

# Direct search for axion dark matter with MADMAX



Fabrice Hubaut, Pascal Pralavorio

CPPM/IN2P3 – Aix-Marseille Université  
(Marseille, FRANCE)



- 1- (Short) Theoretical motivations
- 2- Axion Dark matter searches
- 3- MADMAX experiment
- 4- Prototyping (*magnet, receiver, booster*) for first physics
- 5- French contributions, Timeline

# (Short) Theoretical motivations

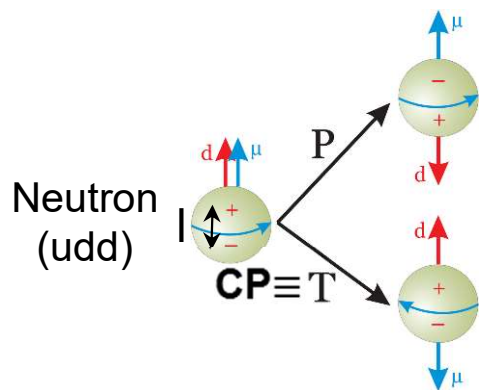
## □ Sources of CP violation in the Standard Model *[one of the Sakharov conditions]*

### ■ CP violation exists in weak interaction: observed in 1964 in kaon system

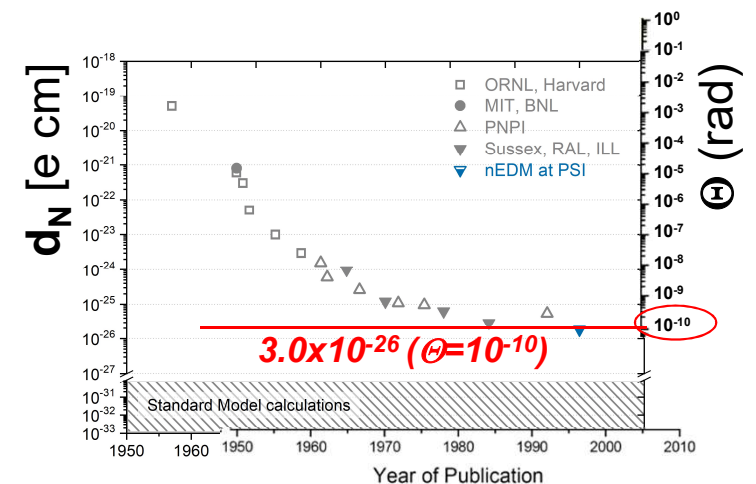
- ✓ Associated phase in quark-mixing CKM matrix measured :  $\delta_{13} \sim 1.2$  rad
- ✓ Phase still to be measured in lepton sector (PMNS matrix) : T2K, DUNE, ORCA, ...

### ■ CP violation in strong interaction ?

- ✓ CP-violating term in QCD Lagrangian (controlled by  $\Theta$ ) is allowed and should exist
- ✓ ... but  $|\Theta| < 10^{-10}$  from neutron electric dipole moment



- Electric dipole moment:  $d_N = e \cdot l$
- If strong CP :  $d_N \sim \Theta \times 10^{-16} \text{ e}\cdot\text{cm}$
- Experimental results today:  
 $\rightarrow d_N < 3 \times 10^{-26} \text{ e}\cdot\text{cm} \rightarrow |\Theta| < 10^{-10}$

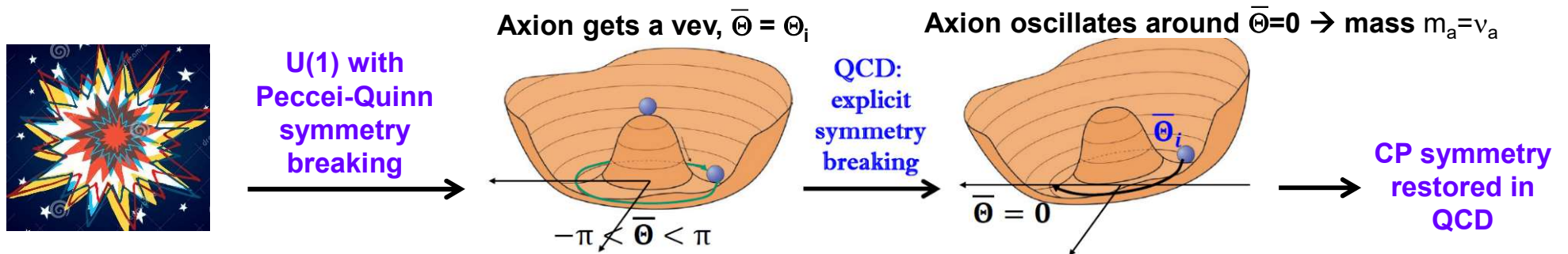


**→ Strong CP Problem = naturalness problem. Why is  $|\Theta|$  so small ?**

# (Short) Theoretical motivations

## □ Solution to Strong CP problem → Axion [motivated by particle physics]

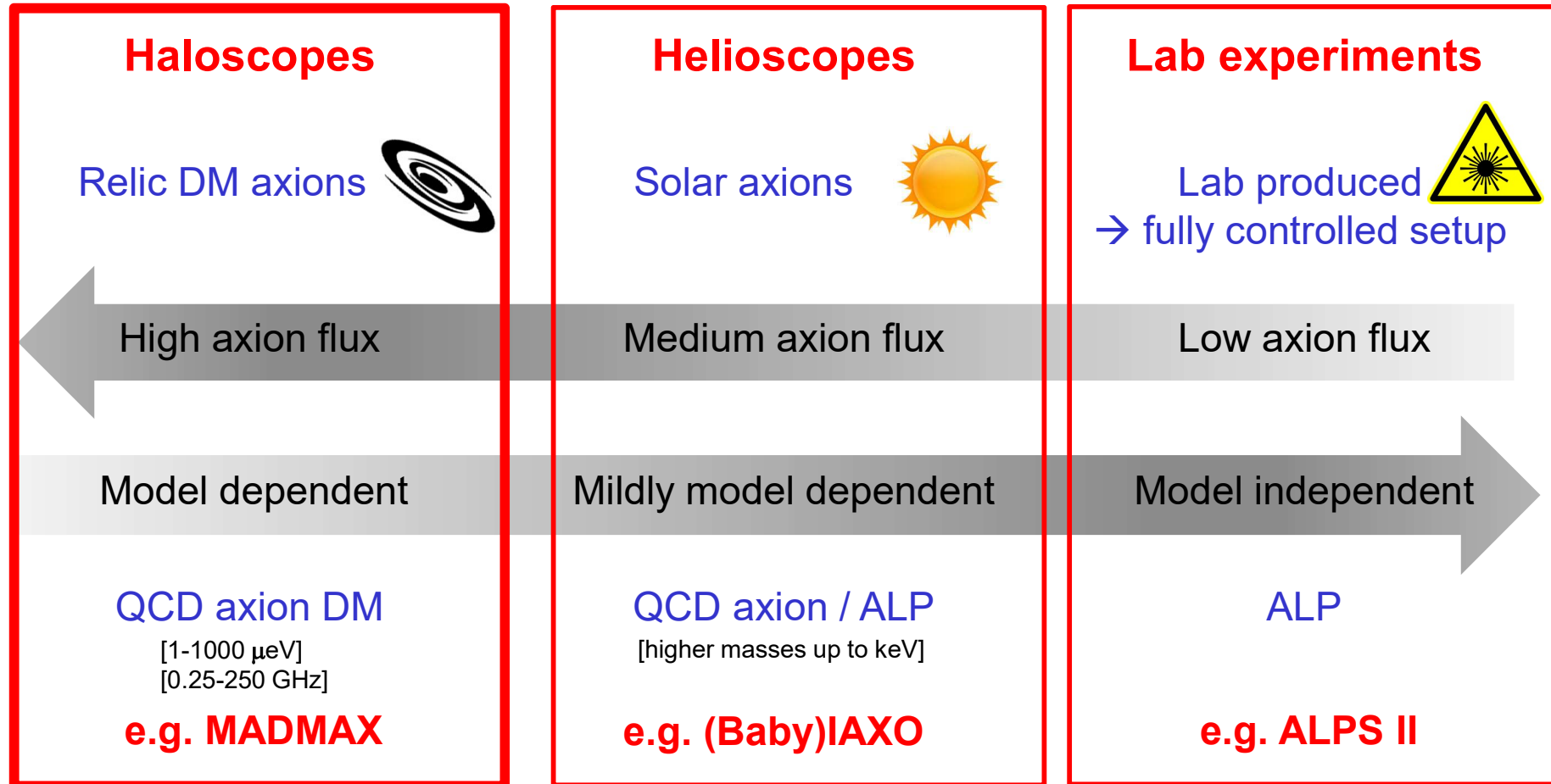
- Mechanism: new global U(1) symmetry (Peccei-Quinn, 1977) spont. broken at scale  $f_a \gg f_{EW}$ 
  - Makes  $\Theta$  a dynamical field ( $\Theta = a/f_a$ ), with  $a$  = pseudo-scalar boson
  - Suppress CP-violating term in Lagrangian ( $\Theta_{\text{eff}} \rightarrow \Theta - a/f_a$ ): explains absence of CP strong
- Consequence: generation of a Goldstone boson = **axion** (Weinberg-Wilczek, 1978)
  - Properties are all known given the scale of symmetry breaking  $f_a$  [mass  $m_a \approx m_\pi f_\pi / f_a \ll eV$ ]
  - Very weak couplings to SM particles (suppressed by  $f_a$ ) and  $\tau_{\text{axion}} > t_{\text{Universe}}$
- Cosmology: Non-thermal axion production at  $T \sim f_a$  (can occur before or after inflation)



→ Axion = natural candidate for DM for  $m_a = 1-10^3 \mu eV$  (i.e.  $f_a = 10^{12}-10^9 \text{ GeV} \gg f_{EW}$ )

Remark: ALP (Axion Like Particle) = scalar not solving strong CP problem but potential DM candidate

# Axion/ALP searches

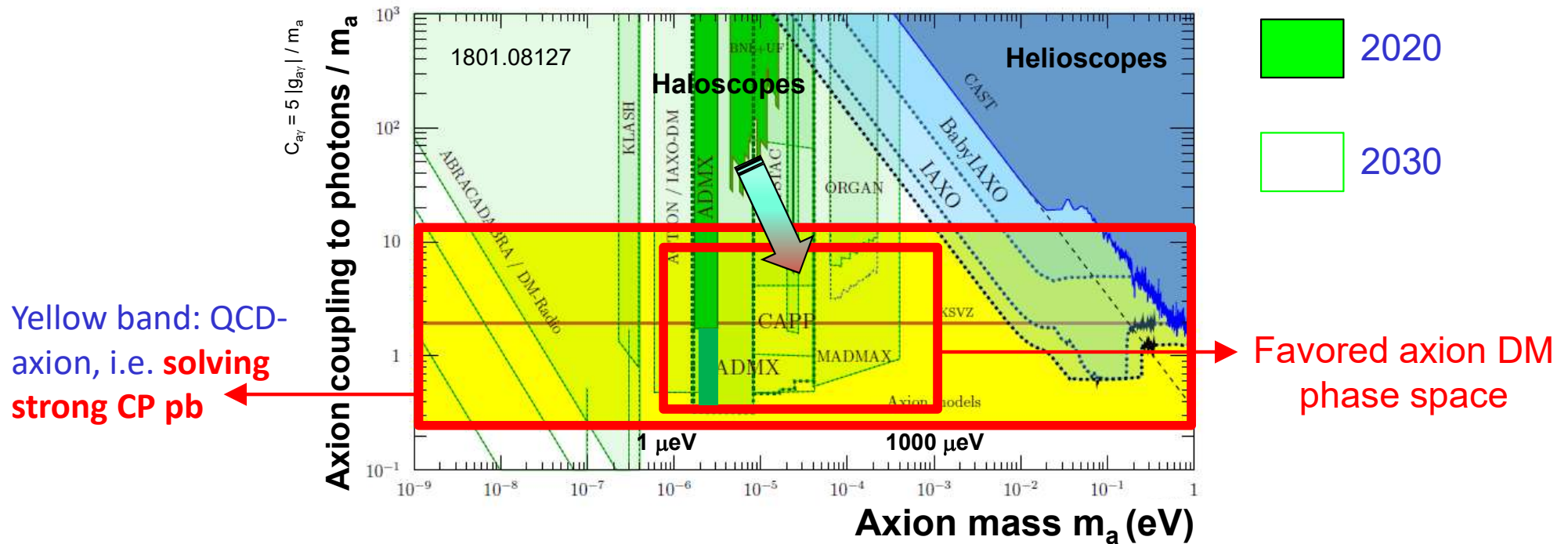


Complementarity between the 3 approaches at the DESY Hub



# Axion Dark Matter searches (1/2)

□ Extraordinary weak coupling of axions to photons ...



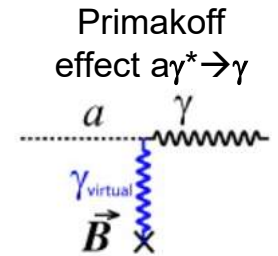
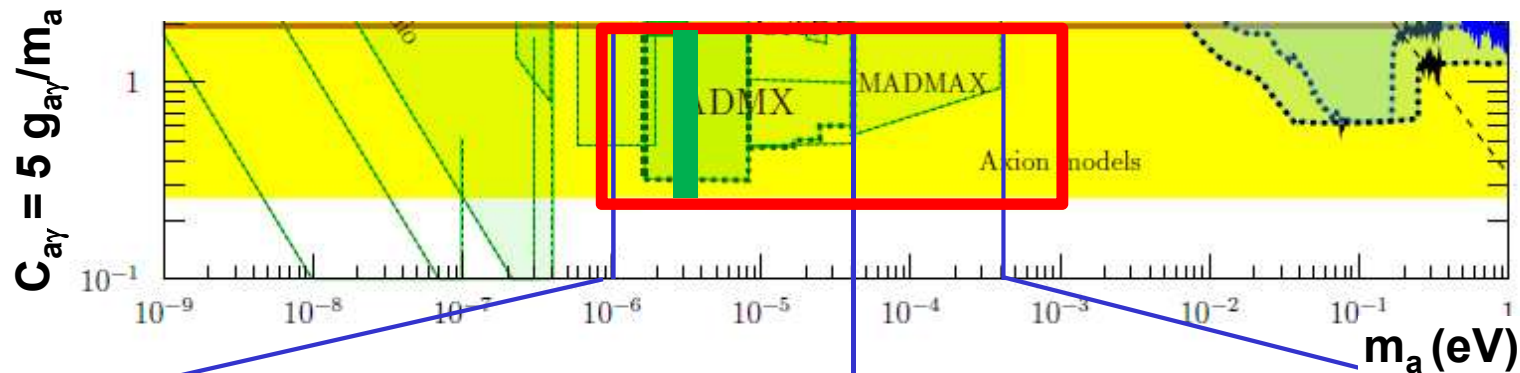
- Only 1 experiment (ADMX) **currently** probe a (very small) part of the favored phase space
- Vast **R&D program** to improve signal sensitivity and expand range of axion mass search (*post-inflationary scenarios suggests  $m_a \sim 100 \mu\text{eV}$* )

**Next decade promising : axion DM most favorable region will be probed**

# Axion Dark Matter searches (2/2)

## □ ... make axions extraordinary challenging to detect

- Convert axions into photons [ $E$  field of  $O(10^{-12} \cdot \frac{B}{10 T})$  V/m] → **high  $B_{\text{field}}$**  [ $B \gg 1 T$ ]
- Boost  $E_{\text{field}}$  [up to detectable  $P \sim 10^{-22}$  W] → **resonant cavities**
- Scan over range of axion mass → **tunable** set-up



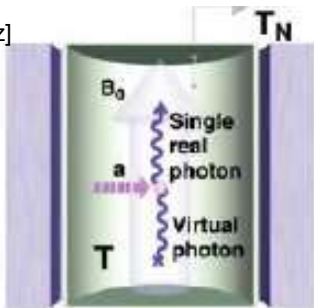
1  $\mu\text{eV}$  [0.25 GHz]

40  $\mu\text{eV}$  [10 GHz]

400  $\mu\text{eV}$  [100 GHz]

**Cavities**

PRL51 (1983) 1415

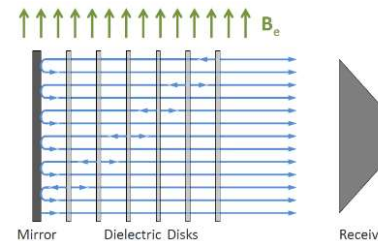


ADMX, HAYSTACK, CAPP

**Multicavities, Very High B**

RADES, ADMX-SideCar, CAPP, GrAHal

*Cavity size too small + high noise*



**MADMAX**

**Dielectric haloscopes**

[novel concept 2013] PRD88 (2013) 115002

[ $v_a = v_\gamma$ ] → Microwave regime

→ New ideas of last decade coming to maturity to scan favored mass range

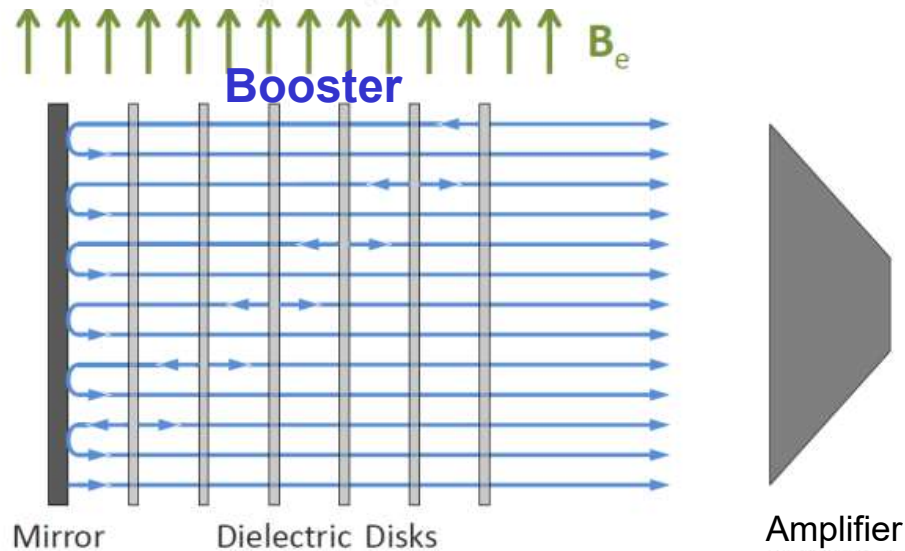
# MADMAX experiment (1/2)

White Paper [EPJC 79 (2019) 186, 1901.07401]

## Principles

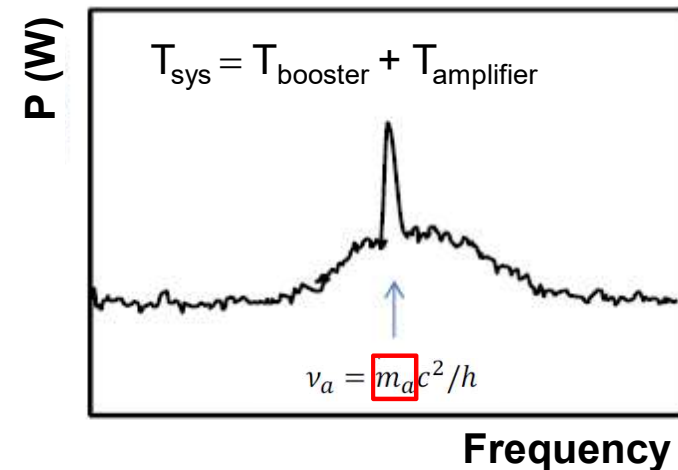
- Constructive interference of coherent photon emission at dielectric layers surface (~leaky resonators cavities): boost ( $\beta^2$ ) wrt only mirror [ $\epsilon, \delta, N$ ]

$$P_{sig} = 10^{-22} \text{ W} \times \left(\frac{\beta^2}{50000}\right) \times \left(\frac{B_e}{10 \text{ T}}\right)^2 \times \left(\frac{A}{1 \text{ m}^2}\right) \times C_{a\gamma}^2$$



$$P_{sig} = 10^{-22} \text{ W} \times \left(\frac{SNR}{5}\right) \times \left(\frac{T_{sys}}{4 \text{ K}}\right) \times \left(\frac{4 \text{ days}}{t}\right)^{1/2}$$

Thermal Noise



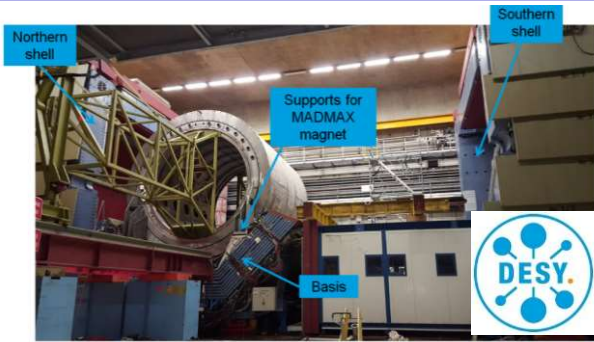
- Axion mass scan : by moving discs with piezo motors ( $\mu\text{m}$  prec.) at 4K under 10 T (50 MHz step)

**MADMAX exploits a new exp. approach to cover an uncharted phase space**

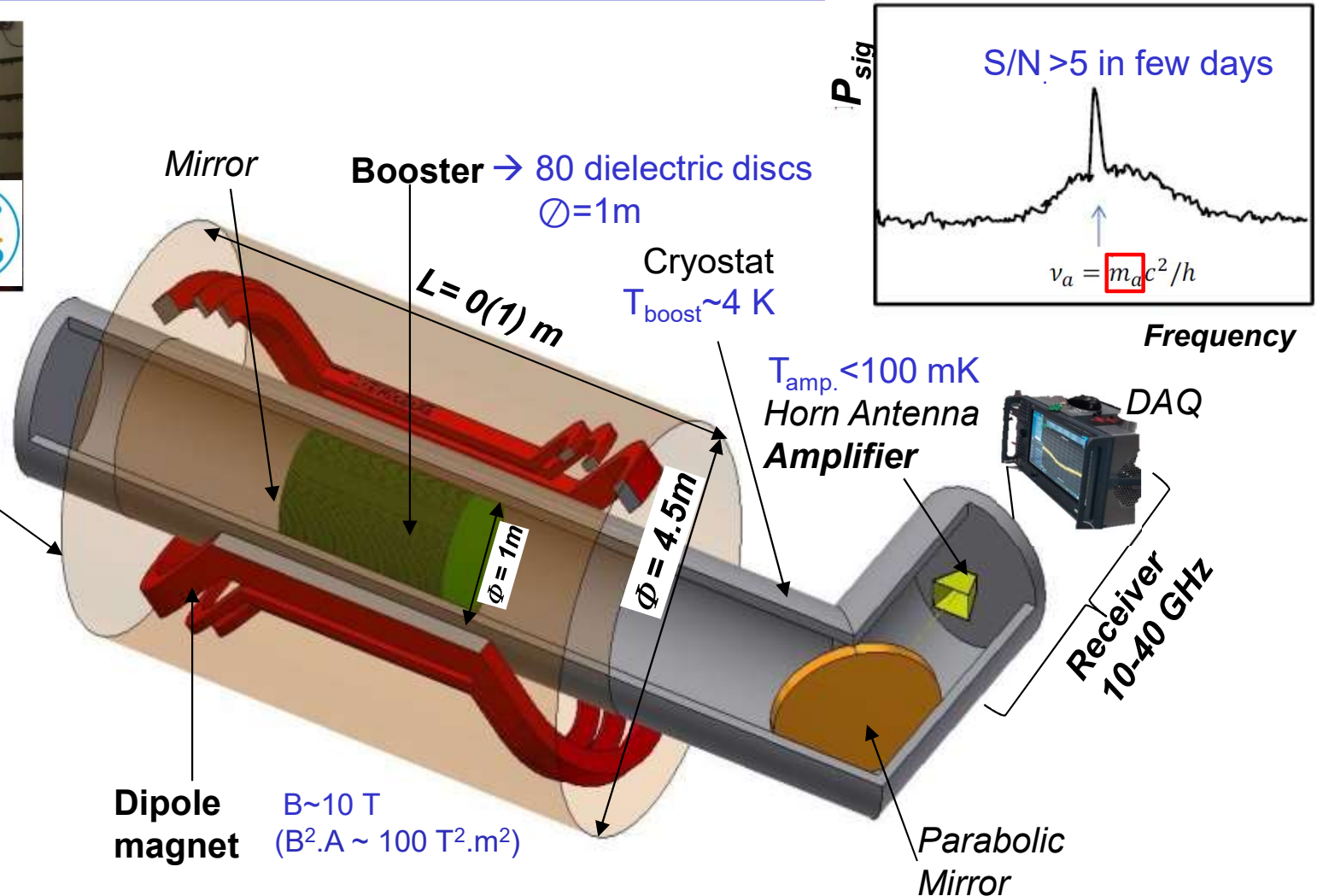
[post-inflationary scenarios suggests  $m_a \sim 100 \mu\text{eV}$ ]

# MADMAX experiment (2/2)

~ 50 people, French (2), German (6) and Spanish (1) institutes



Experiment location: HERA  
in former H1 iron yoke



## 3 main challenges :

- Magnet
- Booster (cold, B field)
- Receiver (amp cold)

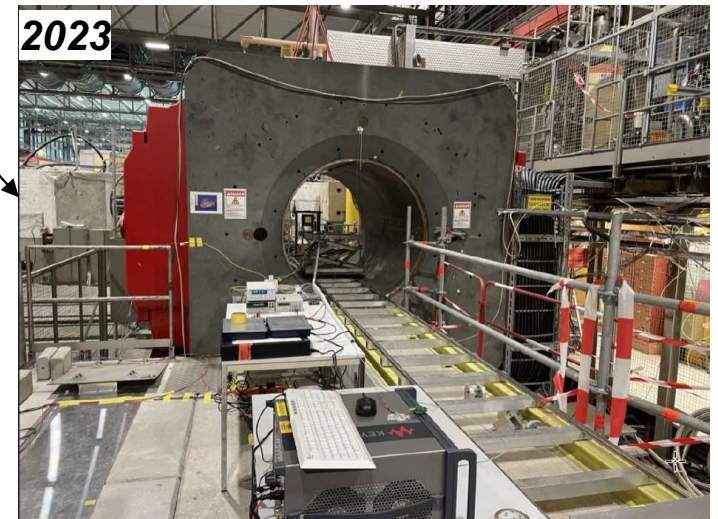
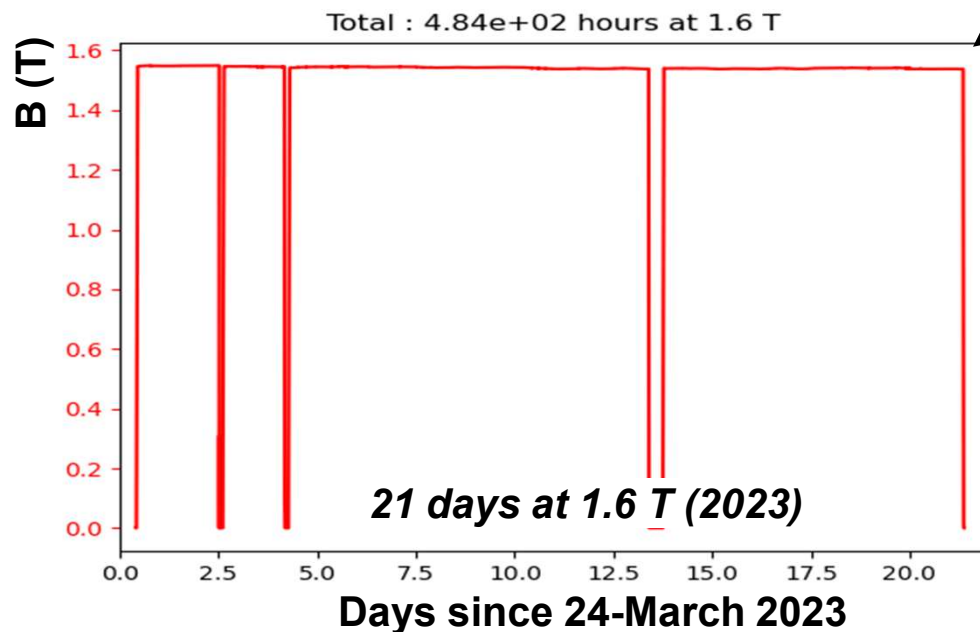
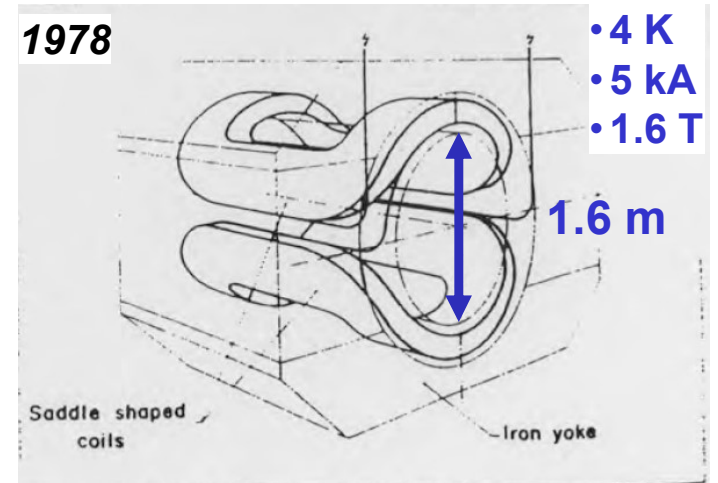
**Start with prototyping phase to validate concept: cutting-edge R&D**



# Need a Magnet

## □ CERN borrows us the world largest warm bore dipole magnet

- **Jun 1978** : Installation in the North Area at CERN
- **Sep 2020** : CERN RB approves usage by MadMax (YETS)
- **Mar 2021** : full refurbishing around magnet area
- **Mar 2022** : installation of new power converters
- **Apr 2022** : magnet recommissioning
- **Mar-Apr 2023**: MADMAX full user of the magnet



# Need a Receiver System

## □ Composed of

### ▪ Low Noise Amplifier (LNA) ...

✓ “Classic” HEMT

### ▪ ... connected to custom-made receiver

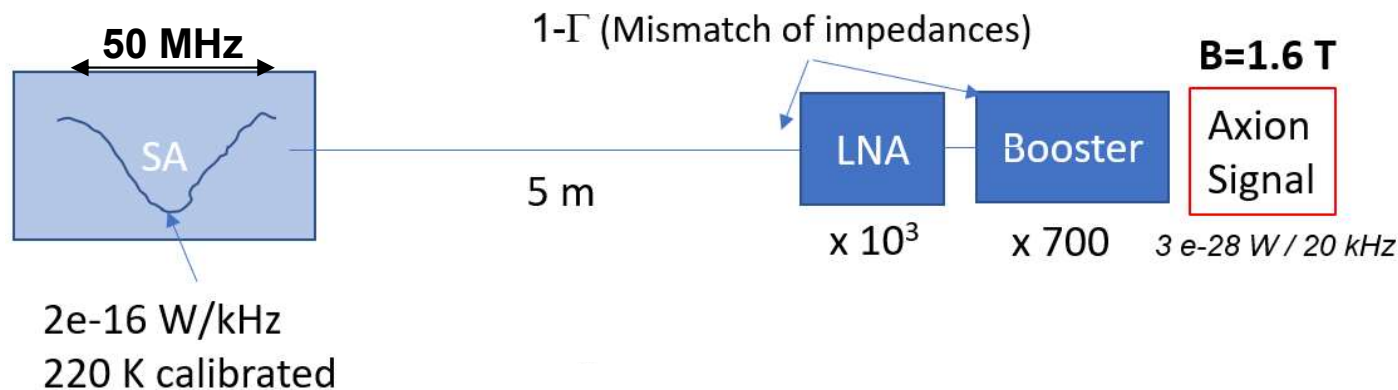
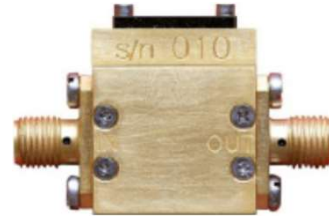
✓ Three mixing stages to down sample from 20 GHz to 50 MHz

✓ Fast Fourier Transform in 4 samplers → 1% dead time

✓ Tested at CERN in 2022 but saturation and time instability

### ▪ ... connected to commercial spectrum analyzer (Keysight)

✓ Tested at CERN in 2023 : stable, no saturation but higher dead time\*



\* Improve dead time next year by adding data streaming

# Develop the booster concept

## □ Address the two main challenges

- Move the disks at  $\mu\text{m}$  level precision at cold and under high B-field
- Understand RF behavior  $\rightarrow$  Calibrate boost factor
- $\rightarrow$  Ultimately do physics !

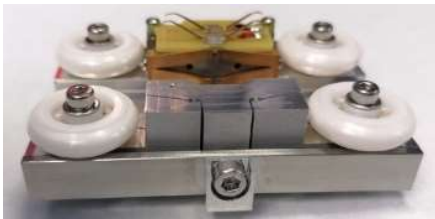
Name	Goal	Type	Made of	Avail.	Test	Room Temp. Cold (10 K)
P200	Piezo-motor + mechanics	Open Booster	1 moveable disk $\phi = 200 \text{ mm}$	2021	<b>2022</b>	
CB100	RF studies + First physics	Closed booster	3 fixed disks $\phi = 100\text{mm}$	2021	2022, 23, <b>24</b>	
CB200	RF studies + First physics	Closed Booster	4 fixed disks $\phi = 200 \text{ mm}$	2022		24
Proto-3	Scan ALP around $100 \mu\text{eV}$	Open Booster	3 moveable disks $\phi = 300 \text{ mm}$	2024		<b>25, 26?</b>

**Gradually building the 'final' booster design**

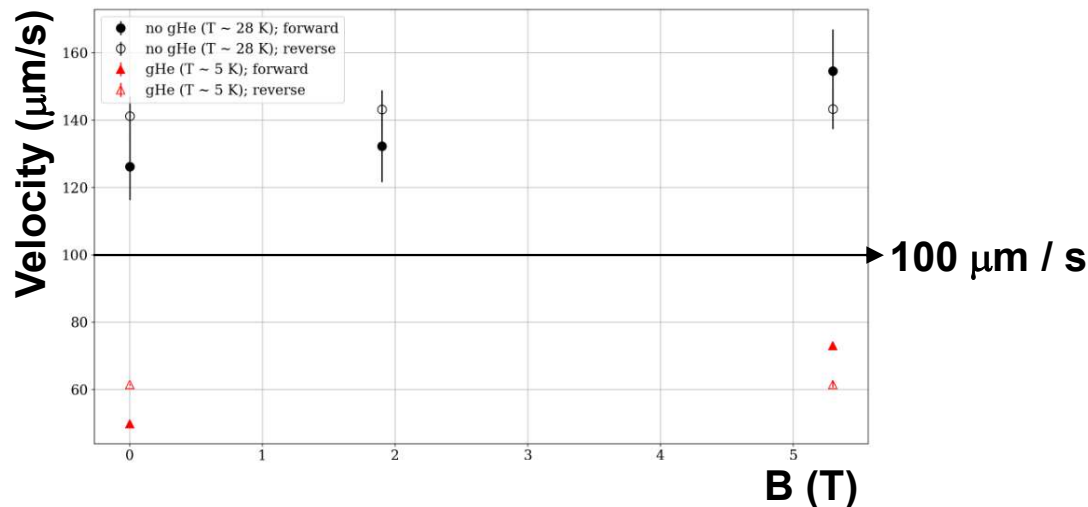
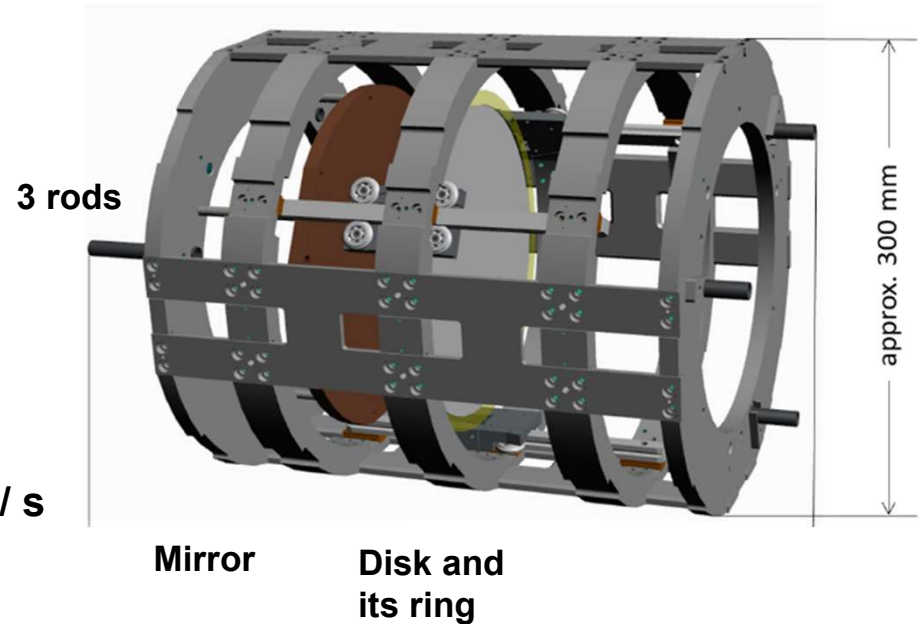
# Testing the disk drive (1/2)

Name	Goal	Concept	Made of	Avail.	ALP Test
P200	Piezo-motor + mechanics	Open Booster	1 <b>moveable</b> disk $\phi = 200$ mm	2021	2022

Test one commercial JPE piezo motor at 5 K and 5.3 T (*ALP magnet in DESY*)



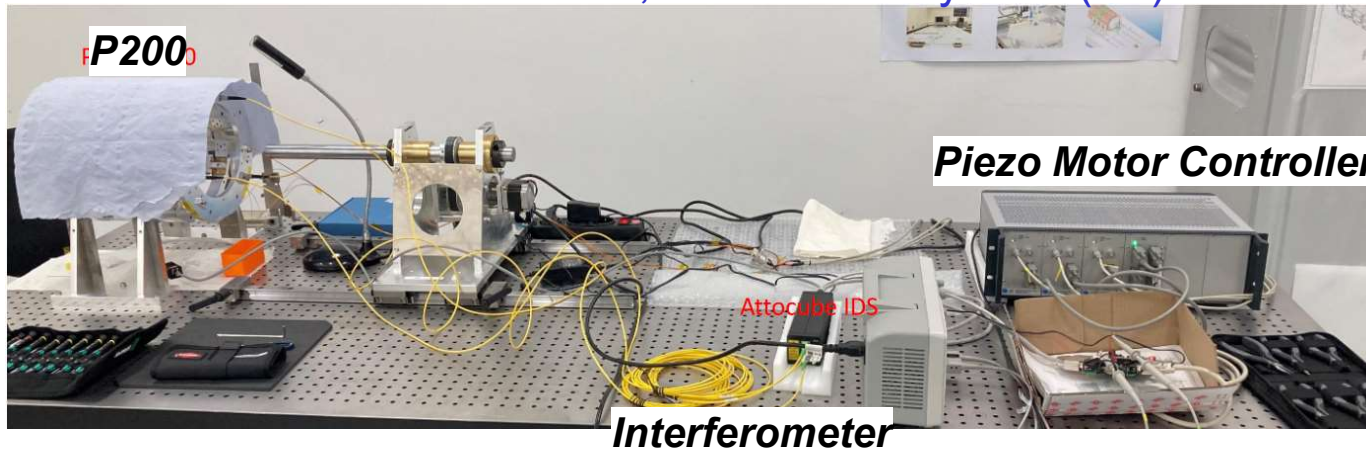
Build full mechanical structure of Open Booster and insert 1 mirror + 1 disk (*3 piezo motors*)



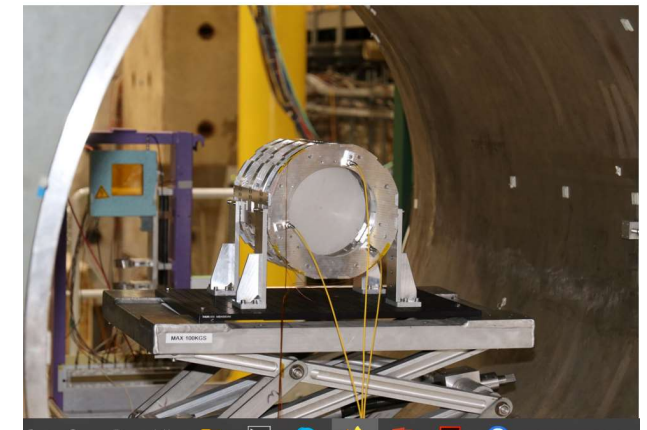
# Testing the disk drive (2/2)

Name	Goal	Concept	Made of	Avail.	Morpurgo Test
P200	Piezo-motor + mechanics	Open Booster	1 moveable disk $\phi = 200$ mm	2021	2022

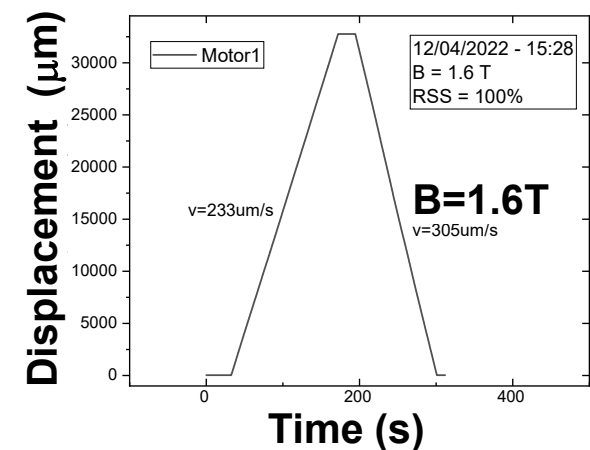
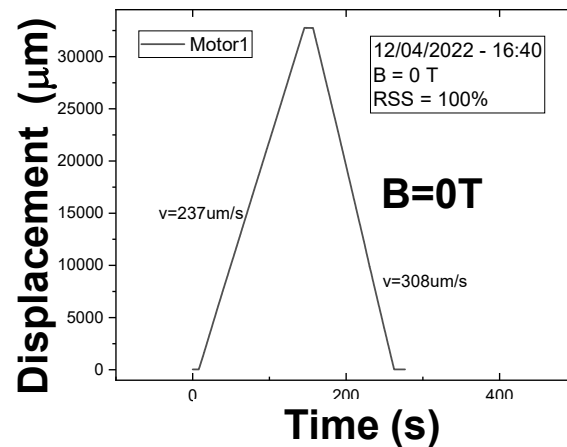
Tested in the lab, in a CERN cryostat (4K)



Tested under 1.6 T

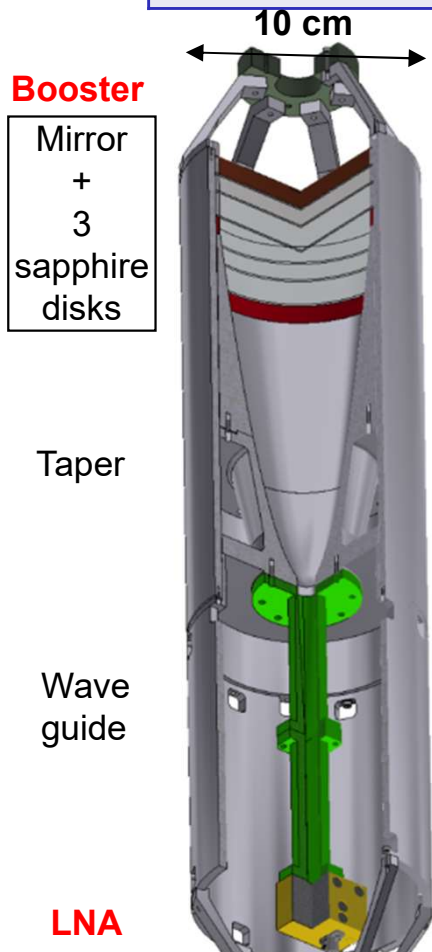


**Validate piezo motors + booster mechanics**

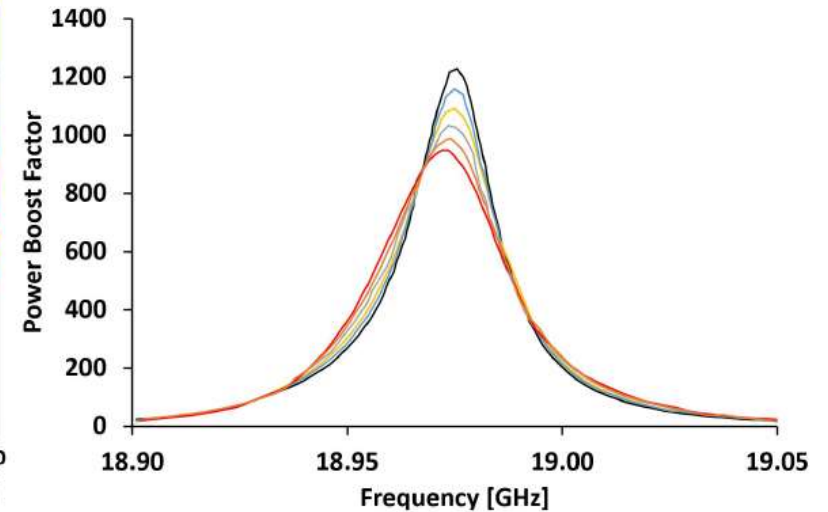
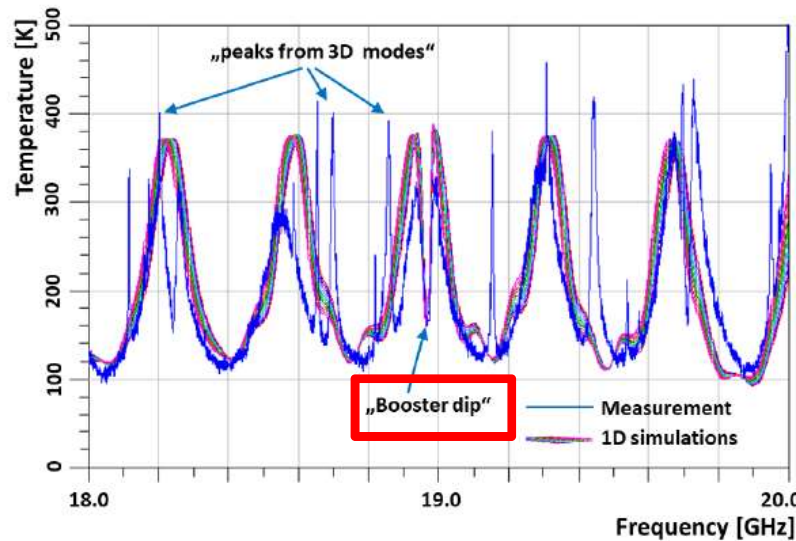


# RF Studies (1/2)

Name	Goal	Concept	Made of	Avail.	Lab Test
CB100	RF studies	Closed booster	3 <b>fixed</b> disks $\phi = 100\text{mm}$	2021	2022



Simulate **LNA** (ADS) and **Booster+taper** system (3D COMSOL)  
 → compare with measured system temperature in 18-20 GHz



Simulation reproduces system temperature  
 →  $\beta^2 = \mathcal{O}(1000)$  at 19 GHz ( $m_{ALP} = 78.5 \mu\text{eV}$ )

# RF Studies (2/2)

Name	Goal	Concept	Made of	Avail.	Morpurgo Test
CB100	RF studies	Closed booster	3 fixed disks $\phi = 100\text{mm}$	2021	2022

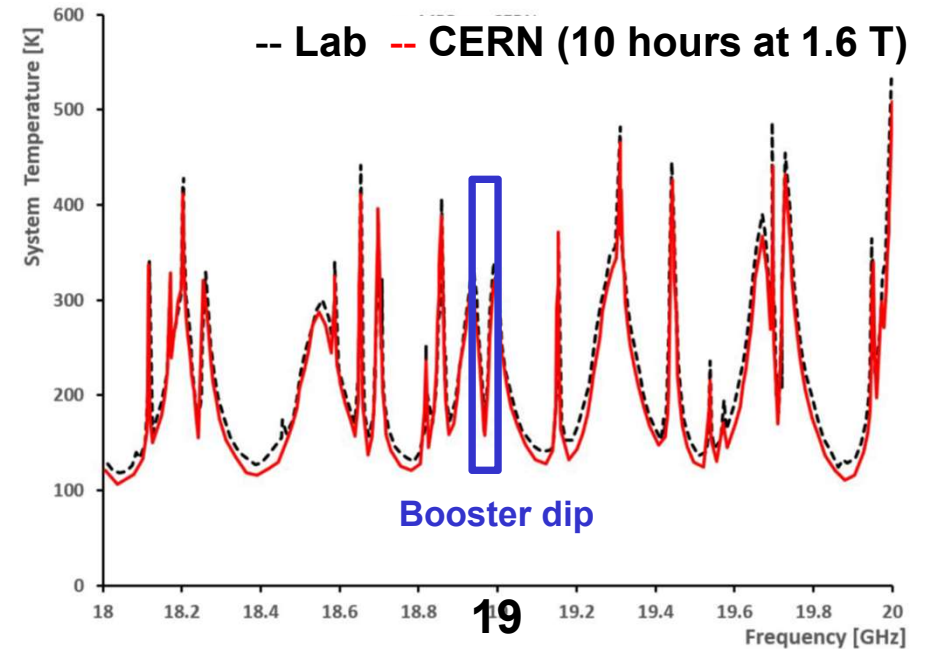
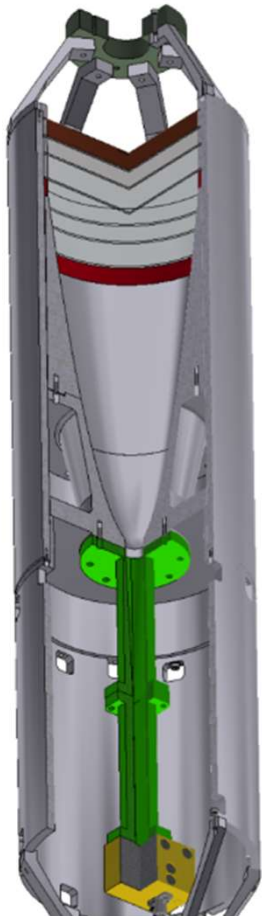
## Booster

Mirror  
+  
3  
sapphire  
disks

Taper

Wave  
guide

LNA

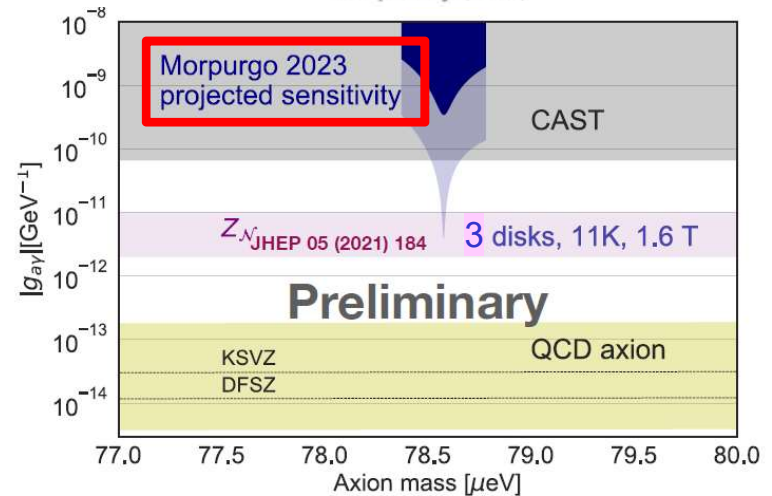
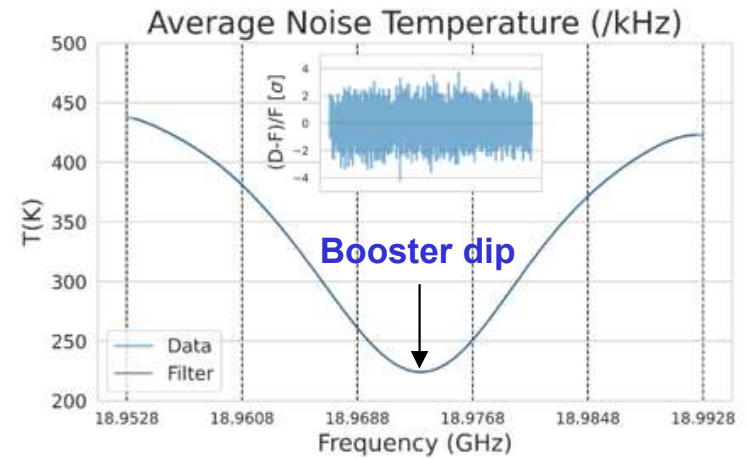
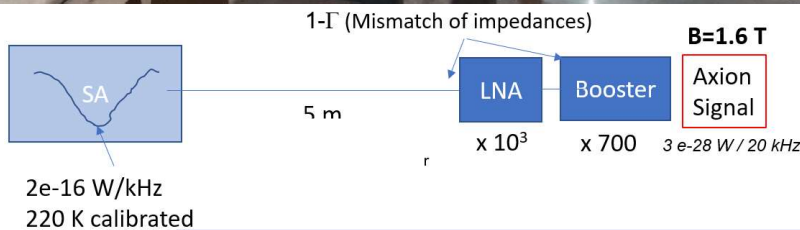
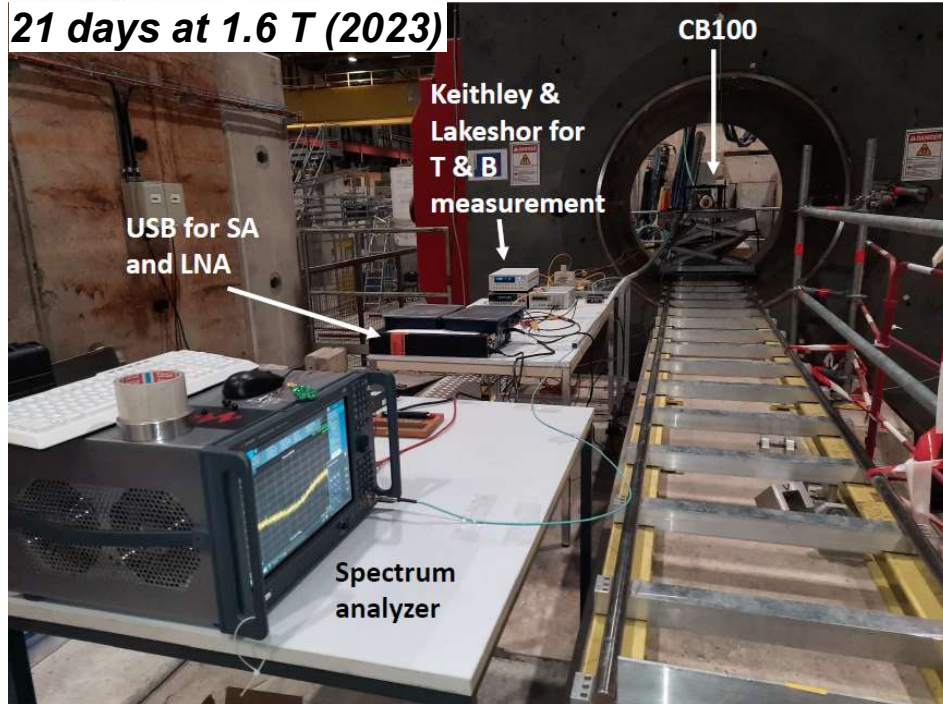


**No impact from the CERN environment**

# First Physics

Name	Goal	Concept	Made of	Avail.	Morpurgo Test
CB100	First physics	Closed booster	3 fixed disks $\phi = 100\text{mm}$	2021	2023

21 days at 1.6 T (2023)



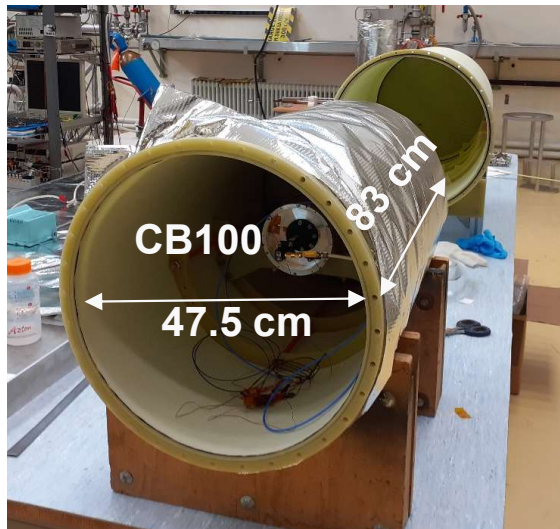


# First ALP Physics

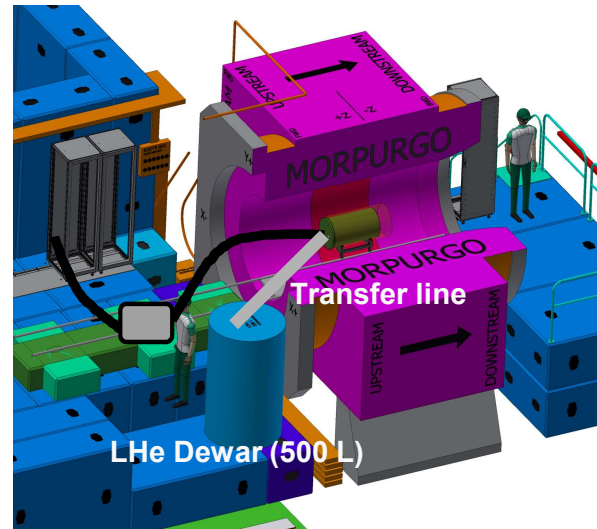
Name	Goal	Concept	Made of	Avail.	Morpurgo Test
CB100	RF studies + First physics	Closed booster	3 fixed disks $\phi = 100\text{mm}$	2021	<b>2024</b>

Develop a 'cheap' cryostat with CERN cryolab to cool the booster + LNA  $\rightarrow$  Validated the principle in 2023

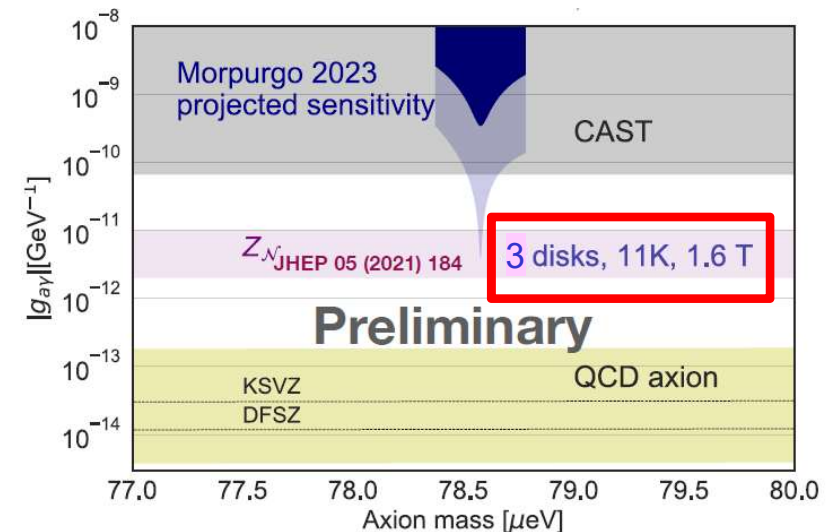
Provided we understand calibration at cold  $\rightarrow$  Improve reach



Vacuum between Inner and Outer vessel



He flow from dewar to cryo  
 $\rightarrow$  cooling in  $\frac{1}{2}$  day  
 $\rightarrow$  Stable  $<10\text{K}$  during  $>10$  hours

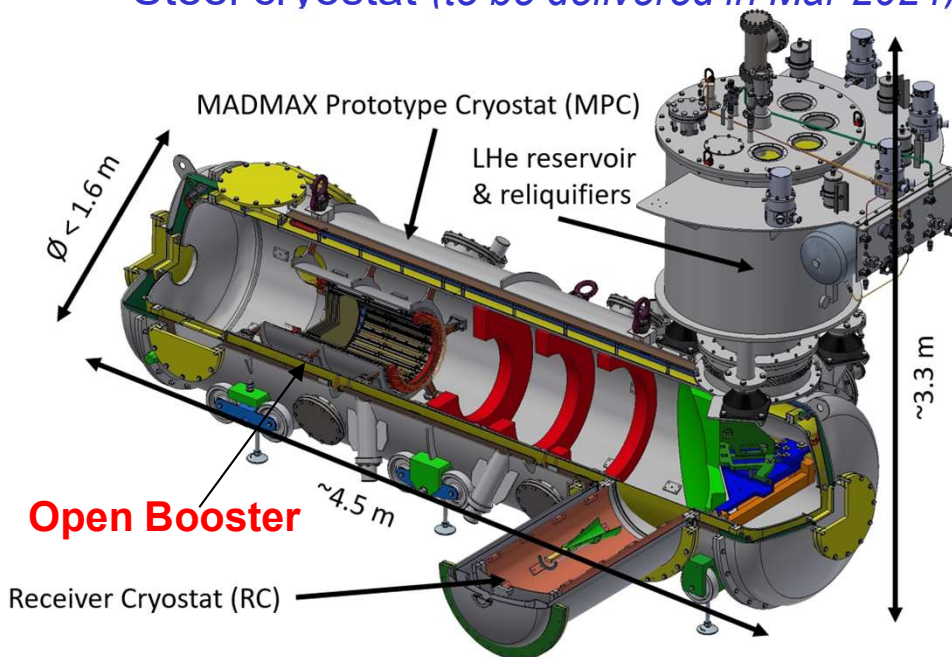


**Short cold run**

# ALP Physics ++

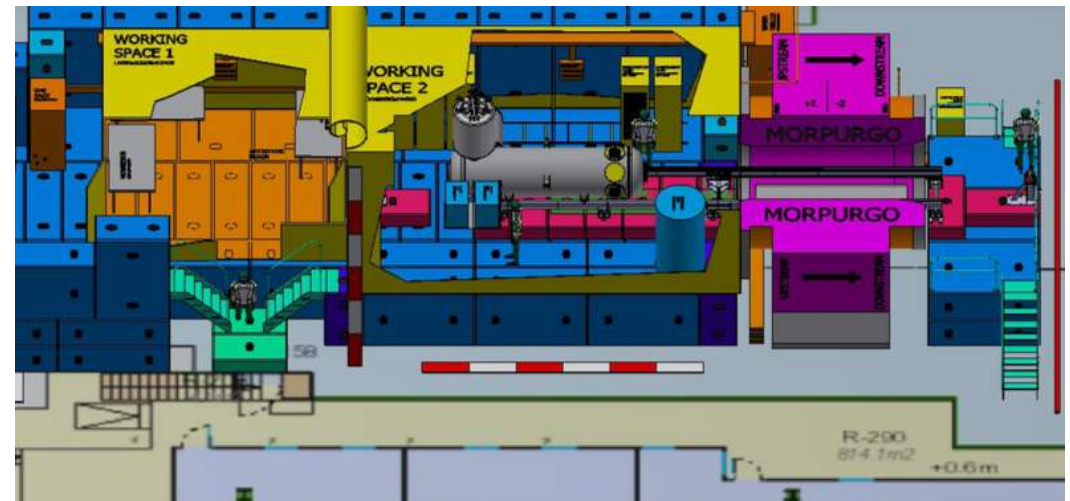
Name	Goal	Concept	Made of	Avail.	Morpurgo Test
Proto-3	Scan ALP around 100 $\mu\text{eV}$	Open Booster	3 moveable disks $\phi = 300$ mm	2024	<b>2025, 26?</b>

Open Booster inserted in a Stainless Steel cryostat (to be delivered in Mar 2024)



Open Booster built with P200 and CB100 experience

Morpurgo CERN area refurbished to host the SS cryostat

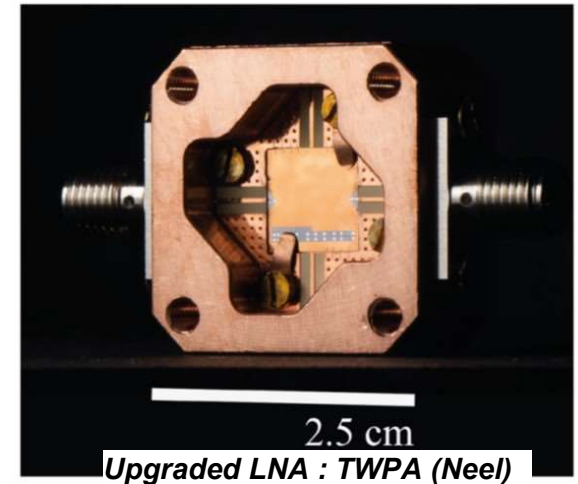
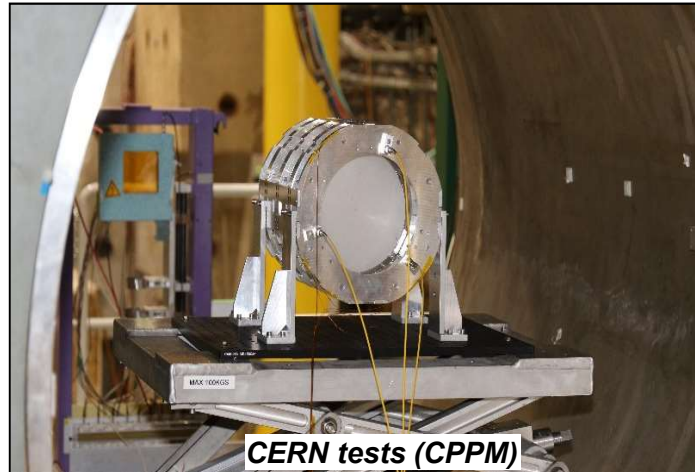


**Long cold run + mass scan**

# MADMAX and France (1/2)

## ❑ Two French institutes joined MADMAX in 2020

- **CPPM** (2 physicists, 3 engineers, 1 PhD) : booster high precision mechanics, CERN infrastructure and coordination. Supported by IN2P3 for 4 years starting Jan 2023
- **Institut Neel** : ultra-low noise amplifier
- + **CEA-IRFU** : work for the final magnet
- + **IRL DMLab** : installed at DESY → MADMAX is one of the supported project

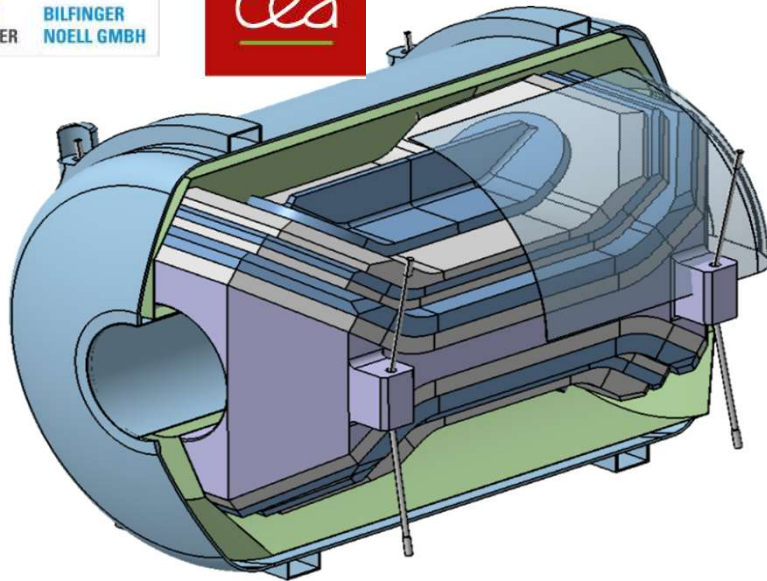


**MADMAX looking for new (French) institutes to join !**

# MADMAX and France (2/2)

## □ Progresses on final magnet

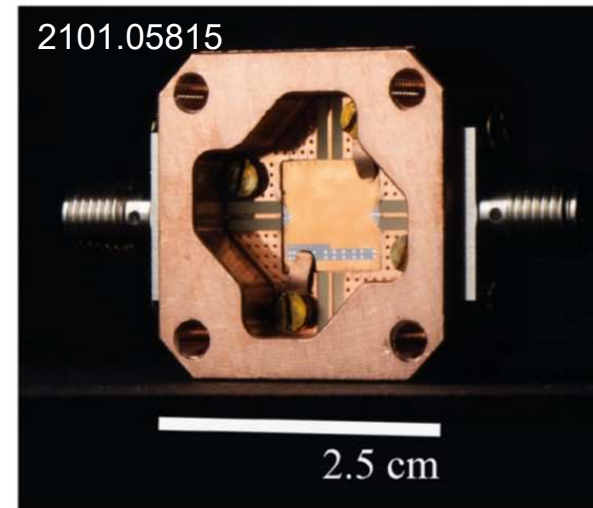
- Design completed: 2x9 skateboard coils with novel copper CICC conductor [NbTi with Cu jacket @ 1.8K]



- Recently demonstrated that coils will be safe in terms of quench protection
- **Next** : Design, manufacture and test a small MADMAX coil (6T)

## □ Progresses on final receiver

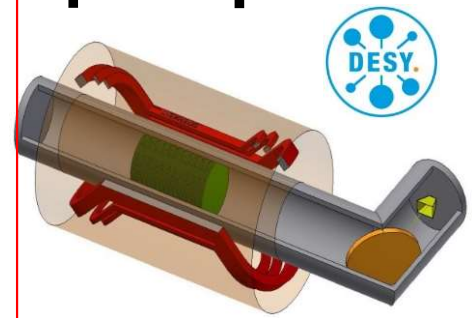
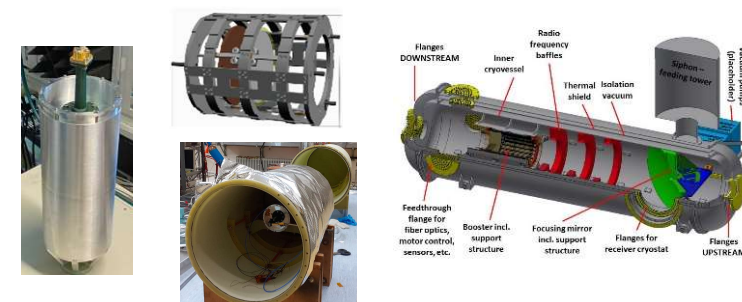
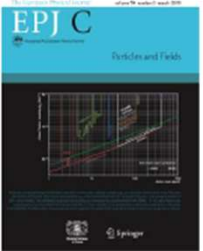
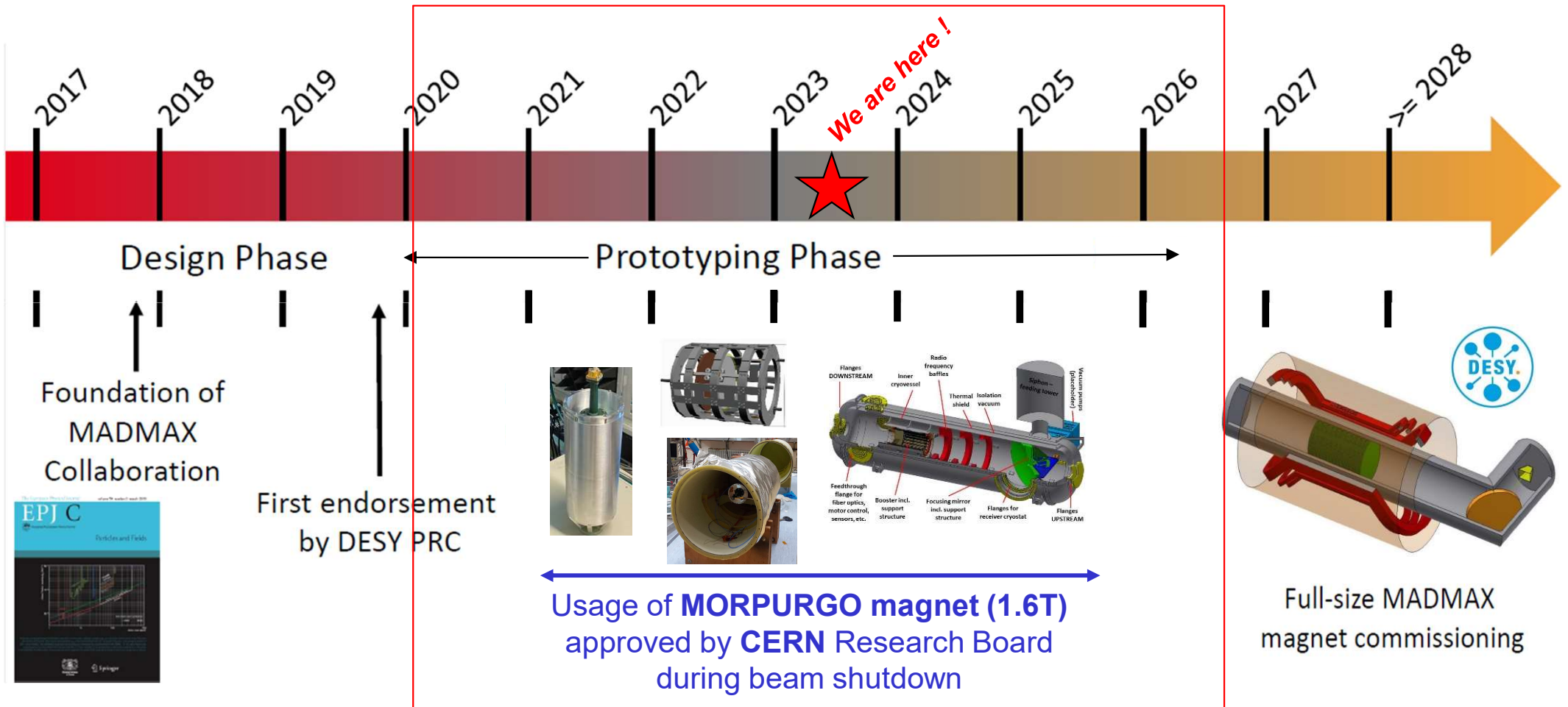
- Very low noise pre-amplifier [ $P_{sig} \sim T_{sys}$ ] HEMT (G=33 dB, 4K added noise) below 40 GHz
- Josephson Junction being developed to further minimize noise (*quantum limit*)



TWPA prototype with  $G > 20$  dB and 1K added noise at 10 GHz

- **Next**: >40 GHz techno. to be developed

# MADMAX timescale



$$g_{\alpha\gamma} \propto \left(\frac{1000-10000}{50000 \beta^2}\right)^{1/2} \times \left(\frac{10-220}{4 \text{ K}} T_{sys}\right)^{1/2} \times \left(\frac{1.6}{B_e}\right) \times \left(\frac{0.01-0.3}{A}\right)^{1/2} \times \left(\frac{4 \text{ days}}{t}\right)^{1/4} \times \left(\frac{\text{SNR}}{5}\right)^{1/2}$$

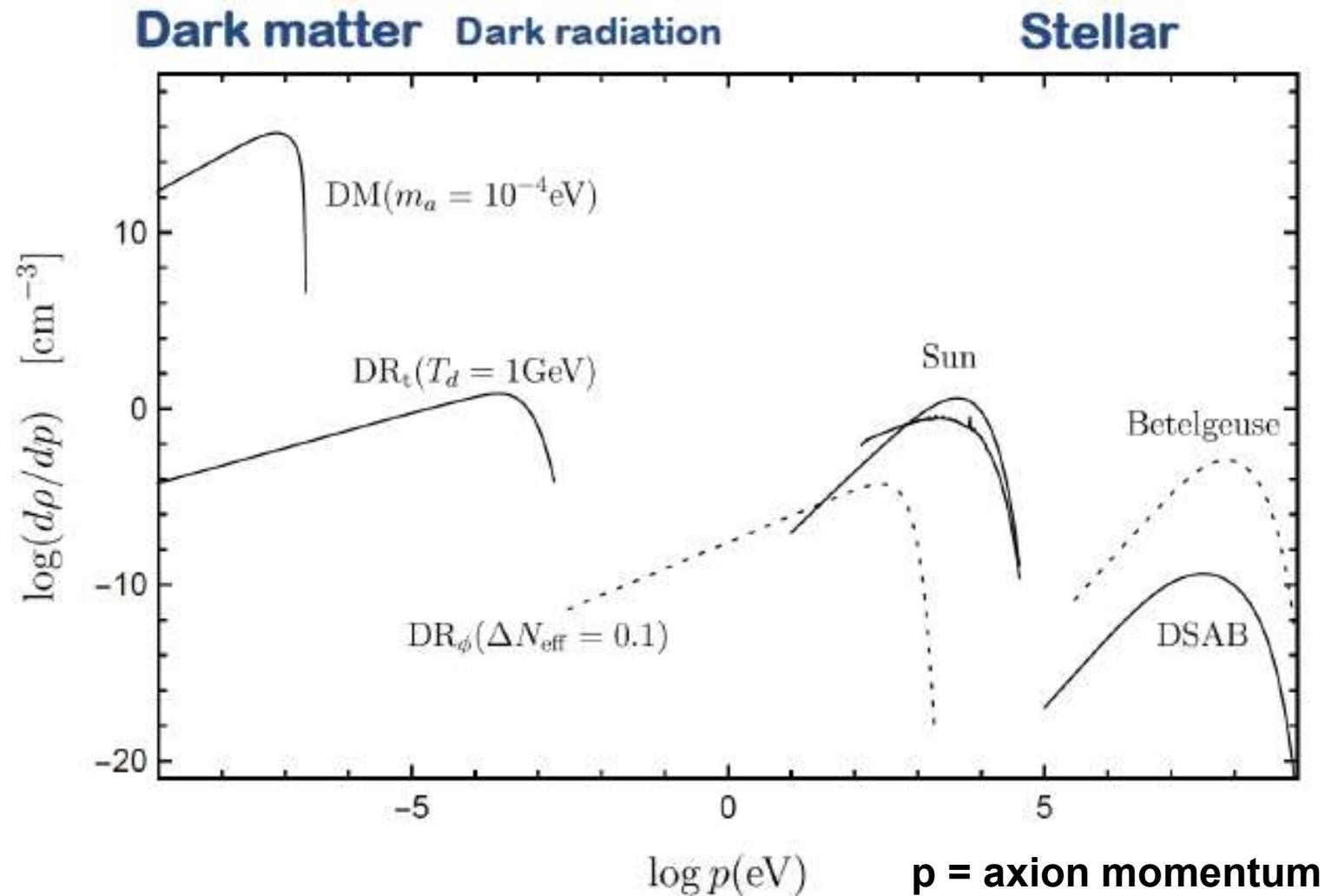


# BACKUP

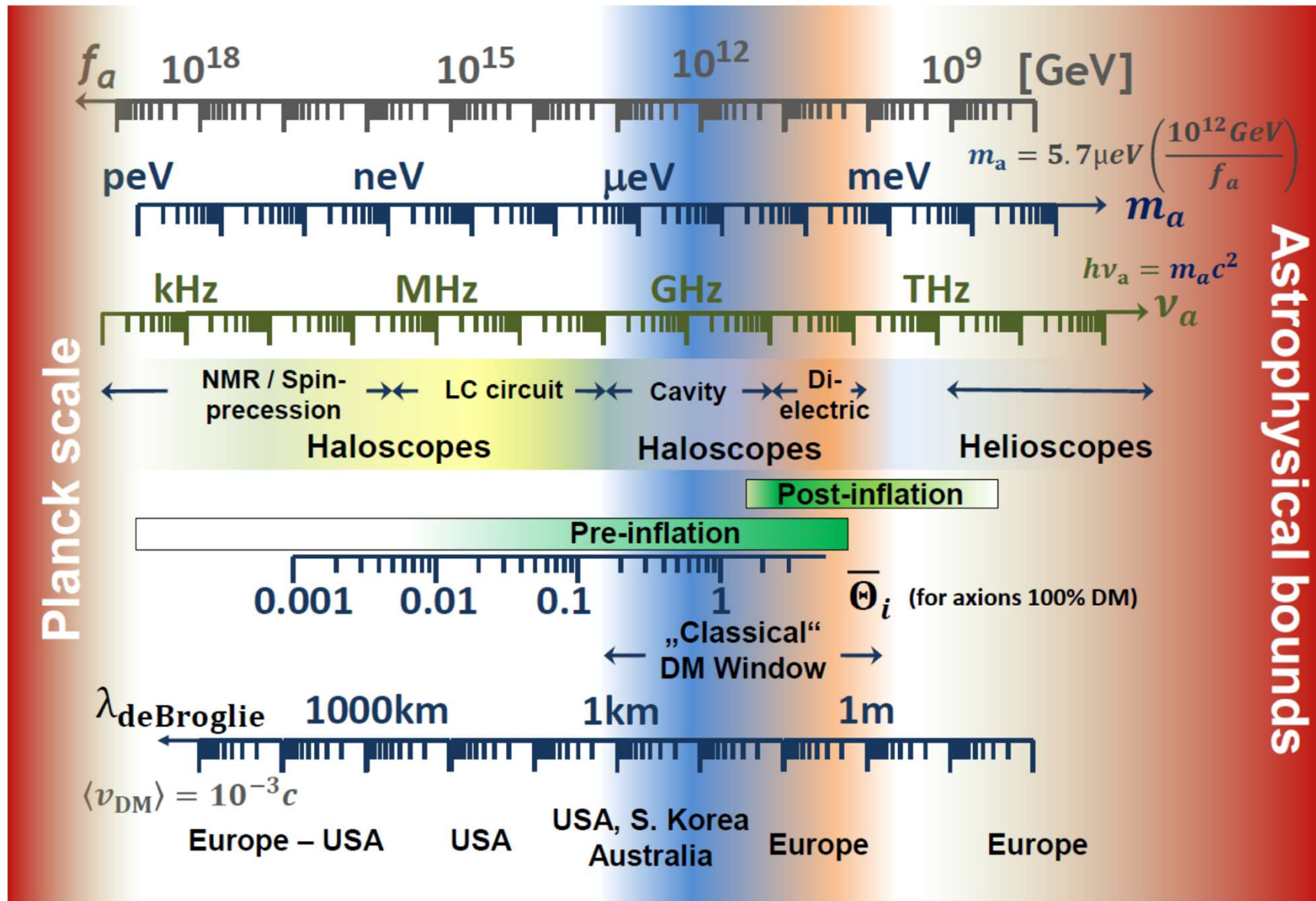
*Collab Week at CPPM (Apr 2023)*



# Sources of axions



# Axion scales

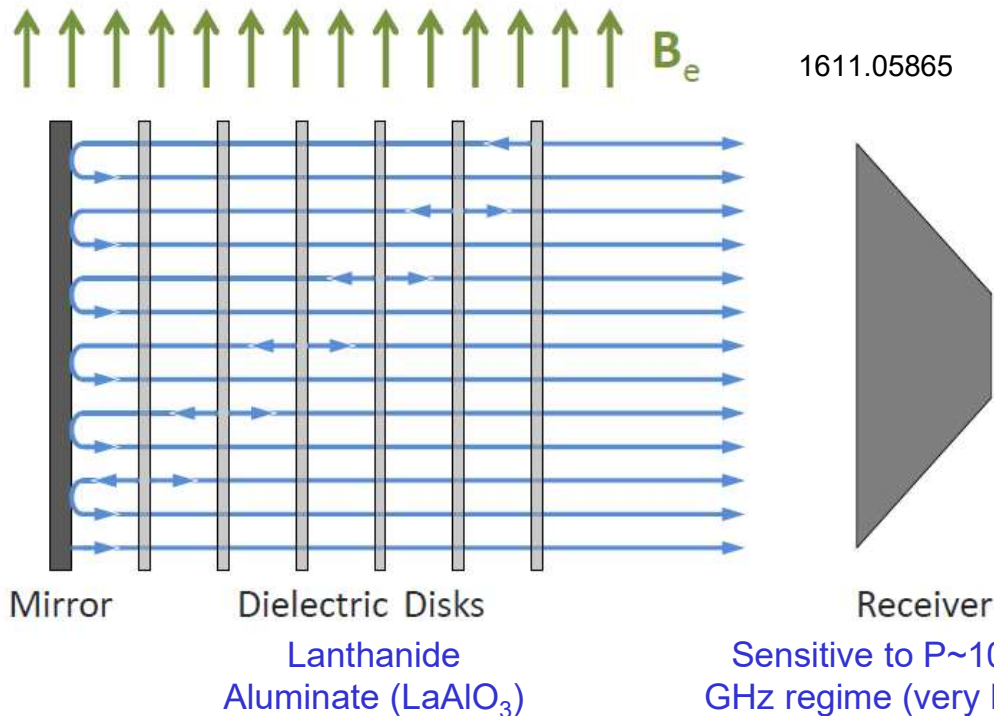




# Dielectric haloscope

## Dielectric haloscope → MadMax experiment

- New experimental concept to alleviate cavity limitation at high  $m_a$  ( $V \sim 1/m_a^3$ )
- Discs + mirror in  $B_e \rightarrow$  wave emission @ interfaces  $\rightarrow$  constructive interferences + resonances



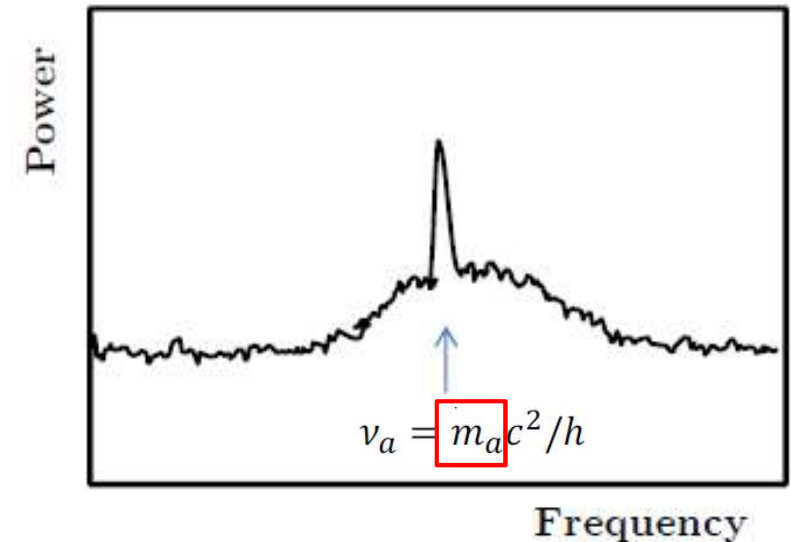
Power / Area

$$P/A = 2.2 \times 10^{-27} \text{ W m}^{-2} \left( \frac{B_e}{10 \text{ T}} \right)^2 C_{a\gamma}^2 \beta^2$$

[if  $\lambda_\gamma/2$  = distance between discs]

Signal boost  $\times 10^5$

$\rightarrow P \sim 10^{-22} \text{ W}$  ( $A = 1 \text{ m}^2$ ,  $B = 10 \text{ T}$ )

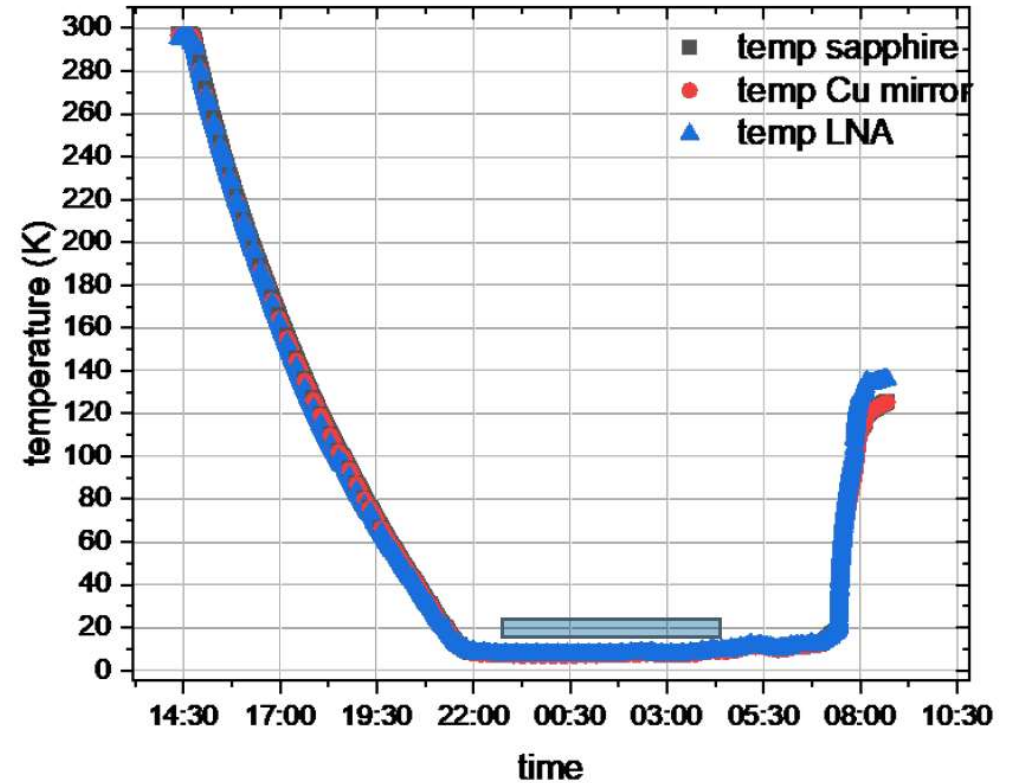
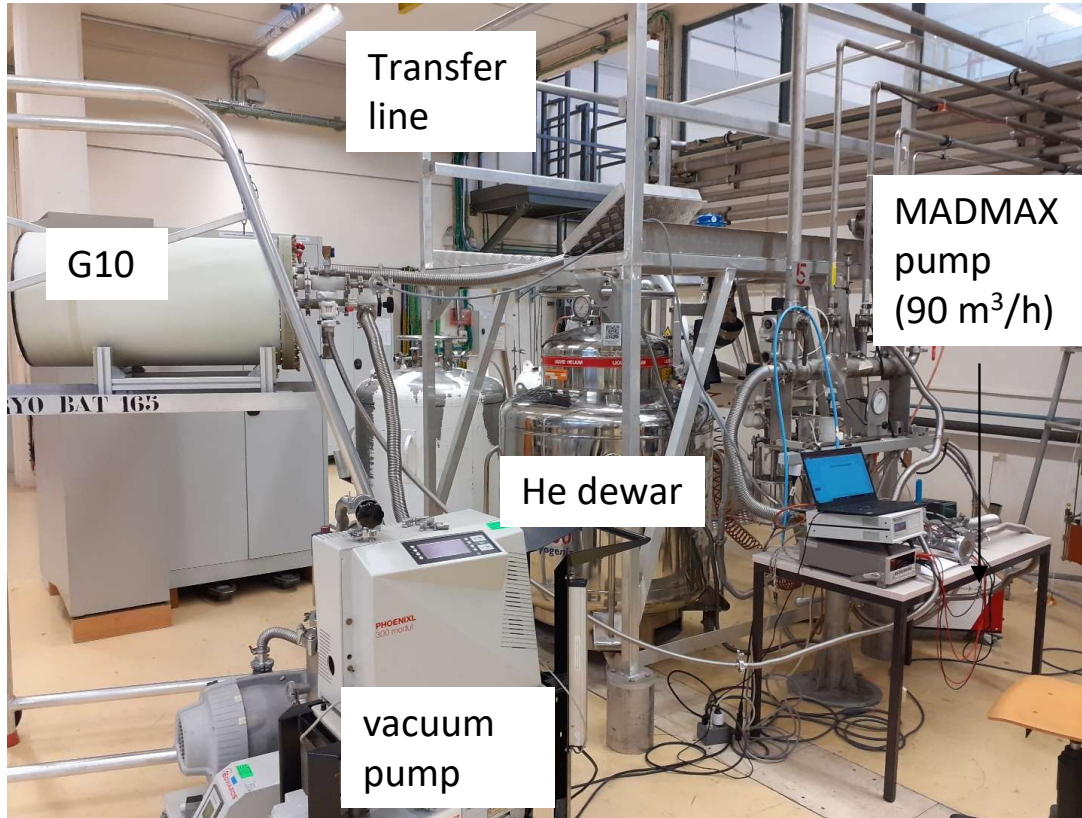


$[E = h\nu_a = m_a c^2 + 0.5 m_a v_a^2, v_a = 10^{-3} c \rightarrow h\nu_a = m_a c^2]$

- Adjustable distance between discs  $\rightarrow$  tune to scan  $m_a$   
[spacing 15 mm for 40  $\mu\text{eV}$  and 1.5 mm for 400  $\mu\text{eV}$ ]

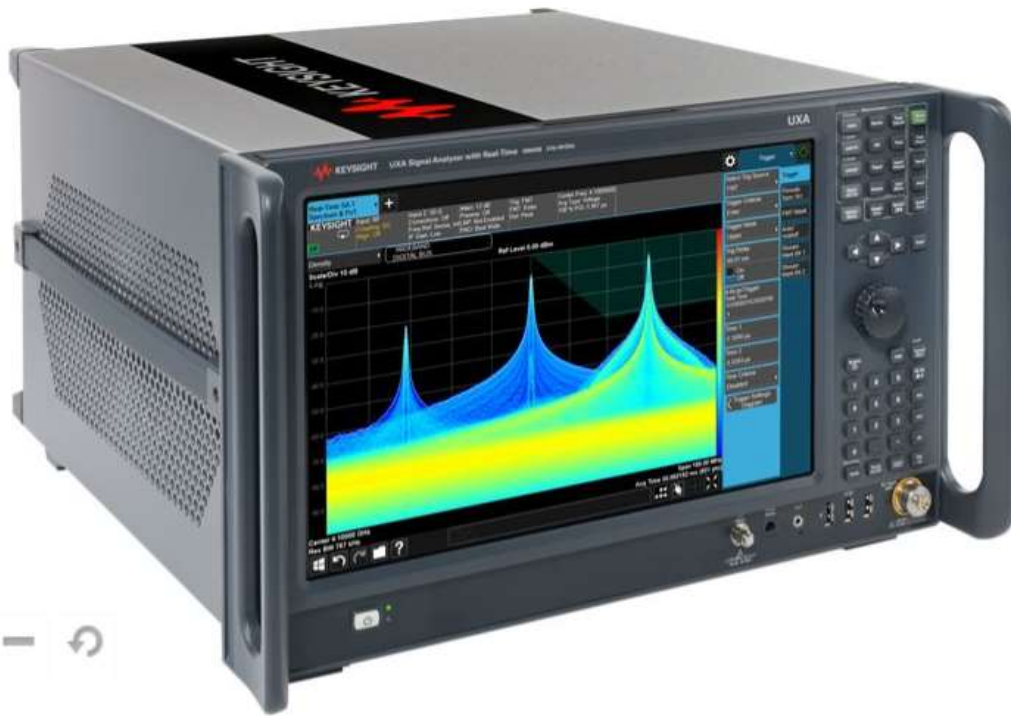
$\rightarrow$  MadMax only capable to explore  $m_a = 40\text{-}400 \mu\text{eV}$  (favored by post-inflation theory)

# G10 cryostat



# Spectrum Analyzer

2023 → Keysight : N9040B

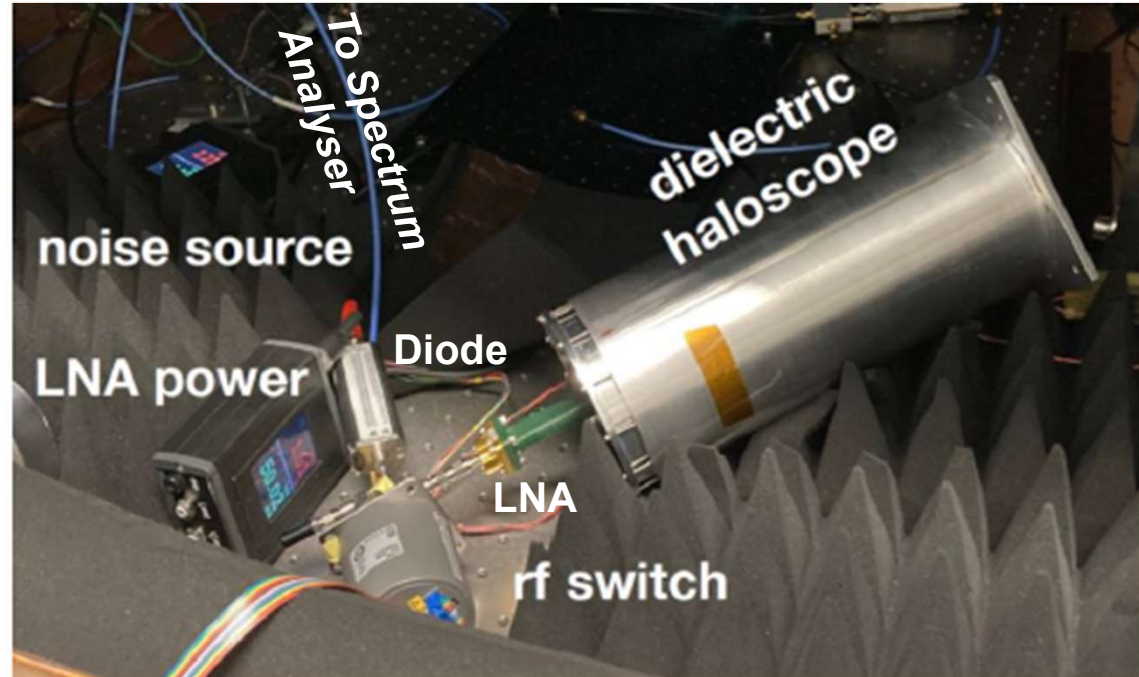
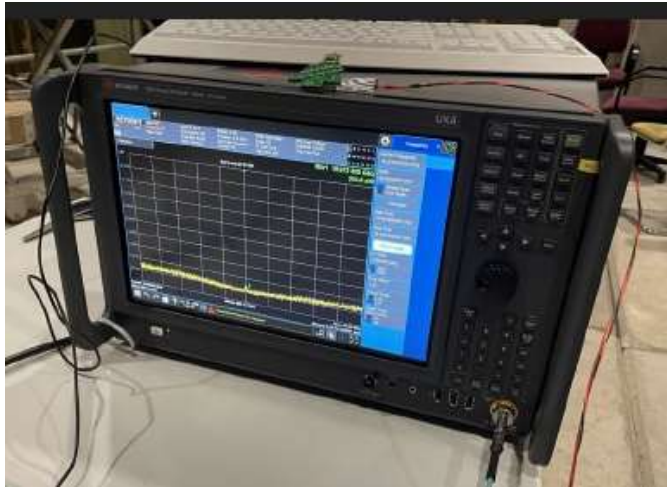


2024 → Rhode&Schwarz FSW26 with streaming option



# System calibration (1/3)

Spectrum Analyser



Diode

FREQ GHz	ENR DB
.01	21.23
0.1	19.89
1.0	19.69
2.0	19.24
3.0	19.29
4.0	19.28
5.0	19.34
6.0	19.23
7.0	19.58
8.0	20.06
9.0	20.00
10.0	20.30
11.0	20.86
12.0	20.95
13.0	20.97
14.0	21.55
15.0	22.14
16.0	22.04
17.0	22.02
18.0	22.19
19.0	22.37
20.0	22.31
21.0	21.84
22.0	23.14
23.0	22.71
24.0	21.63
25.0	20.89
26.0	21.11
26.5	20.12

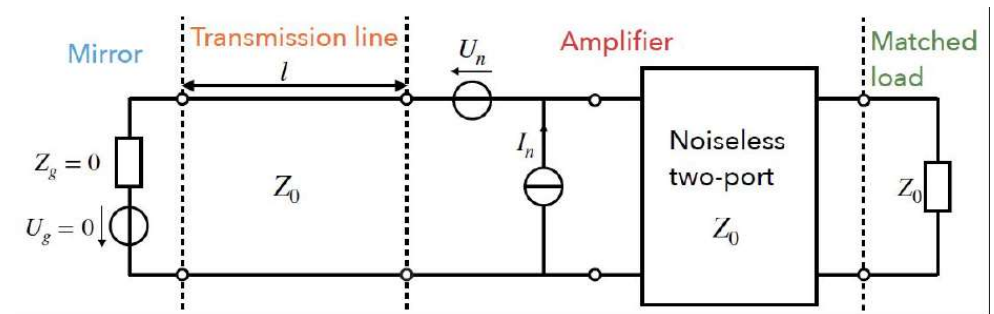
SERIAL NO.  
W149

## 1- Power (W/kHz) to Thermal Noise Temp. (K) :

- Use a well calibrated diode with a 30 dB Attenuator
- T Diode On = Room Temperature + 50K = 345 K
- T Diode Off = Room Temperature = 295 K
- With P (Diode On), P (Diode Off), estimate reflections
- From P (LNA + Booster), P (Diode On), P (Diode Off) deduce **T (LNA + Booster)**

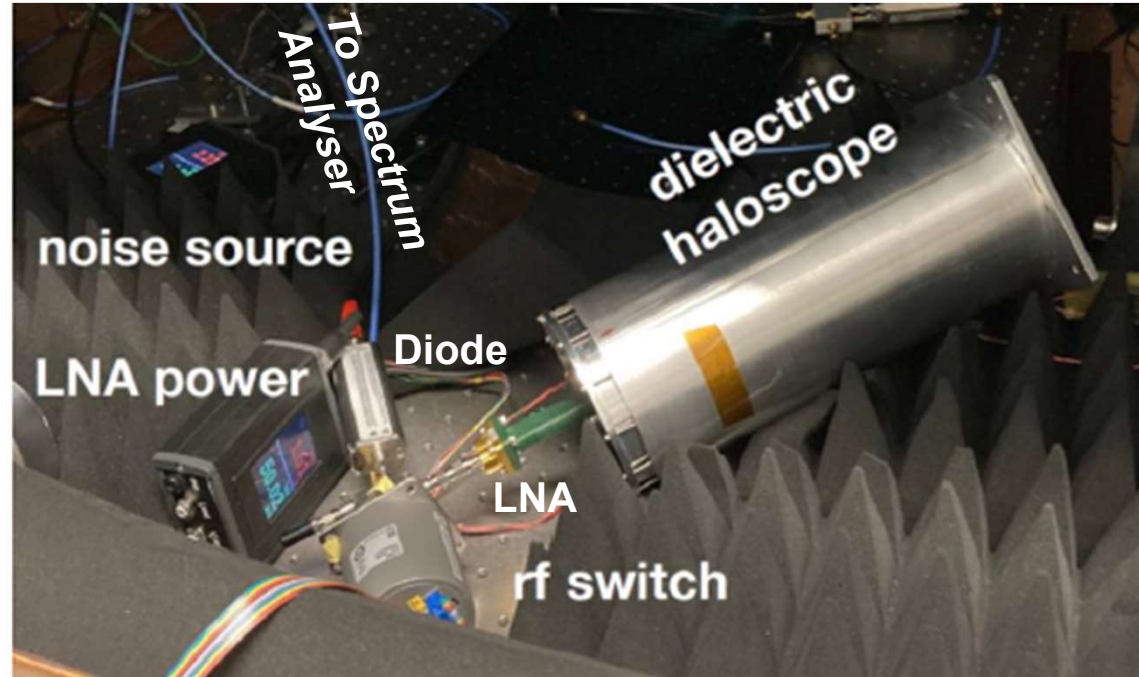
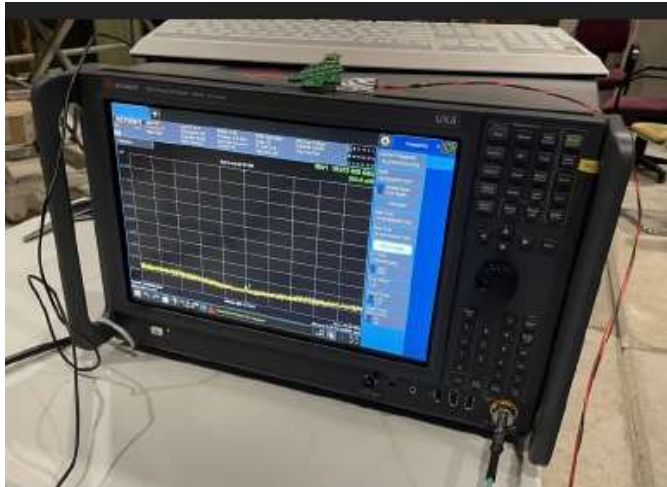
## 2- ADS model ( $I_n$ , $U_n$ ) for LNA Noise

- Short / Open / Load with RF switch



# System calibration (2/3)

Spectrum Analyser



Diode

FREQ GHz	ENR DB
.01	21.23
0.1	19.89
1.0	19.69
2.0	19.24
3.0	19.29
4.0	19.28
5.0	19.34
6.0	19.23
7.0	19.58
8.0	20.06
9.0	20.00
10.0	20.30
11.0	20.86
12.0	20.95
13.0	20.97
14.0	21.55
15.0	22.14
16.0	22.04
17.0	22.02
18.0	22.19
19.0	22.37
20.0	22.31
21.0	21.84
22.0	23.14
23.0	22.71
24.0	21.63
25.0	20.89
26.0	21.11
26.5	20.12

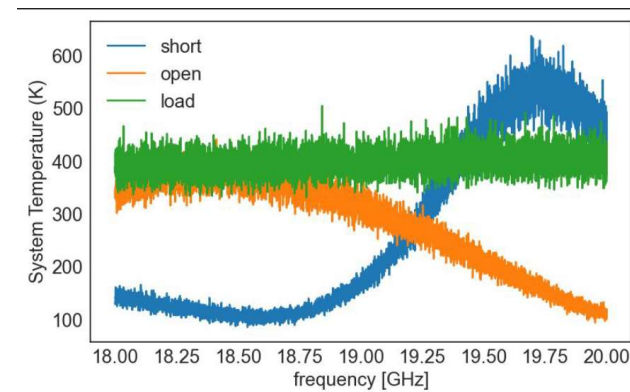
SERIAL NO.  
W149

## 1- Power (W/kHz) to Thermal Noise Temp. (K) :

- Use a well calibrated diode with a 30 dB Attenuator
- T Diode On = Room Temperature + 50K = 345 K
- T Diode Off = Room Temperature = 295 K
- With P (Diode On), P (Diode Off), estimate reflections
- From P (LNA + Booster), P (Diode On), P (Diode Off) deduce T (LNA + Booster)

## 2- ADS model ( $I_n$ , $U_n$ ) for LNA Noise

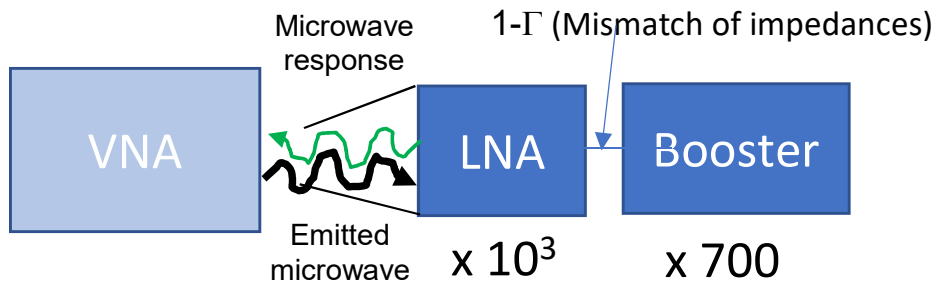
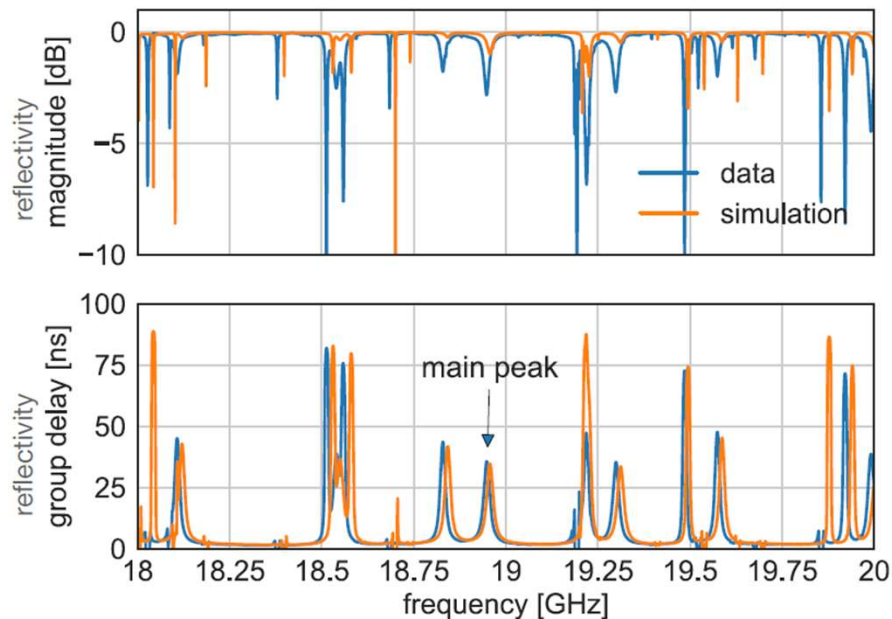
- Short / Open / Load with RF switch



# System calibration (3/3)

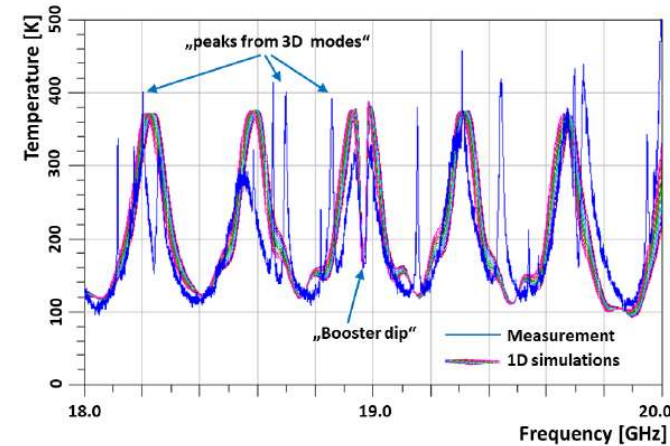
## 3- Reflectivity measurements with a VNA

- Should match the 3D COMSOL simulation of the booster + taper



## 4- Merge ADS and COMSOL simulations to predict T (K)

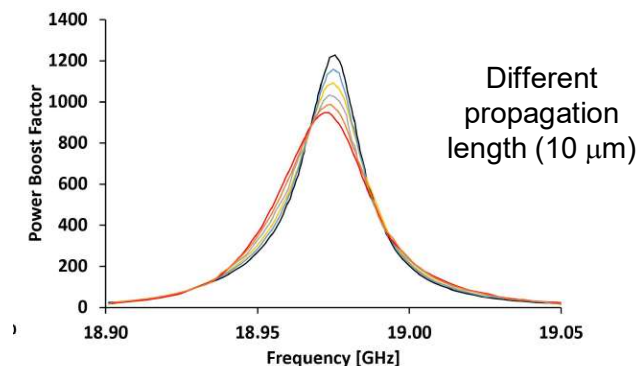
- Should match the measured T (SA)



- Wavy because of coherent and destructive interference (different propagation length) when injecting the LNA noise in the booster

## 5- Deduce the Booster factor from the model

- Including uncertainties



# MadMax activities at CPPM

*CPPM pioneer at IN2P3 in direct searches for axions, world rising activity and in particular in Germany (DESY "Axion Hub")*

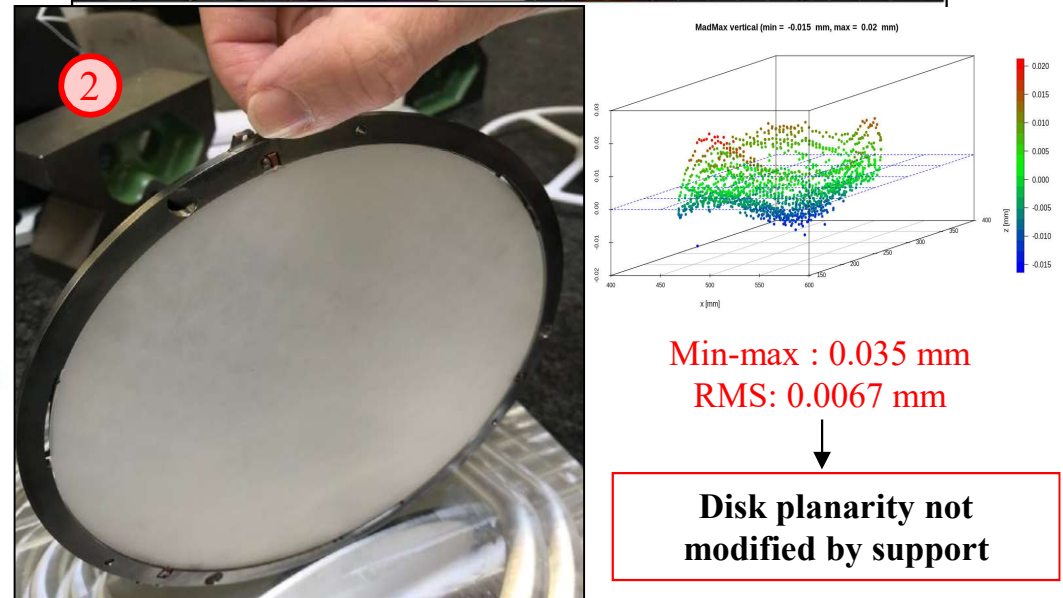
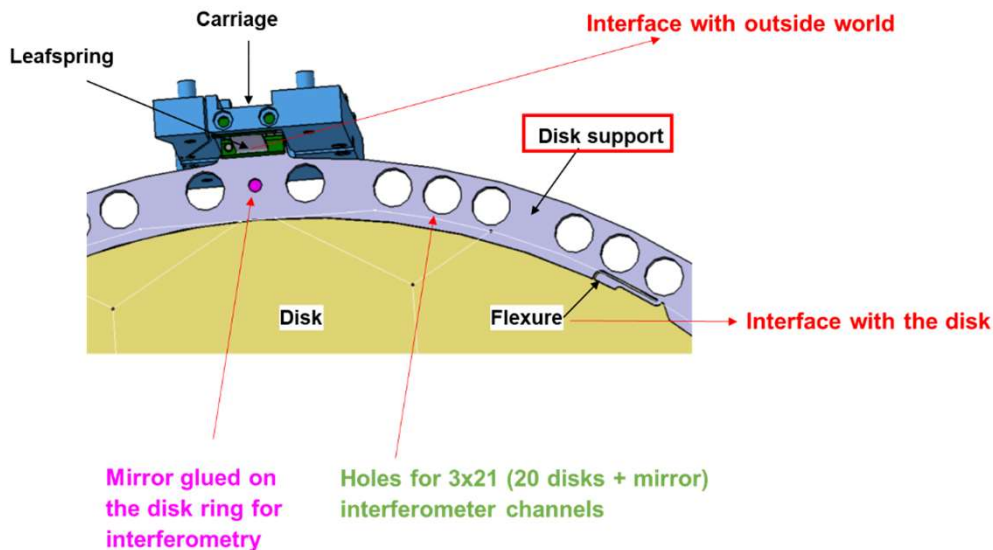
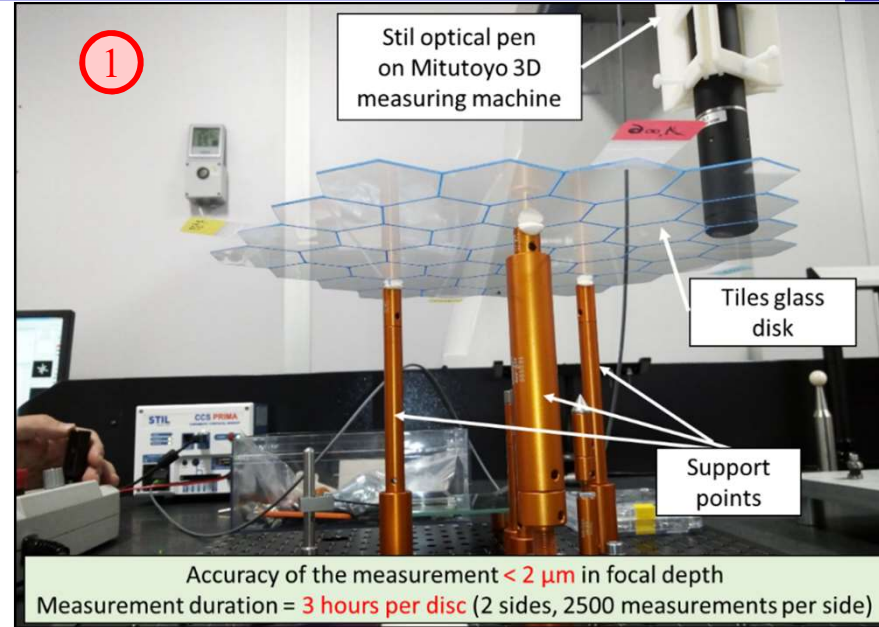
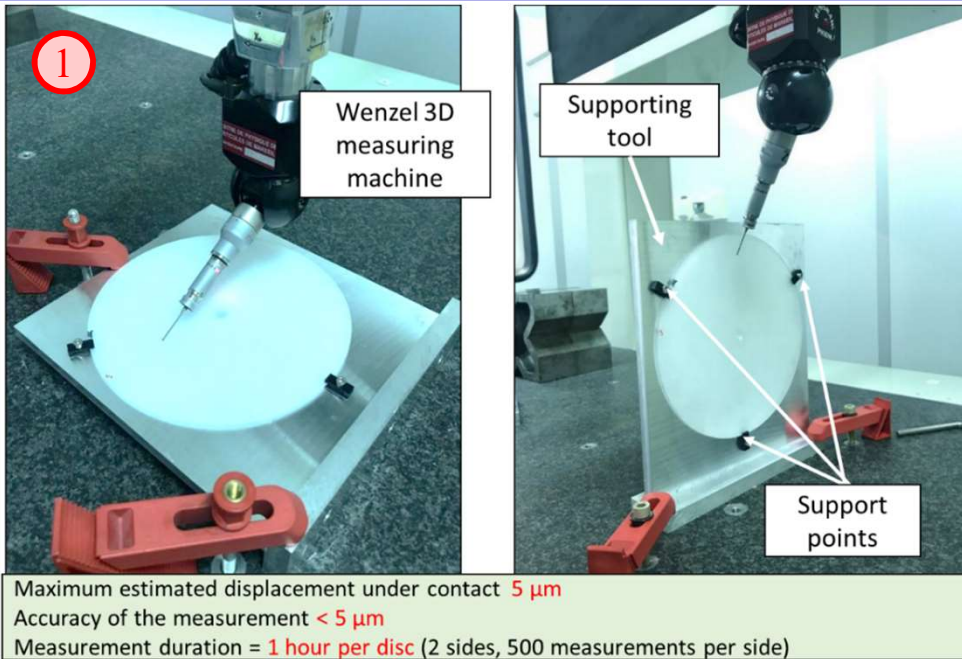
## □ Precision mechanics at CPPM for the prototype boosters

- ① Precision 3D measurements  $O(\mu\text{m})$  for geometry control of the prototype disks
  - CPPM expertise/infrastructure for precision measurements (e.g. ATLAS pixels)
- ② Conception/fabrication of disk support rings
  - Interfaces between disks, piezo motors and interferometer system
  - Cutting edge and challenging R&D → Optimisation of fabrication process to obtain best planarity ( $<10\mu\text{m}$ )

## □ Infrastructures at CERN for protos tests protos in Morpurgo magnet

- ③ Conception, fabrication and installation of mechanical interfaces prototypes-Morpurgo
  - Rails for electric racks, supports for prototypes, rails for big test cryostat
- ④ Coordination of tests at CERN (programme 2021-2025 approved by SPSC) → tests prototypes CB100 (physics) and P200 (meca) in April 2022 in Morpurgo magnet 1.6T

# MadMax activities at CPPM



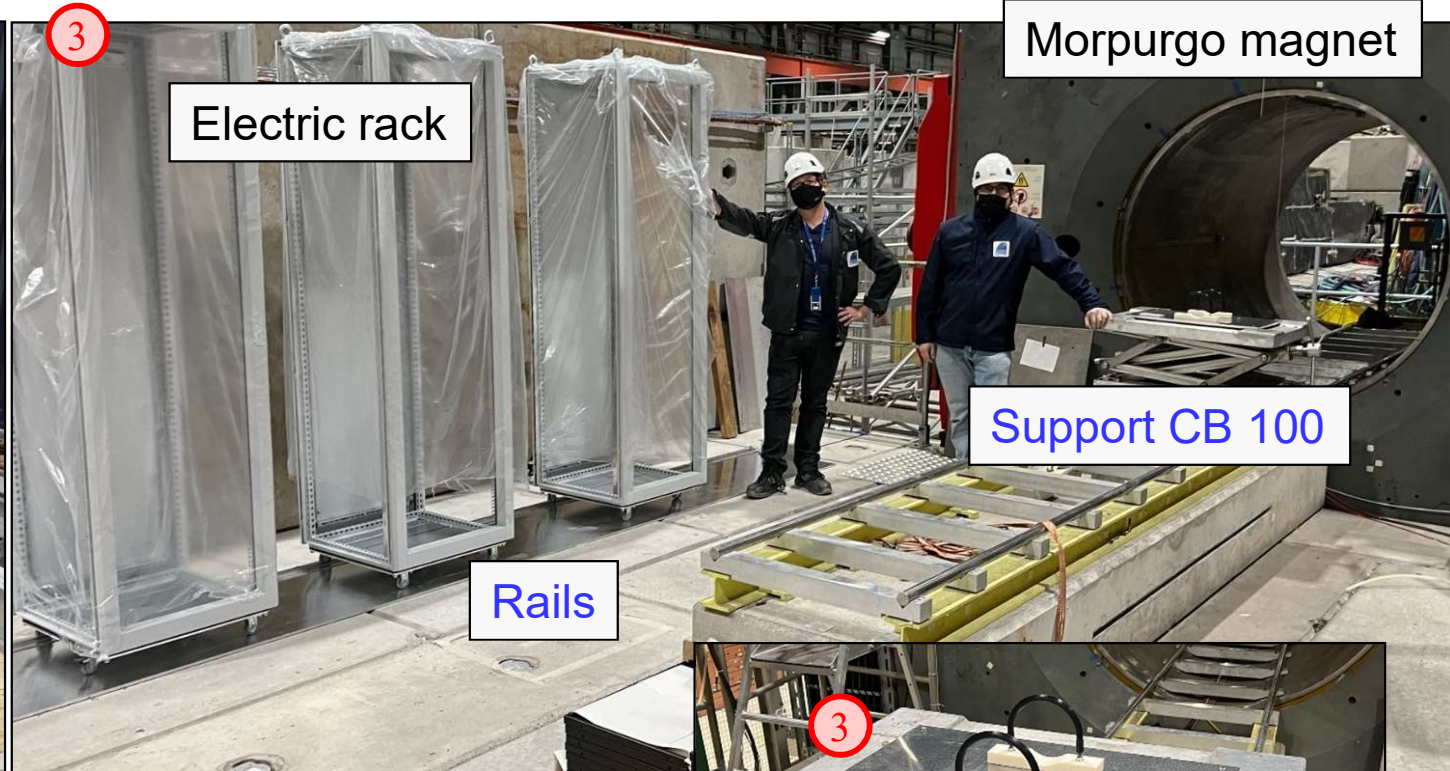


# MadMax activities at CPPM



3

Rails for Cryostat



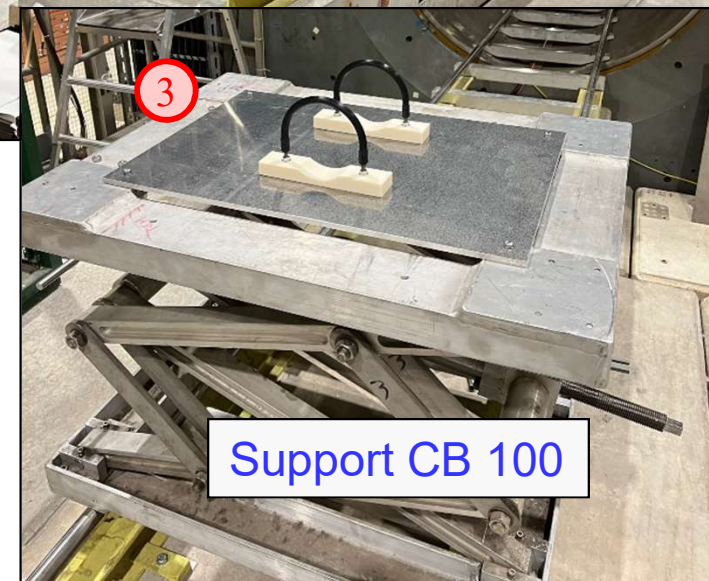
3

Electric rack

Rails

Morpurgo magnet

Support CB 100



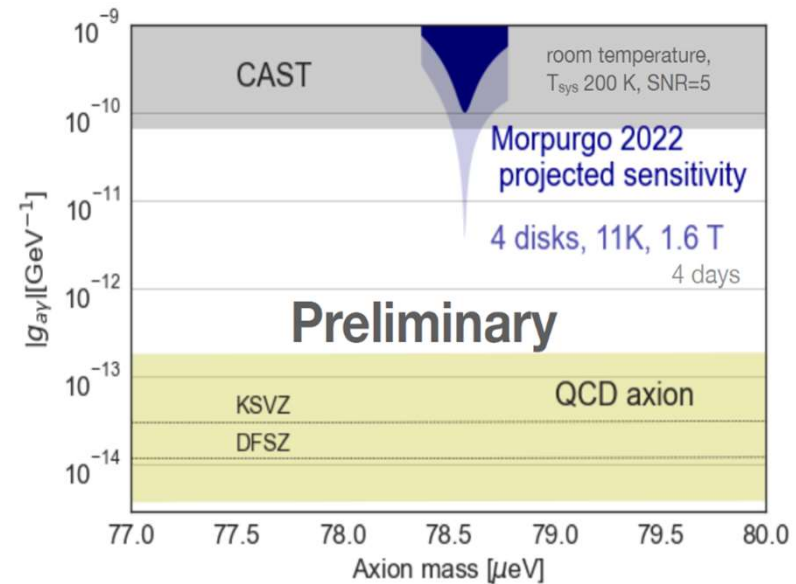
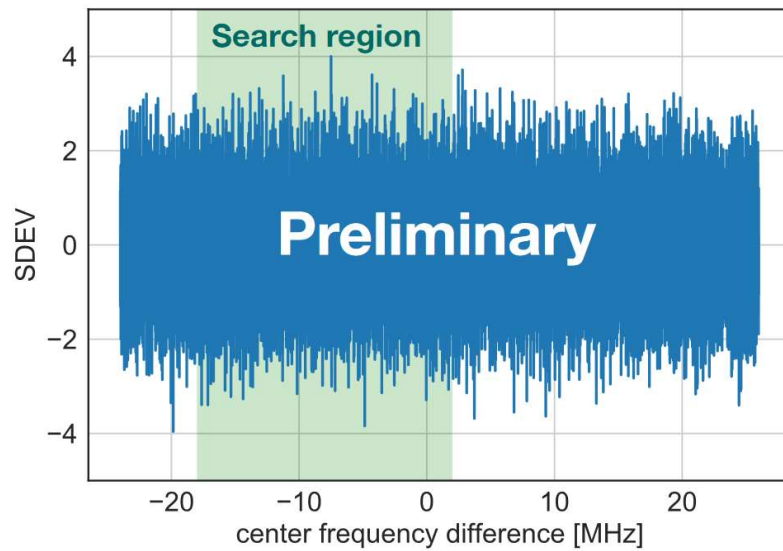
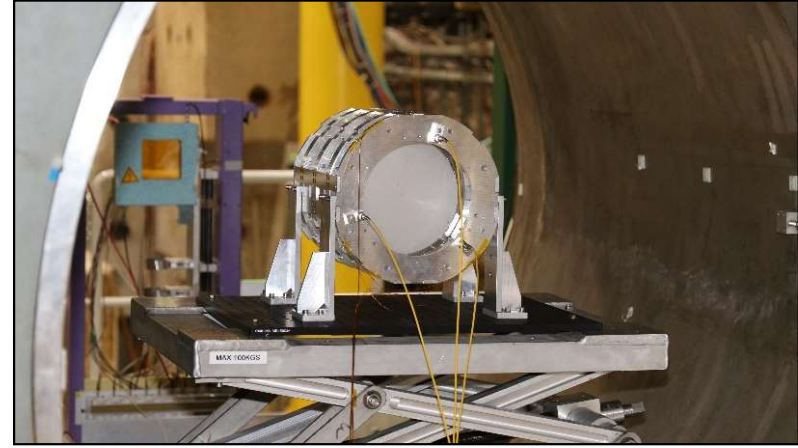
3

Support CB 100

# MadMax activities at CPPM

4

Tests of two first prototypes in April 2022 in Morpurgo (during SPS shutdowns)



→ CPPM drives magnet @ CERN + participates to search for ALP @  $\sim 80 \mu\text{eV}$