

# Long-lived particles and co-scattering

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Based on Alguero, GB, Kraml, Pukhov, 2207.10536, to appear in SciPost Phys

Paris-Saclay, 11/2022

# Outline

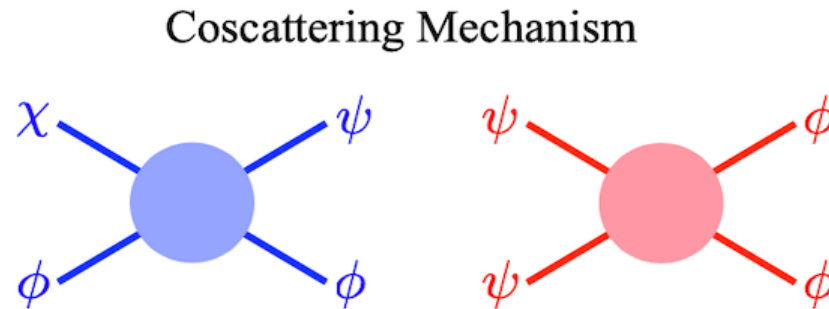
- Introduction
- Coscattering
- A case-study: singlet-triplet
- Remarks and conclusions

# Introduction

- Extensive experimental programs for searches for WIMPs in direct/in-direct searches as well as collider searches
- Many DM production mechanism in early universe: Freeze-out, freeze-in, asymmetric, coscattering...
- Once DM is discovered – can we identify the production mechanism?
- Currently at colliders active searches for long-lived particles
- When DM couplings are very weak and/or spectrum in the dark sector is compressed : possible displaced signatures

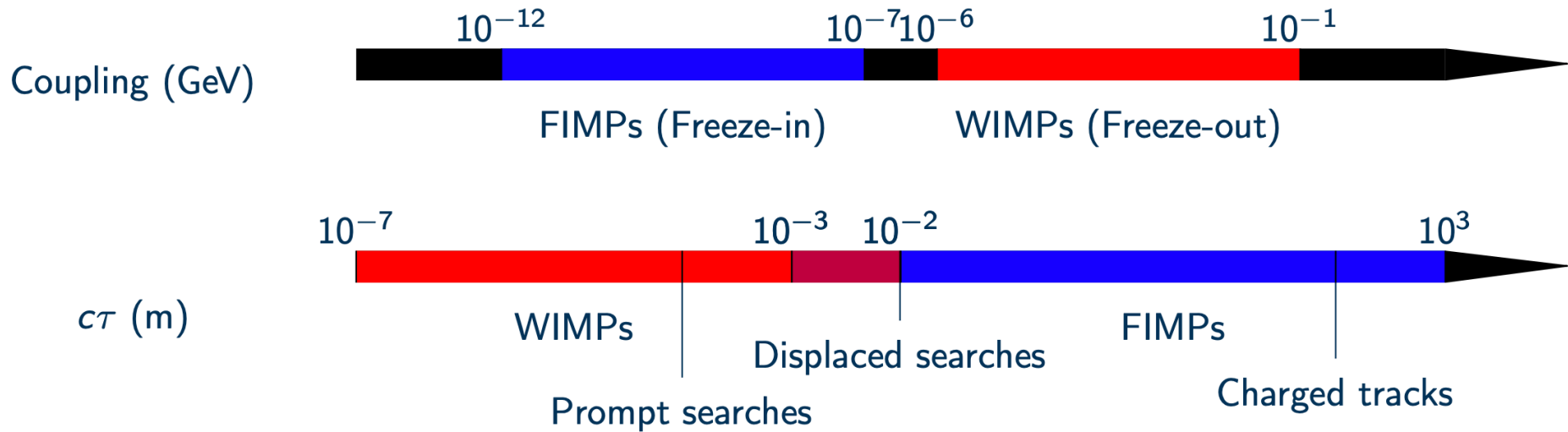
# Co-scattering - Conversion driven FO

- When interactions of DM with SM are very weak, processes  $\chi\chi \rightarrow \text{SMSM}(\phi\phi)$  are inefficient
- If other particles in the dark sector ( $\psi$ ), inelastic scattering can dominate DM production



- M. Garny et al 1705.09292 (conversion-driven FO)
- R.T D'Agnolo et al 1705.08450 (coscattering)
- Note : When red process decouples before blue process : coannihilation sets the relic abundance
- Otherwise FO of inelastic scattering (blue) sets the relic abundance
- Both coannihilation and coscattering require small mass splitting
- In coscattering typically  $\psi$  has long lifetime

# Co-scattering - Conversion driven FO



# A case study : singlet-triplet model

Field	SU(2) <sub>L</sub>	U(1) <sub>Y</sub>	$\mathbb{Z}_2$
$\ell_L$	<b>2</b>	-1/2	+
$e_R$	<b>1</b>	-1	+
$H_1$	<b>2</b>	1/2	+
$\chi$	<b>1</b>	0	-
$\psi$	<b>3</b>	0	-

- Standard model + SU(2) triplet ( $\psi$ ) + SU(2) singlet ( $\chi$ ) both odd under  $\mathbb{Z}_2$
- $\chi, \psi$  can be identified with bino, wino in split-SUSY with heavy Higgsinos
- couplings of  $\chi$  (DM) to SM : very weak but couplings to other dark sector states ( $\psi$ ) not negligible – will drive DM formation

# Lagrangian

$$\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{i}{2} \bar{\chi} \gamma^\mu \partial_\mu \chi + \frac{i}{2} \bar{\psi} \gamma^\mu D_\mu \psi - \frac{1}{2} (m \bar{\chi} \chi + M \bar{\psi} \psi) + \mathcal{L}_5 + \mathcal{L}_{\geq 6},$$

$$\mathcal{L}_5 = -\frac{1}{2} \frac{\kappa}{\Lambda} \bar{\psi} \psi H^\dagger H - \frac{1}{2} \frac{\kappa'}{\Lambda} \bar{\chi} \chi H^\dagger H - \frac{\lambda}{\Lambda} \bar{\chi} \psi^a H^\dagger \tau^a H + \text{l.c.} + \dots,$$

- Only H gets a vev,  $v$
- Mass eigenstates: (singlet  $\sim m$ ; triplet  $\sim M$ )

$$m_{\tilde{\chi}, \tilde{\psi}^0} = \frac{1}{2} \left( m + M \mp \sqrt{(M - m)^2 + 4a^2} \right), \quad \text{where } a = \lambda v^2 / (2\Lambda).$$

- Small mixing singlet-triplet:  $\theta \approx \frac{\lambda v^2}{2\Lambda(M - m)}$

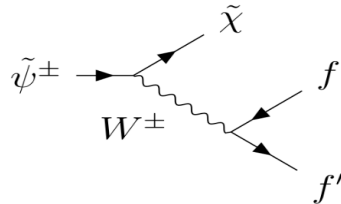
- Mixing also lifts degeneracy between neutral and charged triplet  $\sim 160 \text{ MeV}$  (McKay, Scott, arXiv:1712:00968)

# Interactions

$$\mathcal{L}_{W^\pm \tilde{\psi}^\mp \tilde{\chi}} = -g \sin \theta \bar{\tilde{\chi}} \gamma^\mu W_\mu^+ \tilde{\psi}^- + g \sin \theta \bar{\tilde{\chi}} \gamma^\mu W_\mu^- \tilde{\psi}^+ + \text{h.c.}$$

$$\mathcal{L}_{W^\pm \tilde{\psi}^\mp \tilde{\psi}^0} = g \cos \theta \bar{\tilde{\psi}}^0 \gamma^\mu W_\mu^+ \tilde{\psi}^- .$$

- For small mixing with  $\chi$ ,  $\psi^+$  is long-lived, competing decays : 3-body decay  $\psi^+ \rightarrow \chi f f$  or 2-body decay  $\tilde{\psi}^\pm \rightarrow \tilde{\psi}^0 \pi^\pm$



- Typical lifetime of charged triplet : 1cm- 1 m
- Couplings to Higgs, same order as coupling to gauge bosons:

$$\sim \frac{\lambda v}{\sqrt{2}\Lambda} \bar{\tilde{\chi}} \tilde{\psi}^0 h$$

- Couplings of DM further suppressed: 
$$-\frac{\lambda^2 v^3}{2\sqrt{2}\Lambda^2(M-m)} \bar{\tilde{\chi}} \tilde{\chi} h$$



# Solving for co-scattering (as implemented in micrOMEGAs)

- Small couplings  $\rightarrow$  particles in the dark sector might not be in thermal equilibrium with each other  $\rightarrow$  separate Boltzmann equations
- Assume kinetic equilibrium
- Assign particles to 'dark sectors', within each sector assume that particles are in thermal equilibrium
- Sector 0 : SM; sector 1 : singlet, sector 2 : triplet
- Singlet is lightest, triplet can decay to singlet+SM, triplet has em interacti  
niverse

$$\begin{aligned} \frac{dY_1}{dT} = \frac{1}{3H} \frac{ds}{dT} & \left[ \langle \sigma_{1100} v \rangle (Y_1^2 - Y_1^{eq2}) + \langle \sigma_{1122} v \rangle \left( Y_1^2 - Y_2^2 \frac{Y_1^{eq2}}{Y_2^{eq2}} \right) \right. \\ & + \langle \sigma_{1200} v \rangle (Y_1 Y_2 - Y_1^{eq} Y_2^{eq}) + \langle \sigma_{1222} v \rangle \left( Y_1 Y_2 - Y_2^2 \frac{Y_1^{eq}}{Y_2^{eq}} \right) \\ & \left. - \langle \sigma_{1211} v \rangle \left( Y_1 Y_2 - Y_1^2 \frac{Y_2^{eq}}{Y_1^{eq}} \right) - \frac{\Gamma_{2 \rightarrow 1}}{s} \left( Y_2 - Y_1 \frac{Y_2^{eq}}{Y_1^{eq}} \right) \right], \end{aligned}$$

- Similar Eq. For  $Y_2$

- When coupling of singlet ( $\lambda$ ) is very small : self-annihilation of singlet negligible

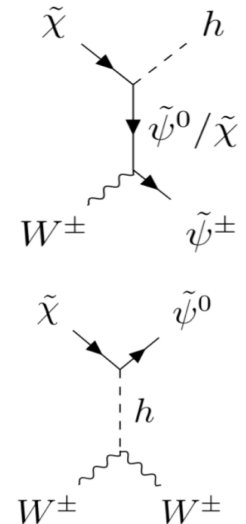
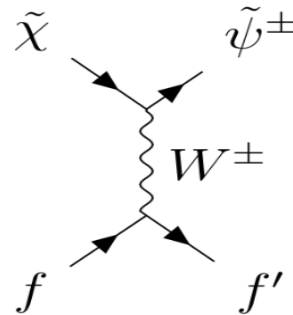
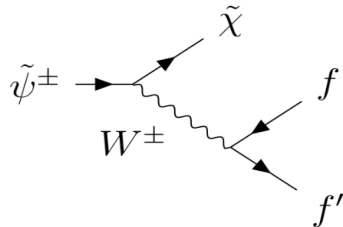
$$\frac{dY_1}{dT} = \frac{-\Gamma_{2 \rightarrow 1}}{HT} \left[ Y_2 - Y_1 \frac{Y_2^{eq}}{Y_1^{eq}} \right],$$

$$\frac{dY_2}{dT} = \frac{s}{HT} \left[ \langle \sigma_{2200} v \rangle (Y_2^2 - Y_2^{eq2}) + \frac{\Gamma_{2 \rightarrow 1}}{s} \left( Y_2 - Y_1 \frac{Y_2^{eq}}{Y_1^{eq}} \right) \right].$$

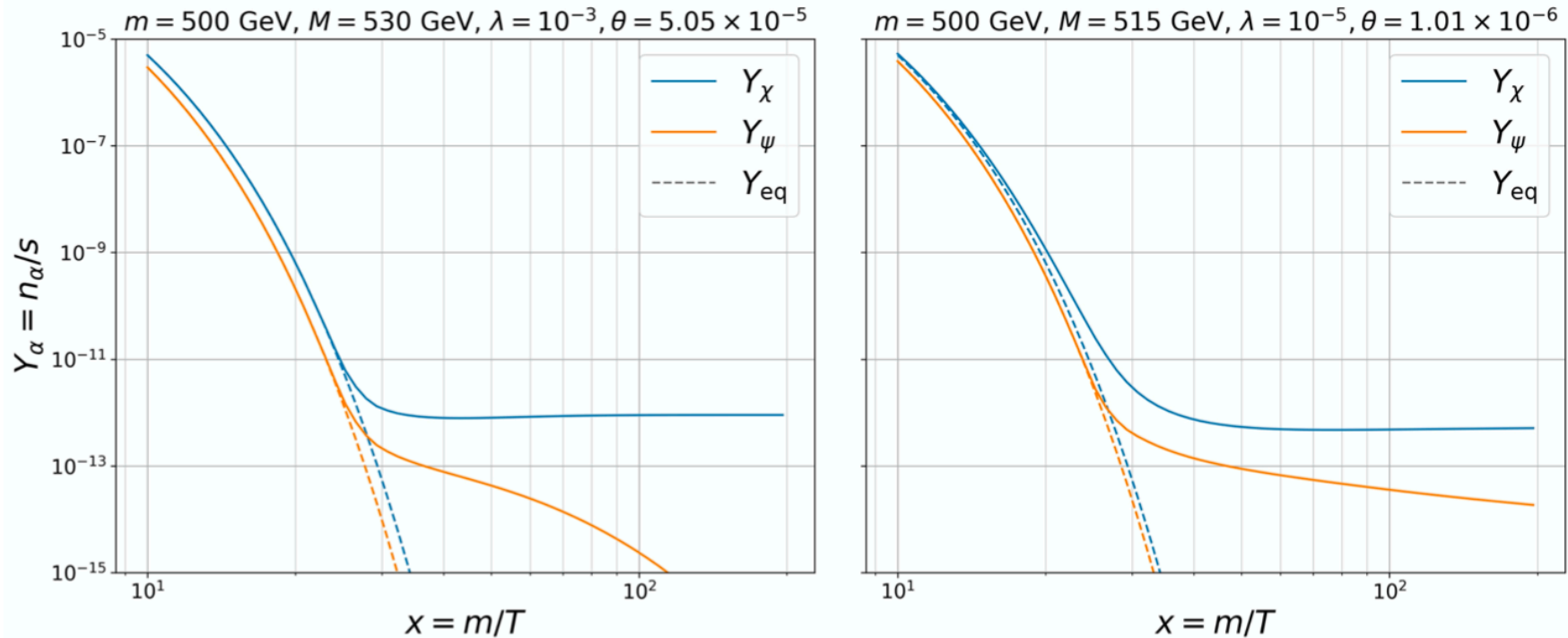
$$\Gamma_{2 \rightarrow 1} = \frac{\sum_{a \in 2} \Gamma_{a \rightarrow 1,0} g_a m_a^2 K_1 \left( \frac{m_a}{T} \right) + \sum_{a \in 1} \Gamma_{a \rightarrow 2,0} g_a m_a^2 K_1 \left( \frac{m_a}{T} \right)}{\sum_{a \in 2} g_a m_a^2 K_2 \left( \frac{m_a}{T} \right)} + \langle \sigma_{2010} v \rangle \bar{n}_0,$$

decays

coscattering

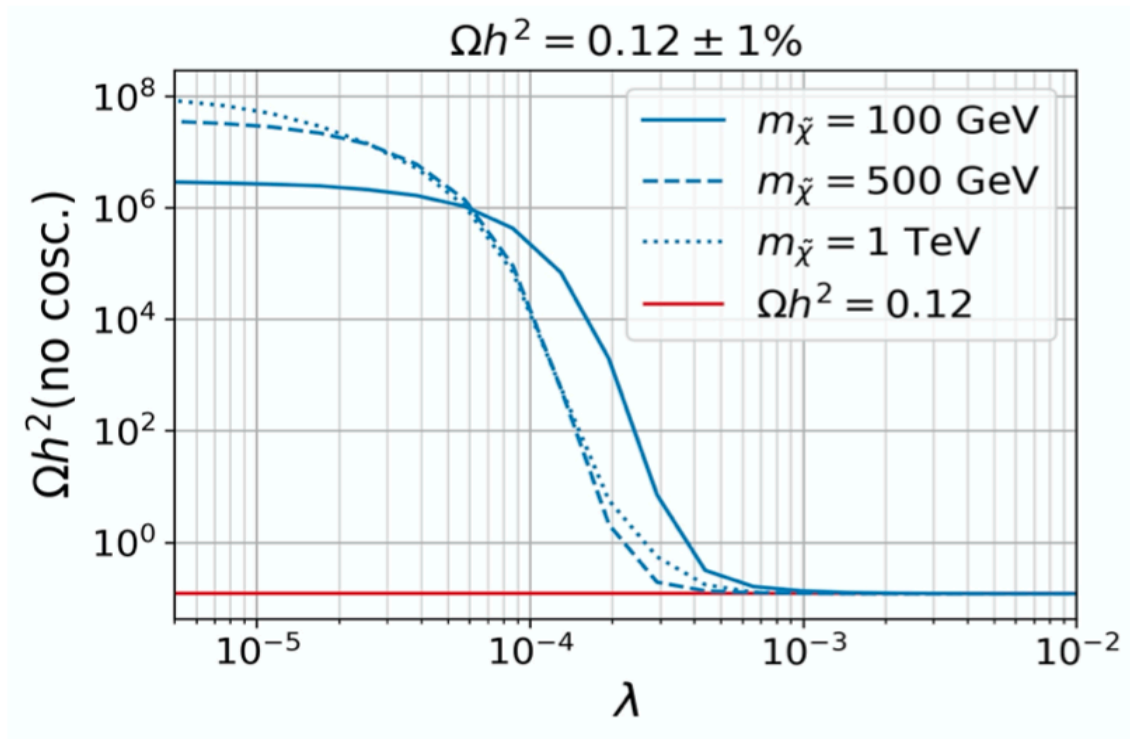


# Abundances



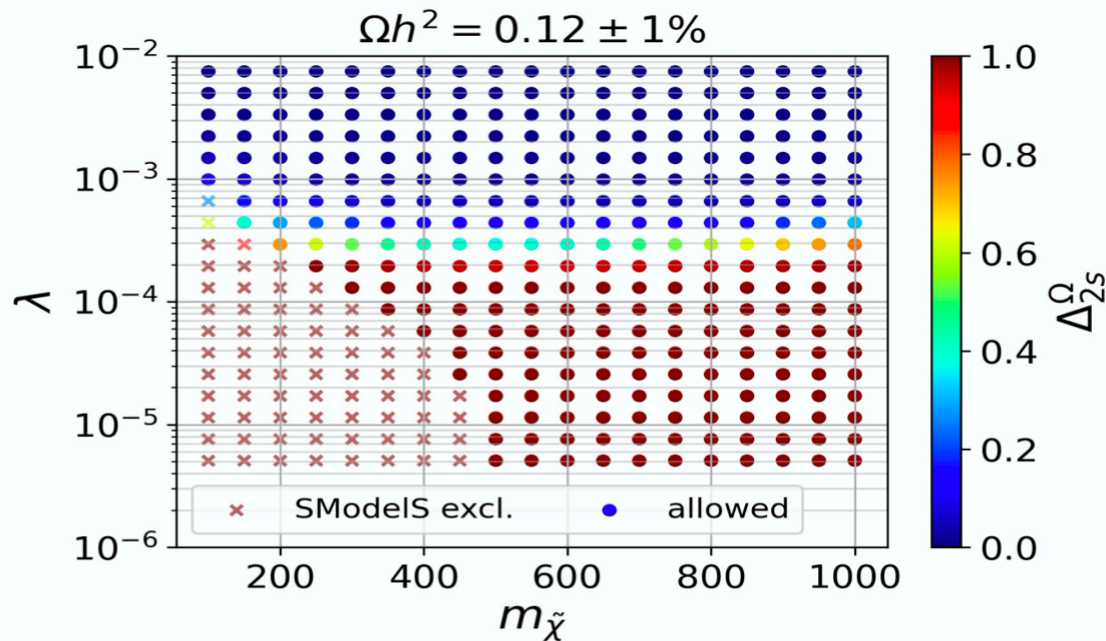
- Solving for relic density: micrOMEGAs5.3 solving momentum integrated Boltzmann eq. including decay/coscattering term
- Small coupling : departure from  $Y_{eq}$

# Importance of co-scattering



- $\Delta m \equiv m_{\tilde{\psi}_0} - m_{\tilde{\chi}}$  adjusted to get  $\Omega h^2 = 0.12$  when all processes included
- Co-scattering dominant at small  $\lambda < \text{few } 10^{-4}$

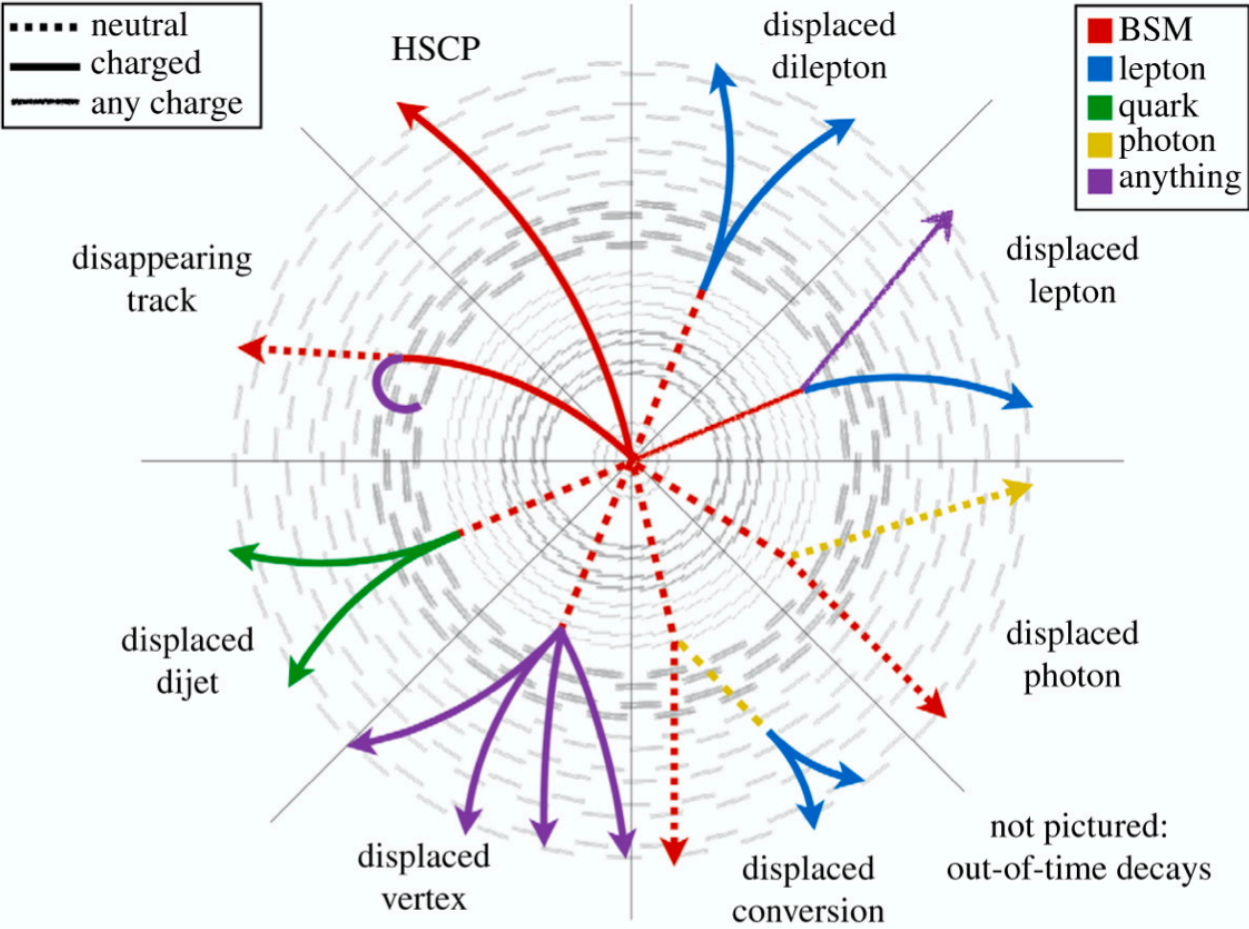
# Importance of co-scattering



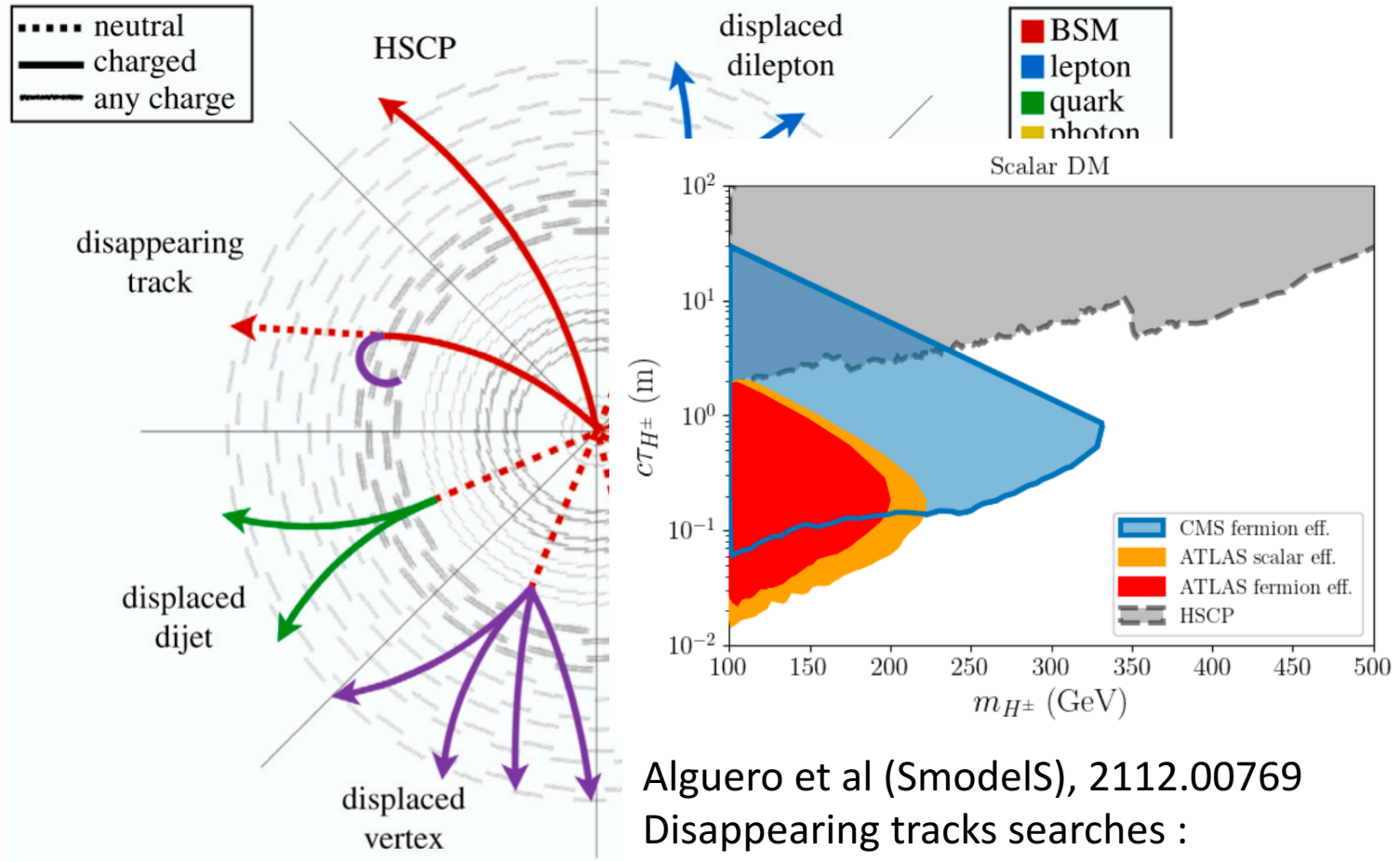
$$\Delta_{2s}^{\Omega} \equiv 1 - \frac{\Omega h^2(2 \text{ sectors})}{\Omega h^2(2 \text{ sectors, no co-scattering})}$$

- $\Delta m \equiv m_{\tilde{\psi}_0} - m_{\tilde{\chi}}$  adjusted to get  $\Omega h^2=0.12$  when all processes included
- Co-scattering dominant at small  $\lambda < \text{few } 10^{-4}$
- $\Delta=0.5$  means  $\Omega h^2$  reduced by factor 2 by co-scattering
- Smodels excl : current disappearing track searches

# LLP at LHC



# LLP at LHC



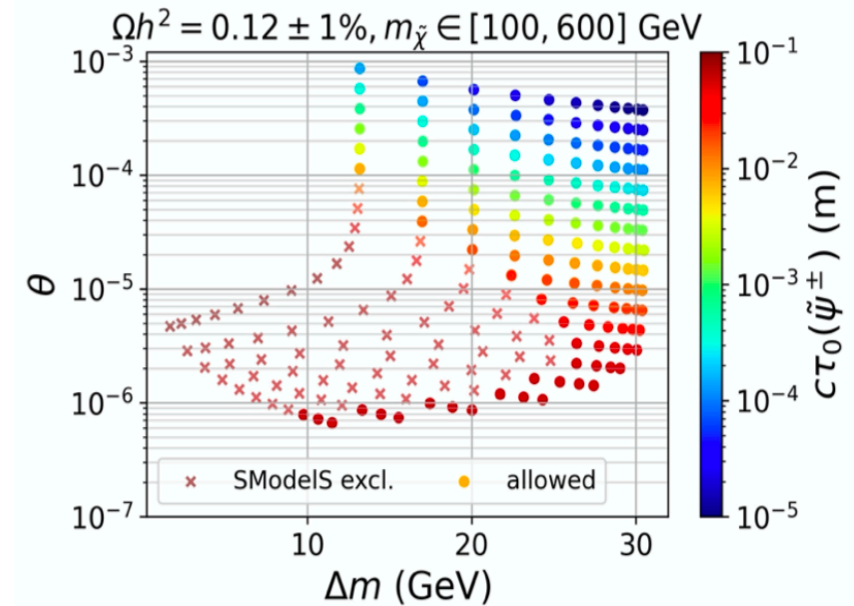
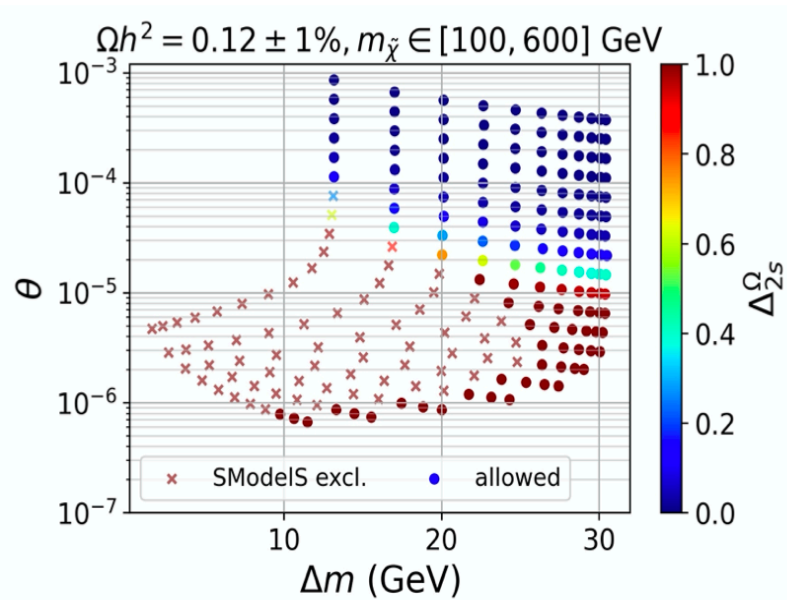
Alguero et al (SmodelS), 2112.00769

Disappearing tracks searches :

ATLAS ( $36\text{fb}^{-1}$ ) , 1712.02118

CMS ( $101\text{fb}^{-1}$ ), 2004.05153

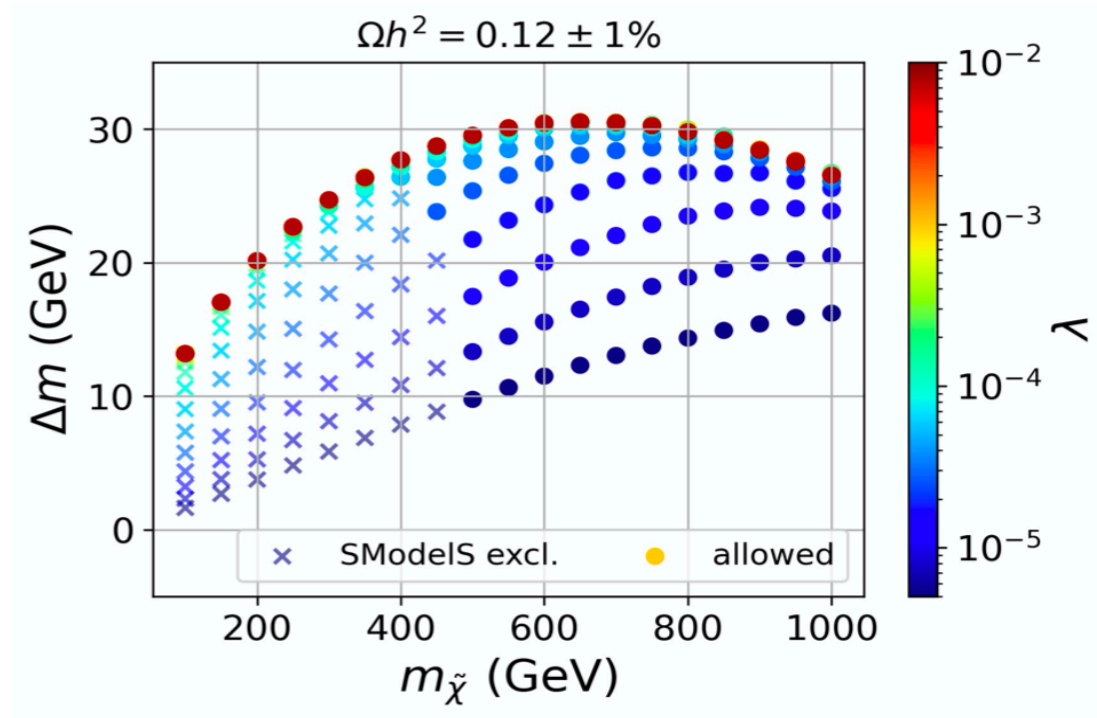
# Coscattering and disappearing tracks



- Smodels excl : current disappearing track searches



# Typical mass splitting



- Singlet –triplet mass splitting : when  $\lambda$  is large ( $10^{-3}$ - $10^{-2}$ ) coannihilation dominates,  $\Delta m \sim 10$ - $30$  GeV (or 13%-3%), smaller mass splitting leads to underabundant DM unless cospattering is included
- When cospattering dominates ( $\lambda < 10^{-4}$ ) mass splitting can be much smaller
- -> at LHC important to cover the region at very small mass splitting

# Remarks

- Departure from equilibrium for very small couplings – loose accuracy using momentum integrated Boltzmann equation
- For very small couplings, early kinetic decoupling, to include these effects must solve full momentum dependent Boltzmann equations

$$\left( \frac{\partial}{\partial t} - H\mathbf{p} \cdot \nabla_{\mathbf{p}} \right) f_{\chi}(p, t) = \frac{1}{E} C[f_{\chi}],$$

- $C(f)$  is collision operator for  $\chi \rightarrow \psi X$

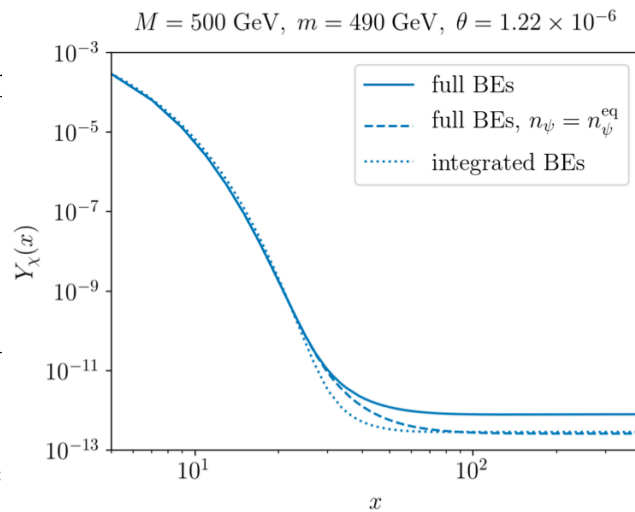
$$\frac{1}{E_{\chi}} C[f_{\chi}] = \sum_{\psi, X, X'} \left( f_{\chi}^{\text{eq}} \frac{n_{\psi}}{n_{\psi}^{\text{eq}}} - f_{\chi} \right) \frac{1}{2E_{\chi}} \int \widetilde{d}p_X \widetilde{d}p_{X'} \widetilde{d}p_{\psi} \delta^{(4)} \left( \sum p_i \right) |\overline{M}|^2 f_X^{\text{eq}}$$

- In the singlet-triplet model, F. Brummer arXiv:1910.01549

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- For very small couplings must solve full Boltzmann equations



ling, to include these effects in equations

- $C(f)$  is collision

$$\frac{1}{E_X} C[f_X] =$$

$$\int \widetilde{p}_X \widetilde{d}p_\psi \delta^{(4)}\left(\sum p_i\right) |\overline{M}|^2 f_X^{\text{eq}}$$

- In the singlet-triplet model, F. Brummer arXiv:1910.01549
- For small couplings correction can be 100%

# Conclusion

- Co-scattering can drive DM formation when DM is very weakly coupled to the SM and there exist other dark states with small mass splitting with DM
- General solution of relic abundance implemented in micrOMEGAs assuming kinetic equilibrium
- To probe cospattering mechanism at colliders important to cover the region with very small mass splitting - in conventional coannihilation this region corresponds to underabundant DM.
- Under investigation : full solution to Boltzmann equation (GB, Chakraborti, Hryczuk et al)