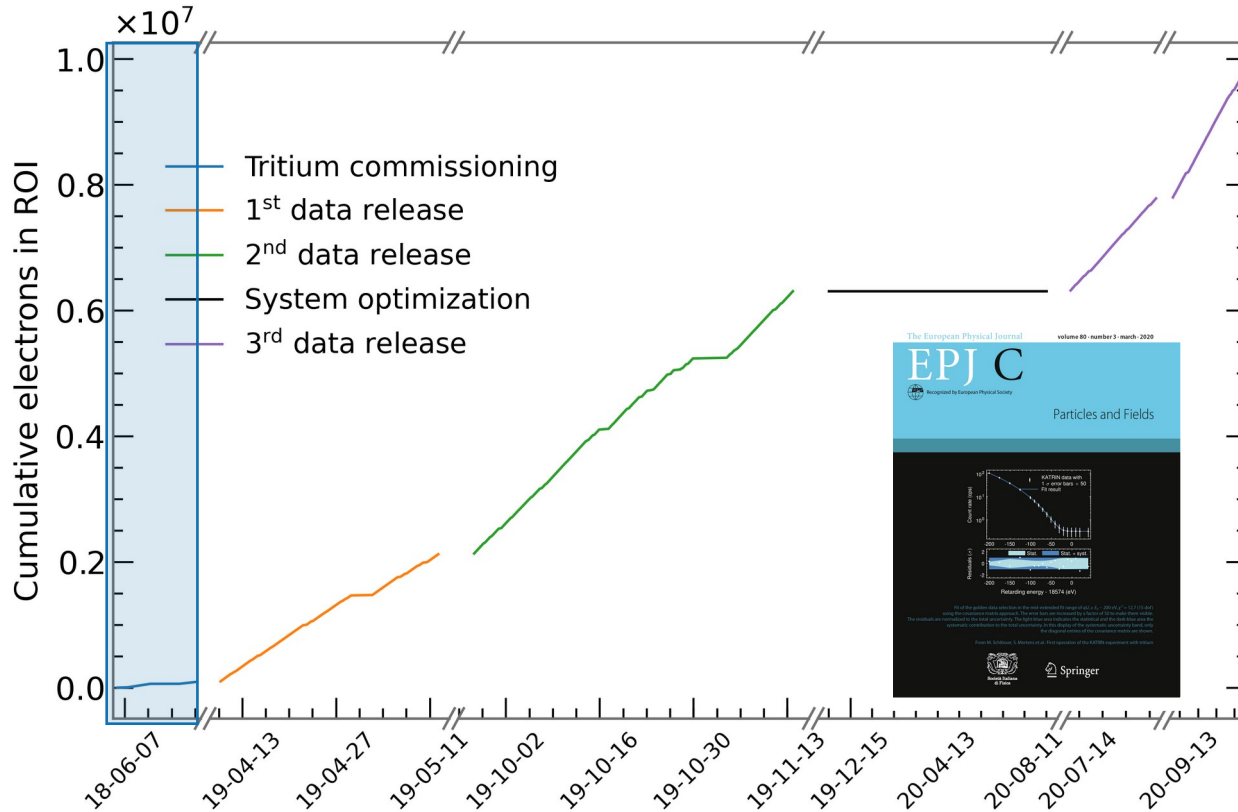
**Forward beam monitor:** monitoring of activity at the 0.1% level

**High voltage system:** monitoring of high voltage at the ppm level (20 mV)

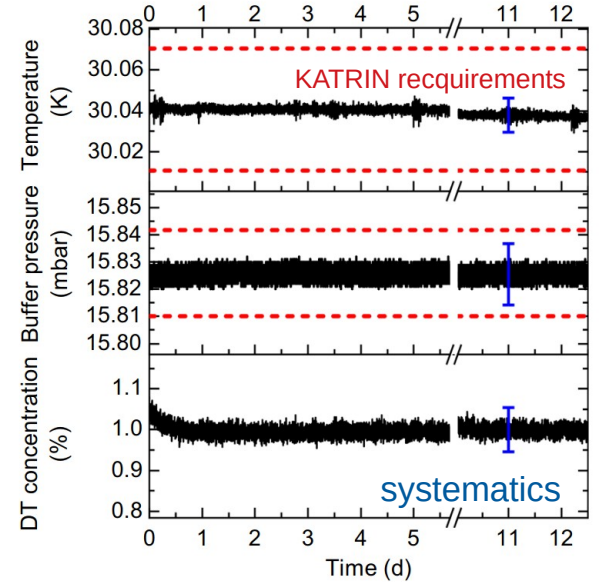




# First tritium campaign

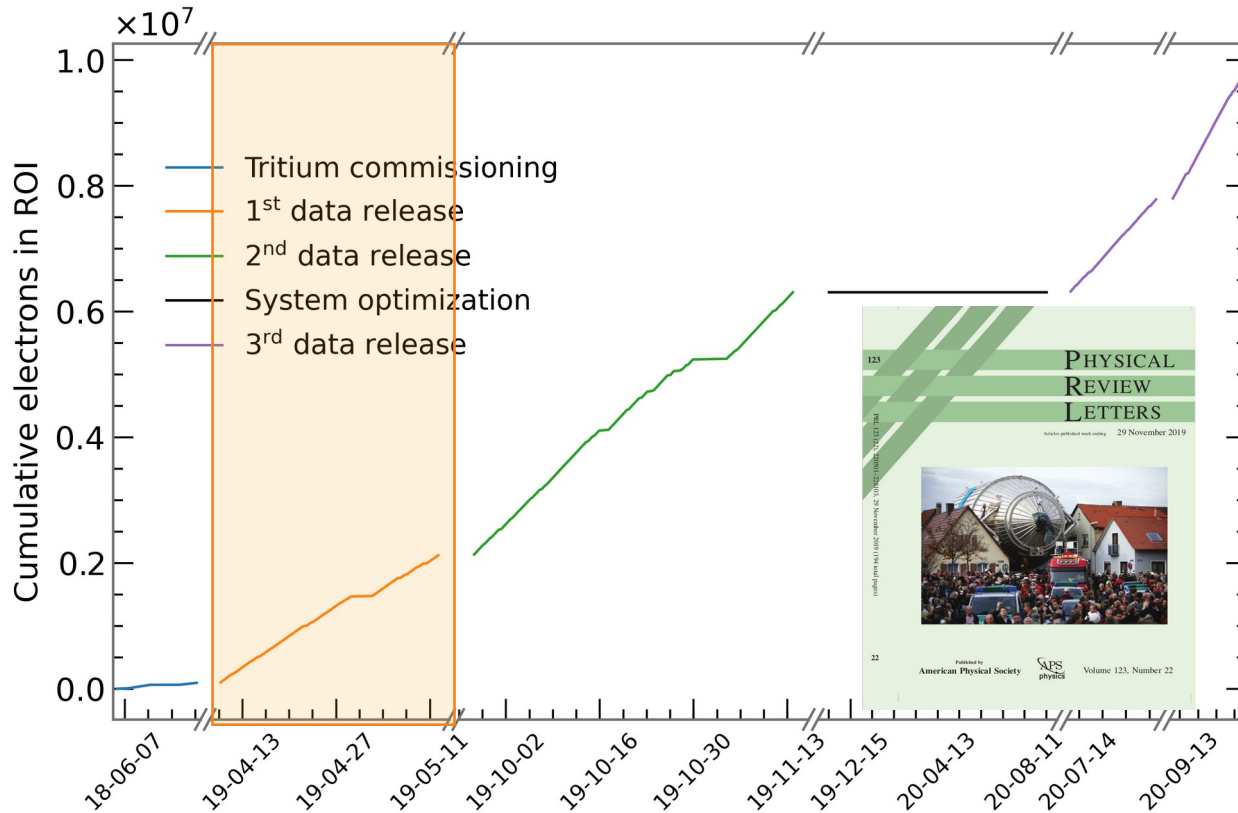


## demonstration of system stability

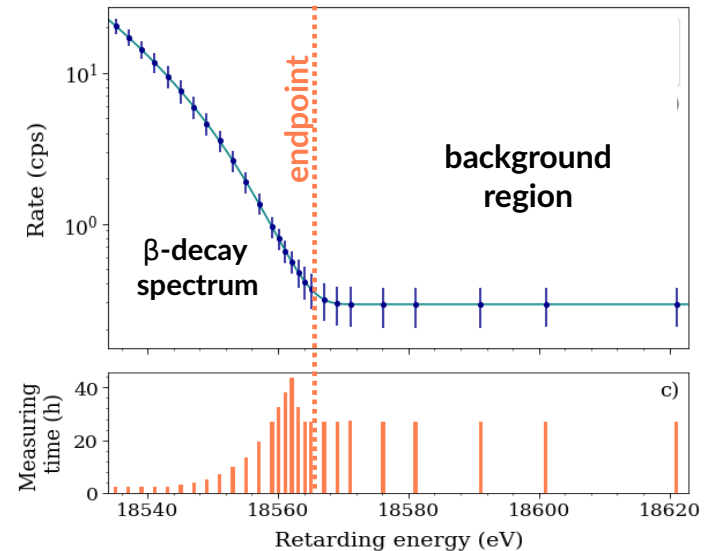


- First operation of the KATRIN experiment with tritium. *Eur. Phys. J. C* 80, 264 (2020)

# 1<sup>st</sup> neutrino mass campaign

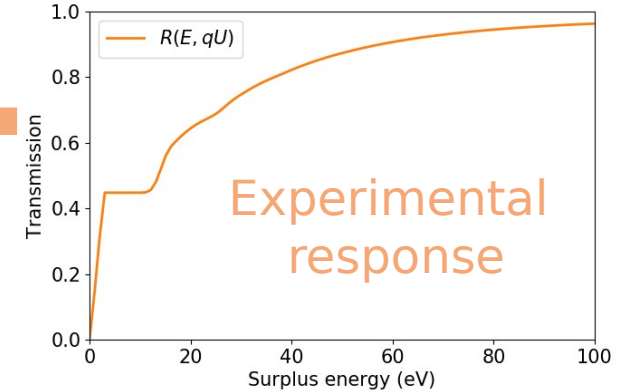
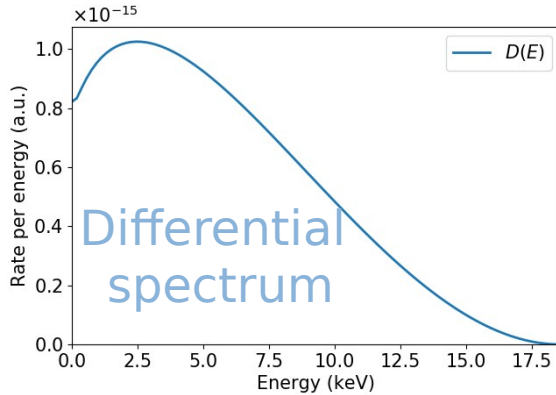


- Measurement time: **22 days**
- Gas density: **22%**
- Isotopic purity: **97.5% tritium**
- Source activity:  **$2.45 \cdot 10^{10}$  Bq**
- Total statistics:  **$2 \cdot 10^6$  e<sup>s</sup>**

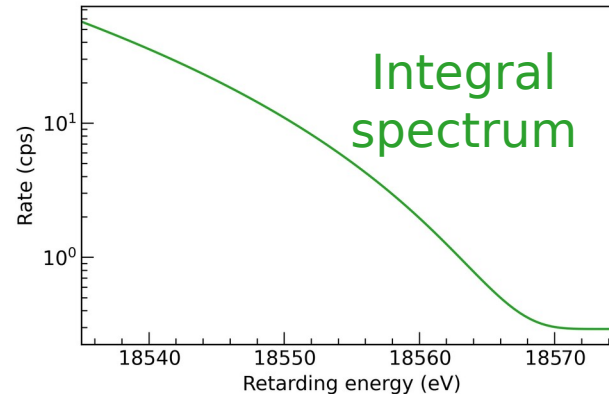


- *Improved Upper Limit on the Neutrino Mass from a Direct Kinematic Method* by KATRIN, KATRIN Collaboration, **Phys. Rev. Lett.** 123, 221802 (2019)

# Model



$$\Gamma(qU) \propto \mathbf{A} \cdot \int_{qU}^{E_0} D(E; m_v^2, E_0) \cdot R(qU, E) dE + \mathbf{B}$$



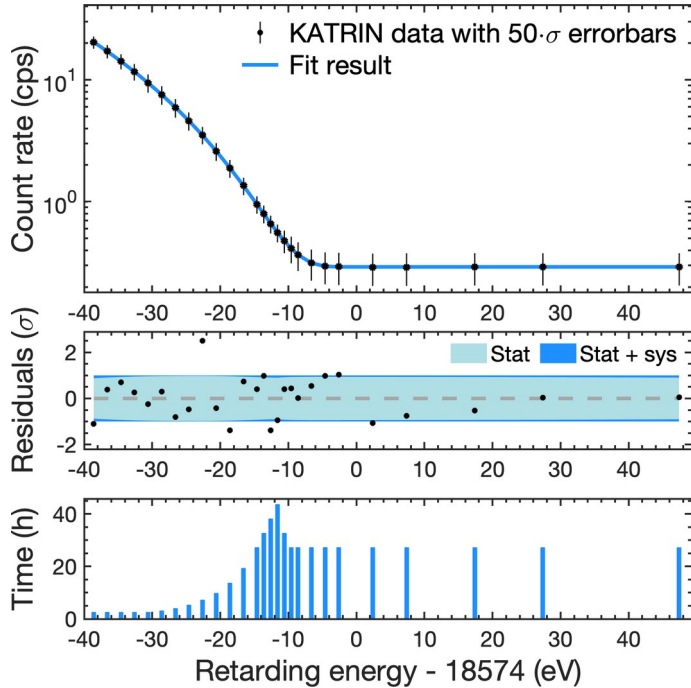
- Fermi theory
- Molecular final states
- Doppler broadening
- Radiative corrections
- ...

- Spectrometer resolution
- Scattering in the source
- Synchrotron radiation
- ...

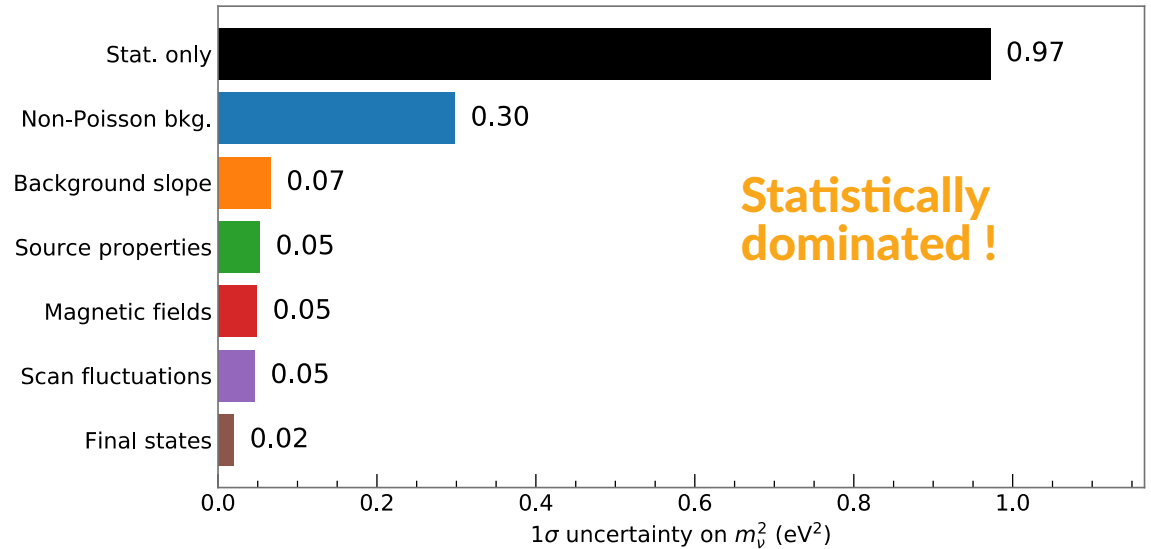
<https://arxiv.org/pdf/2101.05253.pdf>



# 1<sup>st</sup> neutrino mass campaign




PRL 123, 221802 (2019)



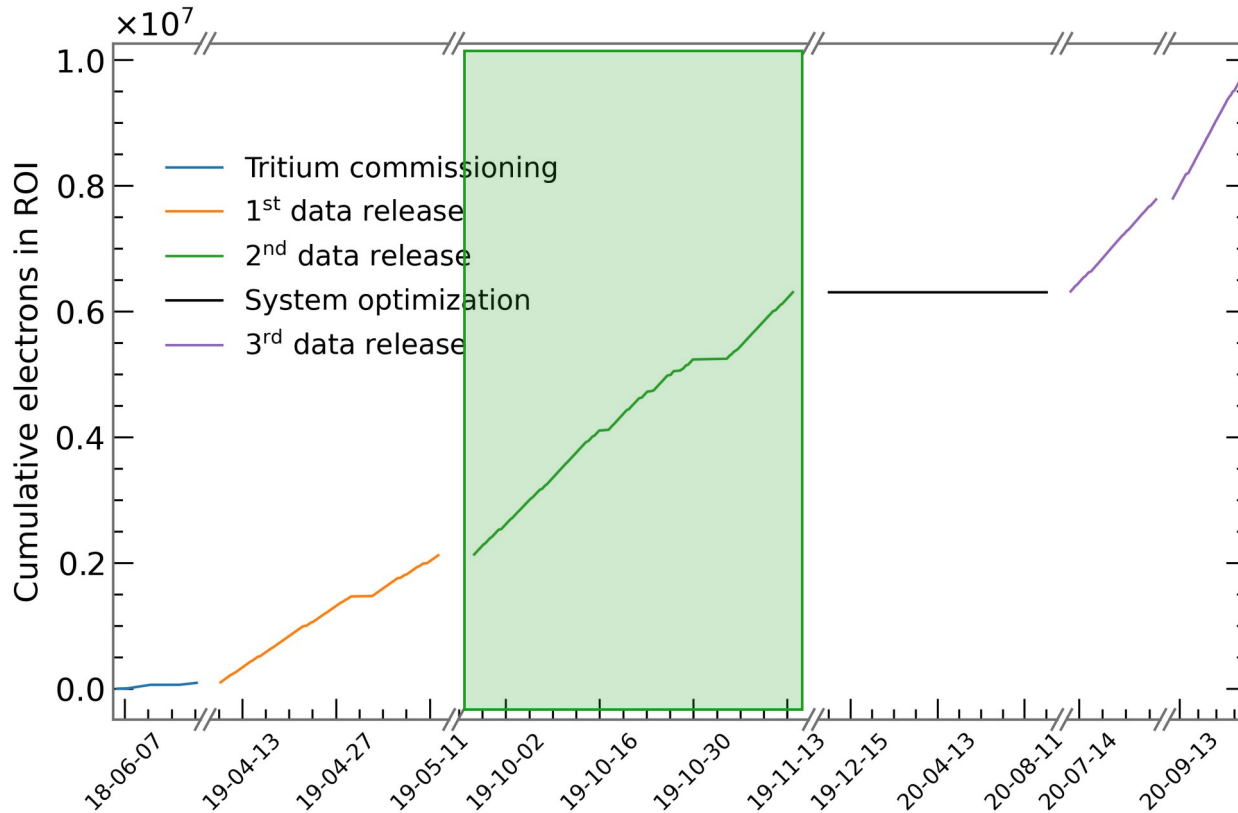
PRL 123, 221802 (2019)

**Neutrino Properties** PDG!

See the note on “Neutrino properties listings” in the Particle Mass  $m < 1.1$  eV, CL = 90% (tritium decay)

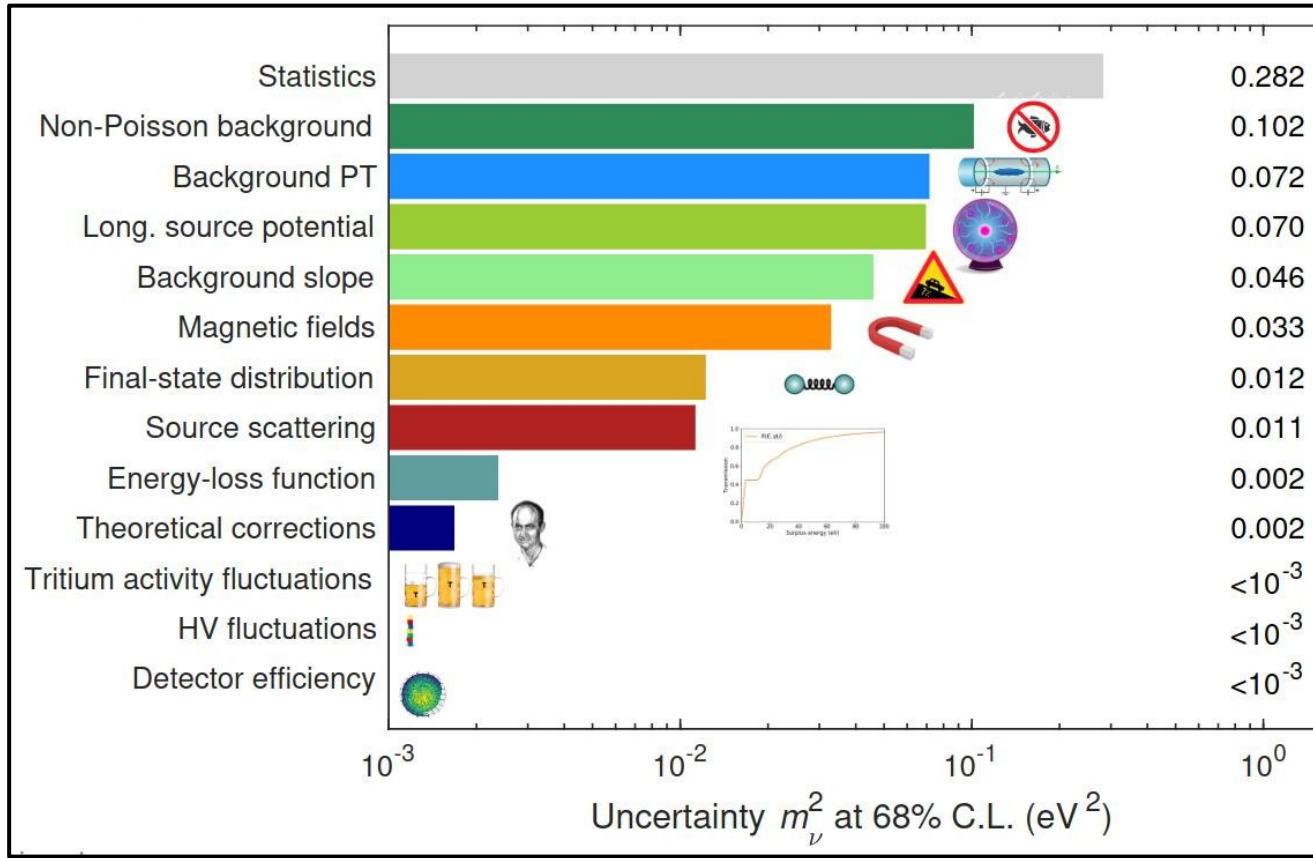
- Best fit value :  $m_\nu^2 = (-1.0 \pm 0.9) \text{ eV}^2$
- Limit :  $m_\nu < 1.1 \text{ eV}$  (90% CL)
-  :  $m_\nu < 0.9 \text{ eV}$  (90% CL)

# 2<sup>nd</sup> neutrino mass campaign



Gas density : 84 %  
Sensitivity :  $m_\nu < 0.7$  eV (90% CL)  
Electrons in ROI :  $4.2 \cdot 10^6$   
Scan time : 31 days

# 2<sup>nd</sup> neutrino mass campaign



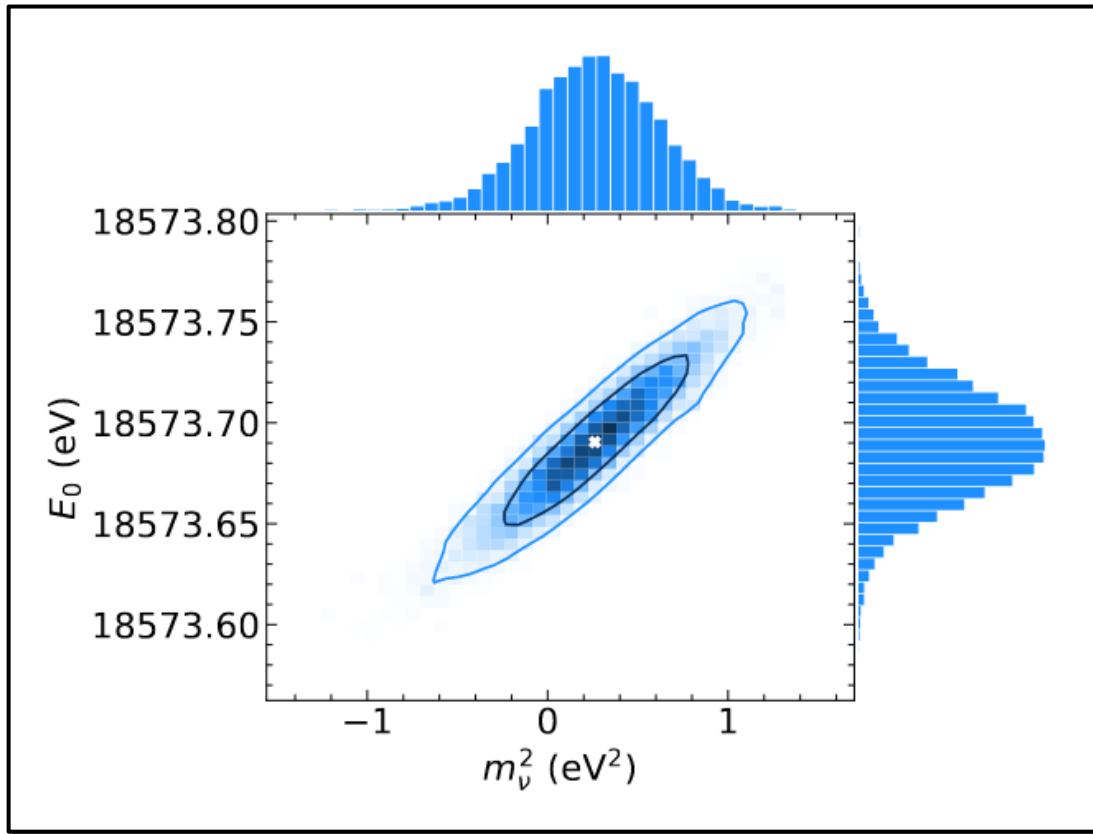
Gas density : 84 %  
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 Scan time : 31 days

$$\sigma_{\text{stat}} \sim 0.28 \text{ eV}^2$$

$$\sigma_{\text{syst}} \sim 0.15 \text{ eV}^2$$

$$\sigma_{\text{tot}} \sim 0.35 \text{ eV}^2$$

# 2<sup>nd</sup> neutrino mass campaign



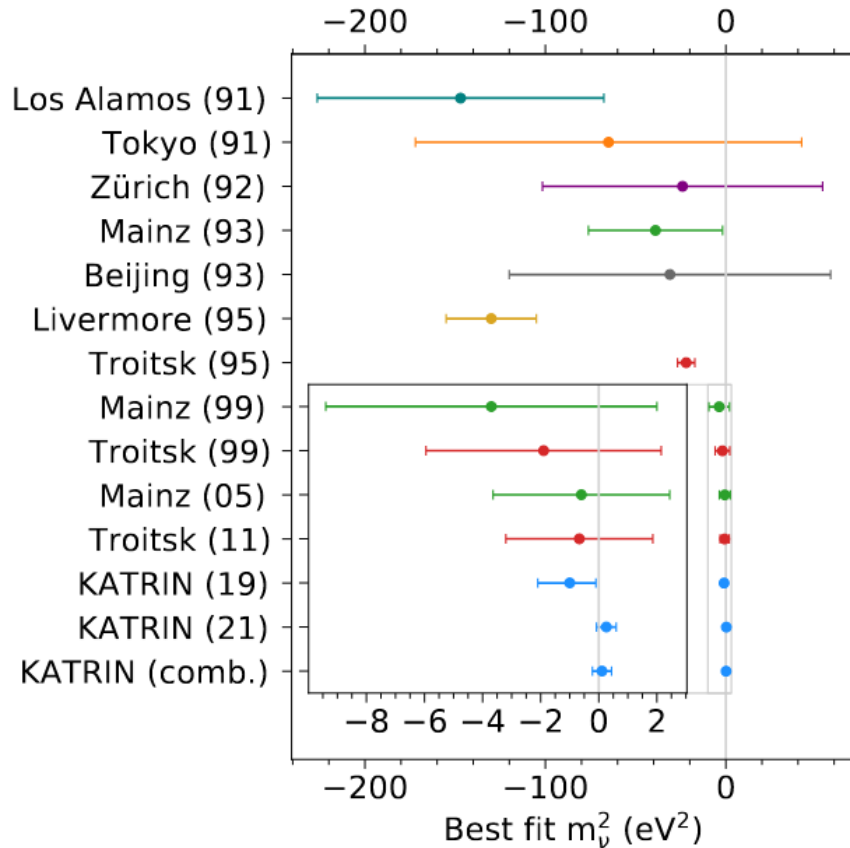
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$$\begin{aligned}\sigma_{\text{stat}} &\sim 0.28 \text{ eV}^2 \\ \sigma_{\text{syst}} &\sim 0.15 \text{ eV}^2 \\ \sigma_{\text{tot}} &\sim 0.35 \text{ eV}^2\end{aligned}$$

Results:

$$\begin{aligned}m_\nu^2 &= 0.26 \pm 0.35 \text{ eV}^2 \\ m_\nu &< 0.9 \text{ eV (90% CL)}\end{aligned}$$

# 2<sup>nd</sup> neutrino mass campaign



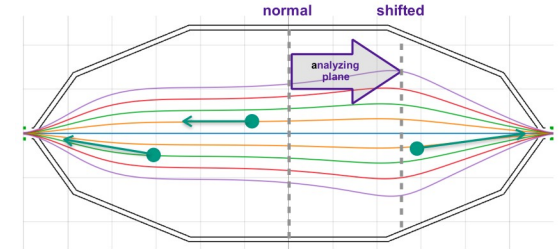
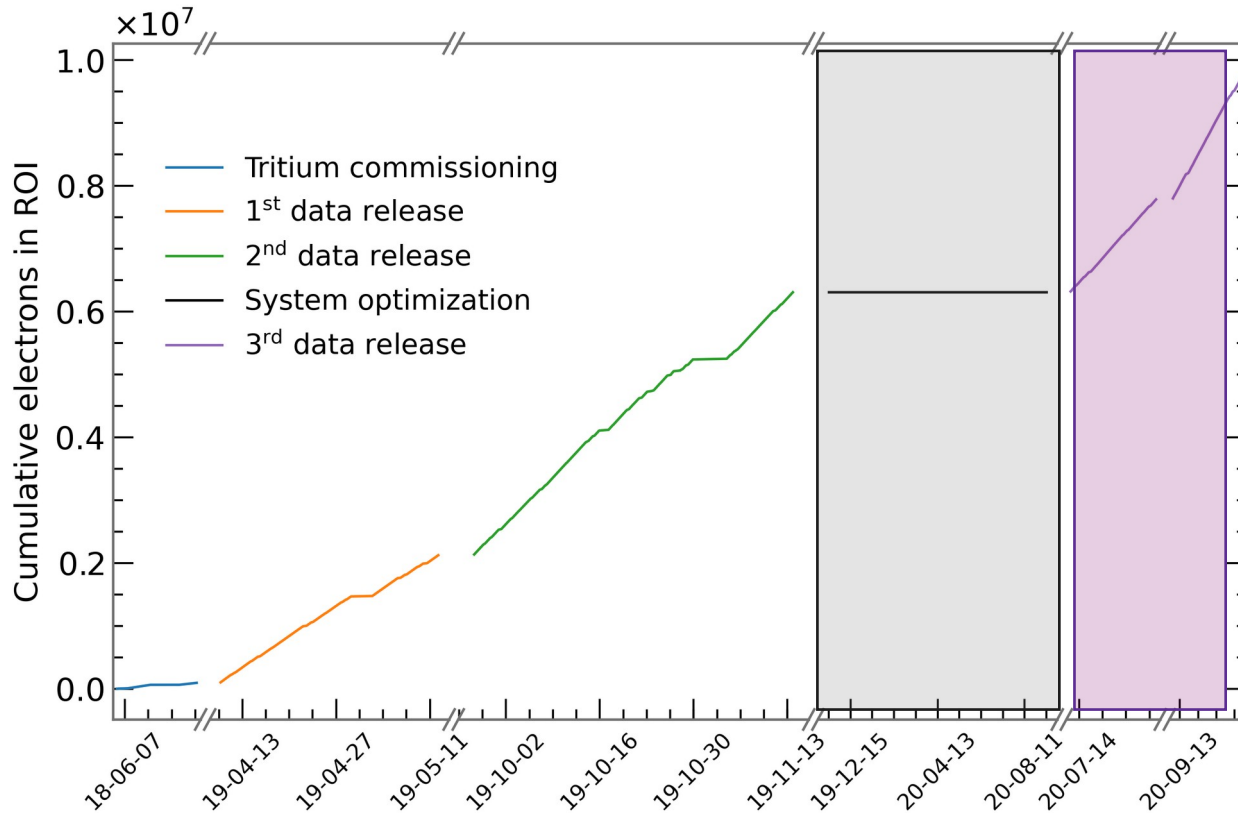
Combining the two campaigns :  
 $m_\nu < 0.8$  eV (90% CL)

First direct neutrino-mass measurement with sub-eV sensitivity

M. Aker,<sup>1</sup> A. Beglarian,<sup>2</sup> J. Behrens,<sup>3,4</sup> A. Berlev,<sup>5</sup> U. Besserer,<sup>1</sup> B. Bieringer,<sup>6</sup> F. Block,<sup>3,3</sup> B. Bornschein,<sup>1</sup> Bornschein,<sup>4</sup> M. Böttcher,<sup>6</sup> T. Brunst,<sup>7,8</sup> T. S. Caldwell,<sup>9,10</sup> R. M. D. Carney,<sup>11</sup> L. La Cascio,<sup>3</sup> S. Chilingaryan,<sup>2</sup> W. Choi,<sup>3</sup> K. Debowski,<sup>12</sup> M. Deffert,<sup>3</sup> M. Descher,<sup>3</sup> D. Díaz Barrero,<sup>13</sup> P. J. Doe,<sup>14</sup> O. Dragoun,<sup>15</sup> G. Drexlin,<sup>3</sup>

(submitted arxiv: [2105.08533.pdf](https://arxiv.org/abs/2105.08533))

# Next neutrino mass campaigns

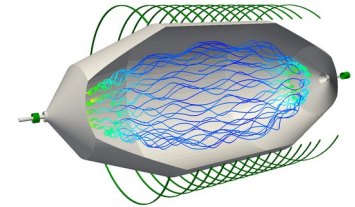


## Plasma investigations

M. Slezak et al. *J. Phys. G* 47 065002 (2020)

## Improve background configuration

F. Fränkle, A. Schaller et al., arxiv :2011.05107v1

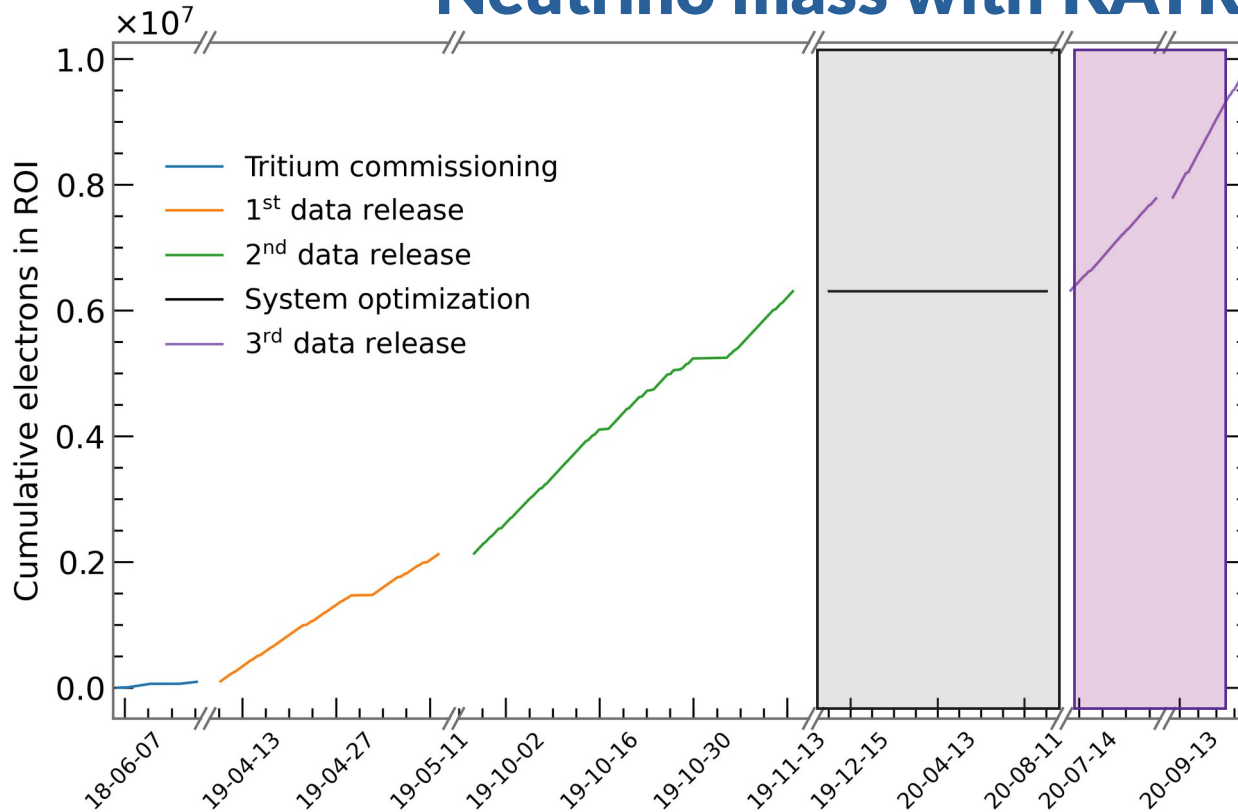


Data taking since Summer 2020 :

→ 1000 days  $\leftrightarrow$   $< 0.3$  eV (90 % CL)

# Conclusion

## Neutrino mass with KATRIN



- 1st tritium campaign :  
→ **very good stability**
- 1st mass campaign  
→ **best limit on the  $m_\nu$**  from direct measure ( $m_\nu < 1.1$  eV)
- 2nd mass campaign  
→ **sub-eV regime**



# How to generate mass from the SM ?

- In the SM, fermions masses = Yukawa coupling between RH-LH and Higgs field
- No RH neutrinos exist in SM → neutrino massless
- To generate mass : need of RH partners
- A lot of different models – all must extend SM

$$- \mathcal{L}_{M_\nu} = M_{Dij} \bar{\nu}_{si} \nu_{Lj} + \frac{1}{2} M_{Nij} \bar{\nu}_{si} \nu_{sj}^c + \text{h.c.}$$

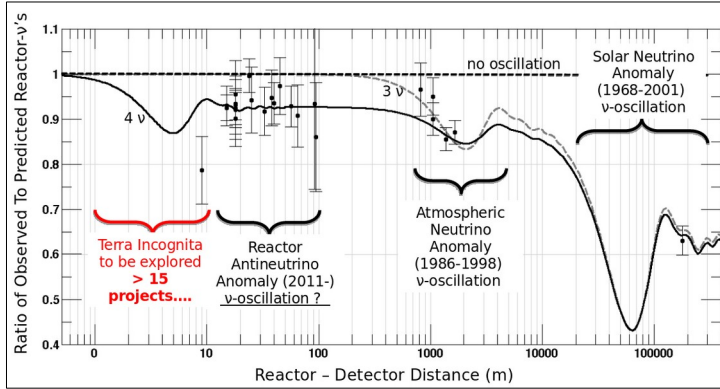
Dirac  
mass term

Majorana  
mass term

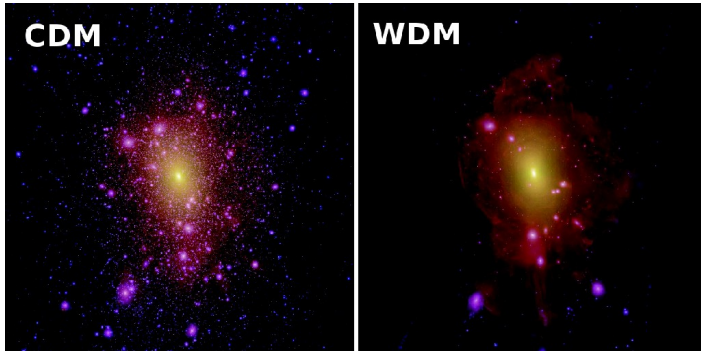
RH neutrinos → Sterile neutrinos

|   |  |  |
|---|--|--|
| 2/3<br>Left<br>Right<br>2.4 MeV<br><b>u</b><br>up                               | 2/3<br>Left<br>Right<br>1.27 GeV<br><b>c</b><br>charm                            | 2/3<br>Left<br>Right<br>171.2 GeV<br><b>t</b><br>top                             |
| -1/3<br>Left<br>Right<br>4.8 MeV<br><b>d</b><br>down                            | -1/3<br>Left<br>Right<br>104 MeV<br><b>s</b><br>strange                          | -1/3<br>Left<br>Right<br>4.2 GeV<br><b>b</b><br>bottom                           |
| < 1 eV<br>Left<br>Right<br><b>N<sub>1</sub></b><br>sterile<br>neutrino<br>~eV ? | < 1 eV<br>Left<br>Right<br><b>N<sub>2</sub></b><br>sterile<br>neutrino<br>~keV ? | < 1 eV<br>Left<br>Right<br><b>N<sub>3</sub></b><br>sterile<br>neutrino<br>~GeV ? |
| -1<br>Left<br>Right<br>0.511 MeV<br><b>e</b><br>electron                        | -1<br>Left<br>Right<br>105.7 MeV<br><b>μ</b><br>muon                             | -1<br>Left<br>Right<br>1.777 GeV<br><b>τ</b><br>tau                              |

# Is there a sterile neutrino ?



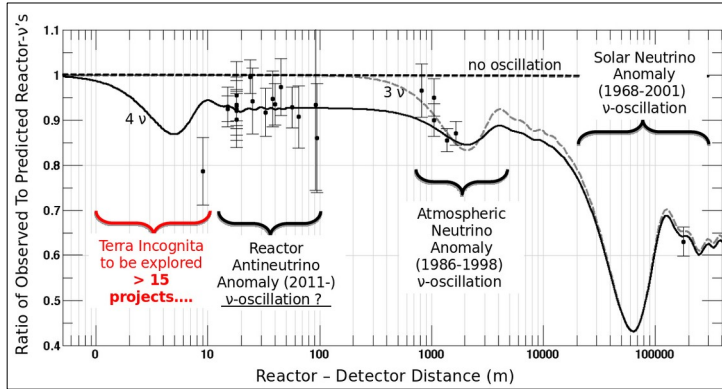
**eV-scale:**  
Resolve anomalies in  
oscillation experiments



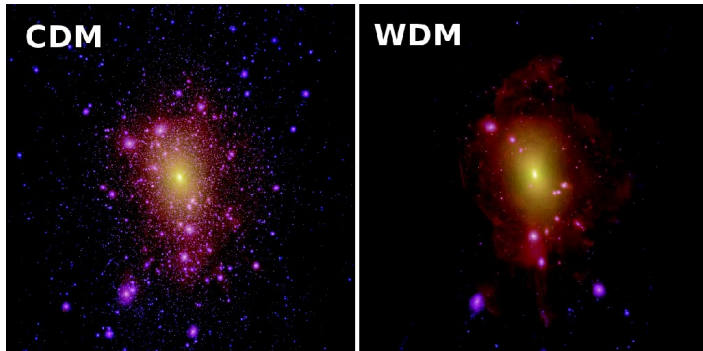
**keV-scale:**  
Dark Matter candidate

|                                 |  |                                  |   |                                 |  |
|---------------------------------|--|----------------------------------|---|---------------------------------|--|
| 2/3<br>Left<br>up<br>Right      | 2.4 MeV<br><b>u</b>                                | 2/3<br>Left<br>charm<br>Right    | 1.27 GeV<br><b>c</b>                              | 2/3<br>Left<br>top<br>Right     | 171.2 GeV<br><b>t</b>                              |
| -1/3<br>Left<br>down<br>Right   | 4.8 MeV<br><b>d</b>                                | -1/3<br>Left<br>strange<br>Right | 104 MeV<br><b>s</b>                               | -1/3<br>Left<br>bottom<br>Right | 4.2 GeV<br><b>b</b>                                |
| 0<br>Left<br>$\nu_e$            | < 1 eV<br><b>N<sub>1</sub></b><br>sterile neutrino | 0<br>Left<br>$\nu_\mu$           | ~eV ?<br><b>N<sub>2</sub></b><br>sterile neutrino | 0<br>Left<br>$\nu_\tau$         | < 1 eV<br><b>N<sub>3</sub></b><br>sterile neutrino |
| -1<br>Left<br>electron<br>Right | 0.511 MeV<br><b>e</b>                              | -1<br>Left<br>muon<br>Right      | 105.7 MeV<br><b>μ</b>                             | -1<br>Left<br>tau<br>Right      | 1.777 GeV<br><b>τ</b>                              |

# Is there a sterile neutrino ?



**eV-scale:**  
Resolve anomalies in  
oscillation experiments

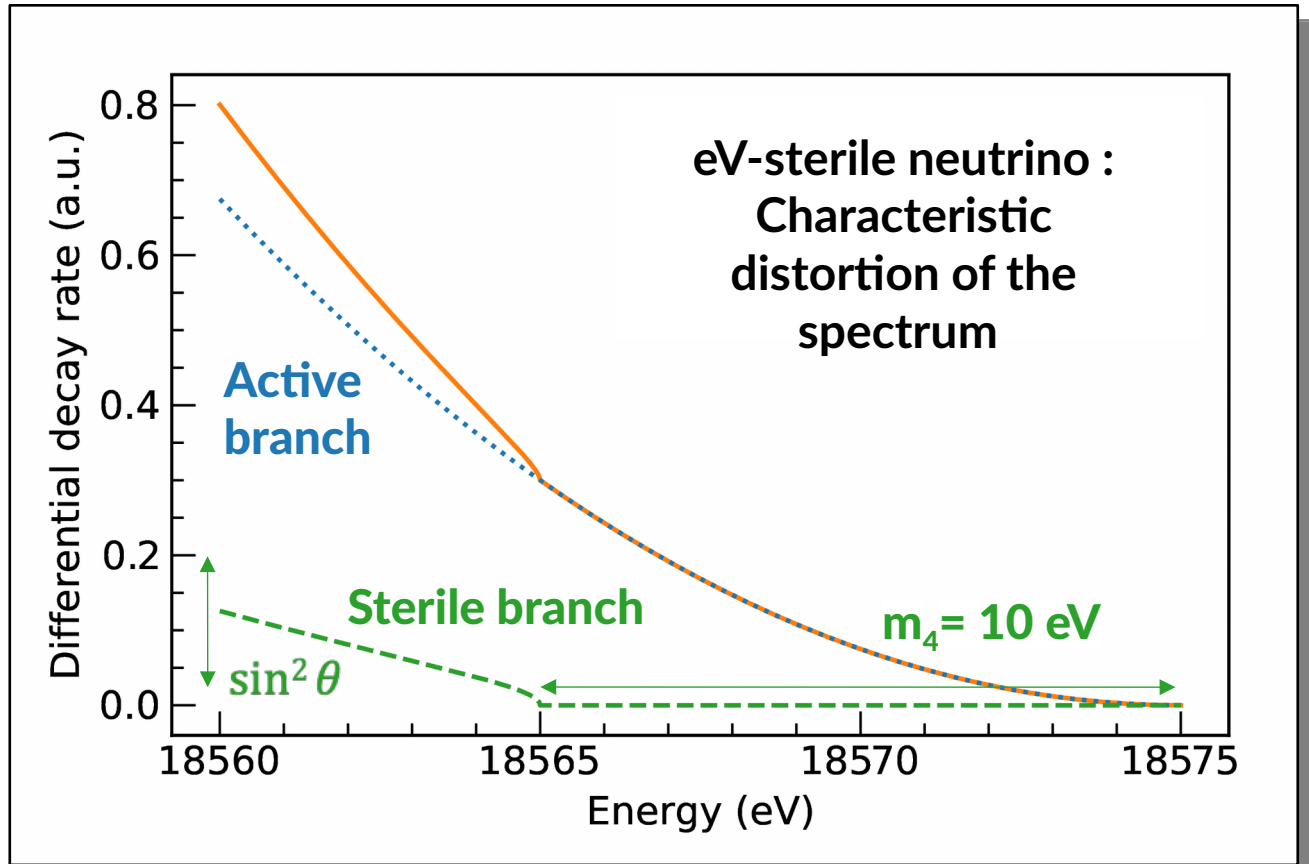
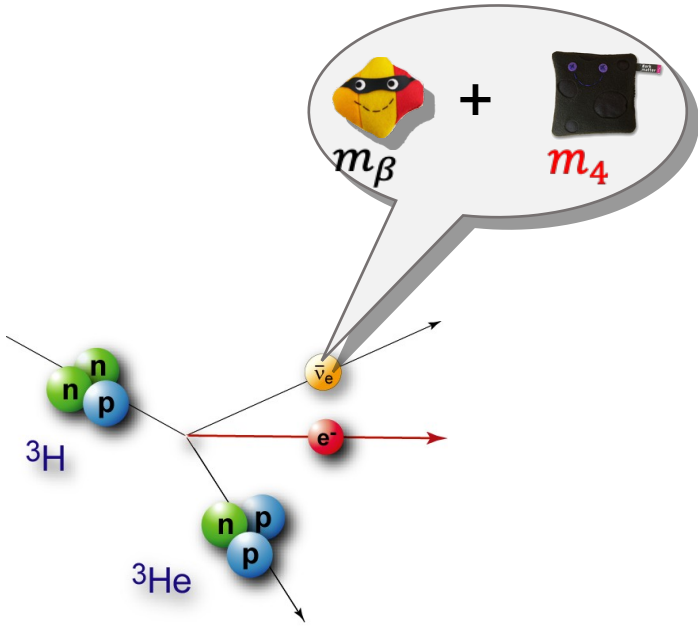


**keV-scale:**  
Dark Matter candidate

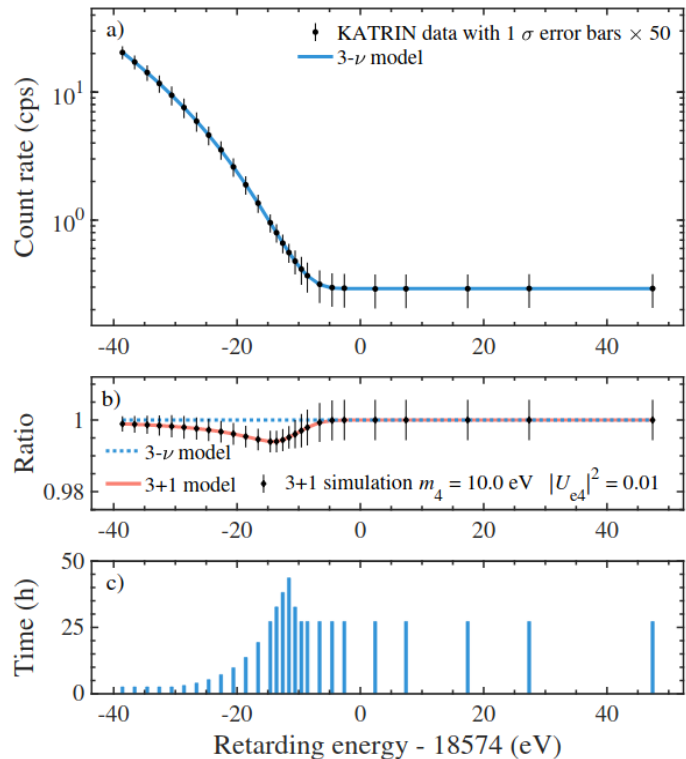


|                                 |  |                                  |   |                                 |  |
|---------------------------------|--|----------------------------------|---|---------------------------------|--|
| 2/3<br>Left<br>up<br>Right      | 2.4 MeV<br><b>u</b>                                | 2/3<br>Left<br>charm<br>Right    | 1.27 GeV<br><b>c</b>  | 2/3<br>Left<br>top<br>Right     | 171.2 GeV<br><b>t</b>  |
| -1/3<br>Left<br>down<br>Right   | 4.8 MeV<br><b>d</b>                                | -1/3<br>Left<br>strange<br>Right | 104 MeV<br><b>s</b>   | -1/3<br>Left<br>bottom<br>Right | 4.2 GeV<br><b>b</b>  |
| 0<br>Left<br>$\nu_e$            | < 1 eV<br><b>N<sub>1</sub></b><br>sterile neutrino | 0<br>Left<br>$\nu_\mu$           | < 1 eV<br>~eV ?<br><b>N<sub>2</sub></b><br>sterile neutrino | 0<br>Left<br>$\nu_\tau$         | < 1 eV<br>~GeV ?<br><b>N<sub>3</sub></b><br>sterile neutrino |
| -1<br>Left<br>electron<br>Right | 0.511 MeV<br><b>e</b>                              | -1<br>Left<br>muon<br>Right      | 105.7 MeV<br><b>μ</b>                                       | -1<br>Left<br>tau<br>Right      | 1.777 GeV<br><b>τ</b>  |

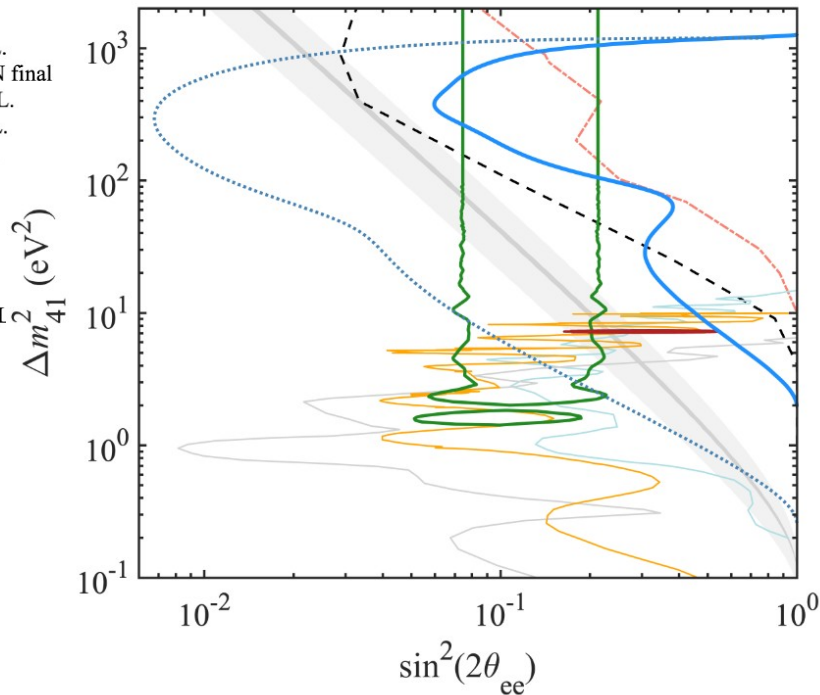
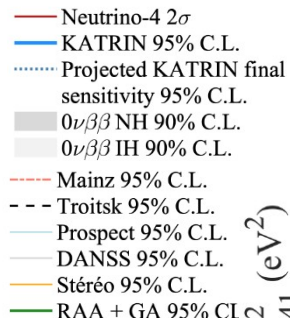
# eV-sterile signature in $\beta$ -decay



# Sterile hunt with KATRIN



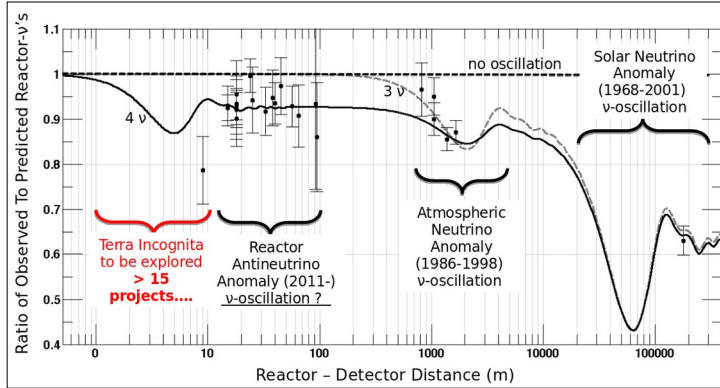
PRL 126, 091803



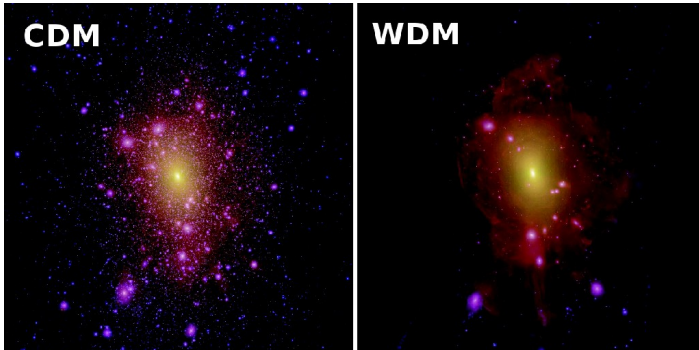
PRL 126, 091803

- **Unique large window at high mass**
- **Complementary with Reactor experiments**
- **Exclude most of the favored phase-space in the next years**

# Is there a sterile neutrino ?



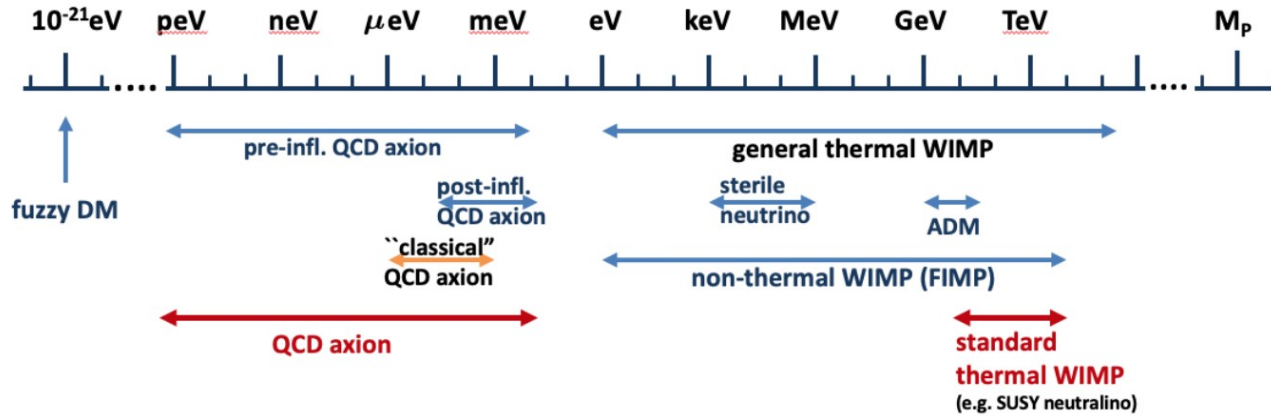
**eV-scale:**  
Resolve anomalies in oscillation experiments



**keV-scale:**  
Dark Matter candidate

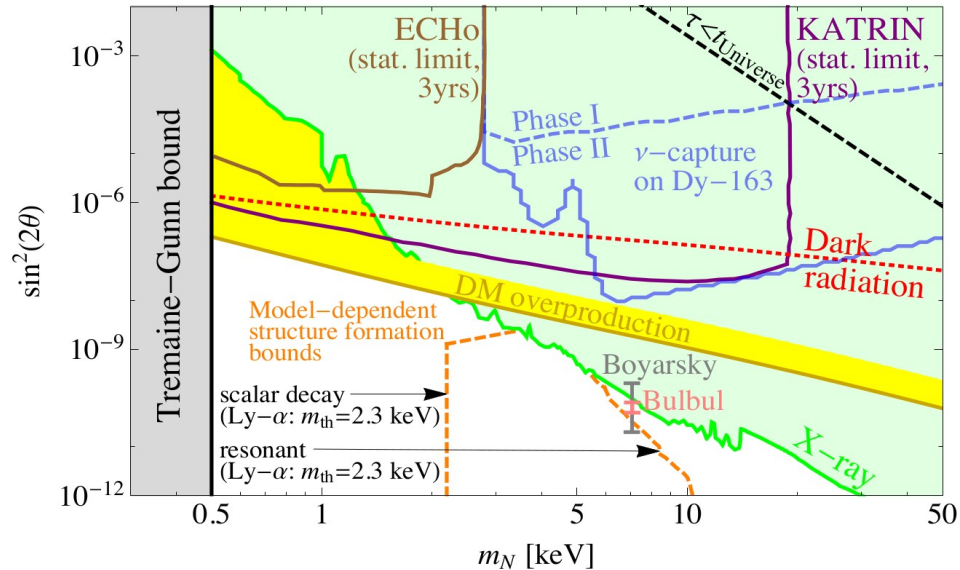
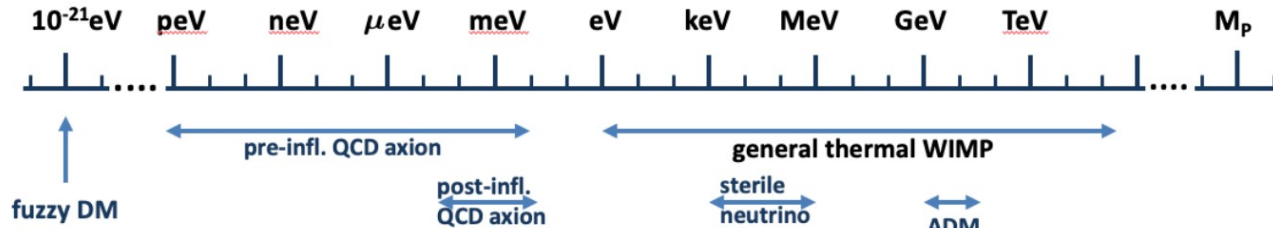
|                                     |  |                                       |   |  |   |
|-------------------------------------|--|---------------------------------------|---|--|---|
| 2/3<br>Left<br>up<br>Right          | 2.4 MeV<br><b>u</b>  | 2/3<br>Left<br>charm<br>Right         | 1.27 GeV<br><b>c</b>  | 2/3<br>Left<br>top<br>Right            | 171.2 GeV<br><b>t</b>   |
| -1/3<br>Left<br>down<br>Right       | 4.8 MeV<br><b>d</b>  | -1/3<br>Left<br>strange<br>Right      | 104 MeV<br><b>s</b>   | -1/3<br>Left<br>bottom<br>Right        | 4.2 GeV<br><b>b</b>   |
| $< 1 \text{ eV}$<br>Left<br>$\nu_e$ | $\sim \text{eV} ?$<br><b>N<sub>1</sub></b><br>sterile neutrino | $< 1 \text{ eV}$<br>Left<br>$\nu_\mu$ | $\sim \text{keV} ?$<br><b>N<sub>2</sub></b><br>sterile neutrino | $< 1 \text{ eV}$<br>Left<br>$\nu_\tau$ | $\sim \text{GeV} ?$<br><b>N<sub>3</sub></b><br>sterile neutrino |
| -1<br>Left<br>electron<br>Right     | 0.511 MeV<br><b>e</b>  | -1<br>Left<br>muon<br>Right           | 105.7 MeV<br><b>μ</b>   | -1<br>Left<br>tau<br>Right             | 1.777 GeV<br><b>τ</b>   |

# keV-sterile neutrinos as Dark Matter

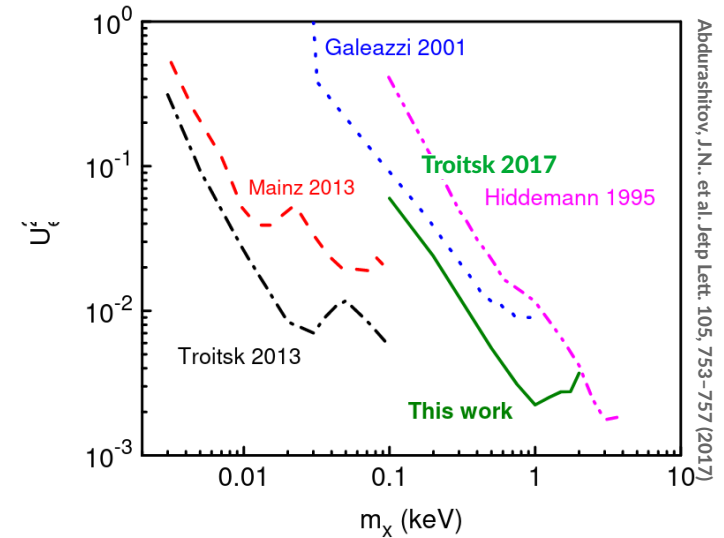
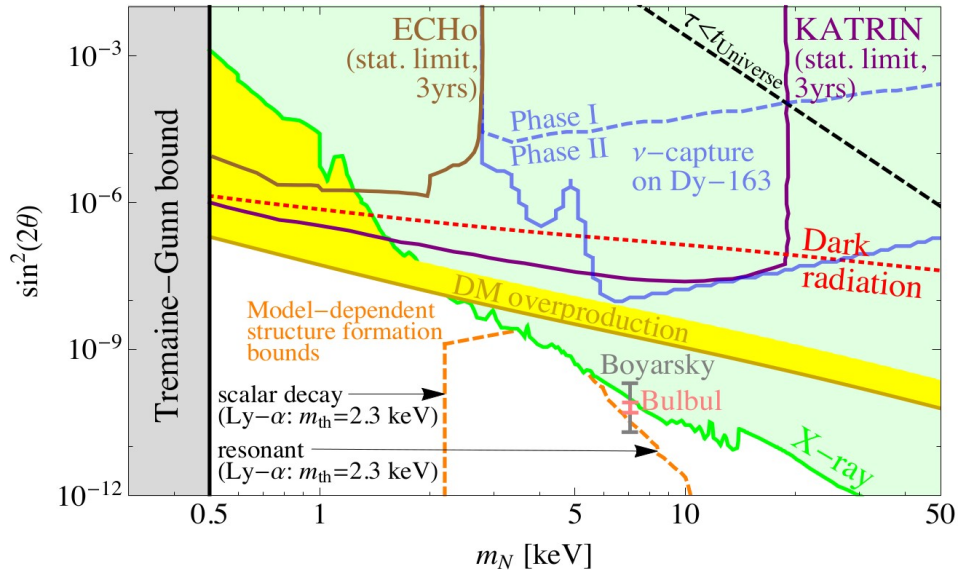
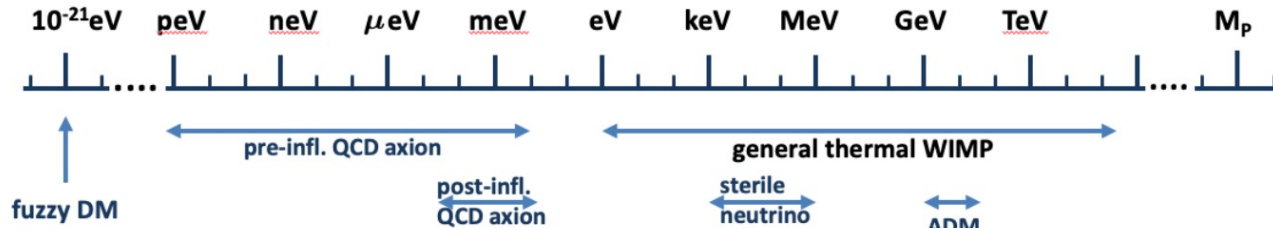




# keV-sterile neutrinos as Dark Matter

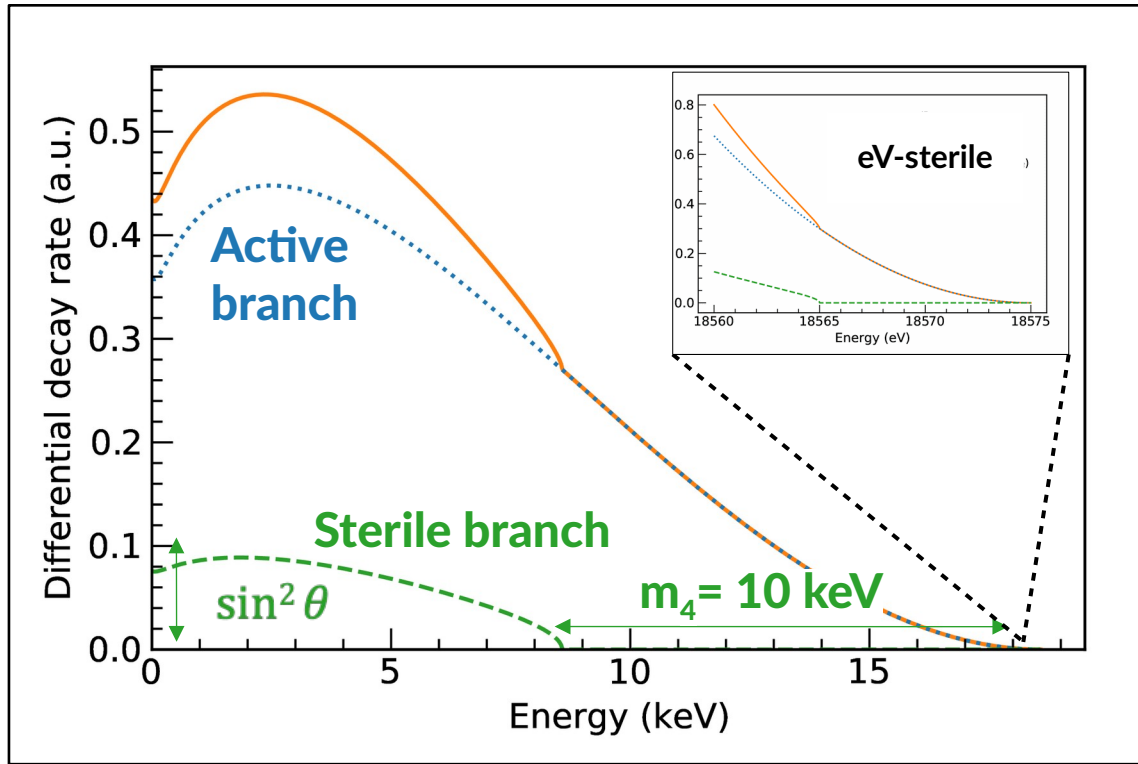


# keV-sterile neutrinos as Dark Matter

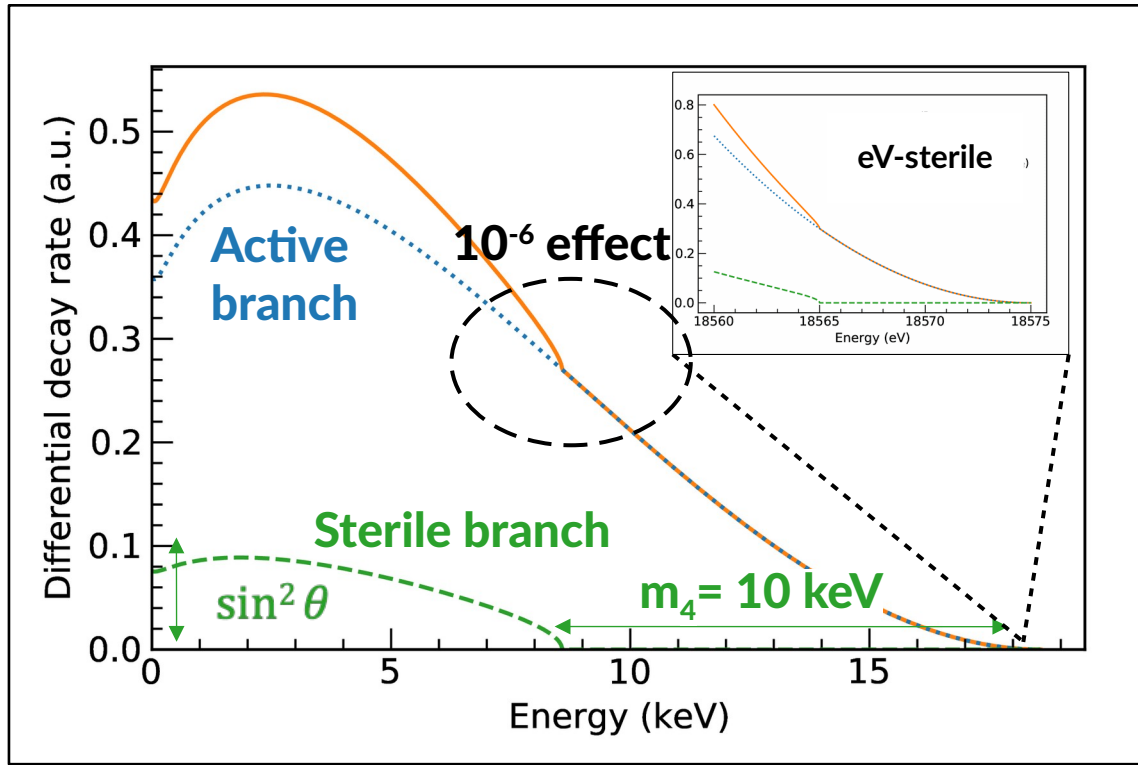


Abdurashitov, J.N., et al. JETP Lett. 105, 753–757 (2017)

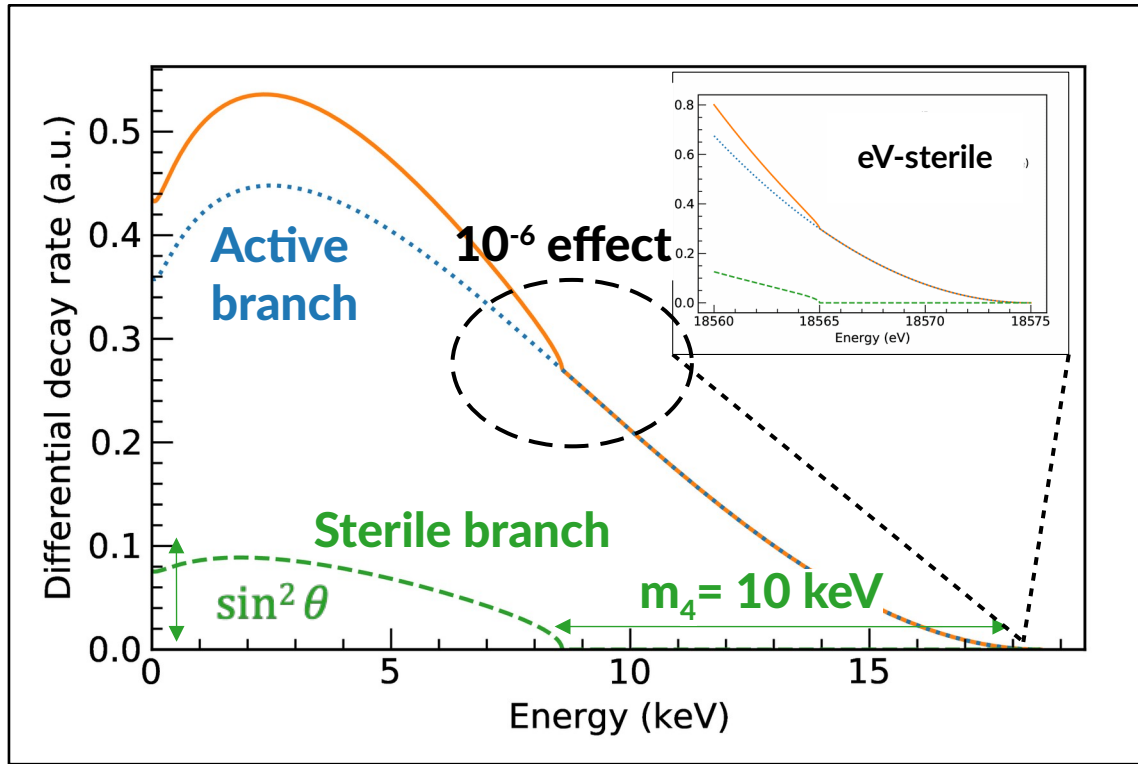
# keV-sterile signature in $\beta$ -decay



# keV-sterile signature in $\beta$ -decay



# keV-sterile signature in $\beta$ -decay



Stringent limit from **astrophysical** and **cosmological** observations ( $\sin^2(\theta) < 10^{-7}$ ):

→ Dramatic **increase of the count rate** (up to  $3 \times 10^8$  Hz)

→ Integral and differential phases (detector with **good resolution**)

→ Highly **pixelised**

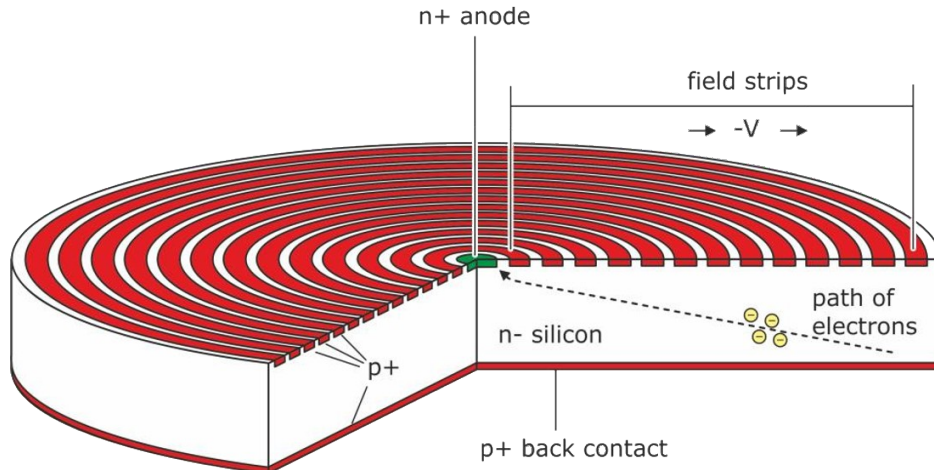
→ **new detector is needed : the TRISTAN project**



# Silicon Drift Detectors

Capability of handling high rates ( $> 3 \times 10^8$  cps)  
+ Excellent energy resolution (300 eV @ 20 keV)

- Silicon Drift Detector (SDD) Technology
- Novelty: large number of pixels (about 3500)
- Novelty: application to high-precision  $\beta$ -spectroscopy

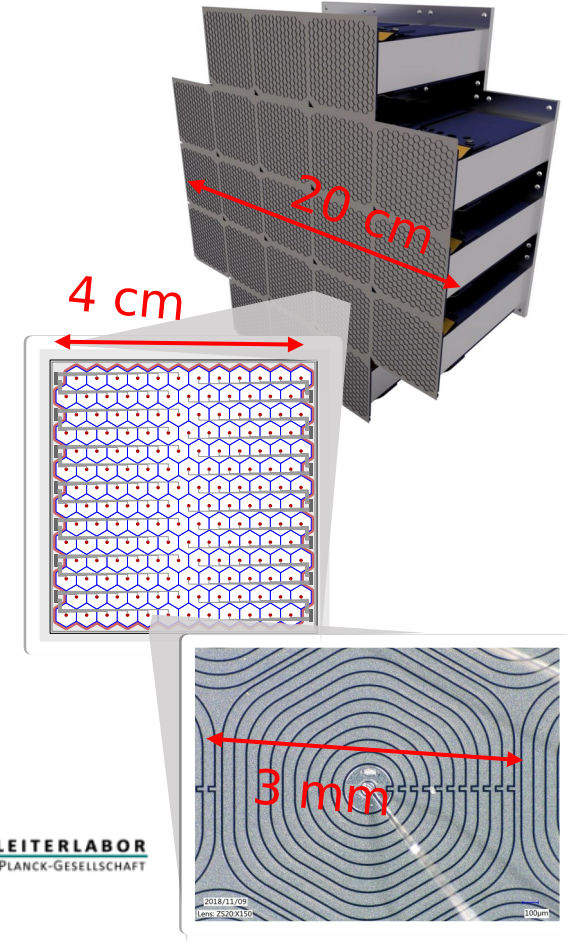
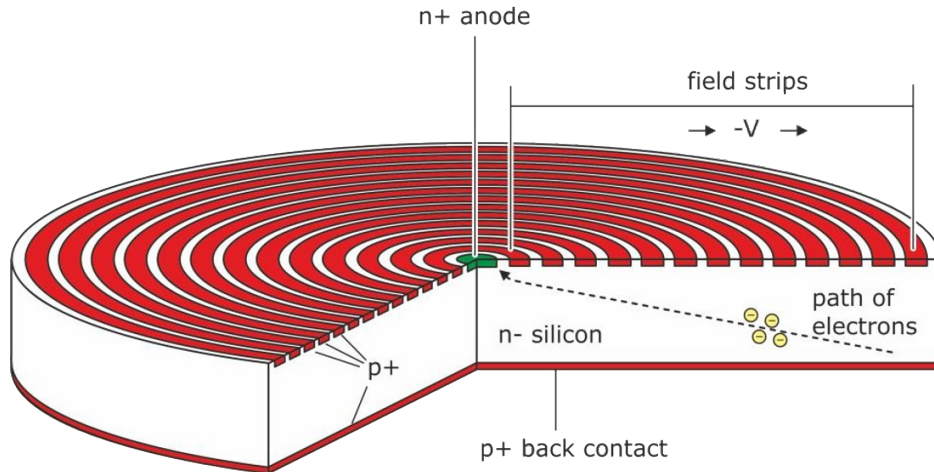




# TRISTAN Project

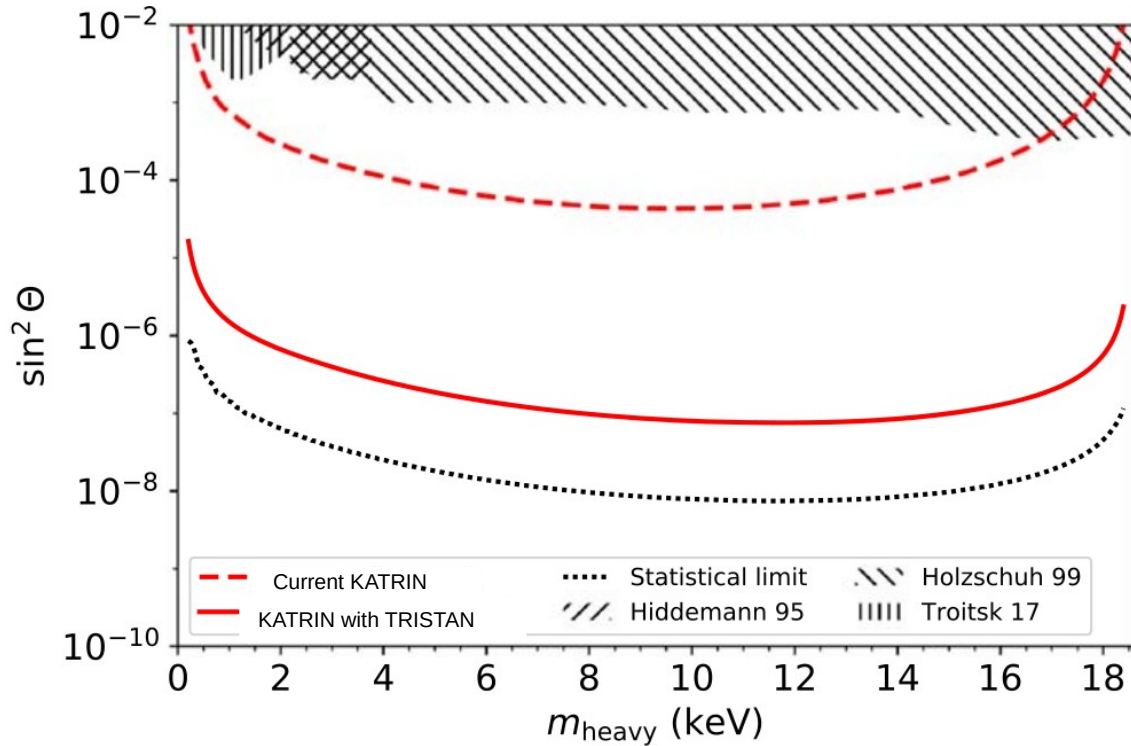
Capability of handling high rates ( $> 3 \times 10^8$  cps)  
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- Silicon Drift Detector (SDD) Technology
- Novelty: large number of pixels (about 3500)
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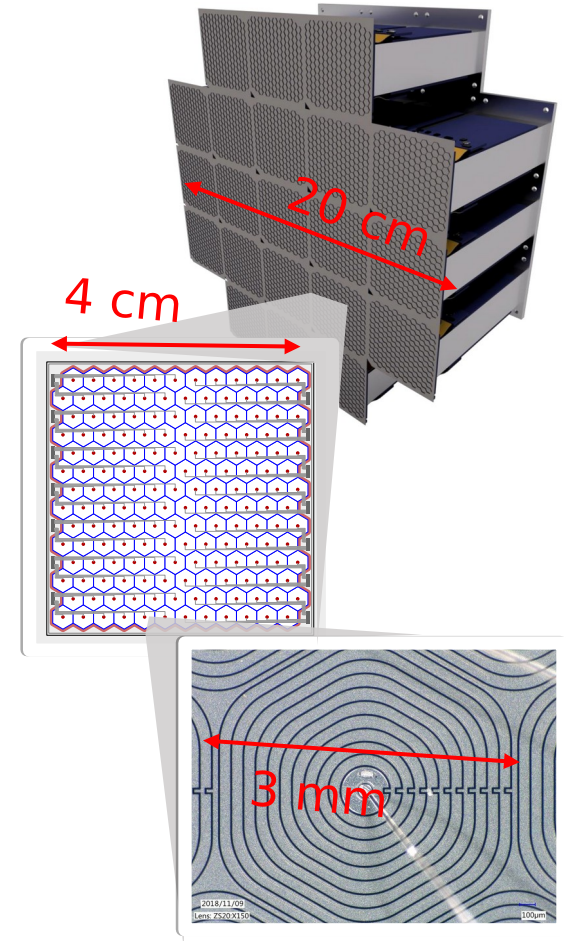


# TRISTAN Project

TRISTAN : Development of a large area SDD array and read-out system to look for keV-sterile neutrino with the KATRIN experiment



J.Phys. G46 (2019) no.6, 065203

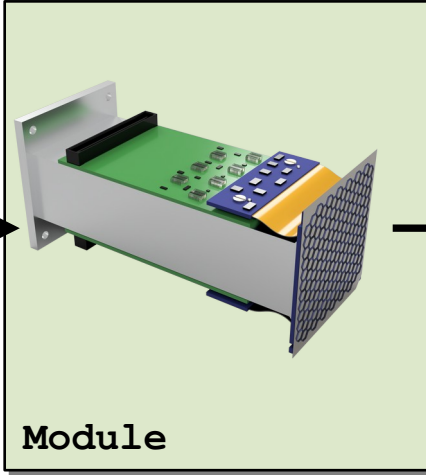


# Staged approach



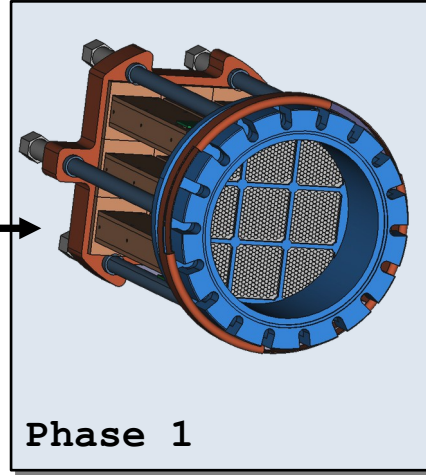
Prototypes

x24



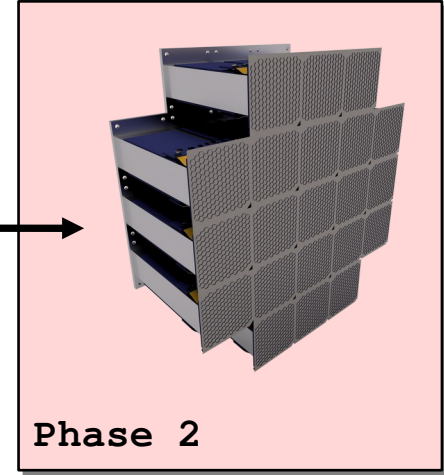
Module

x9



Phase 1

x2.5



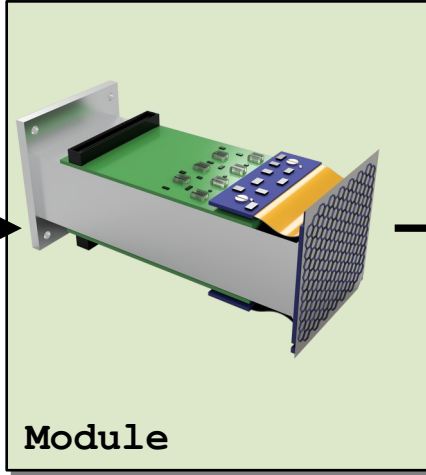
Phase 2

# Staged approach



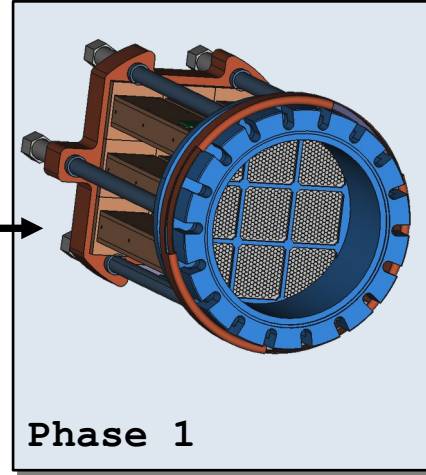
Prototypes

x24



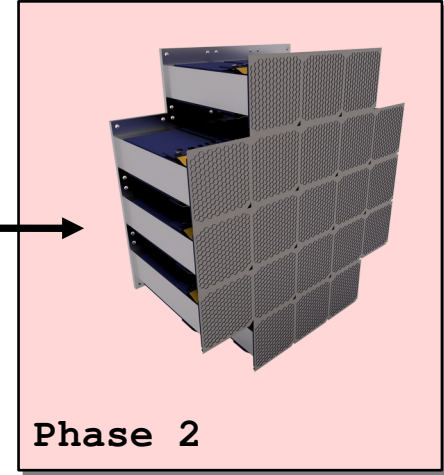
Module

x9



Phase 1

x2.5



Phase 2

## TRISTAN prototype

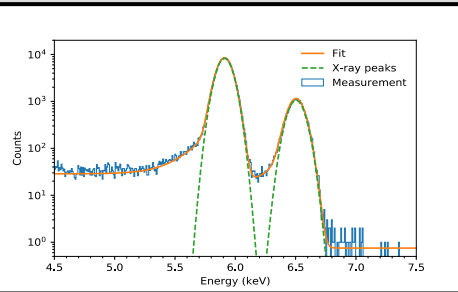
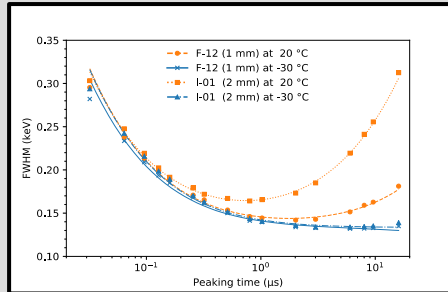
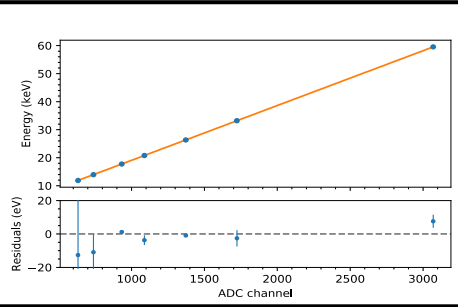
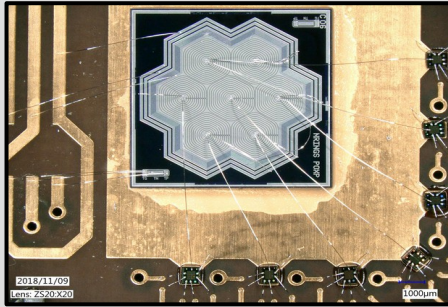
- 7-pixels with external CMOS



# Prototype results

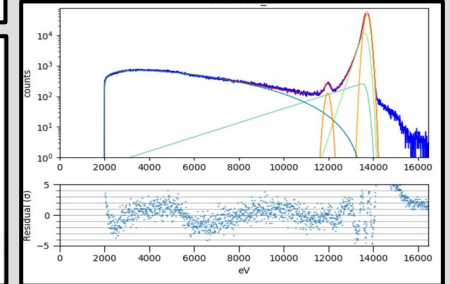
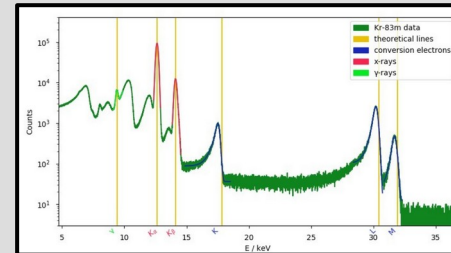
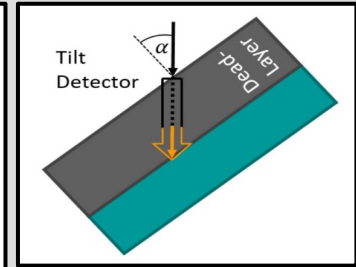
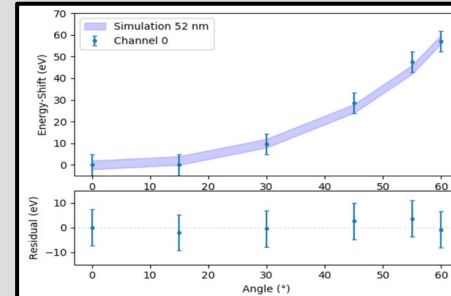
## Photon response

- 130 eV (FWHM) at 6 keV
- <150 eV (FWHM) for <1  $\mu$ s shaping time
- 0.1 % linearity over 60 keV range



## Electron response

- Semi empirical model in construction
- Dead-layer measurements using e-gun (tilting the detector) and  $^{241}\text{Am}$  sources

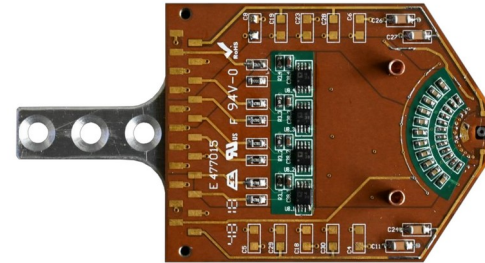
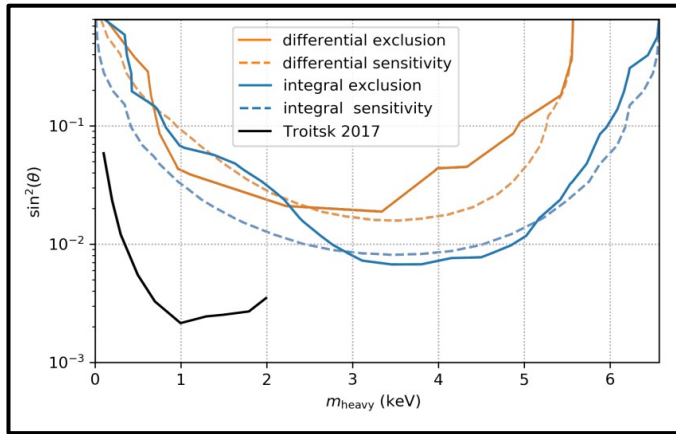
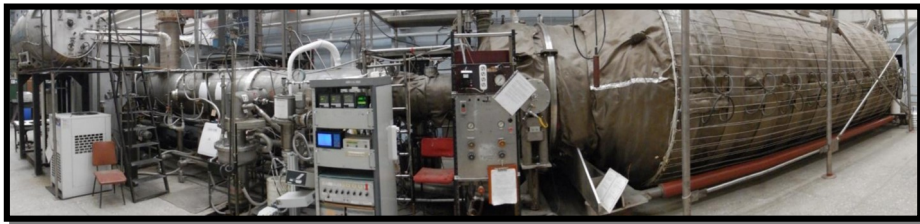


J. Phys. G:46 46 065203 (2019)

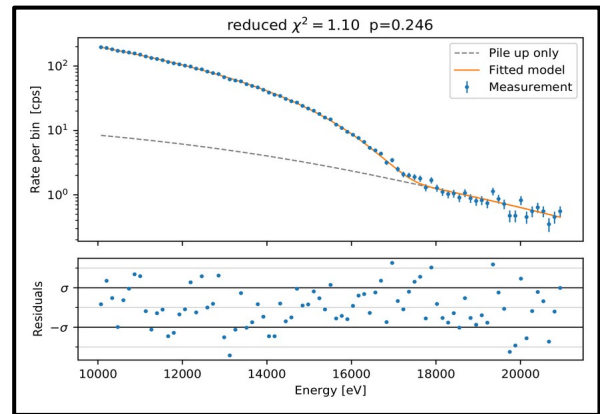
J. Phys. G: 48 015008 (2021)

# Prototype applications

**2017-2019 :**  
TRISTAN 7-pixels prototype implemented in  
**Troitsk** nu mass spectrometer → differential  
and integral measurements



**2019-2020 :**  
TRISTAN 7-pixels prototype implemented in  
**KATRIN** as **Forward Beam Monitor**.  
Monitoring since KNM2



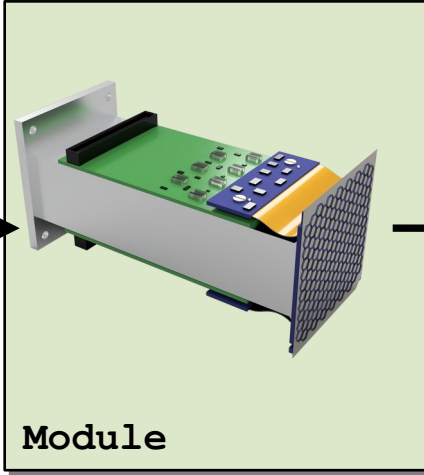


# Staged approach



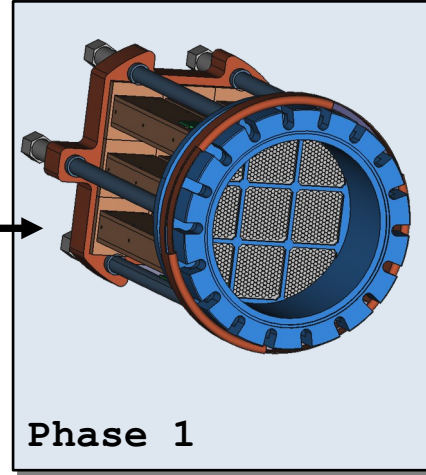
Prototypes

x24



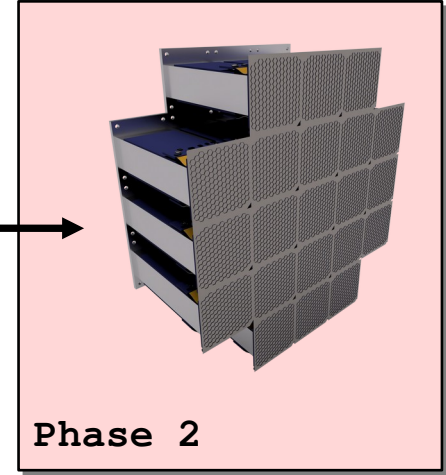
Module

x9



Phase 1

x2.5

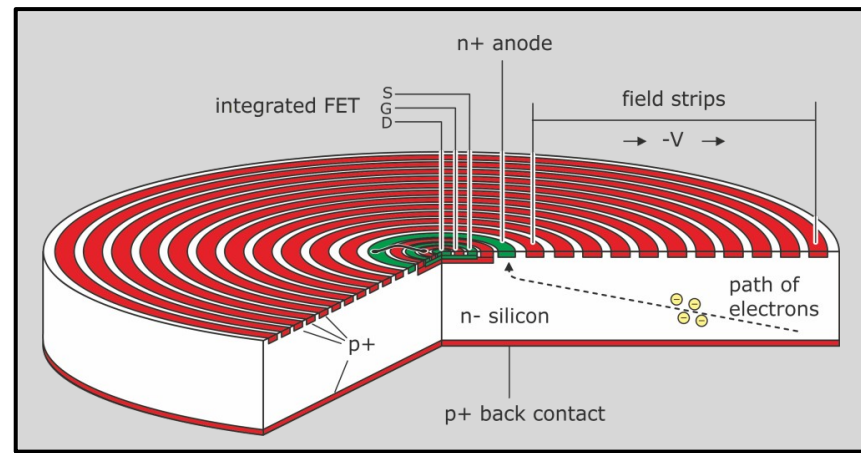
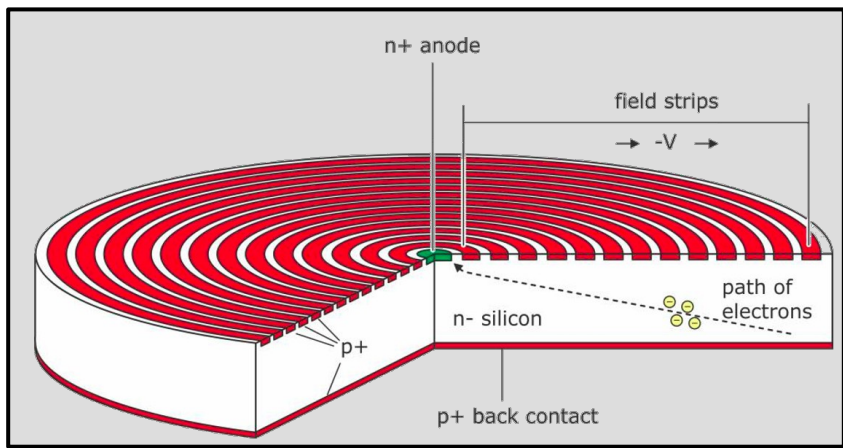
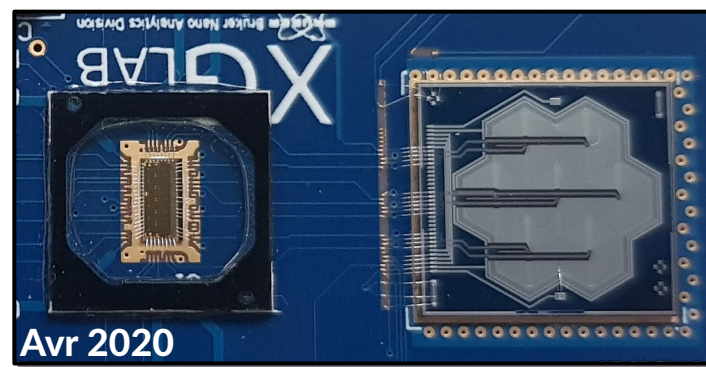
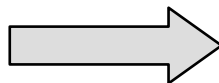
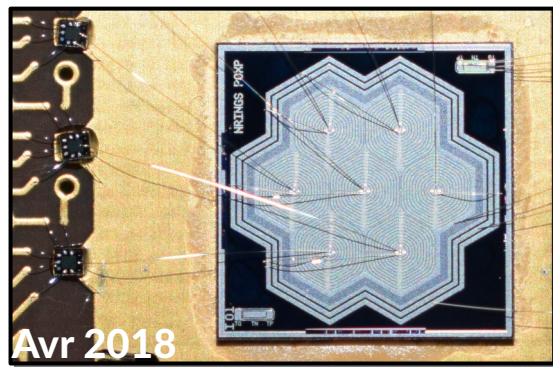


Phase 2

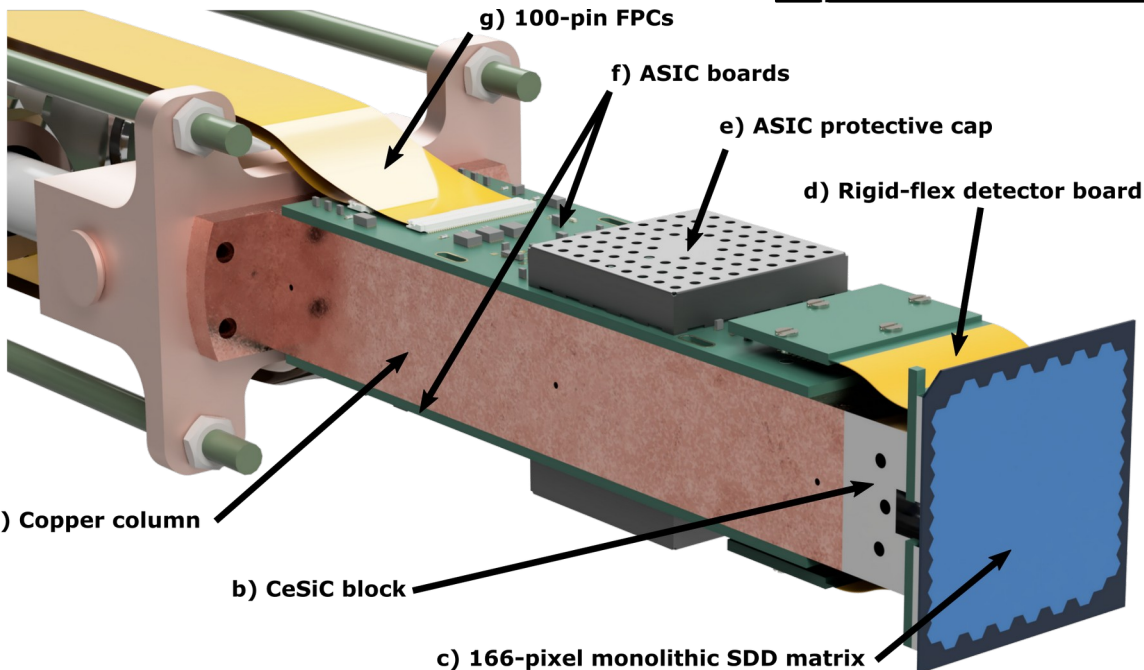
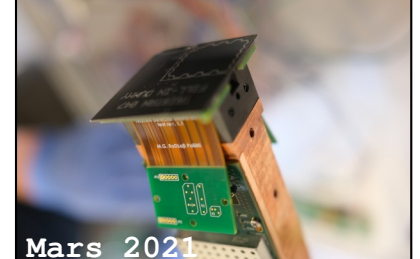
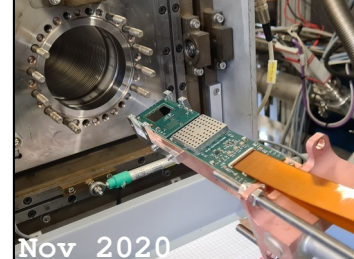
## TRISTAN module

- 166-pixels with integrated JFET

# Optimized design



# Design of the TRISTAN Module



Operating 166-pix SDD in UHV ( $10^{-8}$  mbar), cool ( $-80^{\circ}\text{C}$ ), intense B field ( $\sim 1$  T), low noise ( $< 300$  eV FWHM)

- Large SDD matrix with integrated FET
- CeSiC : carbon-fiber reinforced silicon carbide
- Rigid-flex with high density
- Dedicated ASIC
- 1-m long Kapton flex cable



# TRISTAN Module

1st operation of a 47-pix TRISTAN SDD on Oct 2020

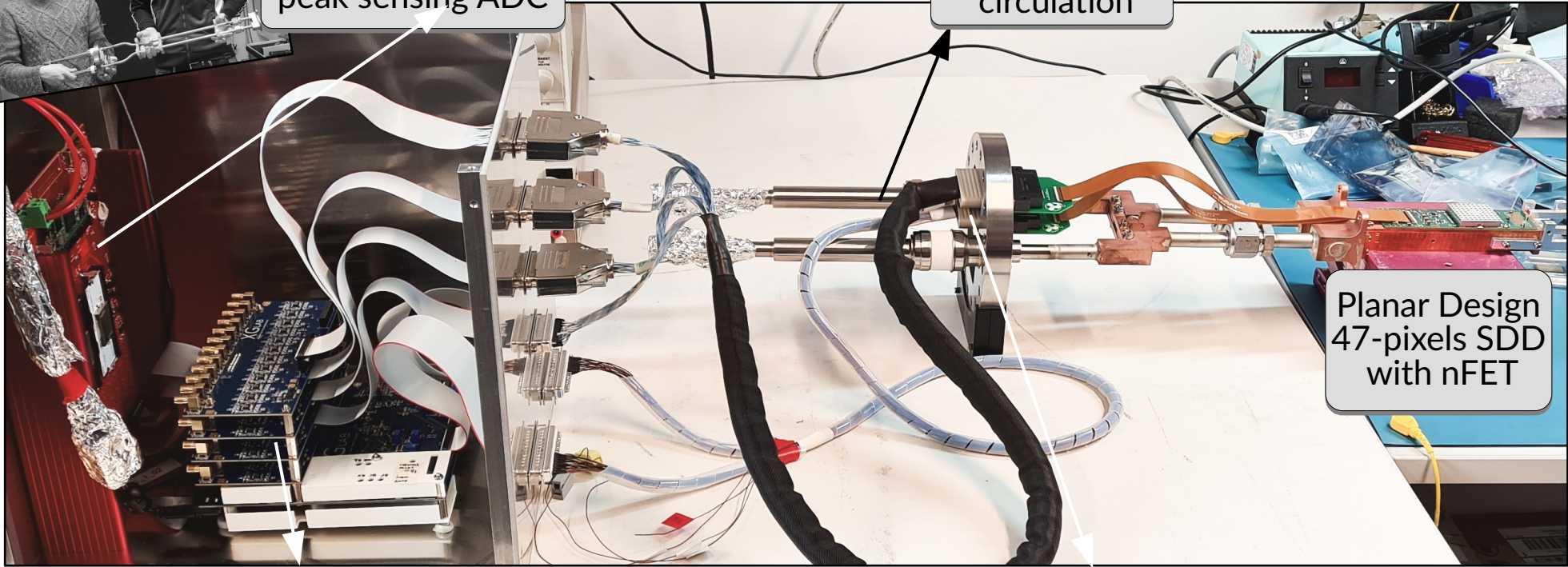
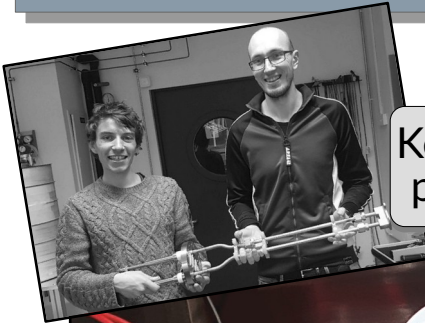
Kerberos, TRISTAN  
peak sensing ADC

Cooling liquid  
circulation

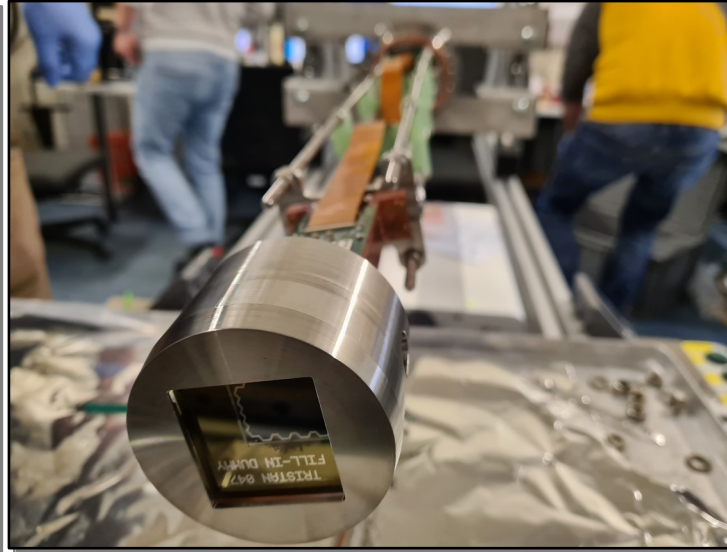
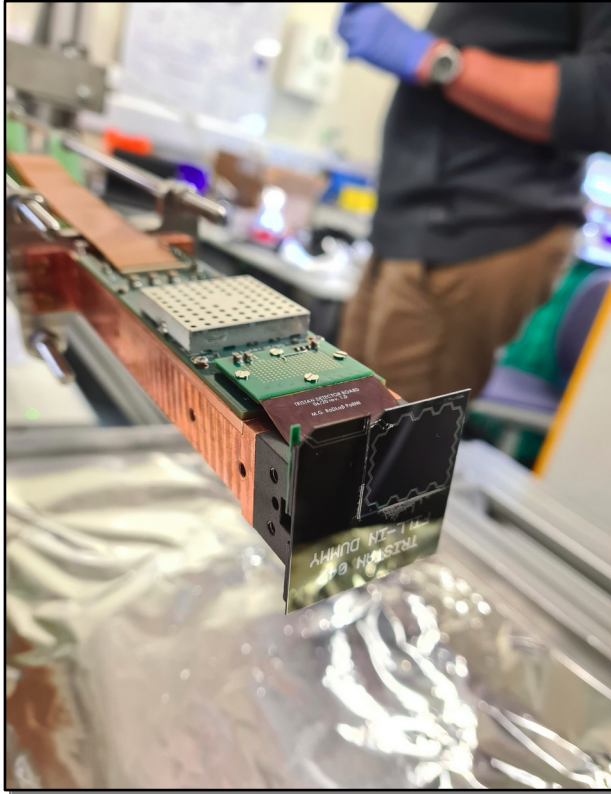
Planar Design  
47-pixels SDD  
with nFET

48-ch bias board

400 pins micro-D  
CF100 flange

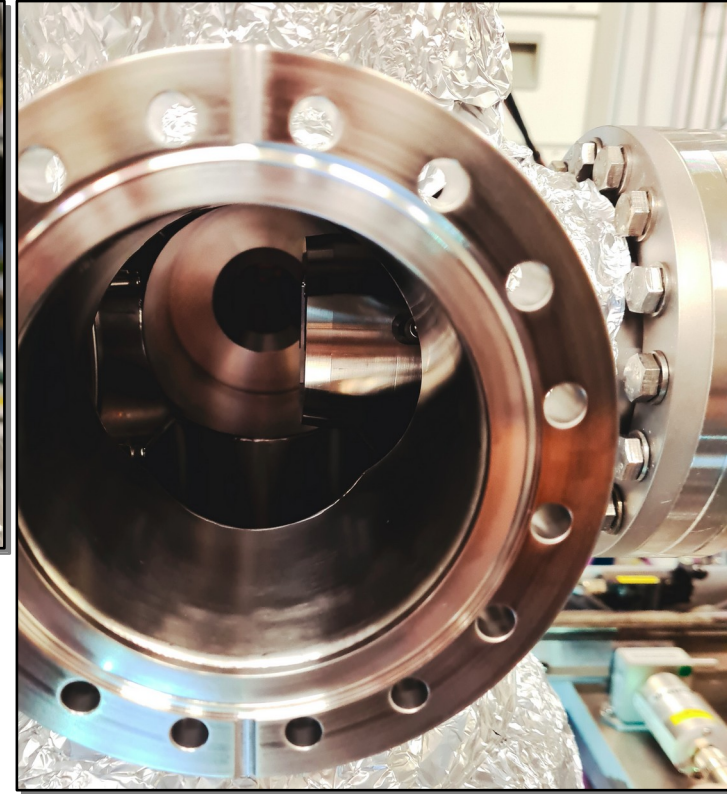


# TRISTAN Module Integration



**$^{83m}\text{Kr}$  electrons** from MoS

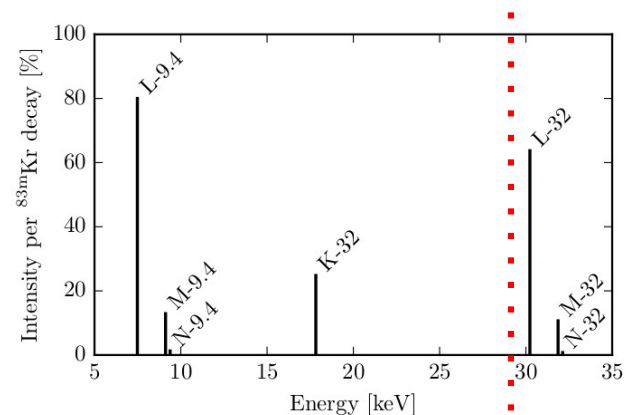
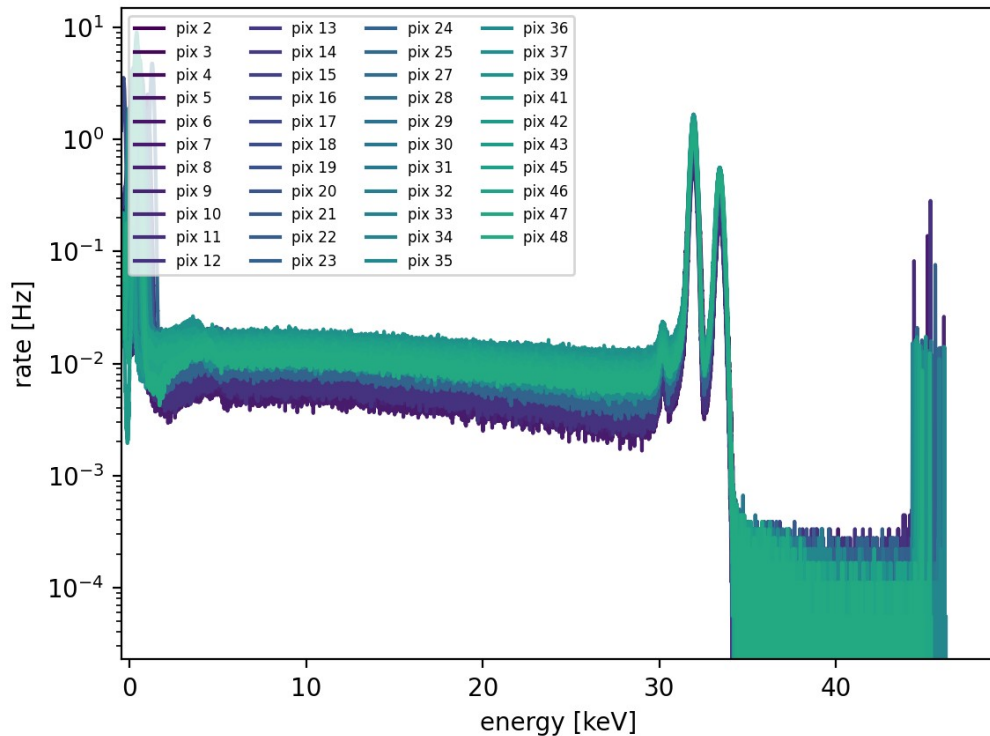
- Monoenergetic
- Vacuum ( $10^{-8}$  mbar)
- Operating at  $-50^{\circ}\text{C}$



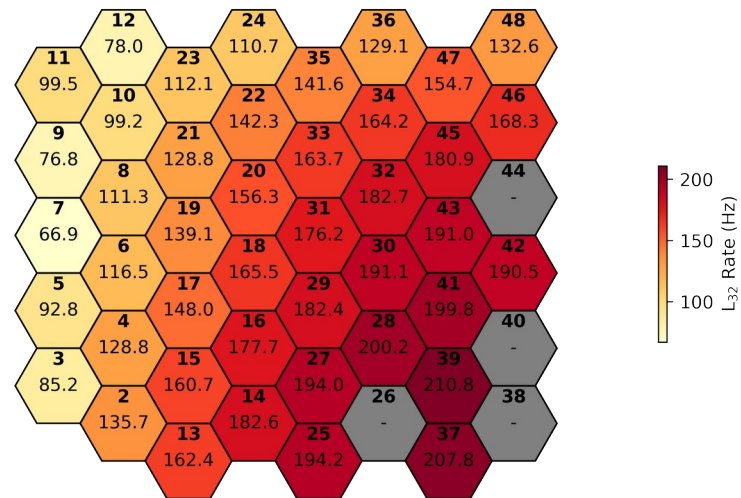
**1st electron light with a 47-pix TRISTAN inside MoS KATRIN on April 2021**



# TRISTAN Module Integration



- Studying of the **electron response**
- Analyzing plane HV to 30 kV to focus on L32 line
- Size of the source  $\sim$  size of the detector



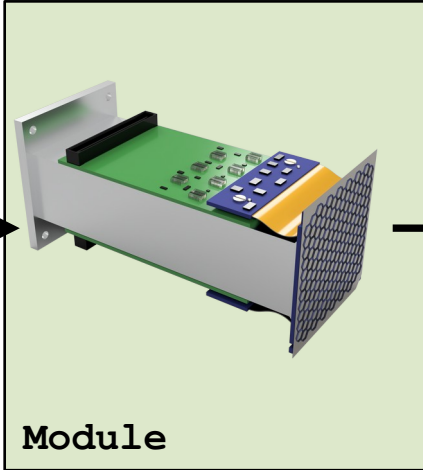


# Staged approach



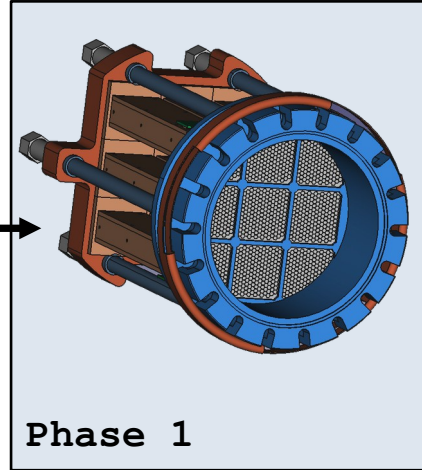
Prototypes

x24



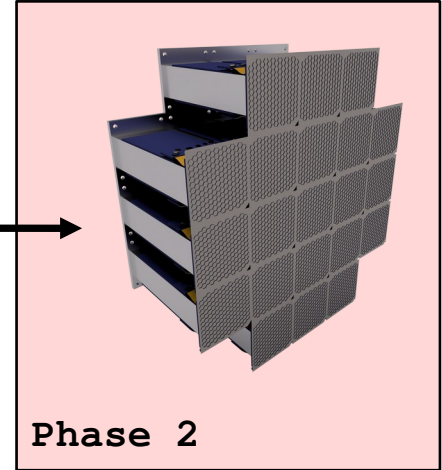
Module

x9



Phase 1

x2.5



Phase 2

## Mini-TRISTAN

- 9 x modules  
→ 1500

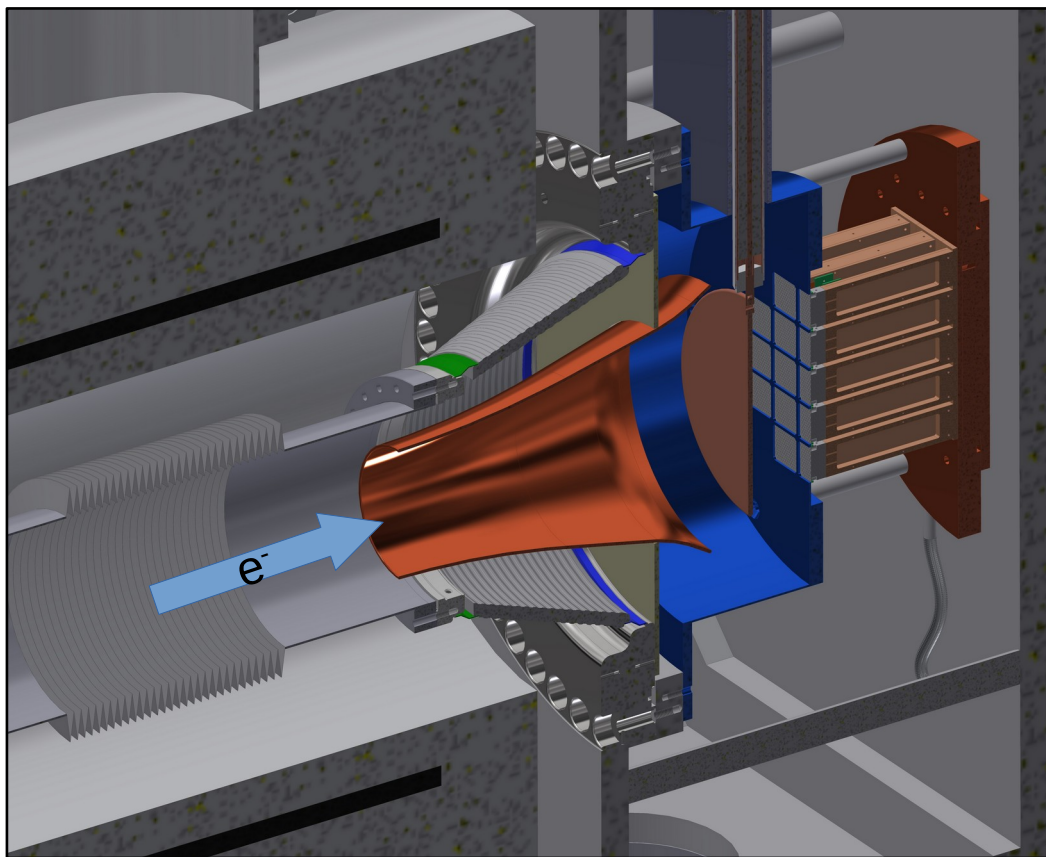
pixels

## Full TRISTAN

- 21 x modules  
→ 3500

pixels

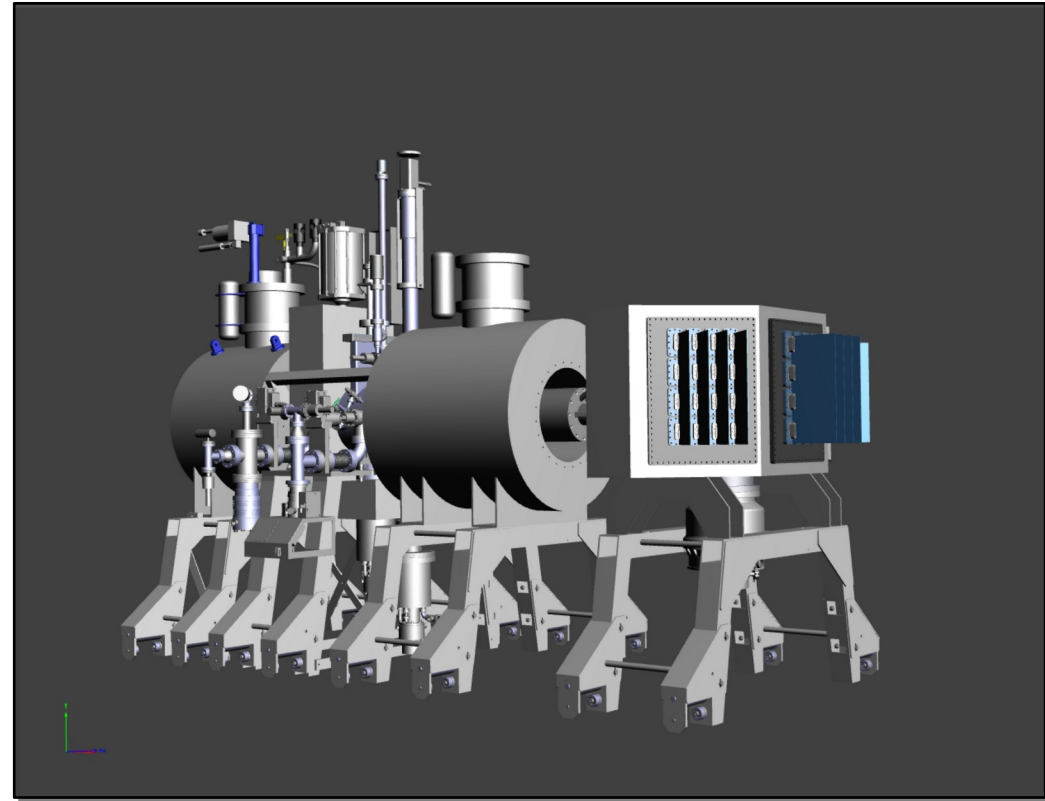
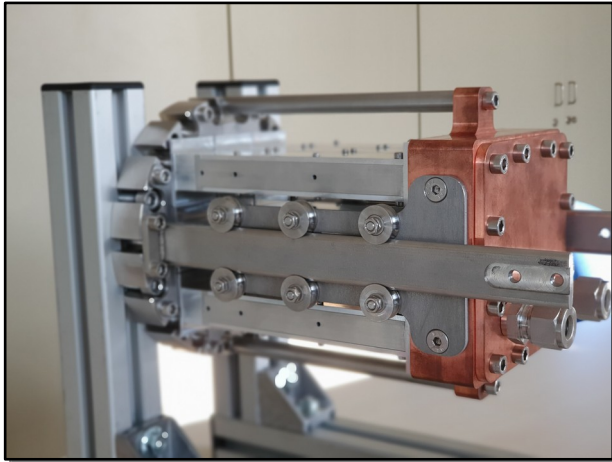
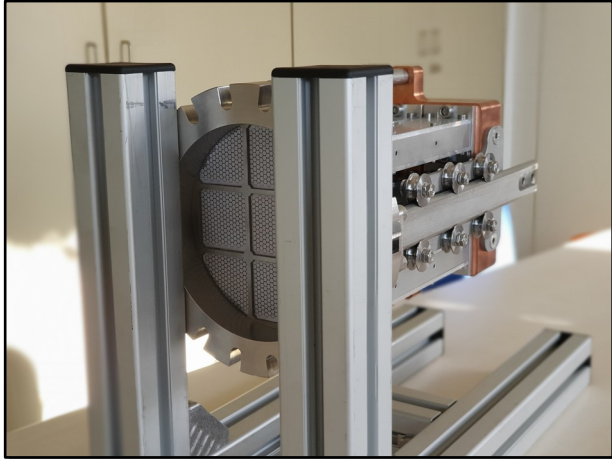
# TRISTAN Final Detector



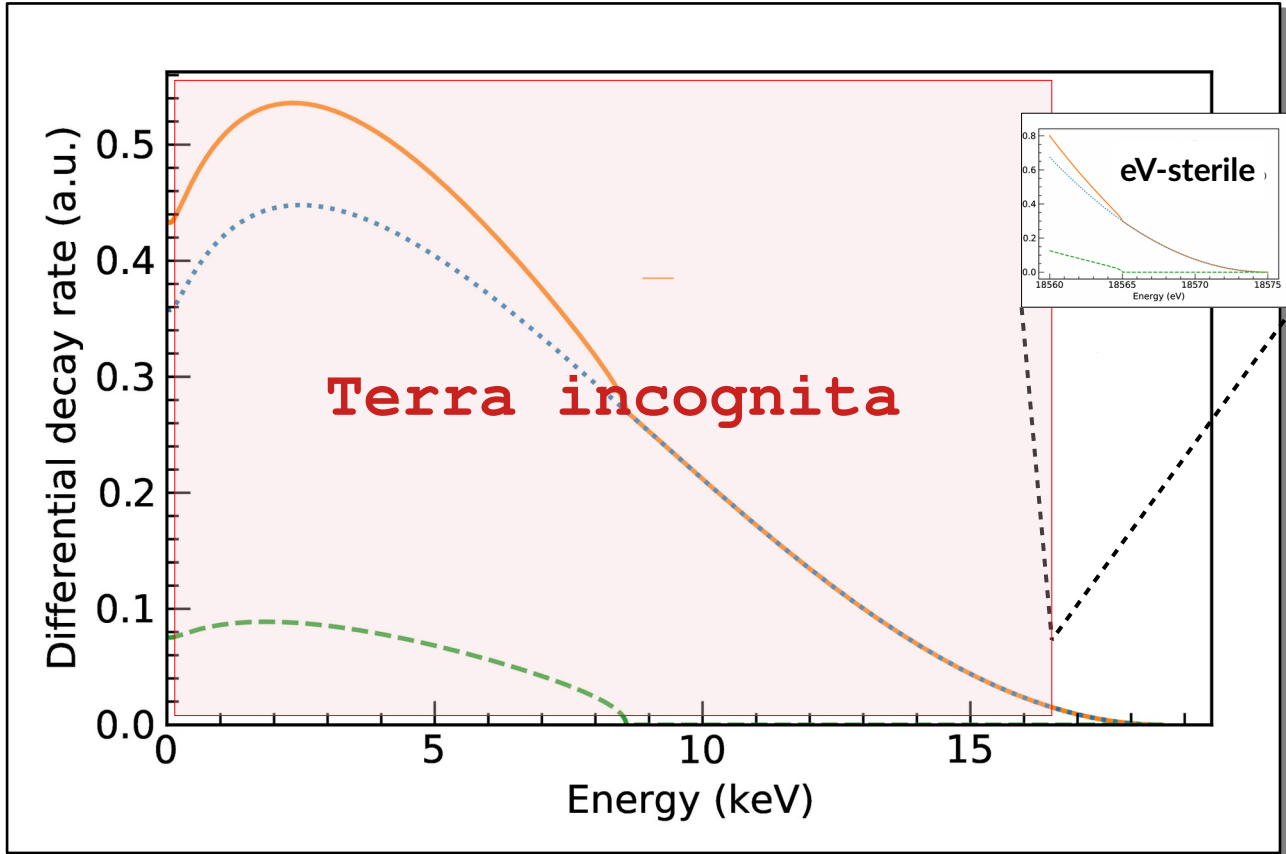
## Next milestones :

- **New DAQ** system
- **Full model** of the tritium spectrum
- **First physics runs** with **9 modules** in the KATRIN beamline (~ 2022)

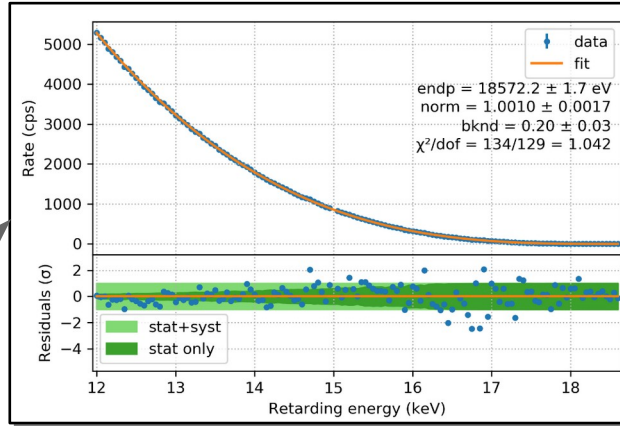
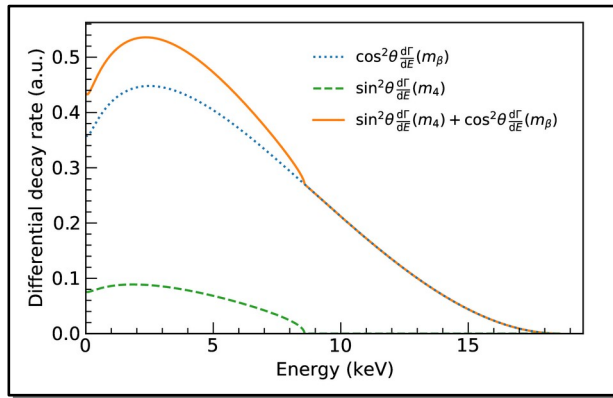
# Integration in KATRIN Beam-line



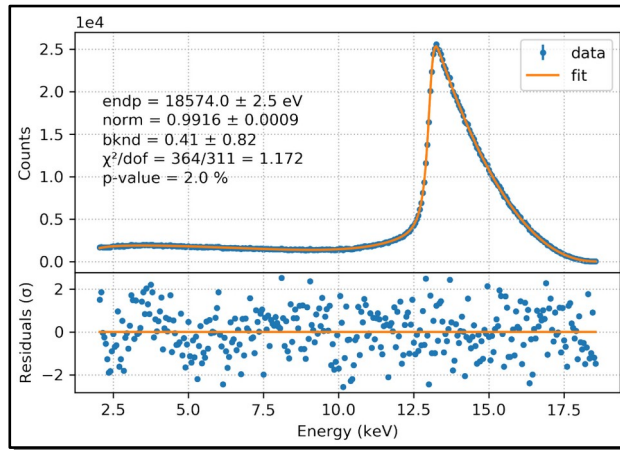
# Deep Tritium Model



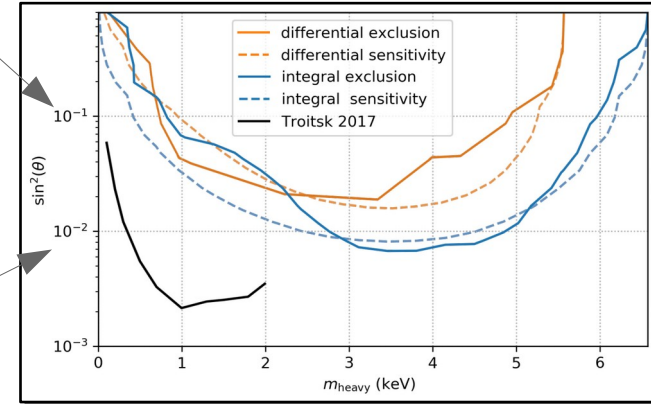
# Deep Tritium Model



JINST 14 P11013 (2019)



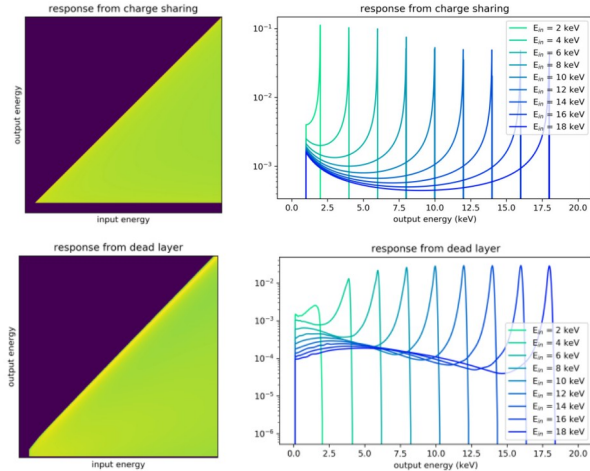
JINST 14 P11013 (2019)



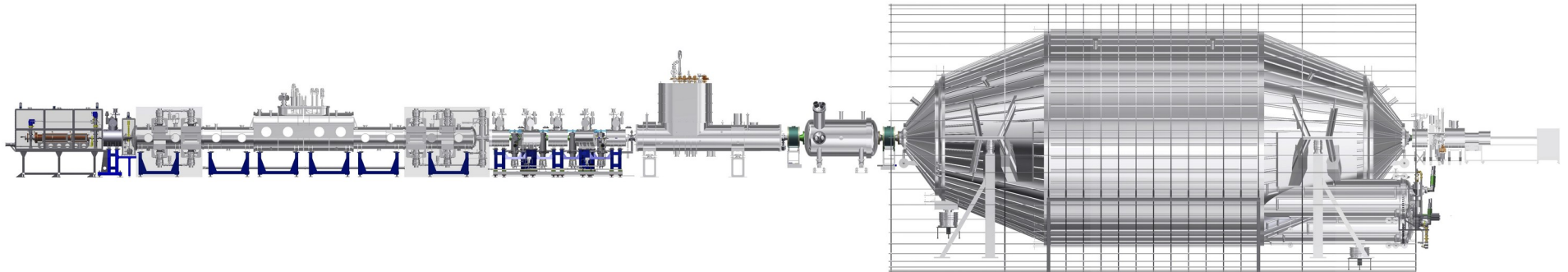
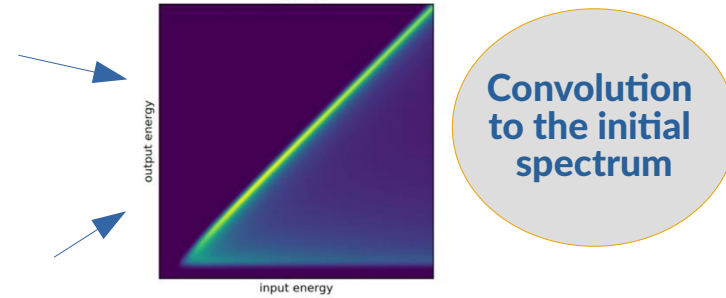


# Deep Tritium Model

Each physical effect described with a response matrix



Combined responses





# Deep Tritium Model : effects to consider

Effect can be different for differential and integral mode

## Rear-Wall

- Back-scattering
- Auger electron emission
- Residual beta-activity

## Source

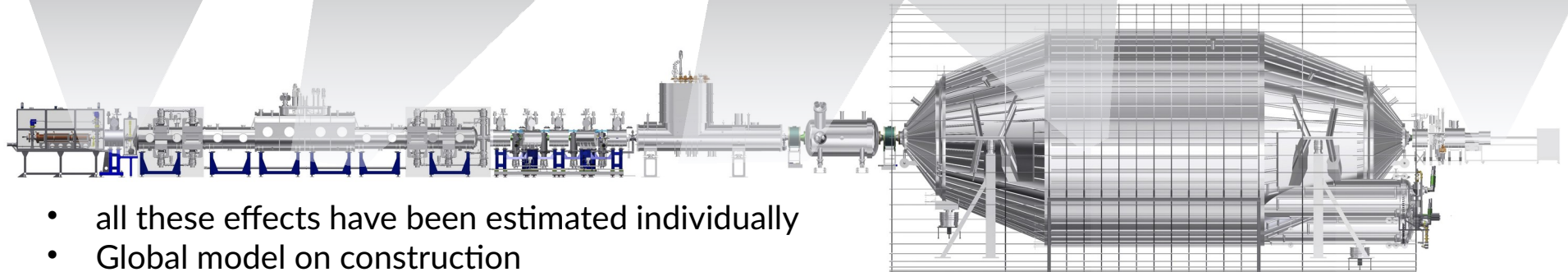
- Scattering
- Magnetic traps
- Plasma effects
- Stability
- Gas composition and impurities

## Transport - Spectrometer

- Non-adiabaticity transport
- Synchrotron radiation
- HV stability
- Background
- B-field stab.

## Detector

- SDD response
- Read-out resp
- ADC NL
- Post acceleration electrode
- Pile-up, backscattering
- Penning traps
- Stability



- all these effects have been estimated individually
- Global model on construction

# Deep Tritium Model : effects to consider

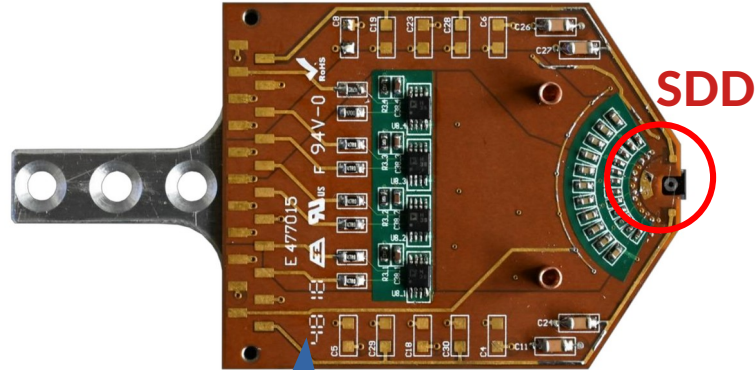
Effect can be different for differential and integral mode

## Rear-Wall

- Back-scattering
- Auger electron emission
- Residual beta-activity

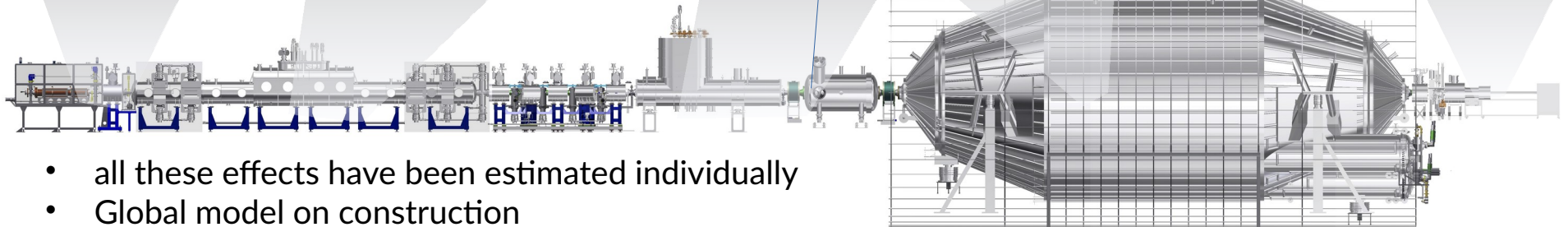
## Source

- Scattering
- Magnetic traps
- Plasma effects
- Stability
- Gas composition and impurities



## Detector

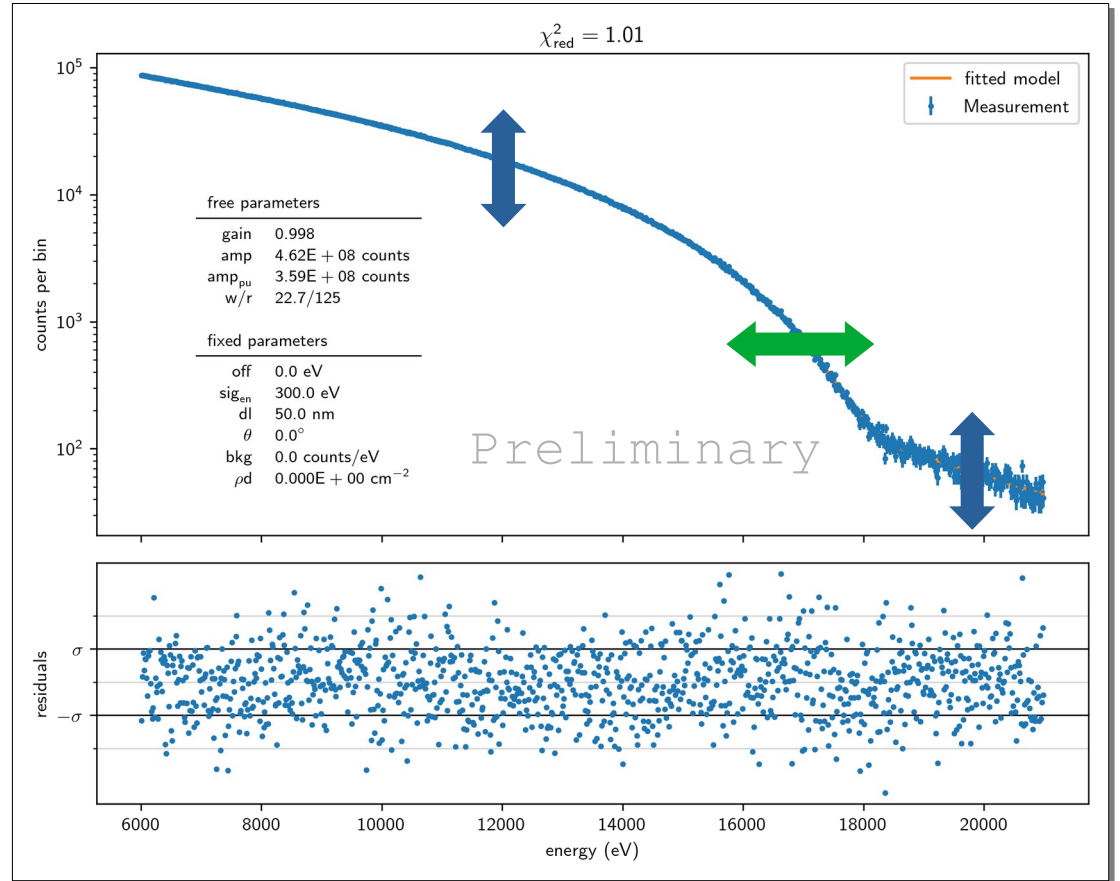
- SDD response
- Read-out resp
- ADC NL
- Post acceleration electrode
- Pile-up, backscattering
- Penning traps
- Stability



- all these effects have been estimated individually
- Global model on construction

# Deep Tritium Model : illustration

- Development of a full model for complete sensitivity studies
- **1h of nominal mass campaign**
- Total countrate: 16079 cps
- Countrate in ROI: 7496 cps
- Counts :  $2.3 \cdot 10^6$  electrons



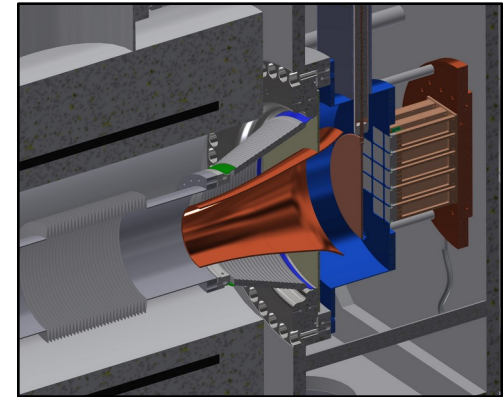
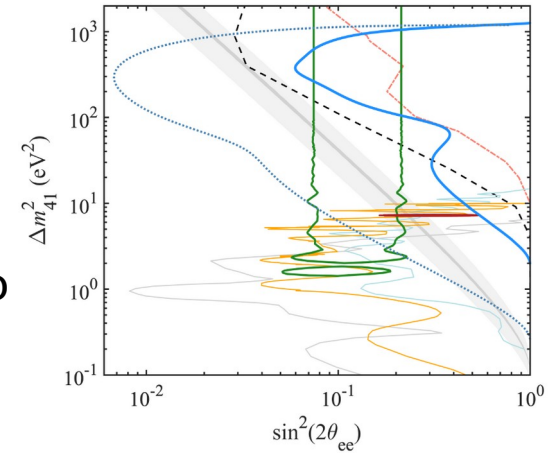
# Conclusion

## Sterile neutrinos with KATRIN

**KATRIN** has now presented a first study on eV sterile neutrinos

### **KATRIN with TRISTAN :**

- feasibility of the SDD technology for the keV-sterile neutrino search has been demonstrated
  - with photons and electrons
  - with tritium in realistic conditions (Troitsk)
- A **first TRISTAN module** is being commissioned in KATRIN
- First physics run with 9 module expected 2022
- A complete **deep tritium model** is being built. New sensitivity studies will be done to reduce systematics.





# Thank you all



11/10/2021

Séminaire PHE, IJCLab - Thibaut Houdy