

université
PARIS-SACLAY

IJClab
Irène Joliot-Curie
Laboratoire de Physique
des 2 Infinis



A2C Astroparticles, Astrophysics
& Cosmology

Early Dark Energy as a solution to the Hubble tension

Adrien La Posta
IJClab

[link](#)

laposta@ijclab.in2p3.fr

CosPT meeting - 06/12/22

FLRW metric

$$ds^2 = -c^2 dt^2 + a^2(t) \left(\frac{dr^2}{1 - Kr^2} + r^2 d\Omega^2 \right)$$

Friedmann equation

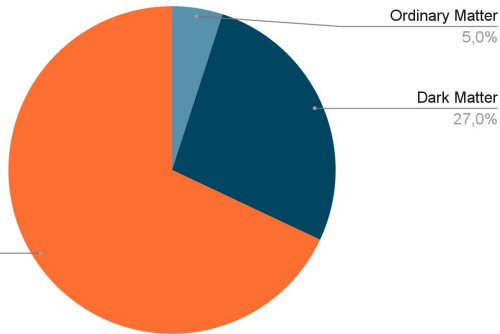
$$H^2(z) = \left(\frac{\dot{a}}{a} \right)^2 = \frac{8\pi G}{3} [(\rho_b^0 + \rho_c^0)(1+z)^3 + \rho_r^0(1+z)^4 + \rho_\Lambda]$$

baryon

CDM

radiation

dark energy



Dark Energy
68,0%

Ordinary Matter
5,0%

Dark Matter
27,0%

10⁻³² seconds

1 second

100 seconds

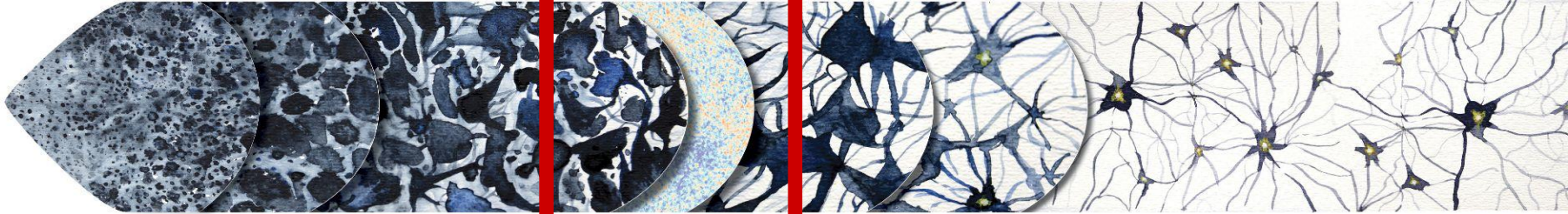
380 000 years

300–500 million years

Billions of years

13.8 billion years

Beginning of the Universe



Inflation

Accelerated expansion of the Universe

Formation of light and matter

Light and matter are coupled

Dark matter evolves independently; it starts clumping and forming a web of structures

Light and matter separate

- Protons and electrons form atoms
- Light starts travelling freely; it will become the Cosmic Microwave Background (CMB)

Dark ages

Atoms start feeling the gravity of the web of dark matter

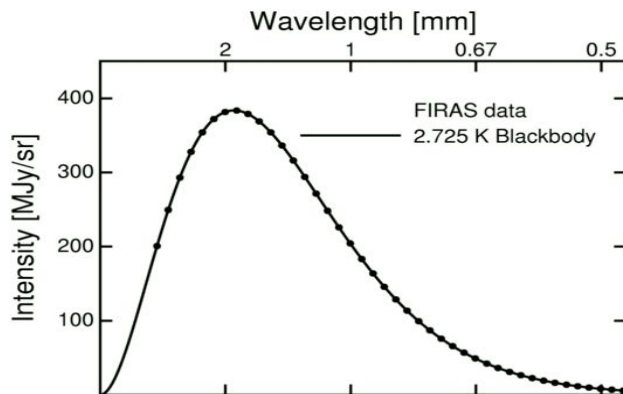
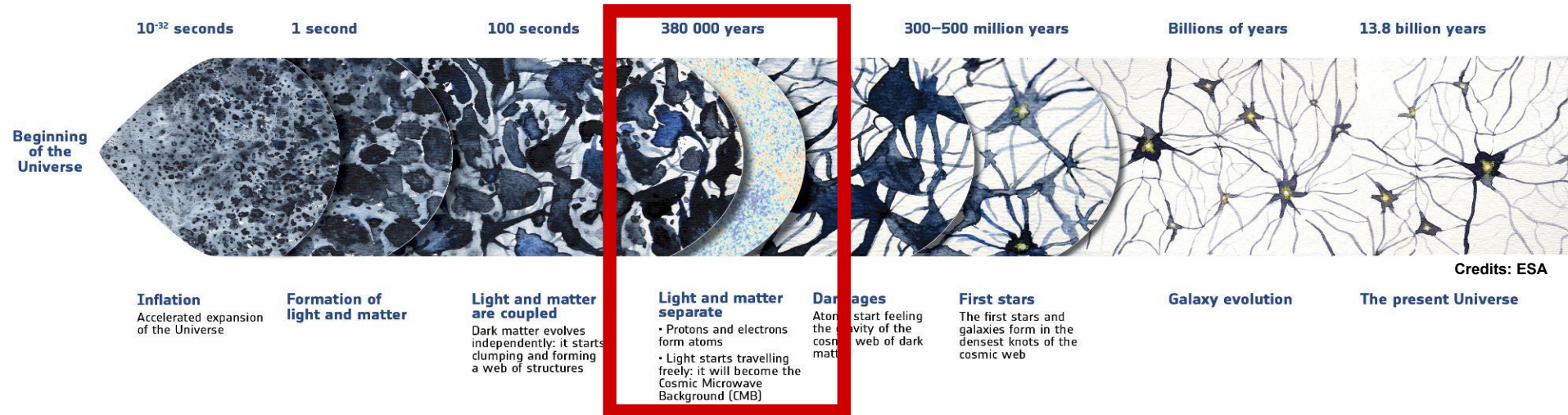
First stars

The first stars and galaxies form in the densest knots of the cosmic web

Galaxy evolution

The present Universe

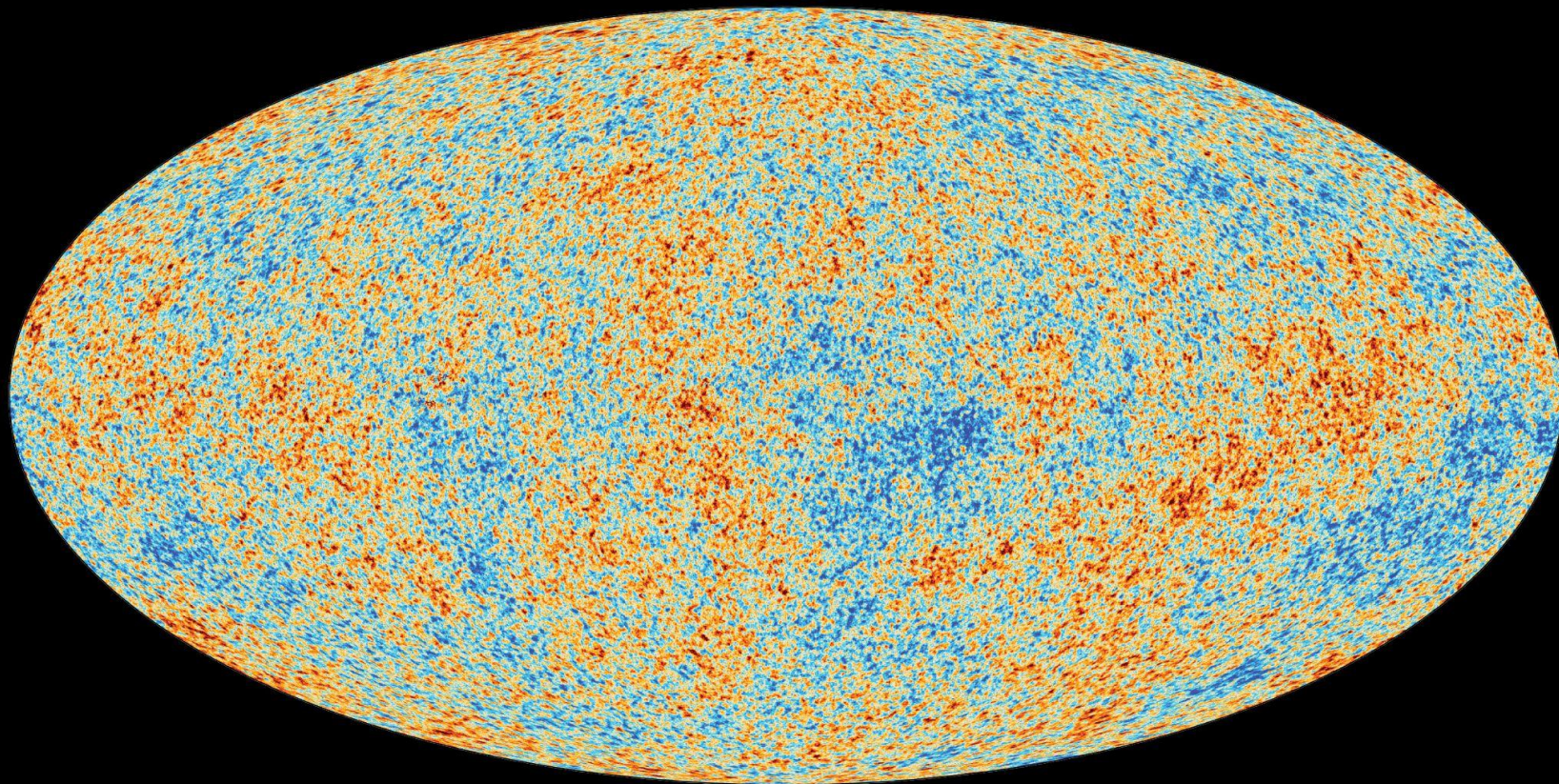
Credits: ESA



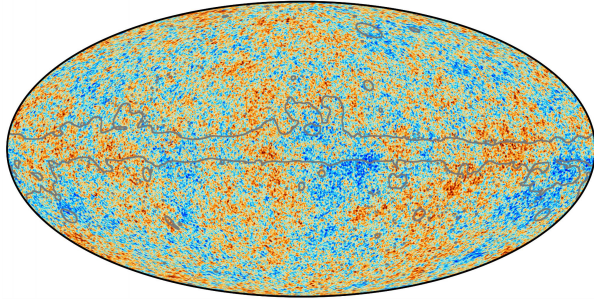
Nearly isotropic blackbody spectrum at $T = 2.725$ K

$$\frac{\delta T}{T} \sim 10^{-5}$$

**CMB temperature as measured by the
Planck satellite**

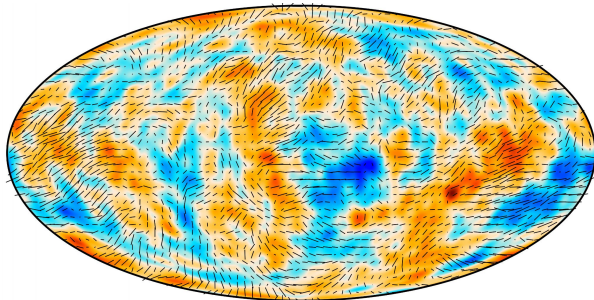


Temperature



-300 300 μK

Polarization E-modes



0.41 μK -160 160 μK

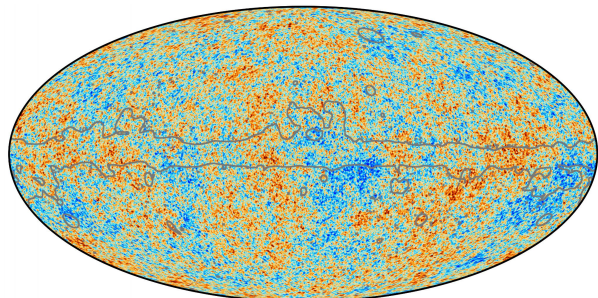
Spherical harmonics

$$\delta T(\hat{n}) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} a_{\ell m}^T Y_{\ell}^m(\theta, \phi)$$

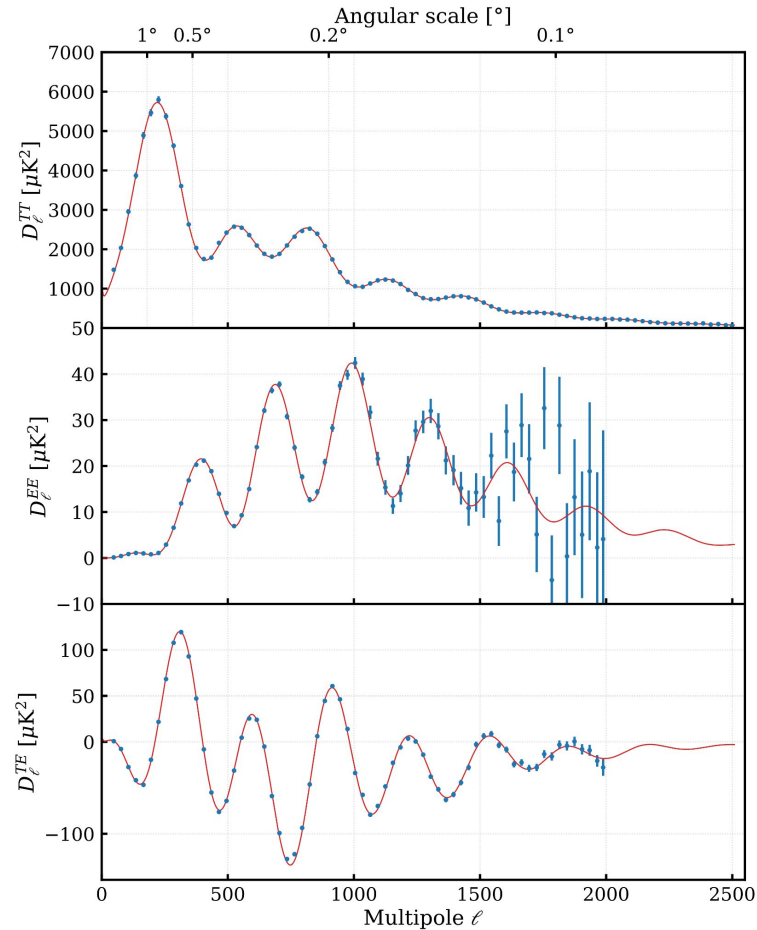
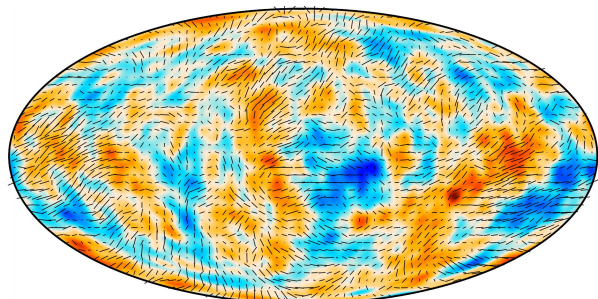
$$\langle a_{\ell m}^T a_{\ell' m'}^{T*} \rangle = \delta_{\ell\ell'} \delta_{mm'} C_{\ell}^{TT}$$

Measuring H_0 from the CMB

Temperature

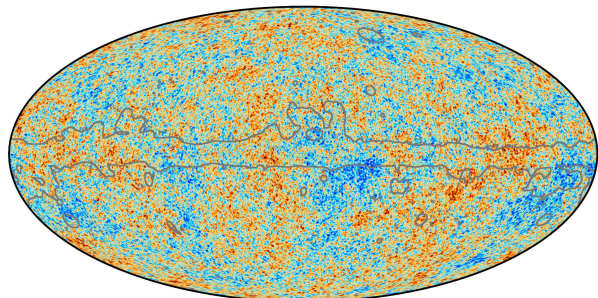


Polarization E-modes



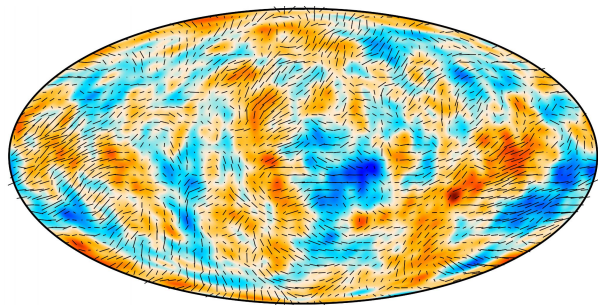
Measuring H_0 from the CMB

Temperature

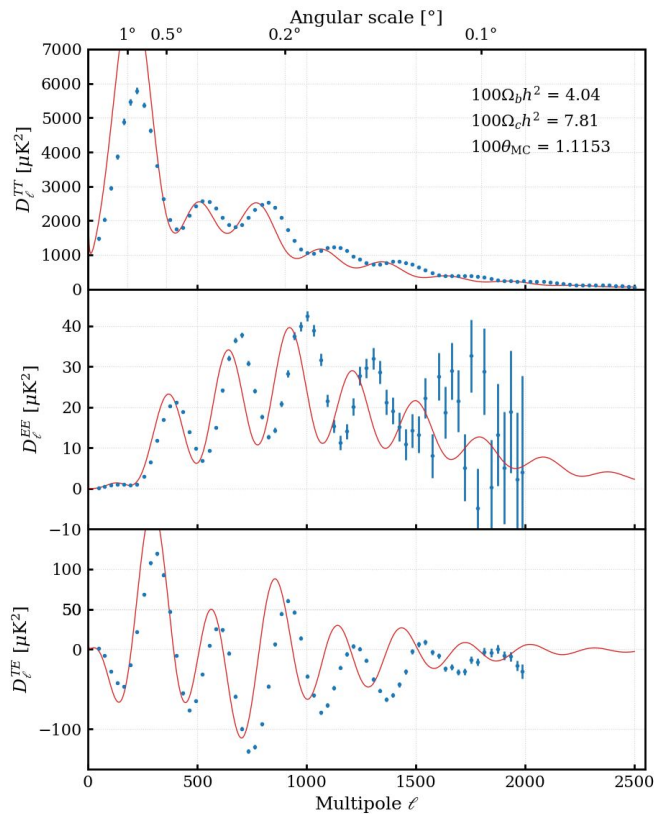


-300 300 μK

Polarization E-modes



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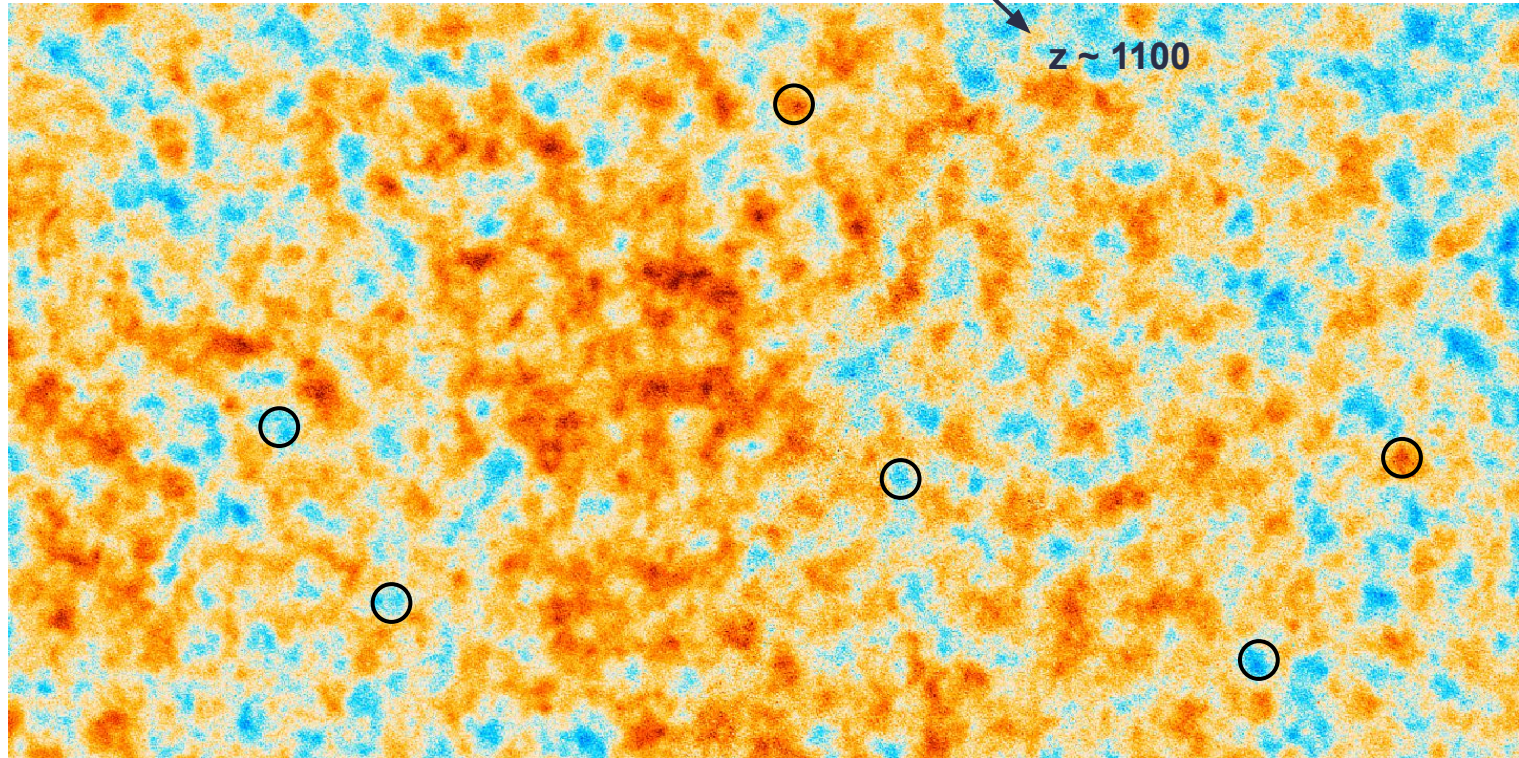


$\rightarrow \theta_* \rho_b^0 \rho_c^0$

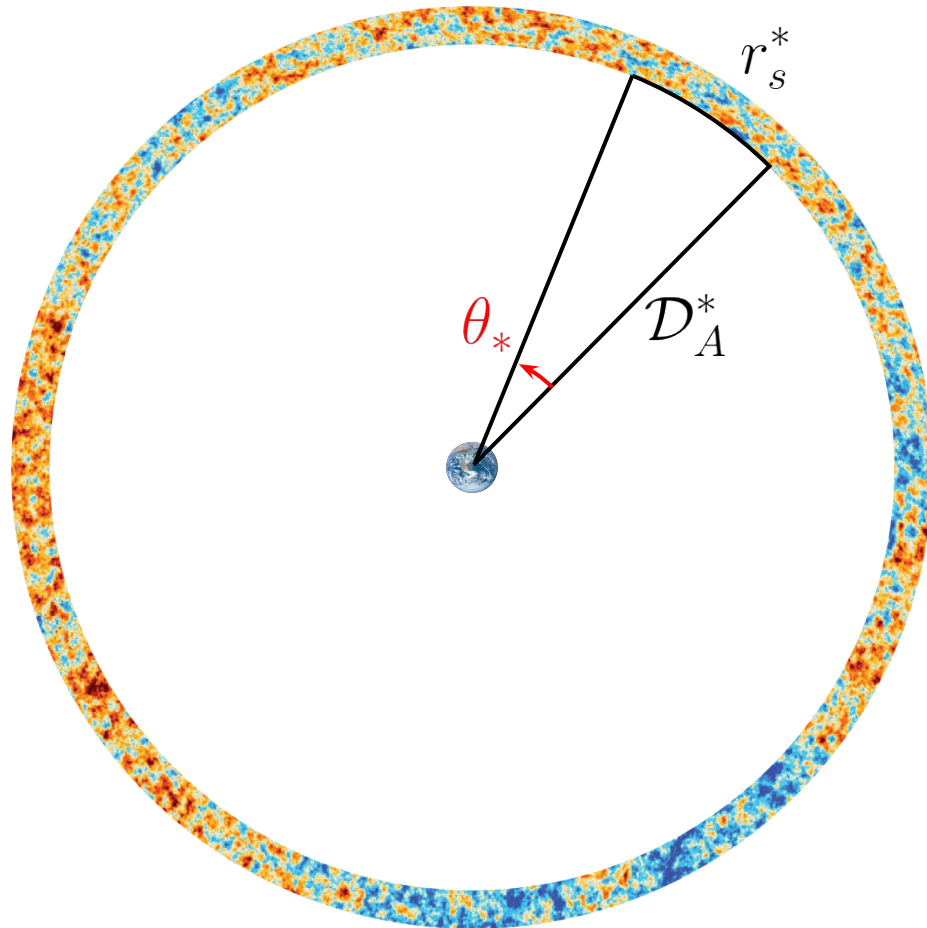
CMB standard ruler : size of the sound horizon at decoupling imprinted in the CMB radiation

Measuring H_0 from the CMB

CMB standard ruler : size of the sound horizon at decoupling imprinted in the CMB radiation

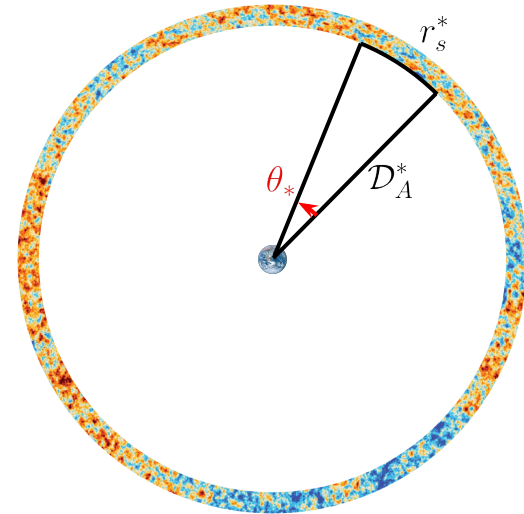


How to measure H_0 from the CMB ?



How to measure H_0 from the CMB ?

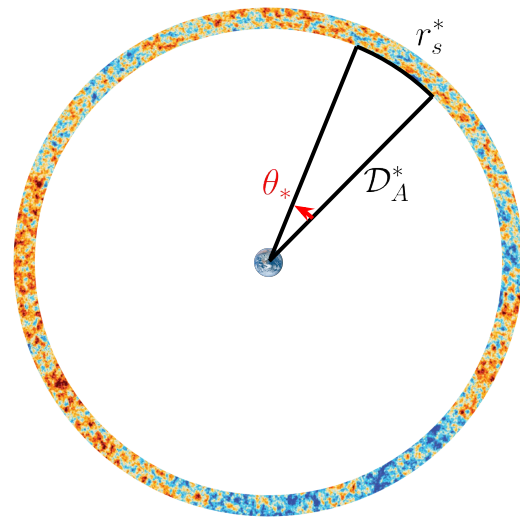
$$\theta_* = \frac{r_s^*}{\mathcal{D}_A^*} \longrightarrow r_s^* = \int_{z^*}^{\infty} \frac{dz}{H(z)} c_s(z)$$



How to measure H_0 from the CMB ?

$$\theta_* = \frac{r_s^*}{\mathcal{D}_A^*} \longrightarrow r_s^* = \int_{z^*}^{\infty} \frac{dz}{H(z)} c_s(z)$$
$$c_s(z) = c \sqrt{\frac{1}{3 [1 + 3\rho_b^0/4\rho_\gamma^0(1+z)^{-1}]}}$$

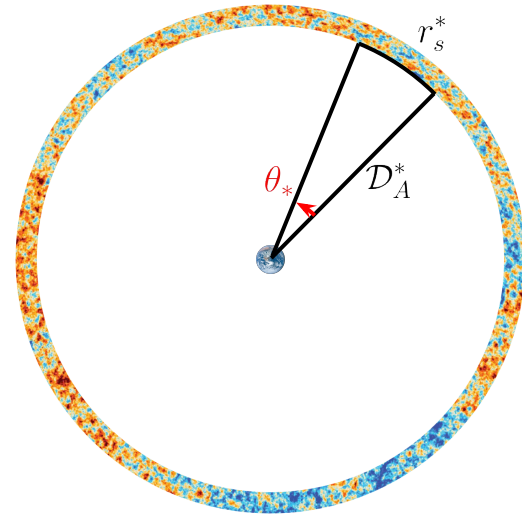
$$H_{\text{early}}^2(z) = \frac{8\pi G}{3} [\rho_r^0(1+z)^4 + (\rho_b^0 + \rho_c^0)(1+z)^3]$$



How to measure H_0 from the CMB ?

Now \mathcal{D}_A^* is known

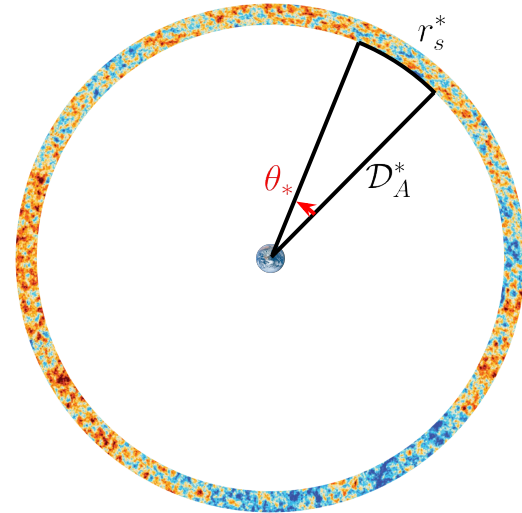
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How to measure H_0 from the CMB ?

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$$\theta_* = \frac{r_s^*}{\mathcal{D}_A^*} \rightarrow \mathcal{D}_A^* = c \int_0^{z^*} \frac{dz}{H(z)}$$

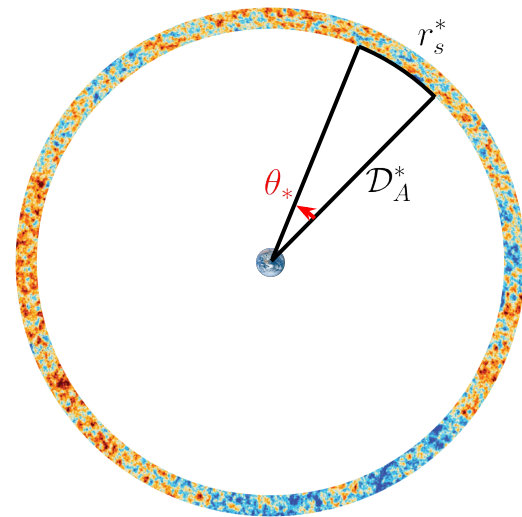


How to measure H_0 from the CMB ?

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$$\theta_* = \frac{r_s^*}{\mathcal{D}_A^*} \rightarrow \mathcal{D}_A^* = c \int_0^{z^*} \frac{dz}{H(z)}$$

$$H_{\text{late}}^2(z) = \frac{8\pi G}{3} [(\rho_b^0 + \rho_c^0)(1+z)^3 + \rho_\Lambda]$$



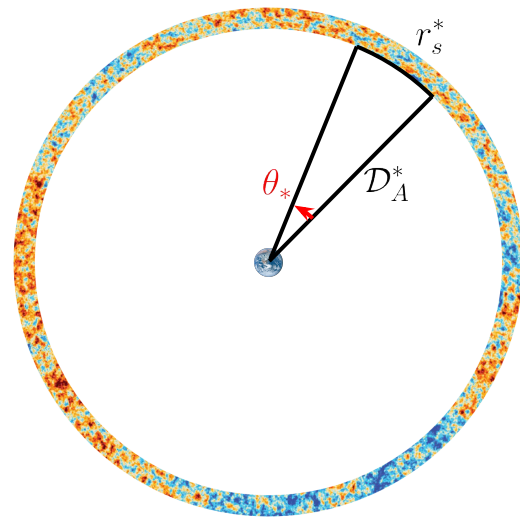
How to measure H_0 from the CMB ?

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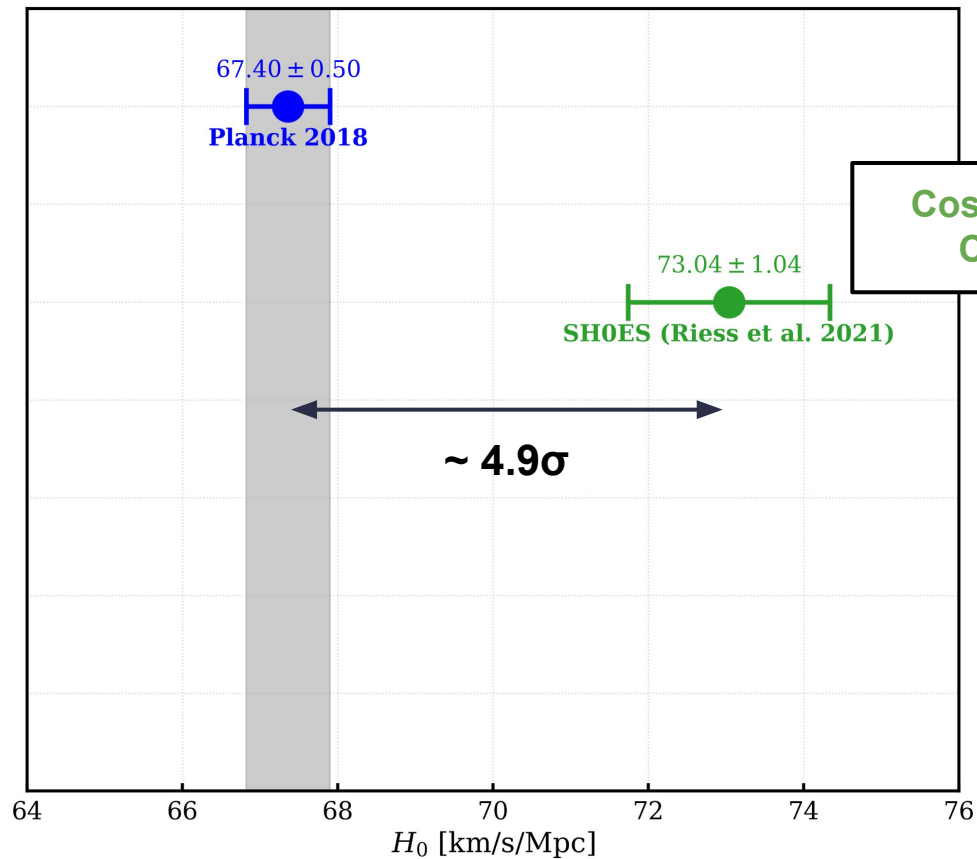
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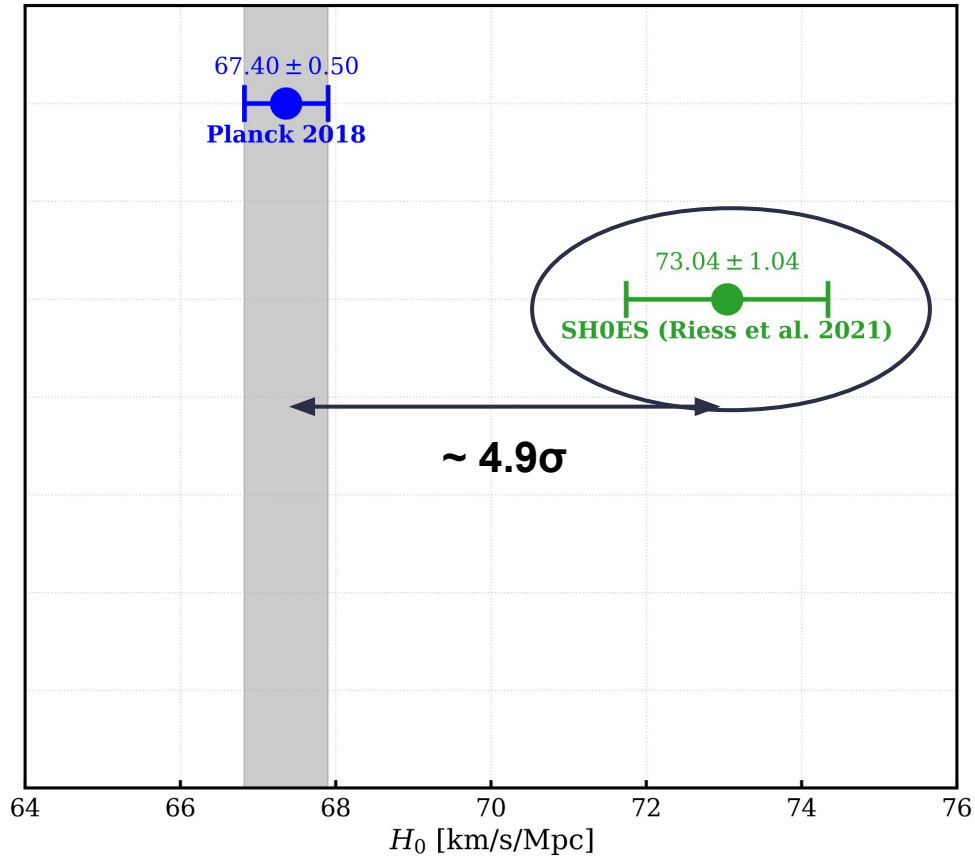
$$H_0^2 = \frac{8\pi G}{3} [\rho_b^0 + \rho_c^0 + \rho_\Lambda]$$



The Hubble tension as of today



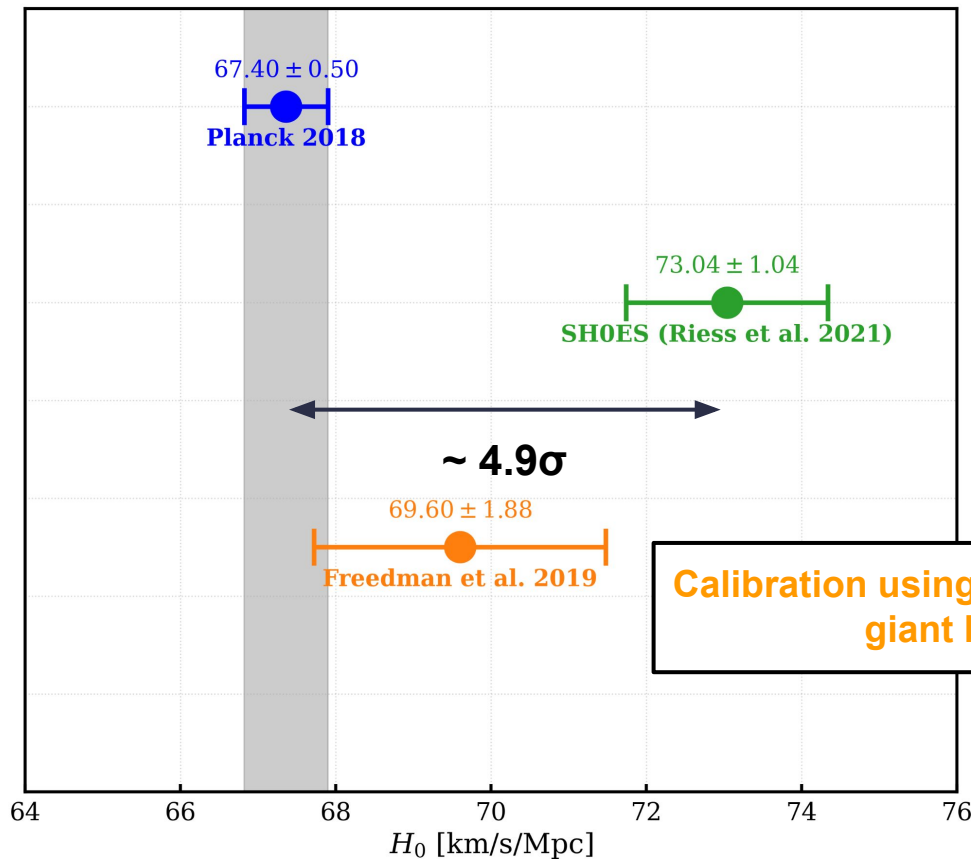
Cosmic distance ladder using
Cepheid calibrated SNIa



Option 1

Astrophysical biases affecting the local measurement of H_0

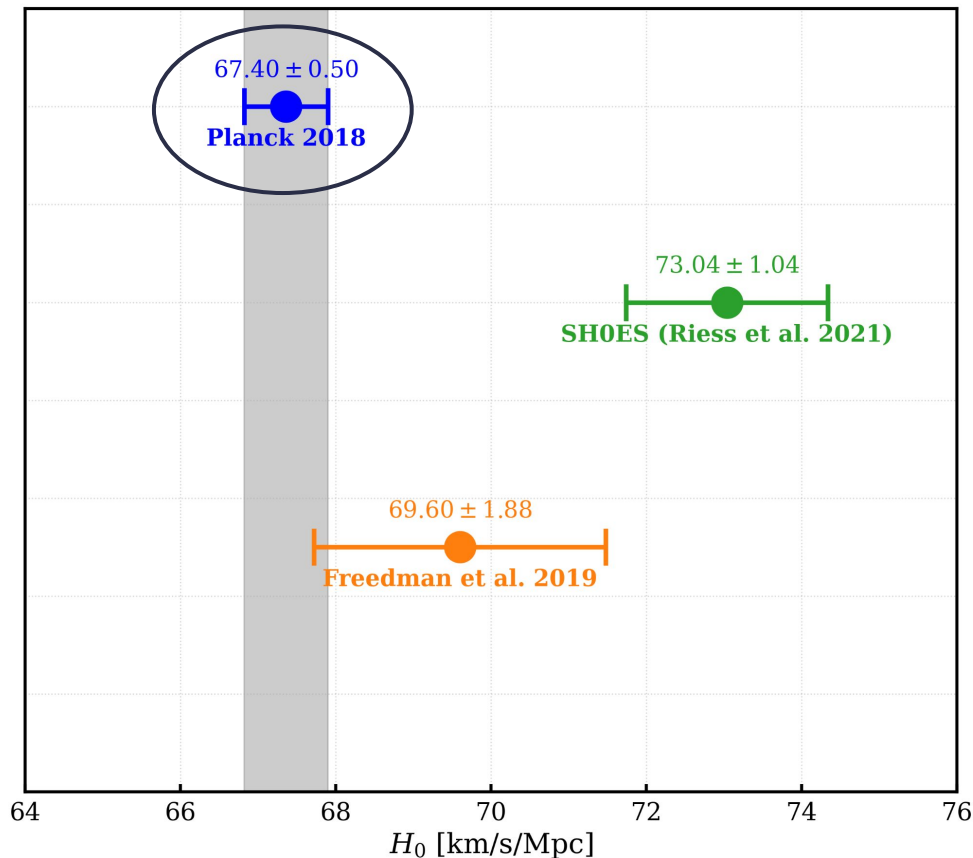
The Hubble tension as of today



Option 1

Astrophysical biases affecting the local measurement of H_0

Calibration using the tip of the red giant branch



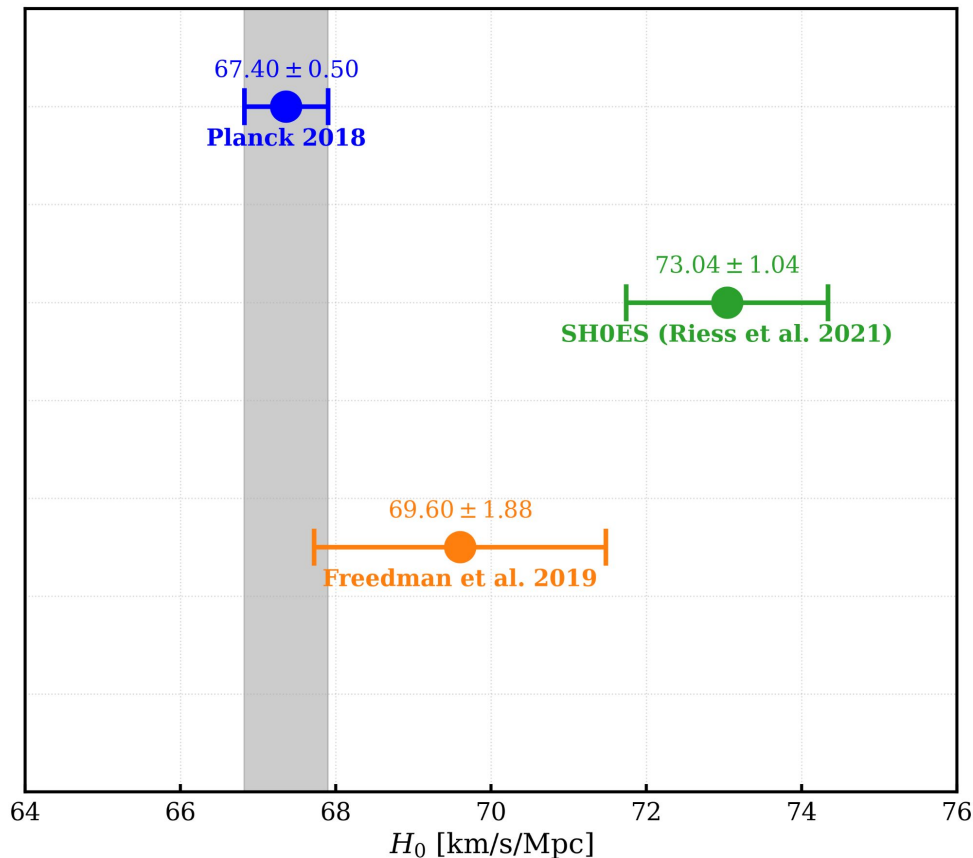
Option 1

Astrophysical biases affecting the local measurement of H_0

Option 2

Instrumental systematic effect biasing the value of H_0 inferred from the CMB

The Hubble tension as of today



Option 1

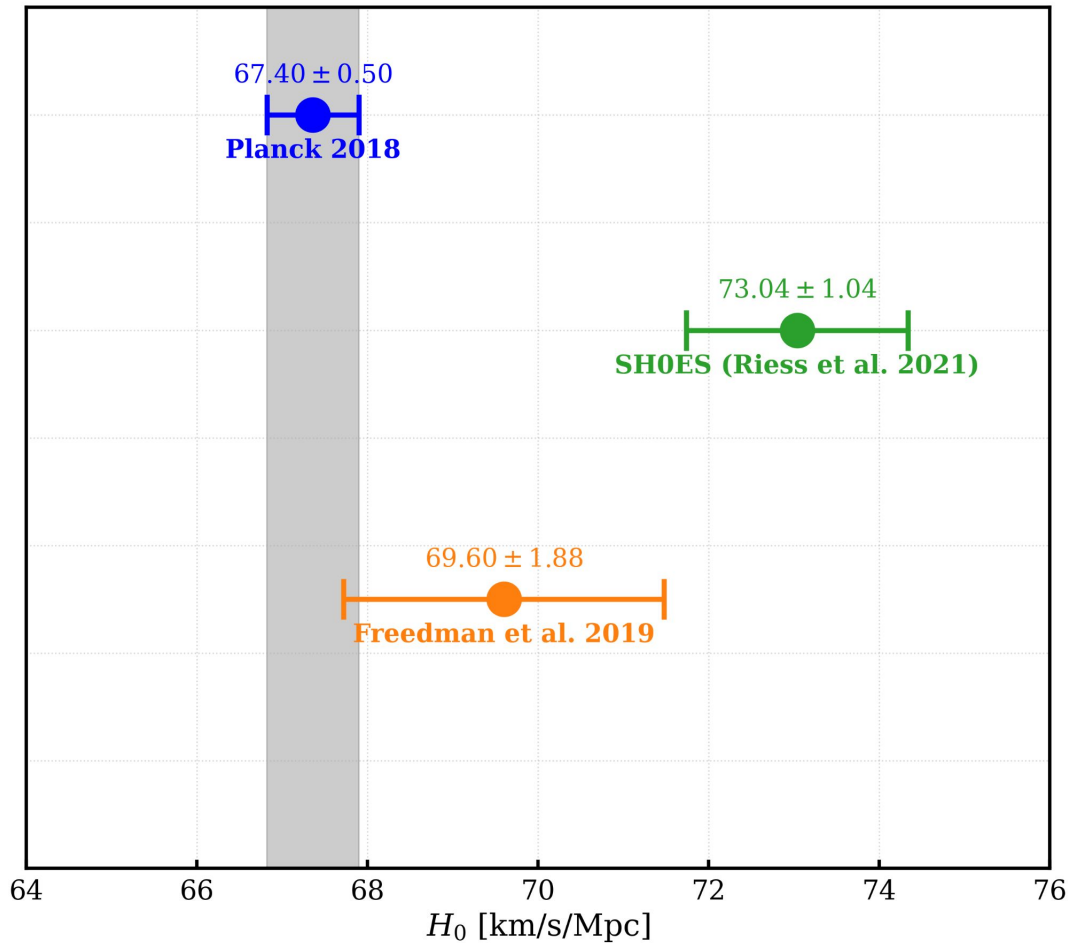
Astrophysical biases affecting the local measurement of H_0

Option 2

Instrumental systematic effect biasing the value of H_0 inferred from the CMB

Option 3

Physics beyond Λ CDM



Option 1

Astrophysical biases affecting the local measurement of H_0

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Instrumental systematic effect biasing the value of H_0 inferred from the CMB

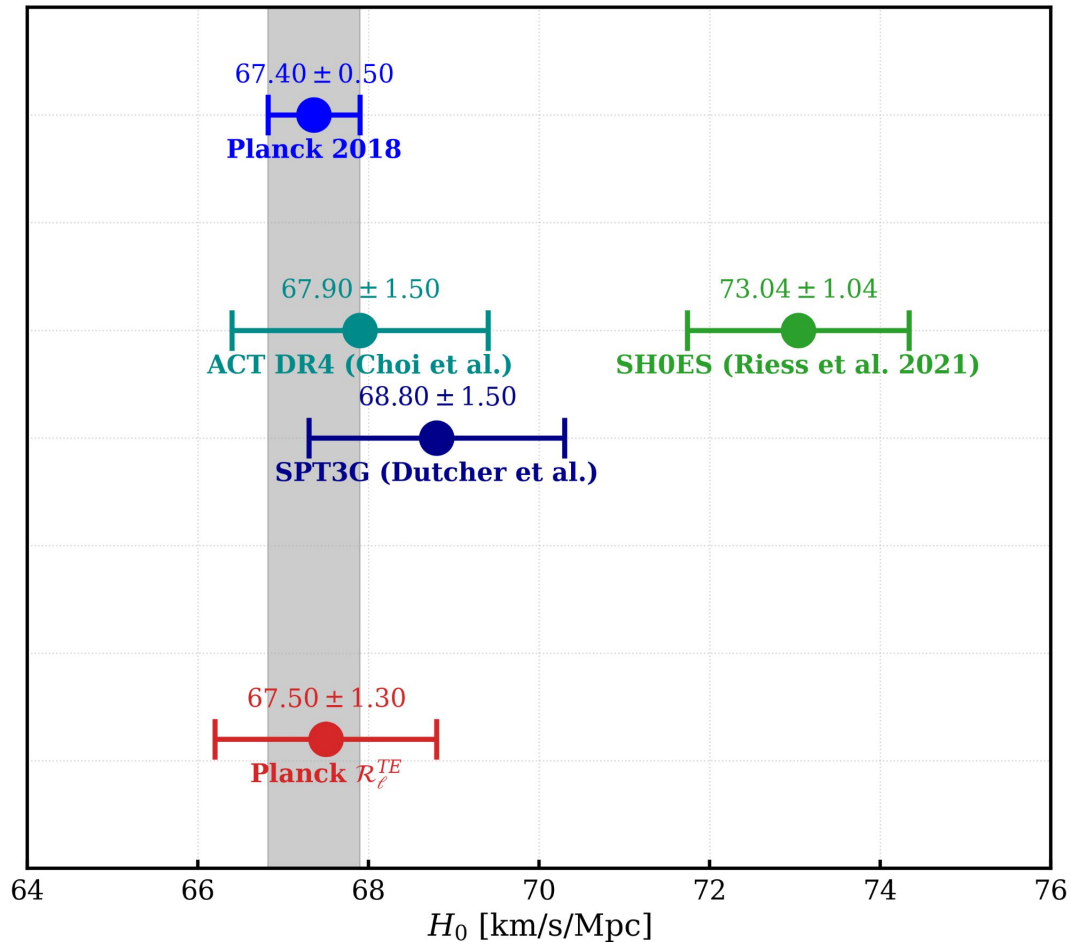
Option 3

Physics beyond Λ CDM

Independent measurements of H_0 from the ground

8





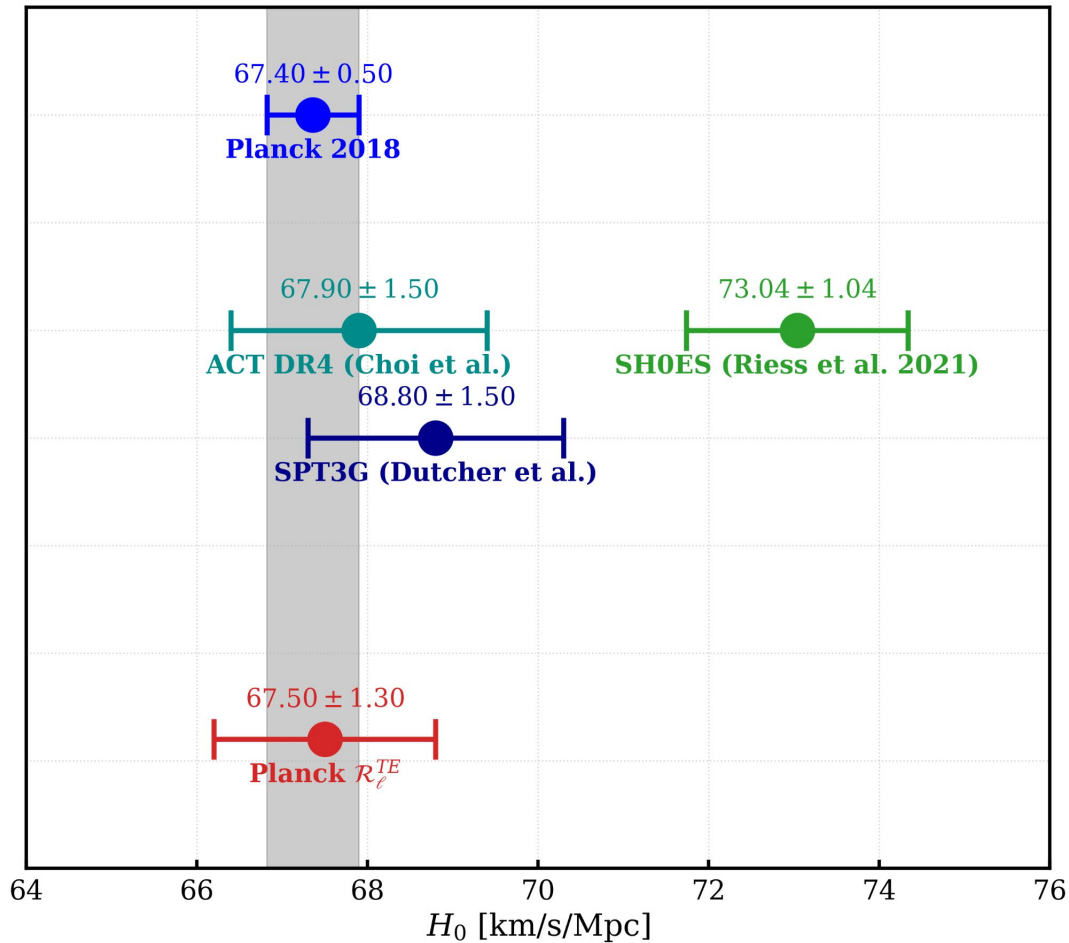
Option 2

Instrumental systematic effect
biasing the value of H_0
inferred from the CMB



Hard to shift the CMB inferred H_0
with a systematic effect :

- Independent measurements from Planck, ACT and SPT
- Constraint from the correlation coefficient, robust against multiplicative systematics



Option 1

Astrophysical biases affecting the local measurement of H_0

Option 2

Instrumental systematic effect biasing the value of H_0 inferred from the CMB

Option 3

Physics beyond Λ CDM

Motivation : obtain a higher value of H_0 from the CMB

Motivation : obtain a higher value of H_0 from the CMB \longrightarrow lower \mathcal{D}_A^*

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$$\theta_* = \frac{r_s^*}{D_A^*} \longrightarrow \text{Decrease } r_s^* = \int_{z^*}^{\infty} \frac{dz}{H(z)} c_s(z)$$

θ_* is fixed by observations

$$\frac{3H_{\text{early}}^2(z)}{8\pi G} = \rho_r(z) + \rho_m(z)$$

One proposed solution : Early Dark Energy


Motivation : obtain a higher value of H_0 from the CMB \longrightarrow lower D_A^*

$$\theta_* = \frac{r_s^*}{D_A^*} \longrightarrow \text{Decrease } r_s^* = \int_{z^*}^{\infty} \frac{dz}{H(z)} c_s(z)$$

θ_* is **Fixed by observations**

$$\frac{3H_{\text{early}}^2(z)}{8\pi G} = \rho_r(z) + \rho_m(z) + \rho_{\text{EDE}}(z)$$

The EDE component is described as a scalar field ϕ (Poulin+ 2019, Smith+ 2019)

Background evolution : $\ddot{\phi} + 3H\dot{\phi} + V'(\phi) = 0$  axion-like potential

$$V(\phi) = m^2 f^2 \left[1 - \cos \left(\frac{\phi}{f} \right) \right]^3$$

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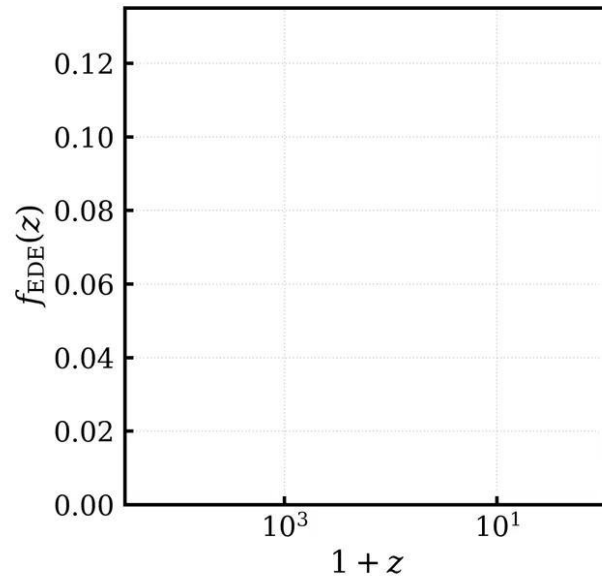
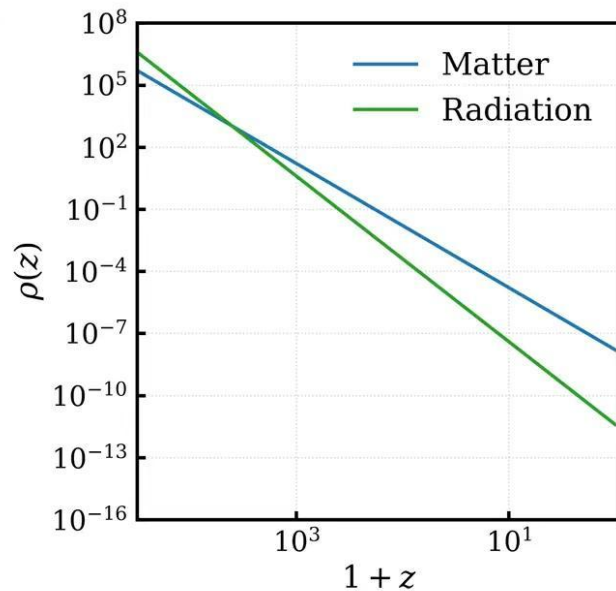
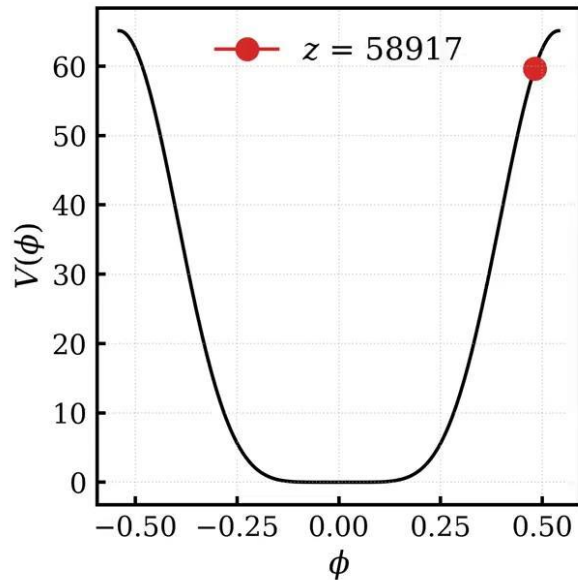
ϕ_i : initial field value

Early Dark Energy : frozen at early times

$$\ddot{\phi} + \boxed{3H\dot{\phi}} + V'(\phi) = 0$$

The field is initially frozen due to
Hubble friction ($H \gg m$)

acts as dark energy ($w = -1$)

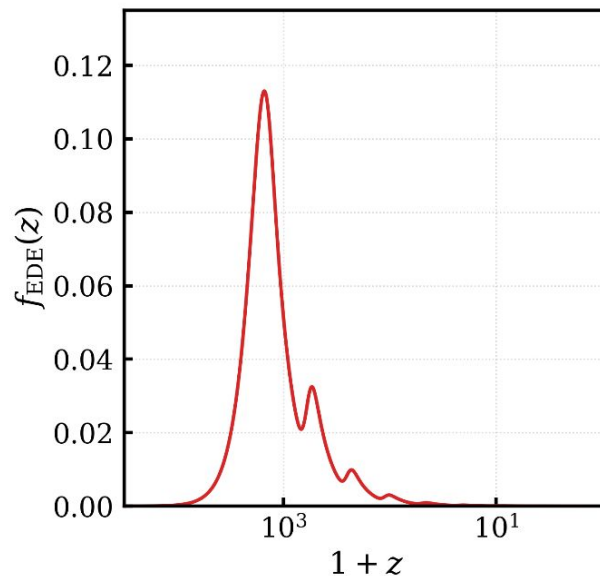
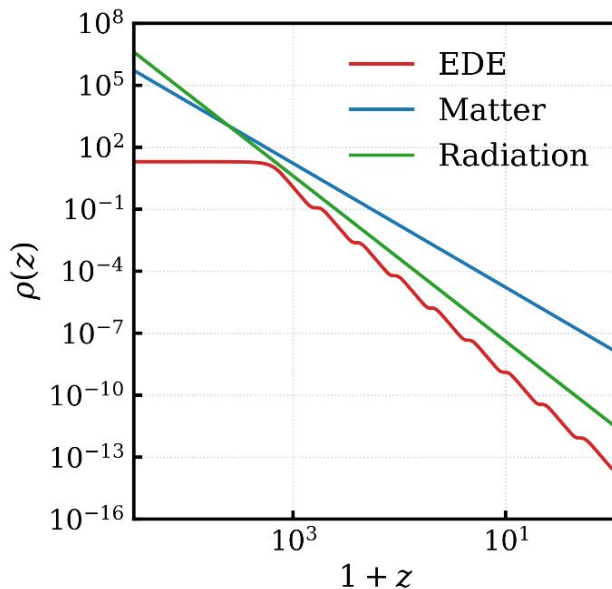
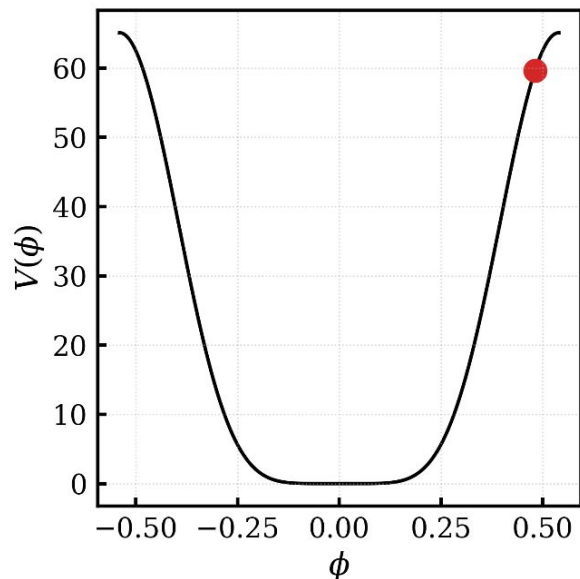


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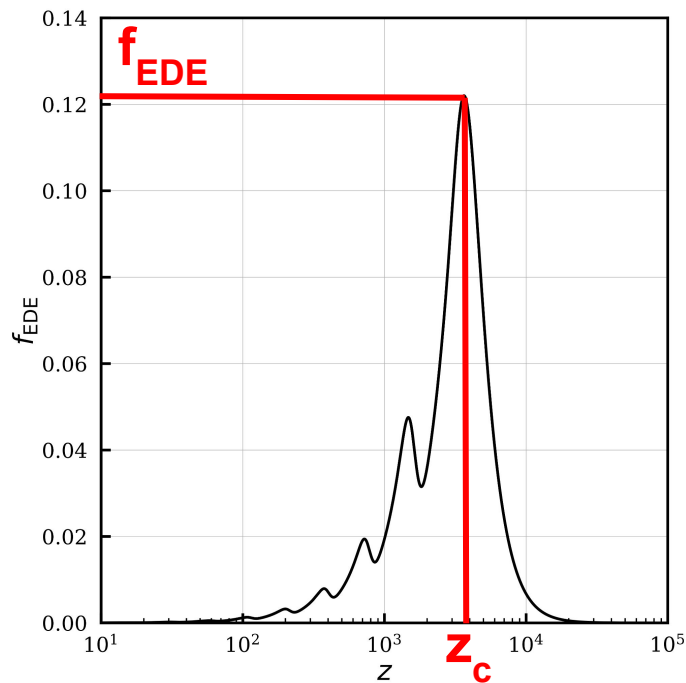


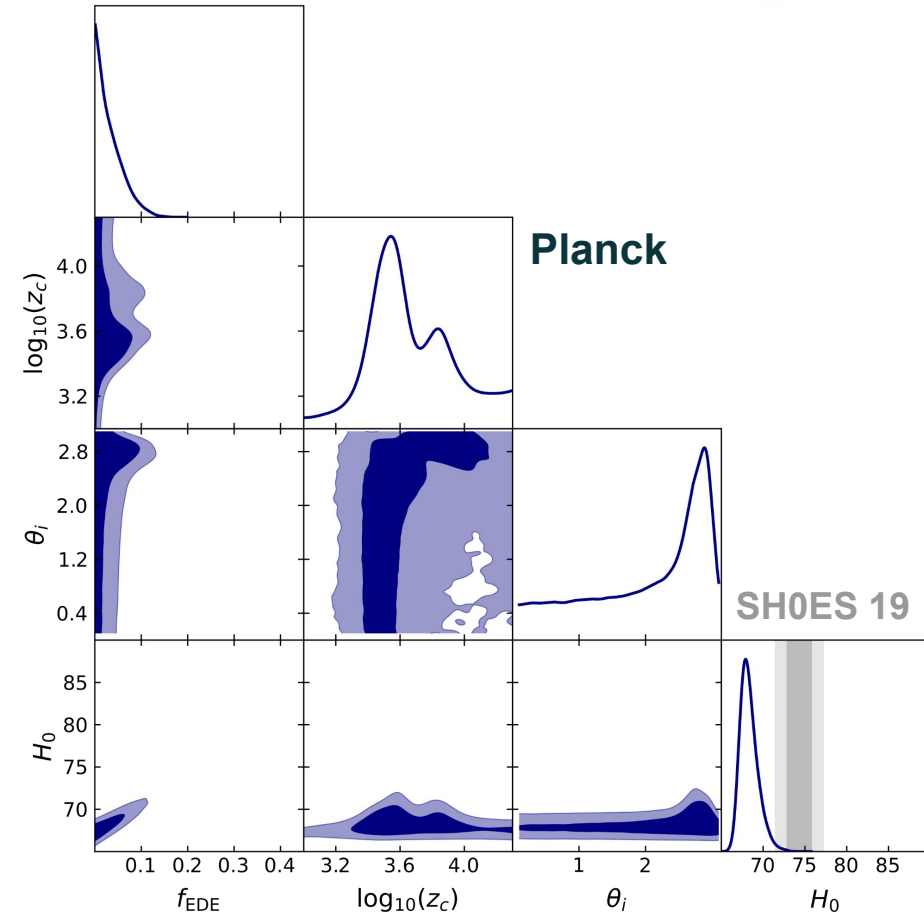
Early Dark Energy : phenomenological parametrization 12

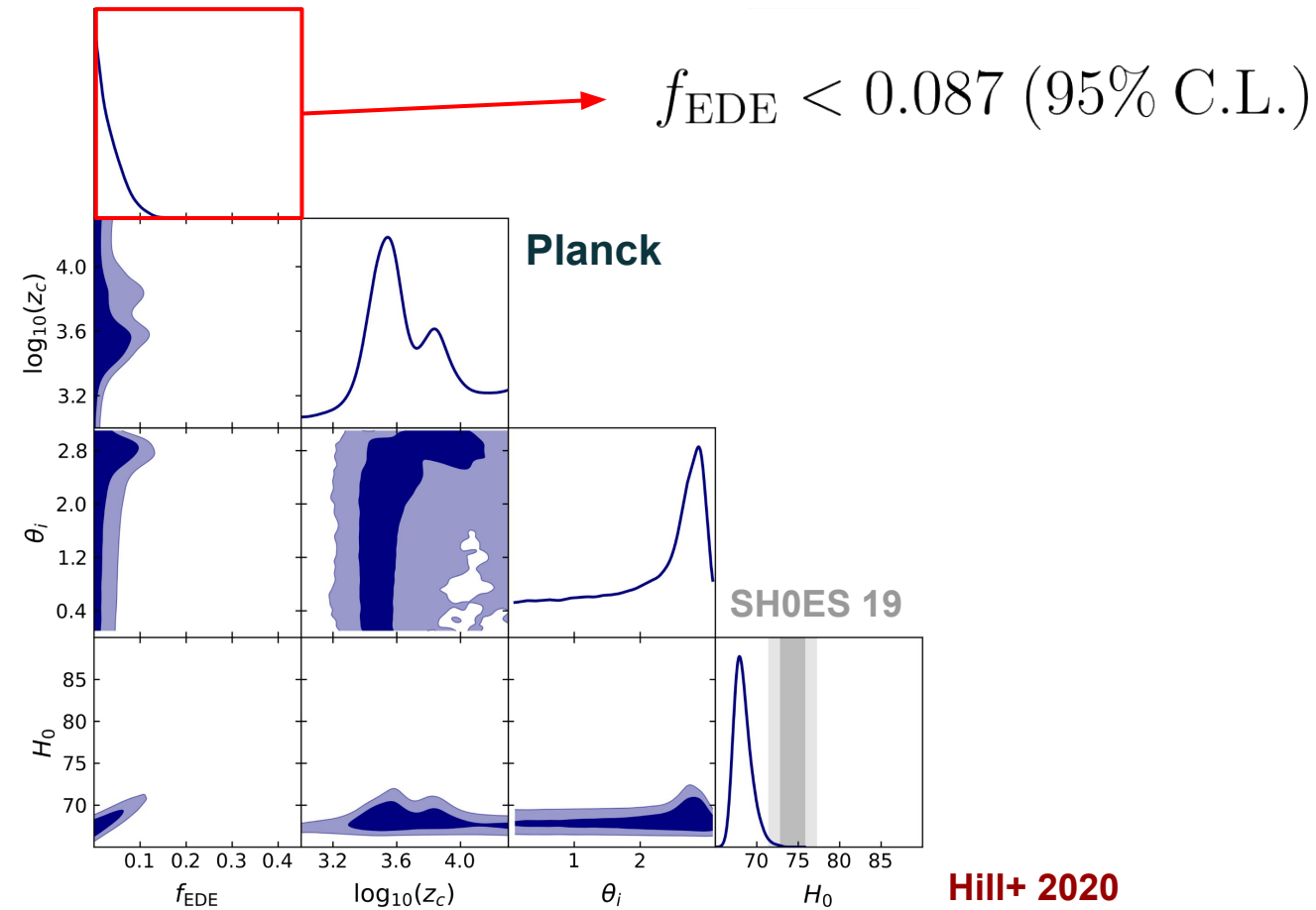
$$(m, f, \phi_i) \longrightarrow (f_{\text{EDE}}, z_c, \phi_i)$$

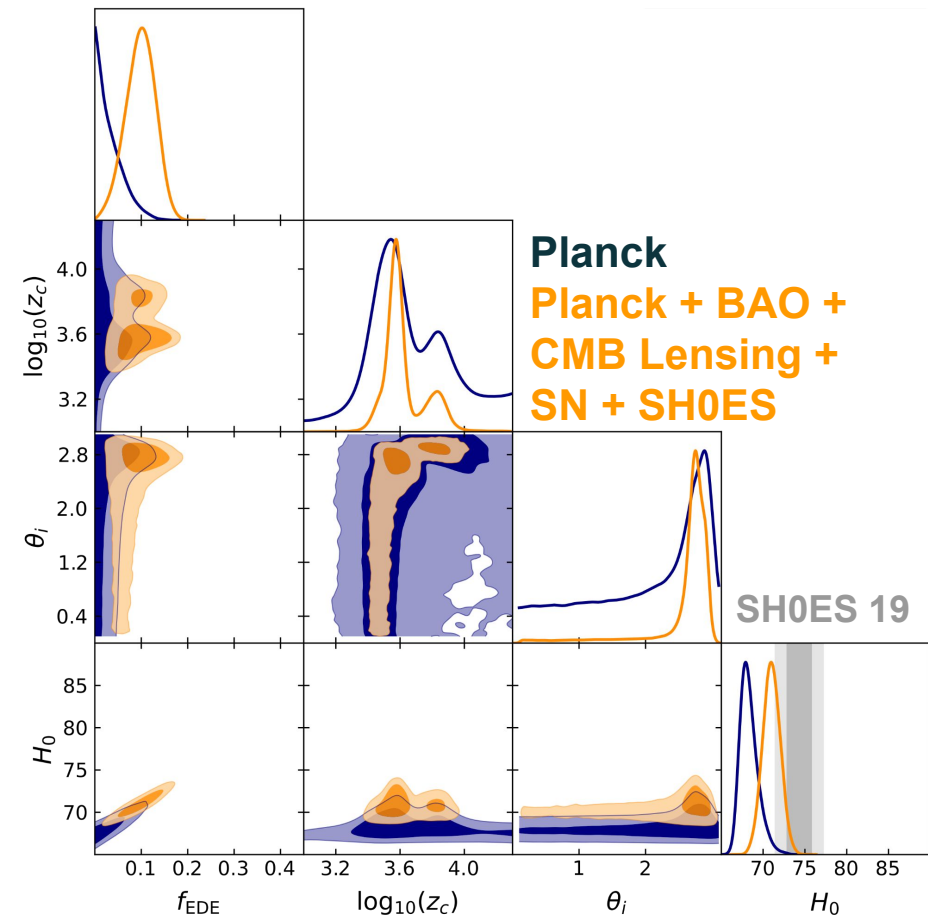
Field parameters

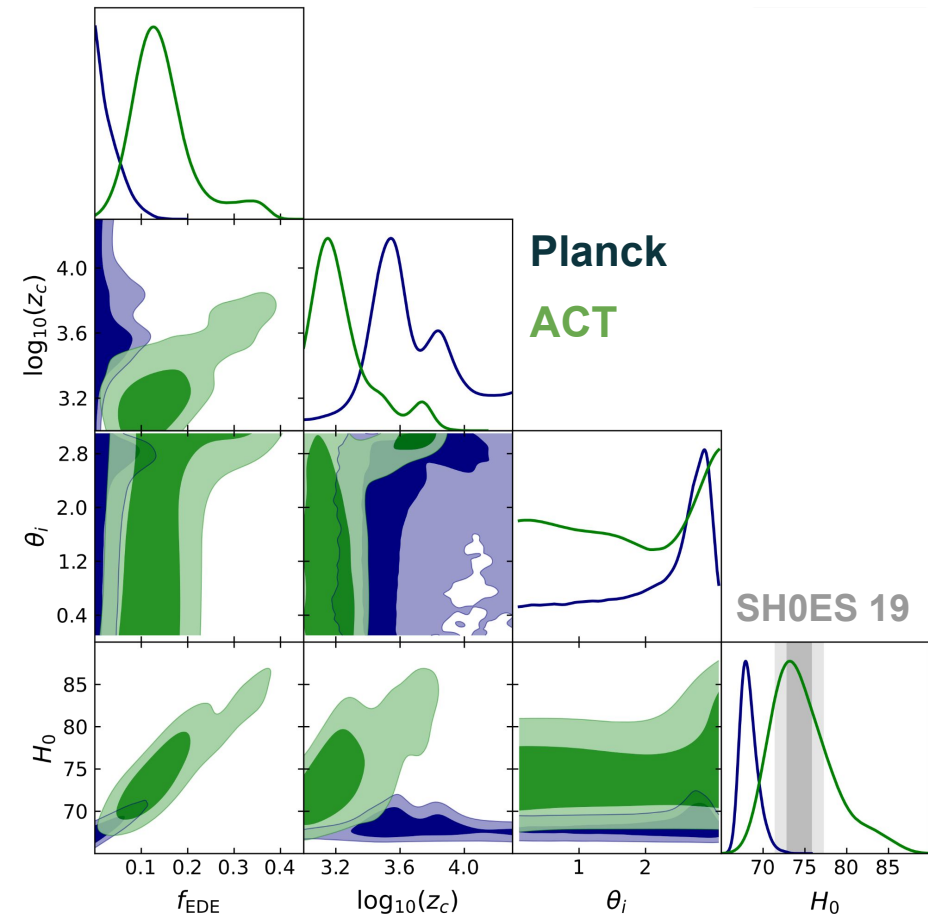
Phenomenological
parametrization

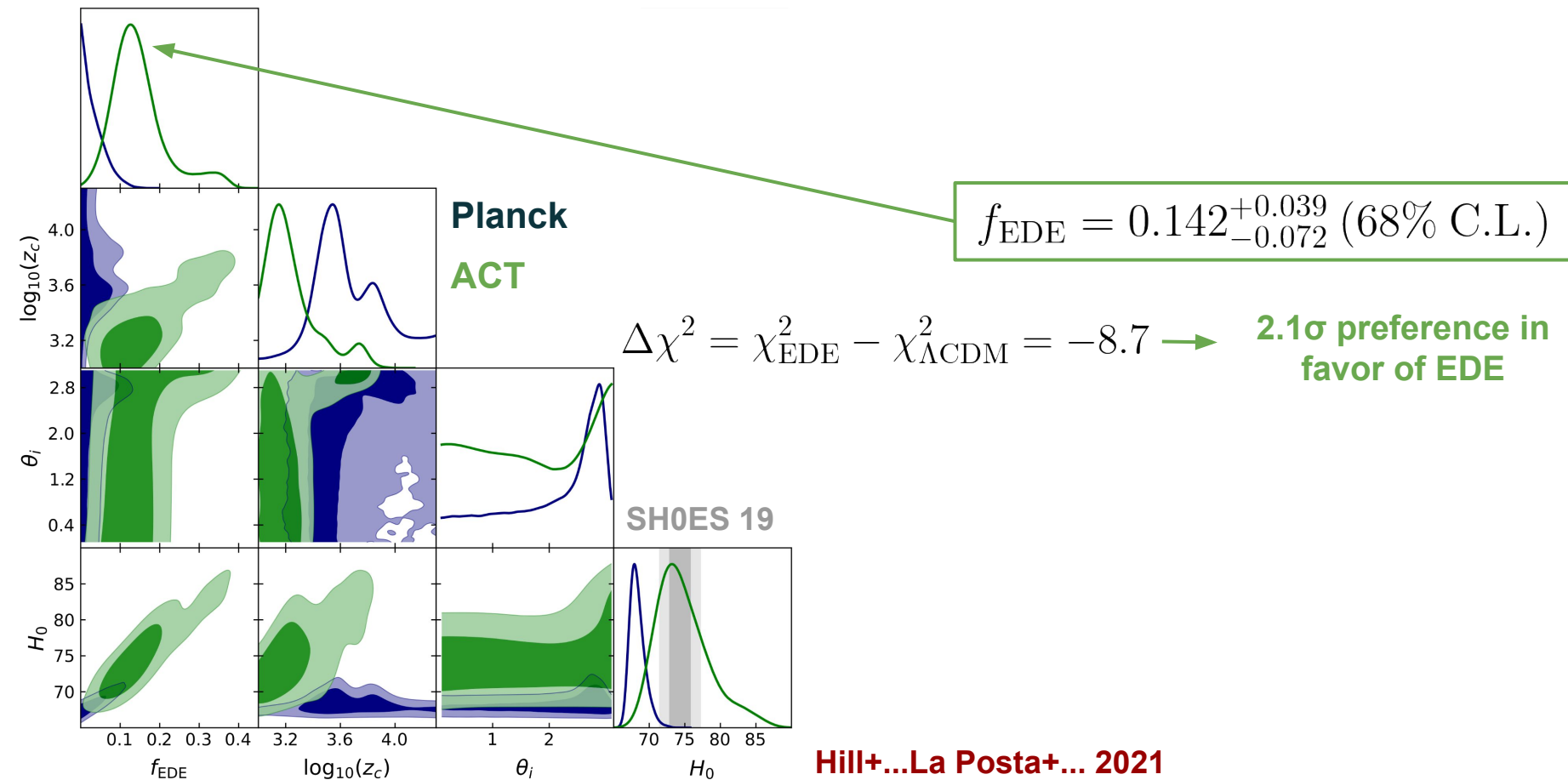


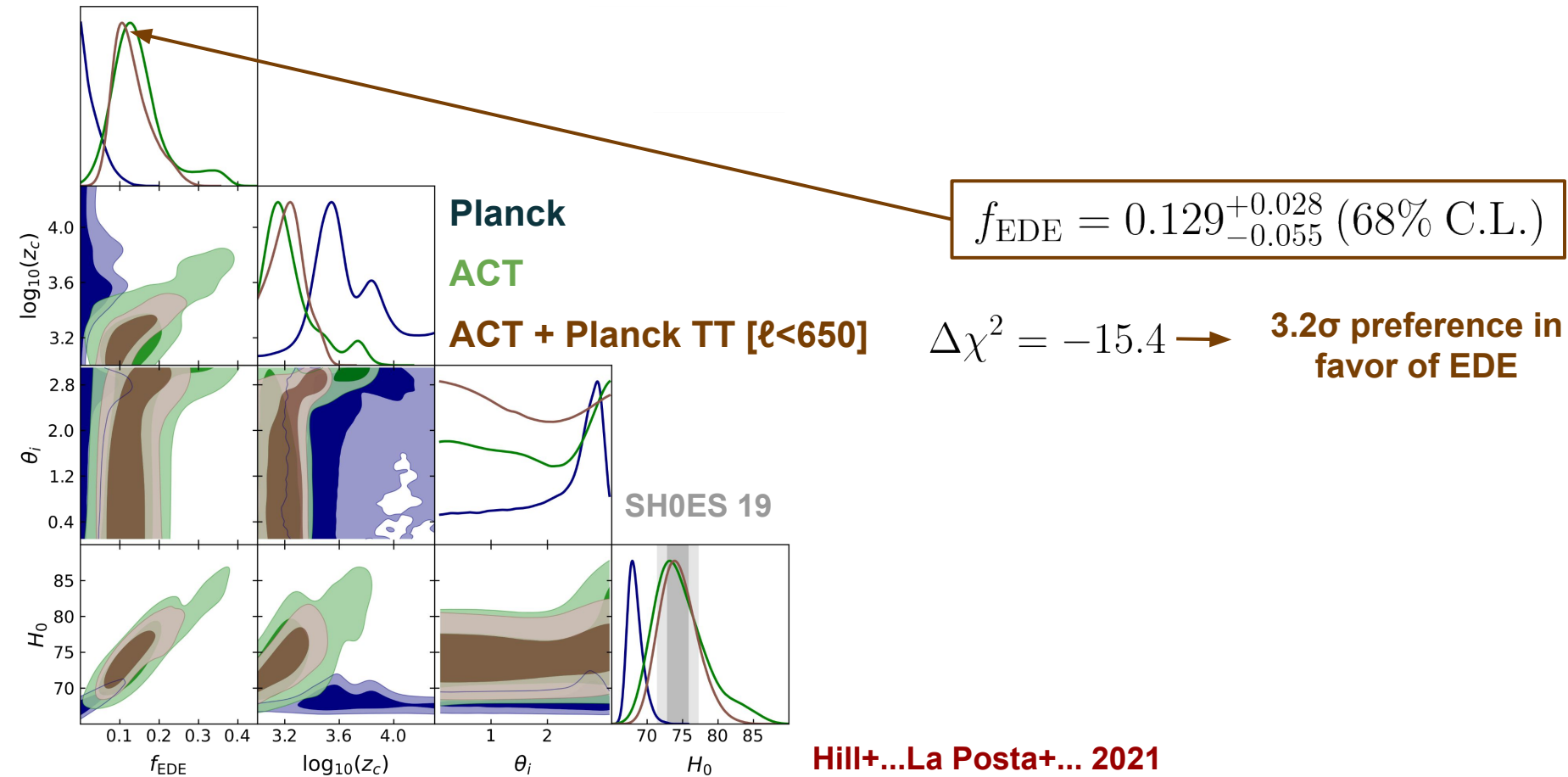






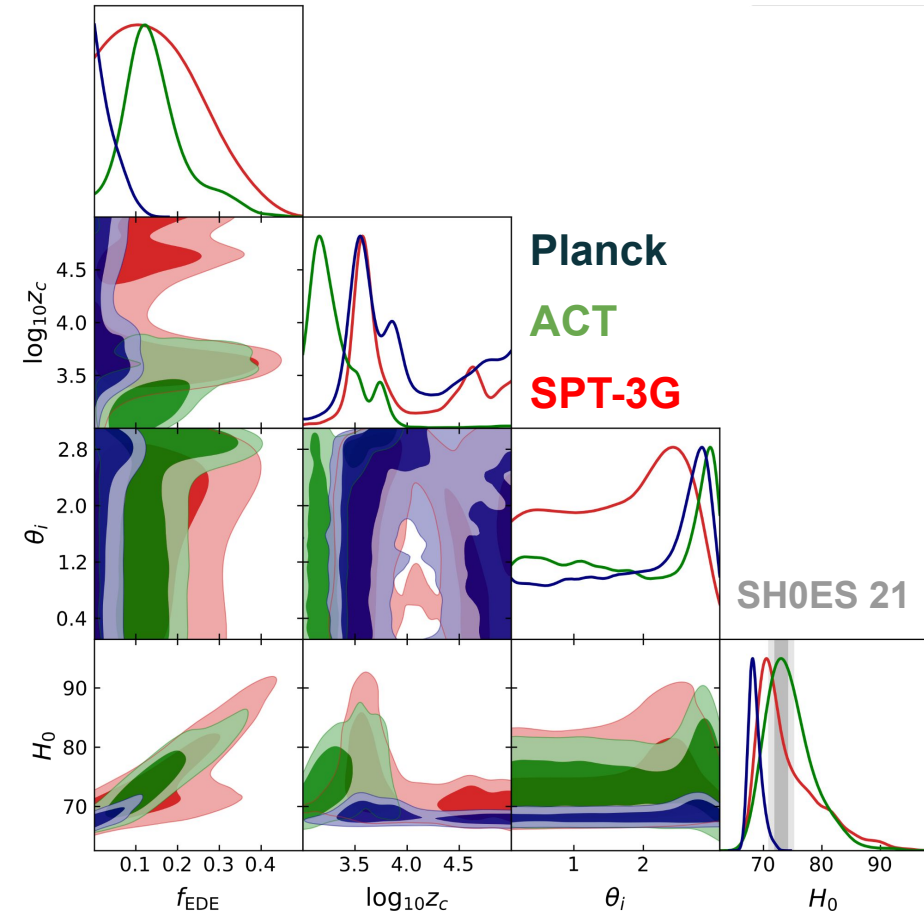


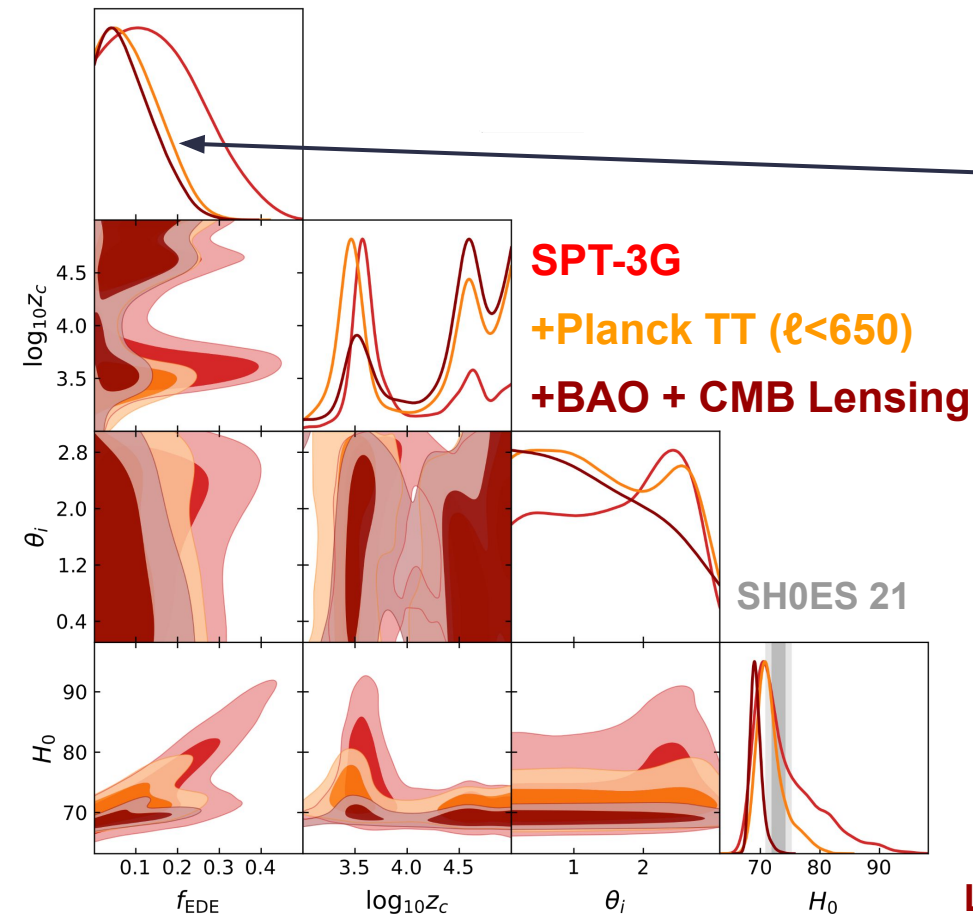




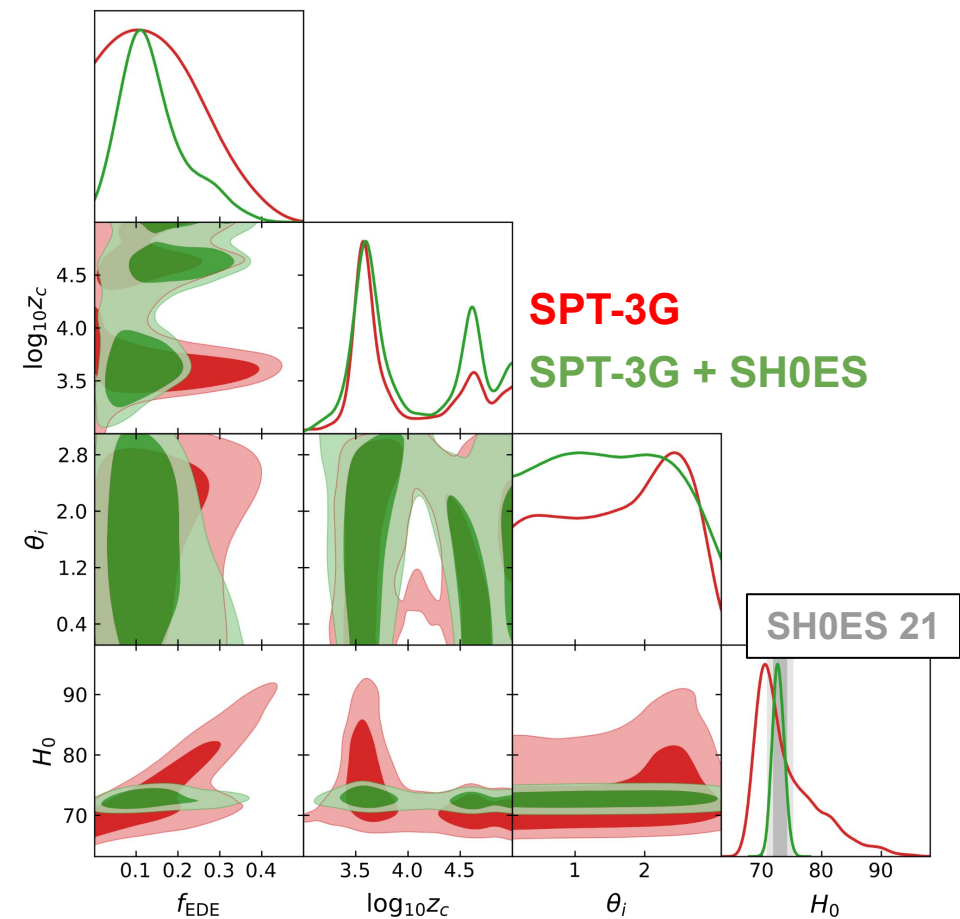
- Planck data alone don't favor high f_{EDE} values (Hill+ 2020)
- Planck data in combination with SH0ES show a preference for non-zero f_{EDE} (Poulin+ 2019, Smith+ 2019)
- ACT data alone favors EDE over Λ CDM (Hill+...La Posta+... 2021)

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- Motivates an analysis of EDE with public SPT-3G data





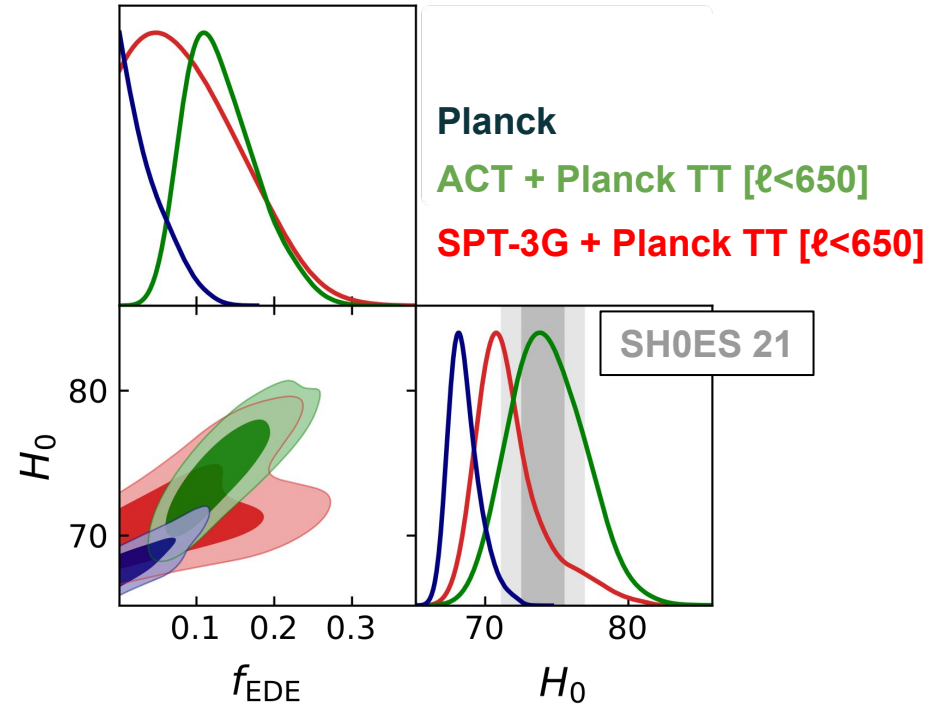
We tighten the constraint on f_{EDE} when we combine **SPT3G** and **Planck TT ($\ell < 650$)** or when we add **LSS probes**



$$\Delta\chi_{\text{SPT-3G}}^2 = -6.3$$

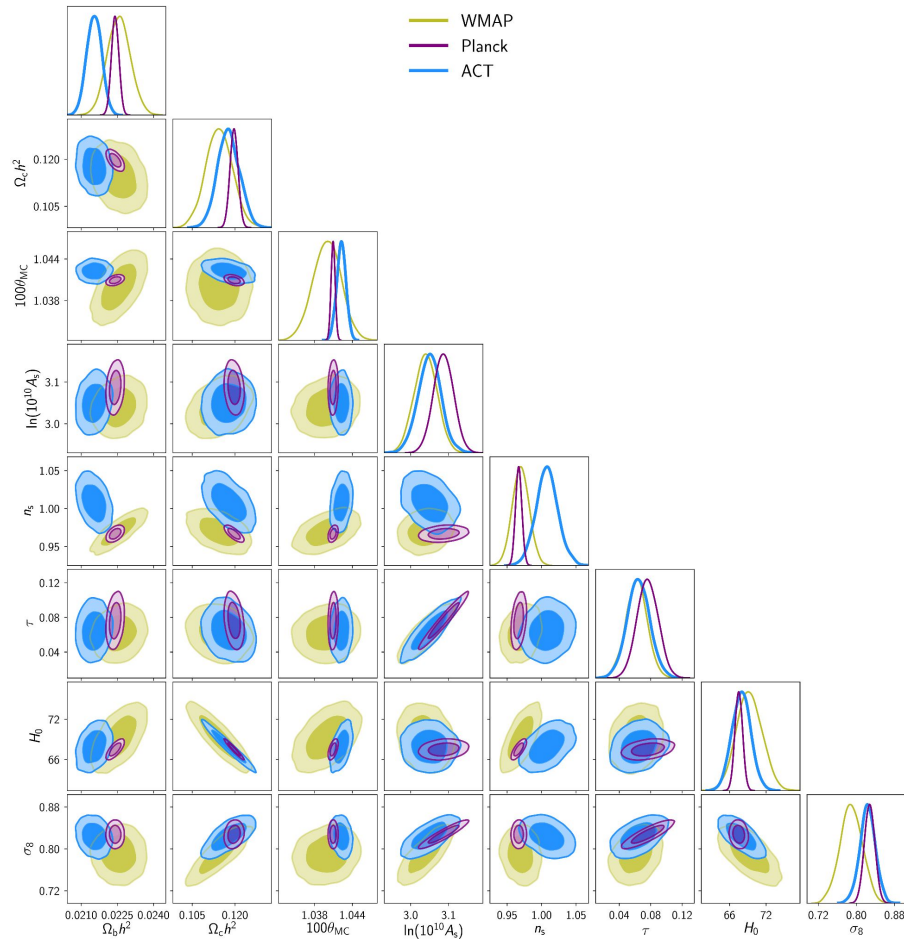
improvement of the fit to
SPT-3G data (with
respect to ΛCDM)

- Planck data alone do not favor high f_{EDE} values
- Planck + SH0ES show a preference for $f_{\text{EDE}} \sim 10\%$
- ACT DR4 data favors EDE over ΛCDM (with $f_{\text{EDE}} \sim 10\%$)
- SPT-3G is not as constraining as ACT and Planck : but sees some degree of EDE when combined with SH0ES

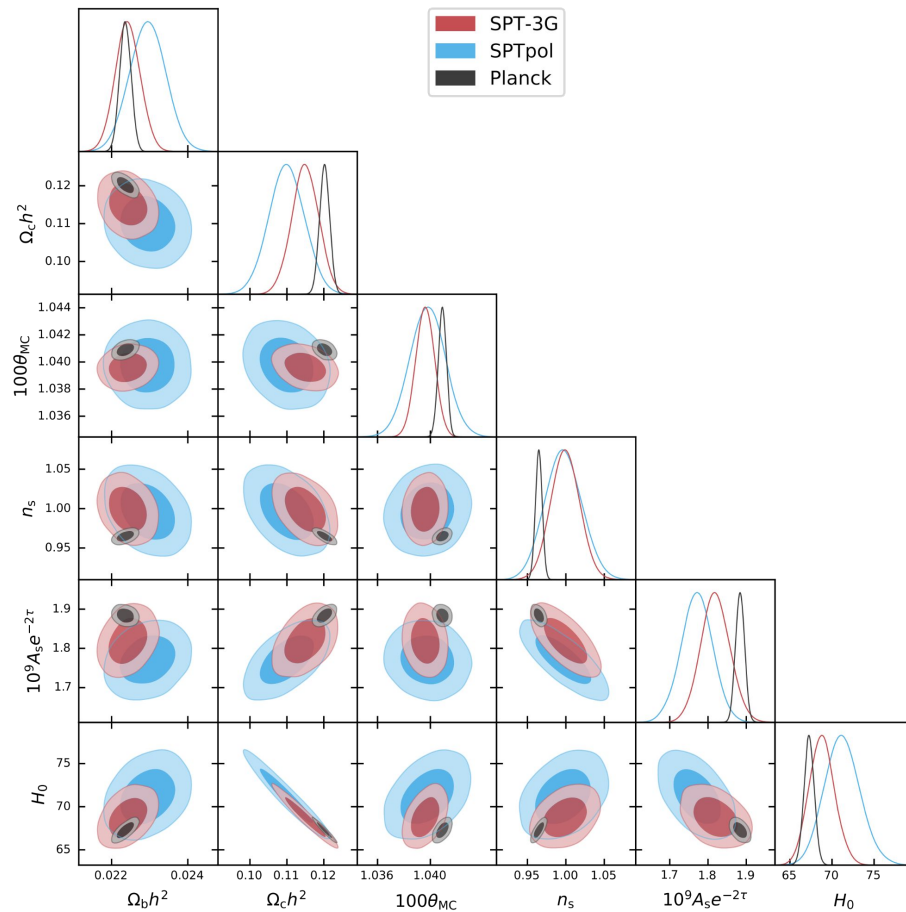


Extra-slides

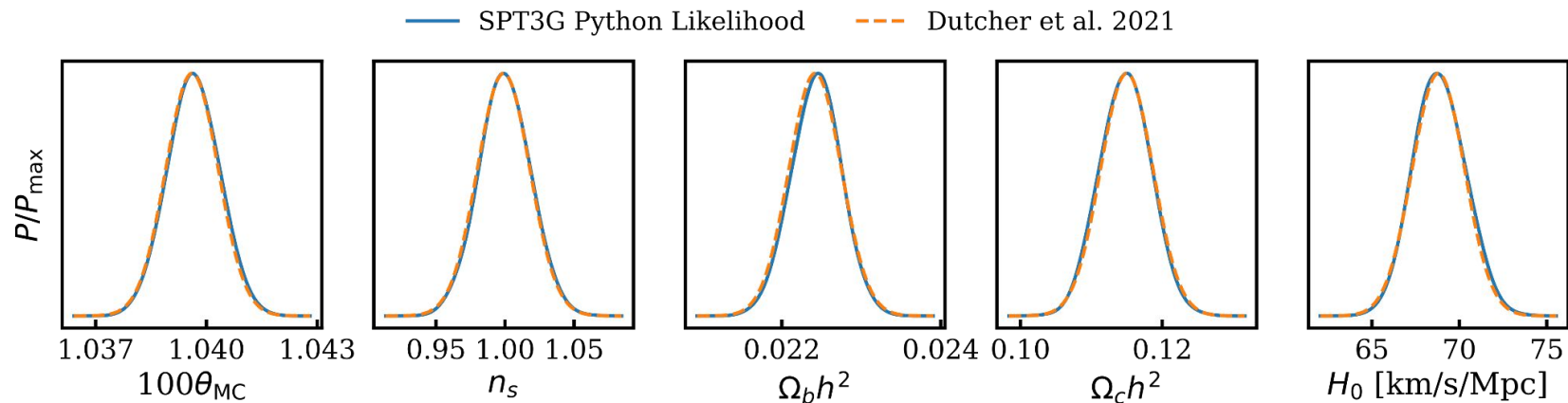
ACTPol cosmology



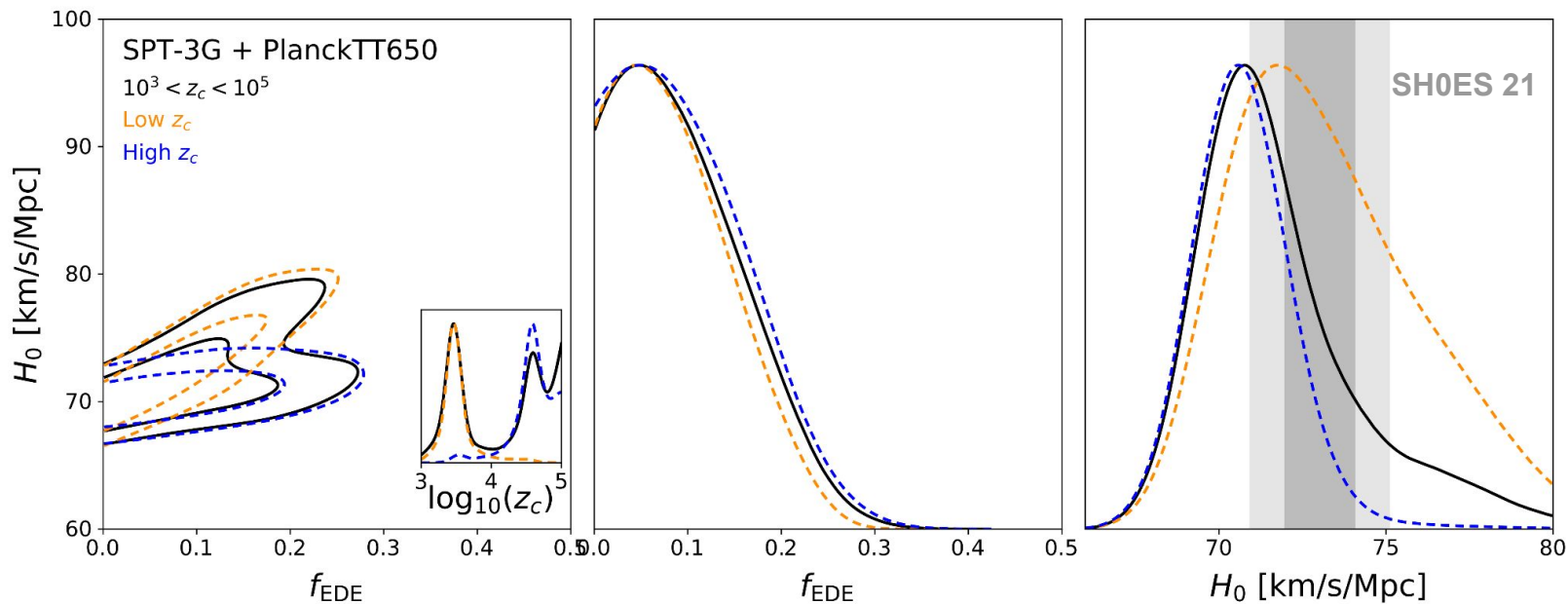
SPT-3G cosmology



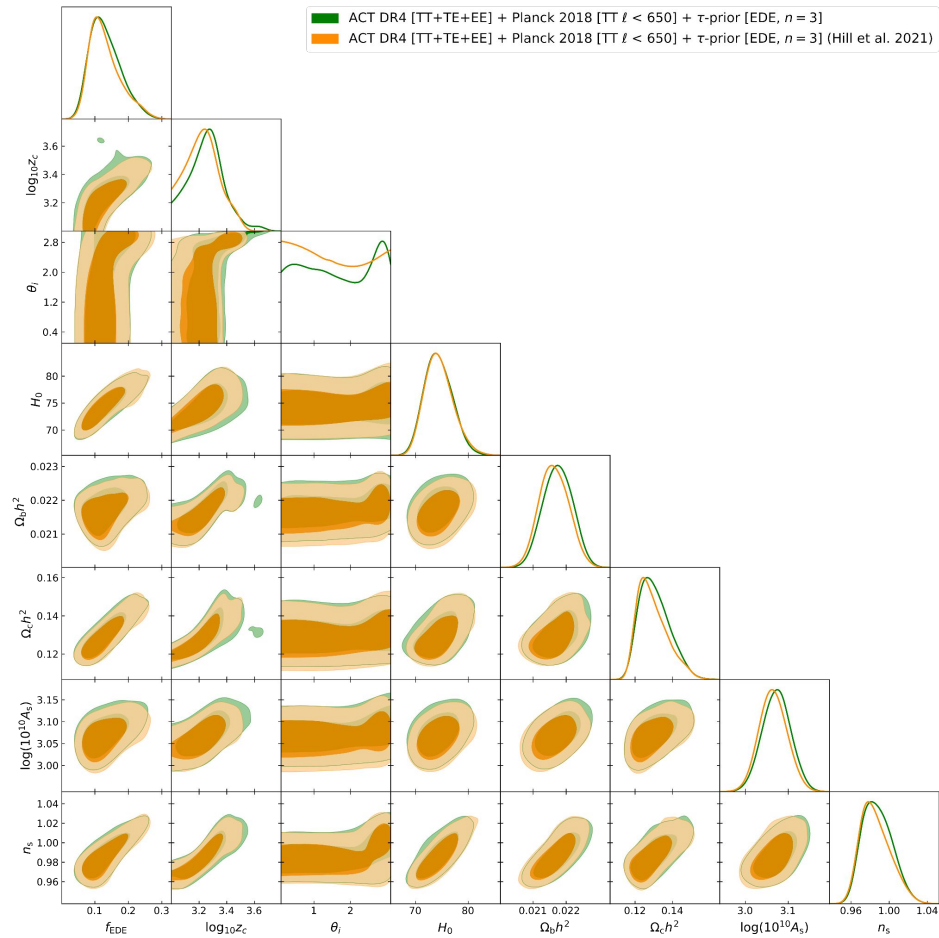
SPT-3G python implementation



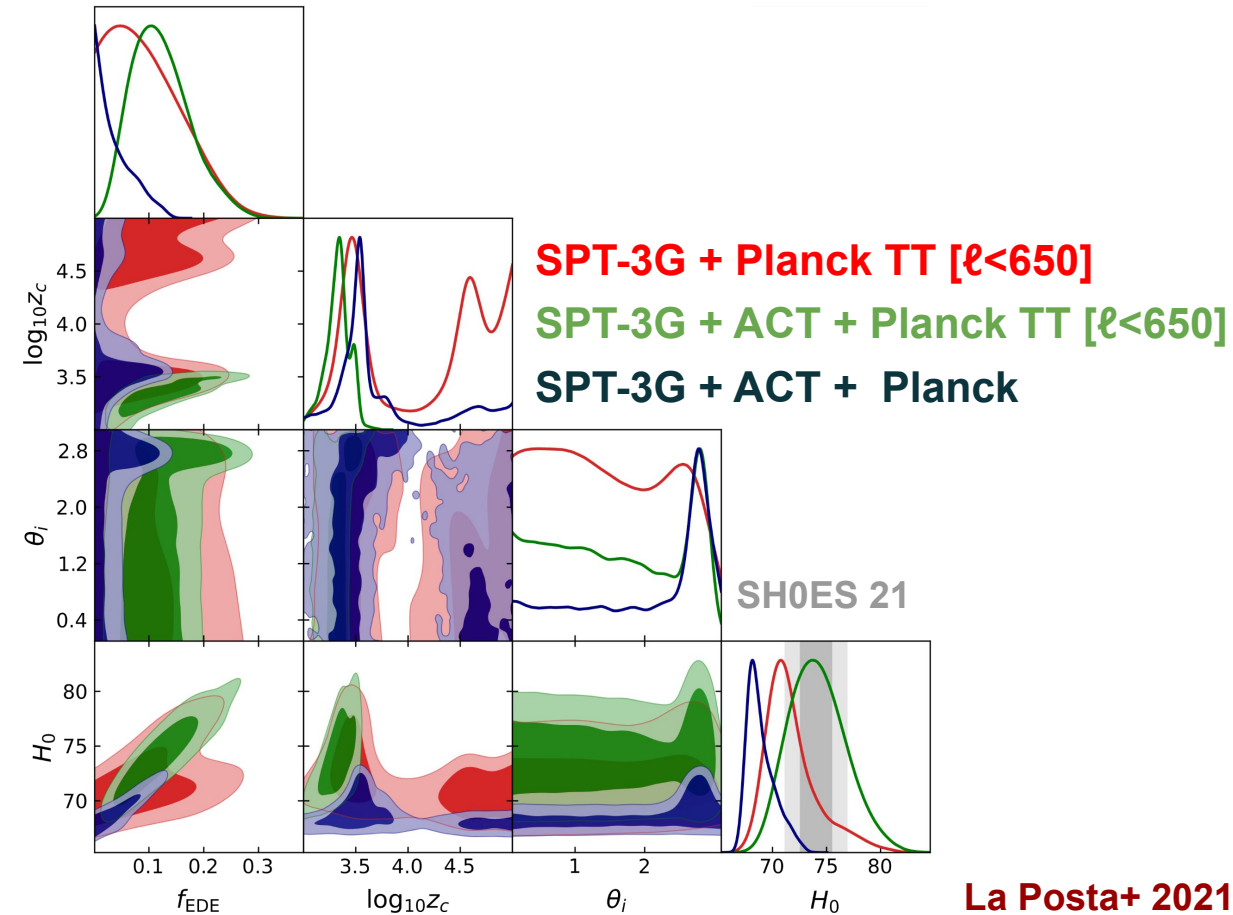
Impact of the z_c prior



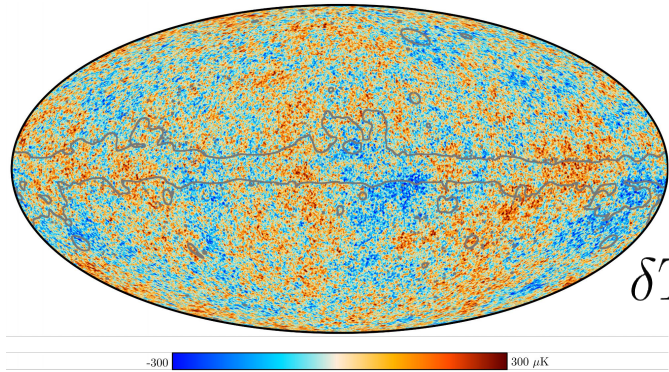
CAMB/CLASS EDE models



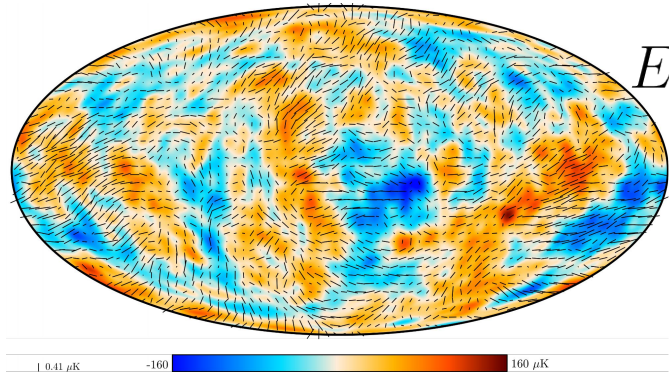
Combining with other CMB datasets



Systematics in the CMB



$$\delta T^{\text{sky}}(\hat{n})$$



$$E^{\text{sky}}(\hat{n})$$

$$\mathcal{I}_{T/E}$$



$$\begin{aligned} \delta T^{\text{obs}}(\hat{n}) &= \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n}) \\ E^{\text{obs}}(\hat{n}) &= \mathcal{I}_E * E^{\text{sky}}(\hat{n}) \end{aligned}$$

Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

$$\mathcal{I}_T = \mathcal{F}_T * c * B_T$$

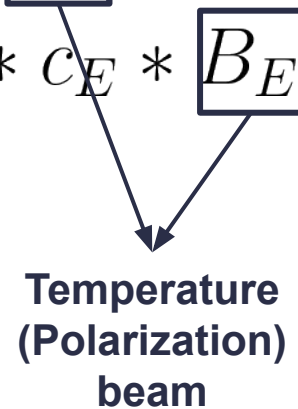
$$\mathcal{I}_E = \mathcal{F}_E * c * c_E * B_E$$

Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

- Finite angular resolution (beams)

$$\mathcal{I}_T = \mathcal{F}_T * c * B_T$$
$$\mathcal{I}_E = \mathcal{F}_E * c * c_E * B_E$$


Temperature
(Polarization)
beam

Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

- Finite angular resolution (beams)
- Calibration

$$\mathcal{I}_T = \mathcal{F}_T * \boxed{C} * B_T$$

$$\mathcal{I}_E = \mathcal{F}_E * \boxed{C} * c_E * B_E$$

Global
calibration



Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

- Finite angular resolution (beams)
- Calibration
- Polarization efficiency

$$\mathcal{I}_T = \mathcal{F}_T * c * B_T$$

$$\mathcal{I}_E = \mathcal{F}_E * c * \boxed{c_E} * B_E$$

Polarization
efficiency

Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

- Finite angular resolution (beams)
- Calibration
- Polarization efficiency
- Transfer functions (map-making)

$$\mathcal{I}_T = \boxed{\mathcal{F}_T} * c * B_T$$
$$\mathcal{I}_E = \boxed{\mathcal{F}_E} * c * c_E * B_E$$

Temperature
(polarization)
transfer function

Systematics in the CMB

$$\delta T^{\text{obs}}(\hat{n}) = \mathcal{I}_T * \delta T^{\text{sky}}(\hat{n})$$

$$E^{\text{obs}}(\hat{n}) = \mathcal{I}_E * E^{\text{sky}}(\hat{n})$$

- Finite angular resolution (beams)
- Calibration
- Polarization efficiency
- Transfer functions (map-making)

**These instrumental effects are
multiplicative in harmonic space**

$$C_{\ell}^{TT,\text{obs}} = (\mathcal{F}_{\ell}^T)^2 c^2 (B_{\ell}^T)^2 C_{\ell}^{TT}$$

$$C_{\ell}^{EE,\text{obs}} = (\mathcal{F}_{\ell}^E)^2 c^2 c_E^2 (B_{\ell}^E)^2 C_{\ell}^{EE}$$

$$C_{\ell}^{TE,\text{obs}} = \mathcal{F}_{\ell}^T \mathcal{F}_{\ell}^E c^2 c_E B_{\ell}^T B_{\ell}^E C_{\ell}^{EE}$$

Correlation coefficient of T and E modes

$$\mathcal{R}_l^{TE} = \frac{\langle a_{lm}^T a_{lm}^{E*} \rangle}{\sqrt{\langle a_{lm}^T a_{lm}^{T*} \rangle \langle a_{lm}^E a_{lm}^{E*} \rangle}} = \frac{C_l^{TE}}{\sqrt{C_l^{TT} C_l^{EE}}}$$

Correlation coefficient of T and E modes

$$\mathcal{R}_\ell^{TE} = \frac{\langle a_{\ell m}^T a_{\ell m}^{E*} \rangle}{\sqrt{\langle a_{\ell m}^T a_{\ell m}^{T*} \rangle \langle a_{\ell m}^E a_{\ell m}^{E*} \rangle}} = \frac{C_\ell^{TE}}{\sqrt{C_\ell^{TT} C_\ell^{EE}}}$$

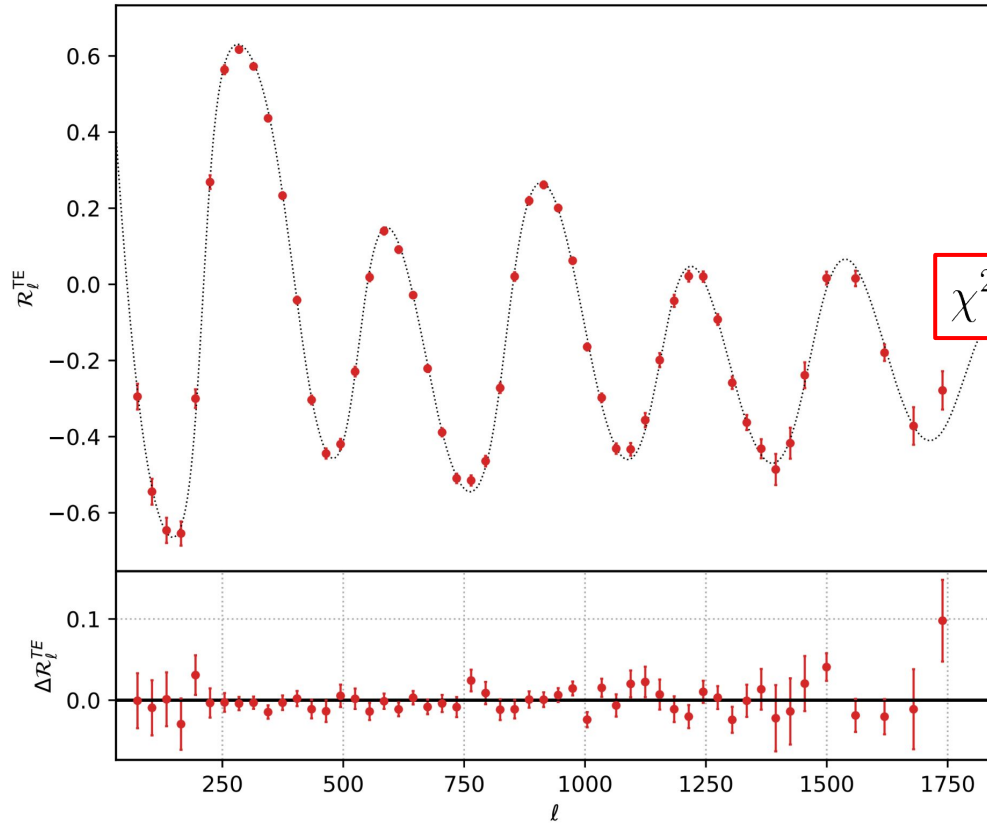
$$\mathcal{R}_\ell^{TE, \text{obs}} = \frac{\mathcal{F}_\ell^T \mathcal{F}_\ell^E c^2 c_E B_\ell^T B_\ell^E C_\ell^{TE}}{\sqrt{(\mathcal{F}_\ell^T)^2 c^2 (B_\ell^T)^2 C_\ell^{TT} \times (\mathcal{F}_\ell^E)^2 c^2 c_E^2 (B_\ell^E)^2 C_\ell^{EE}}}$$

Correlation coefficient of T and E modes

$$\mathcal{R}_l^{TE} = \frac{\langle a_{lm}^T a_{lm}^{E*} \rangle}{\sqrt{\langle a_{lm}^T a_{lm}^{T*} \rangle \langle a_{lm}^E a_{lm}^{E*} \rangle}} = \frac{C_l^{TE}}{\sqrt{C_l^{TT} C_l^{EE}}}$$

$$\mathcal{R}_l^{TE, \text{obs}} = \frac{\cancel{\mathcal{F}_l^T \mathcal{F}_l^E} \cancel{c^2 c_E} \cancel{B_l^T B_l^E} C_l^{TE}}{\sqrt{(\cancel{\mathcal{F}_l^T})^2 \cancel{c^2} (\cancel{B_l^T})^2 C_l^{TT} \times (\cancel{\mathcal{F}_l^E})^2 \cancel{c^2} \cancel{c_E^2} (\cancel{B_l^E})^2 C_l^{EE}}} = \mathcal{R}_l^{TE}$$

Planck correlation coefficient

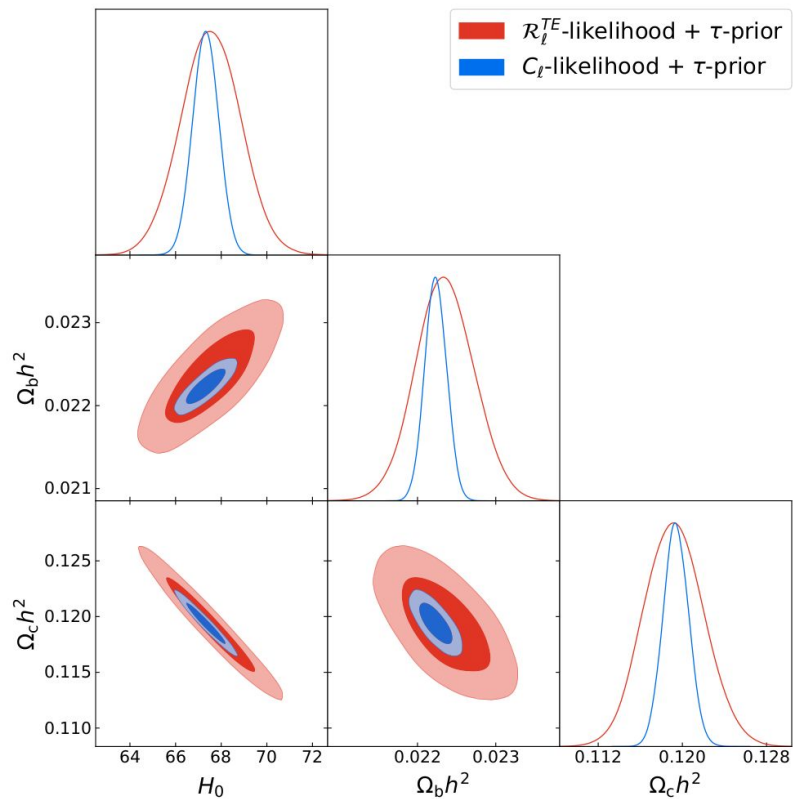


Bestfit from Planck TT,TE,EE

Planck correlation coefficient

$$\chi^2/\text{d.o.f} = 52.16/52 \text{ (PTE} = 0.47)$$

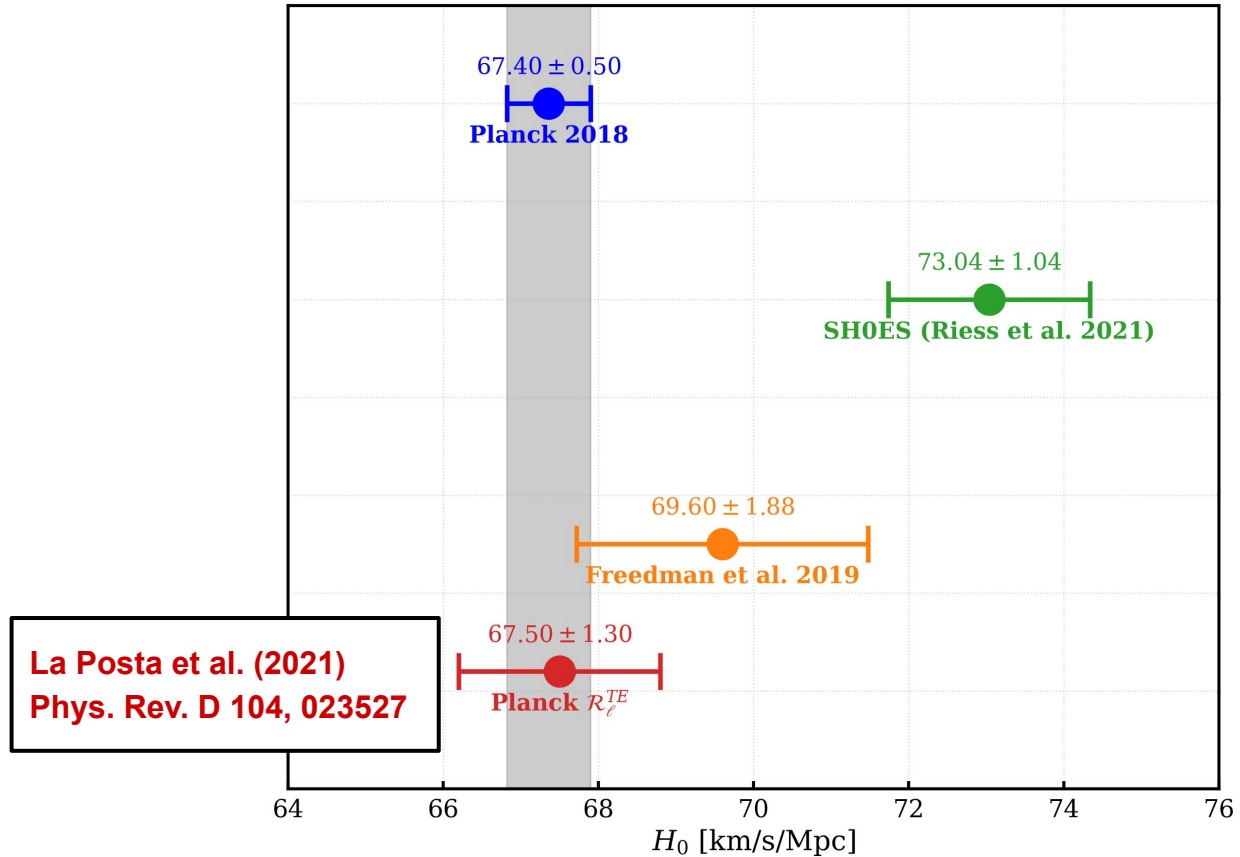
Cosmological results from R^{TE}



3.3 σ away from the latest
SH0ES measurement

$$H_0 = 67.5 \pm 1.3 \text{ [km/s/Mpc]}$$

Hubble tension



Independent measurements of H_0 from the ground

Atacama Cosmology Telescope

6m telescope in the Atacama desert
(Chile ~5000m high)

ACT DR4 (**Choi+ 2020, Aiola+ 2020**)

data collected from 2013 to 2016

Cosmological analysis on ~5400 deg²

observed at 98 and 150 GHz

South Pole Telescope

10m primary mirror
(South Pole ~2800m high)

SPT-3G results (**Dutcher+ 2021**)

4 month period in 2018

Cosmological analysis on ~1500 deg²

observed at 95, 150 and 220 GHz