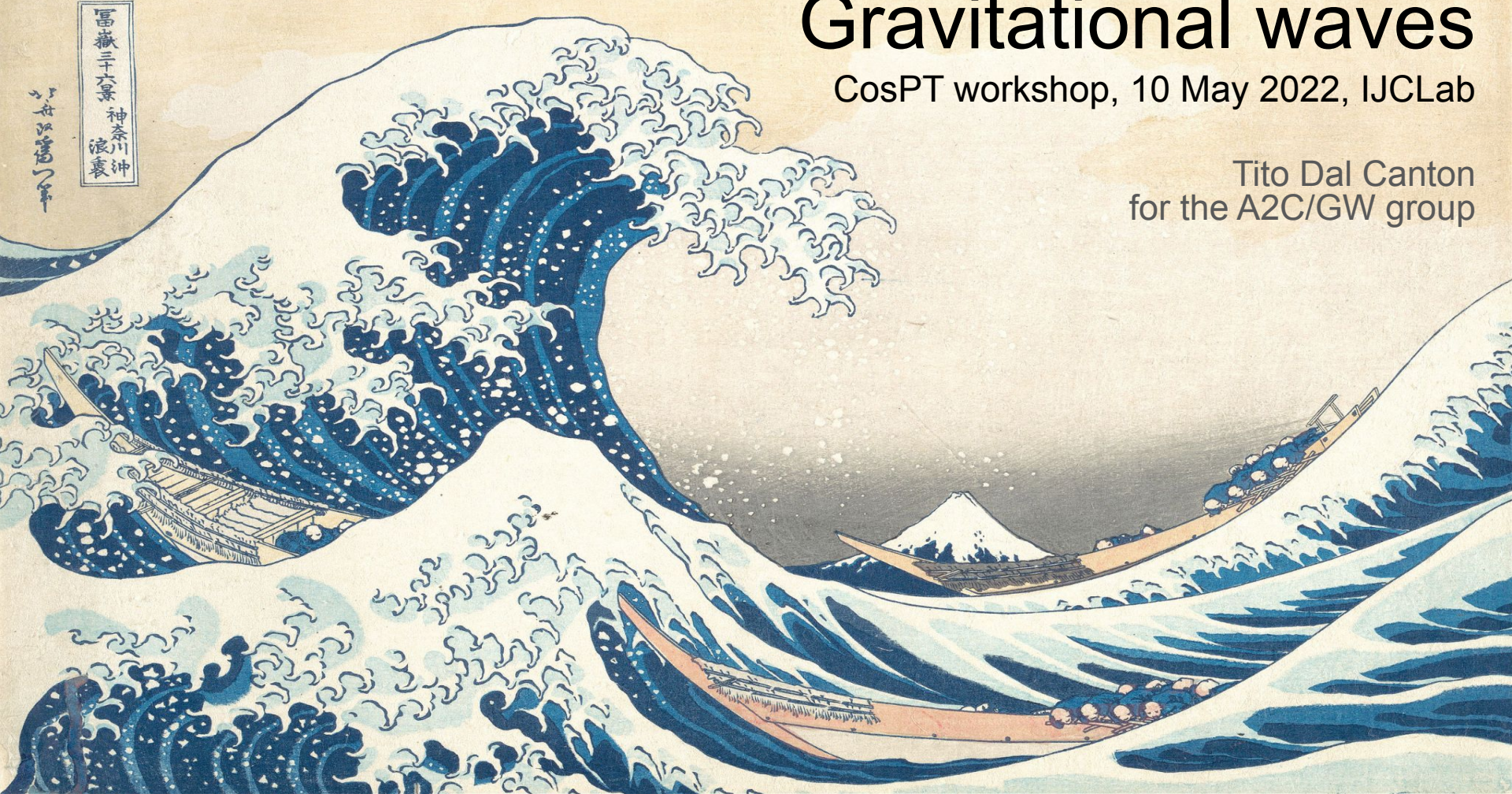


Gravitational waves

CosPT workshop, 10 May 2022, IJCLab

Tito Dal Canton
for the A2C/GW group



富嶽三十六景 神奈川浪裏

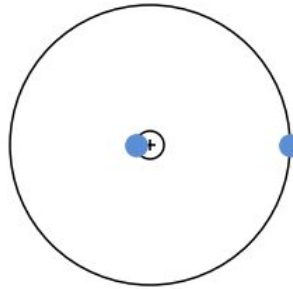
富嶽三十六景 神奈川浪裏

Gravitational-wave theory

$$R_{\mu\nu} - \frac{1}{2}Rg_{\mu\nu} = \frac{8\pi G}{c^4}T_{\mu\nu}$$

Flat spacetime Small perturbation

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$



Mass $\sim 10 M_{\text{Sun}}$
Velocity $\sim c$
Mass quadrupole Q



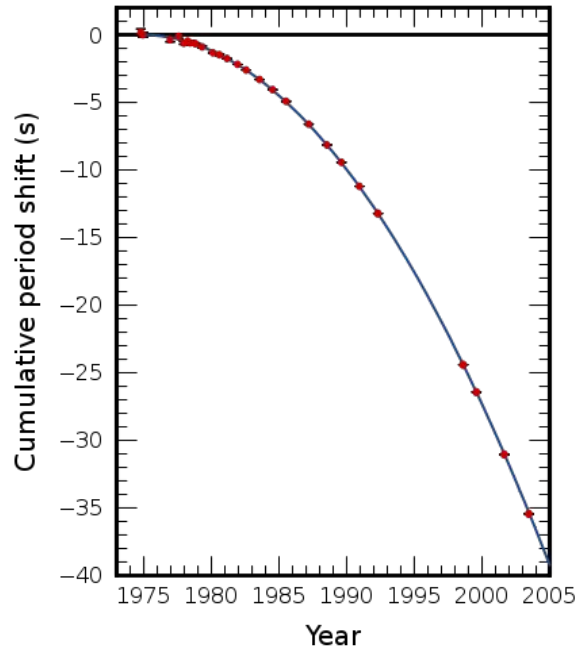
$r \sim 100 \text{ Mpc}$



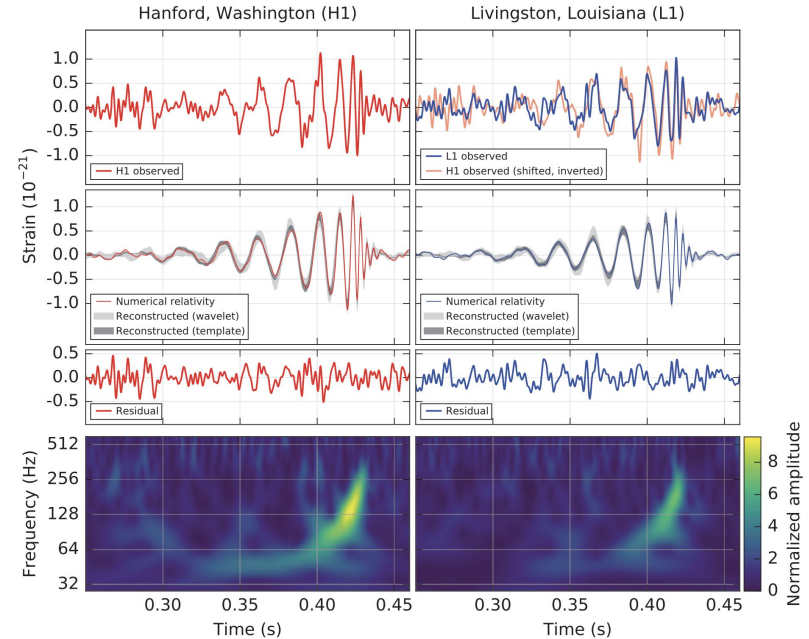
$$h_{ij} \sim \frac{G}{c^4} \frac{\ddot{Q}}{r} \sim 10^{-21}$$

Observational evidence of gravitational waves

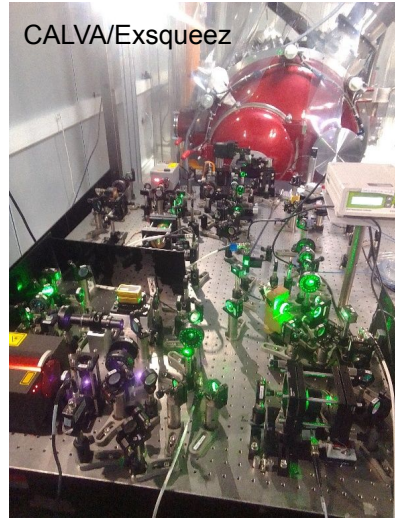
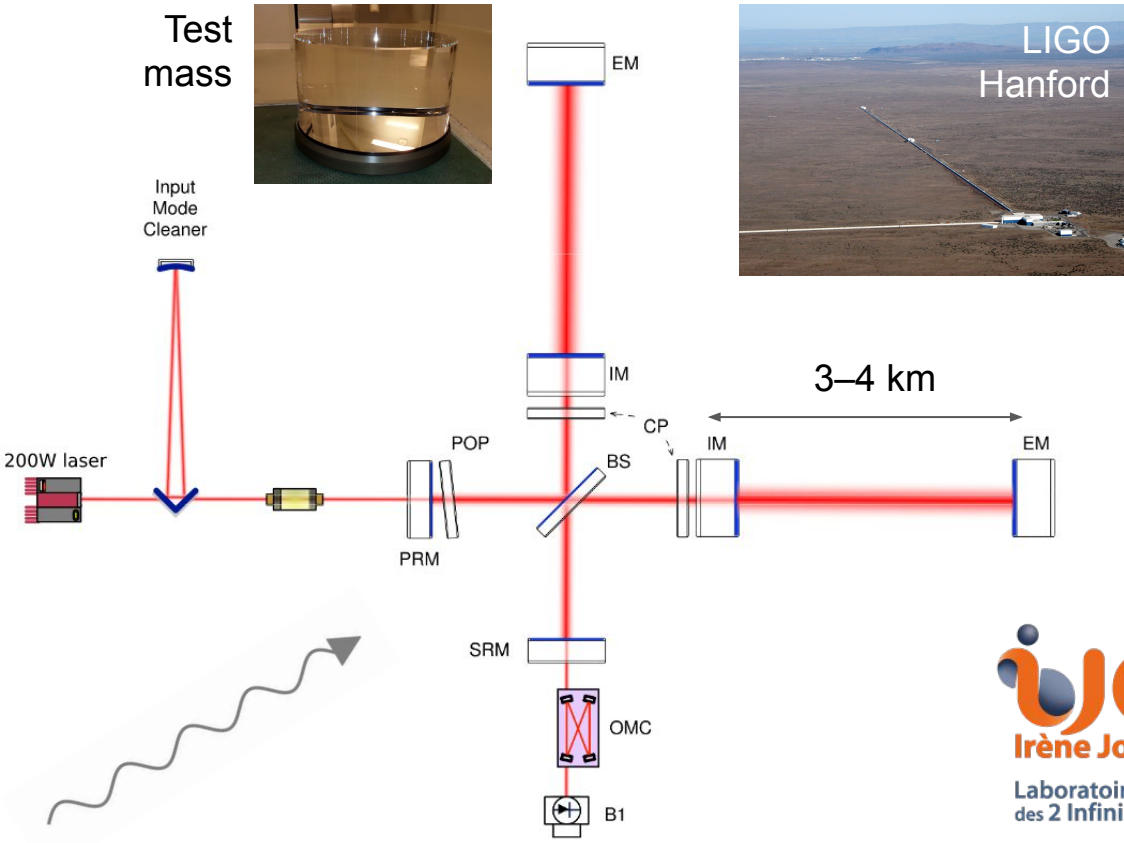
Orbital decay of binary pulsars



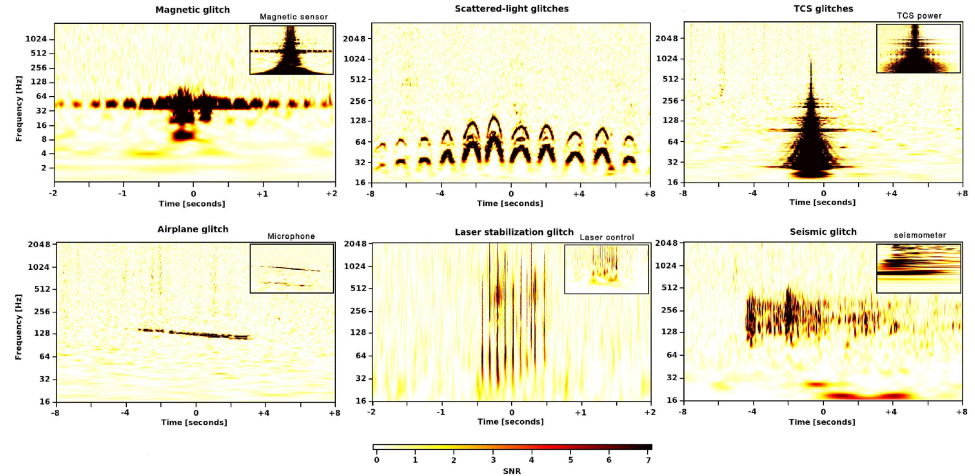
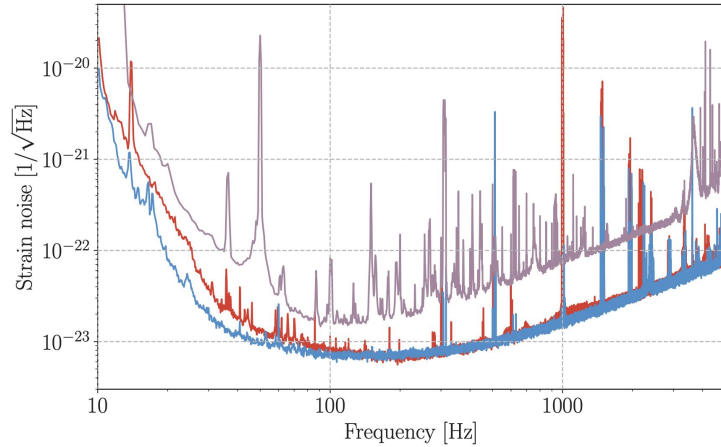
Direct local measurement of $h(t)$



Gravitational-wave detectors

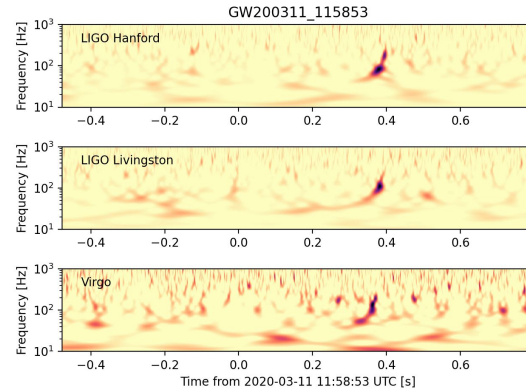
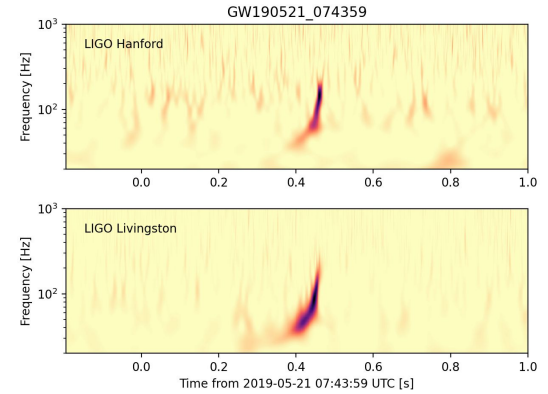
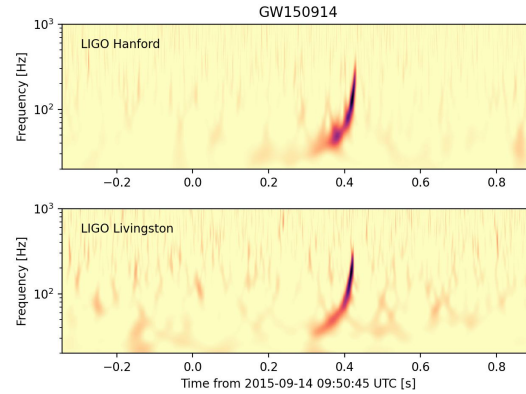
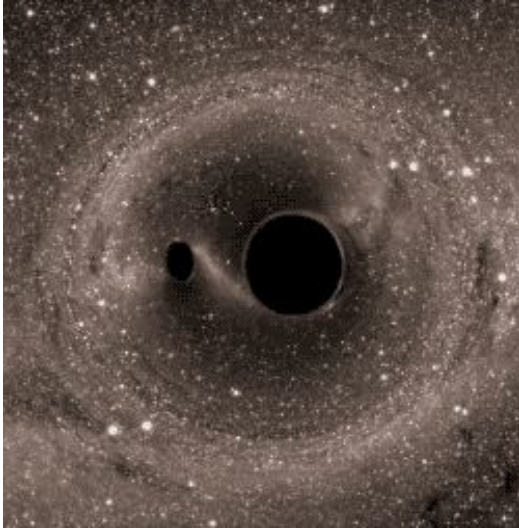


Gravitational-wave data analysis



Noise and data-quality characterization
Searching for GW signals in the data
Astrophysical interpretation, parameter inference
Population inference
Followup or cross-match with other sectors (GRBs, optical, neutrinos...)

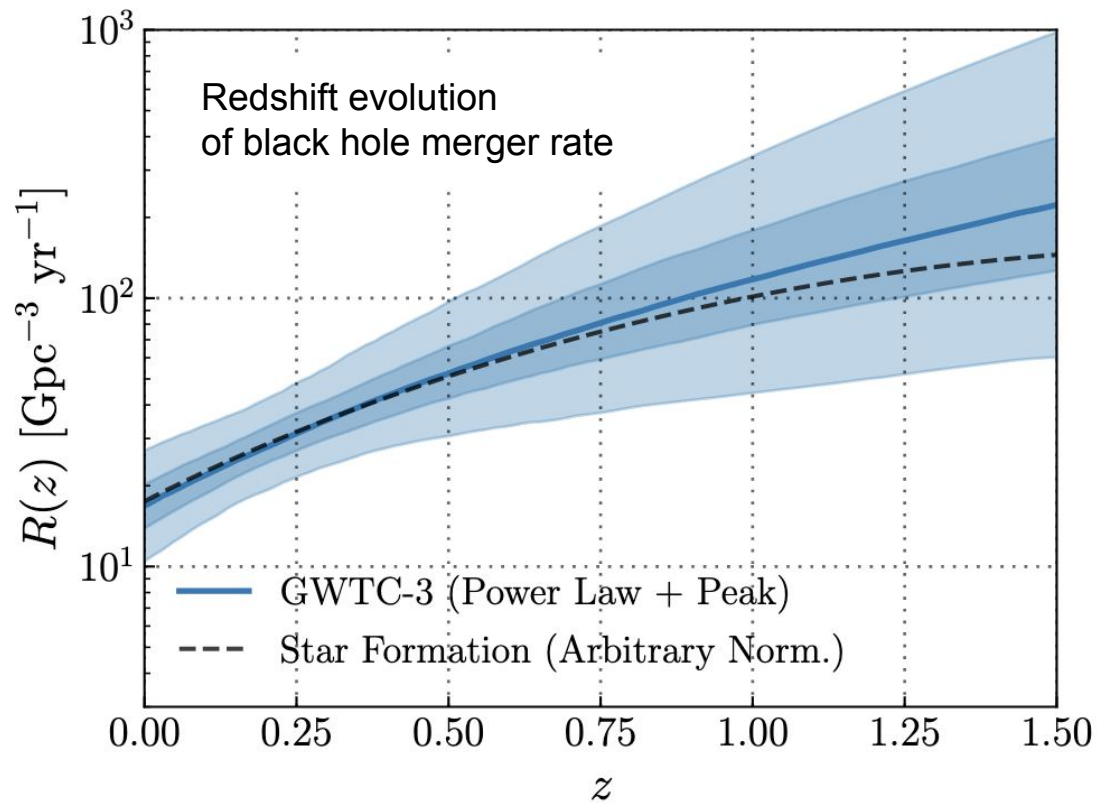
GWs from binary black hole mergers



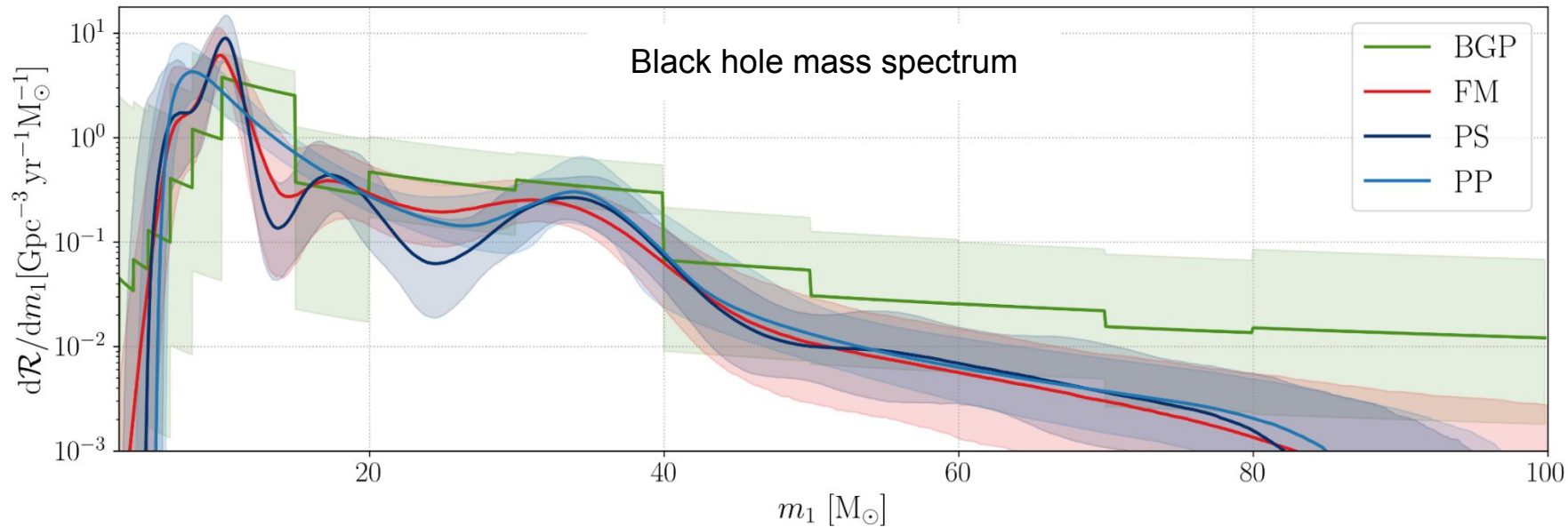
Binary black holes
exist and merge!

~80 observations
in GWTC-3

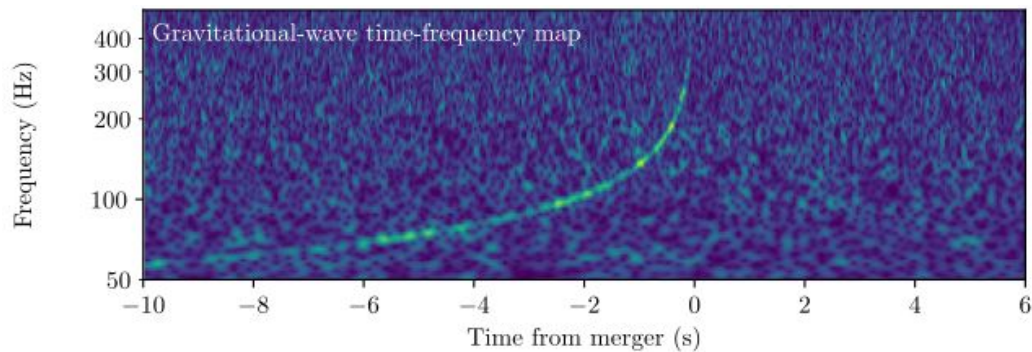
GWs from binary black hole mergers



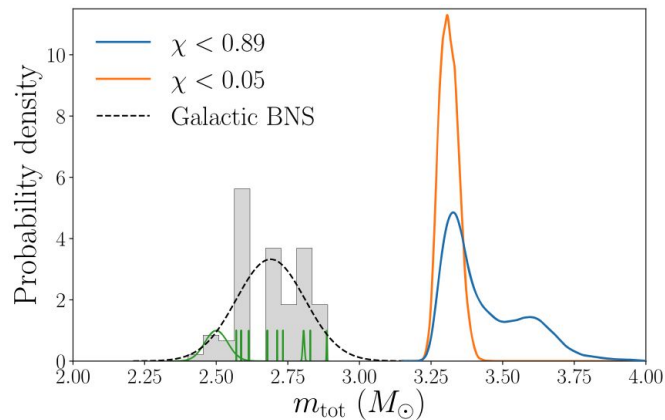
GWs from binary black hole mergers



GWs from binary neutron star mergers

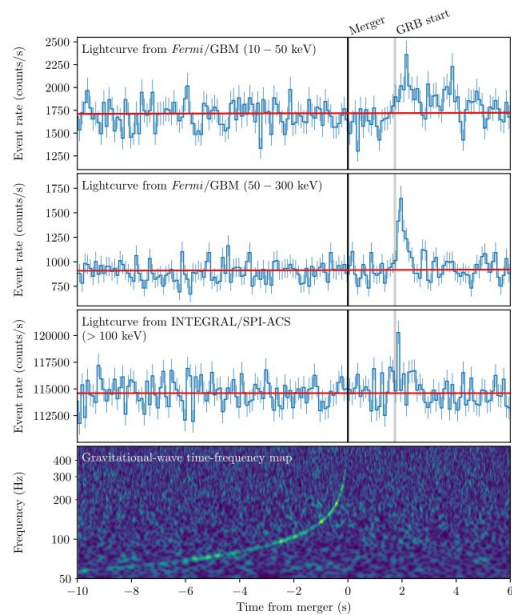


2 observations in GWTC:
GW170817 and GW190425

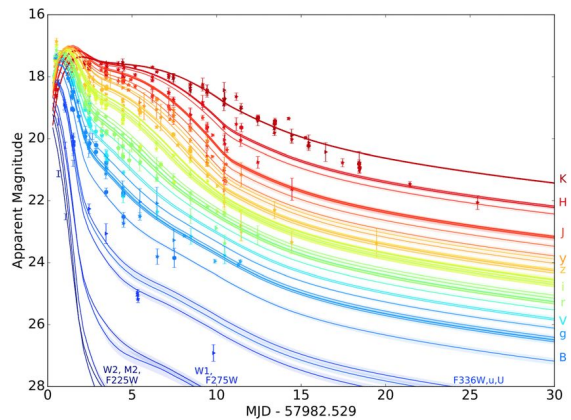


GWs from binary neutron star mergers

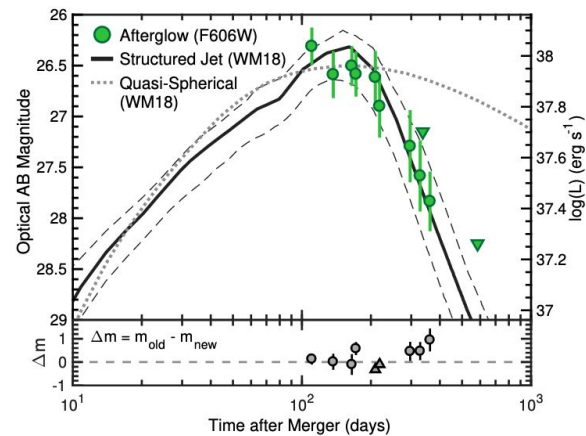
Short GRB



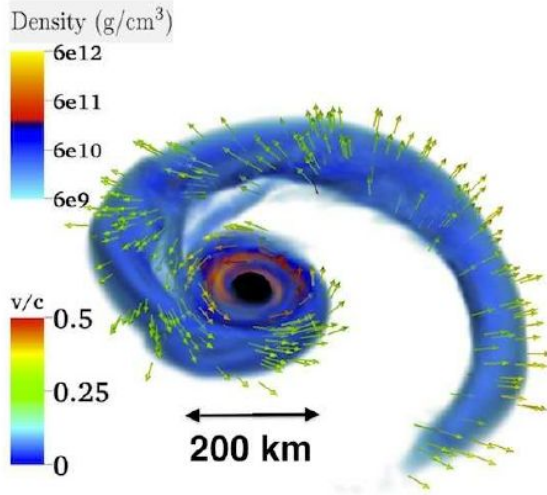
Kilonova



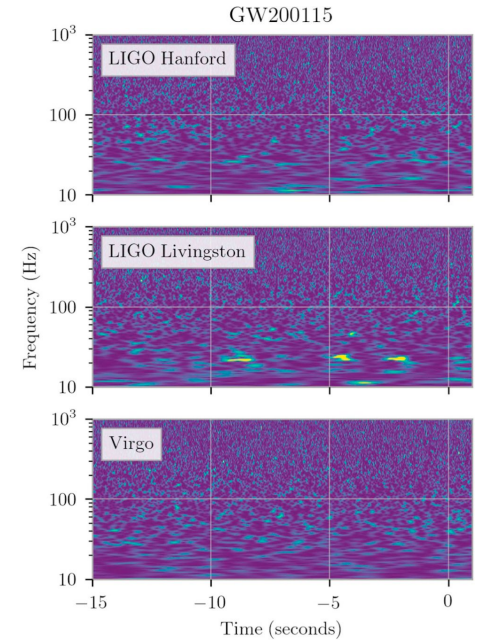
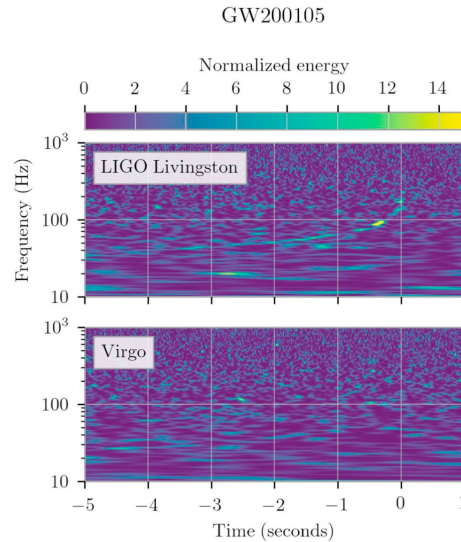
Long-term optical afterglow



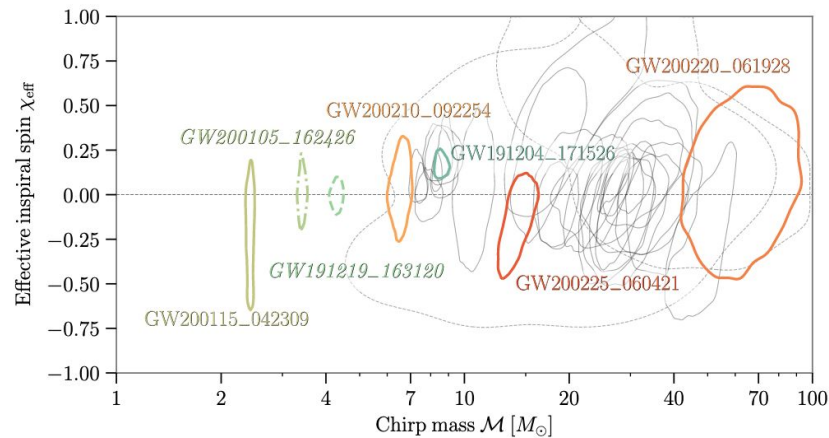
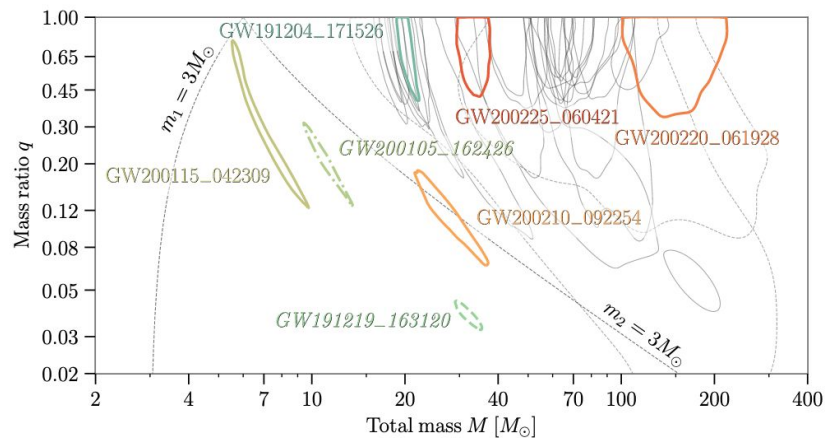
GWs from neutron-star–black-hole mergers



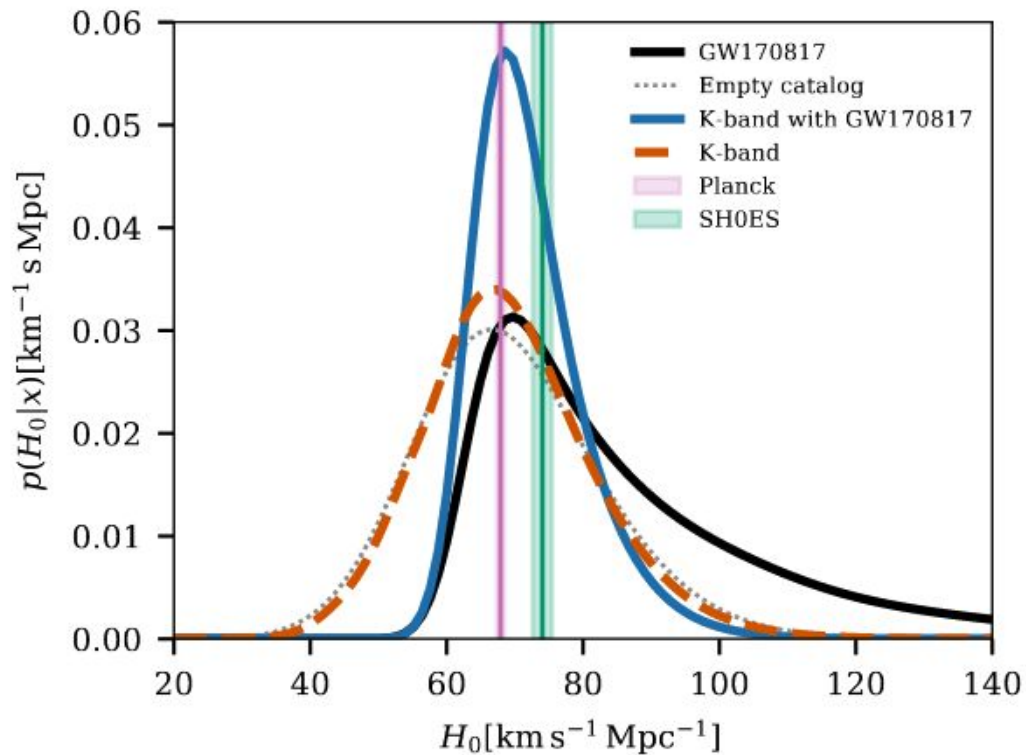
~2 observations in GWTC-3



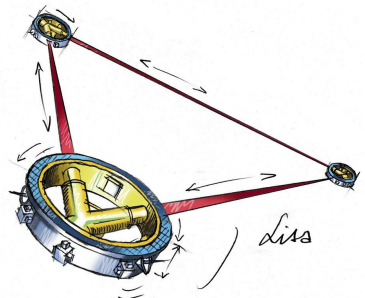
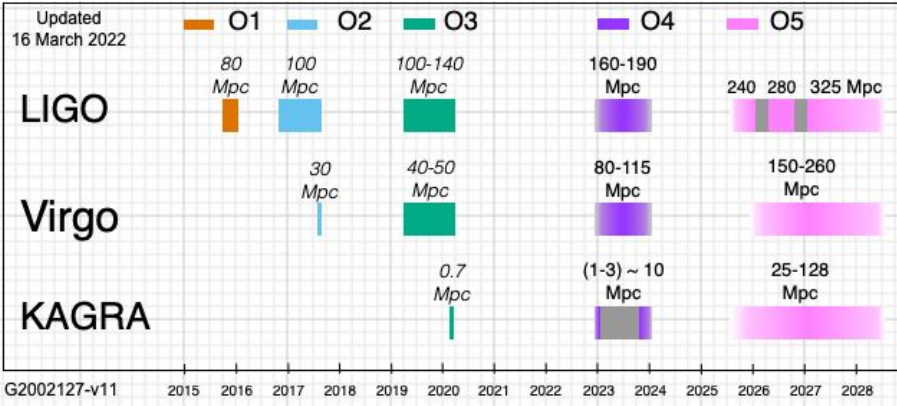
Group picture of recent compact binary mergers



H_0 measurement with NS and BH mergers



Evolution of GW detectors



Far future (mid-2030s):
space detectors



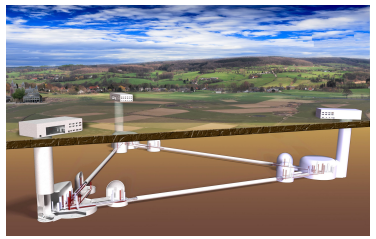
Present:
second-generation
ground-based
detectors



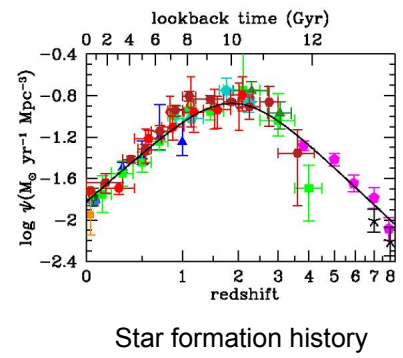
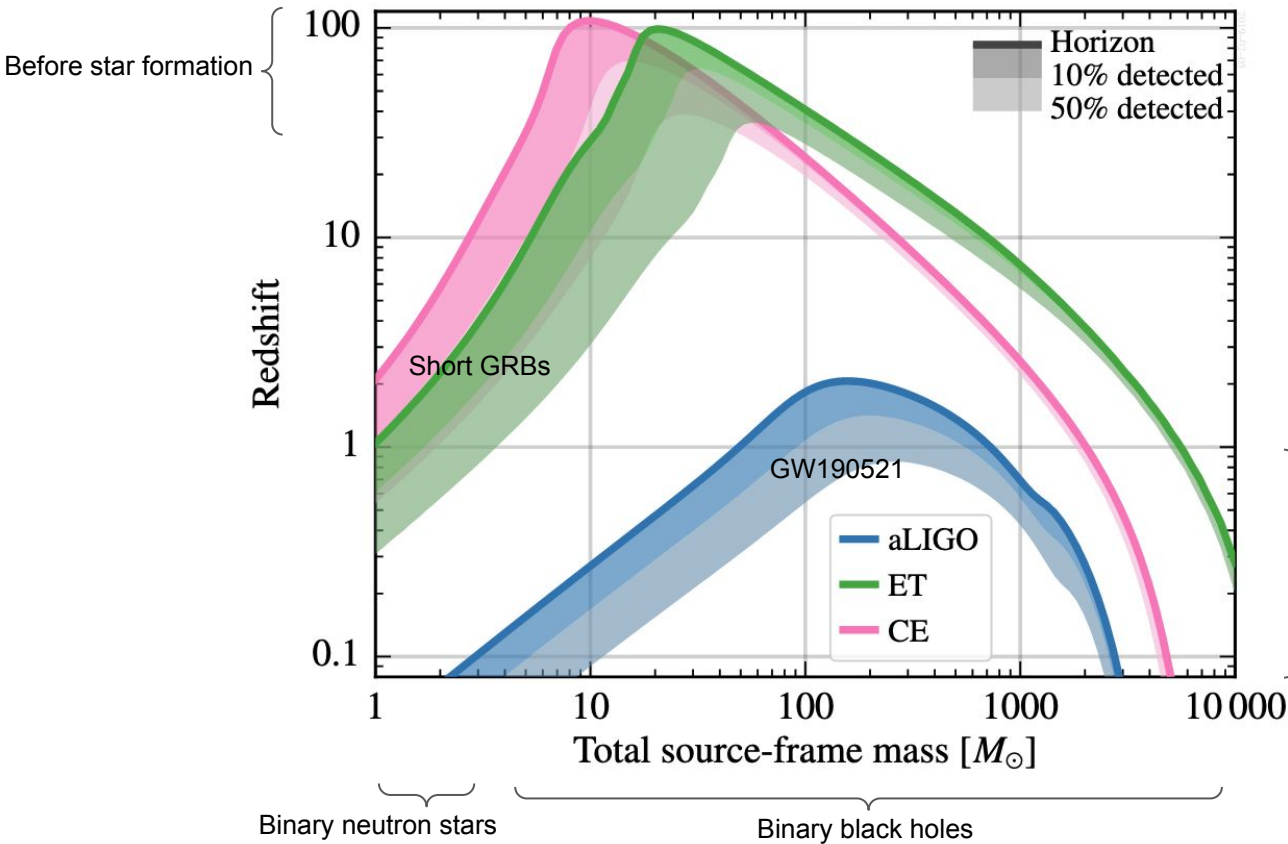
Near future:
“post-O5 era”
(mid-2020s
to early 2030s)



Far future (mid 2030s):
Third-generation
ground-based
detectors

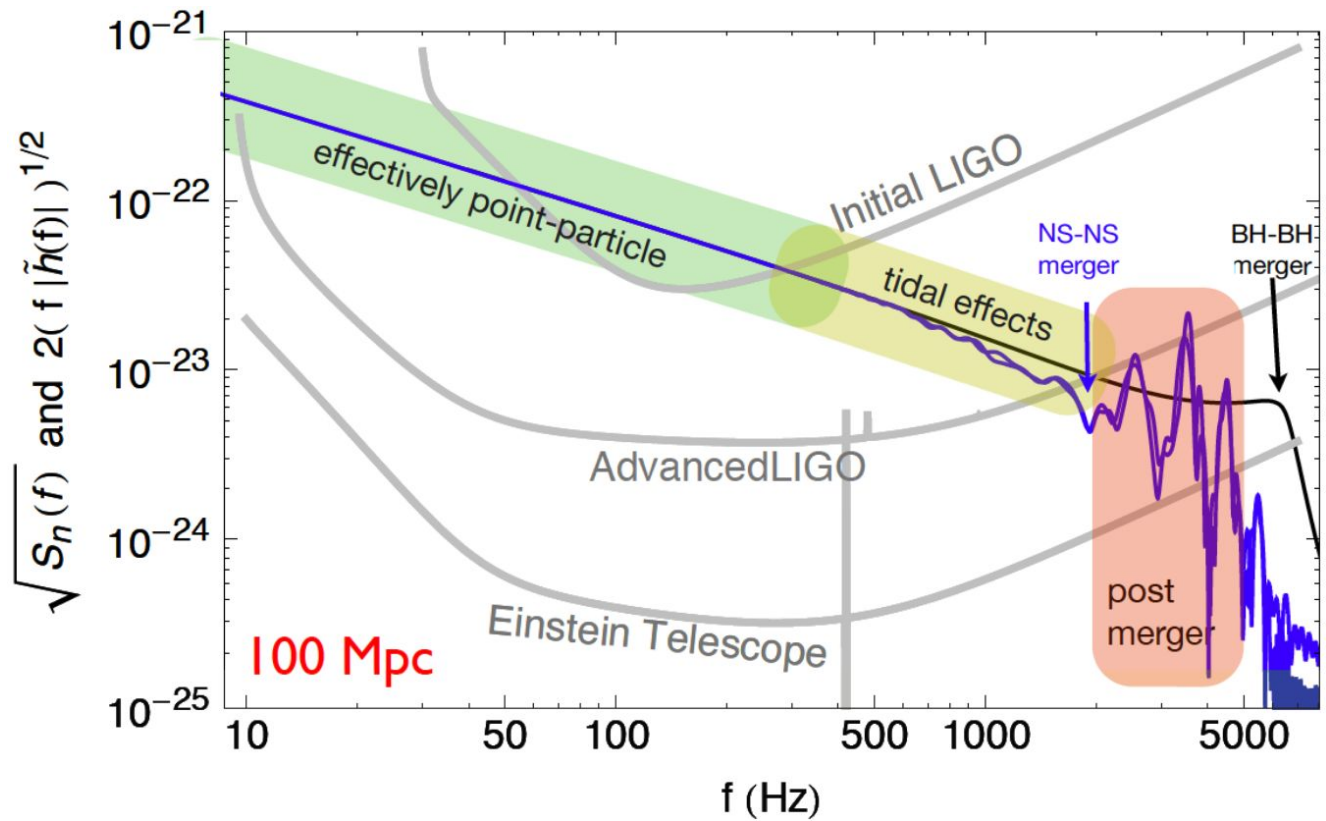


Einstein Telescope and Cosmic Explorer

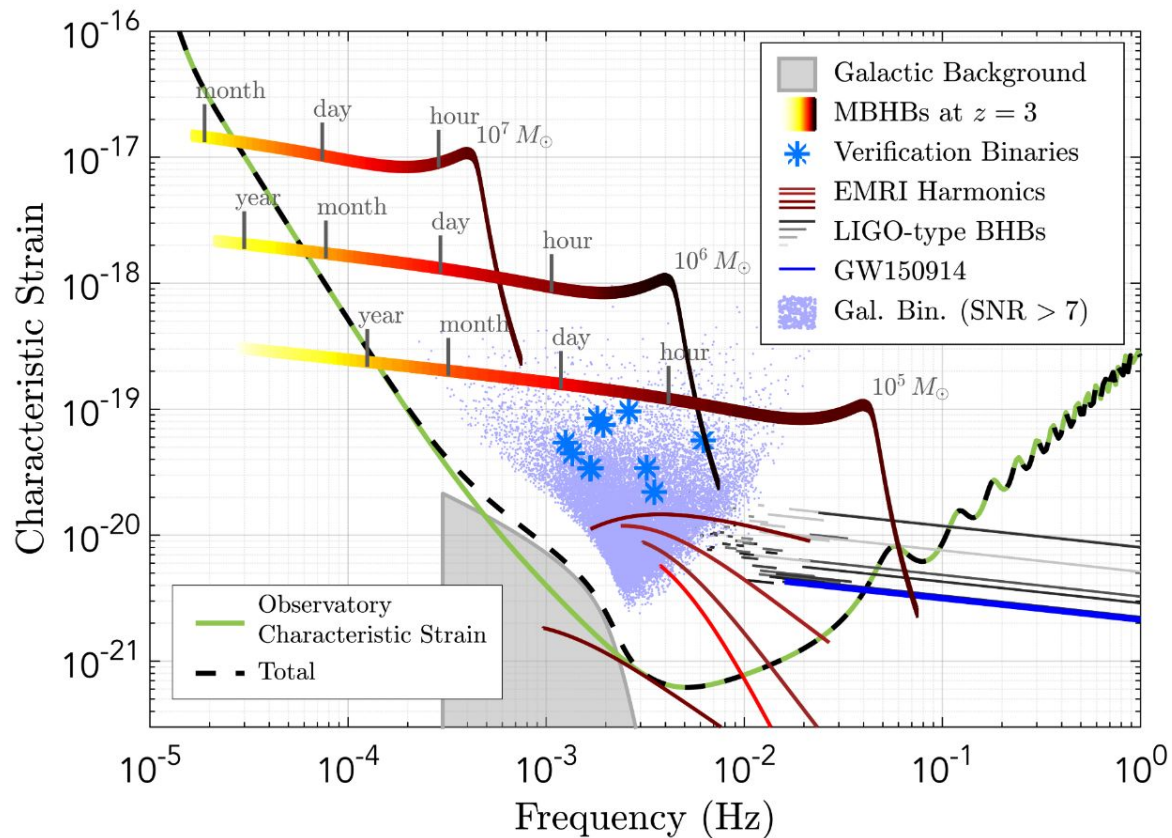


Signals with
S/N ~ 200
(~10x present)

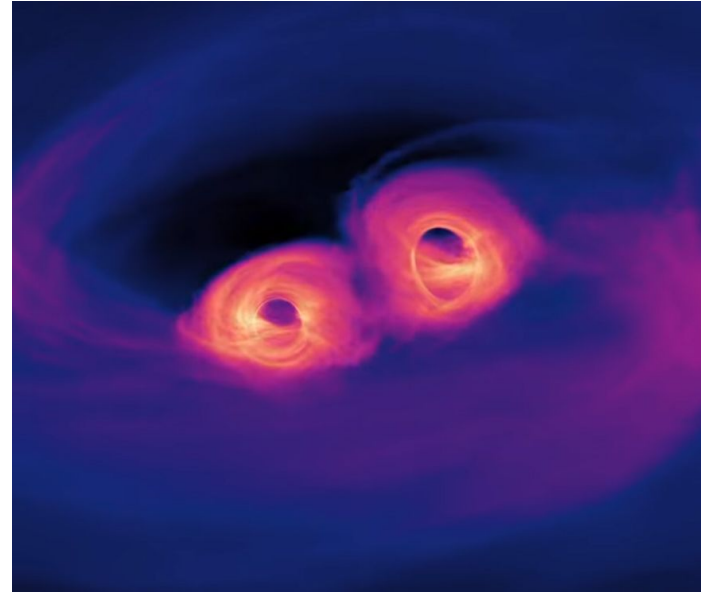
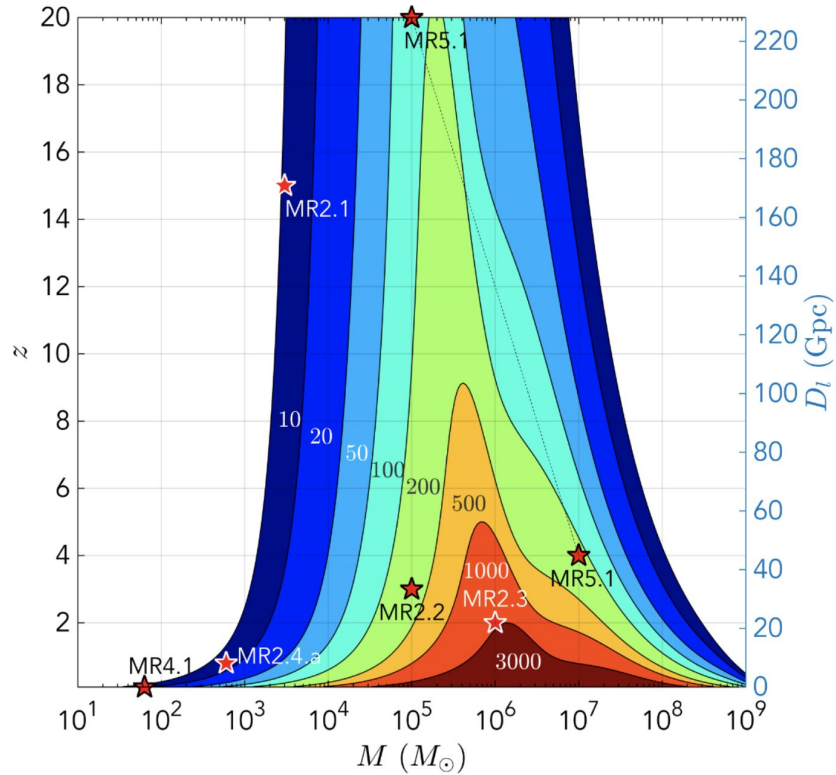
Einstein Telescope and Cosmic Explorer



Laser Interferometer Space Antenna



Laser Interferometer Space Antenna



Thank you!

