

# Data Formats for Gamma-Ray Astronomy (GADF)

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### Overview

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M. Linhoff | 2022-11-22

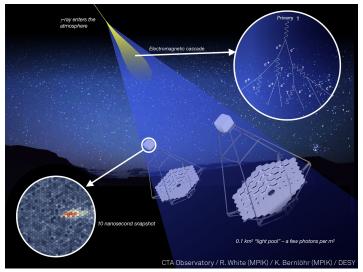


# The Gamma-ray Astronomy Data Formats Initiative (GADF)

- Ground-based gamma-ray observatories are mostly run by particle physicists as closed collaborations
  - No or very limited public data access
  - Mostly proprietary analysis software
  - Custom data formats, usually strongly coupled with the proprietary software
  - Mostly based on CERN's ROOT
- Satellites and the Cherenkov Telescope Array (will) operate as open observatories
- The current generation of telescopes want a better archival format
- Strong scientific use-cases for multi-instrument analysis
- ⇒ Effort for a common, software-independent data format
- ⇒ Development of Open Source analysis software ("science tools")

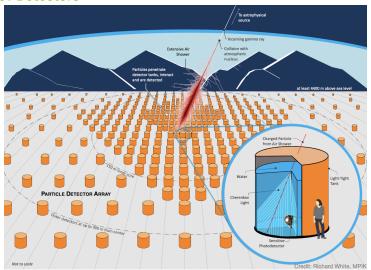


# Imaging Air Cherenkov Telescopes

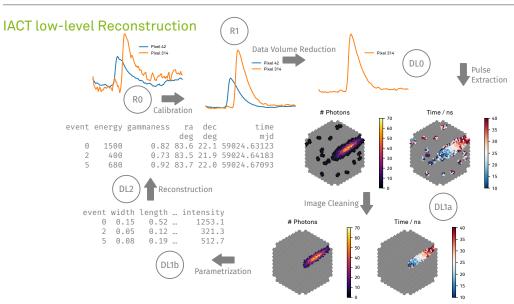




#### Water Cherenkov Detectors







## Gamma-ray Astronomy High Level Analysis

- We need to go back from the estimated quantities in the event lists to physically meaningful quantities
- Energy spectra, Light Curves, Flux Sky Maps, ...
- For this, we need the instrument response to solve the inverse problem, in general:

$$\underbrace{g(\hat{E},\hat{\alpha},\hat{\delta})}_{\text{Observed distribution}} = \underbrace{\prod_{R(\hat{E},\hat{\alpha},\hat{\delta}|E,\alpha,\delta)}^{\text{Instrument Response}} \cdot \underbrace{f(E,\alpha,\delta)}_{\text{True }\gamma\text{-Signal}} \, \text{d}E \, \text{d}\Omega + \underbrace{b(\hat{E},\hat{\alpha},\hat{\delta})}_{\text{Background}}$$

■ *R* is usually factorized into three independent components:

$$R(\hat{E}, \hat{\alpha}, \hat{\delta} | E, \alpha, \delta) = A_{\text{eff}}(E, \alpha, \delta) \cdot M(\hat{E} | E, \alpha, \delta) \cdot PSF(\hat{\alpha}, \hat{\delta} | E, \alpha, \delta)$$

 Current GADF version is mainly focused on specifying a file format for event lists and instrument response functions



# A (short History of GADF)

- 2011 Prototypes for the CTA data format and science tools
- 2016 Establishment of the Gamma-ray Astronomy Data Formats (GADF) initiative
  - First preliminary release version (0.1), mainly focused on IACTs
- 2018 Version 0.2 released
  - Support implemented in the science tools Gammapy and ctools
  - H.E.S.S. releases ≈ 50 h of observations of different sources using the format [doi:10.48550/arXiv.1810.04516]



## A (short History of GADF)

#### 2019

■ FACT, Fermi-LAT, H.E.S.S. MAGIC and VERITAS observations of the Crab Nebula are used to perform the first multi-instrument analysis [doi:10.1051/0004-6361/201834938]

https://github.com/open-gamma-ray-astro/joint-crab

- **ctools** based analysis of the H.E.S.S. data release [doi:10.1051/0004-6361/201936010]
- Comparison of Gammapy and ctools using the H.E.S.S. data release [doi:10.1051/0004-6361/201936452]

#### 2022

- Steering committee formed (GADF → VODF, see next talk)
- Version 0.3 released with support for wide-field instruments like HAWC
- Comparison of HAWC analysis using Gammapy + GADF data and HAWC proprietary software [doi:10.1051/0004-6361/202243527]



# **Current GADF Specifications**

- GADF specifies a common file format for high-level gamma-ray data
- FITS with binary tables for event lists, instrument response functions, pointing information, ...
- Based on the Fermi-LAT format and the OGIP standard from X-ray astronomy
- Developed on GithHub at https://github.com/open-gamma-ray-astro/gamma-astro-data-formats
- Published at https://gamma-astro-data-formats.readthedocs.io
- Also published and citable via Zenodo: [doi:10.5281/zenodo.1409830]



#### Hand-Over to VODF

- GADF had no clear organizational structure
  - No clear resolution of contentious issues
  - No clear roadmap
- Forming of a steering committee with one representative per participating instrument in 2022
- Inclusion of Neutrino observatories → renaming to VODF
- Release 0.3 of GADF planned to be last release before first VODF release



## Some questions discussed but not resolved in GADF

- Which standard takes precedence? Old OGIP vs. current FITS
- Current standard is often very vague instead of descriptive
- Additional IRF parametrizations
- Uncertainties / validities of IRFs
- Multiple IRFs for one observations
- Different IRFs for different event categories ("event types")
- Additional specifications for simulated datasets
- Interoperability with other entities, especially the VO