Accelerator-based coherent lightsources in extreme wavelength range

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Journées Accélérateurs SFP Roscoff, octobre 2023







Curriculum Vitae

2011/09 - 2014/09 Doctorat dans le groupe de dynamique non-linéaire du laboratoire PhLAM à Lille: étude des instabilités spatio-temporelles apparaissant dans les paquets d'électrons relativistes dans les anneaux de stokage.

2014/10 - 2017/04 Post-doc dans le groupe accélérateur du laser à électrons libes (LEL) FERMI en Italie: étude de la dynamique du faisceau d'électrons relativistes et génération de rayonnement cohérent dans les EUV.

2017/05 - 2017/09 Post-doc à Synchrotron SOLEIL sur le projet COXINEL de laser à électrons libres sur accélération plasma: génération de rayonnement LEL à partir de faisceaux d'électrons générés par accélération laser-plasma.

2017/10 - Present Chargée de Recherche CNRS au laboratoire PhLAM.

2017 International Young FEL prize 2022 Médaille de Bronze CNRS

Research group

PhLAM

part of DYSCO team (DYnamique des Systèmes COmplexes)

- Serge Bielawski
- Christophe Szwaj
- Clément Evain (Prix Jean-Louis Laclare 2017)
- Marc Le Parquier (IR CERLA)
- Christelle Hanoun (PhD student 3rd year)
- Quentin Demazeux (PhD student 2nd year)

• In collaboration with several accelerator teams:

- ▷ Synchrotron SOLEIL, France
- KARA, Karlsruhe Institute of Technology, Germany
- \triangleright DESY, Hamburg, Germany
- ▷ FERMI, Trieste, Italy
- ▷ ELBE/HZDR, Dresden, Germany
- COXINEL project, (ERC Sycnhrotron SOLEIL) France
- TWAC project (EIC IJCLab), France

Accelerator-based lightsources: Storage Rings vs. Free-Electron Lasers

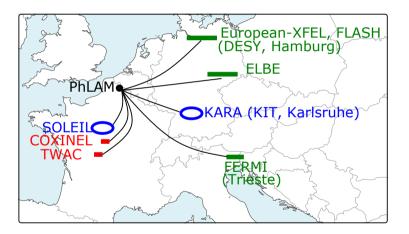
Storage rings: e.g. Synchrotron SOLEIL (France), KARA (Karlsruhe, Germany)...



Free-Electron Lasers: e.g. FERMI (Trieste, Italy), European XFEL (Hamburg, Germany)...



 \rightarrow Relativistic e-beam of few GeV — Radiation from THz to X-rays.



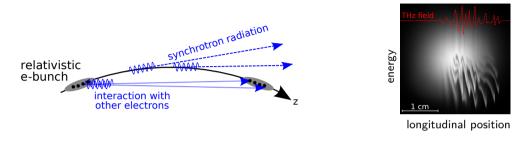
Projects:

- CNRS Momentum
 METEOR
- ANR ULTRASYNC
- EIC TWAC

Collaborations on:

- Ultra-fast electron bunch and THz diagnostics
- Electron bunch dynamics
- New accelerator generations
- Coherent sources

Relativistic electron beam behind synchrotron radiation emission



BUT... microbunching instability

- interaction between e^- and their radiation \Rightarrow microbunching instability
- Formation of microstructures (from mm to μ m scale)
 - \Rightarrow source of intense coherent THz radiation
 - \Rightarrow degradation of electron beam properties
- Irregular evolution in space and time

⇒ \odot Major limitation for the operation of coherent FELs at ultra short wavelengths. ⇒ \odot New source of coherent radiation in the THz domain in storage rings.

PhD thesis: modeling of the electron bunch dynamics

Equation Vlasov-Fokker-Planck (1D)

 $\frac{\partial f}{\partial \theta} - p \frac{\partial f}{\partial q} + \left[q - l_c E_{wf} \right] \frac{\partial f}{\partial p} = 2\varepsilon \frac{\partial}{\partial p} \left(pf + \frac{\partial f}{\partial p} \right)$ 80 1 0.5 0 -0.2 THz signal (a.u.) 0 50 100 150 200 ó Slow time (µs) d energy

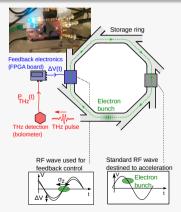
longitudinal position q

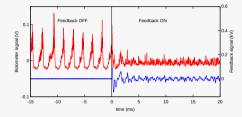
Simulations: SOLEIL @I = 15 mA, ≈ 20 min/32 proc., national GENCI center 7

Control of the microbunching instability at Synchrotron SOLEIL

[**Evain**, Szwaj, Roussel, Rodriguez, Le Parquier, Tordeux, Ribeiro, Labat, Hubert, Brubach, Roy, Bielawski, Nature Physics **15**, 635 (2019)]

Feedback control inspired from chaos control strategy

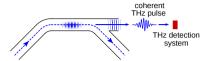




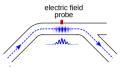
- Apparently simple: required additional hardware: \$400 FPGA board
- Involved strategy borrowed from chaos control theory: the so-called OGY method [Ott, Grebogi & Yorke, PRL, 64, 1196 (1990)]
- Stabilization of a pre-existing solution: major consequence: the required power \rightarrow 0 once the transient have disappeared (power involved at SOLEIL is a fraction of MEGAWATT !)

PhD thesis: observation challenges

Far field: detection of coherent THz radiation

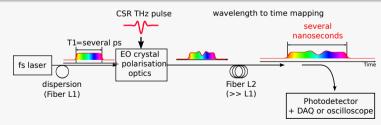


Near field: detection of electron bunch shape



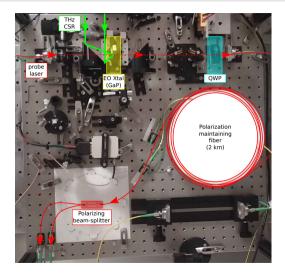
Needs of: (i) speed (sub-ps resolution), (ii) single-shot operation, (iii) >MHz rep. rate, ...

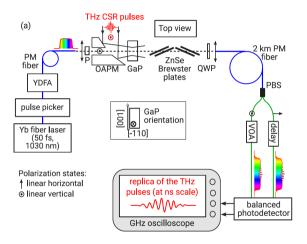
The "PhLAM strategy": electro-optic sampling [Zhang et al., APL (1998)] + photonic time-stretch [Jalali et al., Electronics Letters (1998)]



Detection of a "copy" of the THz pulse slowed down by a factor $M = 1 + L_2/L_1$. (Typical: $L_1 = 10$ m, $L_2 = 2$ km $\rightarrow M \sim 200$, i.e. 5 GHz on the oscilloscope $\equiv 1$ THz at the input).

Time-stretch electro-optic detection at Synchrotron SOLEIL

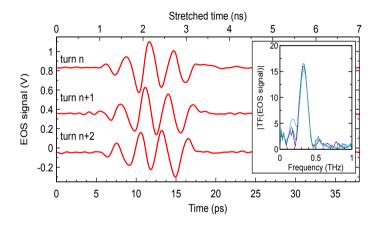


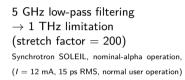


[Roussel et al., Scientific Reports 5, 10330 (2015)]] [Szwaj et al., Rev. Sci. Instrum. 87, 103111 (2016)], [Evain et al., PRL 118, 054801 (2017)]

Successive single-shot CSR pulses recordings

THz CSR electric field from 1 bunch at every turn (i.e. every $pprox \ \mu$ s)



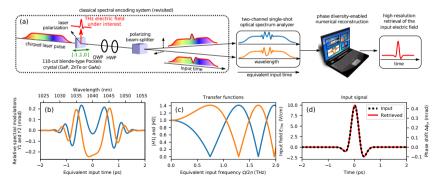


CNRS Momentum METEOR: Microscopie temporelle d'objets relativistes

METEOR solves a 20 year old problem on the temporal resolution of single-shot EO detection of electric field:

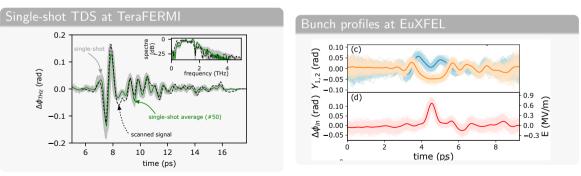
$$au = \sqrt{t_{\textit{laser}} imes t_{\textit{window}}}$$

e.g. $t_{laser} = 100$ fs and $t_{window} = 10$ ps $\rightarrow \tau = 1$ ps » t_{laser} \odot \rightarrow [Roussel, et al. Phase Diversity Electro-optic Sampling: A new approach to single-shot terahertz waveform recording. Light Sci Appl **11**, 14 (2022)]



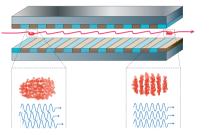
Applications of phase-diversity electro-optic sampling

- ${\scriptstyle \bullet}\,$ TeraFERMI –> low rep. rate, high peak power coherent THz radiation on FEL
- $\bullet~{\sf EuXFEL}$ –> access to single shot electron bunch profile with near field setup
- Synchrotron SOLEIL -> high rep. rate with high average THz power (access to the microbunching dynamics)
- FELBE -> THz FEL at high rep. rate (access to the enveloppe and carrier of the radiation !)
- ... Opens door to single-shot Time-Domain Spectroscopy



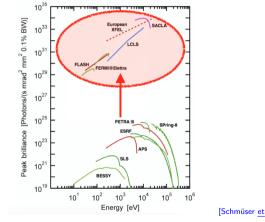
Research activities on Single-pass Free-Electron Laser: collaboration with FERMI (Italy)

Principle: amplification, up to saturation, of the radiation produced by a relativistic electron beam travelling in an periodic magnetic field (i.e. undulator)



[McNeil & Thompson, Nat. Photon. 239 (2010)]

- $\, \bullet \,$ relativistic electrons: energy $\sim \, {\rm GeV}$
- high quality ebeams
- $\bullet\,$ high peak current $\sim\,$ kA
- ${\scriptstyle \bullet} \,$ bunch length $\sim 1-100 \; {\rm fs}$
- $\bullet \ {\rm peak} \ {\rm power} \sim {\rm GW}$

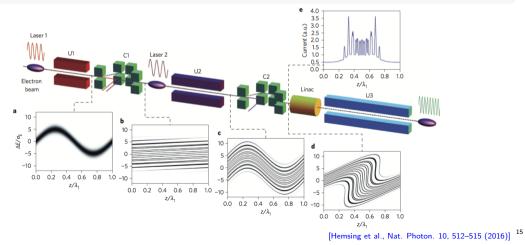


al., Free-Electron Lasers in the Ultraviolet and X-Ray Regime, Springer, (2014)]

Challenge: generation of coherent, ultra-short pulse in the x-rays domain

The beam-echo effect: Echo-Enabled Harmonic Generation (EEHG)

Strongly nonlinear harmonic up-conversion process based on a two-seed laser interaction $\rightarrow 1/\lambda_R = n/\lambda_1 + m/\lambda_2$ [G. Stupakov, Phys. Rev. Lett. 102, 074801 (2009)]



more beamlets in the phase-space \rightarrow higher harmonic in the bunching factor \odot

$\mathsf{Benefits} \text{ of } \mathsf{EEHG}$

- small energy modulation needed,
- from UV to soft x-ray in one stage,
- tunable,
- less sensitive to ebeam quality, ...

Challenges

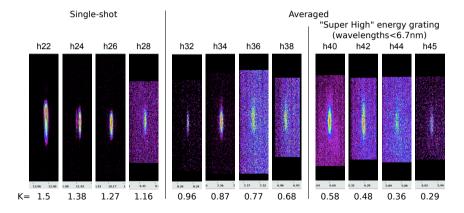
- preservation of fine phase space structure ?
- sensitivity to intrabeam scattering, diffusion and laser quality ?

Experimental demonstration of EEHG: time-lapse

- 2009 first theory of beam-echo effect [G. Stupakov, Phys. Rev. Lett. 102, 074801 (2009)], [D. Xiang and G. Stupakov, Phys. Rev. STAB, 12, 030702 (2009)]
- 2010 3rd harmonic observation + coherent emission (<u>NO FEL !</u>) [D. Xiang et al., Phys. Rev. Lett. 105, 114801 (2010))]
- 2012 3rd harmonic + amplification [Z. T. Zhao et al., Nature Photonics 6, 360–363 (2012)]
- 2014 15th harmonic (NO FEL !) [E. Hemsing et al., Phys. Rev. STAB 17, 070702 (2014)]
- 2016 75th harmonic (NO FEL !) [E. Hemsing et al., Nature Photonics 10, 512–515 (2016)]
- \rightarrow still no FEL based on EEHG in the soft x-ray domain $\ref{eq:field}$
 - 2017 calculation of EEHG @FERMI in the soft x-ray domain [P. R. Rebernik, E. Roussel, G. Penn, G. De Ninno, L. Giannessi, G. Penco, E. Allaria, Echo-Enabled Harmonic Generation Studies for the FERMI Free-Electron Laser. Photonics 2017, 4, 19]

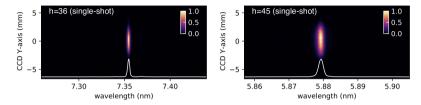
EEHG @FERMI: medium harmonic range

• "low-energy" beam (\approx 900 MeV, i.e. low gain for h > 30)

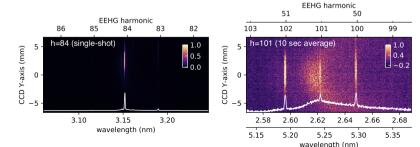


EEHG @FERMI: high harmonics

 \bullet with e-beam at max. energy (≈ 1.5 GeV), very high harmonics up to $\underline{101}$

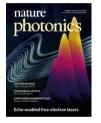


using high sensitivity EUV CCD from Andor:



FERMI FEL collaboration

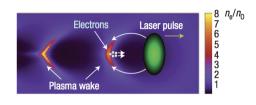
- [2019-2021] MOU between FERMI and PhLAM
 - \rightarrow FEL physics: demonstration of EEHG



[P. Ribic et al, Nat. Photonics (2019)]

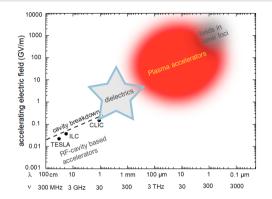
- $\rightarrow\,$ electron beam dynamics: study of microbunching instability
- [2019-2021] CNRS Momentum METEOR: application to TeraFERMI
- [2021-2023] Elettra Distinguished Young Scientist
 - ightarrow study and characterization of the microbunching instability and its impact on the FEL radiation properties
 - $\rightarrow\,$ development of FERMI-FEL for the generation of harmonics below 2 nm.

Towards compact accelerator: Laser-plasma accelerators - Dielectric accelerators



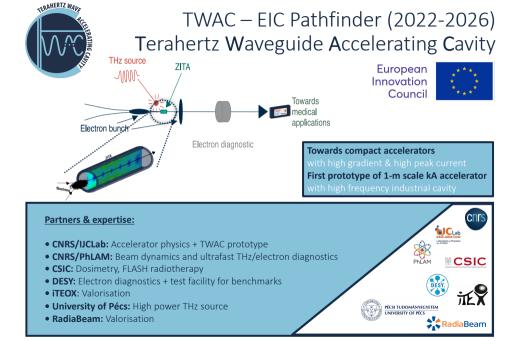
• Accelerating electric field in ionized plasma:

$$E_0(\mathrm{V/m})\simeq96\sqrt{n_0(\mathrm{cm}^{-3})}$$



plasma density of $n_0 \approx 10^{18} \text{ cm}^{-3}$ $\rightarrow E_0 \simeq 100 \text{ GV/m} \gg 10' \text{ s MV/m}$ in conventional radio-frequency (rf) linear accelerators (LINACs) + potential to produce extremely short electron bunches., $\tau_b \ll 100$ fs.

[E. Esarey, et al., RevModPhys.81.1229 (2009)], [B. Hidding, et al., arXiv:1904.09205 (2019)]



Le projet COXINEL: de l'accélération laser-plasma vers un laser à électrons libres injecté par un laser externe

Eléonore Roussel on behalf of the COXINEL team

Univ. Lille, CNRS, UMR 8523 - PhLAM - Physique des Lasers, Atomes et Molécules, Lille, FRANCE

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Towards compact Free-Electron Lasers

2 The COXINEL project



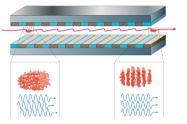




Single-pass Free-Electron Laser

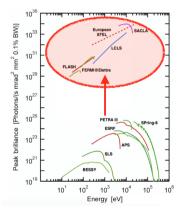
Principle amplification, up to saturation, of the radiation produced by a relativistic electron beam travelling in a periodic magnetic field (i.e. undulator)

Radiation resonance condition: $\lambda_R = \frac{\lambda_u}{2\gamma^2} \left(1 + \frac{\kappa^2}{2}\right)$



[McNeil & Thompson, Nat. Photon. 239 (2010)]

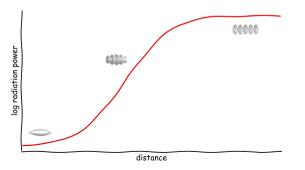
- ${\, \bullet \,}$ relativistic electrons: energy $\sim \, {\rm GeV}$
- high quality ebeams
- $\bullet\,$ high peak current $\sim\,$ kA
- $\bullet\,$ bunch length $\sim 1-100~\text{fs}$
- $\bullet \ \text{peak power} \sim \text{GW}$



[Schmüser <u>et al.</u>, Free-Electron Lasers in the Ultraviolet and X-Ray Regime, Springer, (2014)]

Free-electron laser amplification

- ullet electron beam / optical wave interaction o energy exchange
- microbunching at radiation wavelength $\lambda_R \rightarrow$ coherent emission
- $\bullet\,$ coherent emission exponentially amplified $\rightarrow\,\mathsf{FEL}$
- ${\, \bullet \,}$ loss of energy, increase of energy spread ${\, \rightarrow \,}$ out-of-resonance, saturation



- FEL Pierce parameter $ho \propto \left(I/(\sigma_x \sigma_y)\right)^{1/3}$
- Power $P \propto \exp{(z/L_g)}$
- Gain length $L_g \propto \left(1+\sigma_{\delta}^2/
 ho^2
 ight)/
 ho$
- Saturation length $L_s \approx 22 \times L_g$
- For high-gain FEL: $\sigma_{\delta} \ll \rho$

[K.J. Kim and M. Xie, NIMA 331, 359 (1993)]

Some existing FEL facilities

Example (1): FERMI (Italy)

- $\bullet~200$ m LINAC +~100 m undulator line
- energy $\sim 1.5~{\rm GeV} \rightarrow \gamma \sim 3000$
- ${\it K}_{\it u} \sim [1-7.5]$ and $\lambda_{\it u} = 55/35$ mm
- ullet ightarrow $\lambda_R \sim$ [4 100] nm

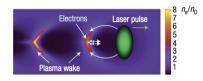
Example (2): LCLS (USA)



- 1 km LINAC + 132 m undulator line
- $K_u = 3.5$ and $\lambda_u = 30$ mm
- energy \sim 14 GeV $\rightarrow \gamma \sim$ 28000
- $\rightarrow \lambda_R = 1.4$ Å



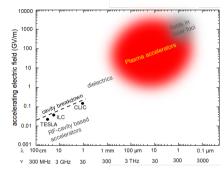
Towards compact accelerator: Laser-plasma accelerators (LPAs)



• Accelerating electric field in ionized plasma:

$$E_0({
m V/m})\simeq 96\sqrt{n_0({
m cm}^{-3})}~~{
m with}~n_0pprox 10^{18}{
m cm}^{-3}~{
m sm}^{-3}$$

 \rightarrow $E_0 \simeq 100~GV/m \gg 10's~MV/m$ in conventional RF LINAC + potential to produce extremely short e- bunches < 100 fs.



[E. Esarey, et al., RevModPhys.81.1229 (2009)], [B. Hidding, et al., arXiv:1904.09205 (2019)]



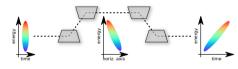
LPA-based FEL: strategies to fulfil the requirement $\sigma_{\delta} \ll \rho$ for high-gain FEL

transverse-gradient undulator

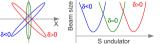


[Z. Huang, et al., Phys. Rev. Lett. 109, 204801 (2012)]

demixing chicane

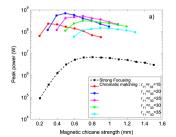


• super-matching



[A. Loulergue et al., New Journal of Physics 17, 023028 (2015)]

[M.E. Couprie, et al., J. of Phys. B: AMO Phys. **47**, 234001 (2014)], [A.R. Maier et al., Phys. Rev. X **2**, 031019 (2012)]

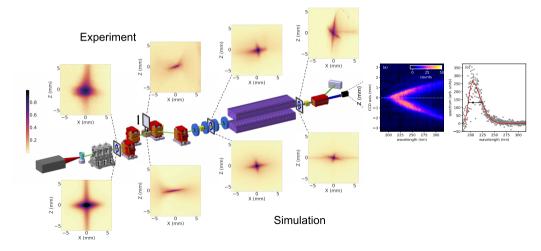


7

COXINEL: Towards a LPA-based FEL demonstrator

COherent X-ray source INferred from Electrons accelerated by Laser

 \rightarrow Free Electron Laser demonstrator using an electron beam produced by laser-plasma acceleration.



COXINEL collaboration

- ERC Advanced Grant COXINEL, PI: M.E. Couprie (Synchrotron SOLEIL) 2014-2019
- COXINEL: collaboration SOLEIL LOA PhLAM
 - → electron beam transport
 [T. André et al., Nature Communications 9, 1334 (2018)]
 [D. Oumbarek Espinos et al., Plasma Phys. Control. Fusion 62, 034001 (2020)]
 - → spontaneous emission of undulator radiation

 [A. Ghaith et al., Scientific Reports 9, 19020 (2019)]
 [E. Roussel et al., Plasma Phys. Control. Fusion 62 074003 (2020)]
 - → theory on LPA-based FEL [M. Labat, ..., E. Roussel, New J. Phys. 22 013051 (2020)]

but NO FEL... divergence was too high, charge density was too low !

• COXINEL#2 (2021-2022): collaboration SOLEIL - HZDR (Dresden) + PhLAM, LOA

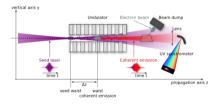
 $\rightarrow\,$ demonstration of first seeded FEL amplification based on LPA [M. Labat et al., Nature Photonics 17, 150-156 (2023)]

Wavelength (nm) Wavelength (nm) 158 162 158 162 (a) (b) 0.5 0.5 (pe y (mm) SEED 0.0 0.0 -0.5 0 -0.52000 2e-;

FAR-FIELD

NEAR-FIELD

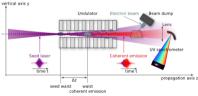
Typical observable: **spatio-spectral distribution** on a 2D UV spectrometer



[M. Labat, ..., E. Roussel, New J. Phys. 22 013051 (2020)]

NEAR-FIELD FAR-FIELD Wavelength (nm) Wavelength (nm) 158 162 158 162 (a) (b) 0.5 -0.5 (pe (mm) SEED 0.0 0.0 ž g -0.5 -0.5 2e-2 2000 0.5 -0.5 (mrad) y (mm) SR 0.0 0.0 -0.5 0 -0.5 UV

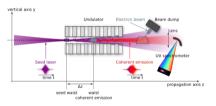
Typical observable: spatio-spectral distribution on a 2D UV spectrometer



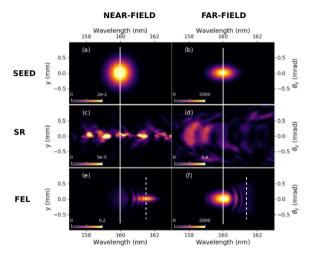
NEAR-FIELD FAR-FIELD Wavelength (nm) Wavelength (nm) 158 162 158 162 (a) (b) 0.5 0.5 (mrad) Typical observable: y (mm) SEED 0.0 0.0 spatio-spectral distribution -0.5 0 -0.5 on a 2D UV spectrometer 2e-2 (d) vertical axis v 0.5 0.5 Electron beam - Beam dump Undulator (mrad) y (mm) SR 0.0 0.0 -0.5 0 -0.5 Coherent emission Seed laser (e) time time t 0.5 - 0.5 (mrad) propagation axis z y (mm) Δz seed waist waist FEL 0.0 0.0 coherent emission -0.5 0 -0.5 2000 160 162 158 160 162 158 Wavelength (nm) Wavelength (nm)

[M. Labat, ..., E. Roussel, New J. Phys. 22 013051 (2020)]

Typical observable: **spatio-spectral distribution** on a 2D UV spectrometer

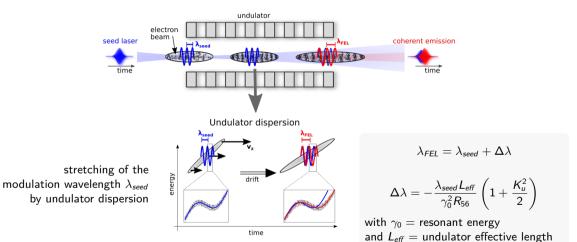


- FEL emission is **red-shifted** w.r.t. seed wavelength
- Presence of interference fringes



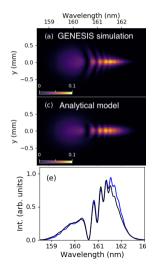
[M. Labat, ..., E. Roussel, New J. Phys. 22 013051 (2020)]

Origin of the red-shifted emission



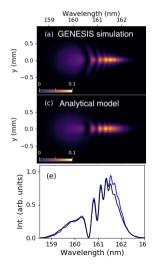
Interference fringes from two coherent pulses

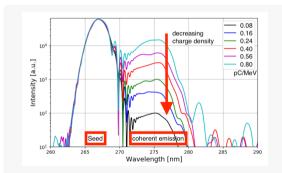
In case of low gain, the FEL pulse is the sum of 2 pulses at \neq wavelength, \rightarrow FEL = seed + coherent emission



Interference fringes from two coherent pulses

In case of low gain, the FEL pulse is the sum of 2 pulses at \neq wavelength, \rightarrow FEL = seed + coherent emission

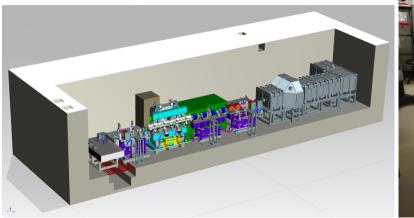




thanks to red-shift + fringes \rightarrow detection of very low FEL signals

COXINEL @ HZDR (Dresden, Germany)

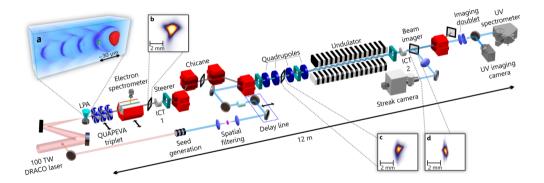
 \rightarrow Oct. 2021: COXINEL beamline moved into 111c LPA cave of HZDR



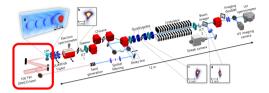


 \rightarrow "A perfect fit"

COXINEL beamline

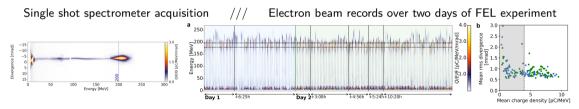


HZDR Laser Plasma Accelerator



- 100 TW-class arm of IR DRACO laser
- Tailored self-truncated ionization-induced injection scheme
- Beam loading to limit energy spread

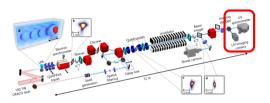
[A. Irman et al., Plasma Phys. Control. Fusion 60, 044015 (2018)],
 [J.P. Couperus et al., Nat. Comm. 8, 487 (2017).]



- Charge density: 5–10 pC/MeV
- Divergence: < 1 mrad-rms
- ${\scriptstyle \bullet }$ Emittance: $\approx 1 \text{ mm.mrad}$
- Stability: >8 hours stable operation + day-to-day reproducible properties

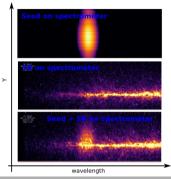
Photon & FEL diagnostics

Spatio-spectral overlap between seed and SR



Streak camera

- Model: FESCA-100 (Hamamatsu)
- Radiation extracted at undulator exit using an Aluminium mirror
- $\rightarrow\,$ Adjustment of seed / SR temporal overlap

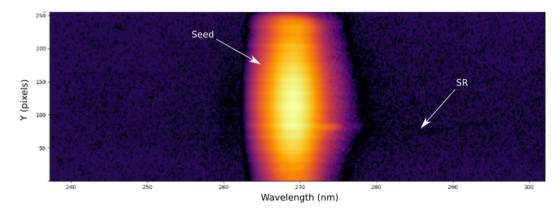


2D UV spectrometer

- iHR320 from Horiba/Jobin-Yvon with a 2D UV-camera
- \rightarrow SR spatio–spectral distribution (Y vs λ)
- $\rightarrow\,$ Seed / SR spatio–spectral alignment

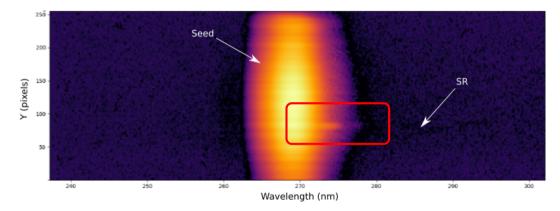
First FEL signal

During the first delay scan...



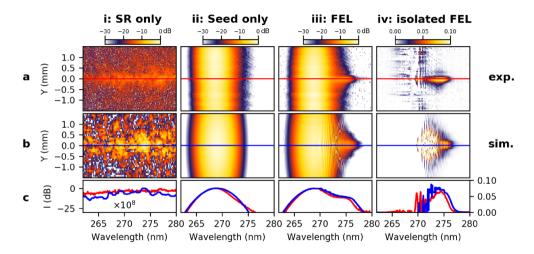
First FEL signal

During the first delay scan...



First FEL signal ! Tiny... but red-shifted (as expected) \rightarrow \odot

After some optimization...

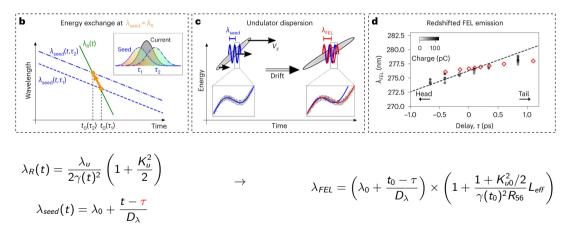


 \rightarrow FEL observation confirmed with shifted by \approx 5 nm (red–shift and fringes as in theory)

[M. Labat et al., Nature Photonics 17, 150-156 (2023)]

FEL spectral control

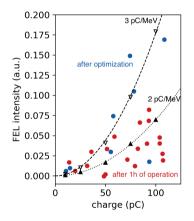
A specificity of chirped seed FEL: spectral control with seed delay



Resonance only occurs at t_0 where $\lambda_R(t_0) = \lambda_{seed}(t_0)$

Coherent emission

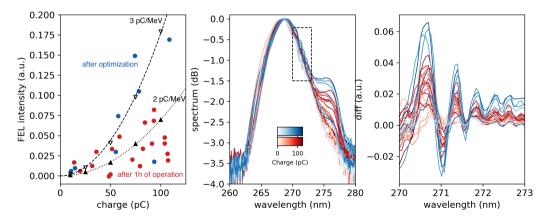
A basic proof of temporal coherence is the quadratic charge dependency of FEL pulse energy.



[M. Labat et al., Nature Photonics 17, 150-156 (2023)]

Coherent emission

- A basic proof of temporal coherence is the quadratic charge dependency of FEL pulse energy.
- **(a)** A strong proof of temporal coherence is **phase-locked interference fringes**.



[M. Labat et al., Nature Photonics 17, 150-156 (2023)]

Conclusion

Achievements on LPA-based FEL

- Several teams have been racing since a decade to demonstrate the feasability of an LPA based FEL
- Recently: 3 teams succeeded in a lapse of 1 year !
 - SASE FEL based on laser plasma accelerator [Wang, W., Feng, K., Ke, L. et al., Nature 595, 516-520 (2021)]
 - SASE FEL based on beam-driven plasma wakefield accelerator [Pompili, R., Alesini, D., Anania, M.P. et al., Nature 605, 659-662 (2022)]
 - seeded FEL based on beam-driven plasma wakefield accelerator [M. Galletti et al. Phys. Rev. Lett. 129, 234801 (2022)]
- ${\scriptstyle \bullet}\,$ The "+" of COXINEL: ${\rightarrow}$ proof of spectral control & temporal coherence !

Next steps

- @ short term
 - ▷ Improve stability and repetition rate of LPA source.
 - ▷ Spectro-temporal reconstruction of the FEL pulses from interference pattern.
- @ long term
 - ▷ FEL towards shorter wavelength.
 - $\,\triangleright\,$ Initial target of COXINEL: 40 nm.

Acknowledgments

