

# Constraining constant and tomographic coupled dark energy with low- and high-redshift probes

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  - Dark energy takes on the form of a scalar field  $\phi$
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- Here we propose a form of parametrisation for  $\beta$ , where it can **vary with redshift**:

$$\beta(z) = \frac{\beta_1 + \beta_n}{2} + \frac{1}{2} \sum_{i=1}^{n-1} (\beta_{i+1} - \beta_i) \tanh[s_i(z - z_i)]$$

where  $\beta_i$  is the amplitude of coupling in each bin,  $s_i$  is the smoothing factor and  $z_i$  is the value of the bin edge

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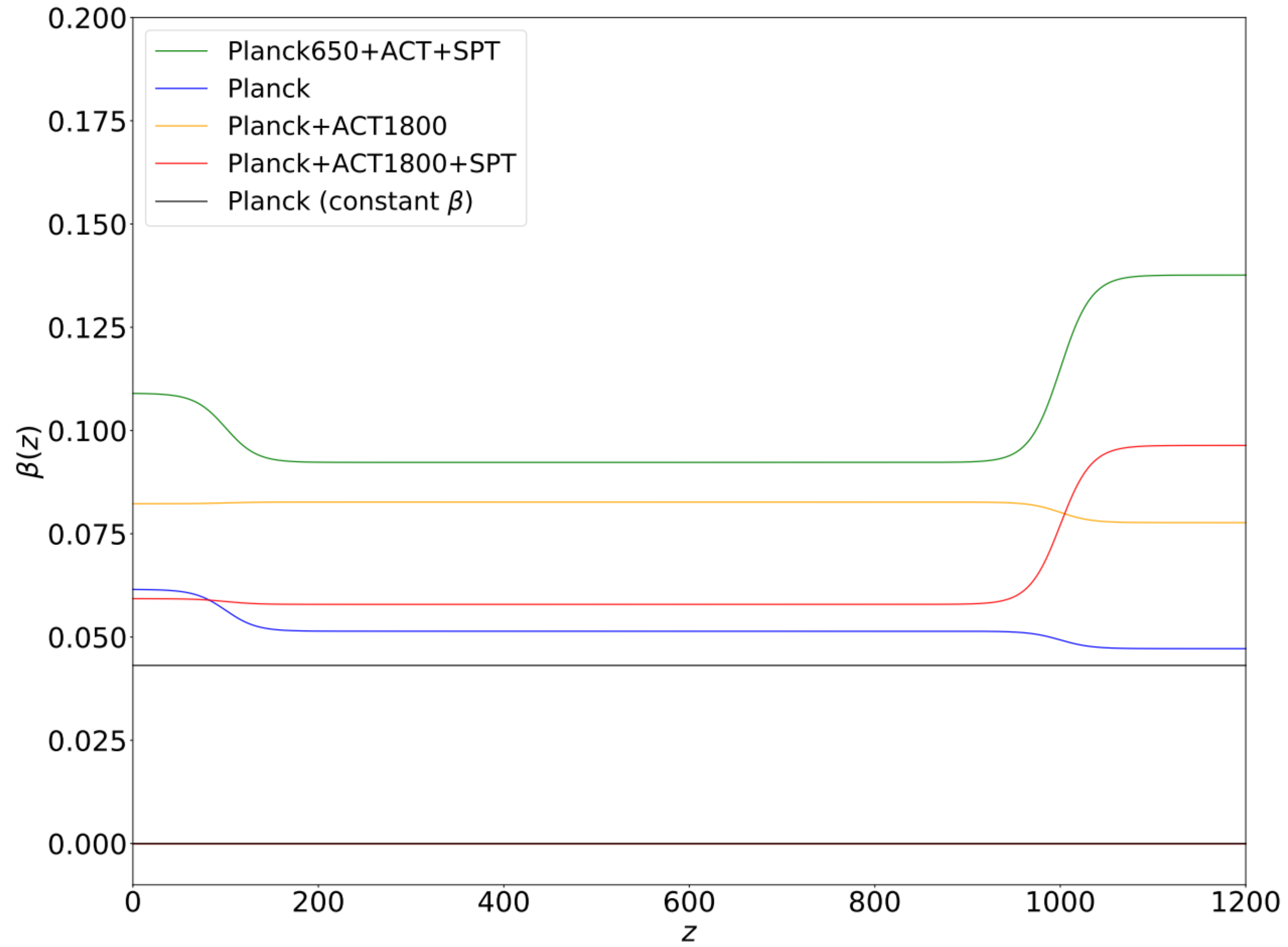
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- We use an array of **high redshift** (CMB) and **low redshift** (BAO, SNeIa, cosmic chronometers, RSD, SHoES) data, as well as **Weak Lensing** and **Galaxy Clustering** to constrain our model

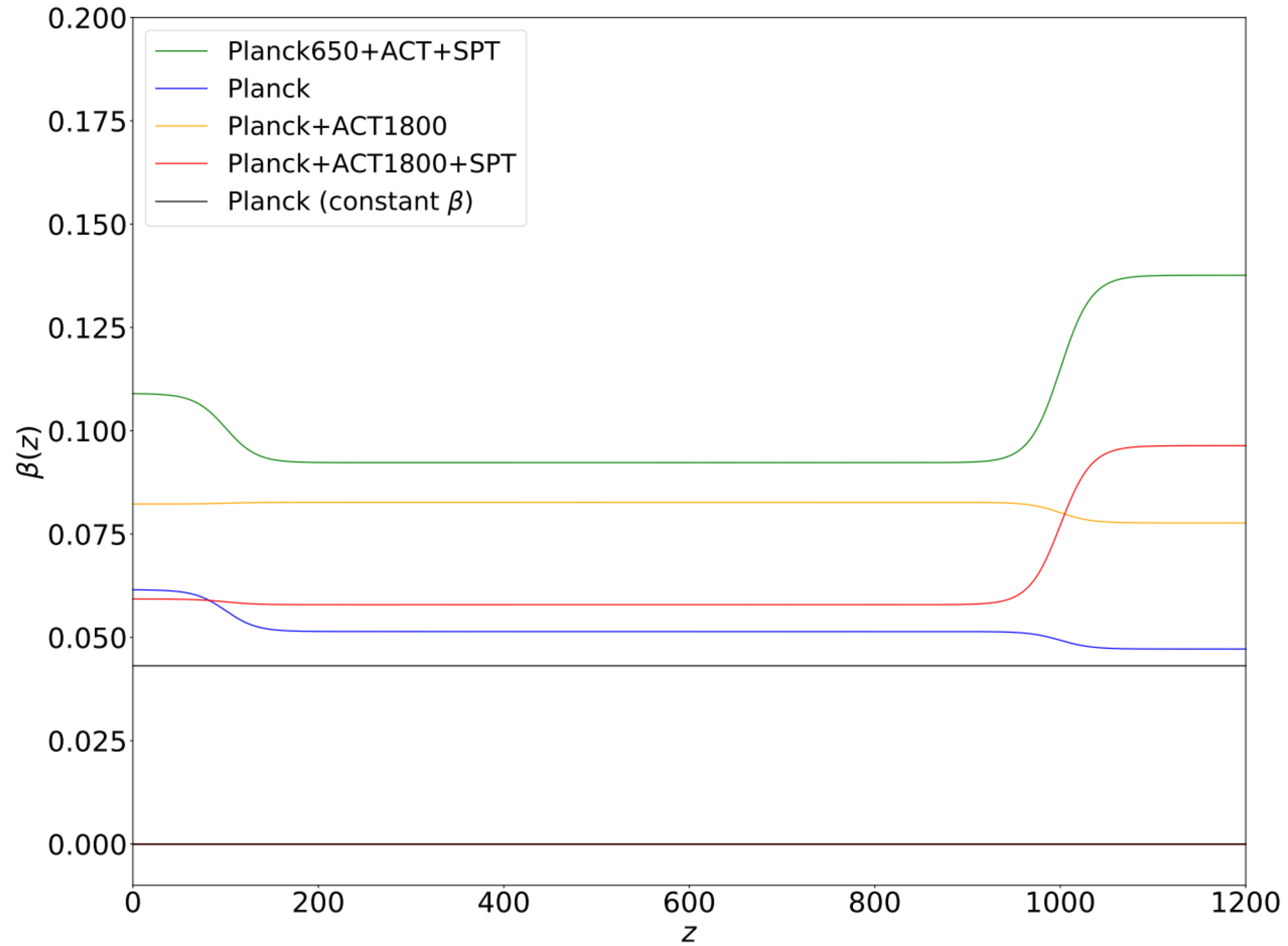
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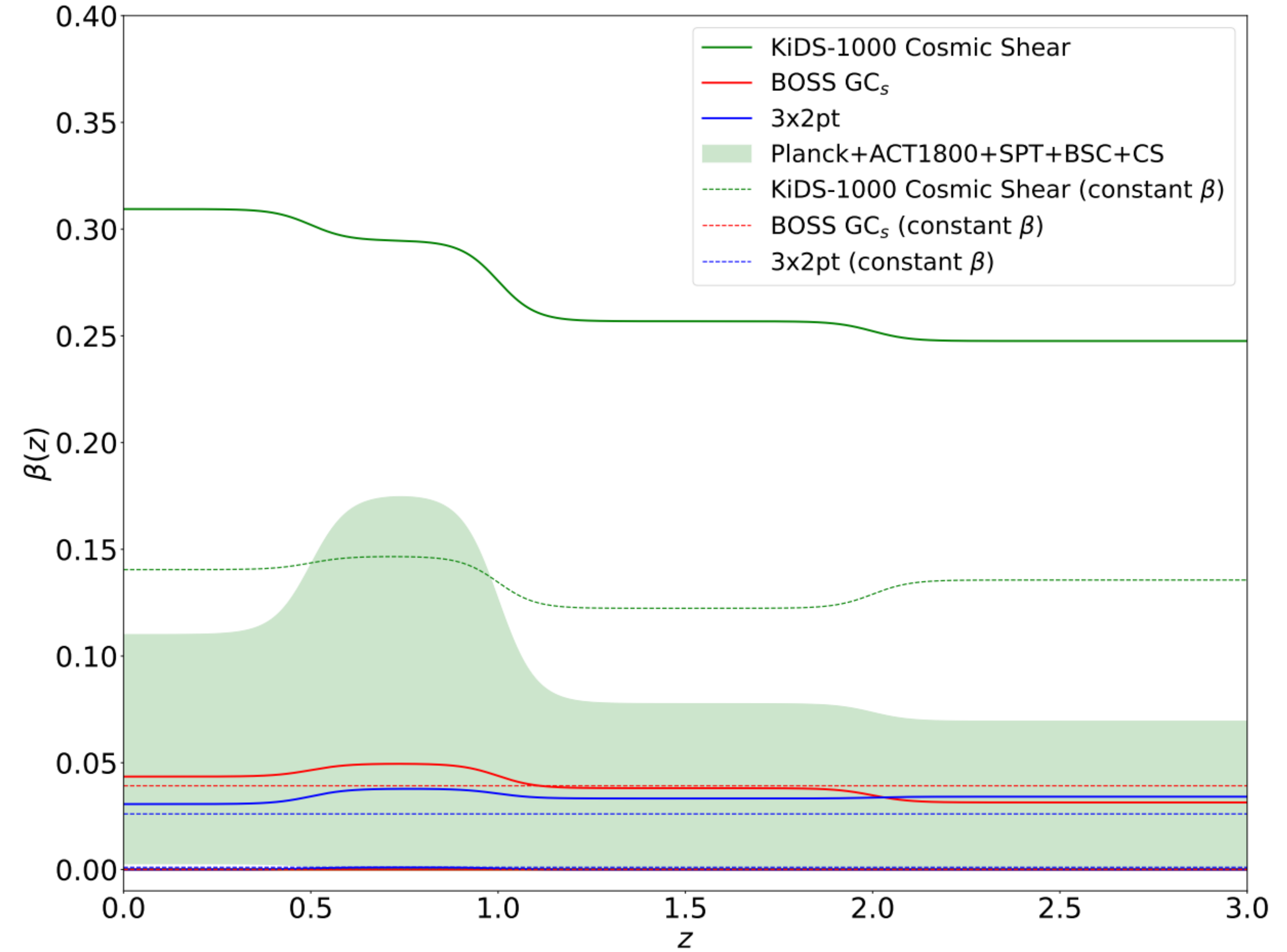
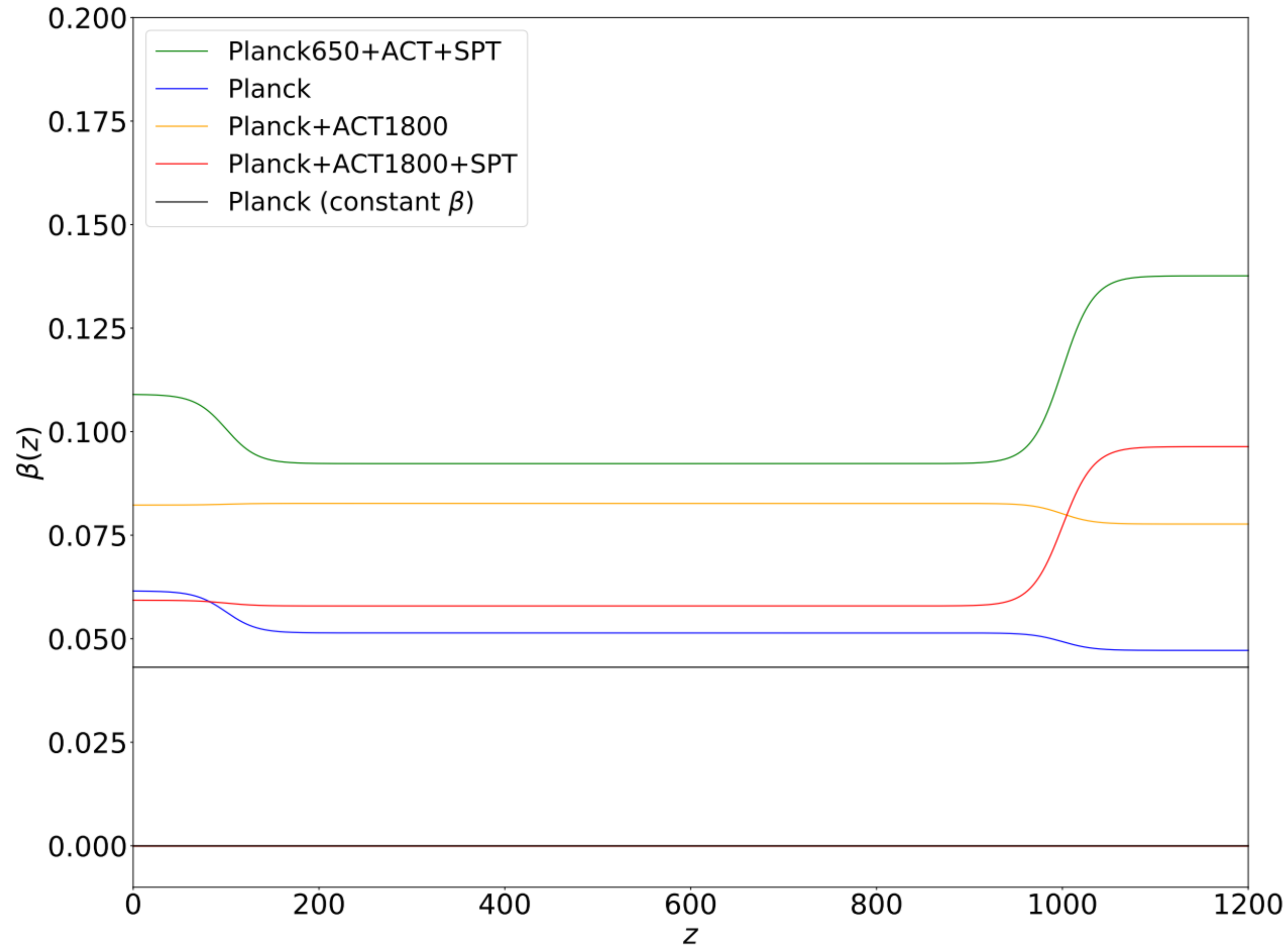


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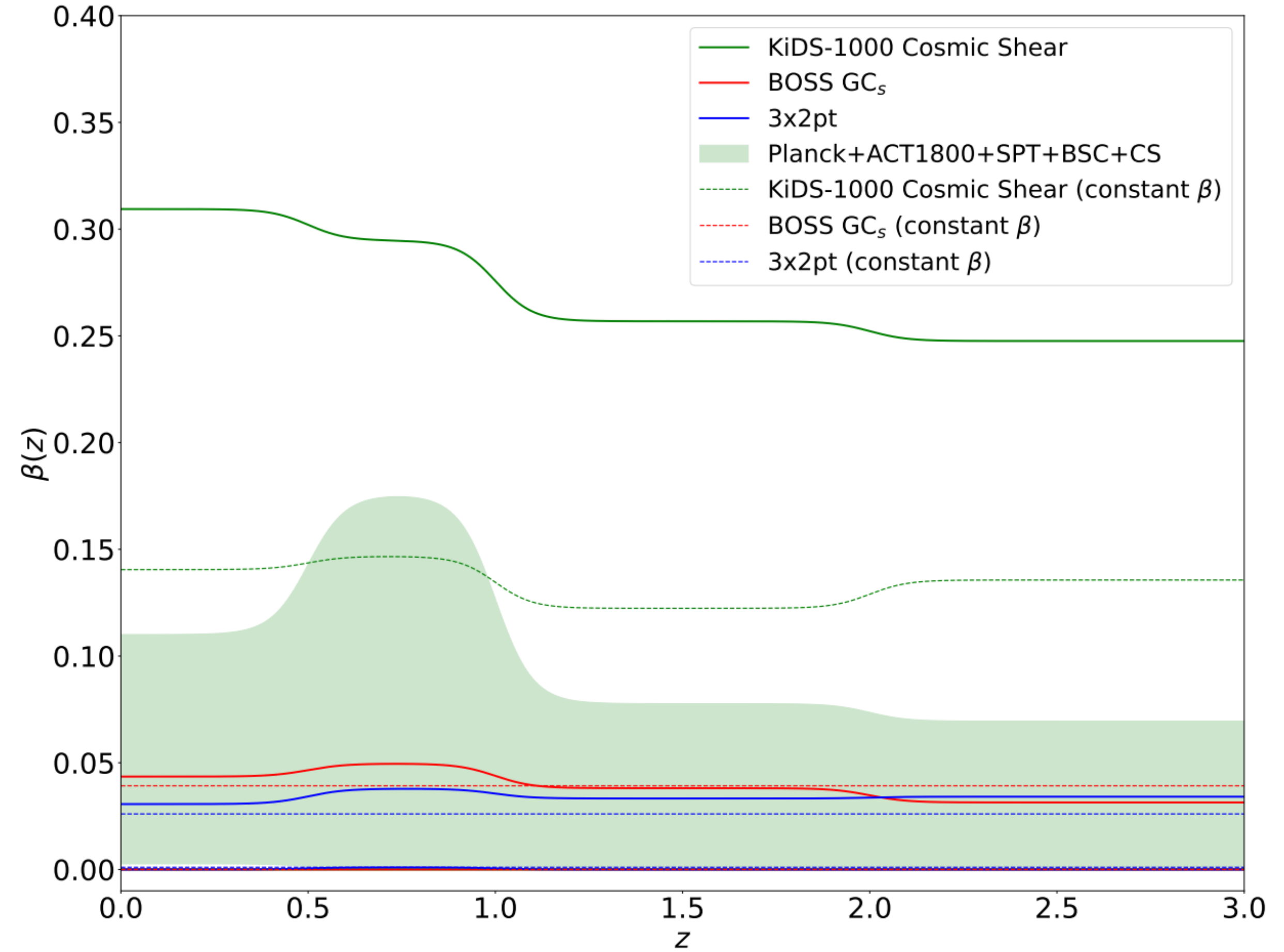
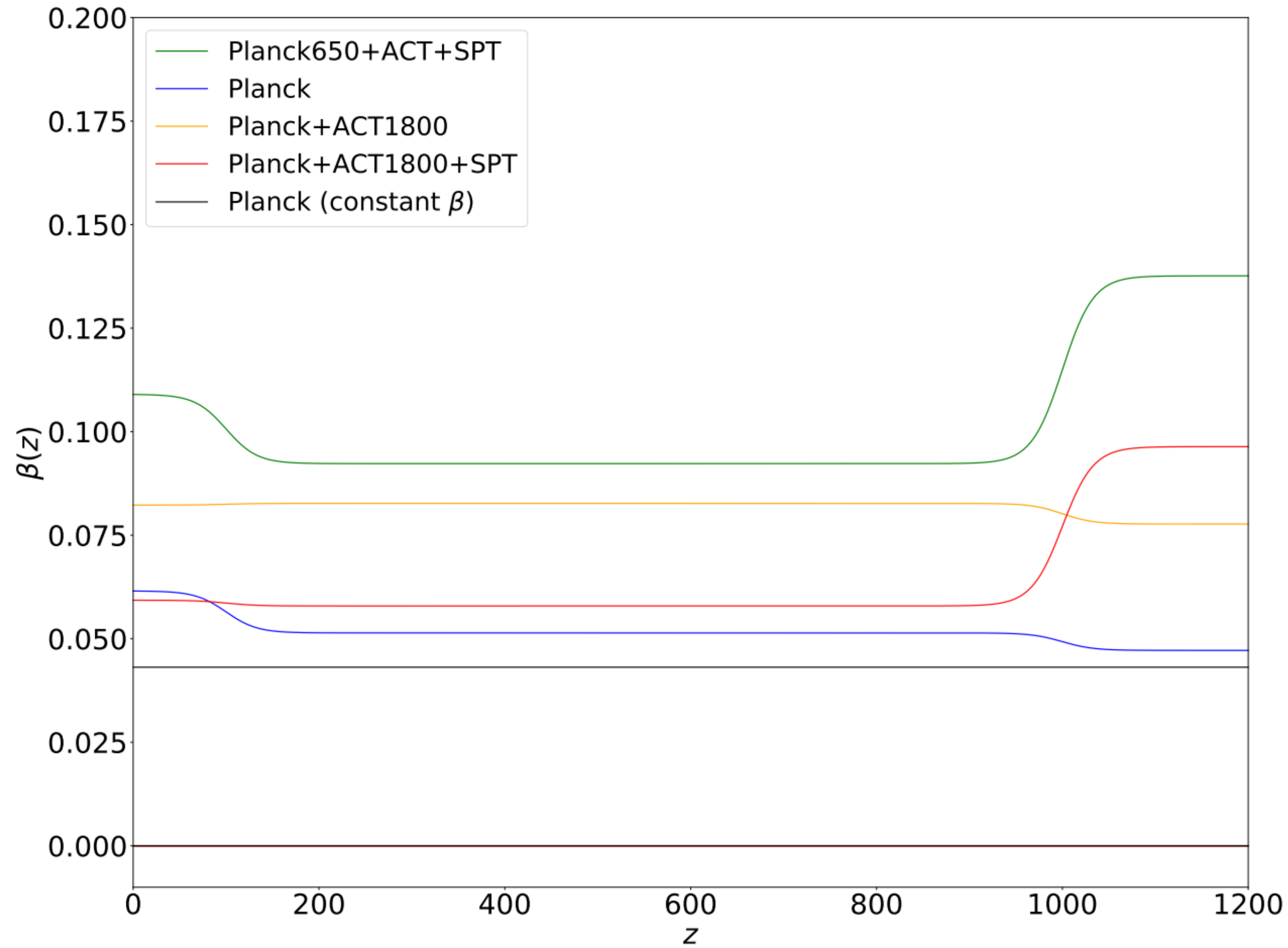
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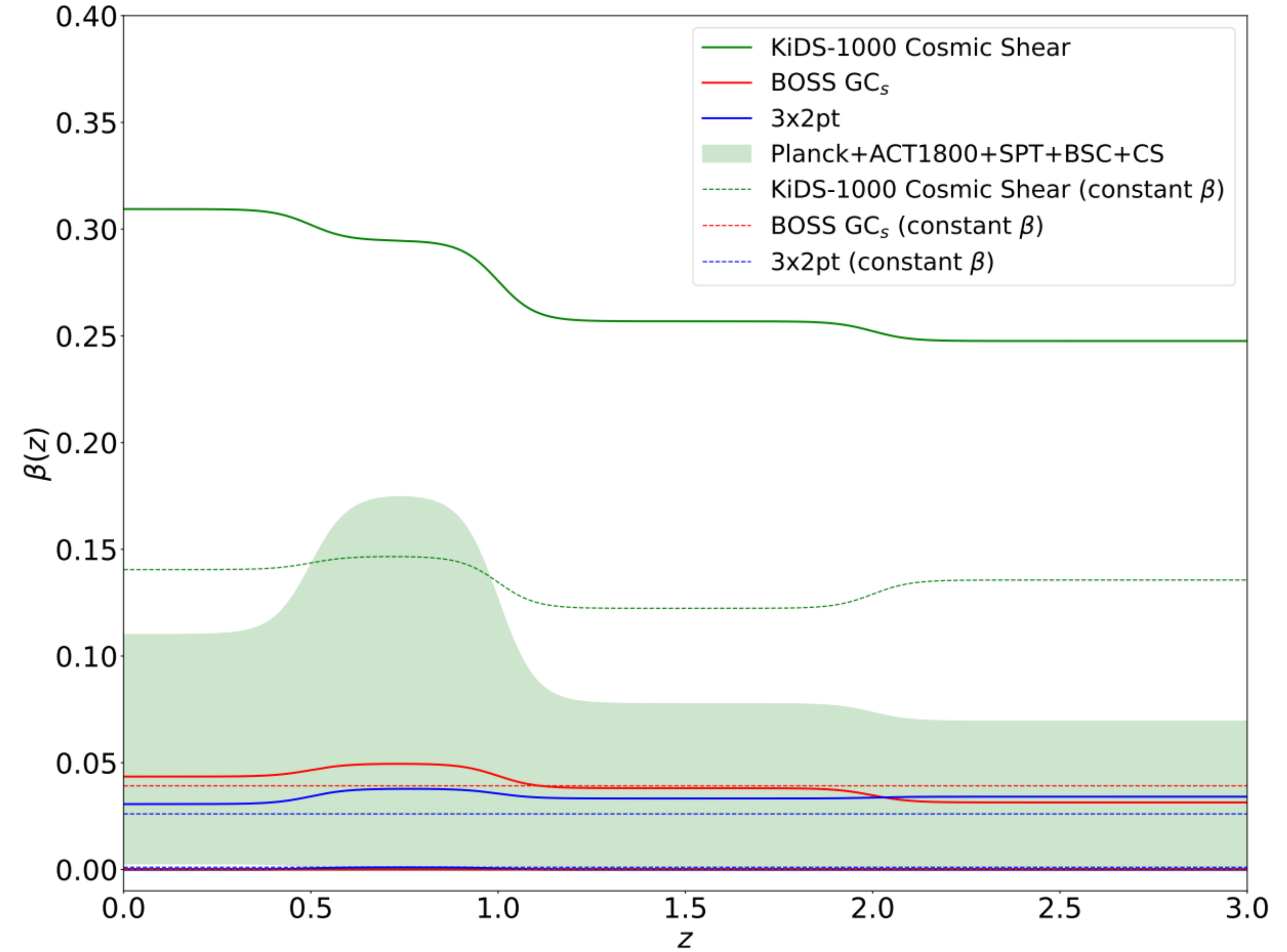
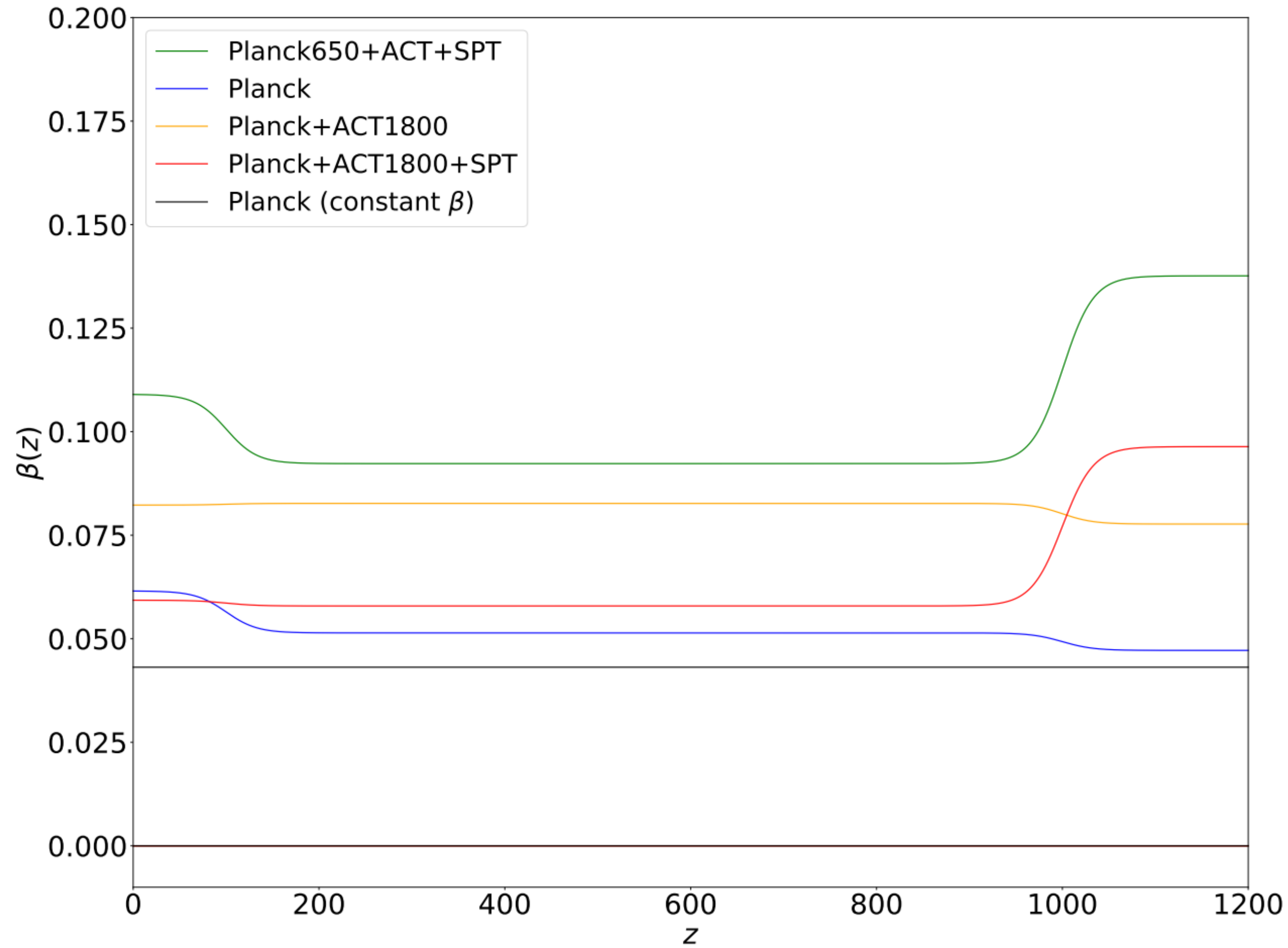
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- GC and 3x2pt: constraining power **comparable with CMB!**
- In a tomographic CDE framework, tension in  $S_8$  (between 3x2pt and CMB) reduces from  $\sim 2.8\sigma$  to  $\sim 1.3\sigma$

**Thank you!**