

Euclid: a mission to unravel the Dark Side of the Universe

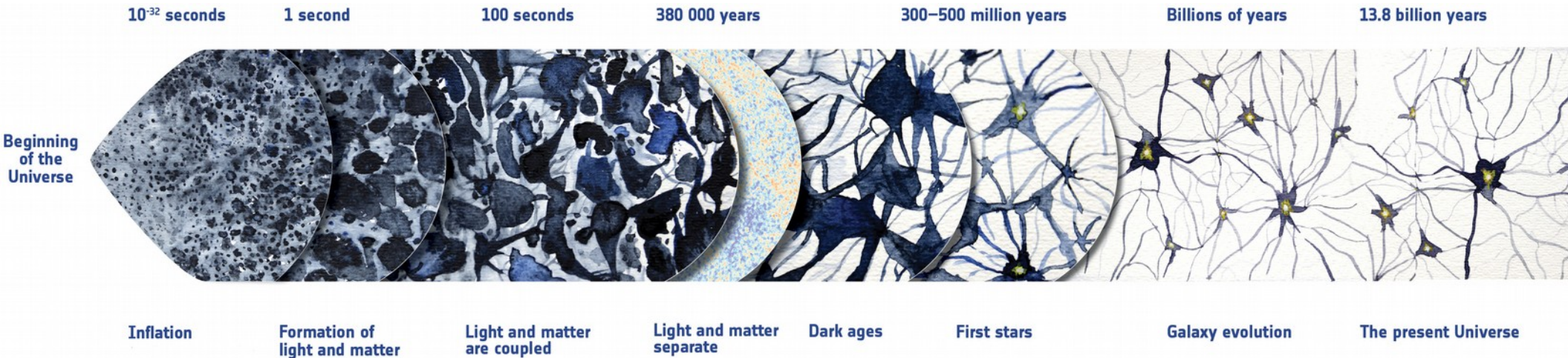
Stéphane Ilić
(IJCLab, France)

Based on the work and efforts of
many members of the Euclid Consortium



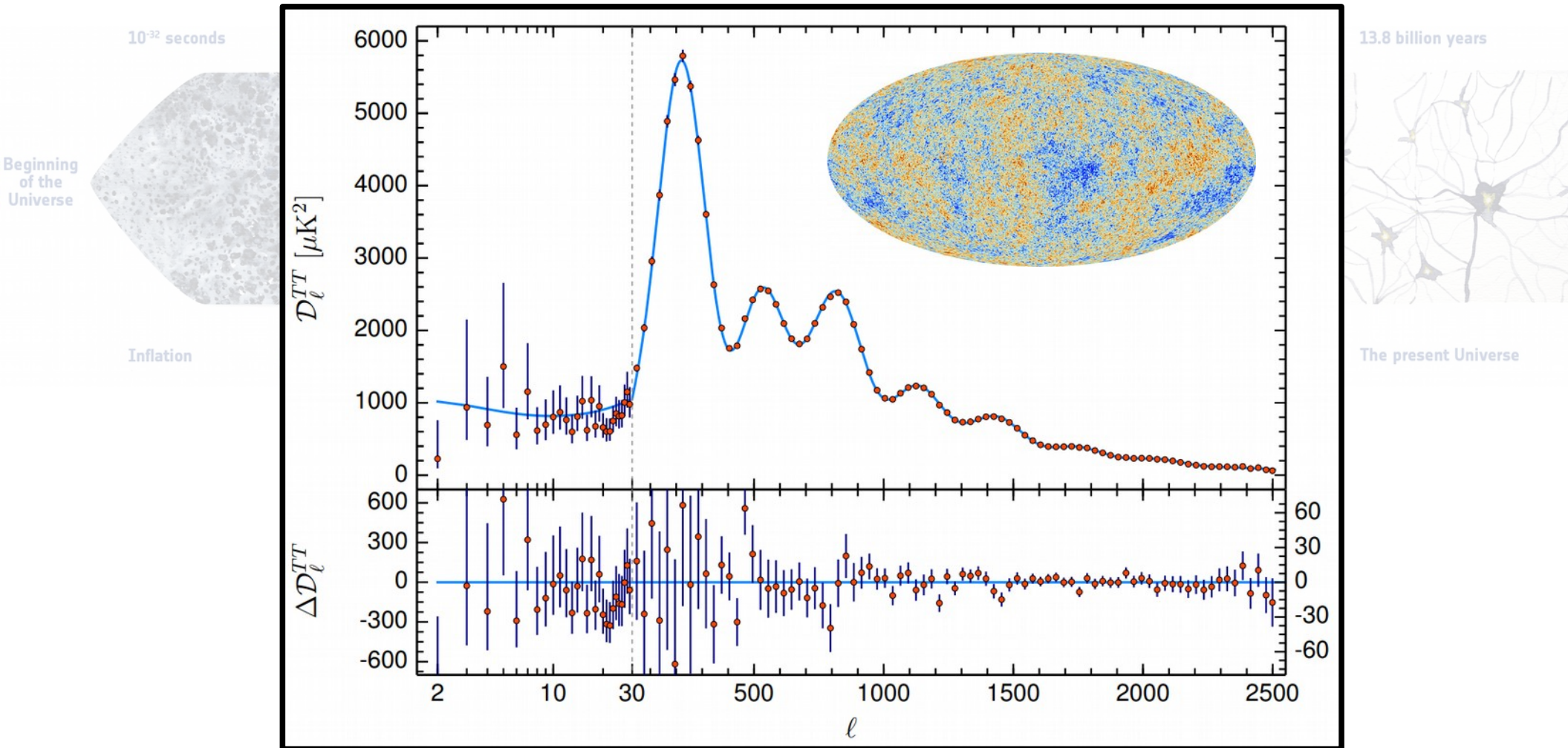
The standard model of cosmology

- The Λ CDM paradigm: a (relatively) simple model, with many successes...



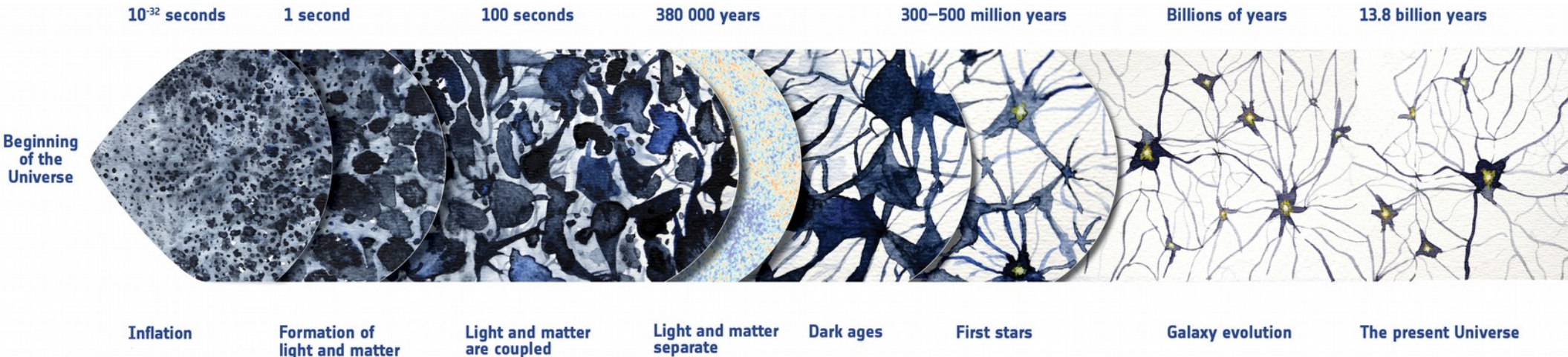
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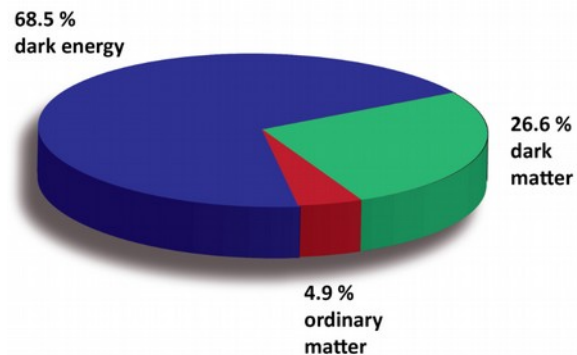
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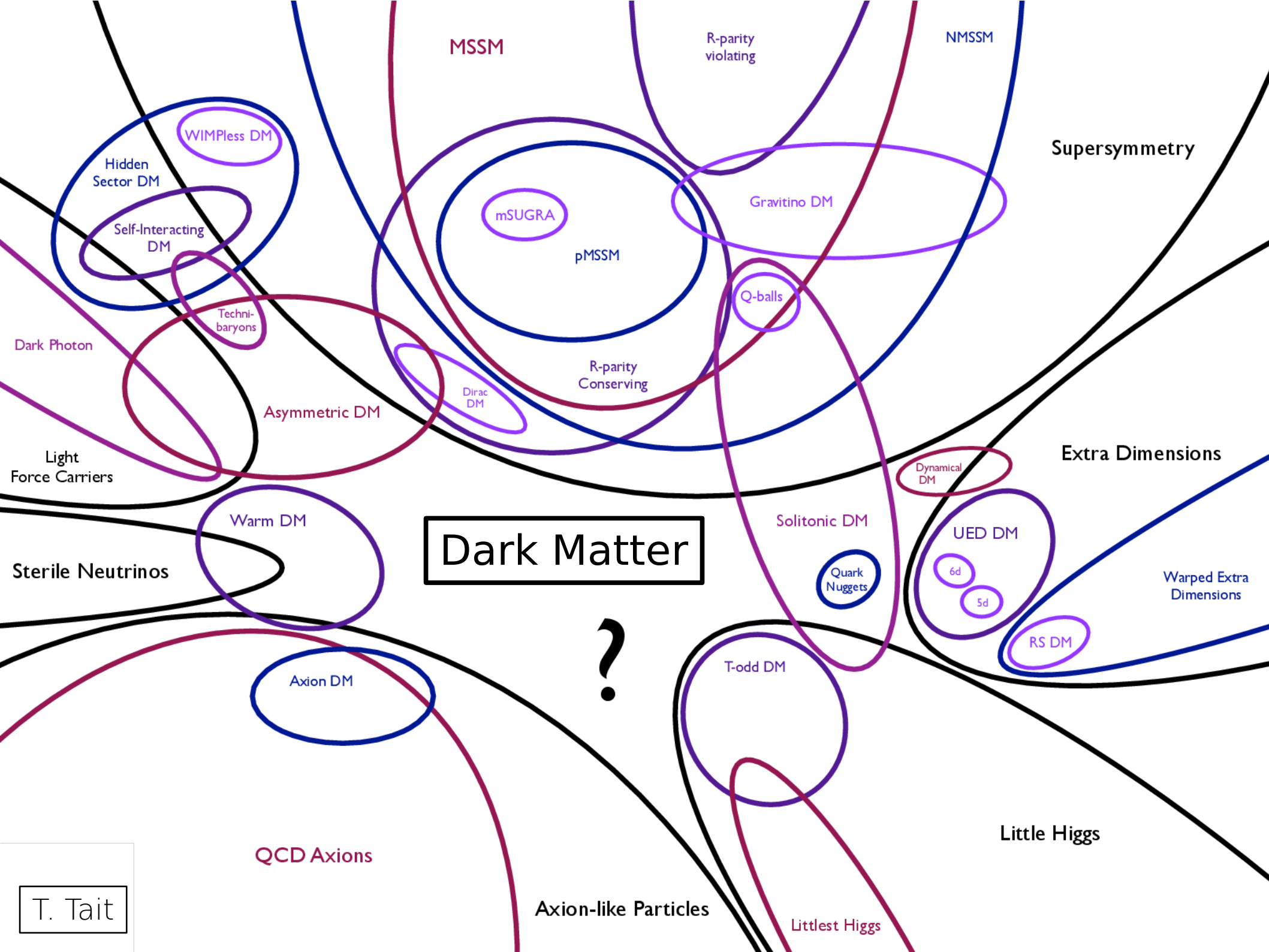
- The Λ CDM paradigm: a (relatively) simple model, with many successes...



- ... but rests on some pillars that are “shrouded in darkness”:

- Primordial Universe, inflation
- Dark matter (“CDM”)
- Dark ages & reionisation
- Dark energy (“ Λ ”)





Dark Matter

?

MSSM

R-parity violating

NMSSM

Supersymmetry

WIMPless DM

Hidden Sector DM

Self-Interacting DM

mSUGRA

pMSSM

Gravitino DM

Q-balls

Techni-baryons

Dark Photon

R-parity Conserving

Dirac DM

Asymmetric DM

Light Force Carriers

Extra Dimensions

Warm DM

Dynamical DM

Solitonic DM

UED DM

Dark Matter

Quark Nuggets

6d

5d

Warped Extra Dimensions

Sterile Neutrinos

RS DM

Axion DM

T-odd DM

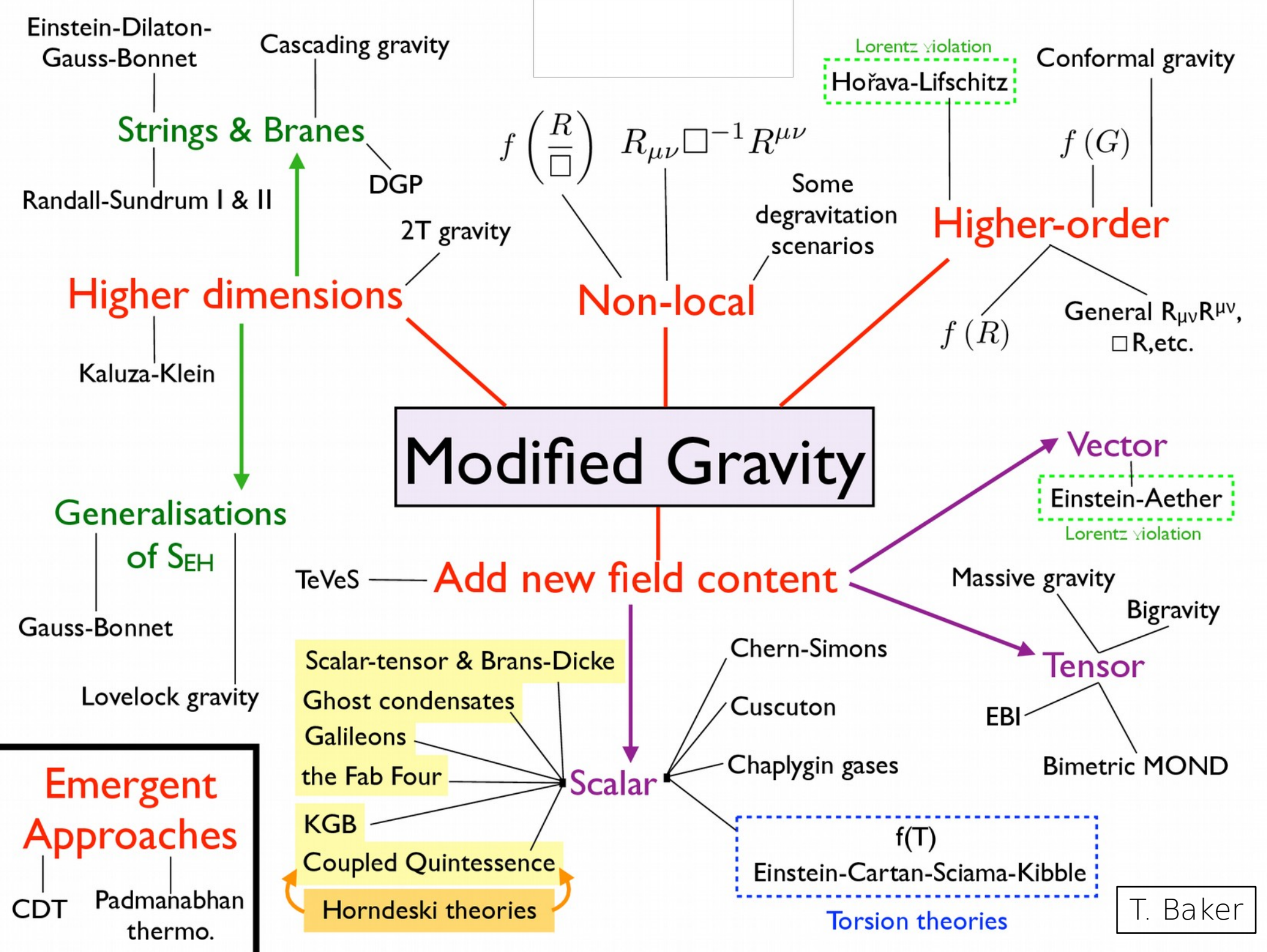
Little Higgs

QCD Axions

Axion-like Particles

Littlest Higgs

T. Tait

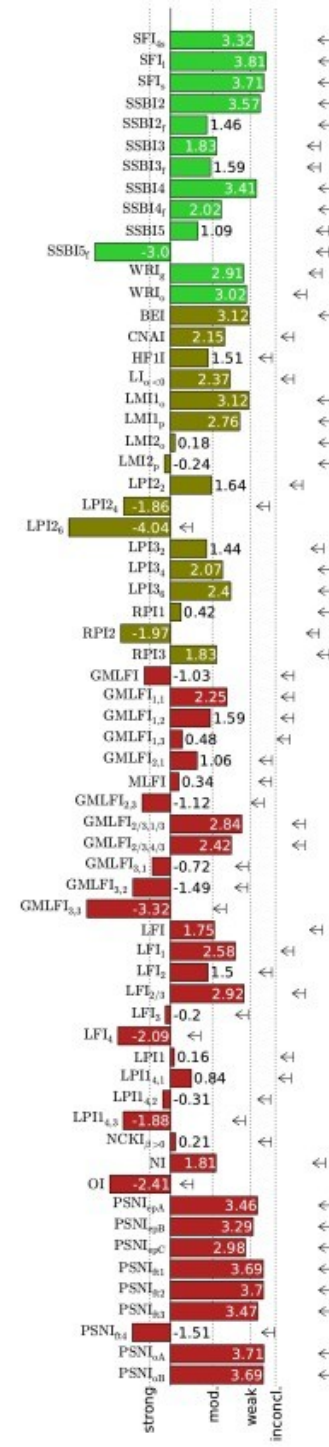


Inflation models

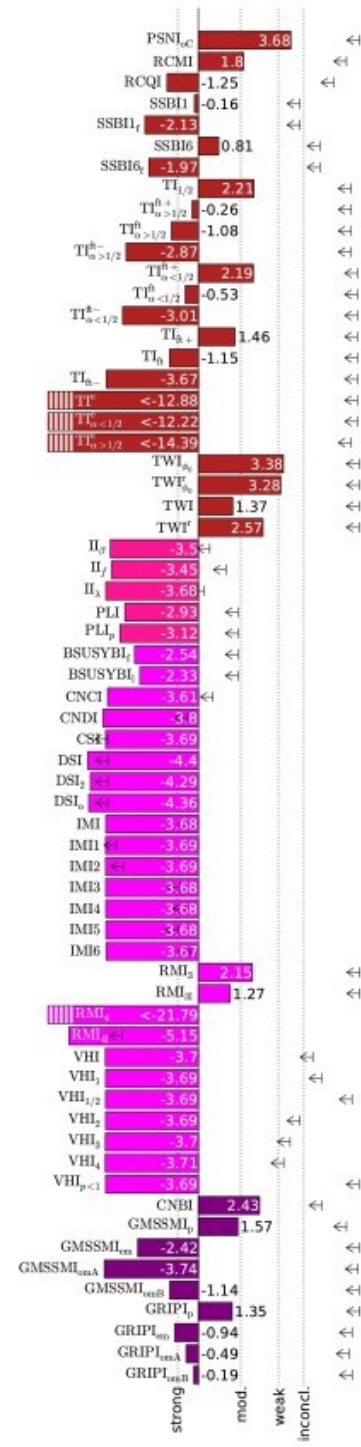


1 1-2 2 2-3 3 1-2-3
ST Classification

Bayesian Evidences $\ln(\mathcal{E}/\mathcal{E}_{SR})$ and $\ln(\mathcal{L}_{max}/\mathcal{E}_{SR})$ for Planck



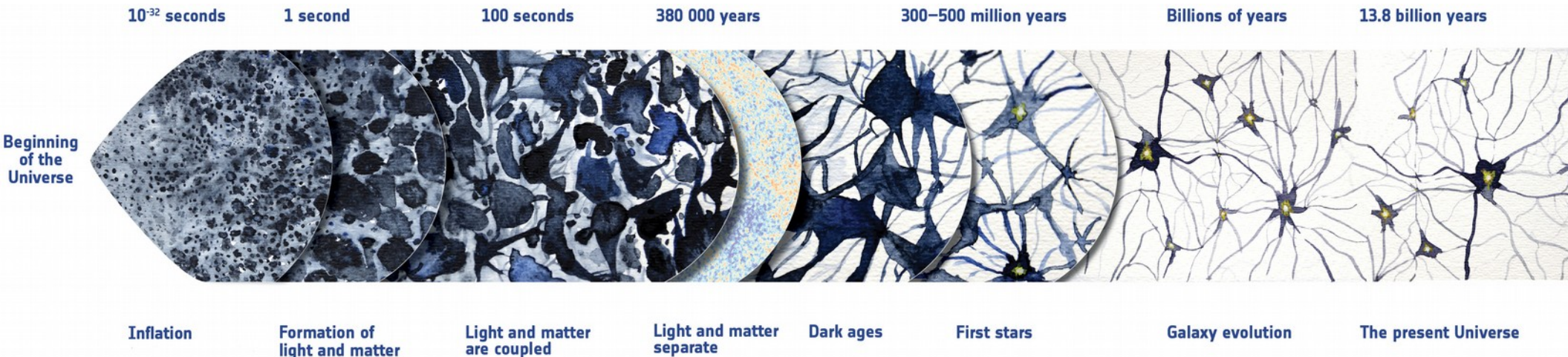
J.Martin, C.Ringeval, R.Trotta, V.Vennin
ASPIC project



Displayed Evidences: 193

The standard model of cosmology

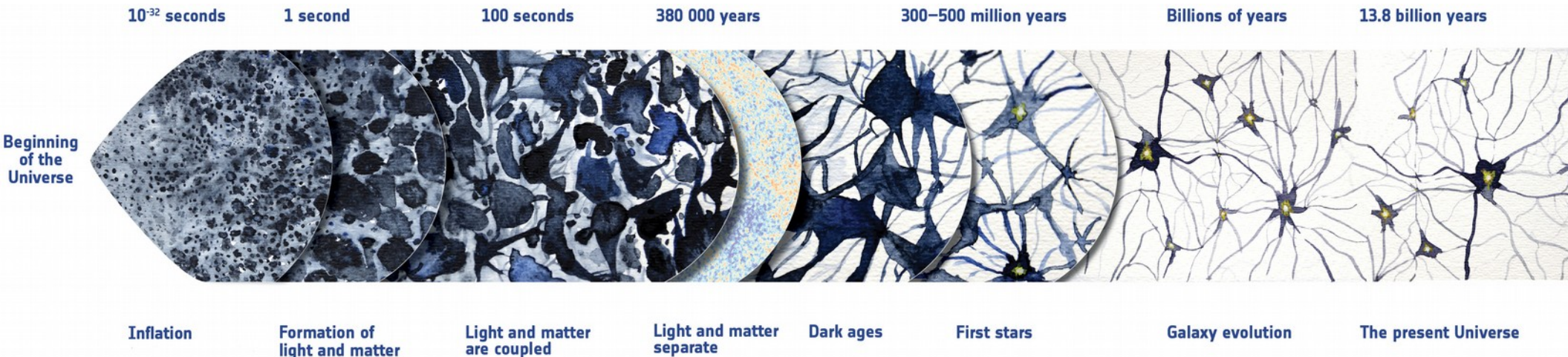
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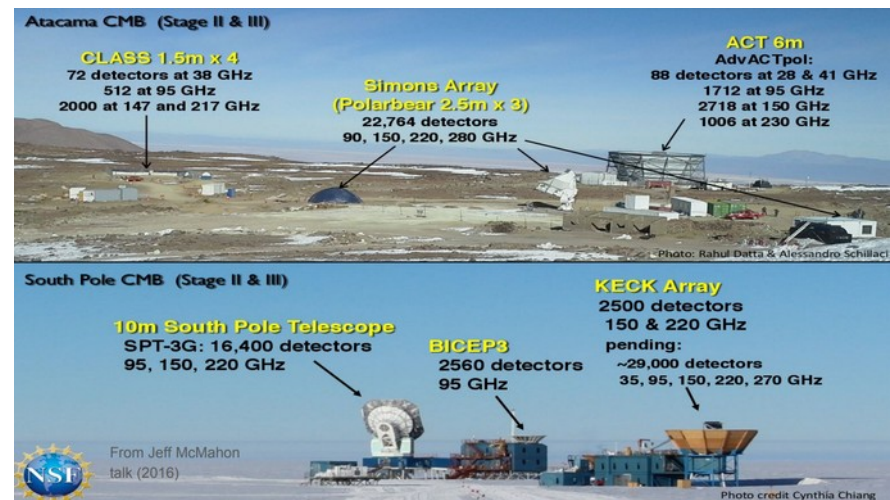
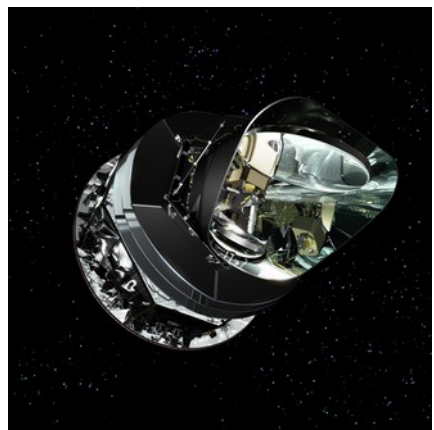
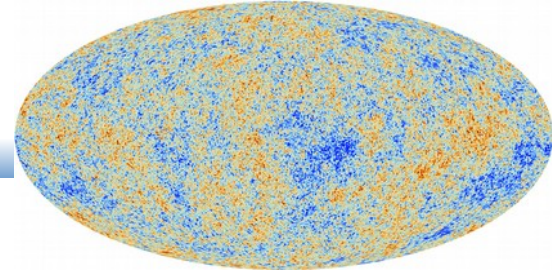
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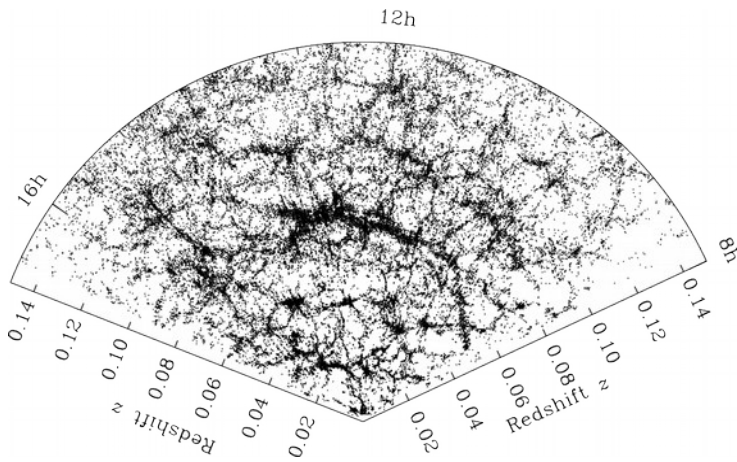
- ... but rests on some pillars that are “shrouded in darkness”:
 - Primordial Universe, inflation
 - Dark matter (“CDM”)
 - Dark ages & reionisation
 - Dark energy (“ Λ ”)
- ... and is shaken by some persistent tensions :
 - H_0 discrepancies
 - σ_8 tensions
 - ISW excesses
 - CMB “anomalies”

Current/upcoming CMB surveys



- Future:
- Simons Observatory
 - LiteBIRD
 - CMB Stage-4

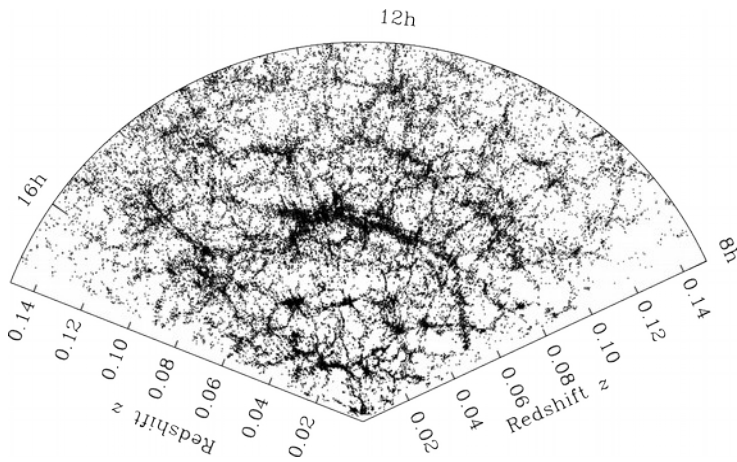
Current/upcoming LSS surveys



DETF classification:

- Stage II: SDSS, KiDS, ...
- Stage III: DES, ...
- Stage IV: DESI, LSST, ...

Current/upcoming LSS surveys



DETF classification:

- Stage II: SDSS, KiDS, ...
- Stage III: DES, ...
- Stage IV: DESI, LSST, ...

The ESA Euclid space telescope mission

→ map the distribution and evolution of galaxies/matter across a large fraction of the observable universe

→ investigate the nature of dark matter and dark energy

“A space mission to map the Dark Universe”

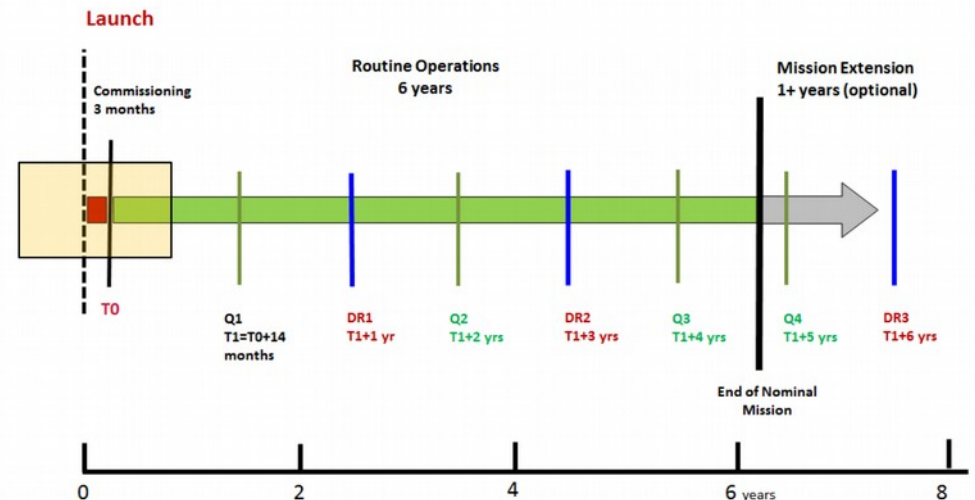
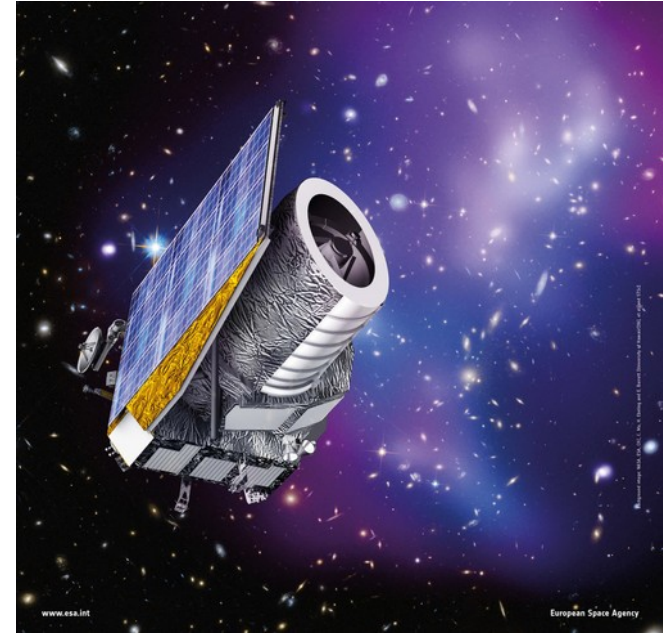
Euclid fact sheet

Technical parameters:

- 1.2 m mirror telescope
- 2 instruments:
 - panoramic visible imager (VIS)
 - near infrared photometer (NISP-P) and slitless spectrograph (NISP-S)

Mission parameters:

- Orbit around L2
- ~15,000 sq. deg., up to $z \sim 2.5$
- Spectro + photo (1.5 B gal) survey
- ~6 years of mission
- Q1 after 17 months, DR1 at 29
- Launch with Soyuz from Kourou
~~with Ariane 6 from Kourou~~
with Falcon 9 from Cape Canaveral
- Launch date: ...



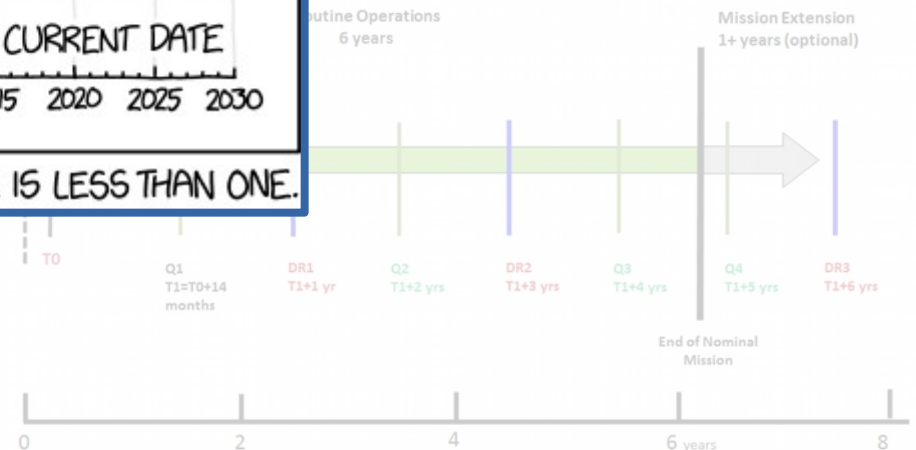
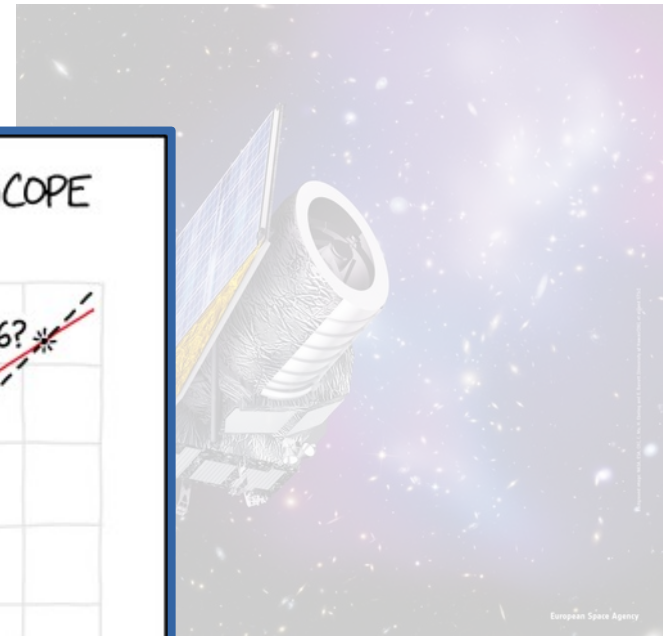
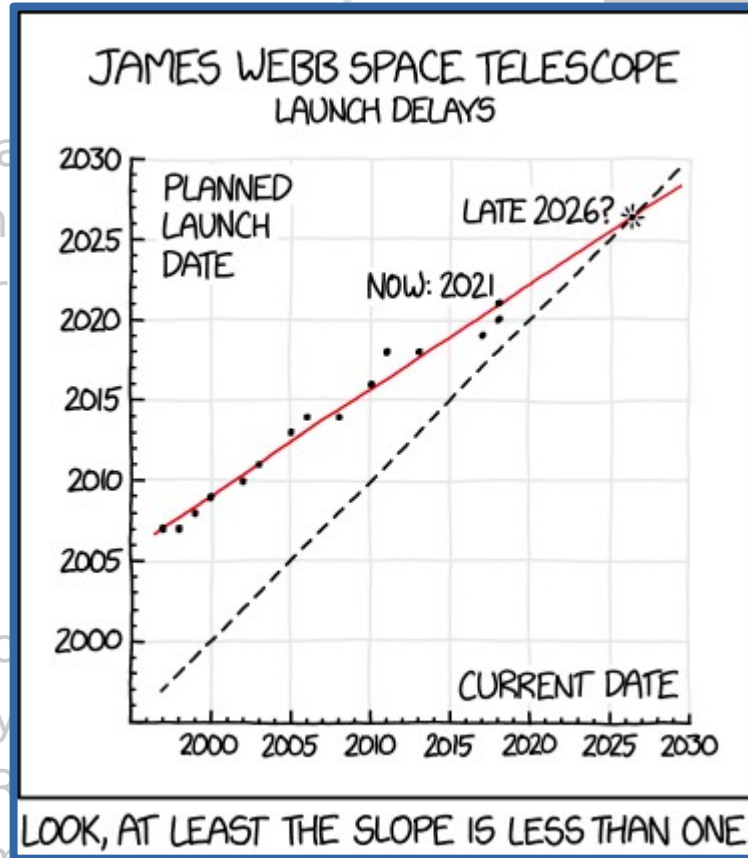
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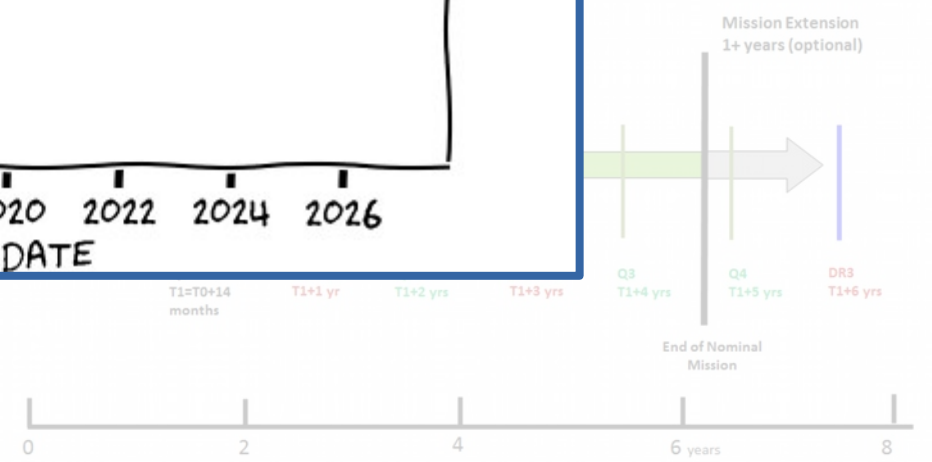
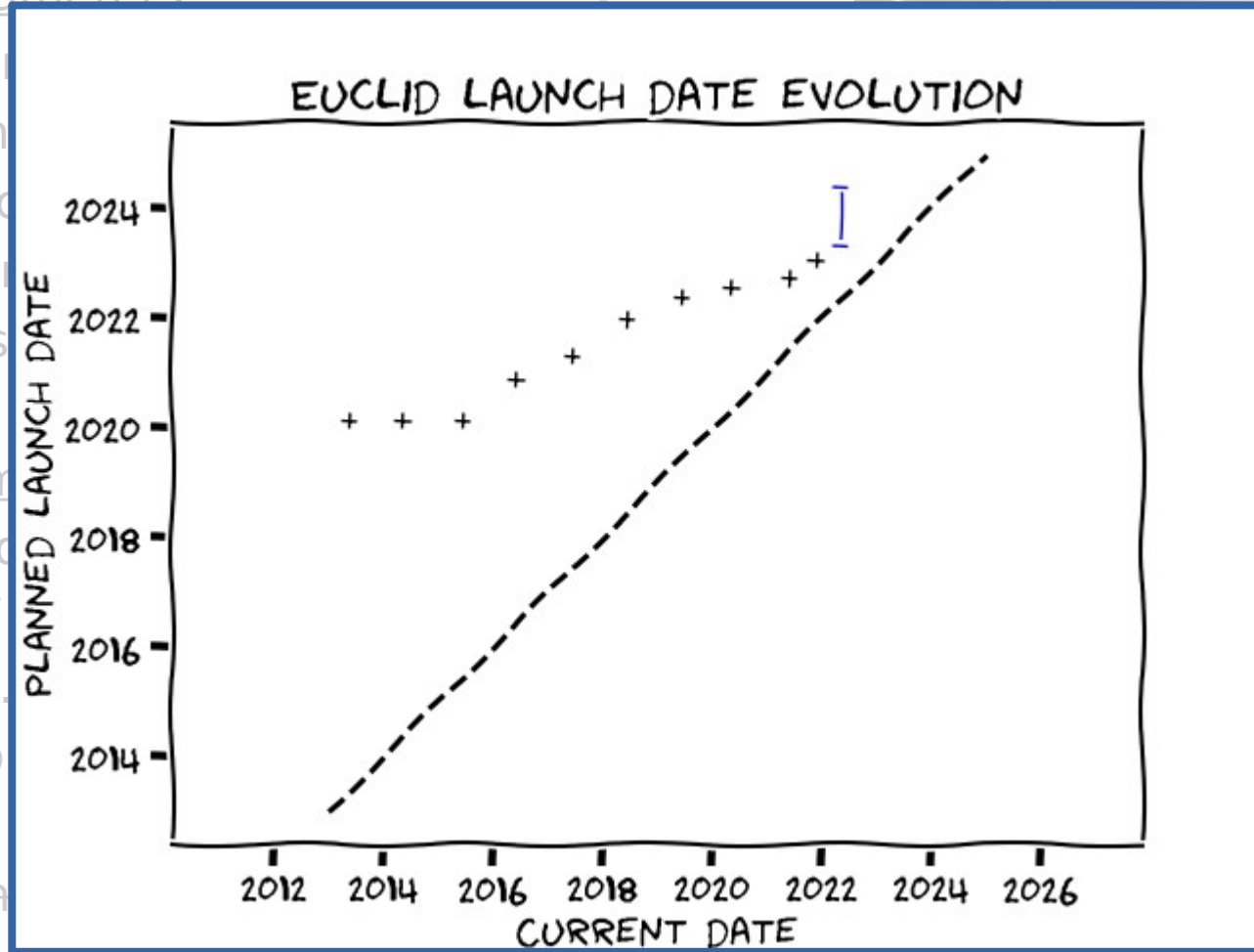
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- panoramic
- near infr
- and slites

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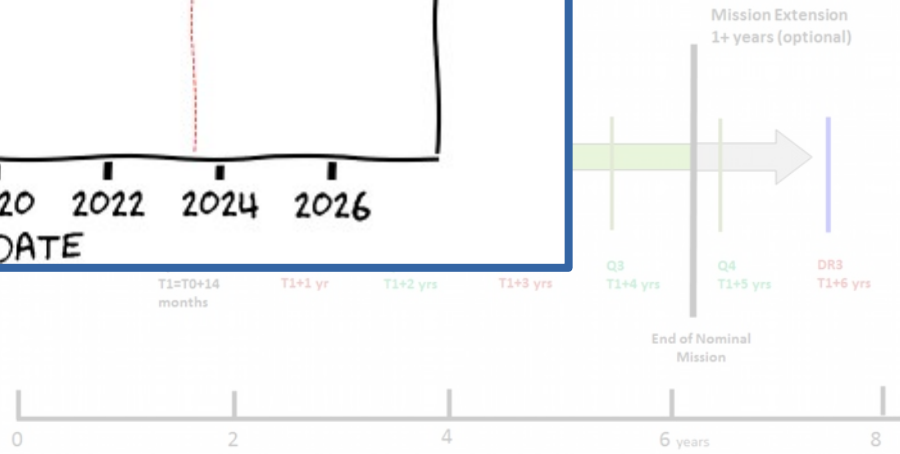
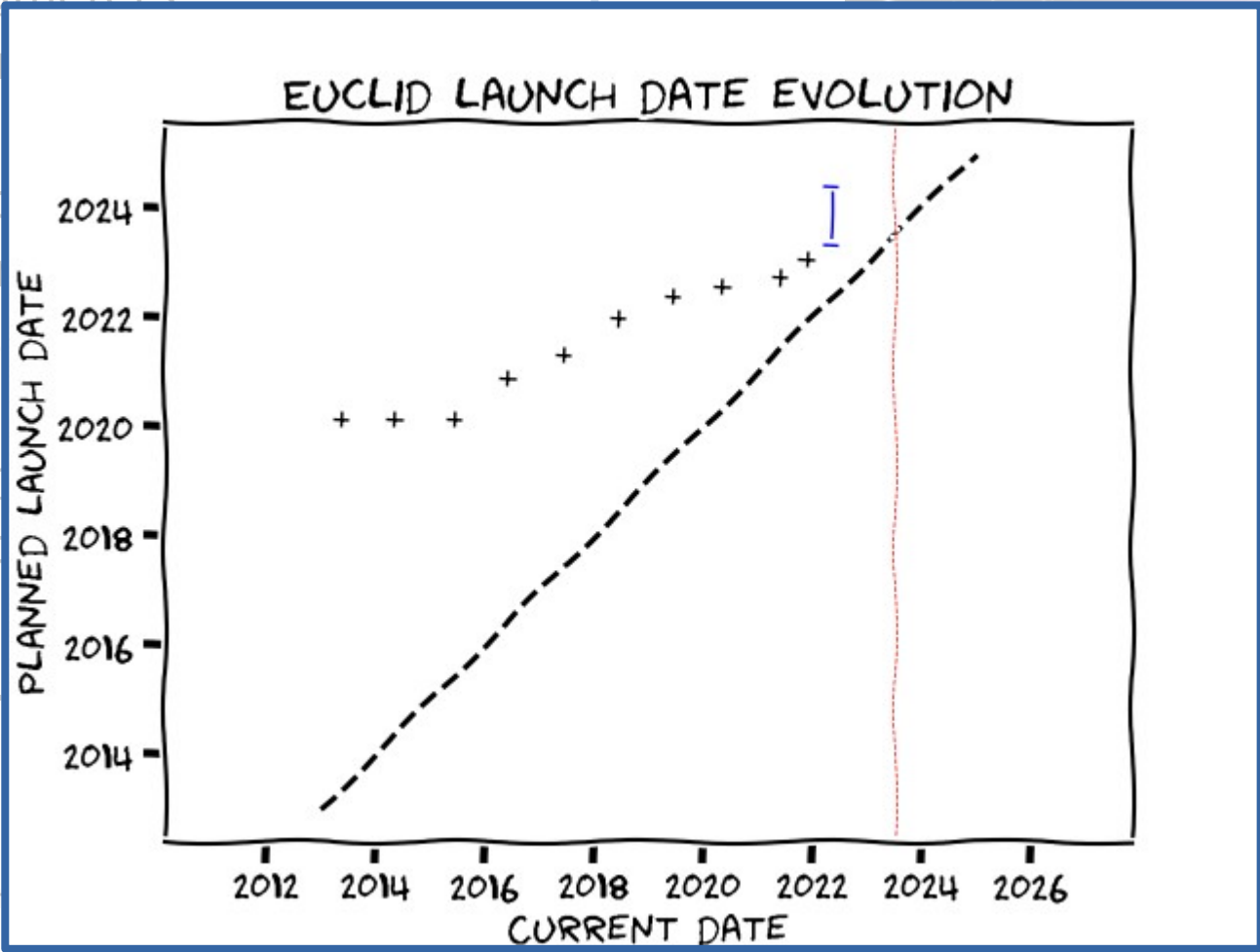
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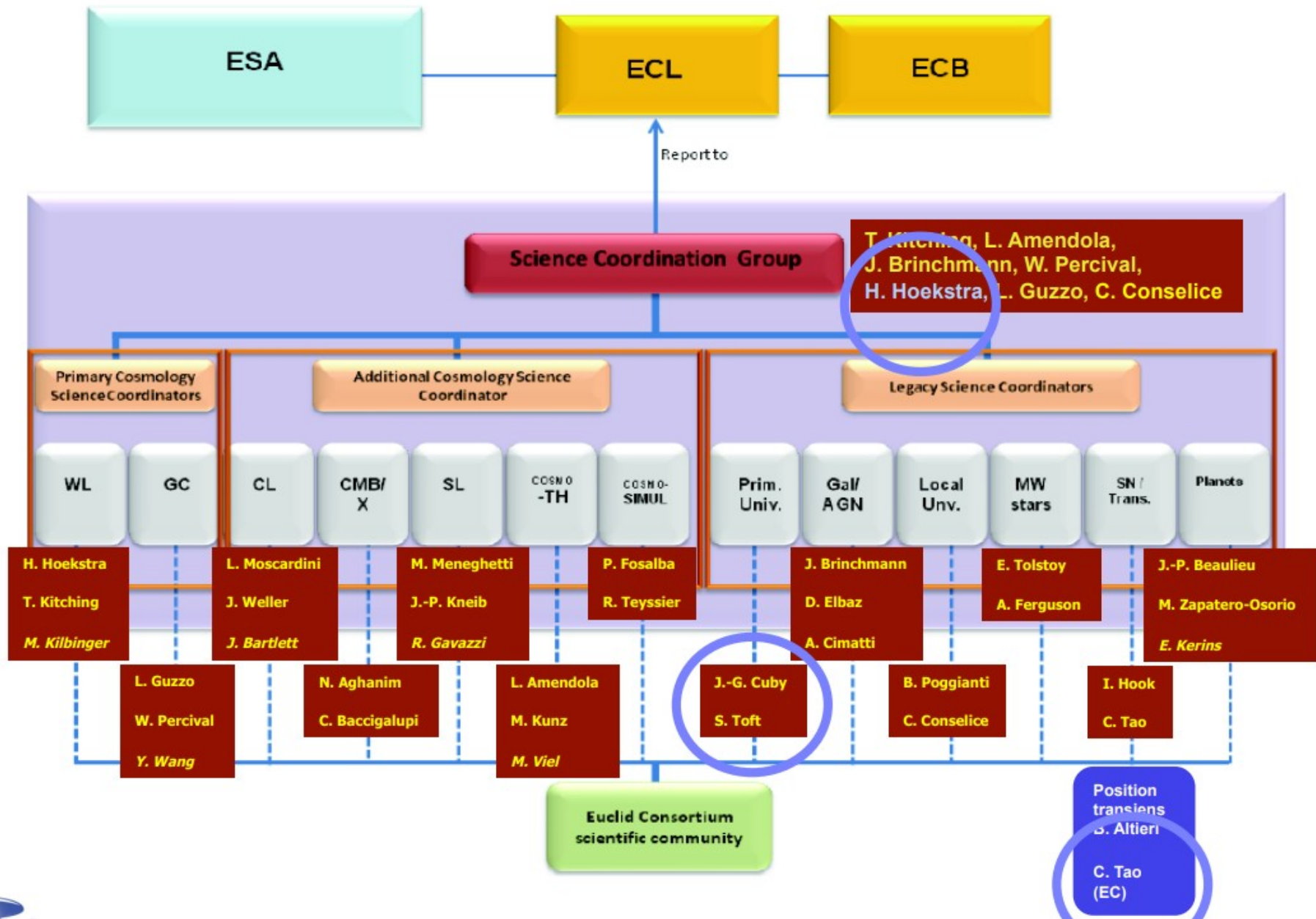
The Euclid Consortium



1500+ members, 200+ institutes, 14+ countries

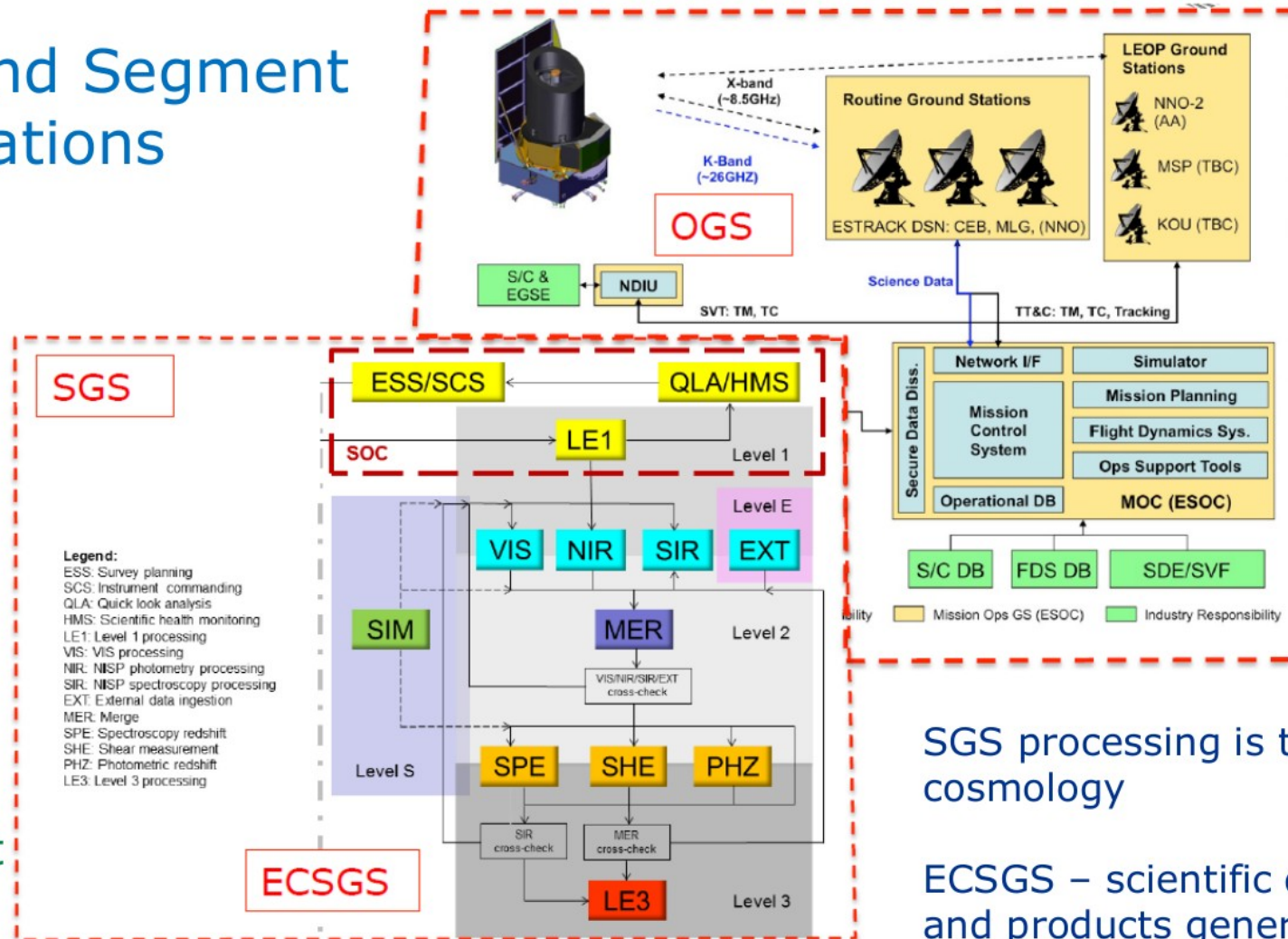
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Science Working Groups (Jan 2016 - update)



The Euclid Consortium

Ground Segment Operations

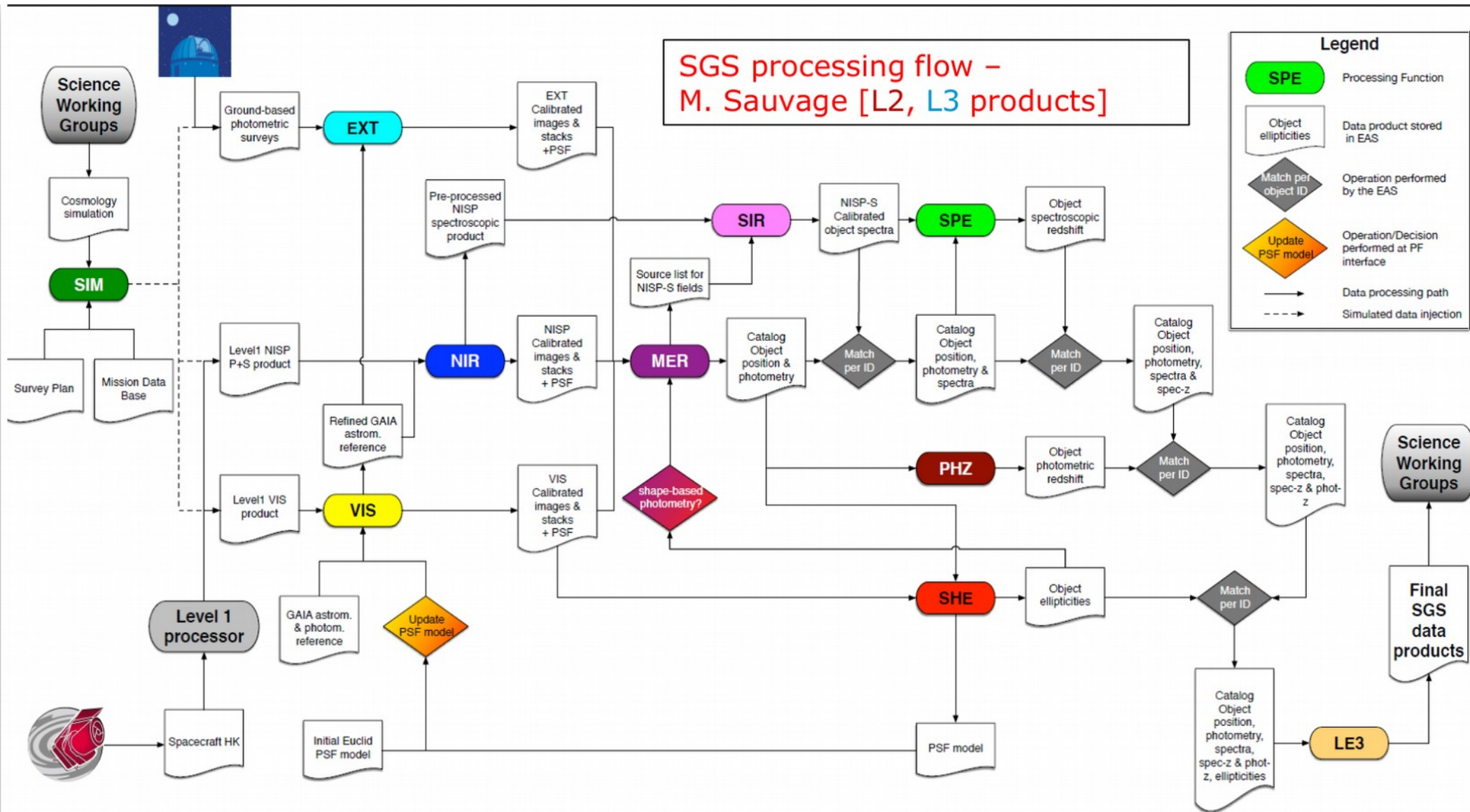


Operations Ground Segment

SGS processing is tailored for cosmology

ECSGS – scientific data processing and products generation are provided by the **Euclid Consortium**

The Euclid Consortium



Euclid publications

Paterson et al.:

Euclid preparation. XXVII. A UV-NIR spectral atlas of compact planetary nebulae for wavelength calibration

Merlin et al.:

Euclid preparation. XXV. The Euclid Morphology Challenge. Towards model-fitting photometry for billions of galaxies

Brettonnière et al.:

Euclid preparation. XXVI. The Euclid Morphology Challenge. Towards structural parameters for billions of galaxies

Camarena et al.:

Euclid: Testing the Copernican principle with next-generation surveys

Cabayol-Garcia et al.:

The PAU Survey & Euclid: Improving broadband photometric redshifts with multi-task learning

Moriya et al.:

Euclid: Discovering pair-instability supernovae with the Deep Survey

Saglia et al.:

Euclid preparation. XX. The Complete Calibration of the Colour-Redshift Relation survey: LBT observations and data release

Keihanen et al.:

Euclid: Fast two-point correlation function covariance through linear construction

Humphrey et al.:

Euclid preparation. XXII. Selection of Quiescent Galaxies from Mock Photometry using Machine Learning

Nesseris et al.:

Euclid: Forecast constraints on consistency tests of the Λ CDM model

Loureiro et al.:

KiDS & Euclid: Cosmological implications of a pseudo angular power spectrum analysis of KiDS-1000 cosmic shear tomography

Martinelli et al.:

Euclid: constraining dark energy coupled to electromagnetism using astrophysical and laboratory data

Scaramella et al.:

Euclid preparation: I. The Euclid Wide Survey

Serrano Borlaff et al.:

Euclid preparation: XVI. Exploring the ultra low-surface brightness Universe with Euclid/VIS

Ilic et al.:

Euclid preparation: XV. Forecasting cosmological constraints for the Euclid and CMB joint analysis

Stanford et al.:

Euclid preparation: XIV. The complete calibration of the colour-redshift calibration (C3R2) survey: data release 3

Martinelli et al.:

Euclid: impact of nonlinear prescriptions on cosmological parameter estimation from weak lensing cosmic shear

Martinelli et al.:

Euclid: Forecast constraints on the cosmic distance duality relation with complementary external probes

Knabenhans et al.:

Euclid preparation: IX. EuclidEmulator2 -- Power spectrum emulation with massive neutrinos and self-consistent dark energy perturbations

Tutusaus et al.:

Euclid: The importance of galaxy clustering and weak lensing cross-correlations within the photometric Euclid survey

Guglielmo et al.:

Euclid preparation: VIII. The Complete Calibration of the Colour-Redshift Relation (C3R2) Survey: VLT/KMOS Observations and Data Release

Pöntinen et al.:

Euclid: Identification of asteroid streaks in simulated images using StreakDet software

Deshpande et al.:

Euclid: On the Reduced Shear Approximation and Magnification Bias for Stage IV Cosmic Shear Experiments

Adam et al.:

Euclid preparation. III. Galaxy cluster detection in the wide photometric survey -- performance and algorithms selection

Inserra et al.:

Euclid: Superluminous supernovae in the Deep Survey

Knabenhans et al.:

Euclid preparation: II. The EuclidEmulator -- A tool to compute the cosmology dependence of the nonlinear matter power spectrum

The science of Euclid

The dark matter and energy can be studied by looking at:

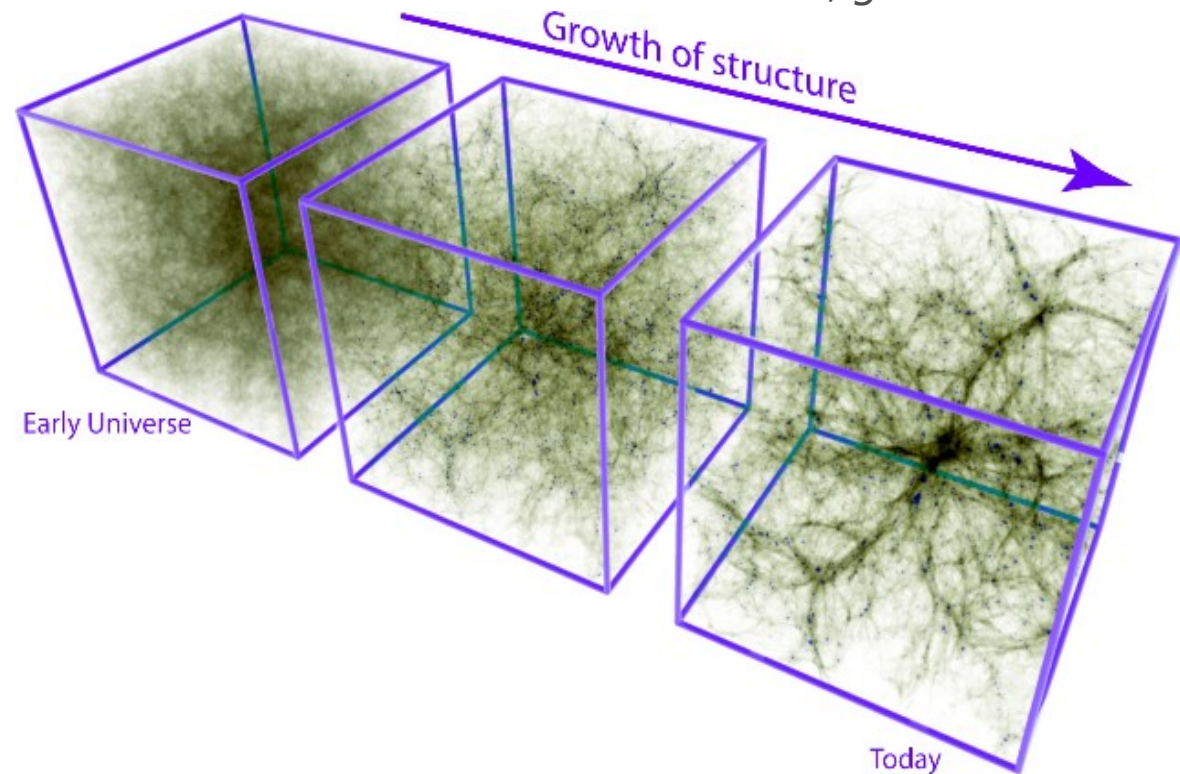
1. The geometry of the universe

→ Measure of position of galaxies as a function of redshift

2. Growth of density perturbations

→ Evolution of structure as a function of cosmic time, *growth rate*

- Structure follows the expansion of the Universe
- Gravity causes structure to evolve



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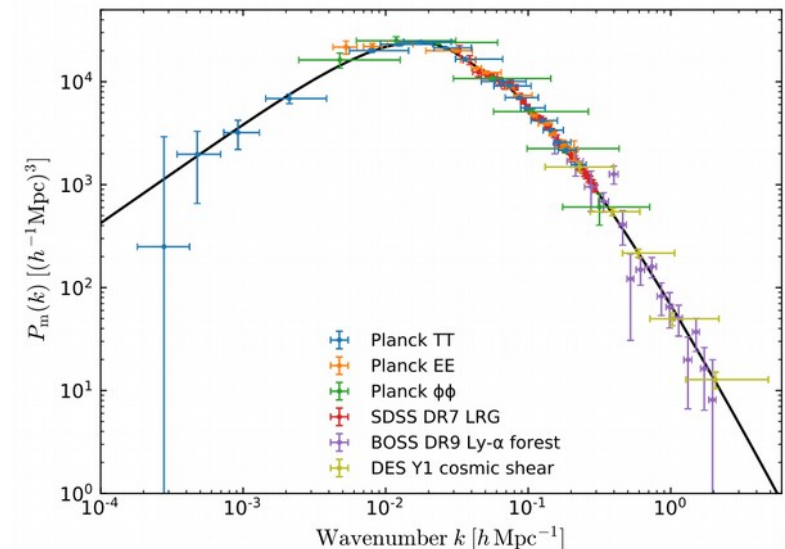
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Euclid *cosmological probes:*

- **Galaxy Clustering**

- **Baryon Acoustic Oscillations (power spectrum at large scales)**

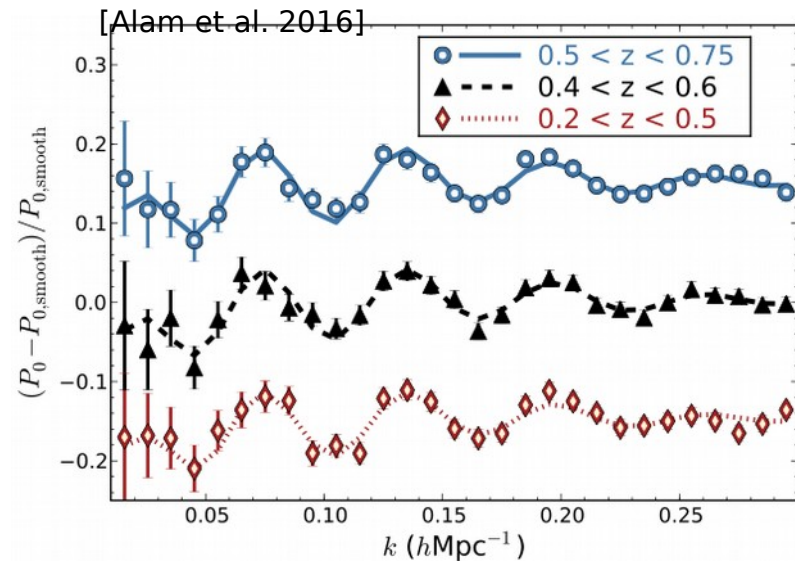
- **redshift-space distortions**



The science of Euclid – galaxy clustering

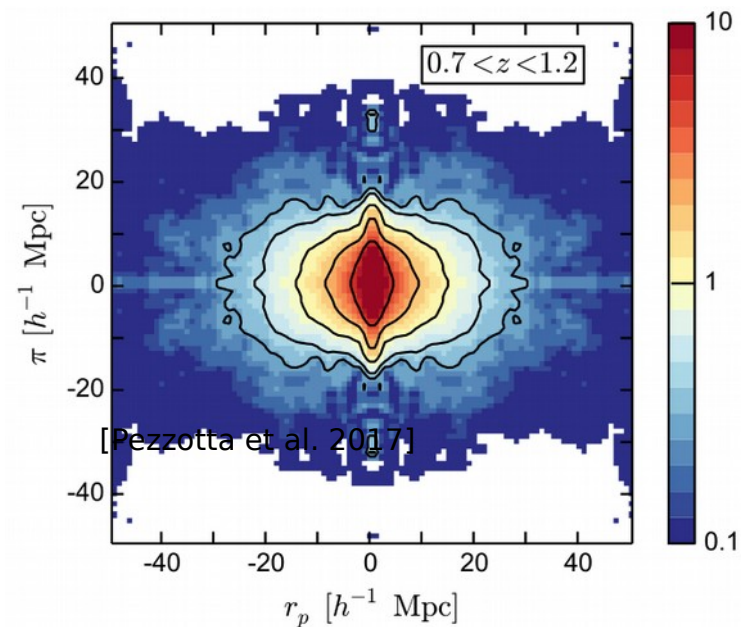
► Baryon acoustic oscillations (BAO):

- provide a **cosmic ruler**
- sensitive to the **expansion** history and the angular-diameter **distance**



► Redshift-space distortions (RSDs):

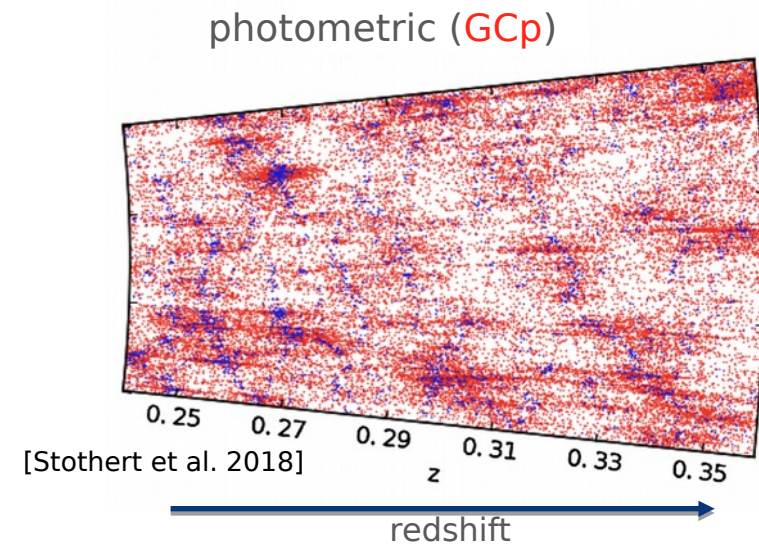
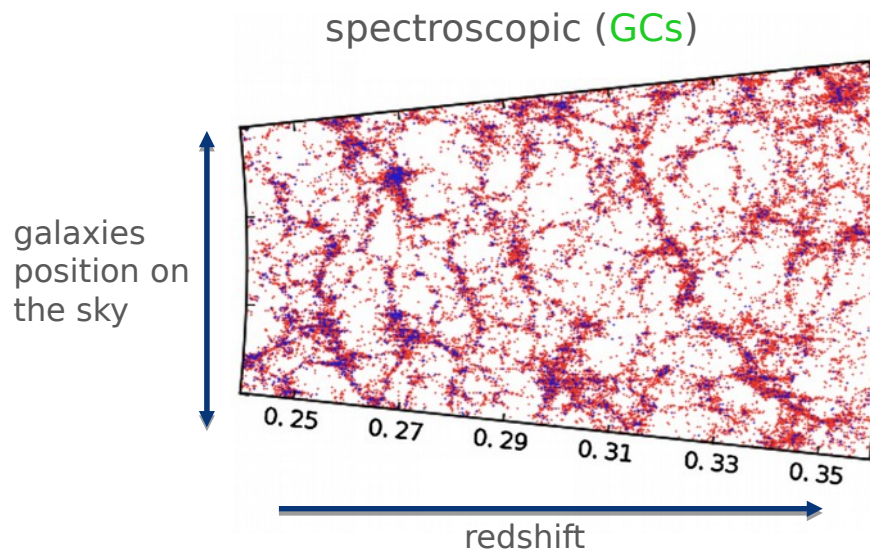
- sensitive to the **growth rate** of structures
- tests of **modified gravity**



The science of Euclid – galaxy clustering

► Spectroscopic vs **photometric** galaxy clustering:

- loss of radial information
- higher number density & different systematic uncertainties
- **source** of cosmological information



The science of Euclid

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→ Measure of position of galaxies as a function of redshift

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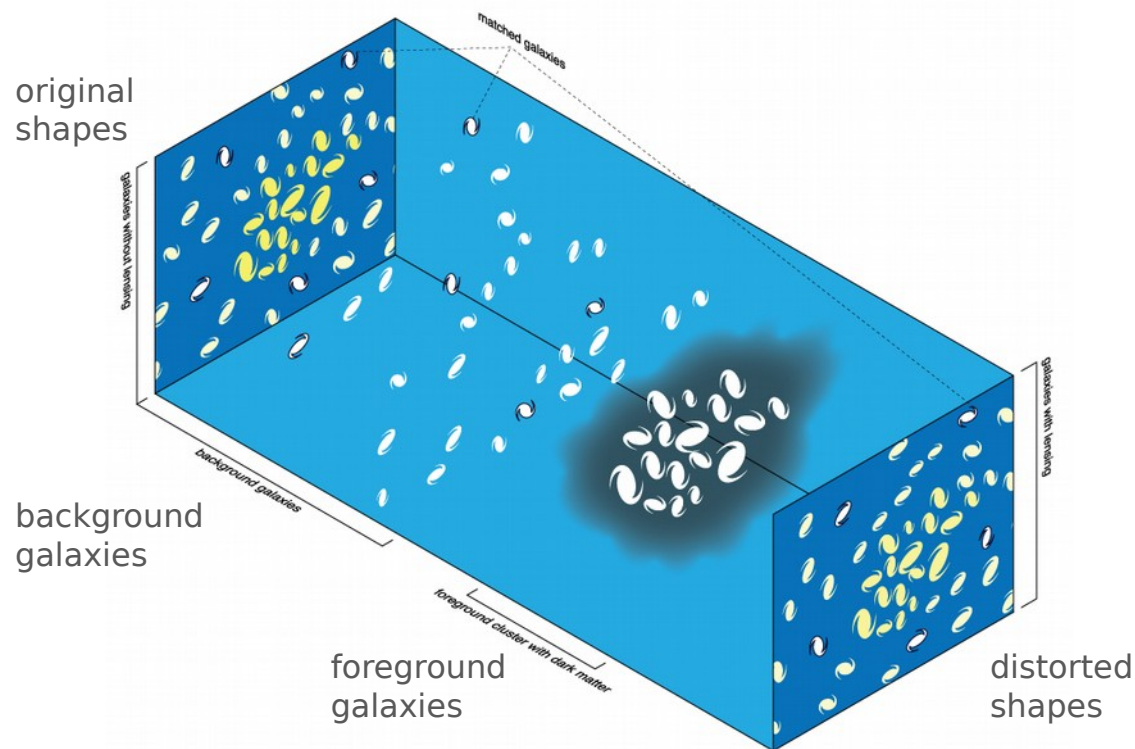
Euclid *cosmological probes*:

- **Galaxy Clustering**
 - **Baryon Acoustic Oscillations (power spectrum at large scale)**
 - **redshift-space distortions**
- **Weak Gravitational Lensing of galaxies by matter**

The science of Euclid – weak lensing

► Weak lensing (WL):

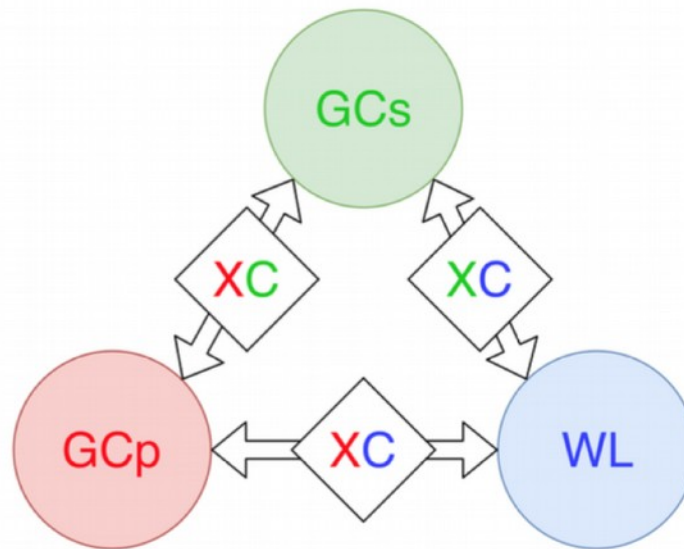
- information about **mass distribution** imprinted on galaxy images
- sensitive to **matter density**, initial conditions, and **growth of structures**



The science of Euclid – combination

- ▶ *Euclid* is the ideal survey for a combined analysis:

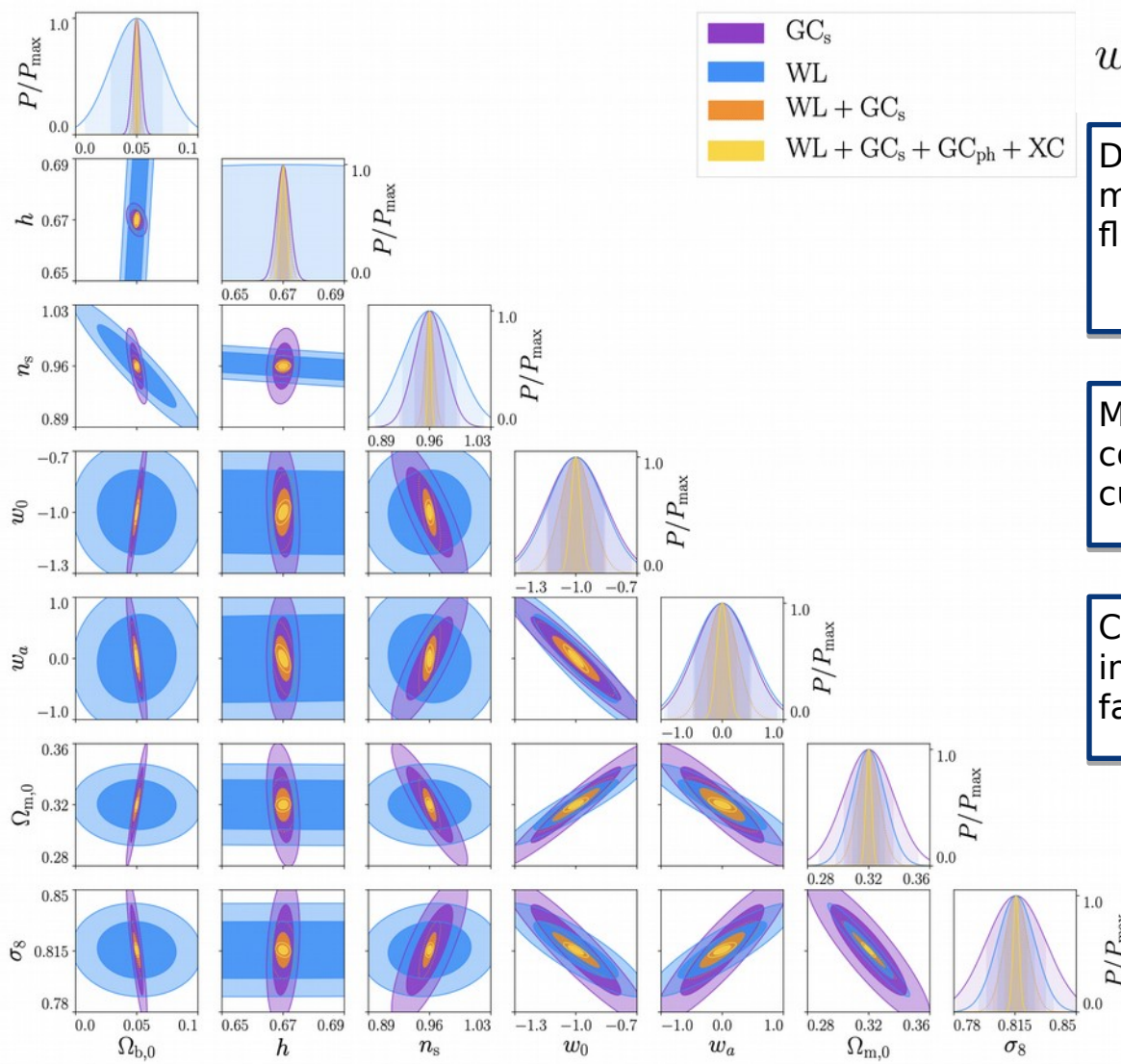
GCs, GCp, WL, and 3 **cross-correlations** (XC)



Forecasts on cosmological constraints

IST:F: Task-force to produce homogenised and validated forecasts:

→ Great complementarity between probes: breaking of degeneracies



- GC_s
- WL
- WL + GC_s
- WL + GC_s + GC_{ph} + XC

$$w(z) = w_0 + w_a \frac{z}{1+z}$$

Dark energy figure-of-merit of 1257 (500) for a flat (non-flat) cosmology

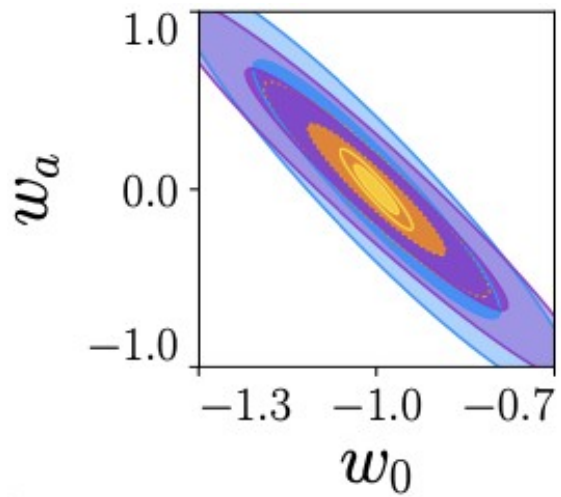
Modified gravity constrained at 5% and curvature at 1%

Cross-correlations improve constraints by a factor 4

[Euclid Collaboration VII. A&A 642, A191 (2020); I. Tutusaus *et al.* A&A 643, A70 (2020)]

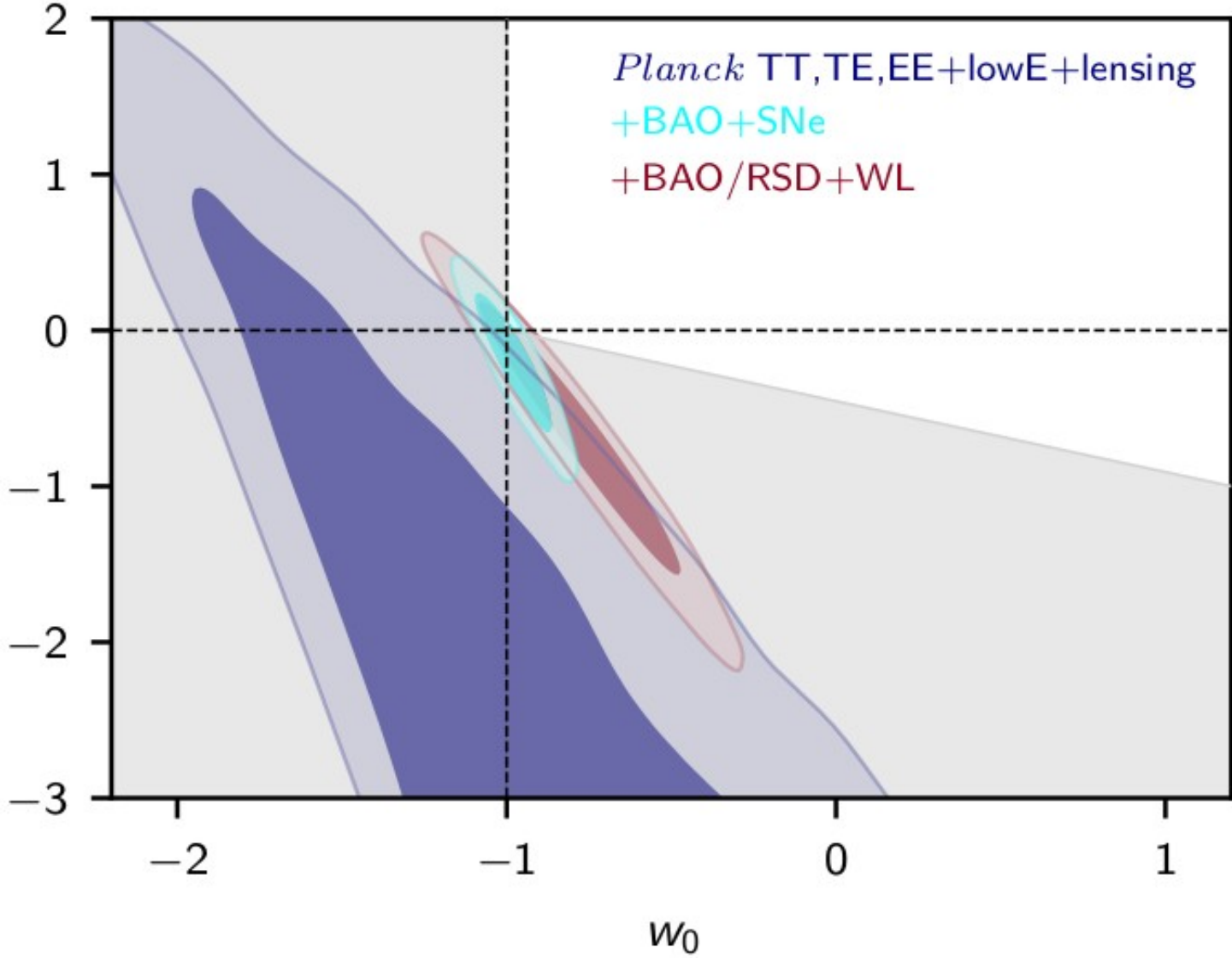
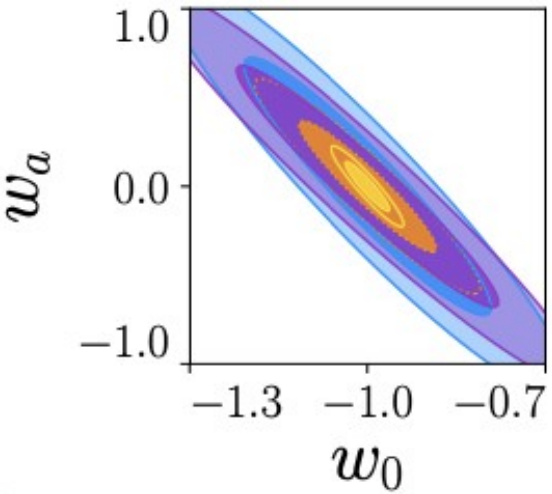
Forecasts on cosmological constraints

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Forecasts on cosmological constraints

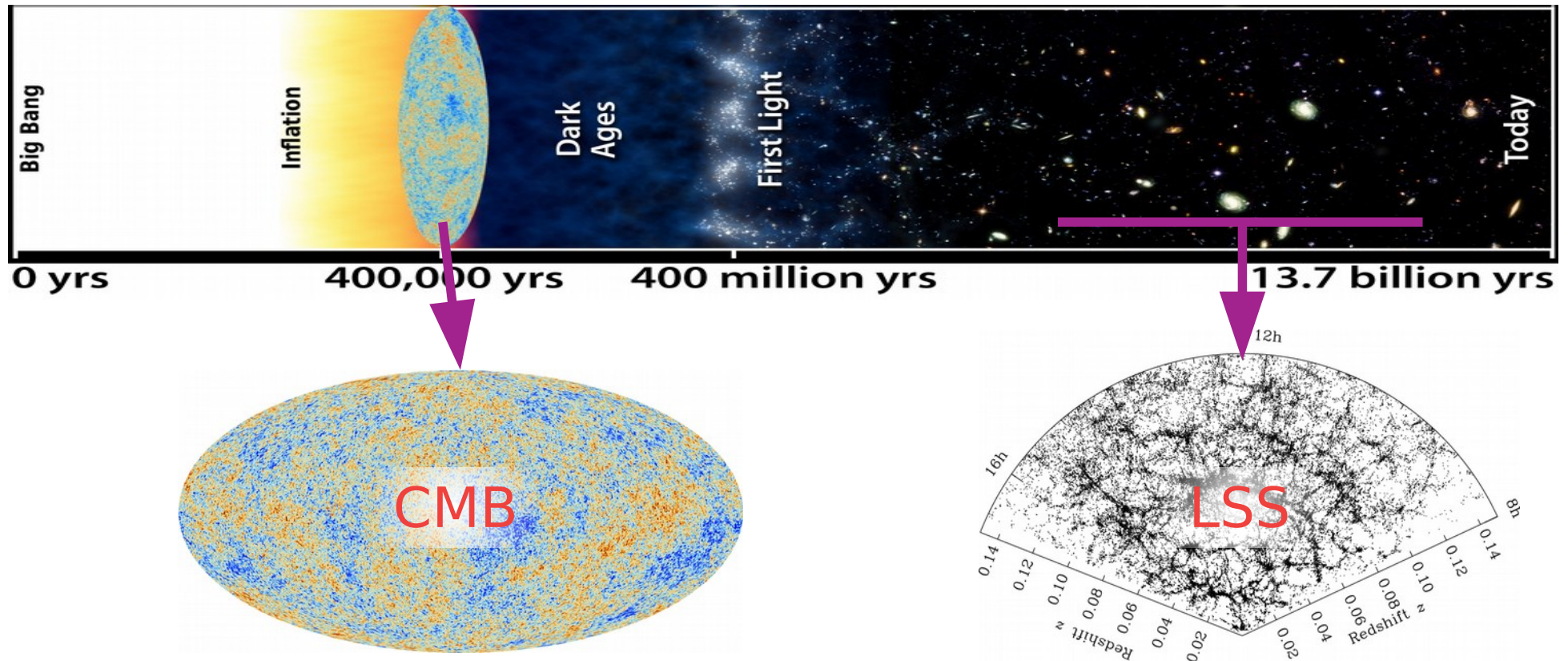
Dark energy figure-of-merit of 1257 (500) for a flat (non-flat) cosmology



Joint analysis with external probes

Joint analysis with external probes

- Probes of different “sectors”:
 - Background evolution: all standard rulers/candles
 - Perturbations: probes of structure growth
- Probes of different epochs:



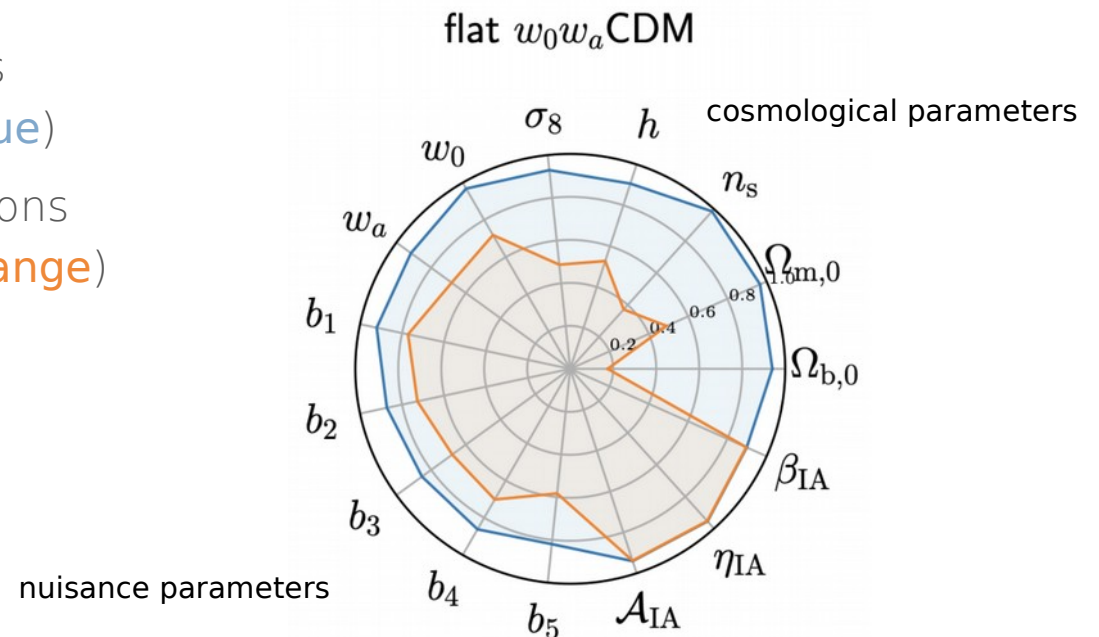
Joint analysis with external probes

Example with CMB data

CMB cross-correlations working group
produce their own set of forecasts

- Euclid main probes + Simons Observatory CMB lensing (**blue**)
- Euclid main probes + all Simons Observatory CMB probes (**orange**)

Improvements up to a
factor of 10

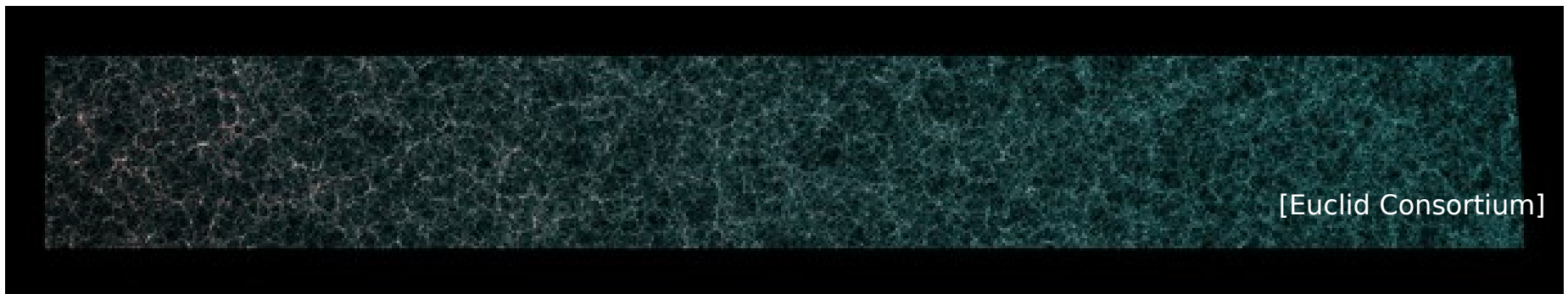


[Euclid Collaboration XV,
Ilic et al, A&A 657, A91 (2022)]

Towards data analysis

Towards data analysis

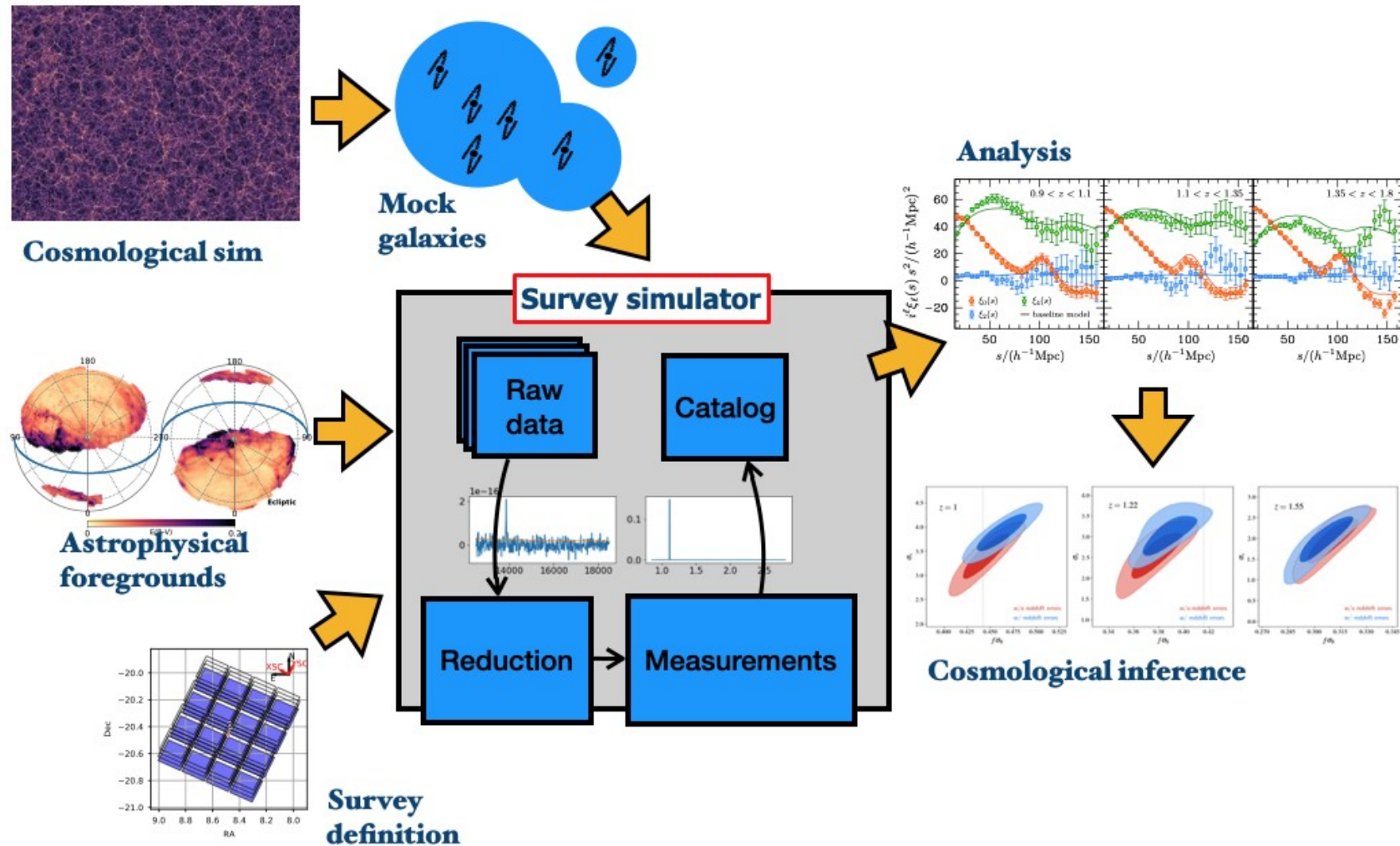
- ▶ Development of cosmological simulations to develop the analysis pipelines:
 - Flagship galaxy catalog [Carretero, Castander, Fosalba, Neissner, Pozzetti, Stadel, Tallada++]:



- WIDE: $10^9 M_{\text{sun}}$ resolution (4.1 trillion particles, 3600 Mpc/h box)
- DEEP: $10^8 M_{\text{sun}}$ resolution (0.9 trillion particles, 1000 Mpc/h box)

Towards data analysis

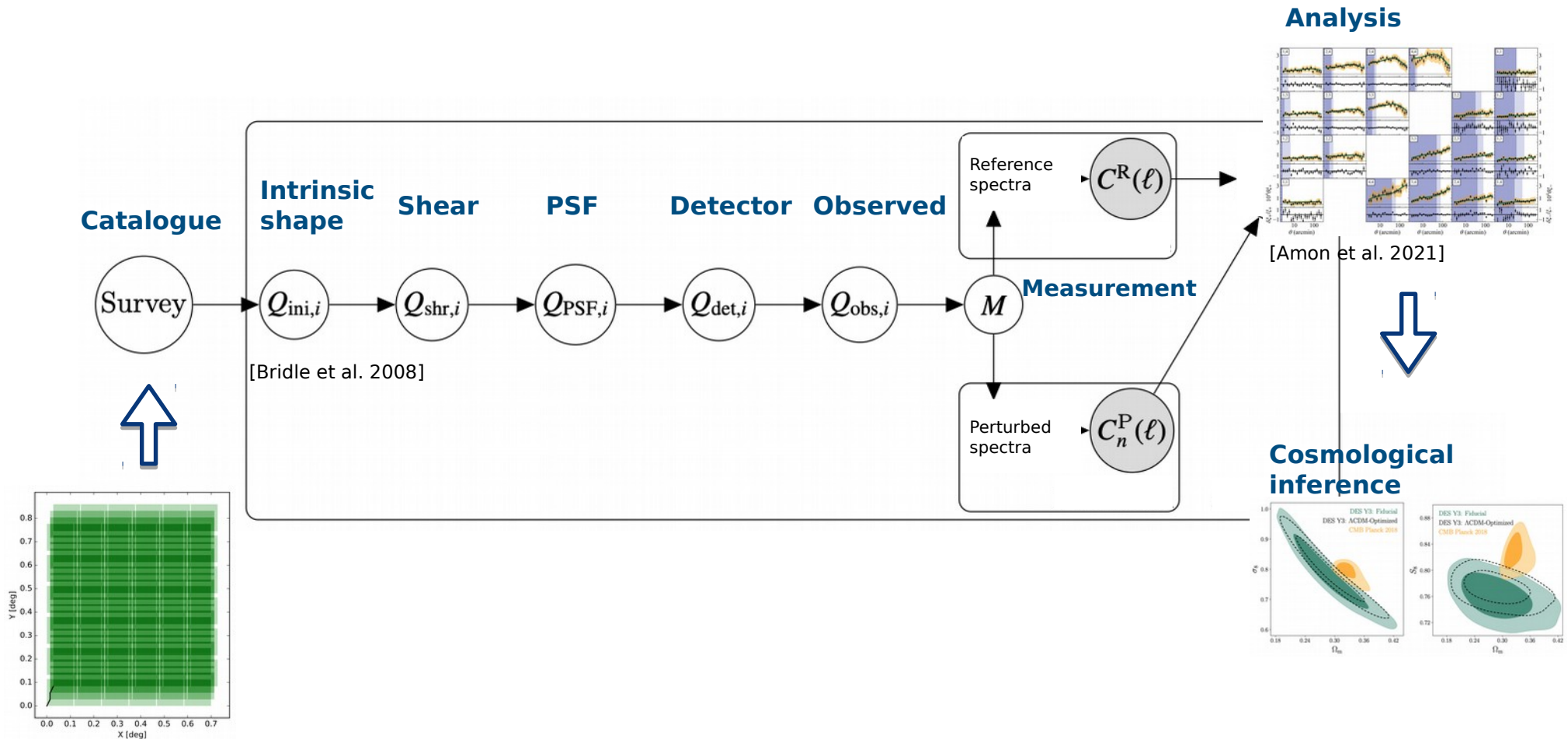
- Complex measurements: End-to-end simulations — spectroscopic galaxy clustering



[Slide courtesy of B. Granett and GC end-to-end group]

Towards data analysis

- Complex measurements: End-to-end simulations — weak lensing & photometric galaxy clustering



Towards data analysis

▶ **Simulations:**

- more volume and resolution, emulators
- end-to-end pipelines

▶ **Modelling** of the observables:

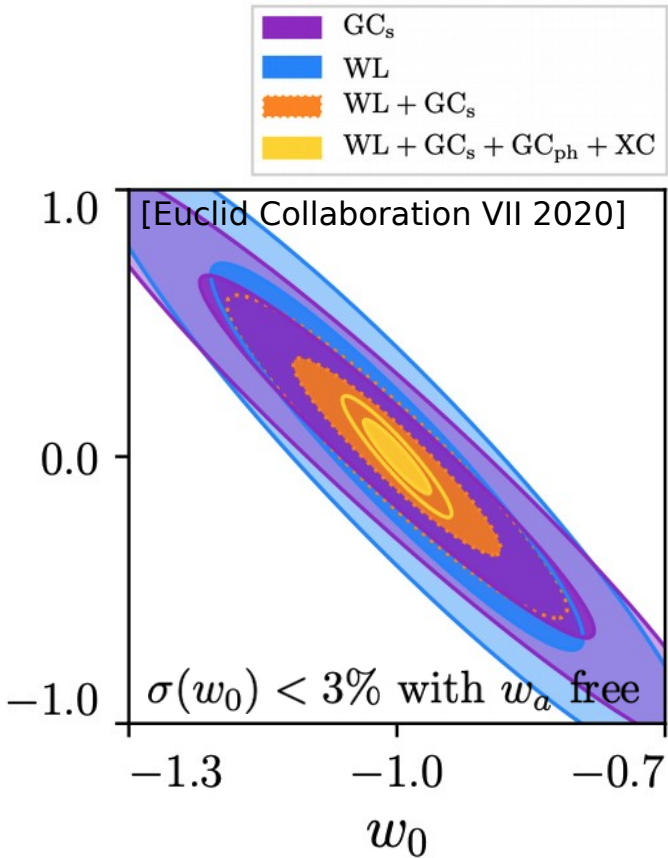
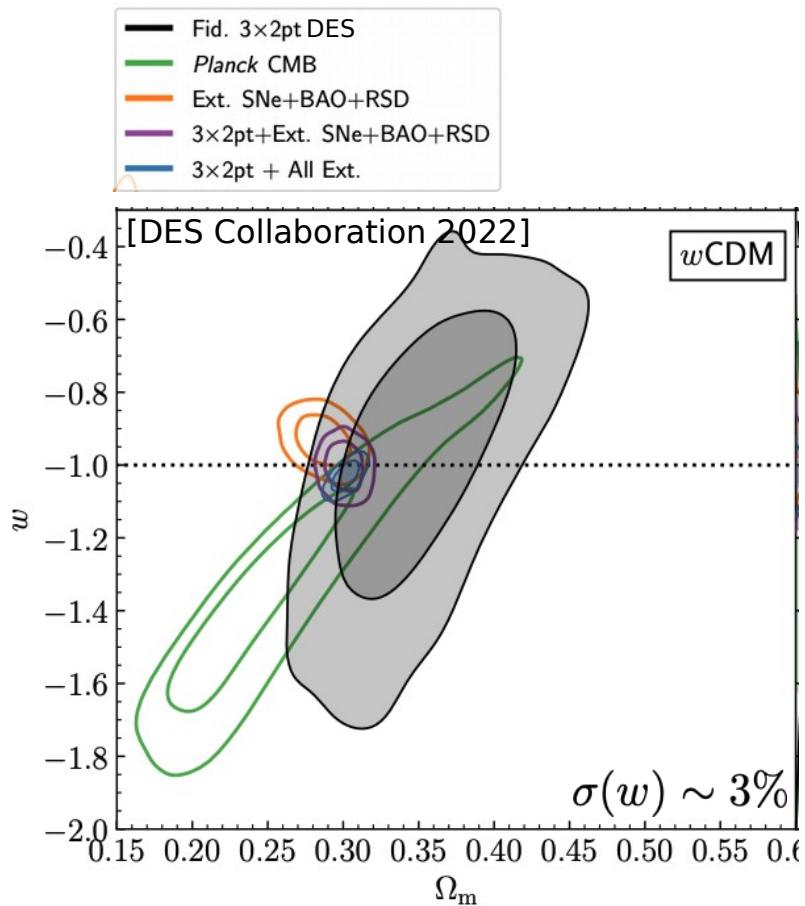
- nonlinear modelling of the matter and galaxy power spectra, including RSDs
- magnification and other relativistic contributions

▶ Towards the **coming data:**

- addition of systematic uncertainties and mitigation techniques

Conclusions

- ▶ *Euclid* will provide unprecedented constraints on **dark matter**, **gravity** at cosmological scales and will constrain **dark energy** better than all current observations together:



The end (for now!)

Thank you for
your attention !