Interoperability in Multimessenger Astrophysics

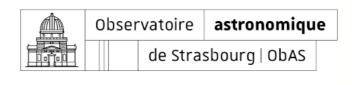
1st Astro-COLIBRI Multimessenger Astrophysics Workshop

Ada Nebot



ASTRONOMIQUES DE **S**TRASBOURG





□ Index

- Interoperability, the VO and the IVOA
 - What, Who, Where, How, When?
- (Some of the) Multimessenger needs
- Does the IVOA cover those needs?
- VOEvent, status and future evolution?

Interoperability

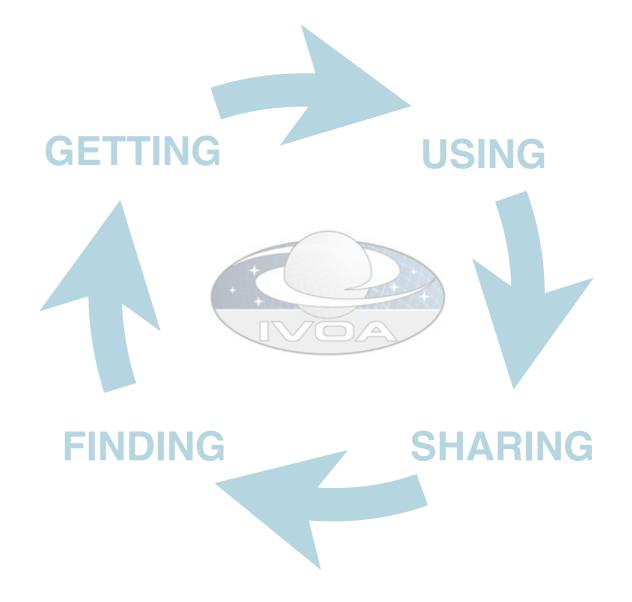
- Few definitions:
 - Interoperability:
 - "The ability of computer systems or software to exchange and make use of information."
 - "The ability of different systems, devices, applications or products to connect and communicate in a coordinated way, without effort from the end user"
 - The Virtual Observatory: "Framework for astronomical datasets, tools, services to work together in a seamless way"

The VO and the IVOA: what?

The VO : "A multi-wavelength digital sky that can be searched, visualised and analysed in new and innovative ways"

What is the International Virtual Observatory Alliance?

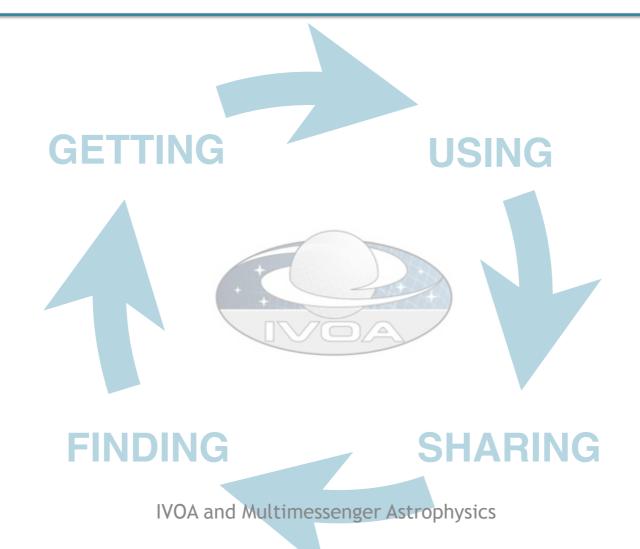
- A science driven organisation that builds the technical standards
- A place for discussing and sharing VO ideas and technology to enable science
- Promoting and publicising the VO



The VO and the IVOA: why?

Clear benefits

- Growth in the scientific return of data
- Capability to discover and fuse multiple data sets
- Application of the VO in planning new observations and observing strategies



The VO and the IVOA: who?

Who is the IVOA?

- 5 Committees: Exec, Tech Coordination, Standards & processes, Media, Science priorities
- 6 Working Groups (WG): Applications, data access, models, grid & web services, registry, semantics
- 8 Interest Groups (IG): Time-domain, radio, solar system, theory, operations, data curation, knowledge & discovery, education

Want to get involved?

- Meetings: 2 interoperability meetings per year
- Don't know where to start? Email any chair/vice-chair of a IG/WG, CSP

http://ivoa.net/

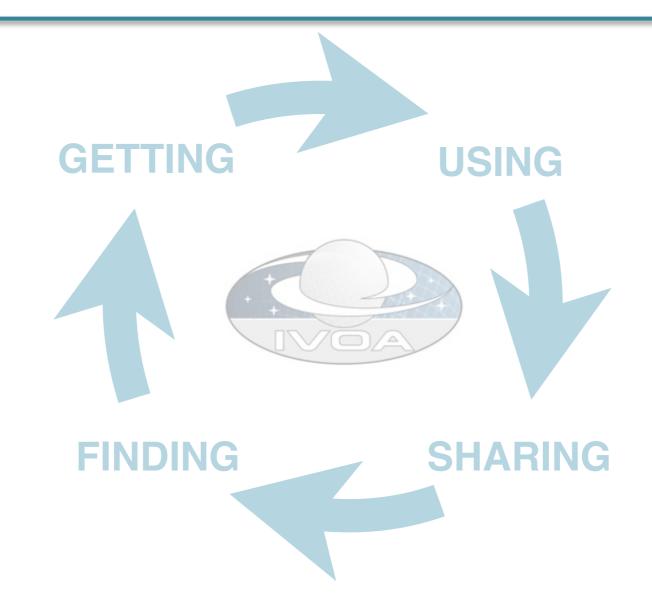
The VO and the IVOA: where?

Existing global framework: populated by major data providers (space and ground based) that is heavily used by the community (e.g. Gaia data access is fully VO)



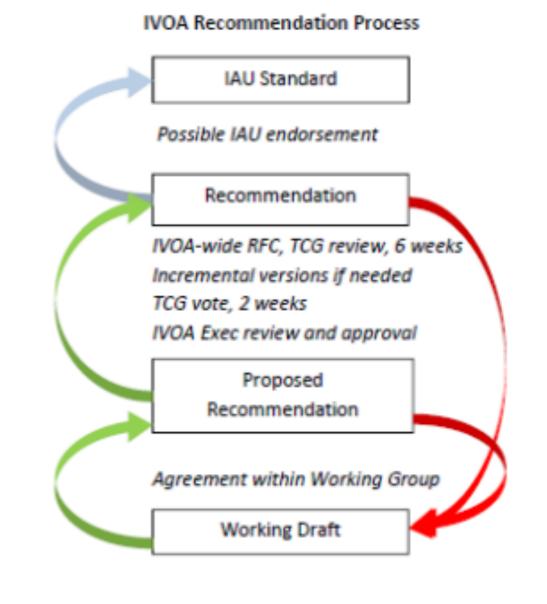
The VO and the IVOA: how?

Through the development and adoption of common standards scientifically driven, as an international community effort where astronomers and software engineers are involved



IVOA development process of standards

- Build IVOA standards to match users needs:
 - Find and report the community needs
 - Find and report gaps in the existing standards
 - Propose new ways to fill the gaps
 - Implement & validate
 - Standardise when consensus is reached



Notes

OK, but where do I start?

 A good starting point to newcomers to the IVOA: the architecture document

https://www.ivoa.net/documents/IVOAArchitecture/20211101/index.html

Things to keep in mind:

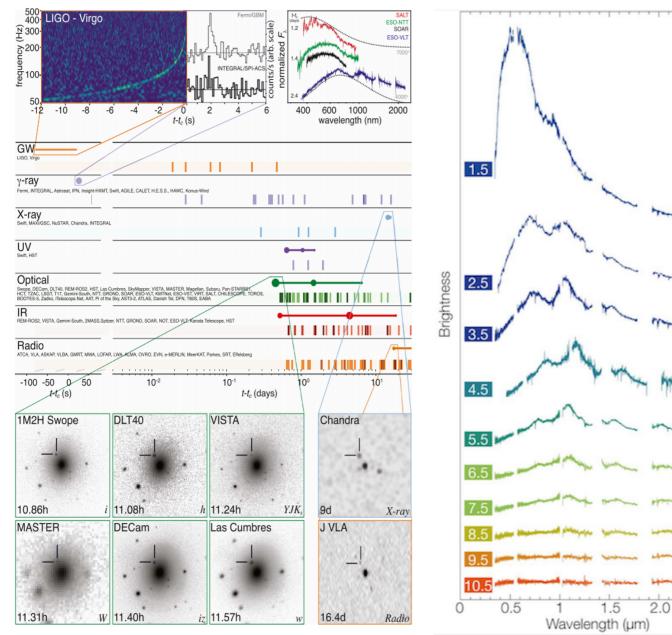
- The IVOA will not answer your scientific questions nor will it ask the questions for you
- The IVOA provides you with common formats and common ways of describing and accessing the data which when adopted will ease your work

Let's see it with an example

Multi-messenger needs?

2.5

HE ASTROPHYSICAL JOURNAL LETTERS, 848:L12 (59pp), 2017 October 20



- Multi-wavelength / messenger approach is needed - different data types
- Follow-up observations and reaction time for that can be crucial - alerts
- Analysis, Visualisation & navigation through the data
- Coordination & transmission of information

The IVOA should match user's needs

Some selected standards

- 1. **VOTable** the format for tabular data for allowing interoperability (coosys, timesys, ucd, utype, VOunits, datalink).
- 2. HiPS more than a format for images tailored for large data volumes
- 3. Search for data:
 - Cone search spatial + temporal search
 - MOC spatial and temporal indexing for large data volumes and more complex areas in the sky
 - TAP + ADQL Table Access Protocol & astronomical data query language
 - **ObsCore & ObsTAP** description of observations
- 4. Planning of observations:
 - ObjVisSAP visibility of object to plan observations
 - **ObsLocTAP** facilitate coordination of observations
 - Facilities / observatory list (under dev.)
- 5. Alerts: VOEvents

6. ... many more! SLAP, SIAP, SSA, Provenance, SAMP... each tailored to specific use cases

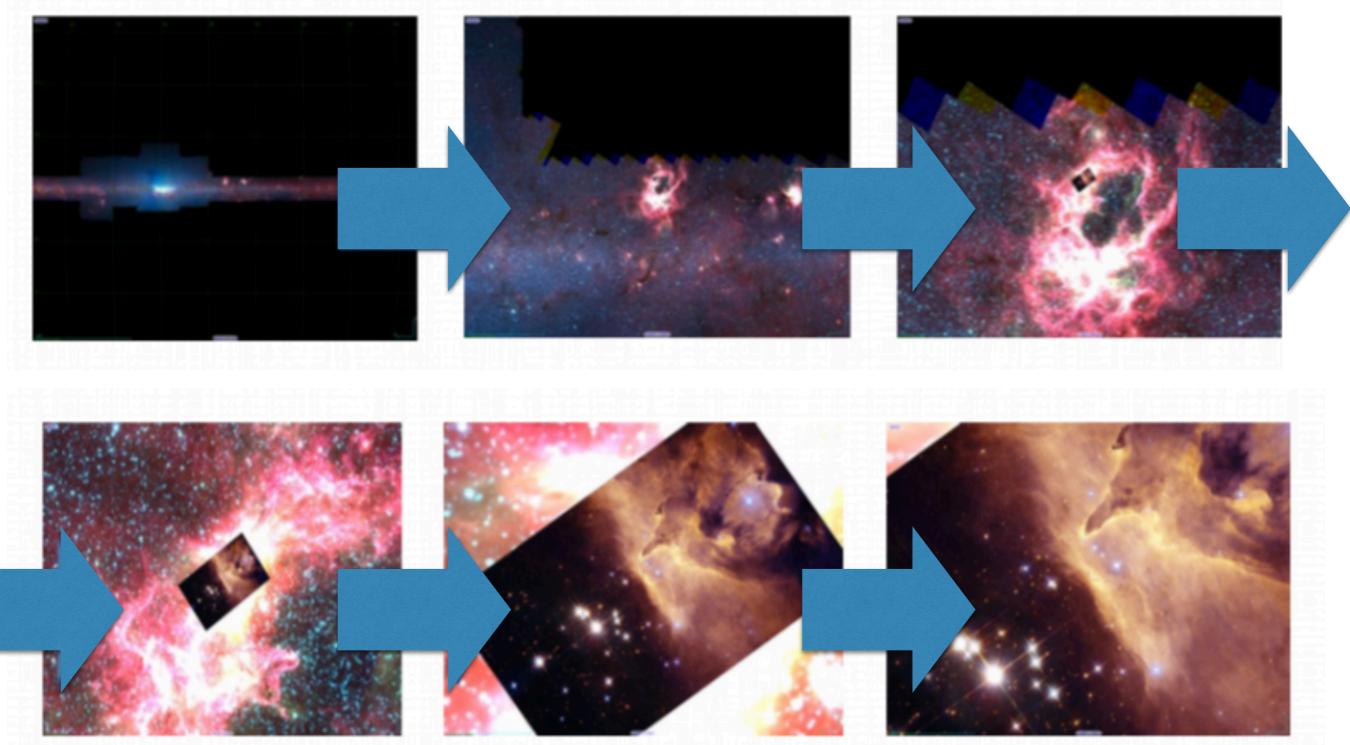
VOTable: format for tabular data

Standardisation of coordinate system annotation (time and space), UCD, utypes, VOUnits, datalink

- **COOSYS** ("ICRS", "eq_FK5",...)
- **TIMESYS** (scale: TT, TAI, ..., refposition: barycenter,... timeorigin: JD, MJD,...)
- Unified Content Descriptor (UCD): controlled vocabulary for describing astronomical data quantities - related to the nature of the values
- UTypes: relationship between the columns and the data model components
- VOUnits: units expressed as a simplified text label (e.g. m.s-2 instead of m s⁻²)
- **Datalink**: links to other associated data

HiPS: Hierarchical image Progressive Survey

• A hierarchical scheme for the description, storage and access of sky survey data (the more you zoom-in the more the details)



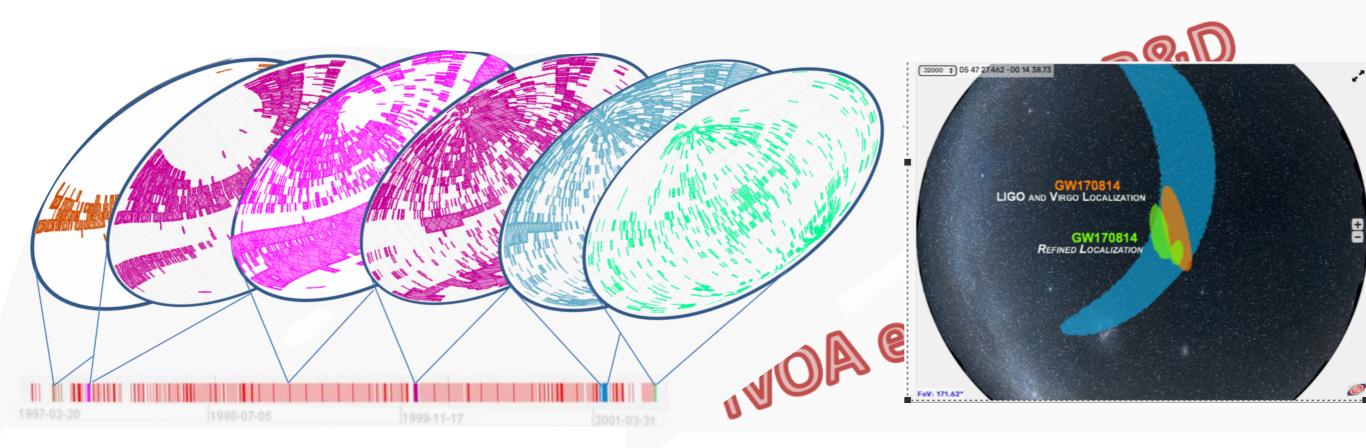
Search: kno

Cone search extension to add a

complicated areas (Simple defin

• MOC : allows to search by tempo

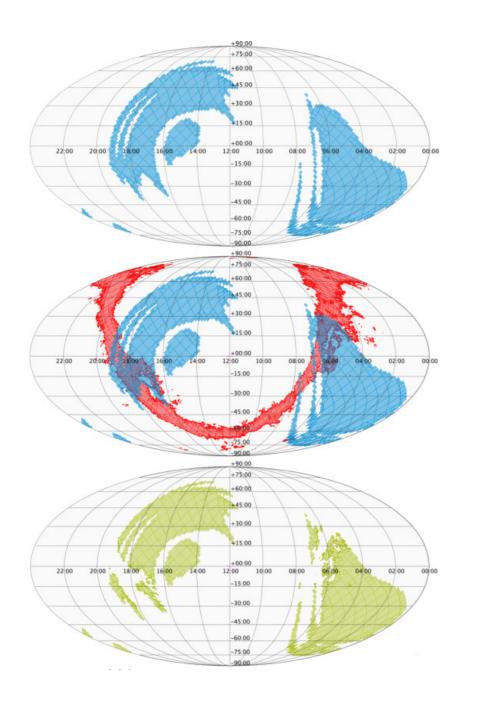
- The STMOC = Space
- Merge together both order to have simultaneously space and time coverage



MOC: Multi-order coverage map

Multi Order Coverage data structure to plan multi-messenger observations

G. Greco, et al. 2022A&C....3900547G



Allows to quickly compute intersections between different surveys

Top: Gravitational-wave sky localisation of GW190425 – 90% credible region

Middle: A selected high Galactic dust extinction region is overlaid. The range of pixel values was set from 0.5 up to the highest value.

Bottom: The resulting sky map defined by the gravitational-wave sky localisation with the dust absorption fields subtracted.

Can be computed for the planning of the observations Library Mocpy a python package affiliated to astropy

TAP & ADQL

- Table Access Protocol (TAP) defines a service protocol for accessing general table data, including astronomical catalogs as well as general database tables. Access is provided for both database and table metadata as well as for actual table data.
- Astronomical Data Query Language (ADQL) Based on Structured Query Language (SQL) with special restrictions and extensions in order to support generic and astronomy specific operations

```
SELECT DISTANCE(
    POINT('ICRS', 266.41683, -29.00781),
    POINT('ICRS', ra, dec)) AS dist, *
FROM gaiaedr3.gaia_source
WHERE 1=CONTAINS(
    POINT('ICRS', 266.41683, -29.00781),
    CIRCLE('ICRS',ra, dec, 0.08333333))
ORDER BY dist ASC
```

Visibility of an object

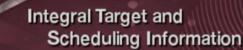
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	with ESO Telescopes > Observing Tools and Services > Calendars and Calculators > Observibility 12 Oct 2017	Target Name M31		(eg; Abell	1750)												
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Observing Facilities Future Facilities and Development Observing with ESO Telescopes Policies and Procedures	See also: Object Observability - Armasses - Daily Almanac - Sky Calendar This tool provides object observability tables based on site, object coordinates and observing period. Times are given for the local time, including daylight saving times when applicable. Select site, object coordinates and observing period; then press Compute.	SIMBAD LOOKUP RES	ULTS:				Staralt is a prog altitude against t a particular nigh the best observi at the bottom of	time for t (Start ng date	a parti rack), o for eac	cular ni or plot l	ight (Star how altitu	ralt), or plo ude change	ot the path es over a y	of your of ear (Sta	robs), or	get a table with	
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Phase 2 Preparation		TARGET DETAILS					Mode										
Phase 3	Site: Paranal Observatory (VLT)	Target Name M31 RA 00:42:44.330	Decimal degrees	Target n or HH:MM:SS.S (eg: 13:30:52.5)	ame or identifier for outp	ut (eg; Abell 1750)	Night	12 Startrac	Octob ck only.	ber	2017	or da	ate when	the loca	al night	starts. Staralt,	
Public Surveys Observing Tools and Services	Dates (yyyy mm dd): From: 2017 10 20 To 2017 11 15			es or DD:MM:SS.S (eg: -01:50:27.0	1)						(Chile)		e with this	formati	•		
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Instrumental Characteristics Archives and Catalogues	RA: 05 23 34.5 Dec: -69 45 22	Select either Revolution Range First Revolutio	on 3369 de	fault is AO17 revolution range: 336	9 to 3551			LA 20.			2005 272						
Calendars and Calculators	compute	Last Revolution 3551									of these:						
Weather Images		Date Range From Date	01 May 2018		May 2018 - 30 Apr 2019				h:mm:s	ss ±dd	:mm:ss						
Astroclimatology Meteo Information	SkyCaic provided by courtesy of John Thorstensen, Dartmouth College. John Thorstensen@dartmouth.edu	To Date	30 Apr 2019				<	name m must be	ust be e in the	a singl	e word w format, o	do not use	different f	formats w		ers. Every entry ent entries. We	
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Different services have different inputs / outputs

Facilitate the work by having some level of standardised input / output

Object Visibility Simple Access Protocol, Aitor Ibarra, Richard Saxton, Jesús Salgado et al. 2020 http://www.ivoa.net/documents/ObjVisSAP/index.html

Coordination of observations



Schedule: All executed Current revolution (1872) Future schedule

Schedule for revolution 1872

(this list is also available in csv-format, click here to download)

Rev	Start time (UTC)	End time (UTC)	Exp. time (s)	Target	Ra (J2000)	Dec (J2000)	Pattern	PI	Propo
1872	2017-10-10 13:29:15	2017-10-10 17:10:51	12600	Gal. Bulge region	17:45:36.00	-28:56:00.0	<u>HEX</u>	Erik Kuulkers	<u>14200</u>
1872	2017-10-10 17:13:34	2017-10-11 07:55:55	50000	Galactic Center	17:52:11.21	-25:21:49.7	<u>5x5 Seq</u>	Joern Wilms	14200
1872	2017-10-11 08:16:46	2017-10-11 11:58:32	12600	Galaxy (I=0, b=0)	17:42:23.76	-29:38:02.4	<u>HEX</u>	Rashid Sunyaev	<u>14200</u>
1872	2017-10-11 12:26:36	2017-10-11 12:56:36	1800	Galaxy (I=0, b=-30)	20:02:16.80	-41:20:31.2	<u>HEX</u>	Rashid Sunyaev	<u>14200</u>
1872	2017-10-11 13:27:21	2017-10-11 14:29:17	3600	Galaxy (I=0, b=-30)	19:59:40.80	-41:05:16.8	<u>HEX</u>	Rashid Sunyaev	<u>14200</u>
1872	2017-10-11 15:00:12	2017-10-11 17:38:07	9000	Galaxy (I=0, b=-30)	19:59:40.80	-41:05:16.8	HEX	Rashid Sunyaev	<u>14200</u>
1872	2017-10-11 18:41:00	2017-10-12 08:01:56	45000	GRS 1915+105	19:15:11.79	+10:56:45.7	<u>5x5 Seq</u>	Jerome Rodriguez	14200

Revolution 1872 to 1872 III Show... show plot

esa

Rashid Sunvaev

Rashid Sunvaev

14200

14200



Short Term Schedule XMM-NEWTON SHORT-TERM SCHEDULE

The Short-term Schedule gives an overview of scheduled observations covering the time range from the past week until the upcoming ~2-4 weeks.

Background: The planning and scheduling procedure is described in Sect. 8.2 of the Policies and Procedures. In addition, the process of scheduling XMM-Newl observations is described in A guided tour to the scheduling of an XMM-Newton orbit.

Description: Each row lists the revolution number (REV#), Observation Identifier (ObsID), target name, pointing coordinates plus position angle (PA), start and stop times, prime instrument, accumulated exposure times (in kiloseconds) for each instrument (without overhead), and name of the Principal investigator (PI). The start and stop times refer to the instrument activities required to perform the observation. The exposure times are accurate times are accurate the same accurate start with the same instrument. Especially for OM, the observation can be split in shorter exposures with different filter/mode. EPIC exposure times in brackets indicate that one or all exposures use the closed filter. Details can be seen when clicking on the ObsID. The row markets the target that is scheduled for the time of the last table update. The creation date is given at the top of the table.

Caveats: The scheduling of an XIMI-Newton revolution may have to be revised (see Sects, 8.2, 8.3, and 5.2.2 of the Policies and Procedures). Contingencies of any type a solar flaring activity may impact at different levels the scheduled programme. The Observation Log Browser can be checked to see what was actually done.

Update frequency: Every 8 hours or when the schedule is updated (new revolution planned or any existing updated). The latest available version can be viewed after clearing the browser buffer from the contents of any previous sessions.

Last updated on: 2017-10-10 12:42:00 UT (Current Rev = 32

Revn #	Obs Id.	Target Name	RA hh:mm:ss	DEC dd:mm:ss	PA ddd.dd	UTC Obs Start yyyy-mm- dd hh:mm:ss	UTC Obs End yyyy-mm- dd hh:mm:ss	Prime Instr.	PN Dur Ks	MOS1 Dur. Ks	MOS2 Dur. Ks	RGS1 Dur. Ks	RGS2 Dur. Ks	OM Dur. Ks	PI
3276	0805150401	ESO 018-G009	08:24:07	-77:46:57	88.63	2017-10-29 19:34:26	2017-10-30 00:54:26	EPIC	16.7	18.1	18.1	18.2	18.2	18.0	Peter Boorman
3276	0801870801	HD 81809	09:27:46	-06:04:17	92.00	2017-10-29 15:00:13	2017-10-29 18:20:13	EPIC	9.5	10.9	10.9	11.0	11.0	10.8	Fabio Favata
3276	0561381201	zeta Puppis	08:03:40	-40:00:36	112.00	2017-10-29 01:21:41	2017-10-29 14:08:21	RGS	44.5	44.9	44.9	45.0	45.0	37.3	Fred Janse XMM- Newton MM
3276	0803950401	SDSS 102714.77+35431	10:27:14	+35:43:17	119.93	2017-10-28 15:44:35	2017-10-28 23:31:15	EPIC	25.5	26.9	26.9	27.0	27.0	26.8	Guido Risaliti
3276	0803240201	J072637.95+394558.0	07:26:37	+39:45:58	91.37	2017-10-28 11:02:32	2017-10-28 14:55:52	EPIC	11.5	12.9	12.9	13.0	13.0	12.9	Nathan Secrest
3275	0801990201	0457-6739	04:57:33	-67:39:06	136.67	2017-10-27 12:22:47	2017-10-28 01:07:47	EPIC	43.4	44.8	44.8	44.9	44.9	43.7	Patrick Kavanagh
3275	0801990401	0449-6903	04:49:34	-69:03:34	138.62	2017-10-26 23:32:47	2017-10-27 12:02:47	EPIC	42.5	43.9	43.9	44.0	44.0	42.8	Patrick Kavanagh
3275	0803952601	SDSS	08:26:19	+31:48:48	101.78	2017-10-26	2017-10-26	EPIC	36.0	37.4	37.4	37.5	37.5	37.3	Guido

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Observing schedules

Short Range Observatory Schedule Download

This is the confirmed schedule of NuSTAR observations. This sequence of observations has been uploaded to the spacecraft and will execute autonomously unless interrupted by a new schedule, Target of Opportunity, or instrument and spacecraft anomalies. This schedule will cover various time ranges depending on the exposure time goal of the observations, but will usually be for a period of at least one week. The times reported here are the start and end of the on-target period (day of year UTC). The estimated exposure time takes into account Earth occutation and the SAA passage time where detector background is increased. The end time of the observation is the start of the silew to the next target. Paese examine the NUSTRA & From Timeline (ATT) for the log of past observations.

Table Header Explanations

obs_start	obs_end	sequenceID	Name	J2000_RA	J2000_Dec	Exp	Notes
2017:281:19:05:02	2017:283:00:30:00	90201021006	Kepler	262.671620	-21.491957	60.6	DDT
2017:283:01:11:23	2017:283:02:40:00	90311211001	Sol_17282_AR2683_POS11	195.15715	-6.38520	3.4	ToO
2017:283:02:40:32	2017:283:04:20:00	90311212001	Sol_17282_AR2683_POS12	195.21879	-6.41062	3.4	ToO
2017:283:04:20:32	2017:283:05:50:00	90311213001	Sol_17282_AR2683_POS13	195.28046	-6.43604	3.4	ToO
2017:283:06:55:11	2017:284:09:20:00	60376001002	2MASXJ19301380p3410495	292.557500	34.180500	55.3	Extragalactic Legacy Survey
2017:284:09:45:09	2017:284:20:35:00	60360008002	SDSSJ152132d21p391206d9	230.3874232	39.2007671	22.0	Extragalactic Legacy Survey
2017:284:21:10:03	2017:285:21:00:00	90301320002	NGC_6440	267.218083	-20.358944	49.5	ToO
2017:285:21:20:06	2017:286:08:20:00	30302020004	GRS_1915p105	288.79813	10.94578	21.9	(2/4) coordinated with XMM and VLT
2017:286:08:35:06	2017:286:19:30:00	60160701002	2MASXJ18560128p1538059	284.00210000	15.63200000	23.3	BAT AGN
2017:286:20:05:11	2017:287:15:05:00	60376007002	UGC06728	176.316800	79.681500	61.4	Extragalactic Legacy Survey
2017:287:15:50:11	2017:288:03:20:00	60368001002	NGC_1144	43.80083	-0.18361	22.0	
2017:288:04:05:09	2017:288:23:00:00	60301004002	ESO_103m35	279.58458	-65.4275	50.3	
2017:288:23:30:08	2017:290:05:45:00	30301026002	AX_J1841d0m0536	280.25179	-5.59625	59.7	phase constrained
2017:290:06:00:04	2017:290:17:00:00	60160670002	2E1739d1m1210	265.47600000	-12.19700000	23.5	BAT AGN
2017:290:17:15:01	2017:291:04:20:00	30363001002	GX_3p1	266.98333	-26.56361	21.8	

Long Range Observatory Schedule Download

This is the latest NuSTAR long-term schedule. Observations have been sorted into one-week intervals, taking into account Sun, Moon, required exposure time, and other constraints. So the date is the Monday of the week in which the observation is schedule to begin. E.g. An observation with a date **2017-12-18** this table is scheduled to have the observation **starting** sometime between **2017-12-18**

0000Z and 2017-12-25 0000Z.

Currently the schedule is driven by the large number of observations coordinated with other observatories and the need to complete the NuSTAR Guest Observer programs. The exposure goal for targets allotted within one week may appear to fill more then the available NuSTAR exposure time in that week (average is 30) as per week) but many observations start in one week and complete in the following week.

Targets of opportunity and any instrument or spacecraft anomalies may also cause the observing times of targets to shift. This long-ten schedule is our present estimate of the future order of observations. Please be aware of the uncertainties.

ToO = Target of Opportunity DDT = Directors Discretionary Time NO3 = NuSTAR GO cycle-3 I15 = INTEGRAL GO cycle-15

x16 = XMM-Newton GO cycle-16
 C18 = Chandra GO cycle-18
 ELS/GLS = Extragalactic/Galactic legacy surveys

		Schedulin Begin UT	ng U End	UT	SU Id	Principal Investigat	Exp #	Target	Science Instrume	Mode	Apertures		Exposure Time(sec)		AL	E
	2017.288	23:00:00	23:3	5:07	14835Z1	Lockwood	Z1-001	DARK	STIS/MA2	TIME-T	F28X50LP	MIRVIS	1300.00			
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-001	WASP-69	COS/NUV	ACQ/SE	F28X50LP PSA PSA PSA PSA PSA PSA PSA PSA F28X50LP F28X50LP	G230L	12.00			
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-002	WASP-69	COS/NUV	ACQ/PE	PSA	G230L	12.00	35	02	0
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-003	WASP-69	COS/NUV	ACQ/PE	PSA	G230L	12.00	35	03	0
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-004	WASP-69	COS/FUV	TIME-T	PSA	G130M	1917.00	35	05	0
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-005	WASP-69	COS/FUV	TIME-T	PSA	G130M G130M	2706.00	35	07	0
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-006	WASP-69	COS/FUV	TIME-T	PSA	G130M	2706.00	35	09	0
	2017,288	23:14:45	06:3	0:55	1476735	Sing	35-007	WASP-69	COS/FUV	TIME-T	PSA	G130M	2706.00	35	0 B	0
	2017.288	23:14:45	06:3	0:55	1476735	Sing	35-008	WASP-69	COS/FUV	TIME-T	PSA	G130M MIRVIS	2706.00	35	0D	0
	2017.289	00:00:00	00:2	8:32	14819JF	Riley	JF-001	DARK	STIS/CCD	ACCUM	F28X50LP	MIRVIS	1100.00	JF	01	0
	2017.289	00:00:00	00:2	8:32	14819JF	Riley	JF-002	DARK	STIS/CCD	ACCUM	F28X50LP	MIRVIS	60.00	JF	01	0
1	2017,289	00:00:00	00:2	8:32	14819JF	Riley	JF-003	DARK			F28X50LP	MIRVIS	60.00			
1	2017,289	00:00:00	00:4	6:10	145333B	Bourgue	3B-001	DARK-NM			UVIS	F373N	900.00	3B	01	0
		00:00:00									UVIS	F373N	900.00	3P	02	ő
1		00:39:46									F28X50LP	F373N MIRVIS	1100.00	JG	01	0
						Riley					F28X50LP	MTRVIS	60.00	JG	01	ő
	2017.299	00:39:46	01:0	8:18	14819.70	Biley	JG-003	DARK			F28X50LP	MTRVIS	60.00	JC	01	õ
						Bourgue					UVIS	MIRVIS MIRVIS F467M F467M	900.00	30	01	0
		00:46:10							WFC3/UVI		INTE	P467M	900.00	30	02	0
						Riley					F28X50LP	MTDUTS	0.00	917	01	0
						Riley					F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6.24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	PROVENTION	MIRVIS	0.00	90	01	0
	2017.209	01:27:12	01:5	6.24	1402190	Riley Riley Riley Riley	90-001	DIAD	STIS/CCD	ACCUM	PROVENTION	MIRVIS MIRVIS MIRVIS MIRVIS	0.00	90	01	0
	2017.209	01:27:12	01:5	0:24	1402190	Riley	90-001	DIAD	STIS/CCD	ACCOM	FZGADULF	MIRVID	0.00	90	01	
	2017.289	01127112	0115	6124	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X5ULP	MIRVIS	0.00	90	01	0
	2017.289	01.27112	0115	6.24	1402190	Riley	90-001	DIAD	STIS/CCD	ACCUM	PROVENT	MIRVID	0.00	90	01	2
	2017.289	01127112	0115	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	0115	0:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	0:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	FZSX5ULP	MIRVIS MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-001	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley Riley Riley Riley Riley Riley Riley Riley Riley Riley Riley Colimowski	90-002	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-002	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:27:12	01:5	6:24	1482190	Riley	90-002	BIAS	STIS/CCD	ACCUM	F28X50LP	MIRVIS	0.00			
	2017.289	01:40:00	02:0	9:22	14518F0	Golimowski	F0-001	BIAS	ACS/WFC	ACCUM	WFC	F502N	0.00	FO	01	0
												F660N				
	2017.289	01:40:00	02:0	9:22	14518F0	Golimowski	F0-002	DARK	ACS/WFC	ACCUM	WFC	F502N	1000.50	FO	01	0
												F660N				
	2017.289	02:09:22	02:3	8:56	14518F1	Golimowski	F1-001	DARK	ACS/WFC	ACCUM		F502N	0.50	F1	01	0
												F660N				
	2017.289	02:09:22	02:3	8:56	14518F1	Golimowski	F1-002	DARK	ACS/WFC	ACCUM		F502N	1000.50	F1	01	0
												F660N				

9-Oct-2017 18:48:29 --- Preliminary HST Observing Timeline Report for SMS: 172888A4 ---SMS Start: 2017.288:22:10:00 (15-OCT-2017 22:10:00), End: 2017.296:00:00:00 (23-OCT-2017 00:00:00

9-Oct-2017 18:48:29 --- Preliminary HST Observing Timeline Report for SMS: 172888A4 ---SMS Start: 2017.288:22:10:00 (15-OCT-2017 22:10:00), End: 2017.296:00:00:00 (23-OCT-2017 00:00:00)

	Scheduling Unit		Principal			Science			Spectral	Exposure			
	Begin UT End UT	SU Id	Investigat	Exp #	Target	Instrume	Mode	Apertures	Elements	Time(sec)	OB	AL	, E.
													-
2017.289	02:38:56 03:08:18	14518F2	Golimowski	F2-001	BIAS	ACS/WFC	ACCUM	WFC	F502N	0.00	F2	01	0
									F660N				
2017.289	02:38:56 03:08:18	14518F2	Golimowski	F2-002	DARK	ACS/WFC	ACCUM	WFC	F502N	1000.50	F2	01	0
									F660N				
2017.289	03:10:31 03:40:05	14518F3	Golimowski	F3-001	DARK	ACS/WFC	ACCUM	WFC	F502N	0.50	F3	01	0
									F660N				
2017.289	03:10:31 03:40:05	14518F3	Golimowski	F3-002	DARK	ACS/WFC	ACCUM	WEC	F502N	1000.50	F3	01	0
									F660N				
2017.289	03:46:00 04:48:35	1483522	Lockwood	Z2-001	DARK	STIS/MA2	TIME-T	F28X50LP	MIRVIS	1300.00	Z 2	01	0
2017.289	03:49:34 05:01:49	1454639	Shanahan	39-001	TUNGSTEN	WFC3/UVI	ACCUM	UVIS1-M512-S	F645N	60.00	39	01	0
2017,289	03:49:34 05:01:49	1454639	Shanahan	39-002	TUNGSTEN	WFC3/UVI	ACCUM	UVIS	F814W	2.00	39	01	0
2017.289	03:49:34 05:01:49	1454639	Shanahan	39-003	TUNGSTEN	WFC3/UVI	ACCUM	UVIS	F438W	360.00	39	01	0
	03:49:34 05:01:49	1454639	Shanahan		TUNGSTEN	WEC3/UVT		UVIS	F438W		39		

What object has been (or will be) observed when and in which wavelength?

Observation Locator Table Access Protocol, Aitor Ibarra, Jesús Salgado et al. 2021 http://www.ivoa.net/documents/ObsLocTAP/^{VOA}

Publishing your data in the VO

Do you have data and you would like to publish it?

https://wiki.ivoa.net/twiki/bin/view/IVOA/PublishingInTheVO

Depending of your needs and your technical experience, there are several ways to publish you data into the VO:

- if you have very little technical expertise, you might want to contact your national VO projects who might already have the infrastructure to host your data and make it available to the VO. See the list of national VO projects and the selection of open data centers offering publication as a service
- if you would like to find your VO services in client applications, please publish your service in a VO Registry.
- if you have some technical expertise but don't want to spend time to development, you might want to use directly some existing VO Publishing toolkits.
- if you prefer VO interfaces to your data yourself, you might find some useful VO software tools and libraries. A good start would be to determine what type of data you want to publish (images, catalogues, spectra, ...). Reading through the IVOA Architecture document might also be useful before going into the more details of the specific IVOA standards that you might need to use

Access data in the VO

There are different ways to access the VO, eg via Aladin, Topcat, python

Page Contents

- Introduction
- Installation
 - Source Installation
- Requirements
- Getting started
 - Data Access
- Registry search
- Using **pyvo**

PyVO

Introduction

This is the documentation for PyVO, an affiliated package for the astropy package.

PyVO lets you find and retrieve astronomical data available from archives that support standard IVOA virtual observatory service protocols.

- Table Access Protocol (TAP) accessing source catalogs using sql-ish queries.
- Simple Image Access (SIA) finding images in an archive.
- Simple Spectral Access (SSA) finding spectra in an archive.
- Simple Cone Search (SCS) for positional searching a source catalog or an observation log.
- Simple Line Access (SLAP) finding data about spectral lines, including their rest frequencies.

Note

If you need to access data which is not available via the Virtual Observatory standards, try the astropy affiliated package **astroquery** (and, of course, ask the data providers to do the right thing and use the proper standards for their publication).

Lots of tutorials on how to access data

https://nasa-navo.github.io/navo-workshop/

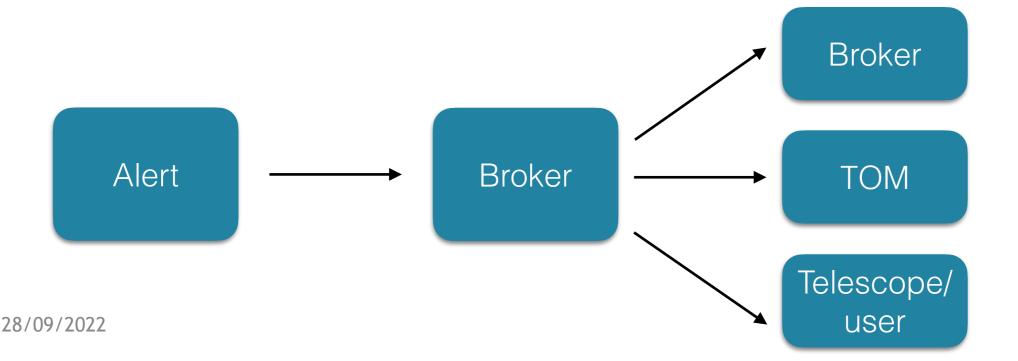
https://github.com/cds-astro/tutorials

https://github.com/cds-astro/mocpy/tree/master/notebooks

□ Alerts

1. VOEvent (REC):

- Content: a core data model for alerts
 - Who, What, WhereWhen, Why, and How
 - Defined by the community (e.g. FRB)
- A format: a serialization of that data model (xml)
- 2. VTP VOEvent Transport protocol (REC):
 - a communication protocol for publishing, and subscribing to alerts



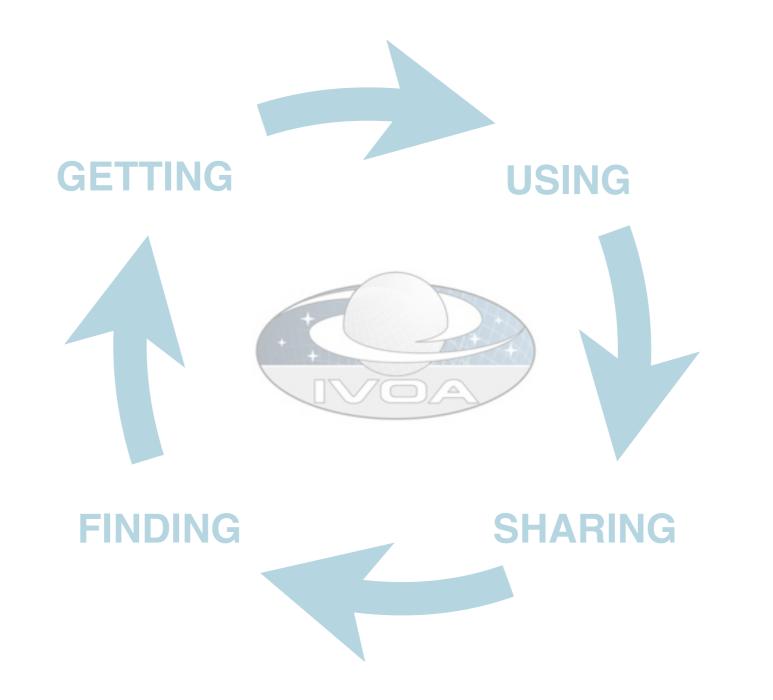
VOEvent: Sky Event Reporting Metadata

- "Defines the content and meaning of a standard information packet for representing, transmitting, publishing and archiving information about a transient celestial event, with the implication that timely follow-up is of interest"
 - Who: Identification of scientifically responsible Author
 - What: Event Characterization modeled by the Author
 - WhereWhen: Space-Time Coordinates of the event
 - How: Instrument Configuration
 - Why: Initial Scientific Assessment
 - **Citations**: Follow-up Observations
 - **Description**: Human Oriented Content
 - Reference: External Content

Open questions

- Are the existing standards enough or do we need changes / new developments?
- In the context of alerts:
 - VOEvent : model (content) + serialization (format)
 - VTP : transport protocol works for low rates (10 Hz) doesn't scale well for very high rates (>10³Hz)
- Is this still what the community needs?
- New technologies and serialisations are used, in particular GCN adopting and moving towards Kafka / Avro (ZTF, Rubin, LVK,...)
- What is the impact on the existing VOEvent infrastructure & community?
- How to find who distributes alerts in the VO land? Register in the registry? Or the community knows where to find alerts?
- Is there a need to standardize event types (Who, What, WhereWhen, Why, and How) the core model?
 - Extending the work on FRBs to cover other Types?
 - If so, smaller working groups within transient sub-fields develop a proposed standard to be iterated upon within the wider user base. Once a standard is developed, it can be submitted to the IVOA and published as an IVOA Note.
 - How do we store information in a way that it can be easily extracted and interpreted later? What types of queries do people want to perform on archived VOEvents? How do we make VOEvent streams findable

□ Thanks!



Some useful links

- <u>https://www.ivoa.net</u>
- Docs : <u>https://www.ivoa.net/documents/</u>
- GitHub : <u>https://github.com/ivoa</u>
- Mailing list : <u>https://www.ivoa.net/members/index.html</u>
- Architecture: <u>https://www.ivoa.net/documents/IVOAArchitecture/</u> 20211101/index.html
- Slack: <u>https://join.slack.com/t/ivoa/shared_invite/zt-17kd0v93b-b32~KReWd1T96gDyYFDLPQ</u>